

Progress on aMCfast

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ERC miniworkshop

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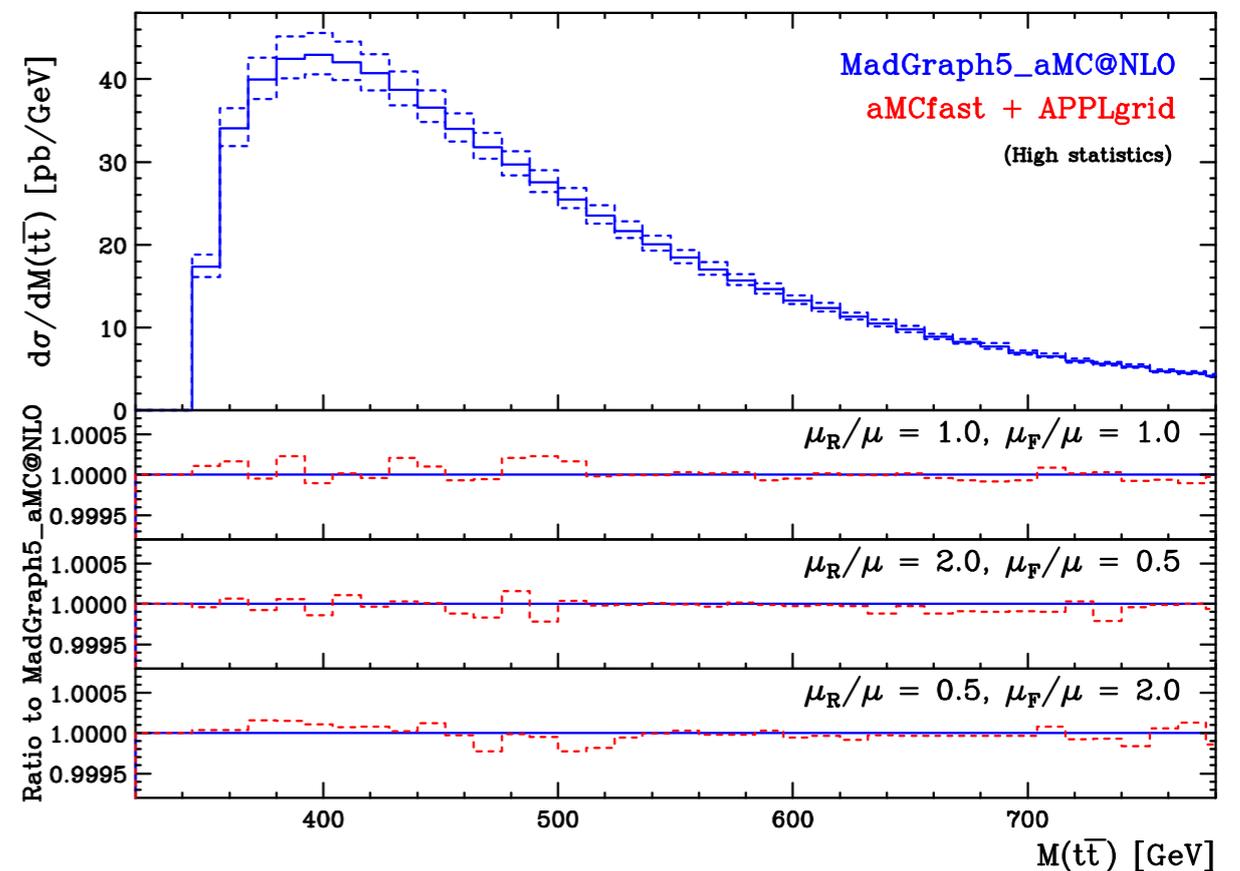
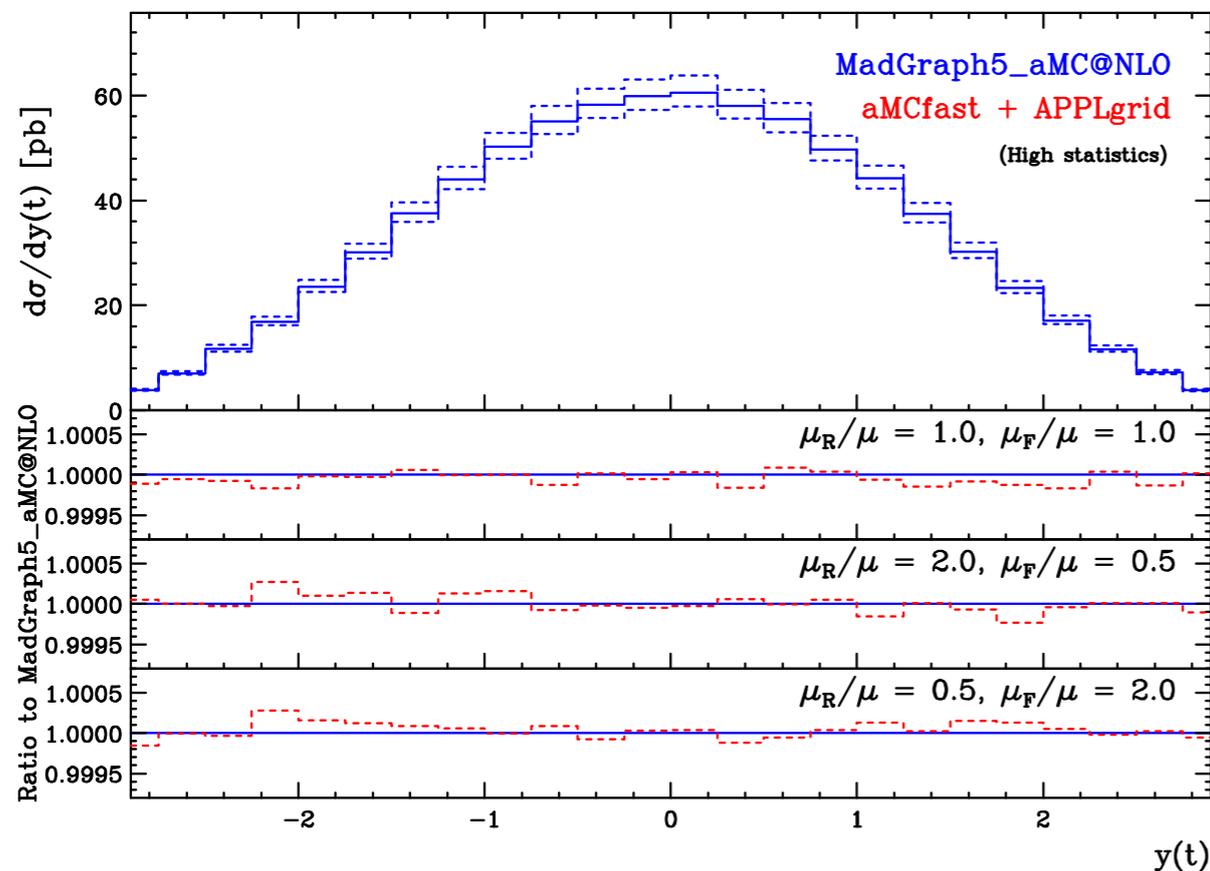
Recap on aMCfast

- The inclusion of hadron-collider data into fits of **parton distribution functions** (PDFs) is a crucial requirements for precision physics at the LHC.
- Presently, hadronic NLO(+PS) calculations are too **time-consuming** to be directly employed in a PDF fit.
- Need for **fast interfaces** to such calculations (*e.g.* APPLgrid).
- **aMCfast** provides an interface that bridges:
 - **MadGraph5_aMC@NLO** [[arXiv:1405.0301](#)]
 - an **automated** cross section calculator of LO and NLO cross sections, with and without matching to parton showers.
 - **APPLgrid** [[arXiv:0911.2985](#)]
 - a framework for the fast computation of cross sections \Rightarrow APPLgrid interpolation grids.
- The **full automation** of MG5_aMC@NLO allows one to produce interpolation grids for any process without any additional work (like *e.g.* for MCFM).

The aMCfast Interface

The NLO case

- At fixed order the the fast interpolation technique is **well-established** and presents no conceptual problems:
 - exact **independence** of the interpolation grids from PDFs and α_s ,
 - commonly used in most of the modern PDF fits.
- The aMCfast interface for NLO computation provides a high accuracy:



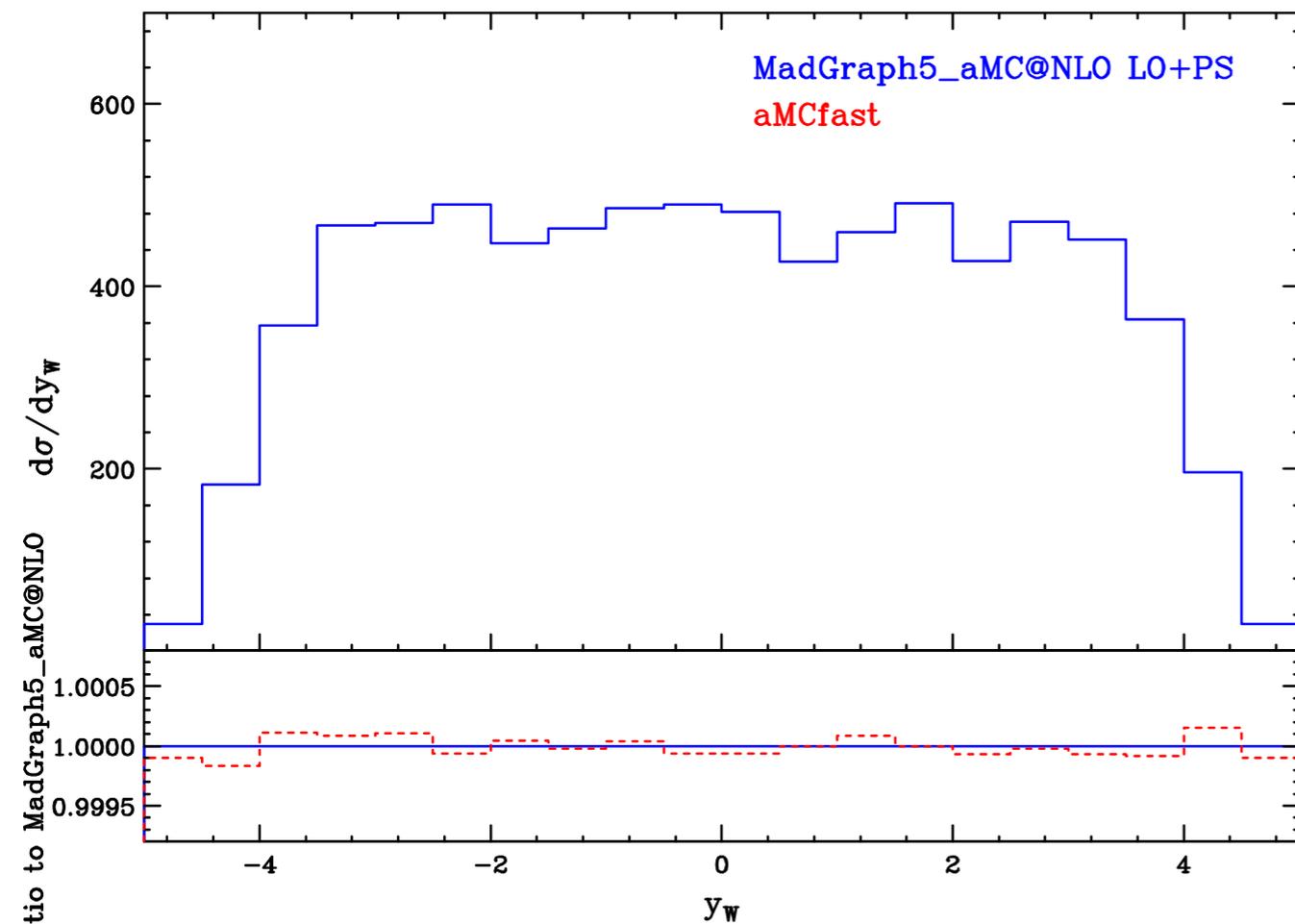
- and has been already employed in e few public studies.

The aMCfast Interface

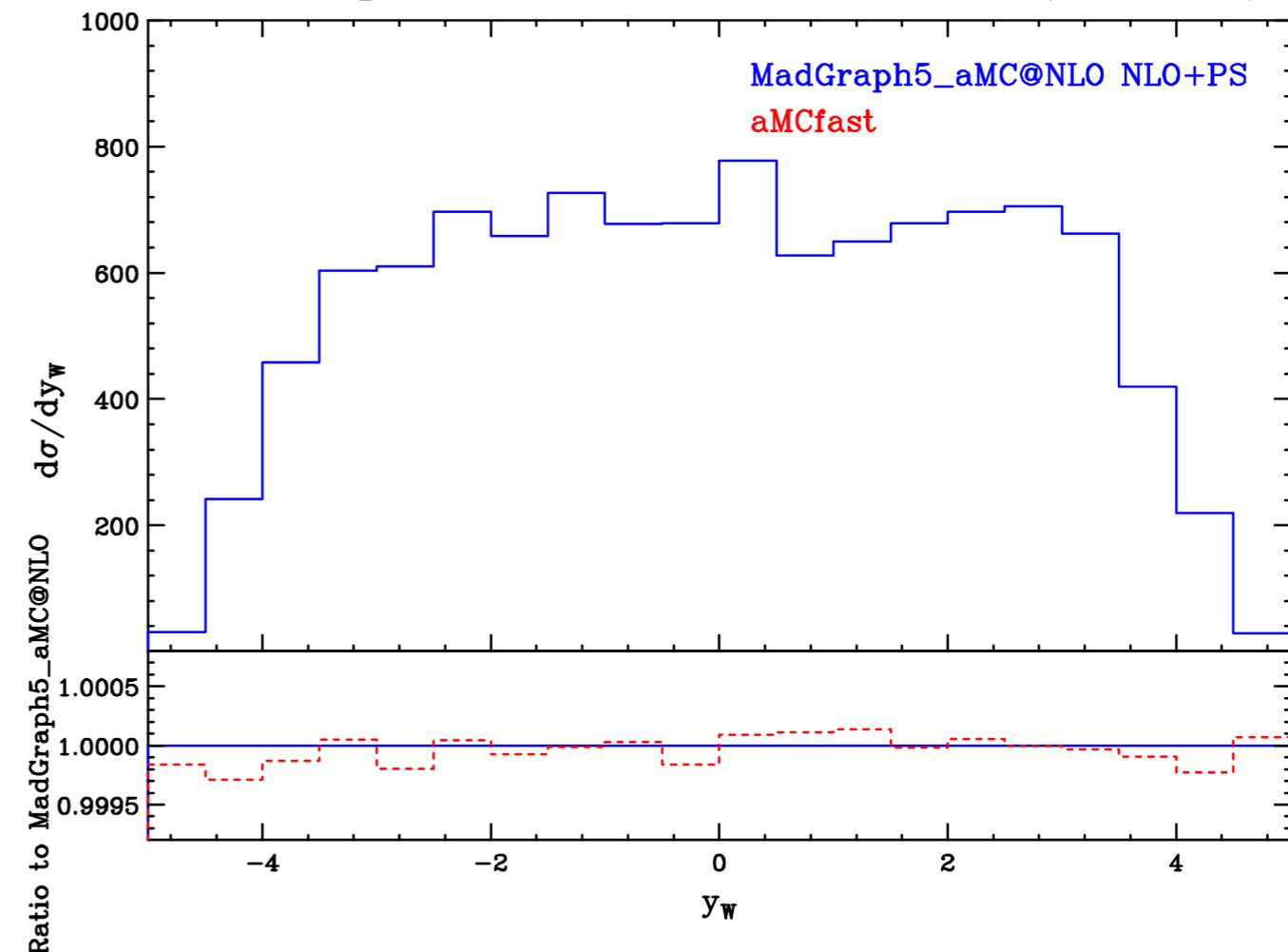
The NLO + PS Case

- Technically speaking, the aMCfast interface at NLO+PS mode of has already been implemented and tested for **all** PSs interfaced to MG5_aMC@NLO.
- W rapidity for $e^+ \nu$ production at LO and NLO + Herwig6:

MadGraph5_aMC@NLO vs. aMCfast (LO+PS)



MadGraph5_aMC@NLO vs. aMCfast (NLO+PS)



- Perfect agreement between reference and reconstructed histograms also in the low statistics regime, as in the fixed-order case.

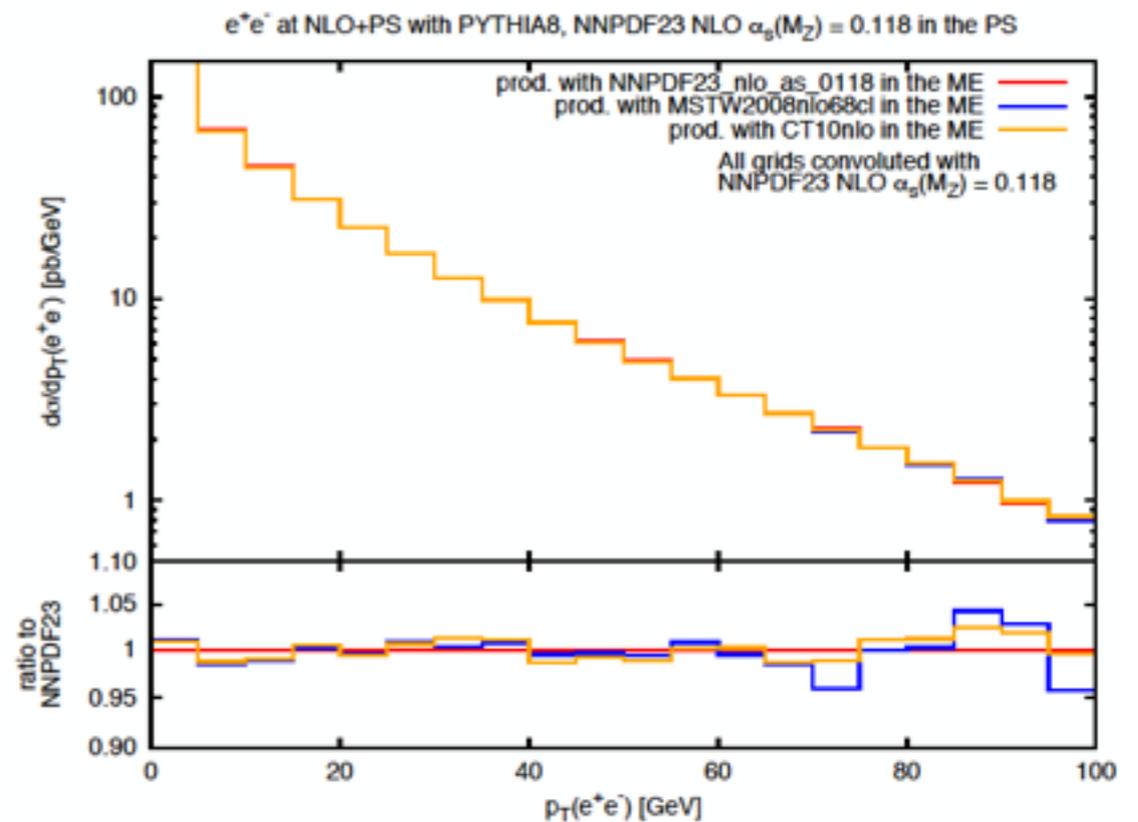
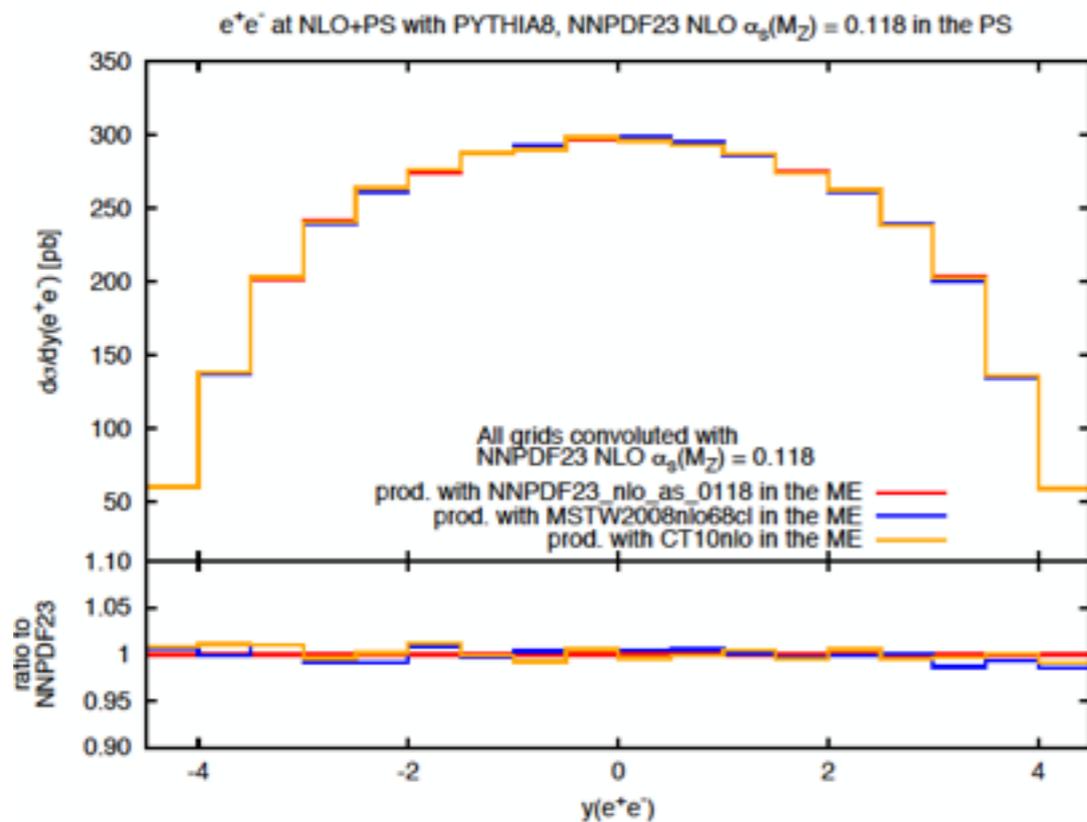
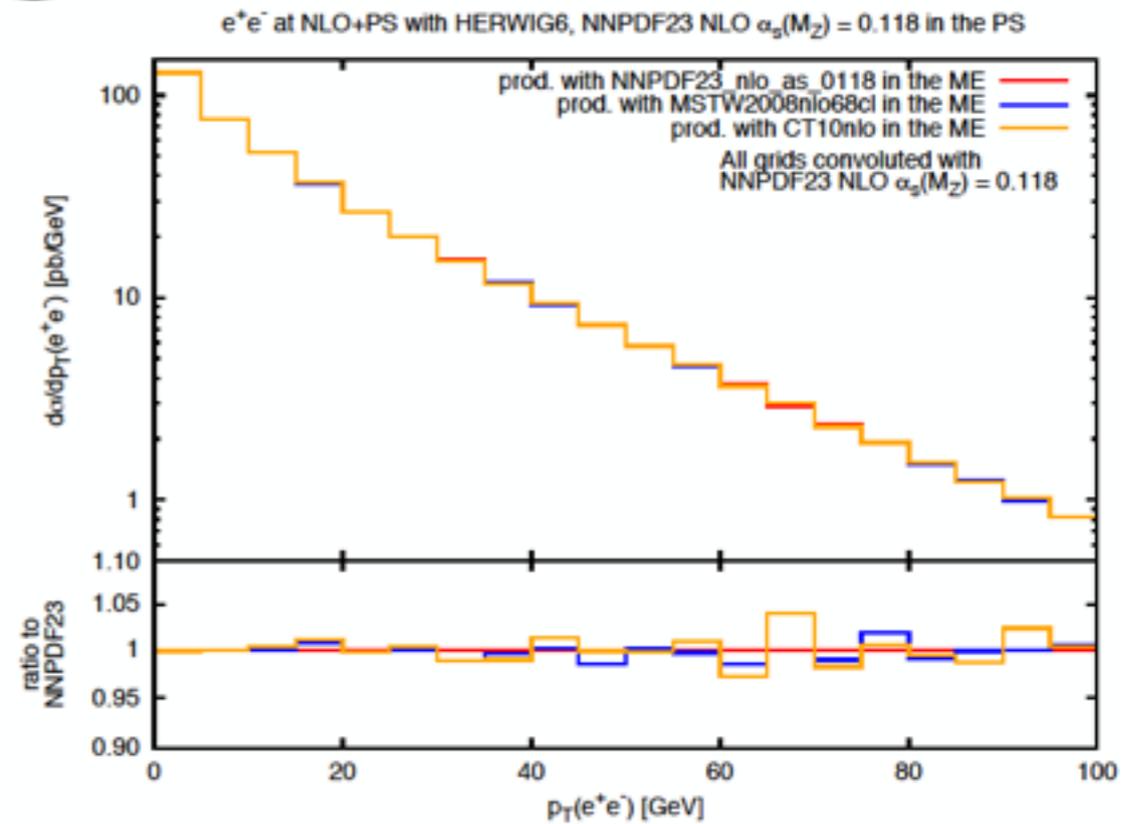
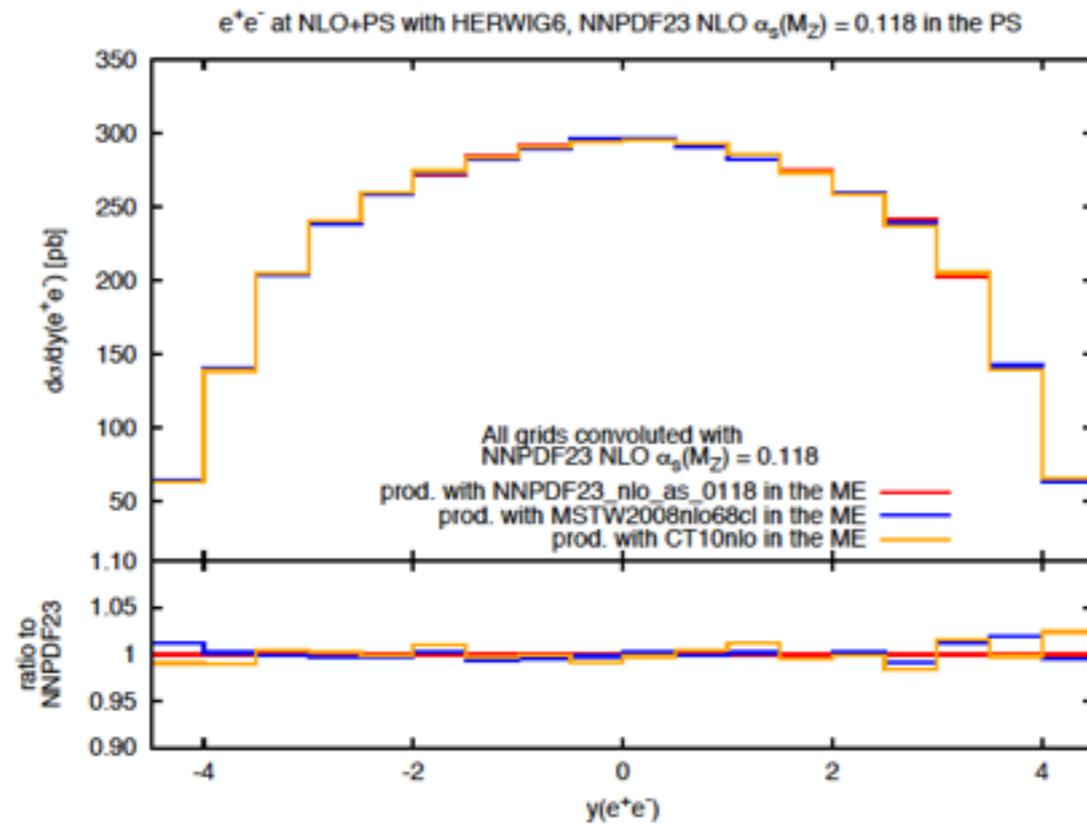
The aMCfast Interface

The NLO + PS Case

- The production of interpolation grids in the presence of PS poses more **conceptual questions** as compared to the fixed-order case.
- There are **two main issues**:
 - 1) **Dependence on PDFs** of the **PS evolution** as a results of different kinematic configurations at the **matrix element** (ME) level when the latter is computed with different PDF sets cannot be removed.
 - 2) **Dependence on PDFs** of the **initial-state PS** (IS-PS) cannot be disentangled:
 - expected to be small as it appears as a ratio of PDFs at the same x but different Q^2 .
- In order for such grids to be safely used in a PDF fit one needs to check **explicitly** that the dependence on the PDFs used for the production is mild:
 - process/observable dependent,
 - parton shower dependent.

