

Self-Seeding implementation at European XFEL

Shan Liu, DESY

on behalf of the HXRSS collaboration team

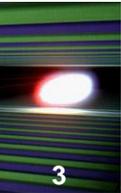


- **HXRSS simulations:** G. Geloni, V. Kocharyan, E. Saldin, S. Tomin, S. Serkez et al.
- **Monochromator:** H. Sinn, X. Dong, L. Samoylova, T. Wohlenberg, D. Shu (ANL), V. Blank (TISNUM), S. Terentiev (TISNUM) et al.
- **Chicane magnets and girder:** W. Decking, S. Liu, T. Wohlenberg, B. Krause, A. Petrov, N. Golubeva, C. Engling et al.

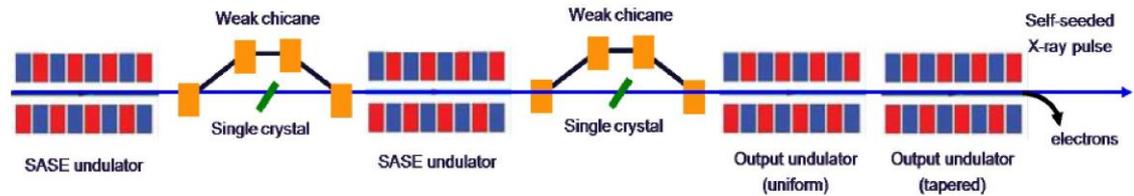
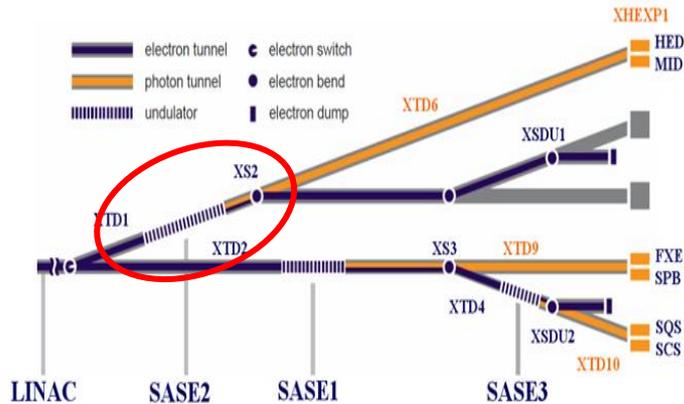
Outlines

- **Why** Hard X-Ray Self-Seeding (HXRSS) at European XFEL?
- **How** HXRSS designed at European XFEL?
- **What** are the challenges for HXRSS?
- Summary and discussions

HXRSS at European EXFEL



G. Geloni, V. Kocharyan, E. Saldin (DESY 10-133)



SASE2 line (3 keV -25 keV) will be first equipped with HXRSS

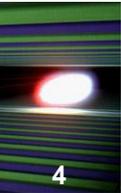
- High repetition-rate
- Long undulators → HXRSS+tapering

Simulations: (1) HXRSS – 9 keV – taper, (2) HXRSS – 14.4 keV- 2nd harm. bunching – taper

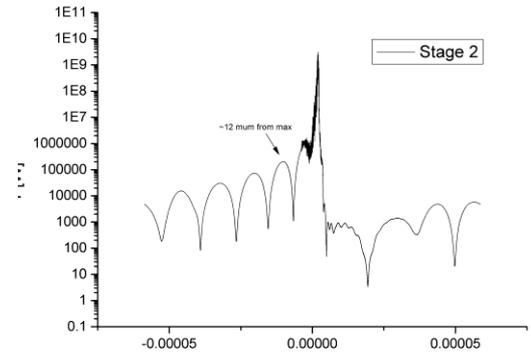
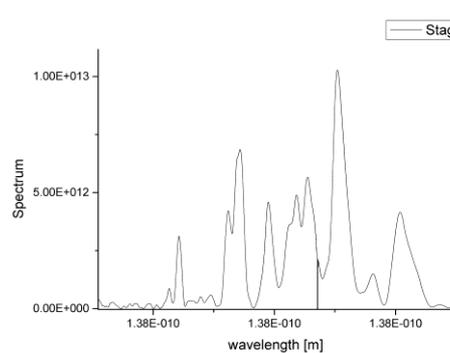