The aMCfast Interface

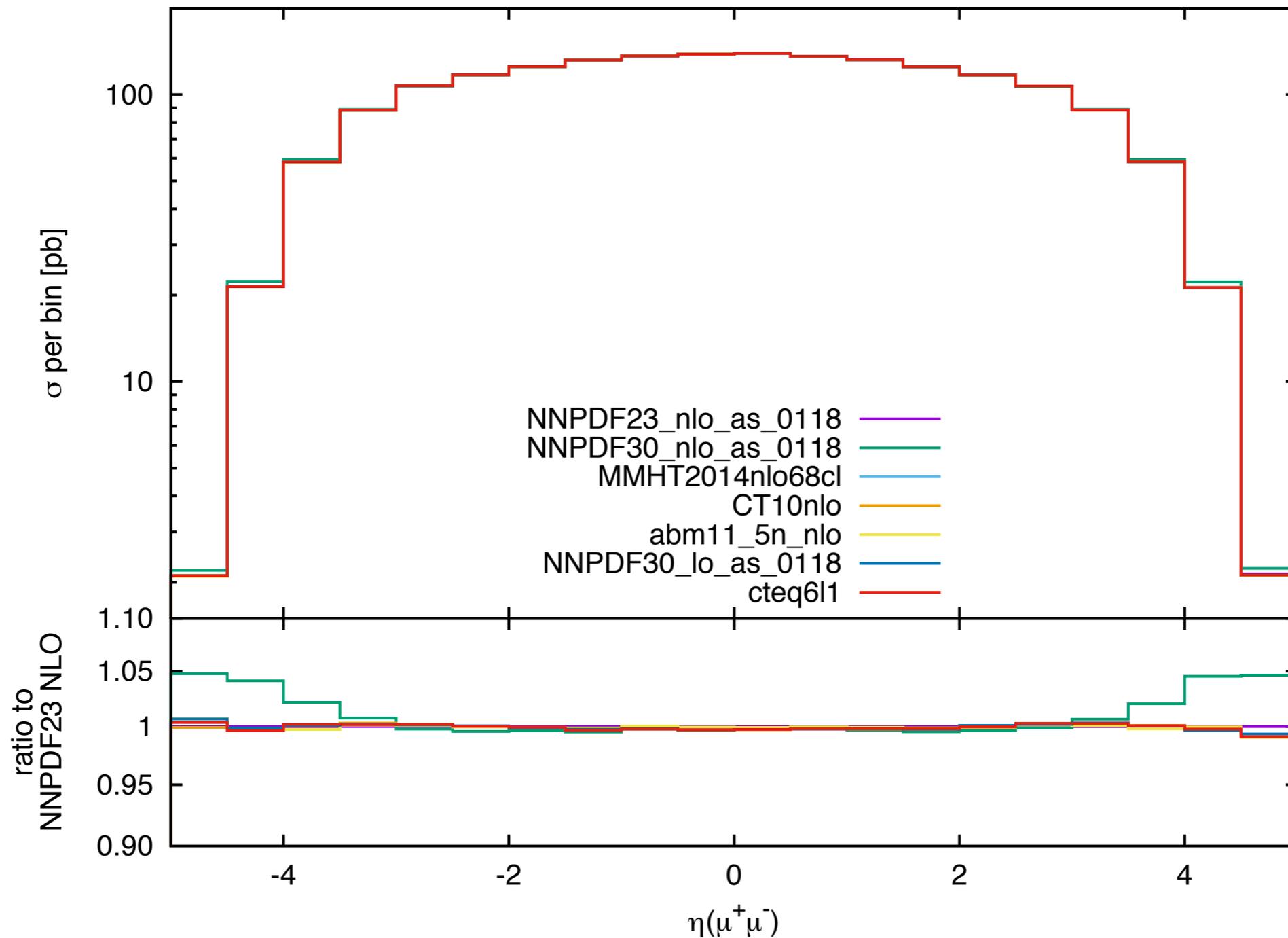
The NLO + PS Case: Dependence on the ME



The aMCfast Interface

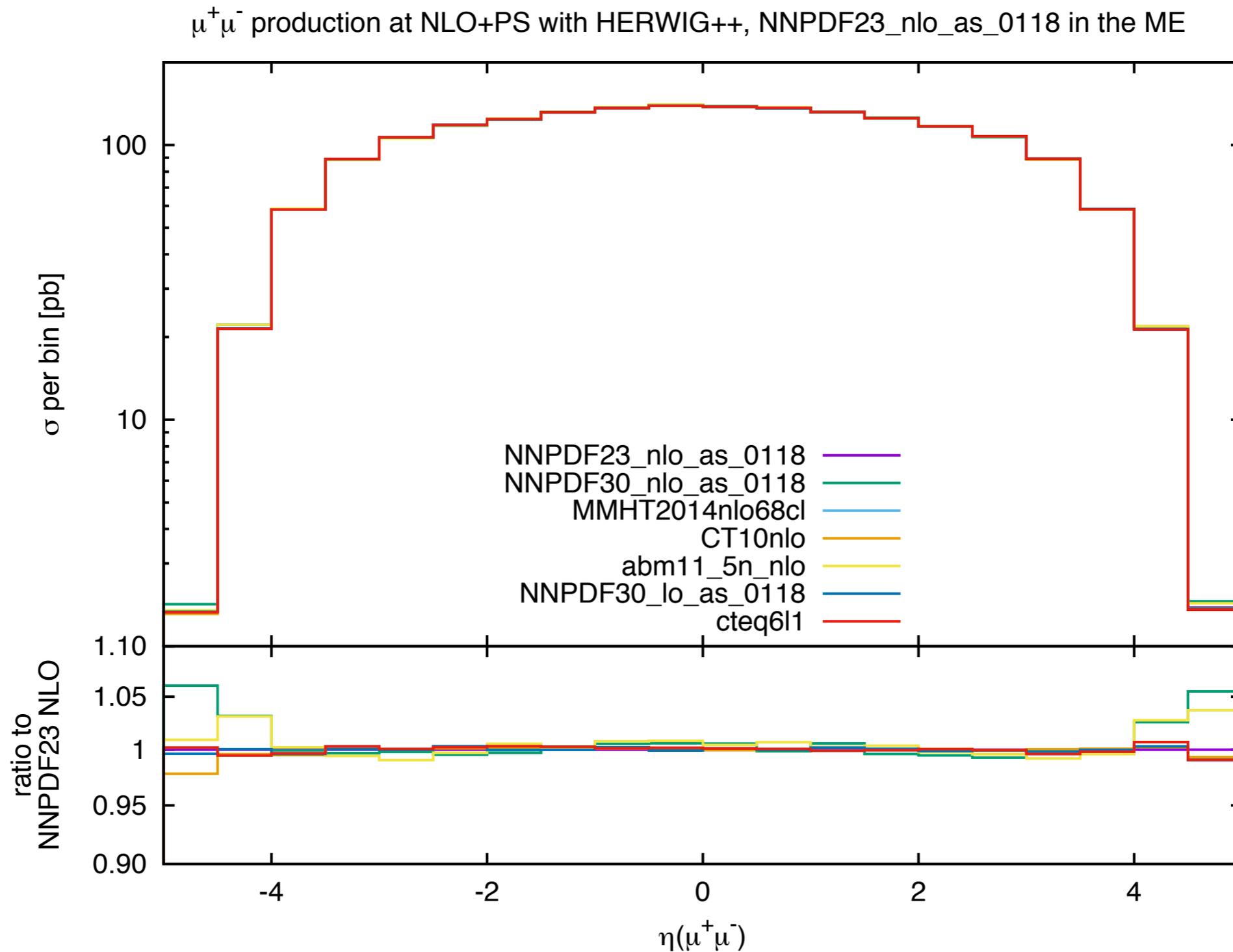
The NLO + PS Case: Dependence on the IS-PS

$\mu^+\mu^-$ production at NLO+PS with HERWIG6, NNPDF23_nlo_as_0118 in the ME



The aMCfast Interface

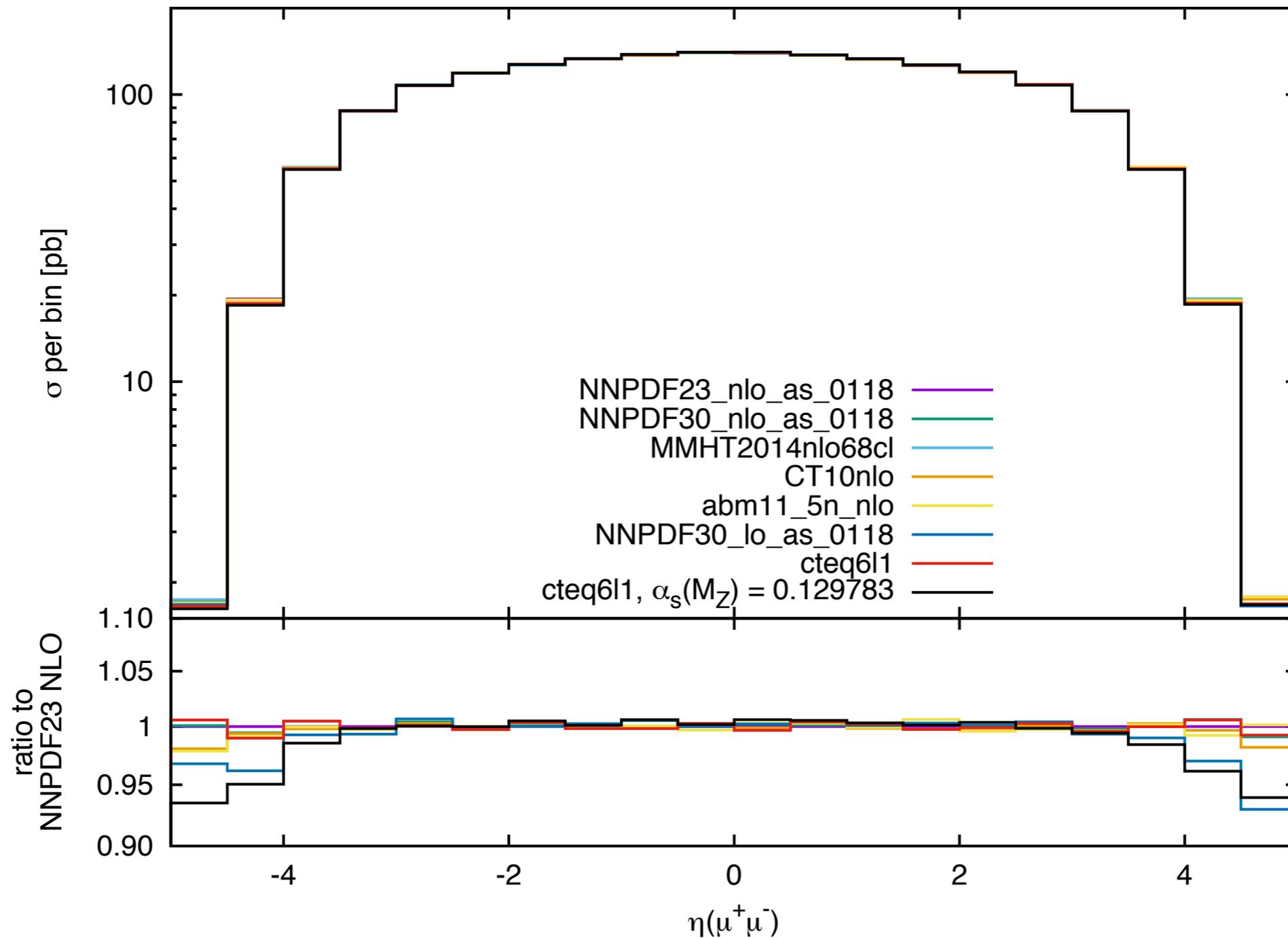
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The aMCfast Interface

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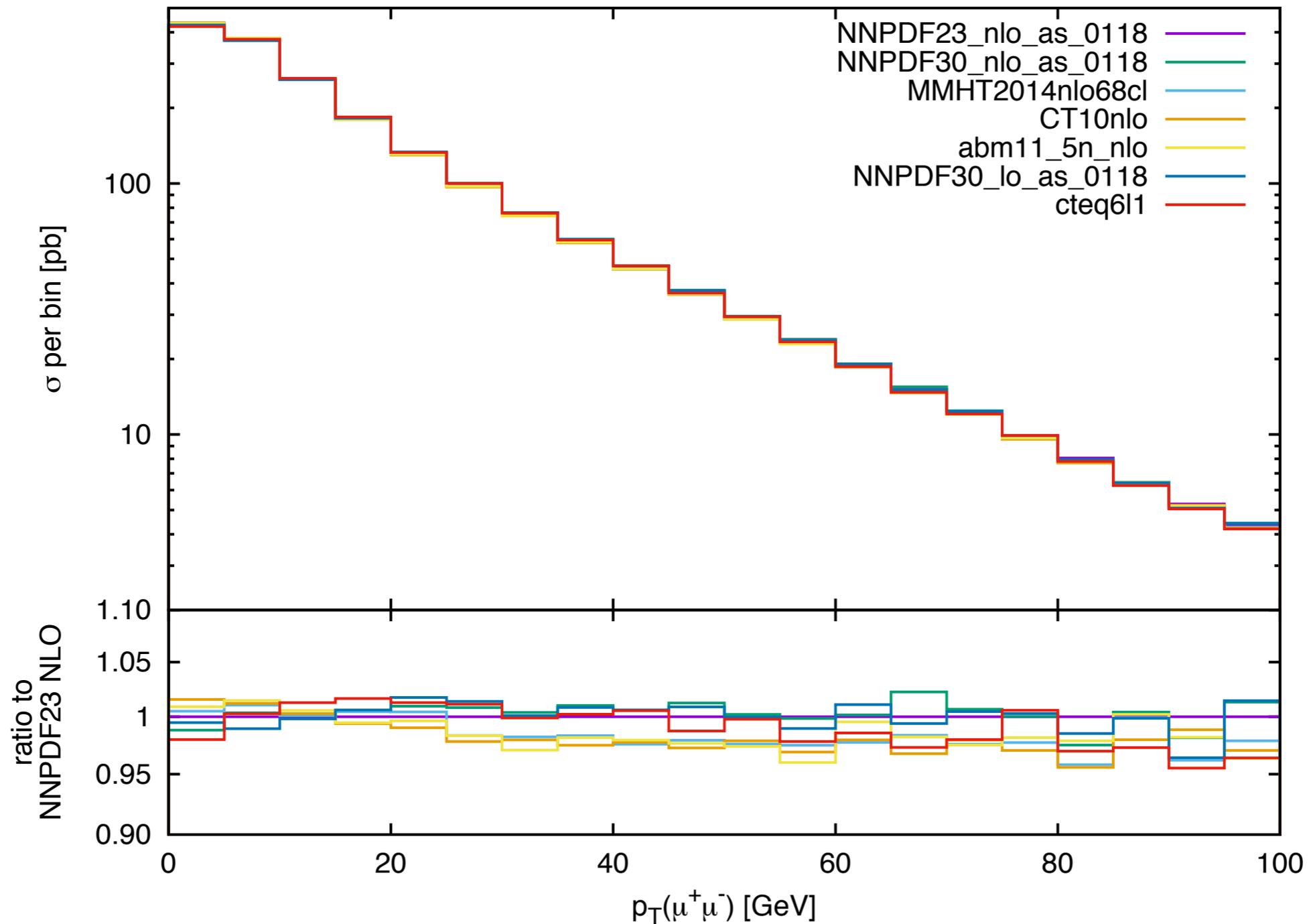
$\mu^+\mu^-$ production at NLO+PS with PYTHIA8, NNPDF23_nlo_as_0118 in the ME



The aMCfast Interface

The NLO + PS Case: Dependence on the IS-PS

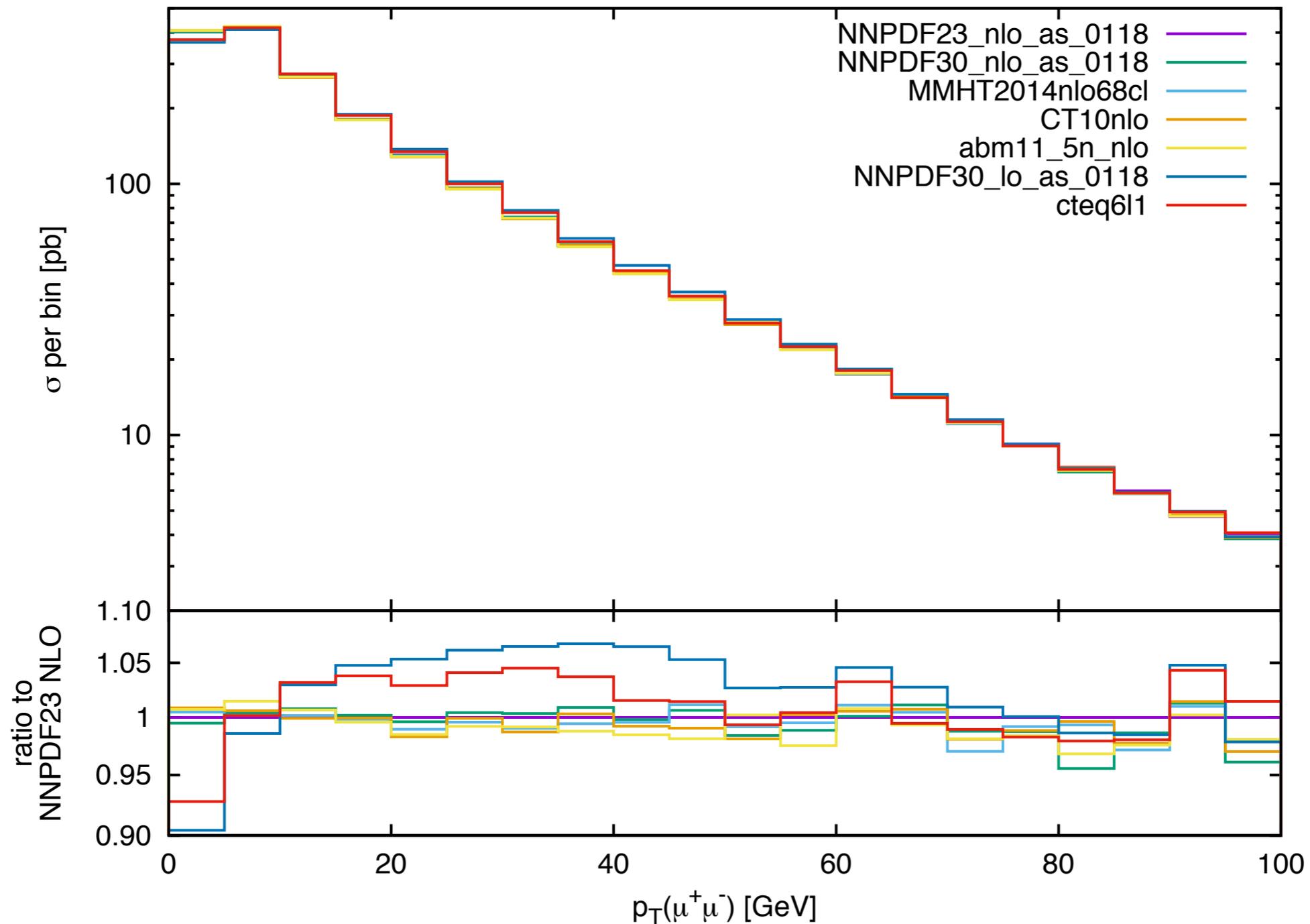
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The aMCfast Interface

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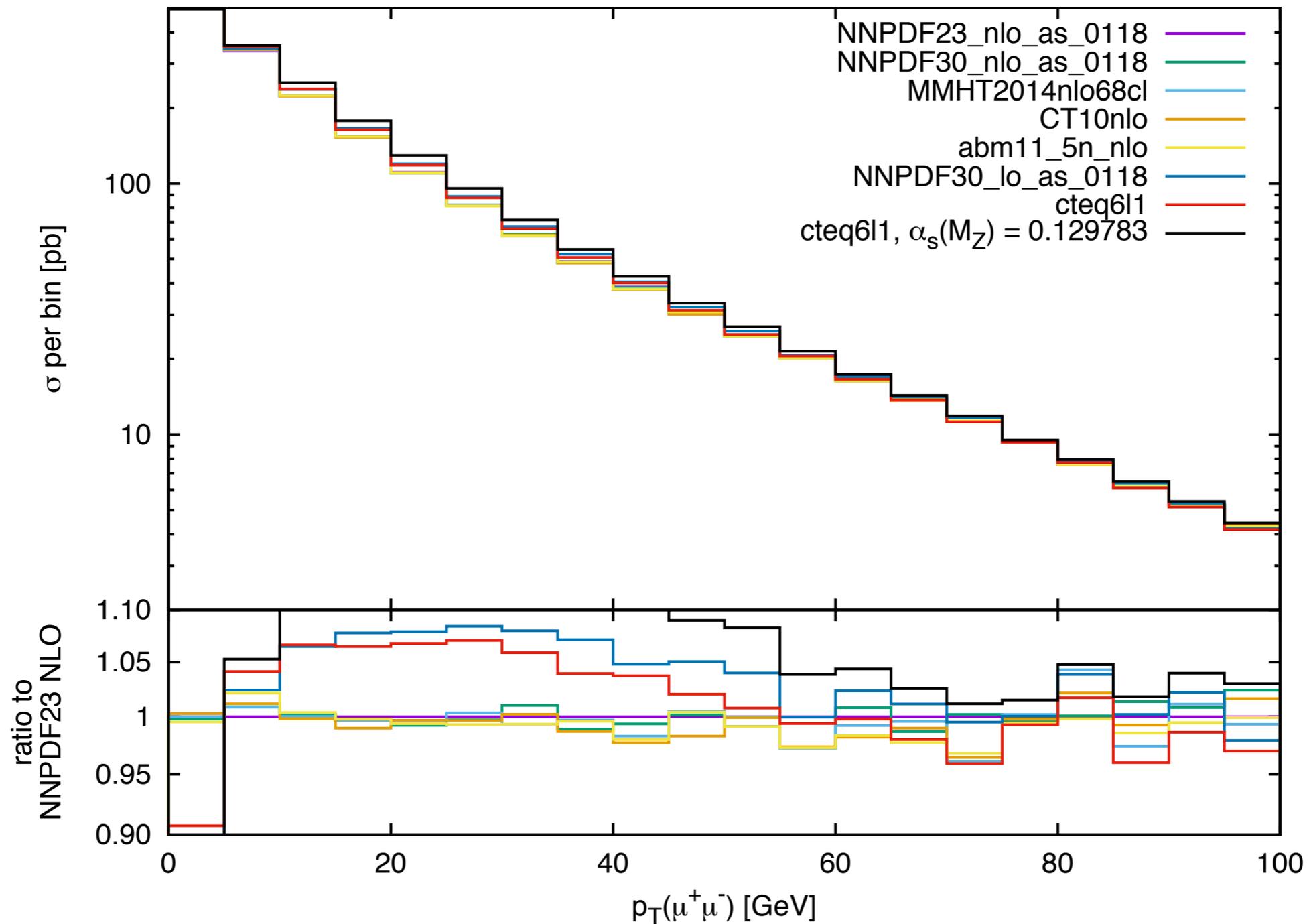
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The aMCfast Interface

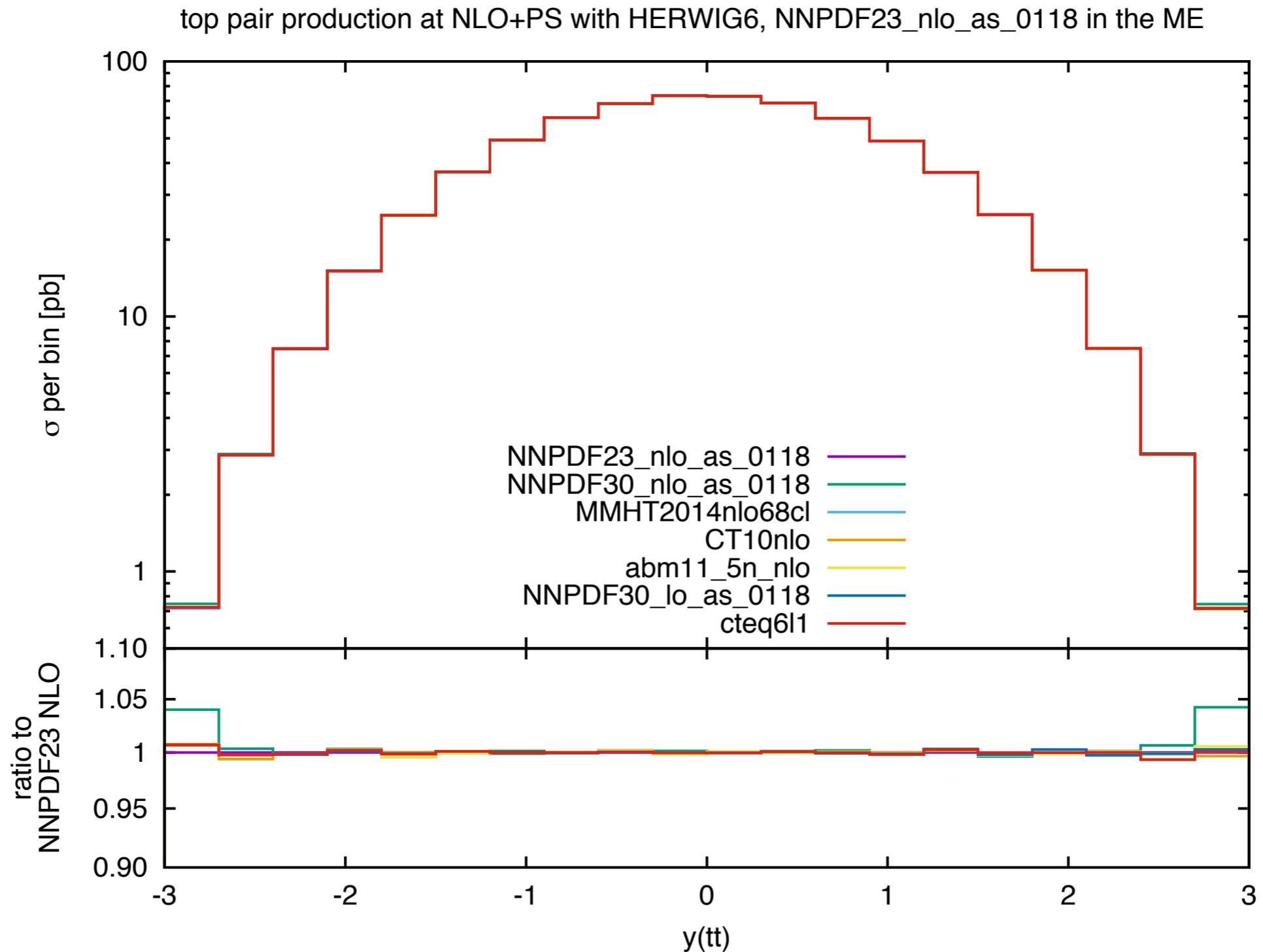
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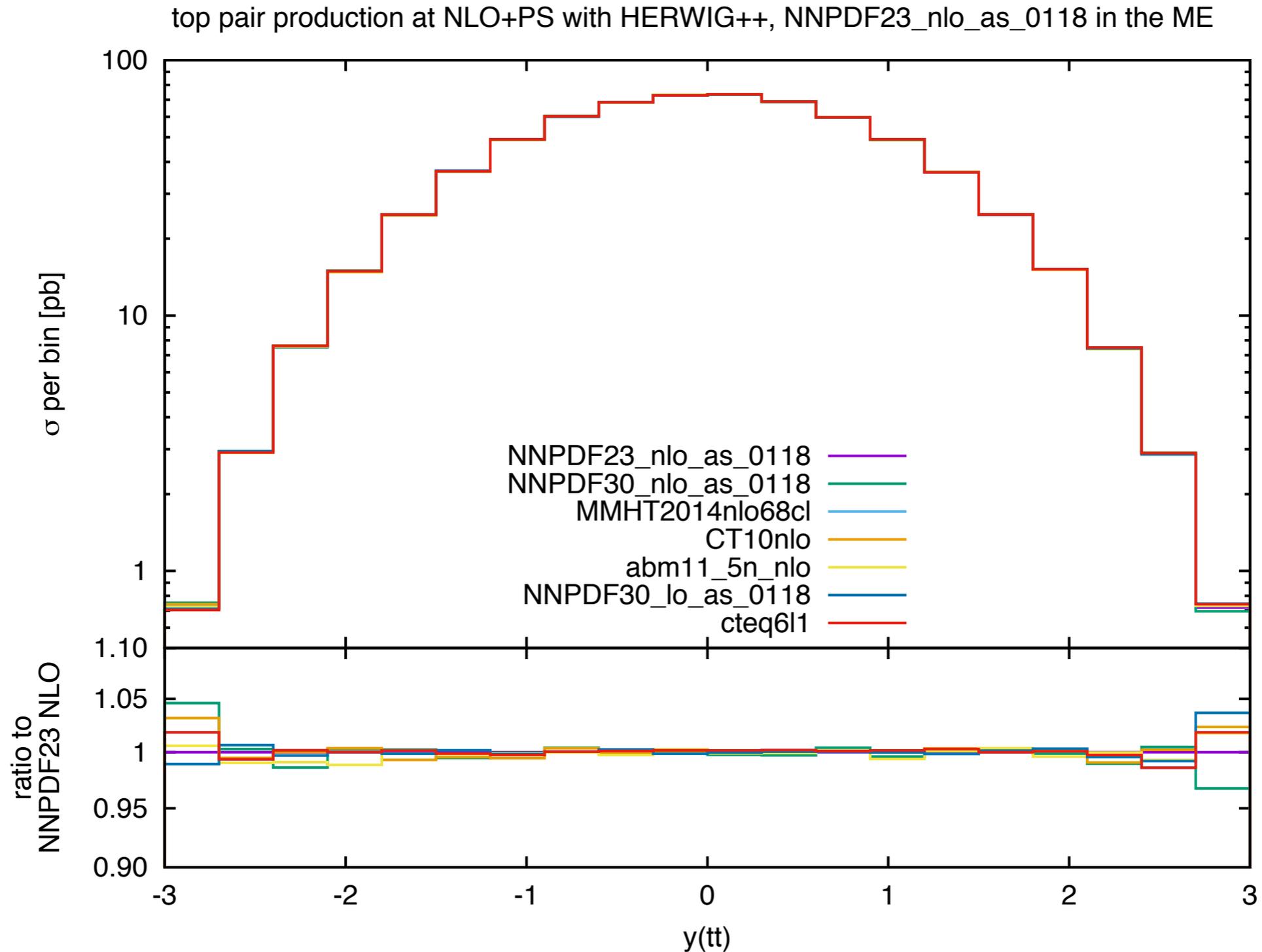
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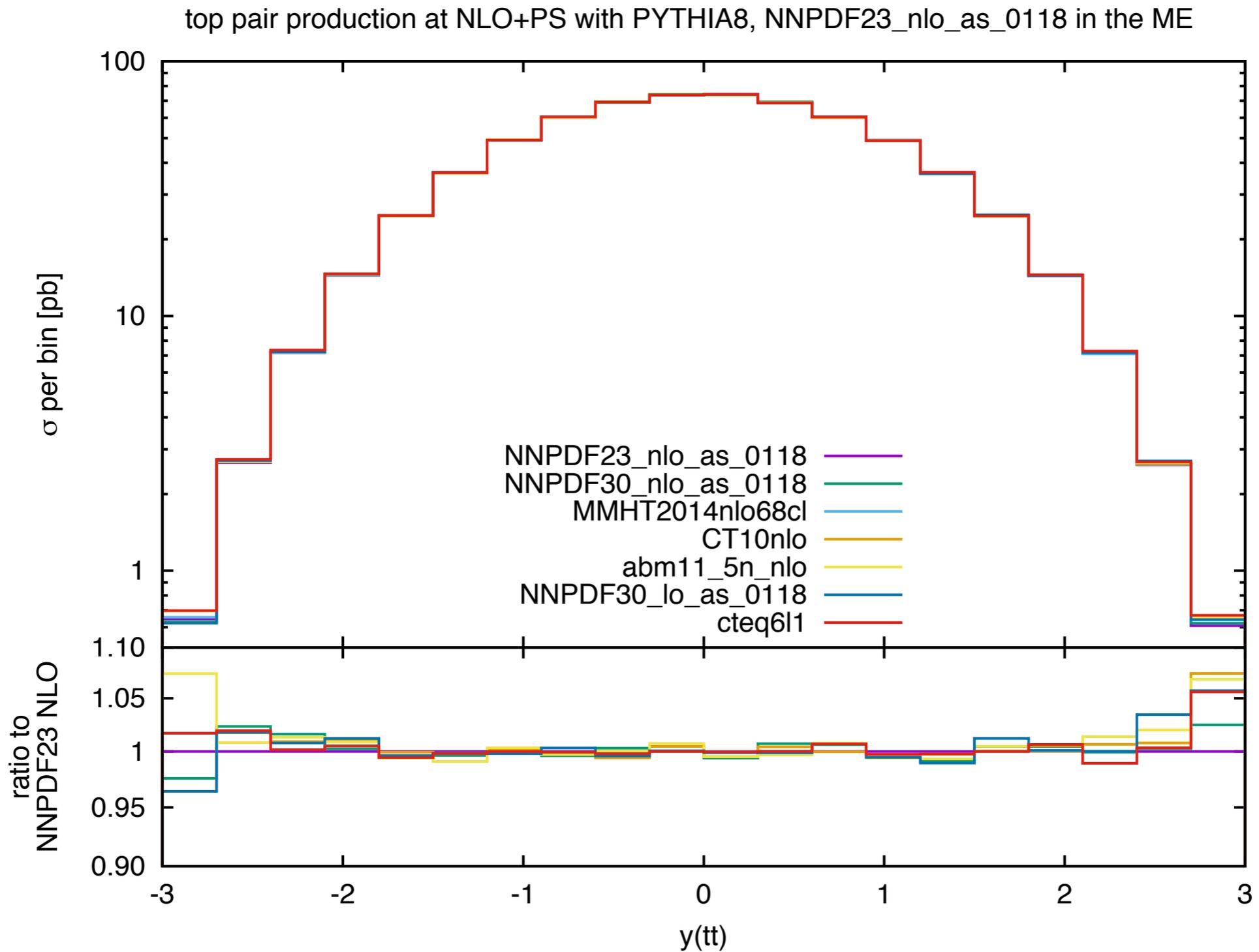
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The NLO + PS Case: Dependence on the IS-PS



The aMCfast Interface

The NLO + PS Case: Dependence on the IS-PS



The aMCfast Interface

The NLO + PS Case

- In general, the **dependence** on the PDF set used in the PS is **contained**:
 - provided that “reasonable” PDFs are used,
 - using LO or NLO PDFs can make a substantial difference.
- However, there are corners (*e.g.* **large rapidity regions**) where such dependence is more marked.
- In addition, in those corners different PSs might behave differently.
- Also, less inclusive observables present a stronger dependence (merging should help).
- In conclusion, even though deviations are typically contained, there is **no guarantee** that the procedure is generally unbiased and thus formally suitable for PDF fits. However, how relevant is the bias?
 - To answer this question we need to estimate the **potential impact** of the PS corrections on PDFs.

The aMCfast Interface

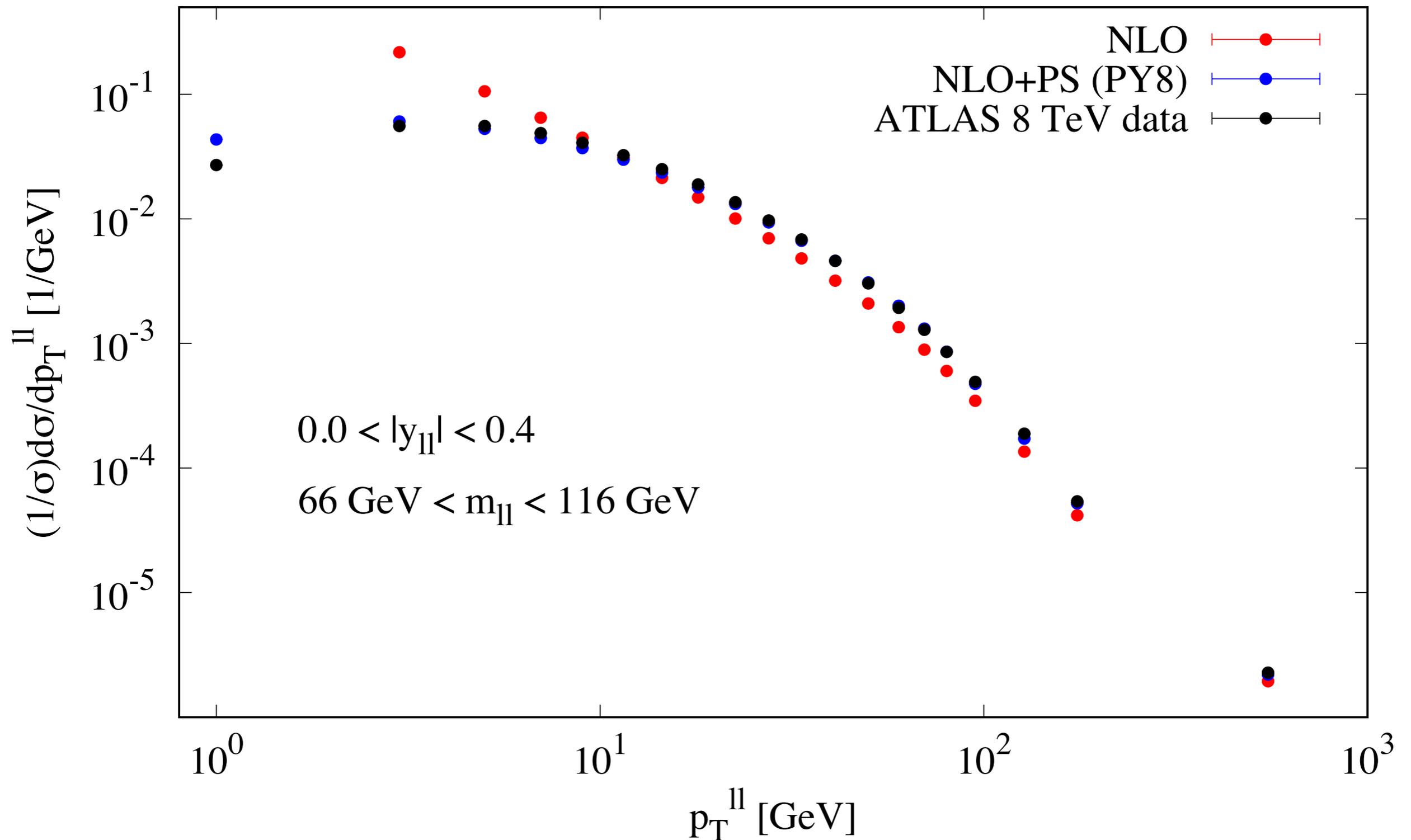
The NLO + PS Case: The ATLAS Z p_T data

- As first step we considered the ATLAS **Z p_T distribution** data at 8 TeV in the lepton-pair invariant mass bin $66 \text{ GeV} < m_{ll} < 116 \text{ GeV}$:
 - 6 rapidity bins in the range $0 < |y_{ll}| < 2.4$ with 20 data points each.
 - Extremely **small uncertainties**.
- Data points in each rapidity bin extend between 0 and 900 GeV in p_T :
 - when fitting PDFs at fixed order low p_T data must be excluded, typically **$p_T > 25 \text{ GeV}$**
 - in addition, since the p_T at LO is identically zero, one needs to supplement the NLO predictions with NNLO k-factors (or alternatively consider Z+j).
 - PS is expected to help in describing low p_T data and thus the cut can be lowered accordingly.

The aMCfast Interface

The NLO + PS Case: The ATLAS Z p_T data

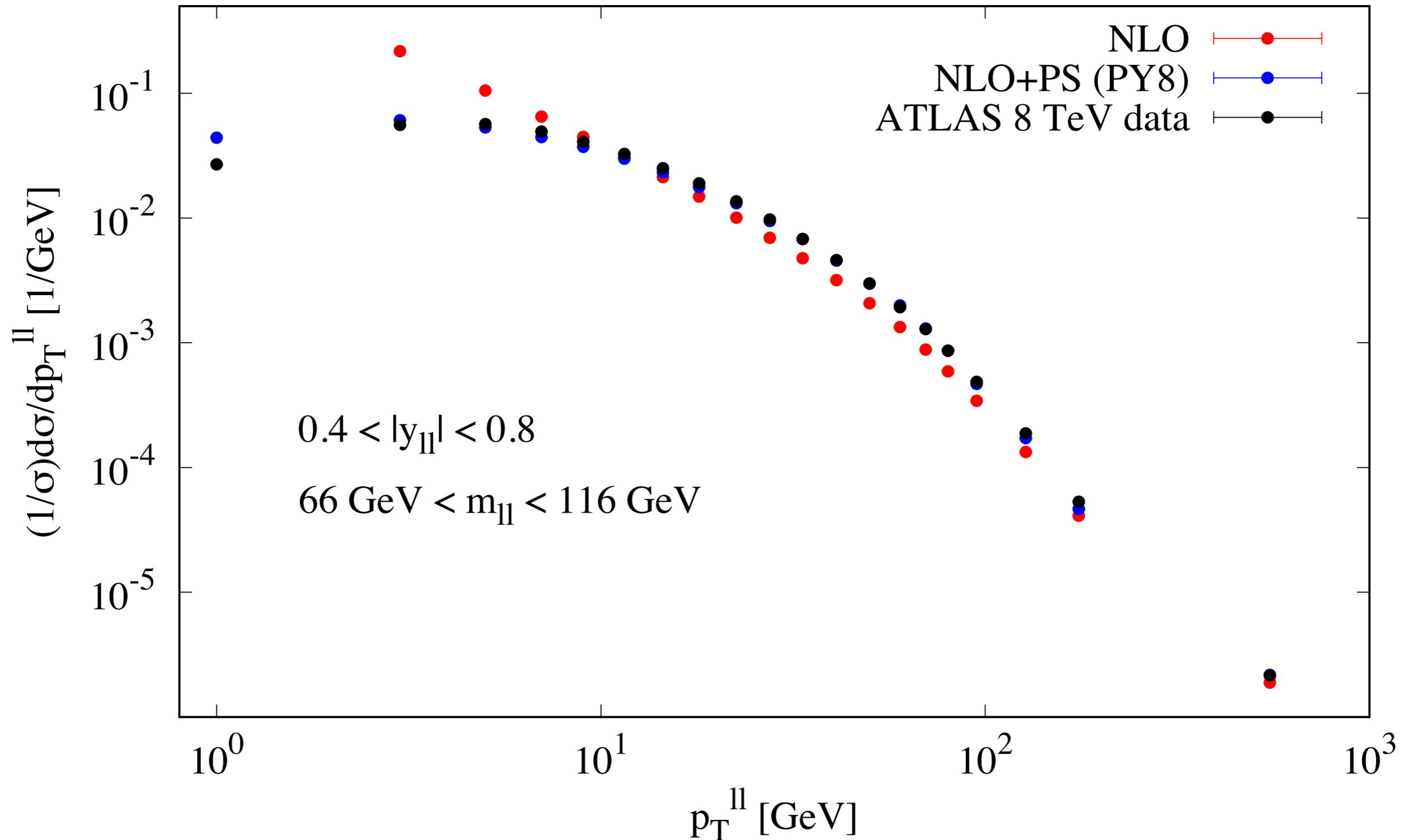
MadGraph5_aMCatNLO + NNPDF23_nlo_as_0118



The aMCfast Interface

The NLO + PS Case: The ATLAS Z p_T data

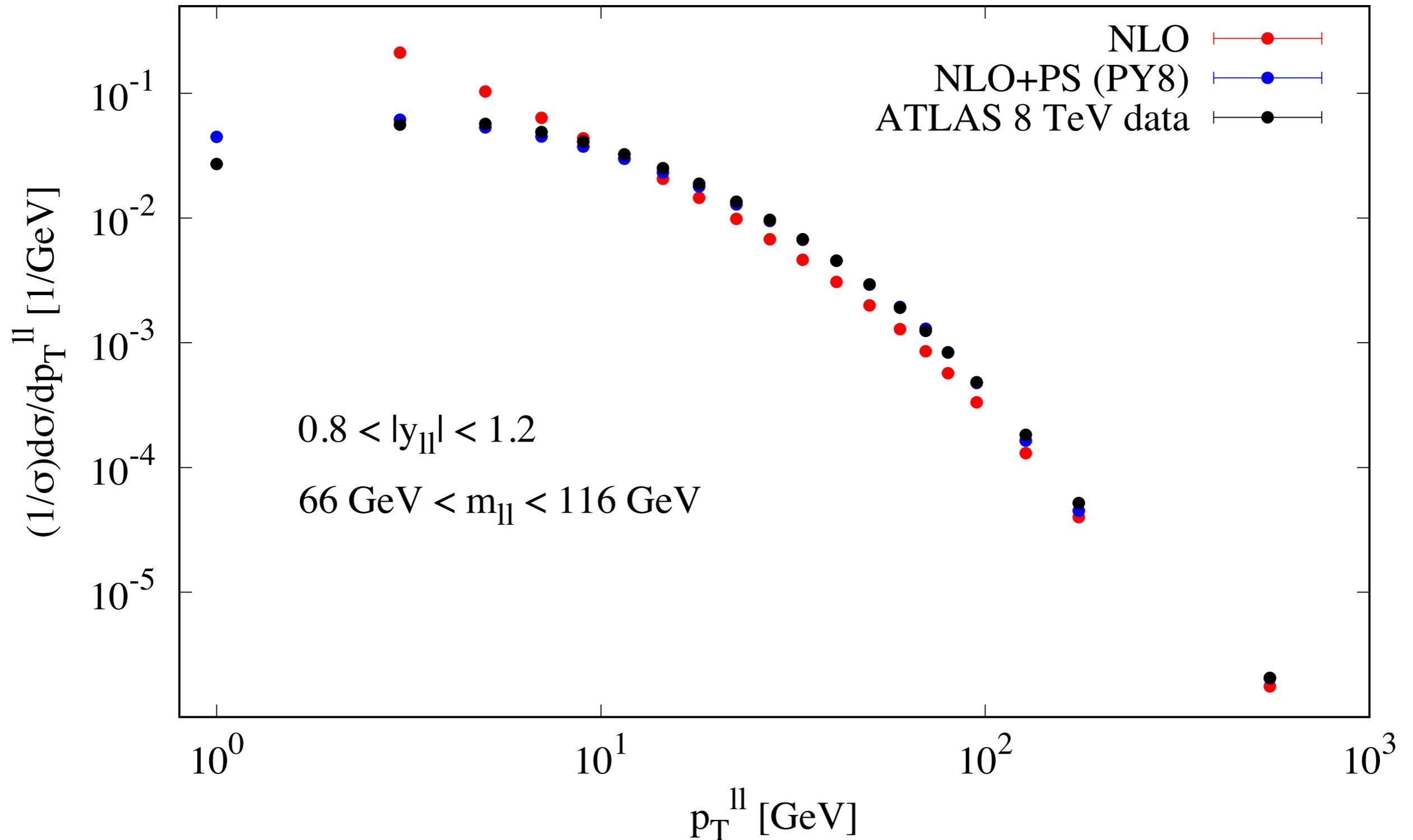
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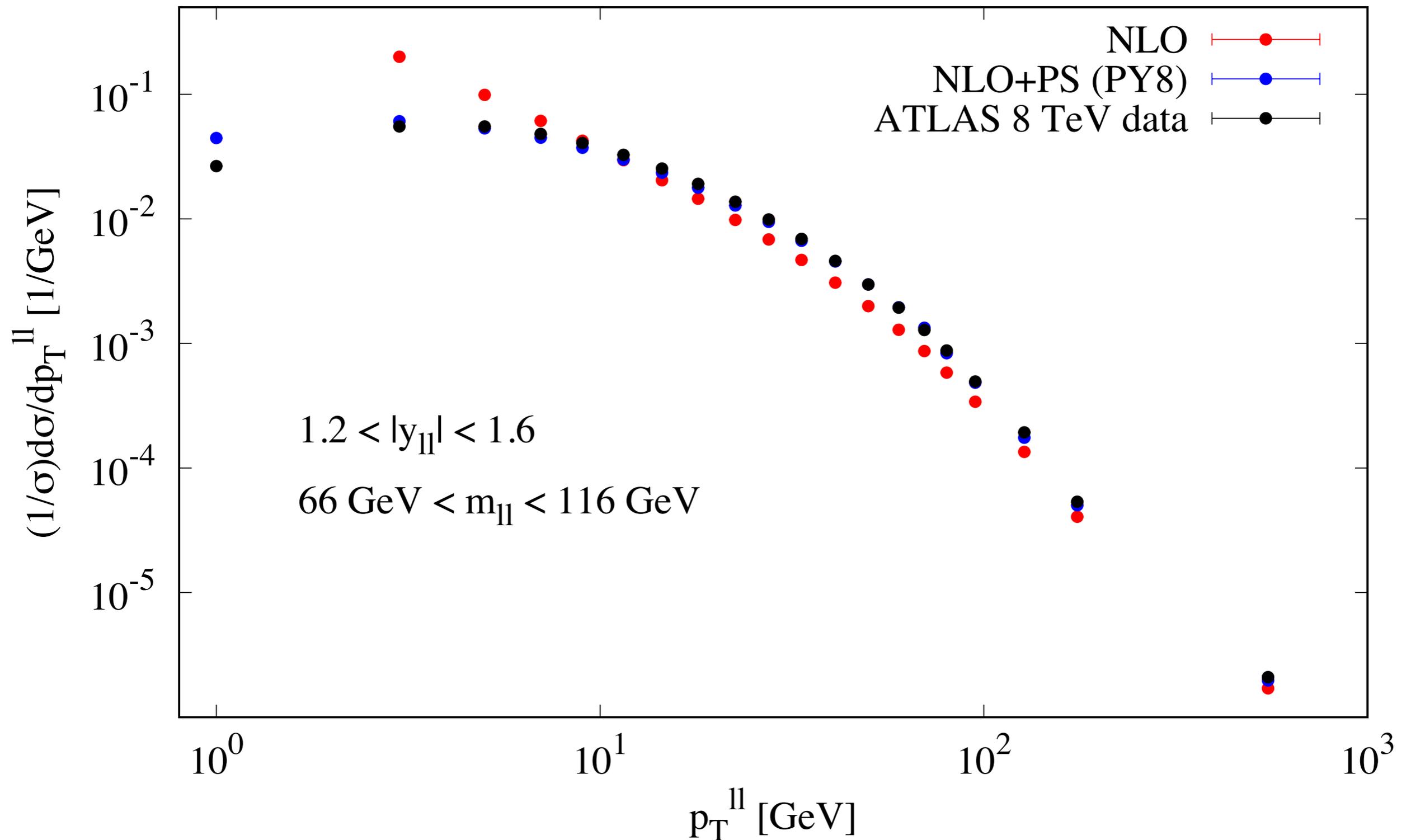
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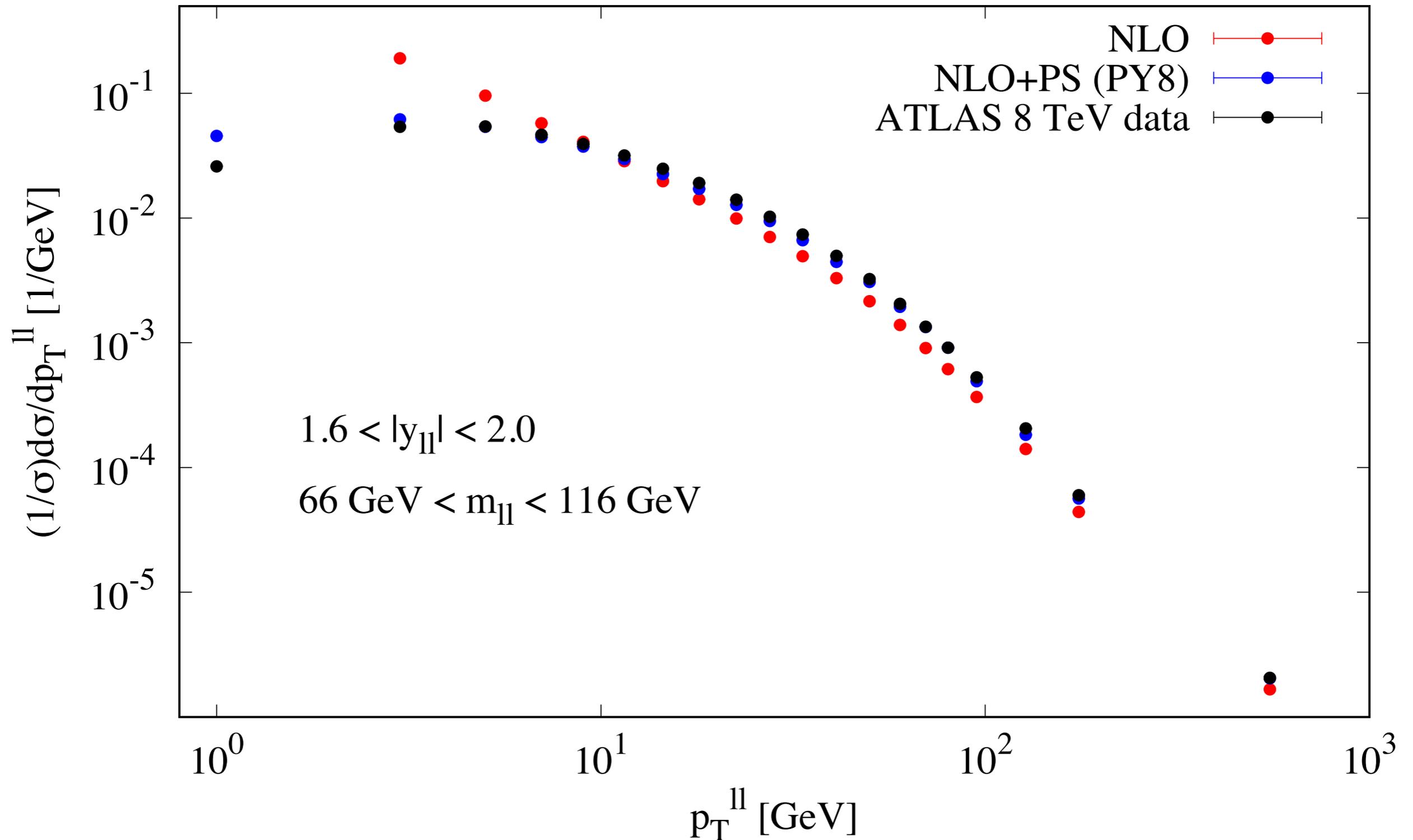
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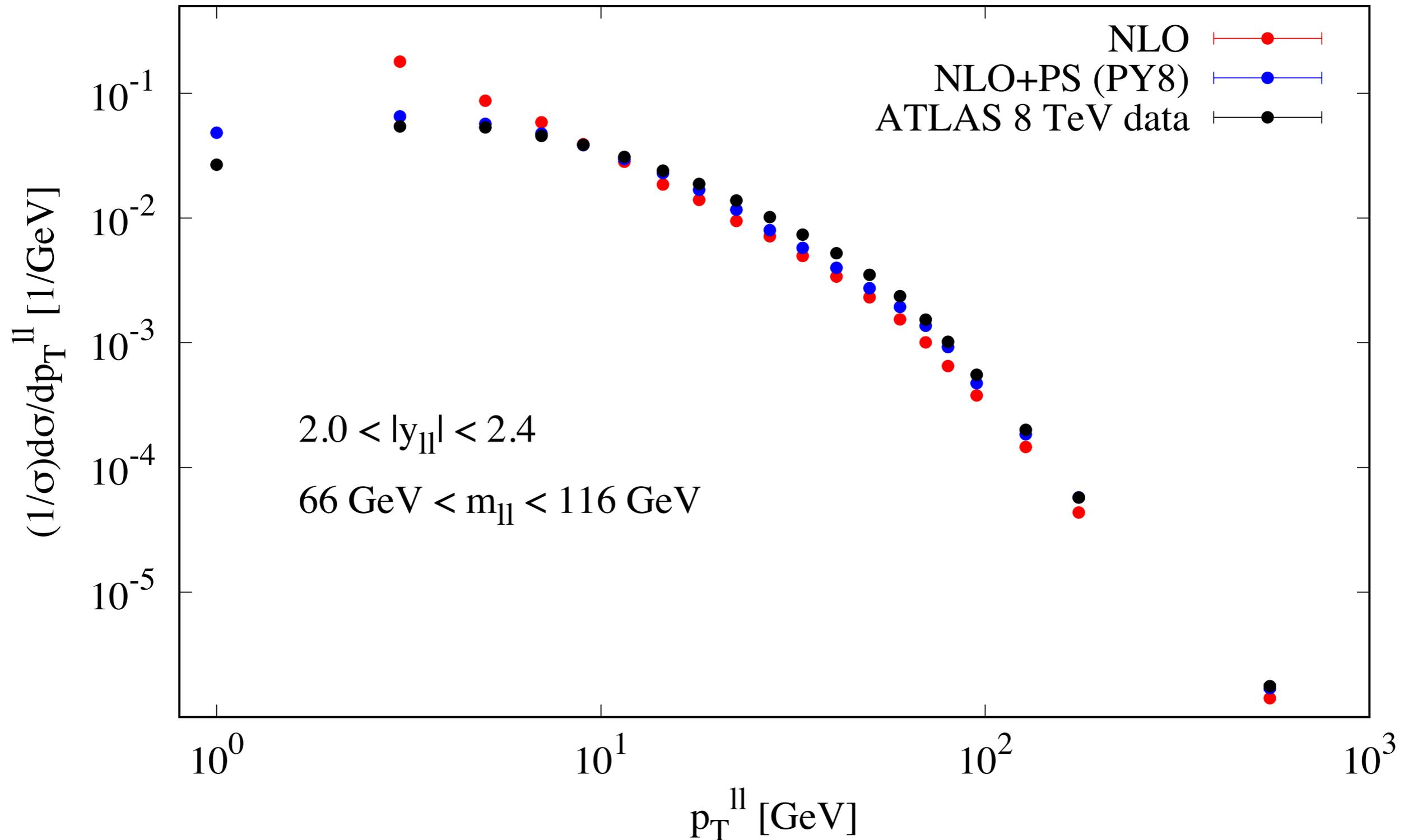
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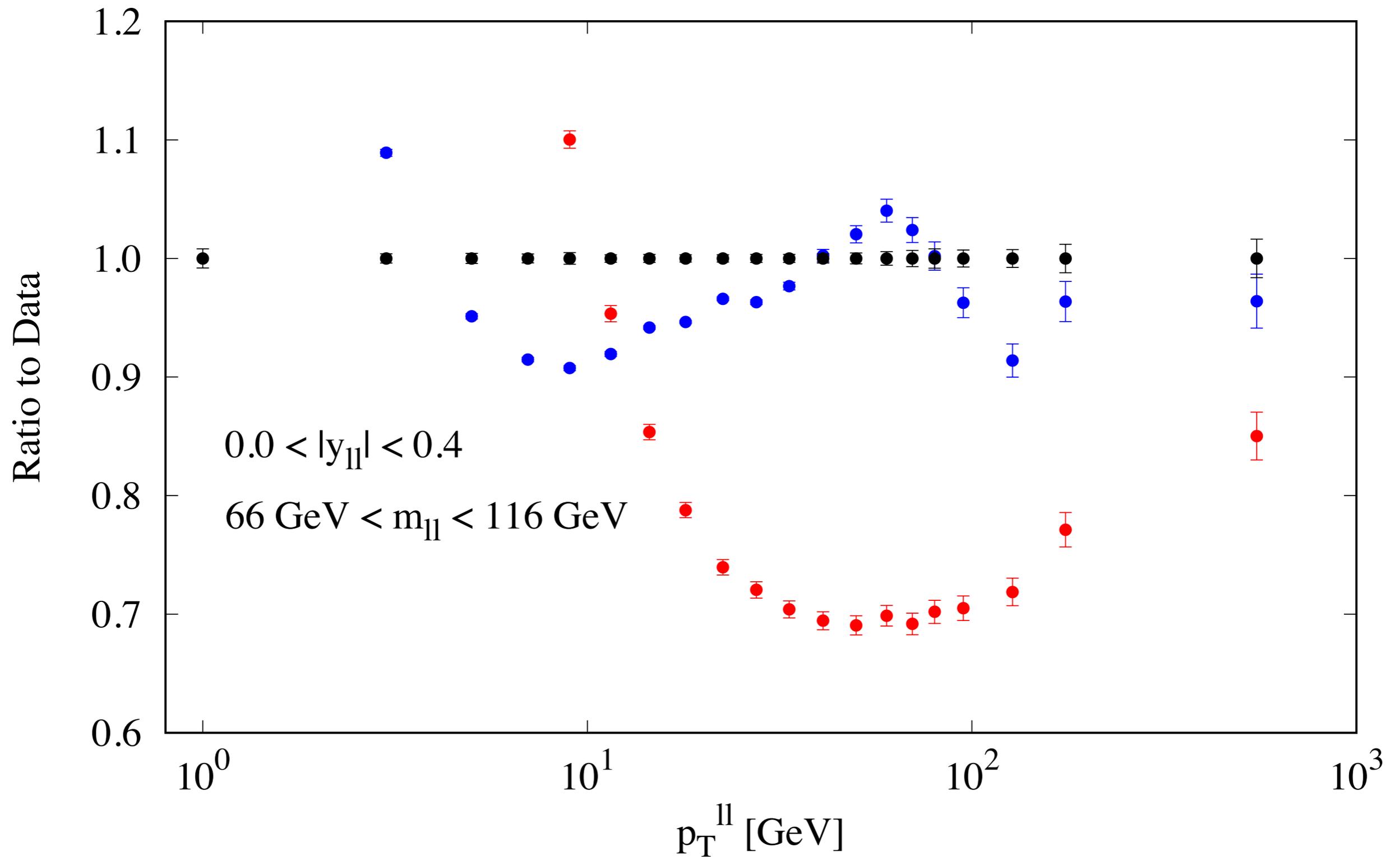
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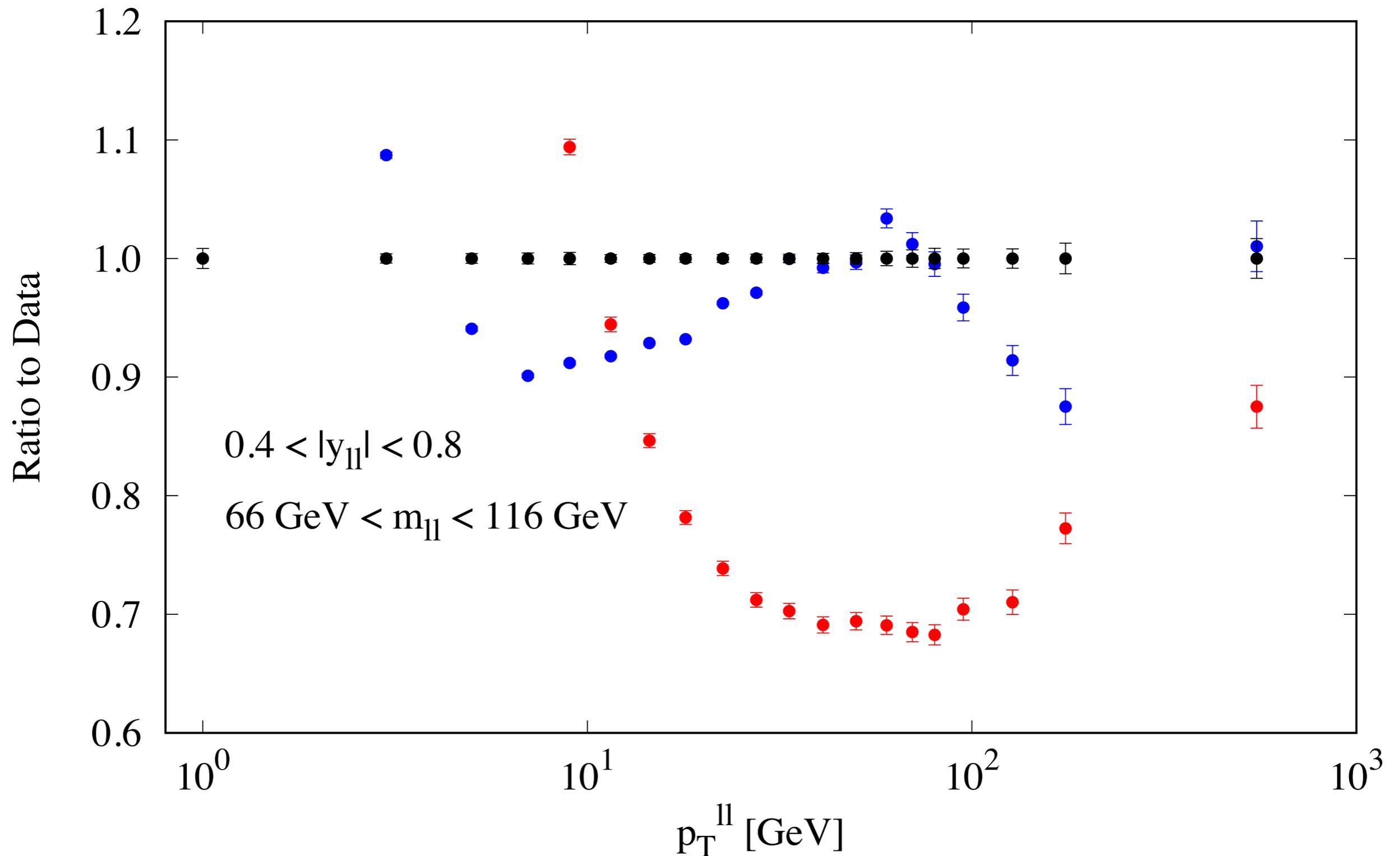
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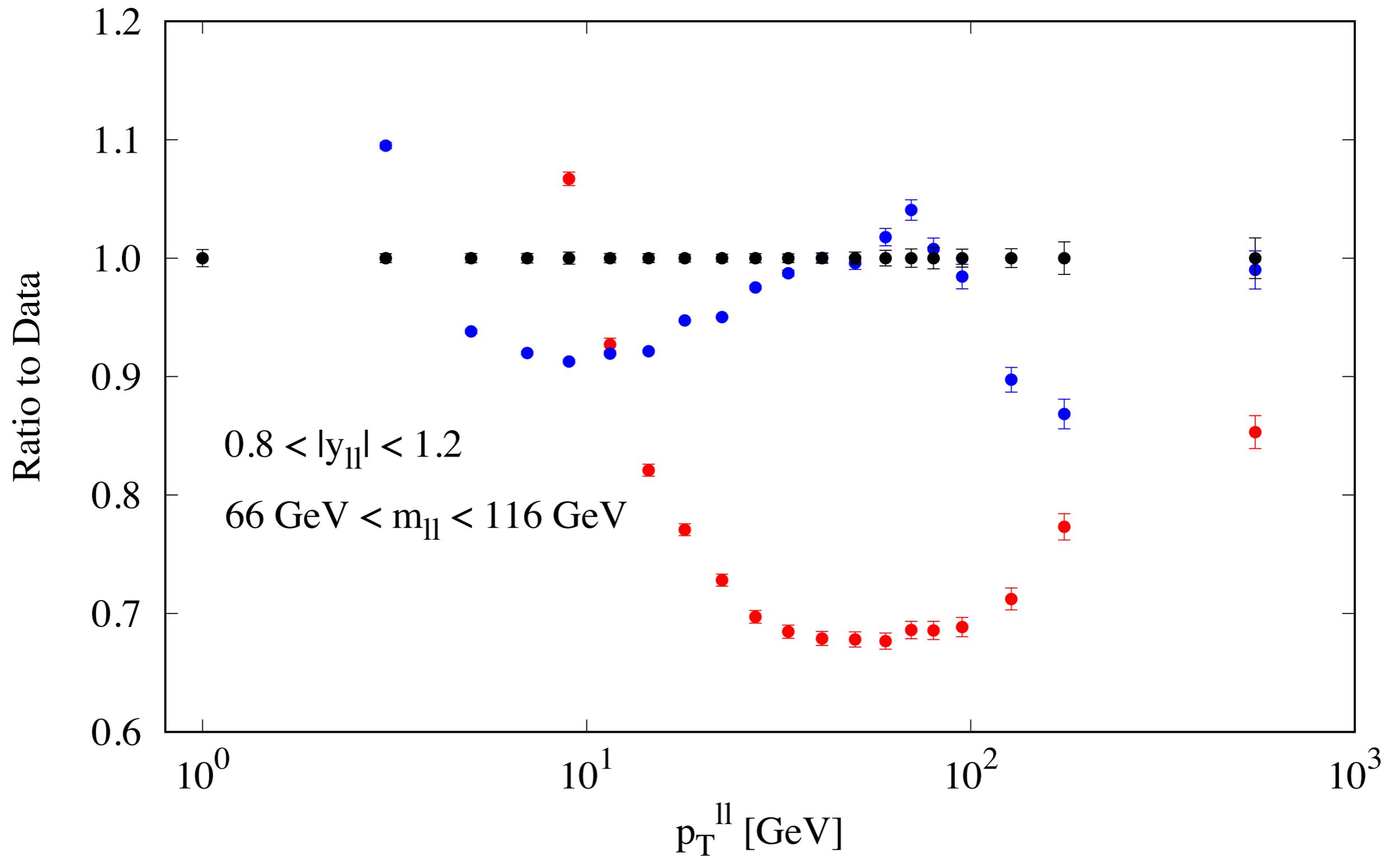
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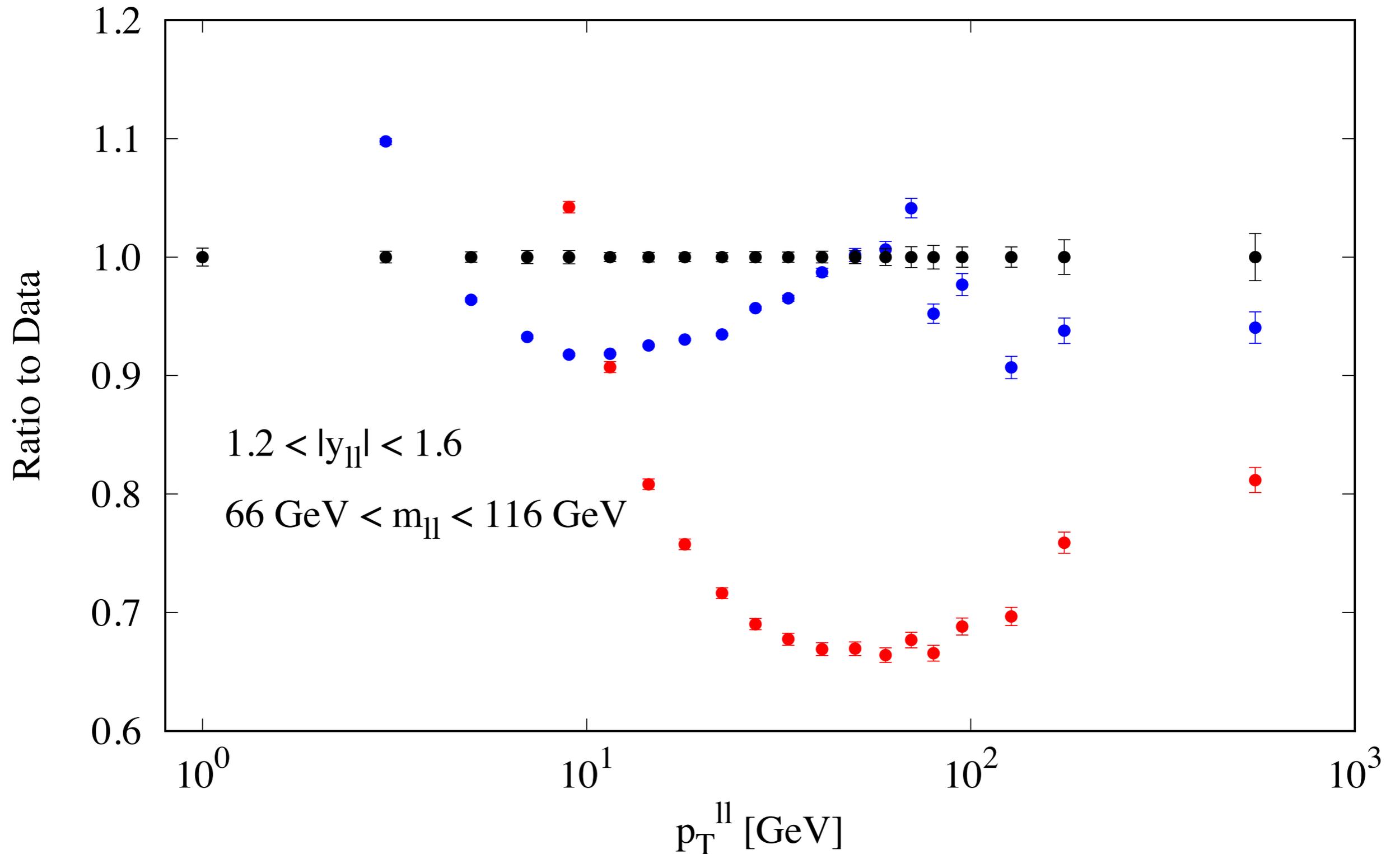
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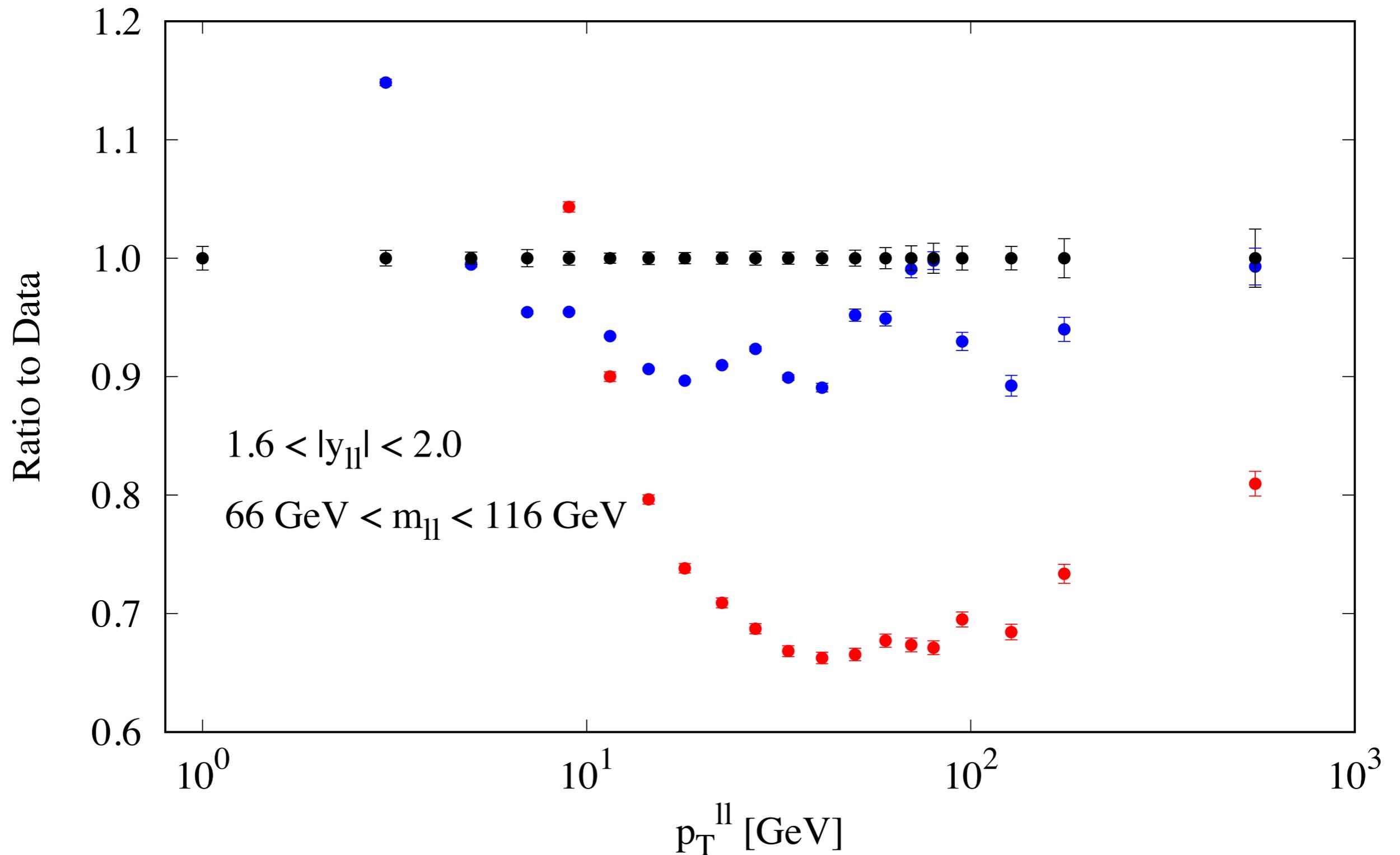
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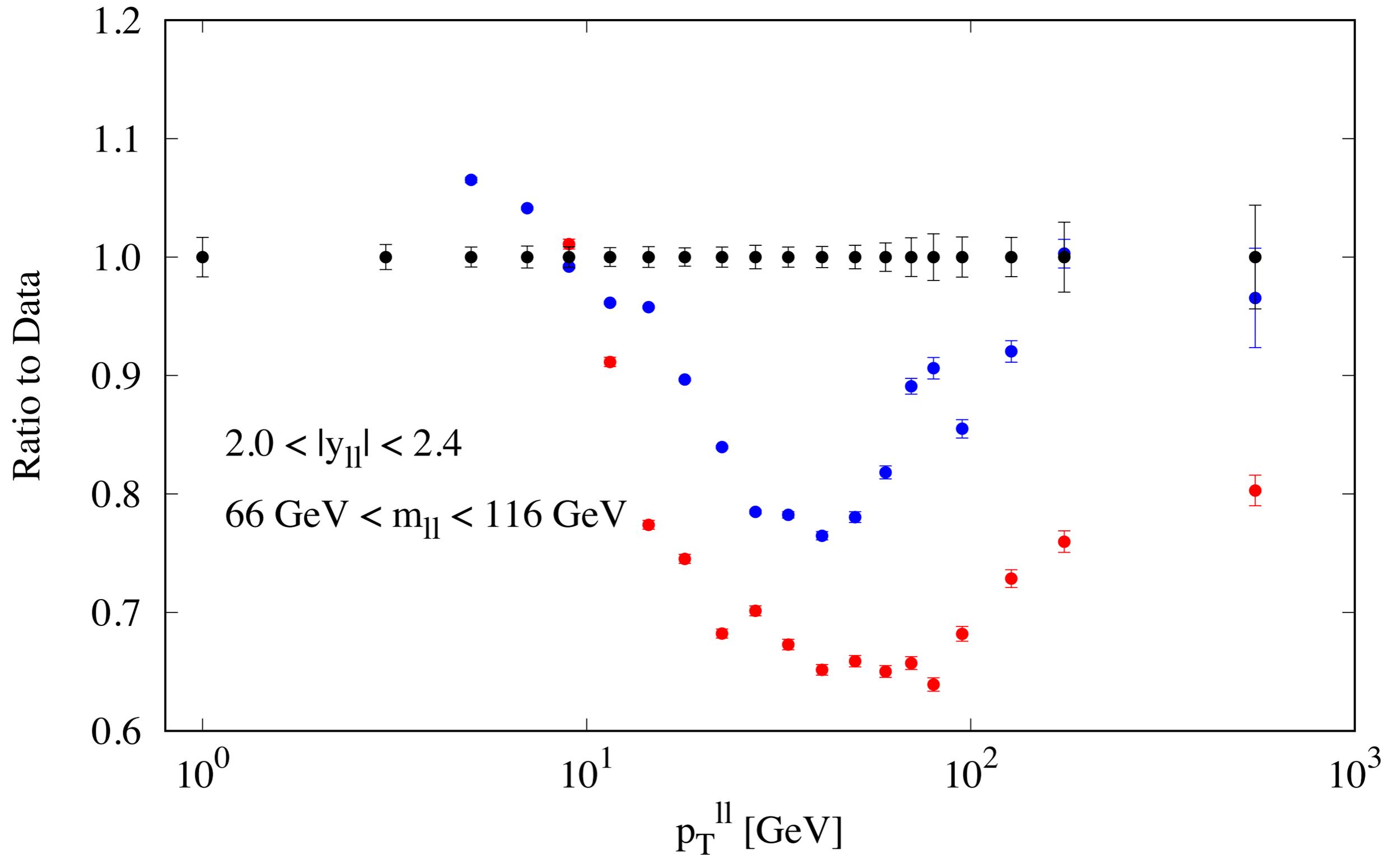
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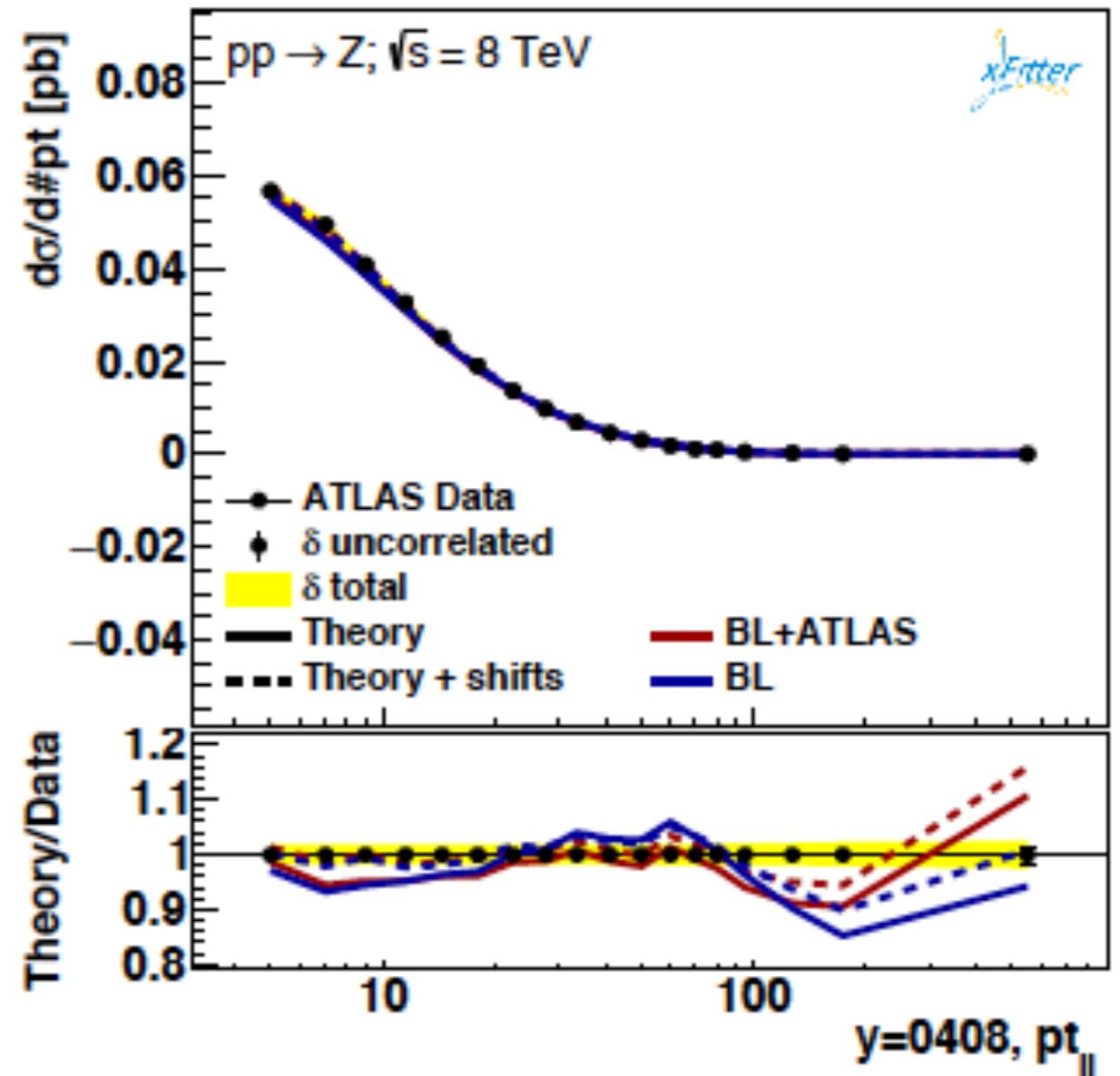
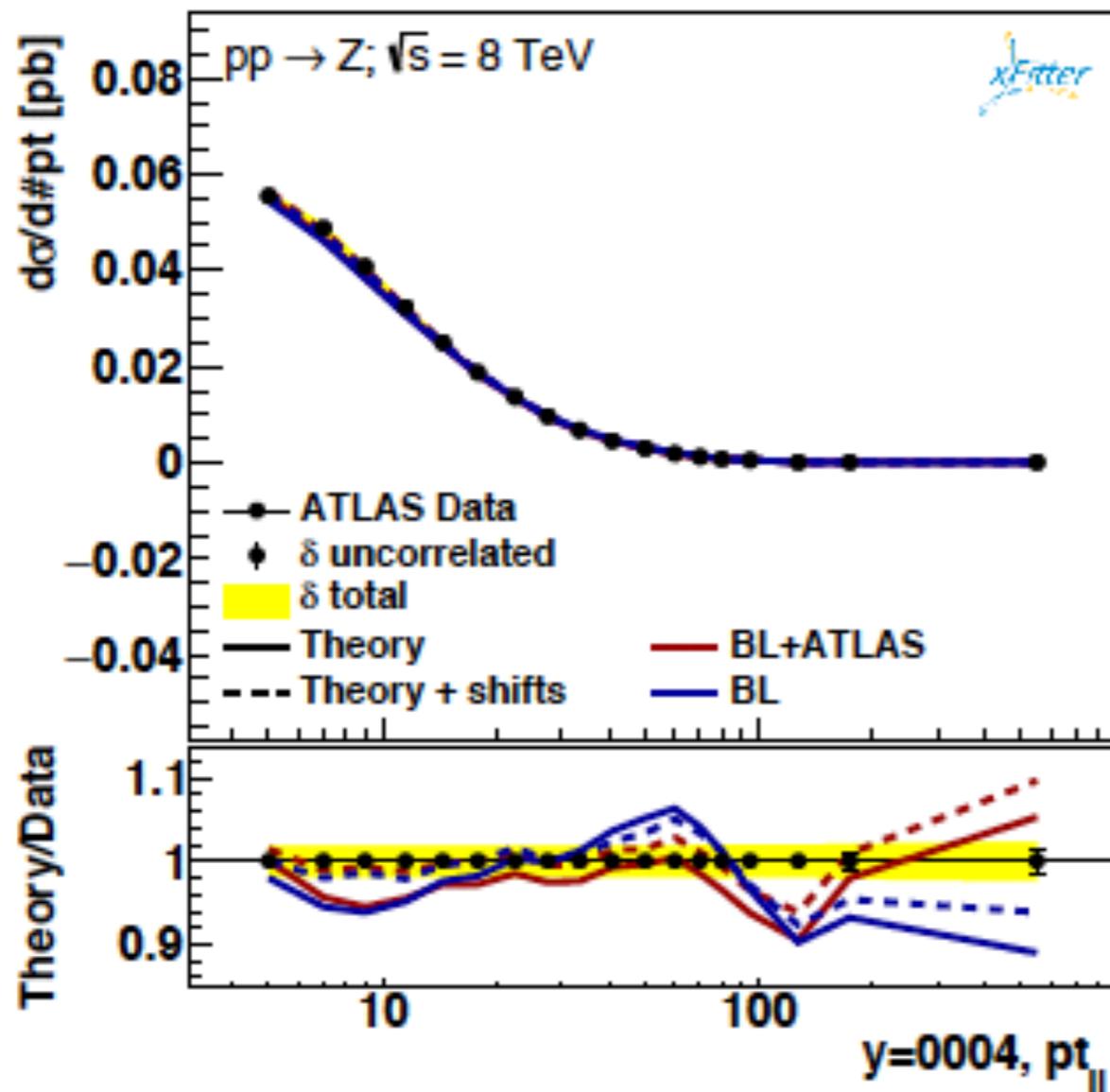
The aMCfast Interface

The NLO + PS Case: Attempting a First PDF Fit

- Actually the PS does help but the experimental uncertainties are so small that a good **description** is **hardly achievable**.
- Try to fit a set of PDFs to this data:
 - use HERA1+2 combined inclusive DIS data for the baseline,
 - include ATLAS Z p_T data on top of them and check the impact on PDF.
 - Exclude the first two p_T bins because they are very off.
- To do so we have used the the open-source fitting code **xFitter** (former HERAFitter).

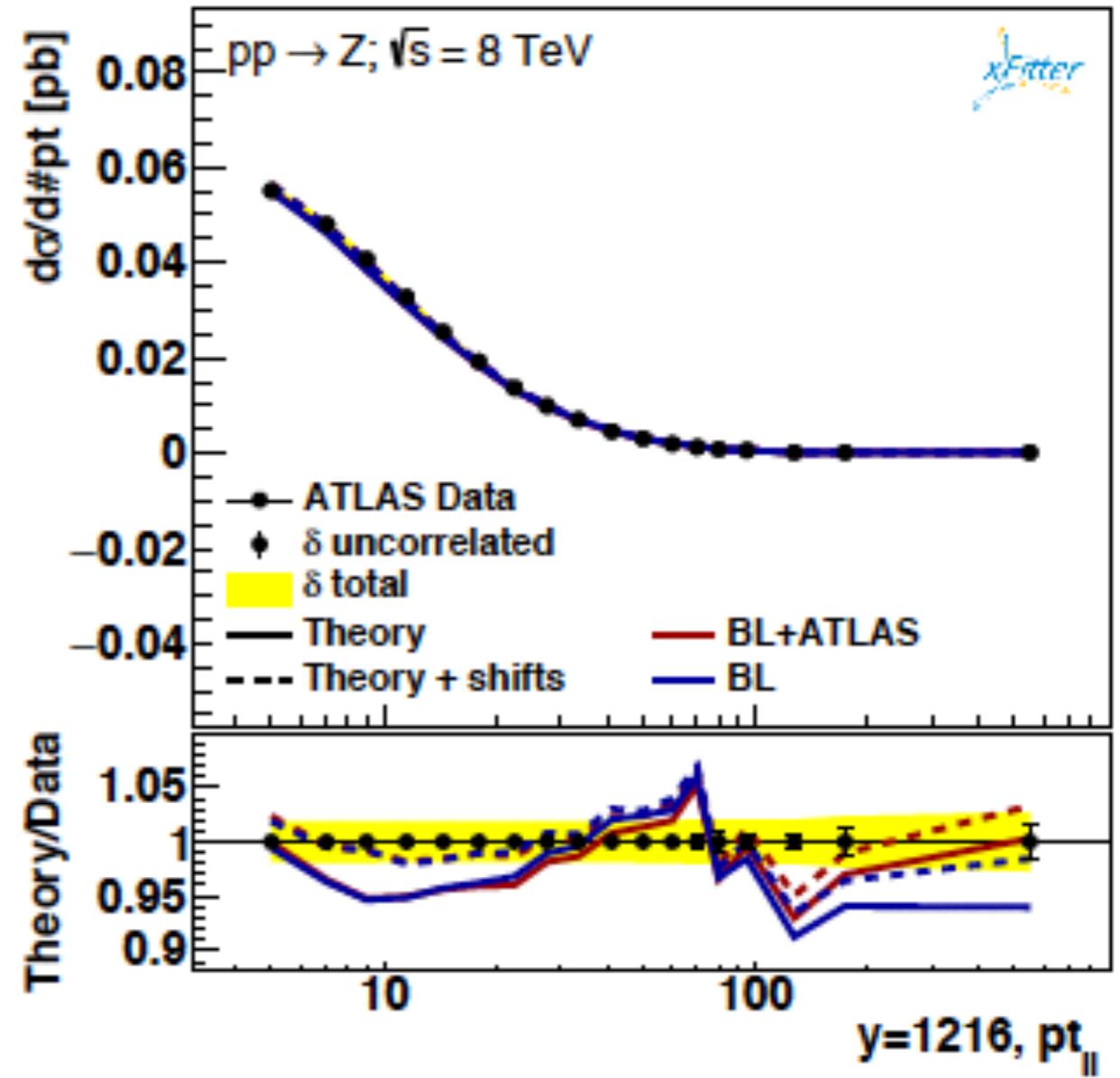
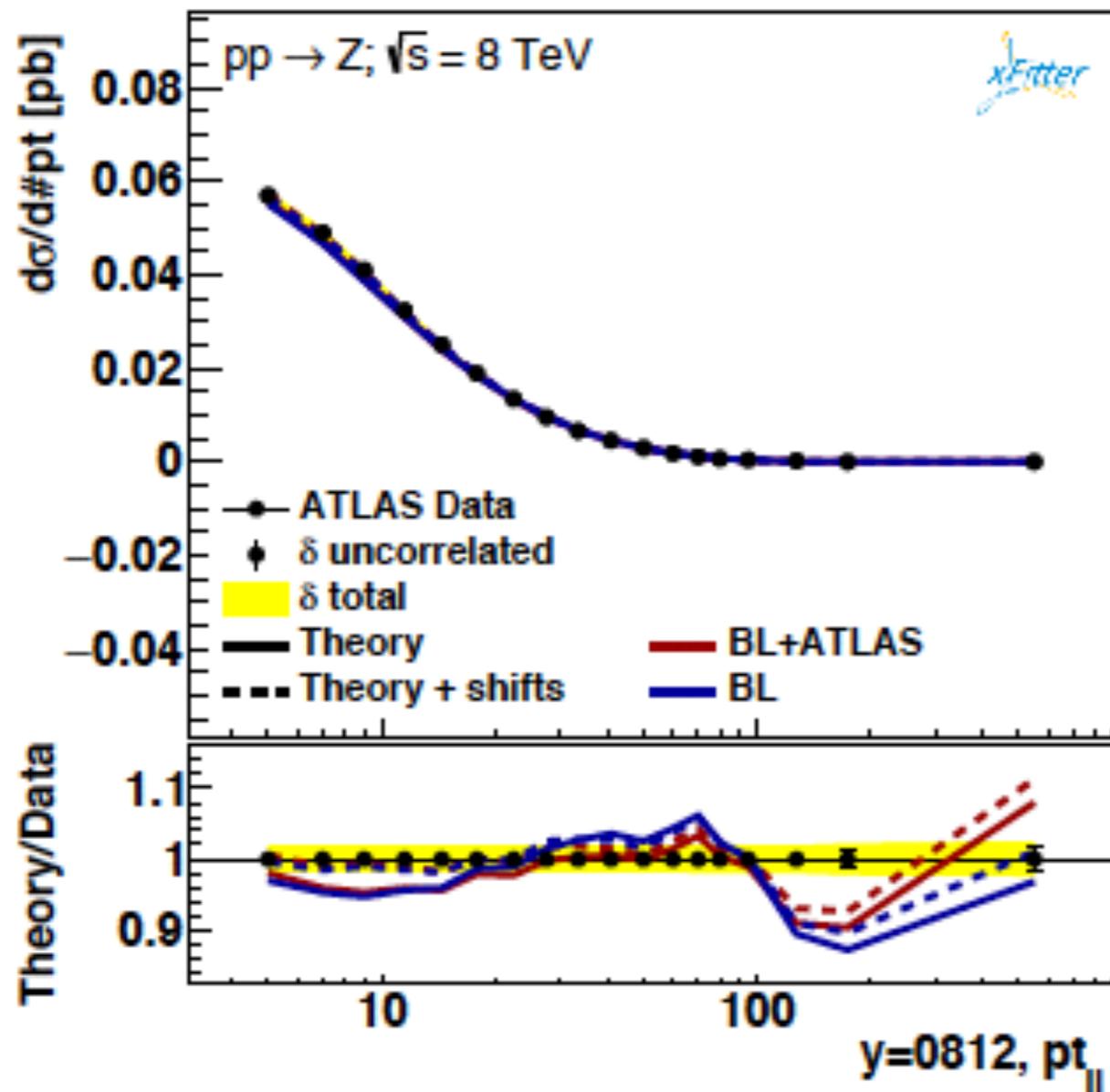
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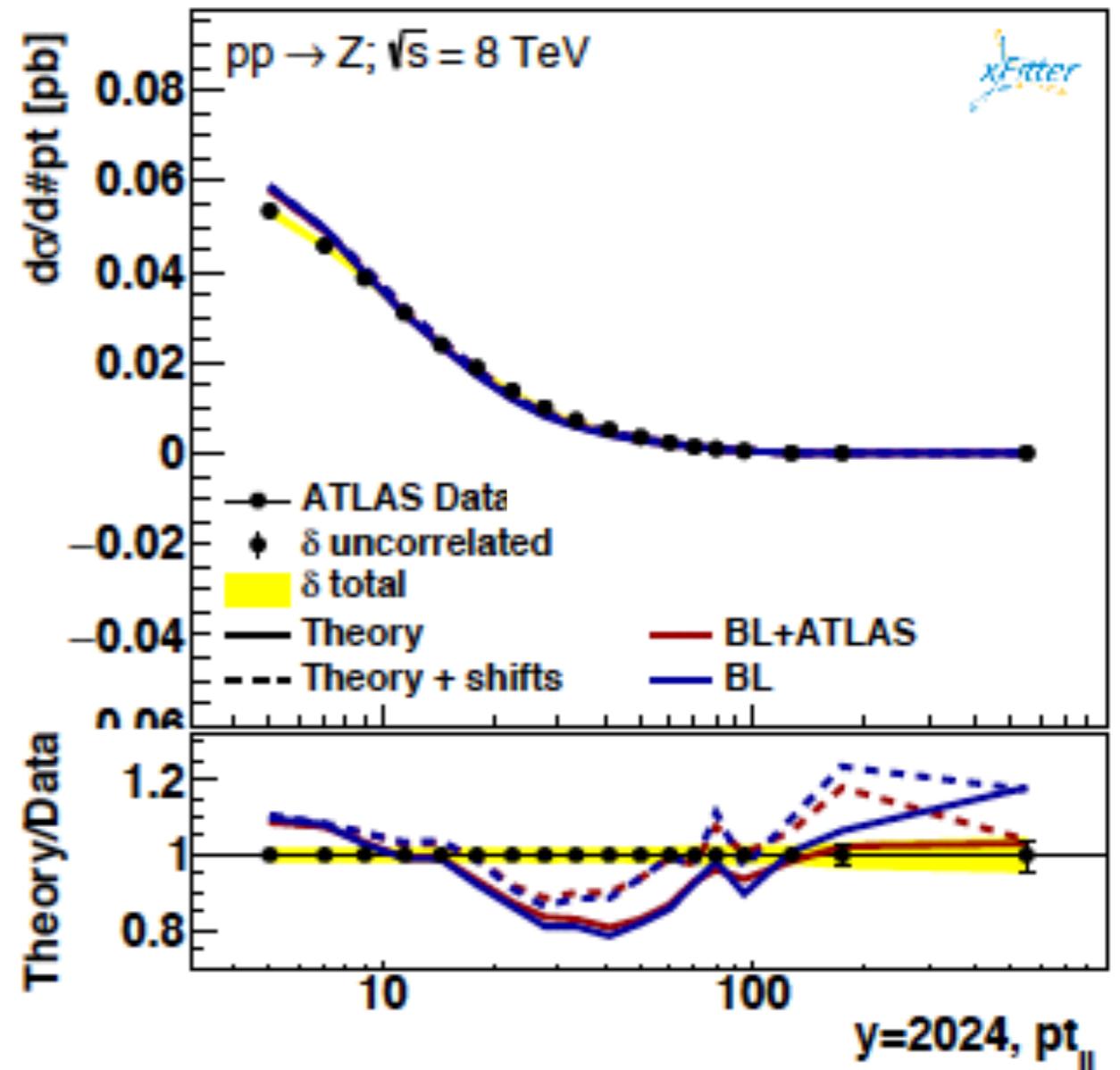
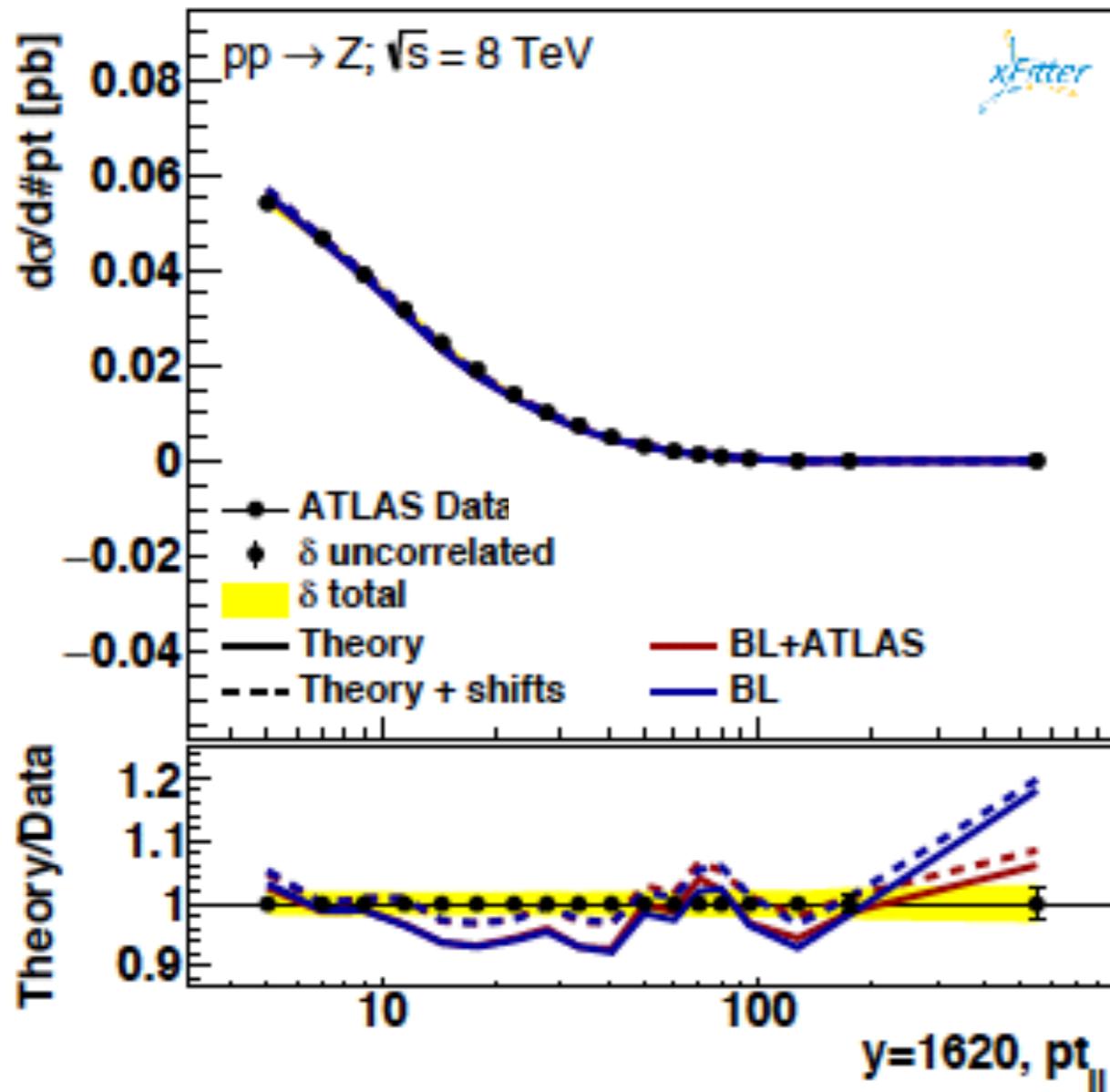
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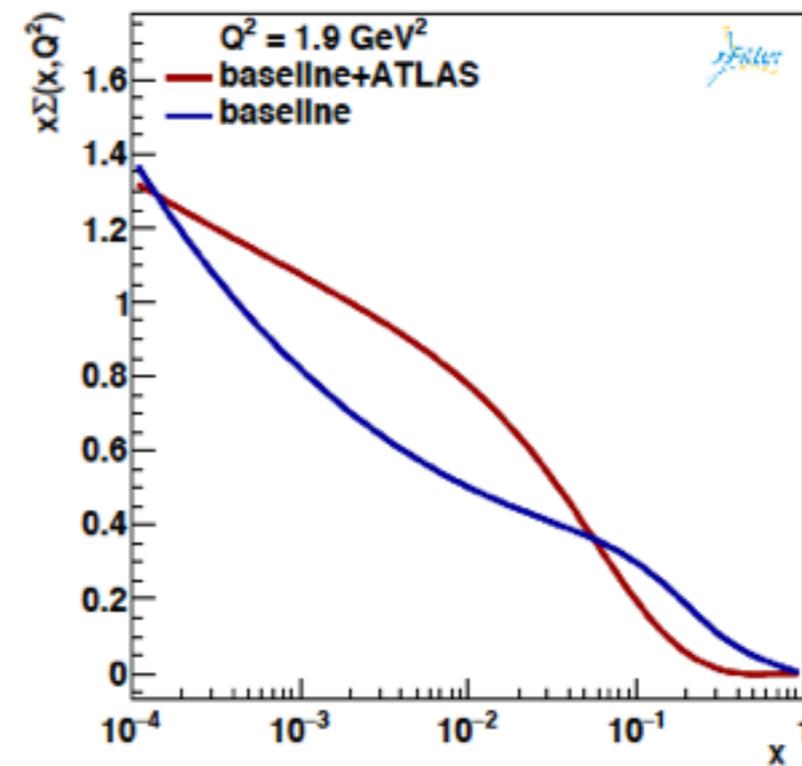
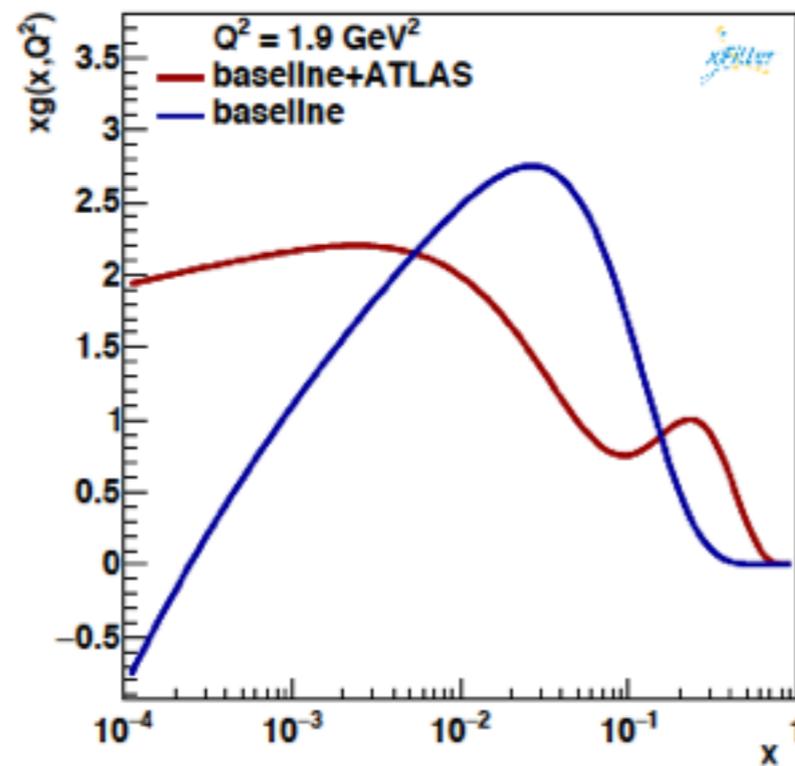
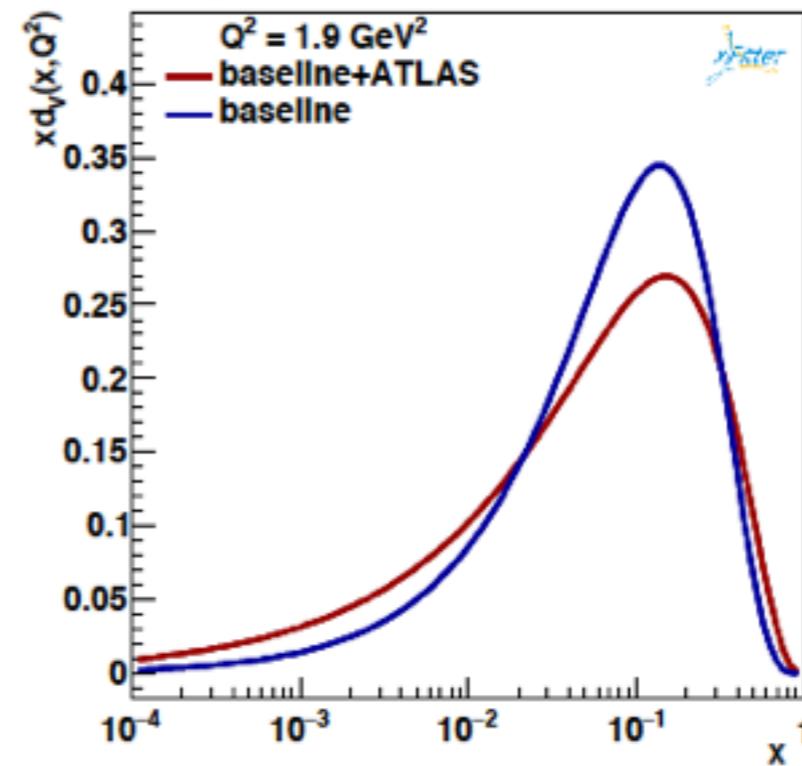
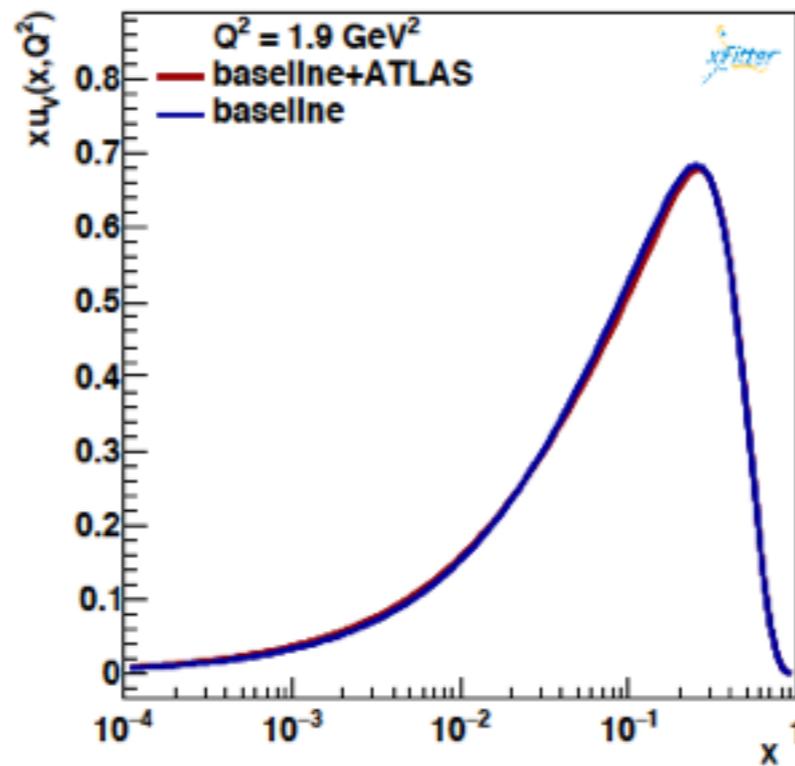
The aMCfast Interface

The NLO + PS Case: Attempting a First PDF Fit



The aMCfast Interface

The NLO + PS Case: Attempting a First PDF Fit



The aMCfast Interface

The NLO + PS Case: Attempting a First PDF Fit

- Compare χ^2 before and after the fit:

Dataset	BL+ATLAS	BL
Average y1216 101	372 / 18	449 / 18
Average y1620 101	516 / 18	583 / 18
Average y0408 101	520 / 18	798 / 18
Average y0812 101	429 / 18	731 / 18
Average y2024 101	1299 / 18	1750 / 18
Average y0004 101	356 / 18	693 / 18

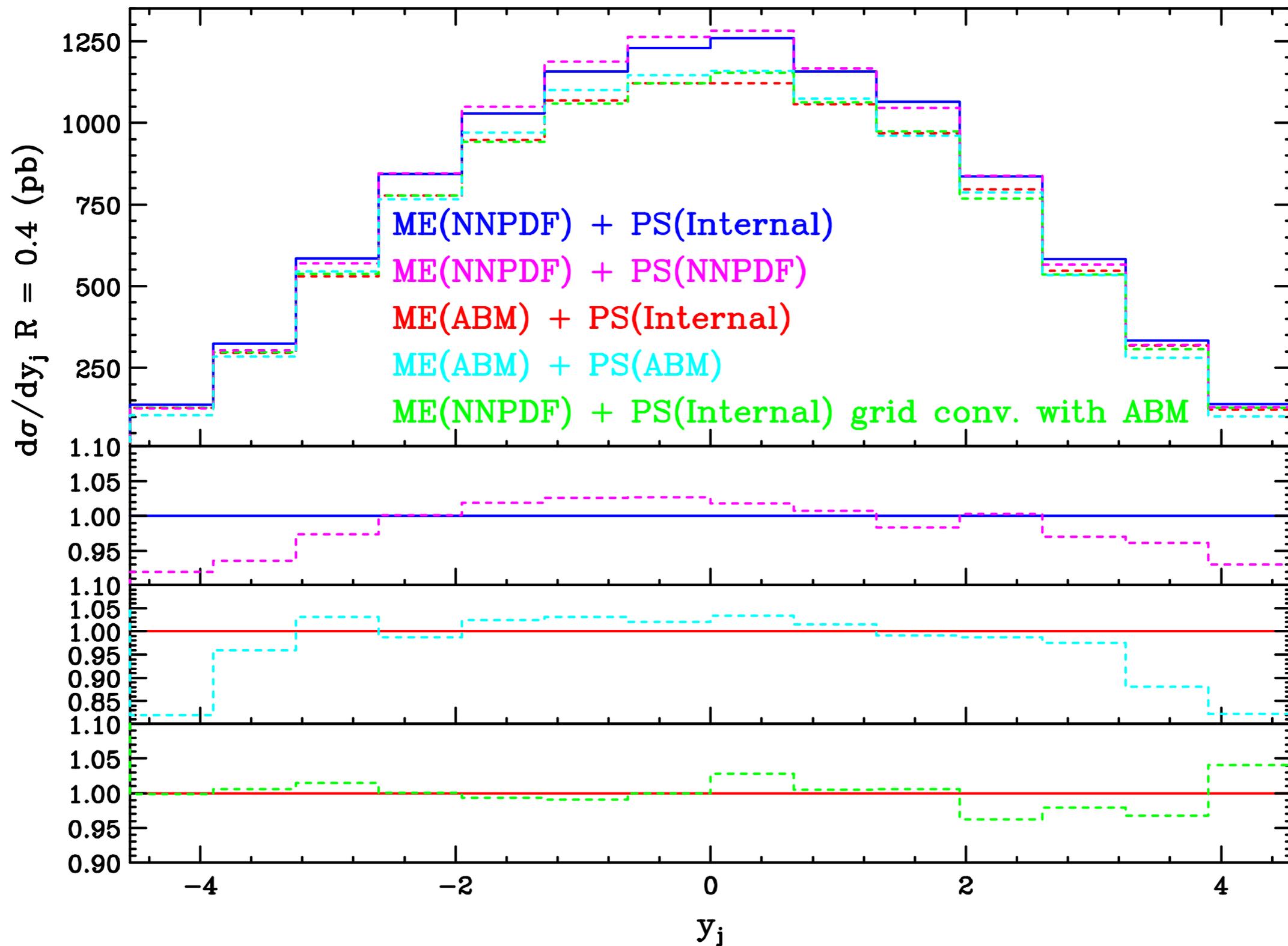
- The fit does help reduce the χ^2 but the description is far from being acceptable.
- Very substantial effect on PDFs.
- Very preliminary results... this is not the last word.

Backup Slides

The aMCfast Interface

The NLO + PS Case: Preliminary Results

Prediction for $e^+ \nu$ production



Fast NLO Computations

The Interpolating Grids

- The basic idea is that of a Lagrange-polynomial expansion:

$$F(z) = \sum_i F(z_i) I_i^{(s)}(z)$$

Grid nodes

Interpolation functions

Fast NLO Computations

The Interpolating Grids

- The basic idea is that of a Lagrange-polynomial expansion:

$$F(z) = \sum_i F(z_i) I_i^{(s)}(z)$$

- Suppose you want to compute numerically the following integral, e.g. by Monte Carlo methods:

$$J = \int_a^b dz F(z) S(z) = \sum_{k=1}^M \Phi_k F(z_k) S(z_k)$$

Normalization factor

Random points in the interval [a,b]

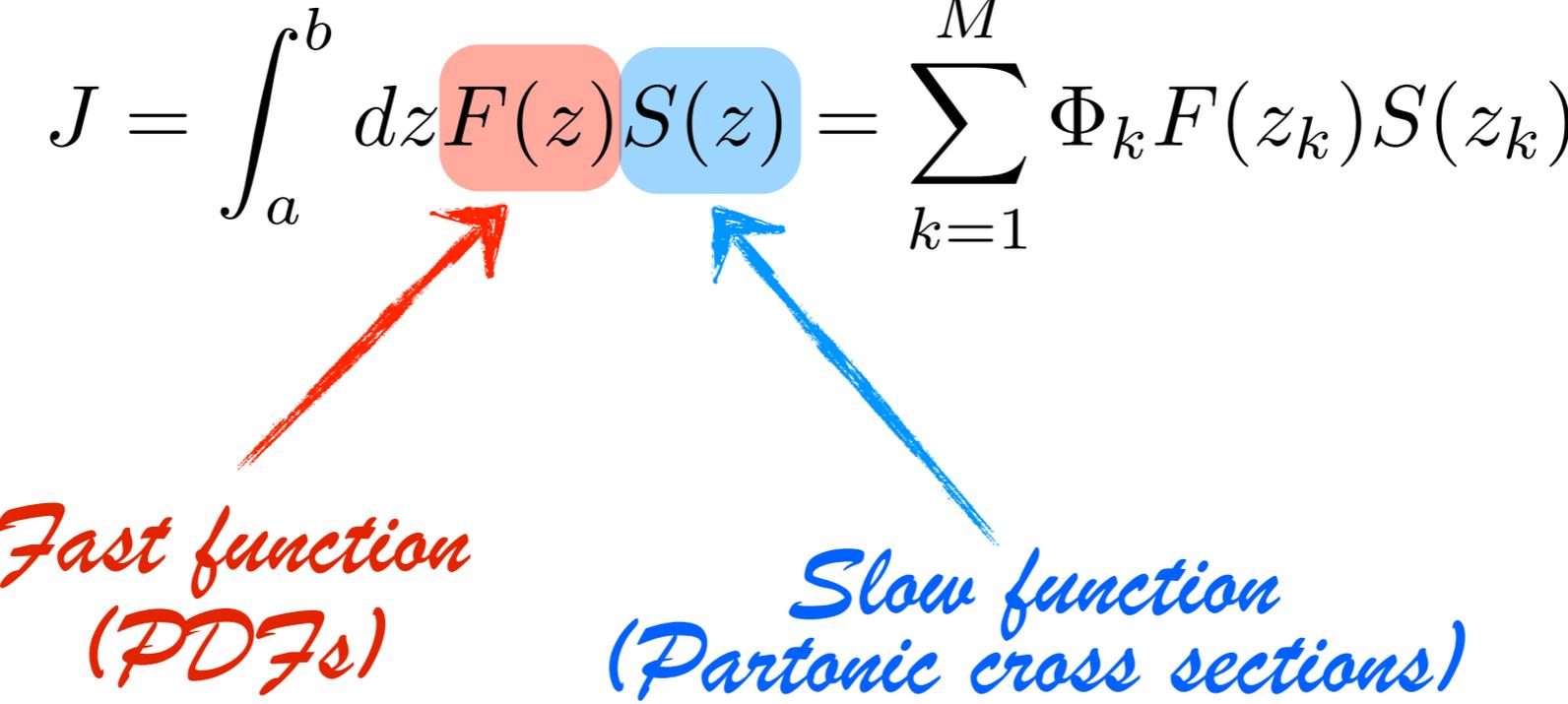
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*Fast function
(PDFs)*

*Slow function
(Partonic cross sections)*

Fast NLO Computations

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- Using the interpolation formula:

$$J = \sum_i F(z_i) G_i \quad \text{with} \quad G_i = \sum_{k=1}^M \Phi_k S(z_k) I_i^{(s)}(z_k)$$


*(1-dimensional) interpolation grid independent of $F(z)$:
precomputed and stored*

Fast NLO Computations

The Interpolating Grids

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$$J = \sum_i F(z_i) G_i \quad \text{with} \quad G_i = \sum_{k=1}^M \Phi_k S(z_k) I_i^{(s)}(z_k)$$

- Once G_i has been precomputed, the *a posteriori* computation of J with any function $F(z)$ will be extremely fast.

Fast NLO Computations

The Hard Cross Sections in aMC@NLO at NLO

- The generalization of this procedure to the realistic case of a hard NLO cross section is straightforward, considering that:

$$d\sigma^{(\text{NLO})} \longleftrightarrow \left\{ d\sigma^{(\text{NLO},\alpha)} \right\}_{\alpha=E,S,C,SC} \leftarrow \text{Event \& Counterevents}$$

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$$d\sigma^{(\text{NLO})} \longleftrightarrow \left\{ d\sigma^{(\text{NLO},\alpha)} \right\}_{\alpha=E,S,C,SC}$$

$$d\sigma^{(\text{NLO},\alpha)} = f_1(x_1^{(\alpha)}, \mu_F^{(\alpha)}) f_2(x_2^{(\alpha)}, \mu_F^{(\alpha)}) W^{(\alpha)} d\chi_{Bj} d\chi_{n+1}$$

PDFs

Partonic cross section

Fast NLO Computations

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$$W^{(\alpha)} = \alpha_s^{b+1}(\mu_R^{(\alpha)}) \left[W_0^{(\alpha)} + W_F^{(\alpha)} \ln \left(\frac{\mu_F^{(\alpha)}}{Q} \right) + W_R^{(\alpha)} \ln \left(\frac{\mu_R^{(\alpha)}}{Q} \right) \right] + \alpha_s^b(\mu_R^{(\alpha)}) W_B \delta_{\alpha S}$$

NLO term

Born term

Fast NLO Computations

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$$d\sigma^{(\text{NLO})} \longleftrightarrow \left\{ d\sigma^{(\text{NLO},\alpha)} \right\}_{\alpha=E,S,C,SC} \quad \text{Slow functions}$$

$$d\sigma^{(\text{NLO},\alpha)} = f_1(x_1^{(\alpha)}, \mu_F^{(\alpha)}) f_2(x_2^{(\alpha)}, \mu_F^{(\alpha)}) W^{(\alpha)} d\chi_{Bj} d\chi_{n+1} \quad \text{Fast functions}$$

$$W^{(\alpha)} = \alpha_s^{b+1}(\mu_R^{(\alpha)}) \left[W_0^{(\alpha)} + W_F^{(\alpha)} \ln \left(\frac{\mu_F^{(\alpha)}}{Q} \right) + W_R^{(\alpha)} \ln \left(\frac{\mu_R^{(\alpha)}}{Q} \right) \right] \\ + \alpha_s^b(\mu_R^{(\alpha)}) W_B \delta_{\alpha S}$$

Fast NLO Computations

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$$d\sigma^{(\text{NLO})} \longleftrightarrow \left\{ d\sigma^{(\text{NLO},\alpha)} \right\}_{\alpha=E,S,C,SC} \quad \text{Slow functions}$$

$$d\sigma^{(\text{NLO},\alpha)} = f_1(x_1^{(\alpha)}, \mu_F^{(\alpha)}) f_2(x_2^{(\alpha)}, \mu_F^{(\alpha)}) W^{(\alpha)} d\chi_{Bj} d\chi_{n+1} \quad \text{Fast functions}$$

$$W^{(\alpha)} = \alpha_s^{b+1}(\mu_R^{(\alpha)}) \left[W_0^{(\alpha)} + W_F^{(\alpha)} \ln \left(\frac{\mu_F^{(\alpha)}}{Q} \right) + W_R^{(\alpha)} \ln \left(\frac{\mu_R^{(\alpha)}}{Q} \right) \right] \\ + \alpha_s^b(\mu_R^{(\alpha)}) W_B \delta_{\alpha S}$$

- 4 **slow functions** \Rightarrow 4 **interpolation grids**.

- The **fast functions** are functions of 4 independent variables $(x_1, x_2, \mu_F, \mu_R) \Rightarrow$ 4-dimensional interpolation grids needed.

Fast NLO Computations

The Hard Cross Sections in aMC@NLO at NLO

- The generalization of this procedure to the realistic case of a hard NLO cross section is straightforward, considering that:

$$d\sigma^{(\text{NLO})} \longleftrightarrow \left\{ d\sigma^{(\text{NLO},\alpha)} \right\}_{\alpha=E,S,C,SC} \quad \text{Slow functions}$$

$$d\sigma^{(\text{NLO},\alpha)} = f_1(x_1^{(\alpha)}, \mu_F^{(\alpha)}) f_2(x_2^{(\alpha)}, \mu_F^{(\alpha)}) W^{(\alpha)} d\chi_{Bj} d\chi_{n+1} \quad \text{Fast functions}$$

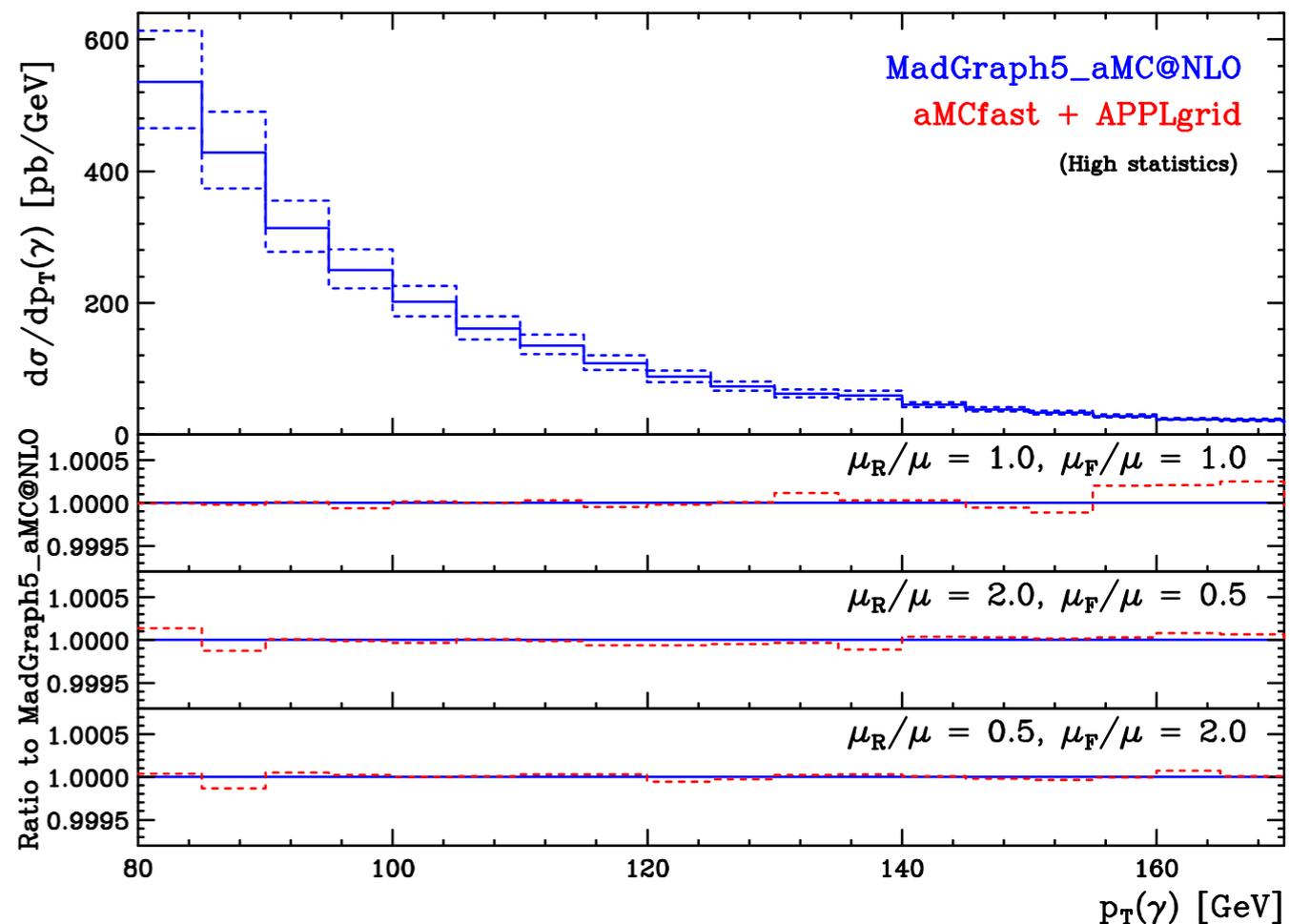
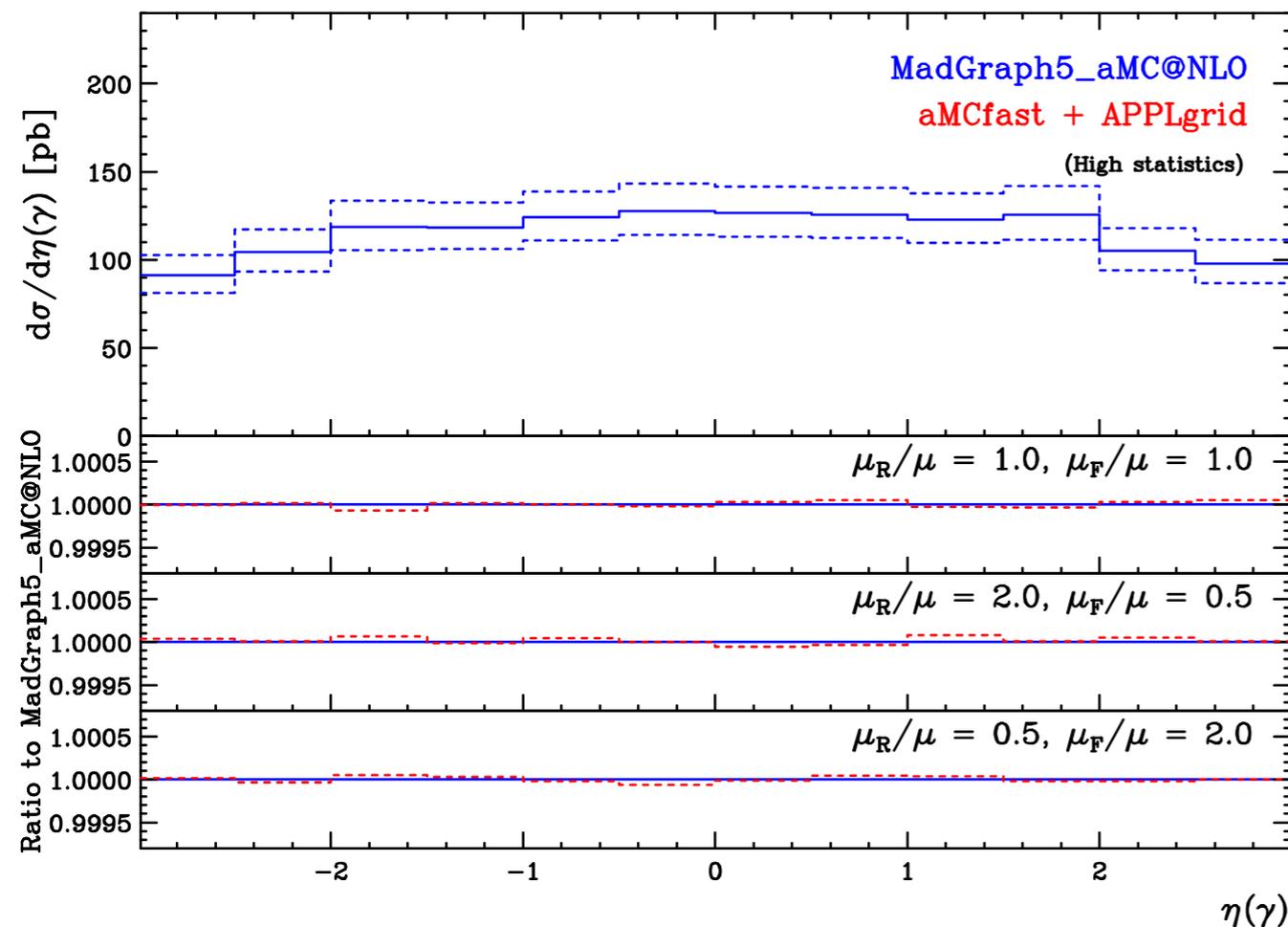
$$W^{(\alpha)} = \alpha_s^{b+1}(\mu_R^{(\alpha)}) \left[W_0^{(\alpha)} + W_F^{(\alpha)} \ln \left(\frac{\mu_F^{(\alpha)}}{Q} \right) + W_R^{(\alpha)} \ln \left(\frac{\mu_R^{(\alpha)}}{Q} \right) \right] \\ + \alpha_s^b(\mu_R^{(\alpha)}) W_B \delta_{\alpha S}$$

- 4 **slow functions** \Rightarrow 4 **interpolation grids**.
- The **fast functions** are functions of 4 independent variables $(x_1, x_2, \mu_F, \mu_R) \Rightarrow$ 4-dimensional interpolation grids needed.
- But assuming $\mu_F \propto \mu_R \Rightarrow$ **3-dimensional** interpolation grids.

The aMCfast Interface

Validation: Photon Production with one Jet

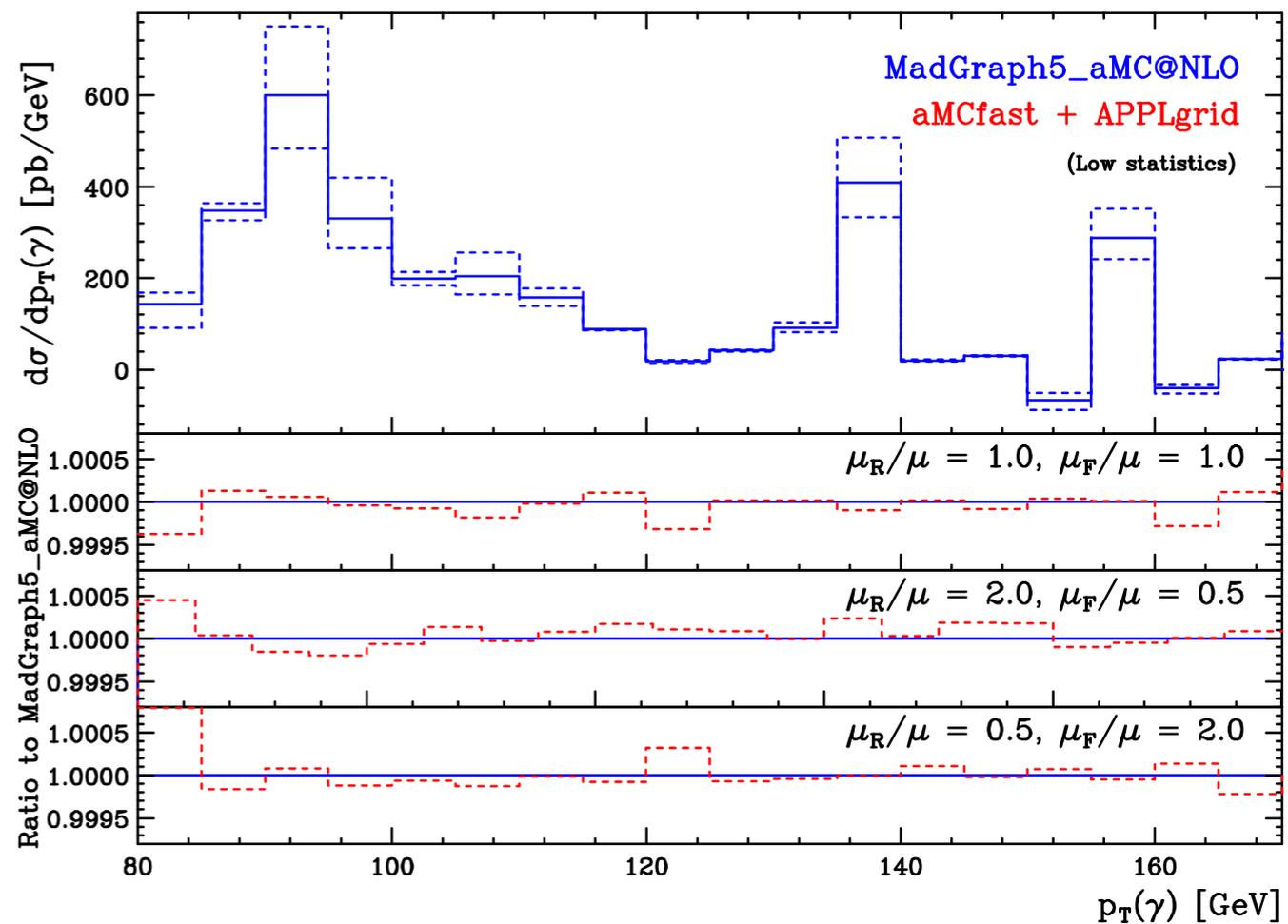
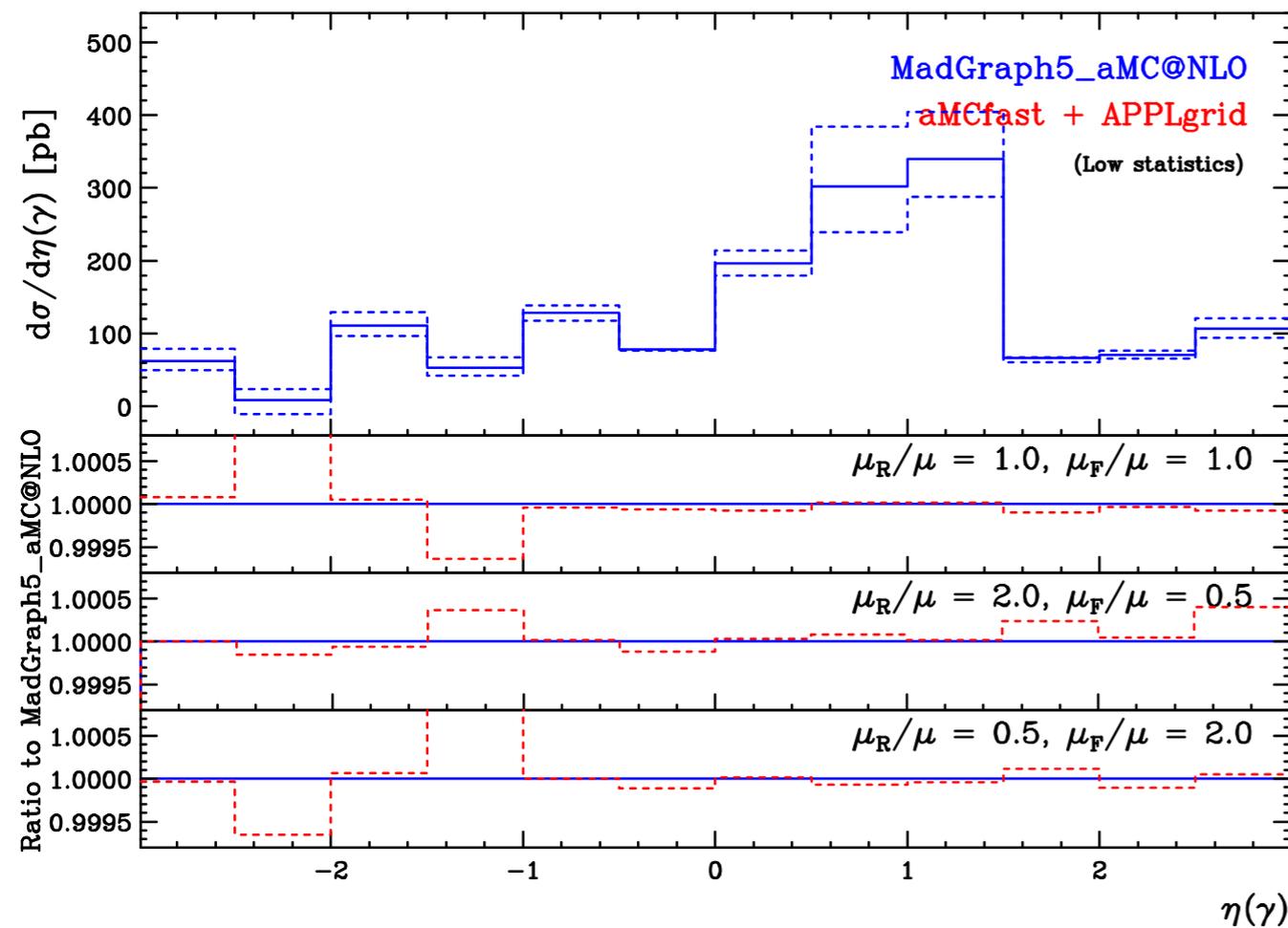
- Important for the gluon in the region relevant for Higgs production in gluon fusion.
- We looked at the following observables:
 - the pseudo-rapidity distribution of the photon (left),
 - the transverse momentum distribution of the photon (right).
- High statistics plots:



The aMCfast Interface

Validation: Photon Production with one Jet

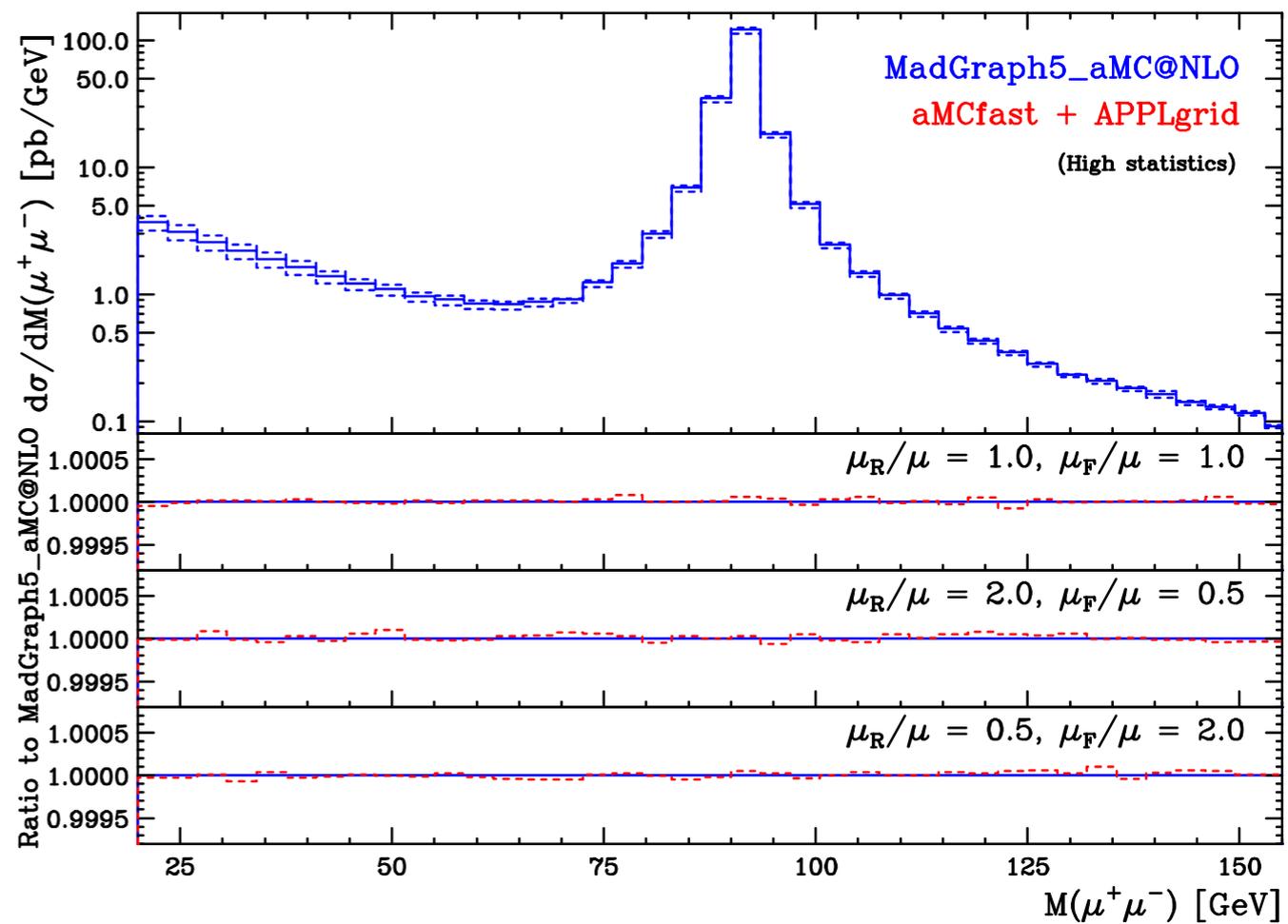
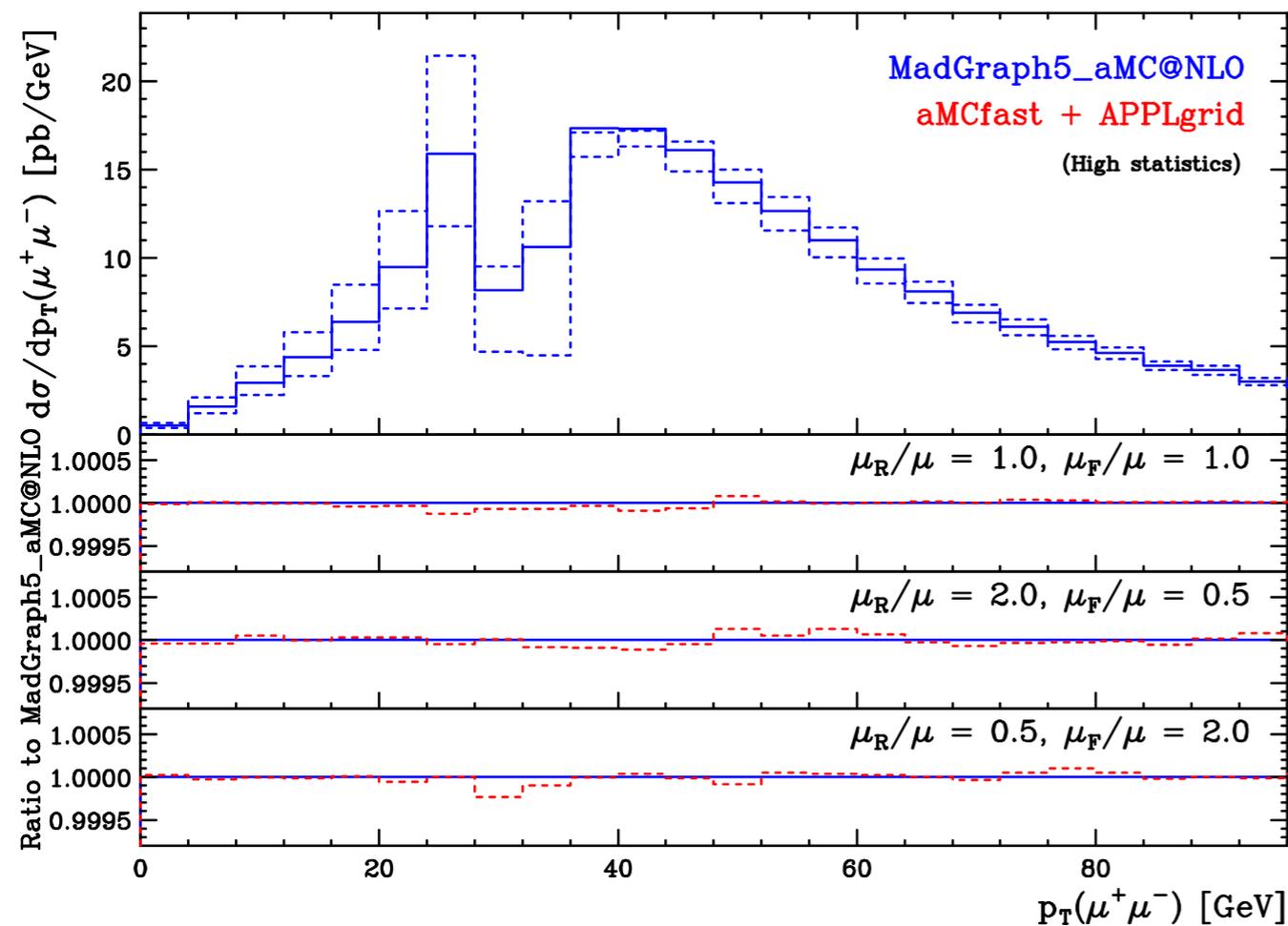
- Important for the gluon in the region relevant for Higgs production in gluon fusion.
- We looked at the following observables:
 - the pseudo-rapidity distribution of the photon (left),
 - the transverse momentum distribution of the photon (right).
- Low statistics plots:



The aMCfast Interface

Validation: Dilepton Production with one Jet

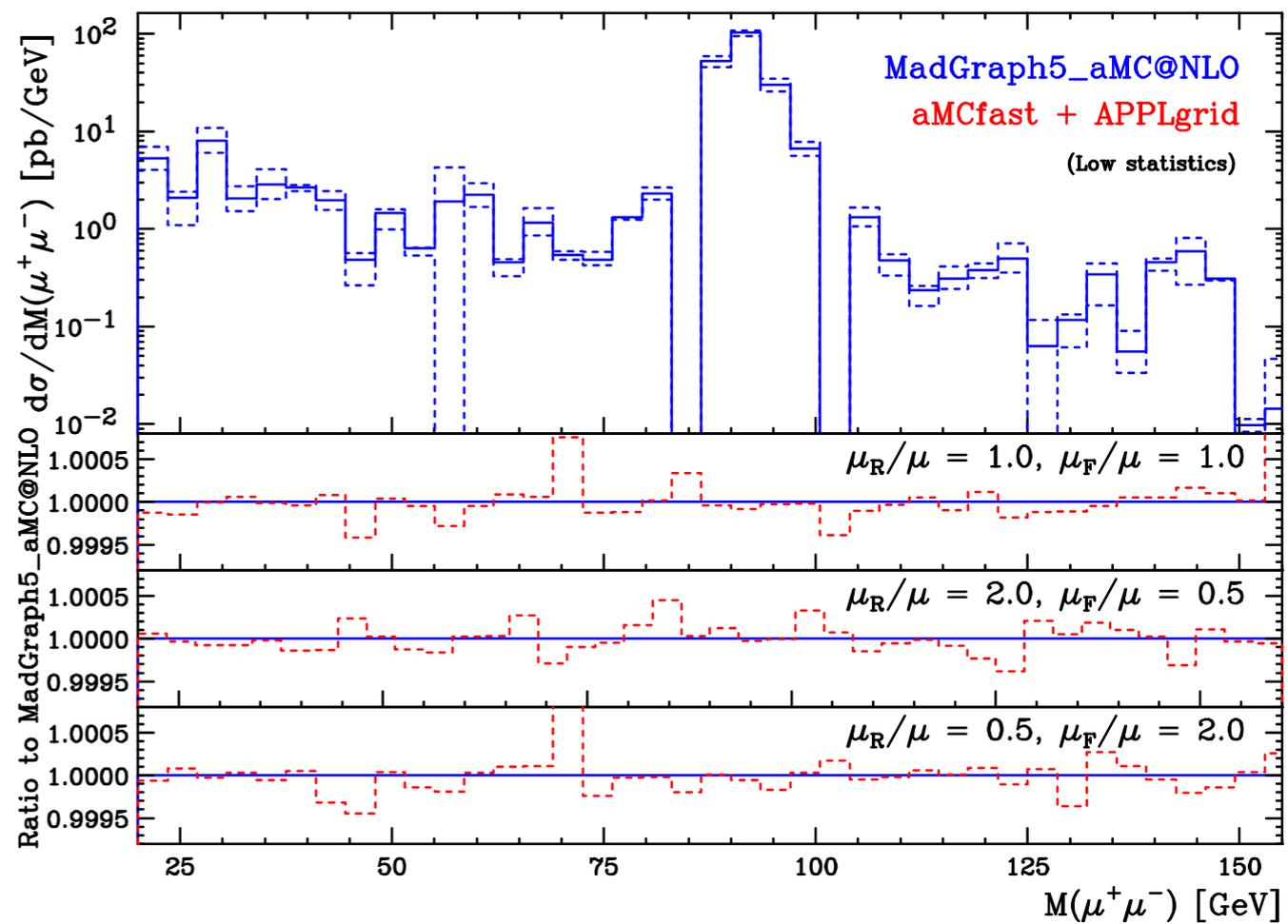
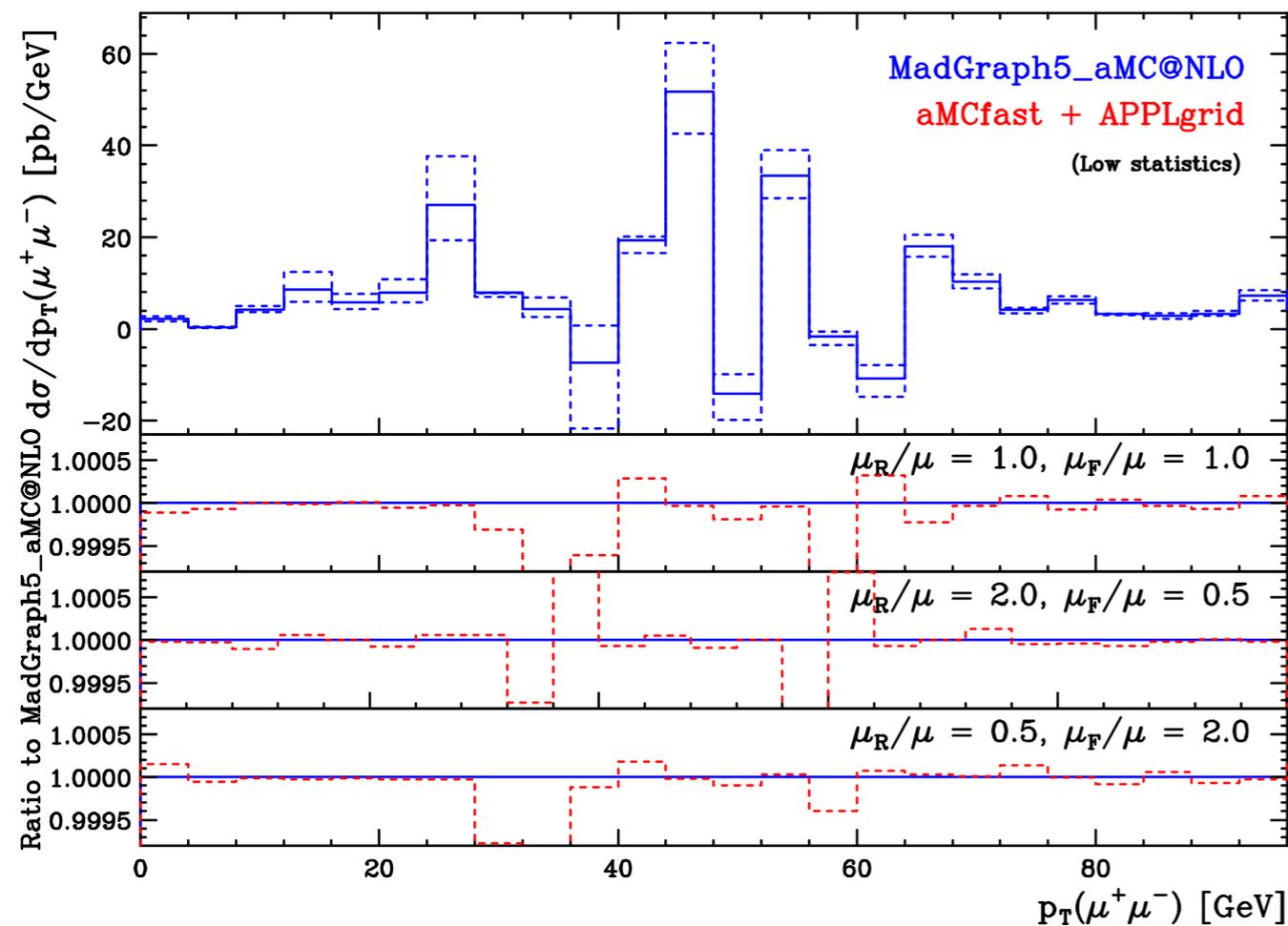
- Relevant for quarks and antiquarks in the large- x region.
- We looked at the following observables:
 - the transverse momentum distribution of the lepton pair (left),
 - the invariant mass distribution of the lepton pair (right).
- High statistics plots:



The aMCfast Interface

Validation: Dilepton Production with one Jet

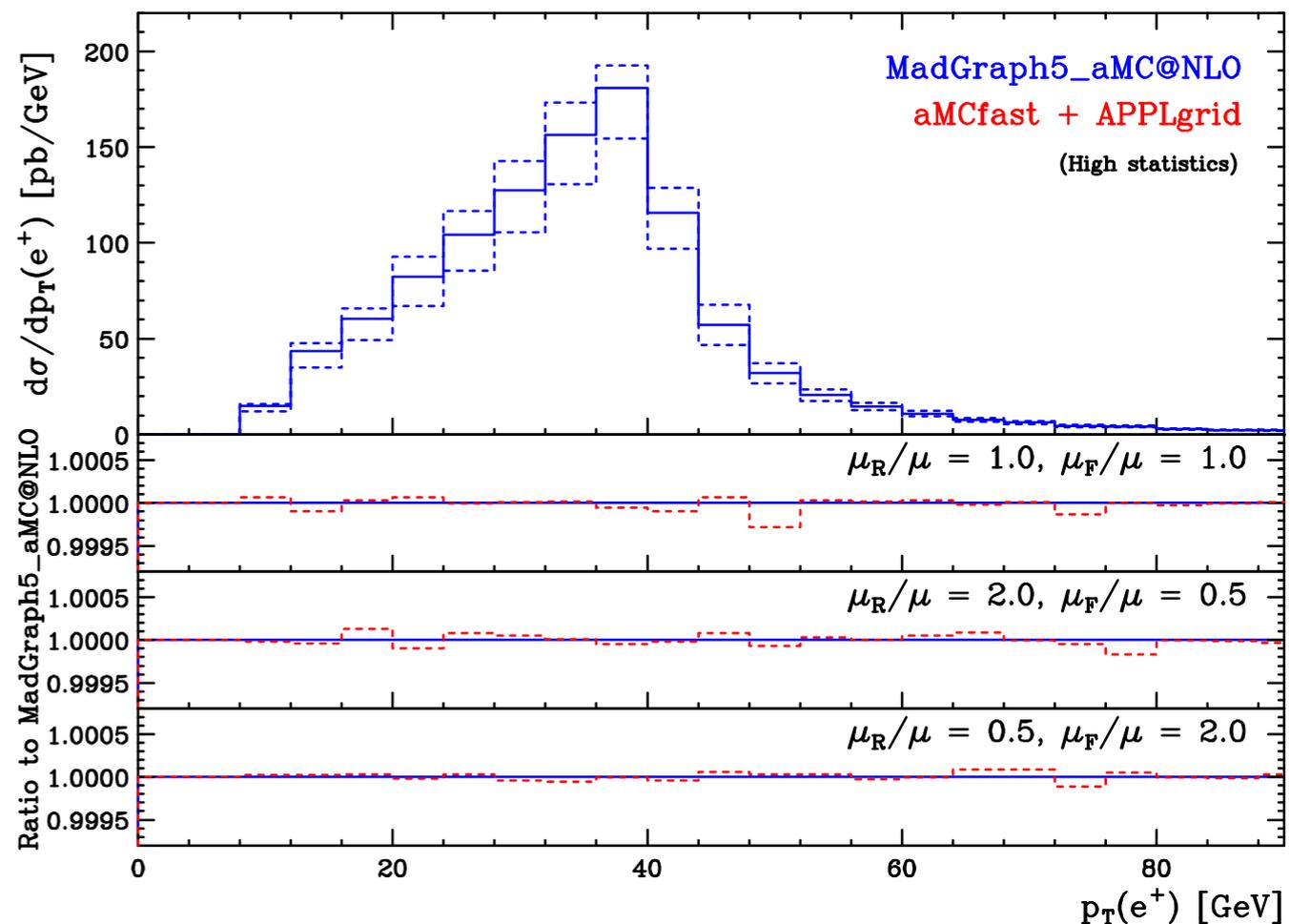
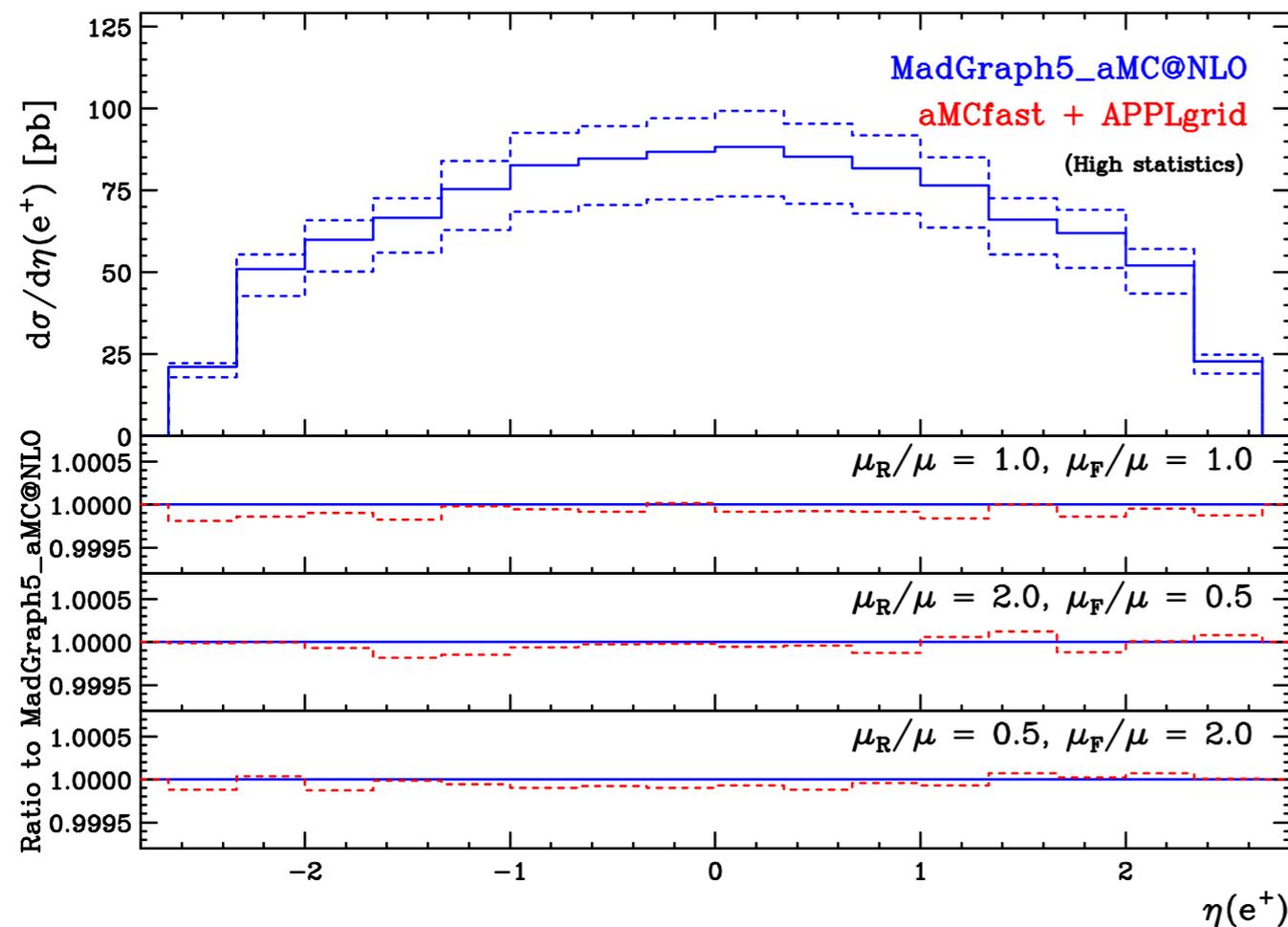
- Relevant for quarks and antiquarks in the large- x region.
- We looked at the following observables:
 - the transverse momentum distribution of the lepton pair (left),
 - the invariant mass distribution of the lepton pair (right).
- Low statistics plots:



The aMCfast Interface

Validation: $W + c$ Production

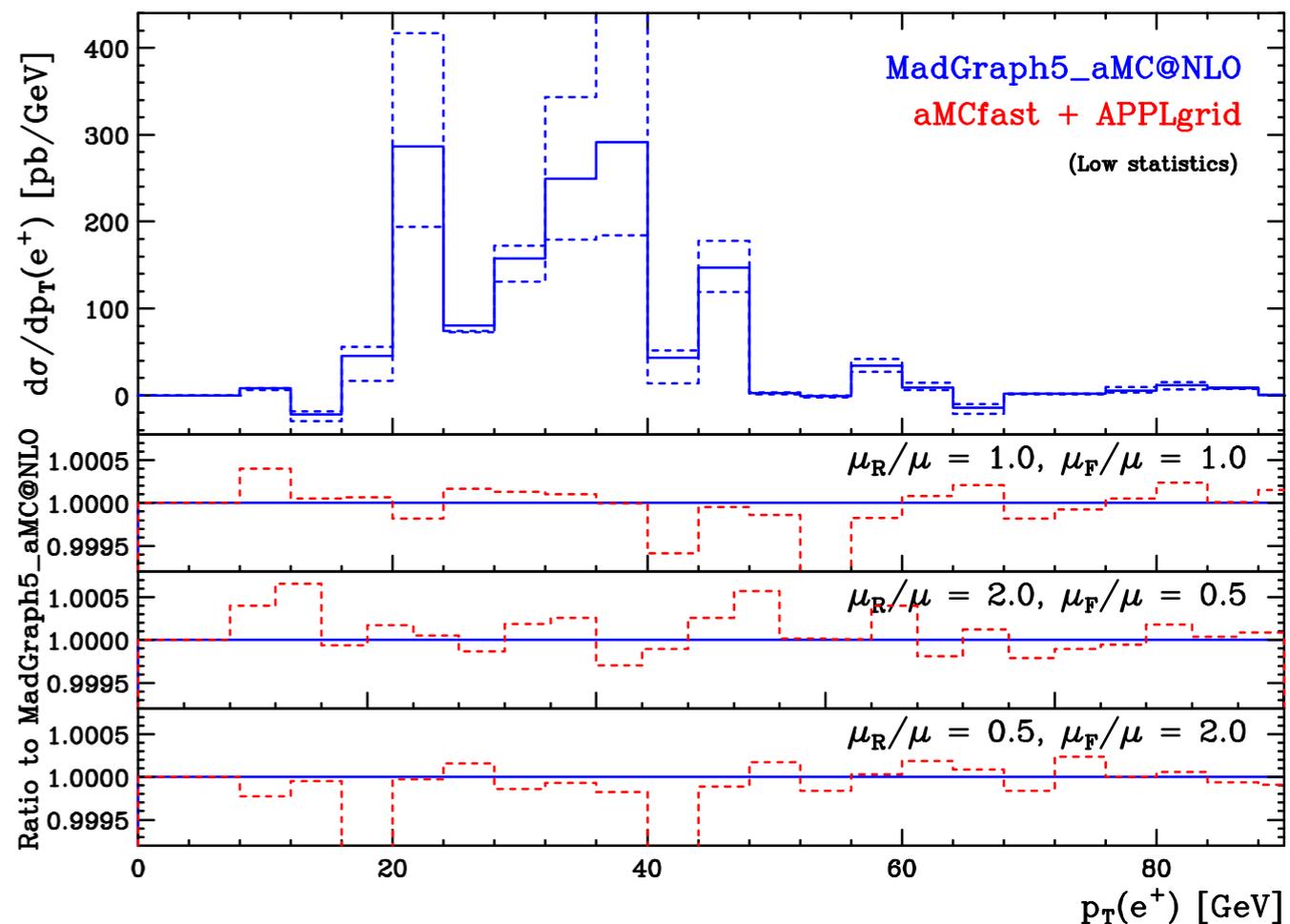
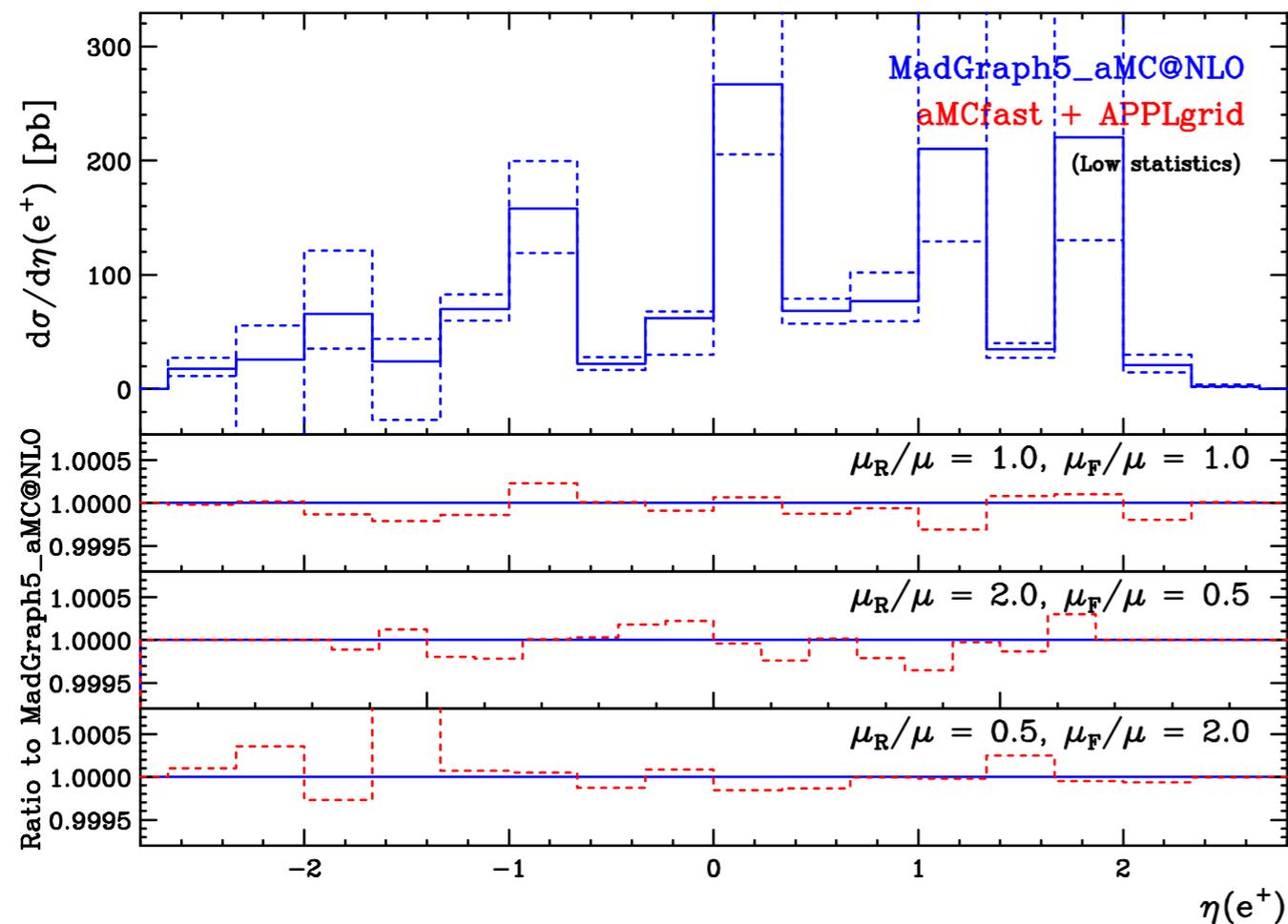
- Important for constraining the strange-quark PDFs.
- We looked at the following observables:
 - the pseudo-rapidity distribution of the lepton (left),
 - the transverse momentum distribution of the lepton (right).
- High statistics plots:



The aMCfast Interface

Validation: $W + c$ Production

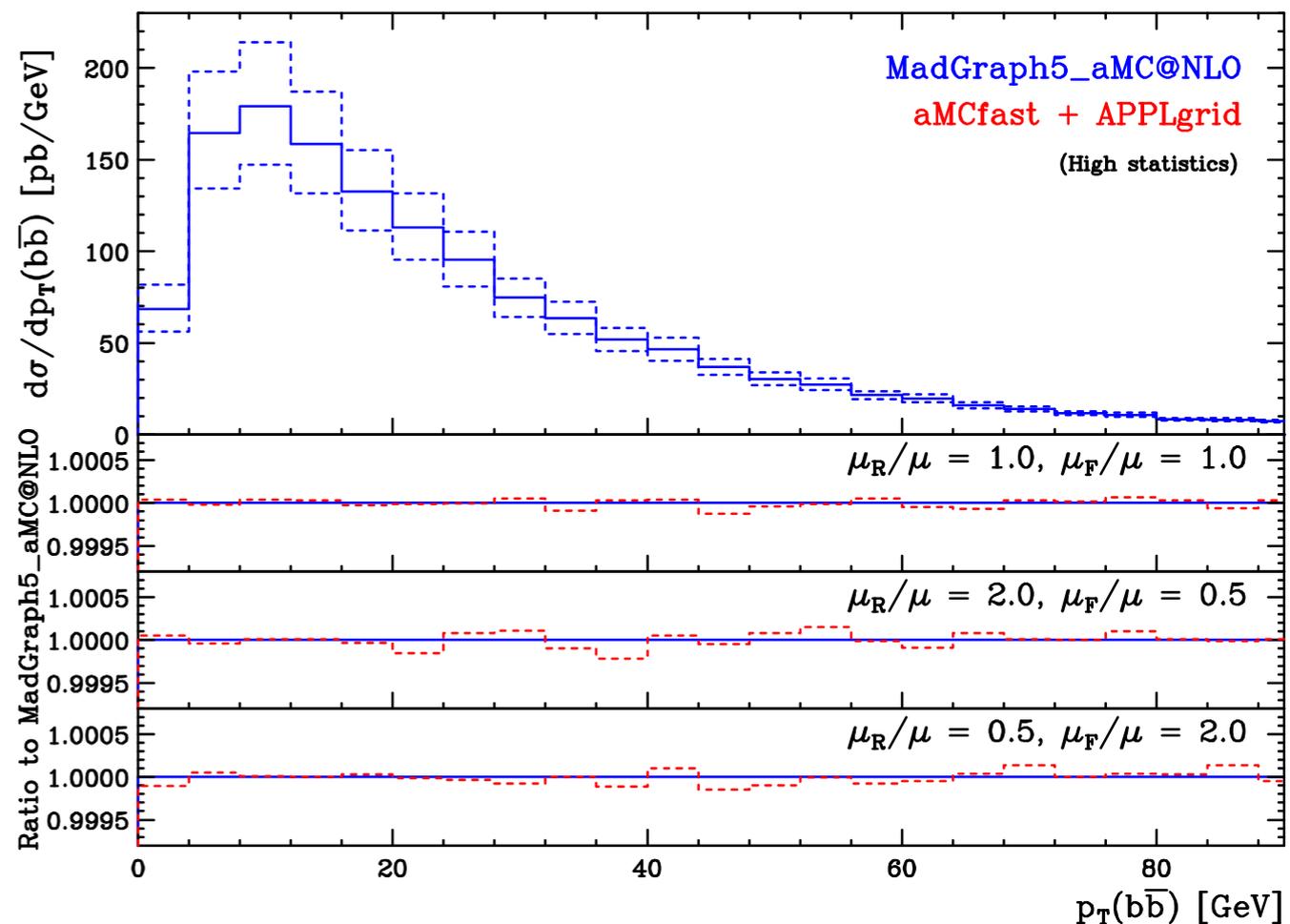
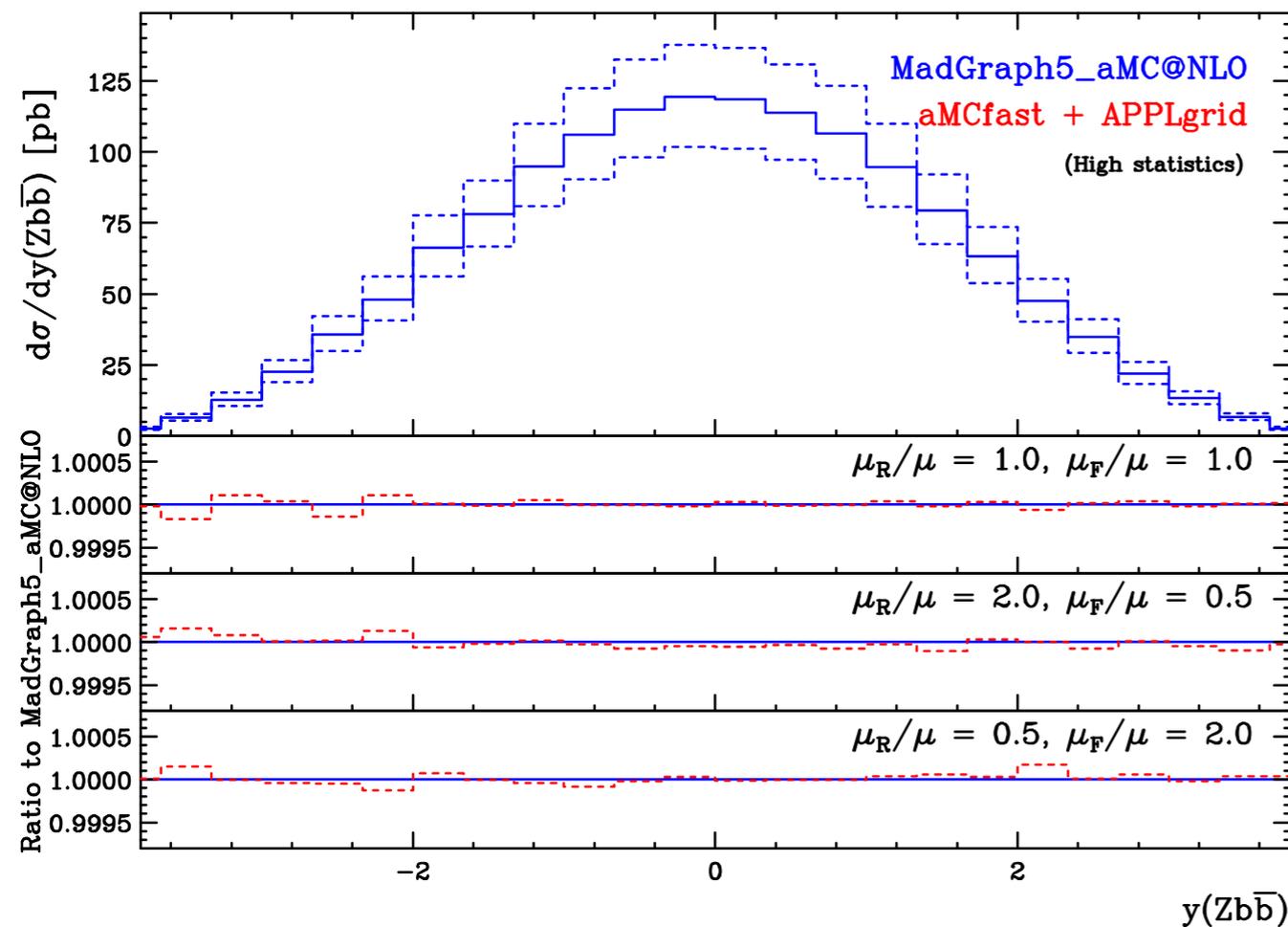
- Important for constraining the strange-quark PDFs.
- We looked at the following observables:
 - the pseudo-rapidity distribution of the lepton (left),
 - the transverse momentum distribution of the lepton (right).
- Low statistics plots:



The aMCfast Interface

Validation: $Z + bb$ Production

- This is just an example of complicated process.
- We looked at the following observables:
 - the rapidity distribution of the Zbb system (left),
 - the transverse momentum distribution of the Zbb system (right).
- High statistics plots:



The aMCfast Interface

Validation: $Z + bb$ Production

- This is just an example of complicated process.
- We looked at the following observables:
 - the rapidity distribution of the Zbb system (left),
 - the transverse momentum distribution of the Zbb system (right).
- Low statistics plots:

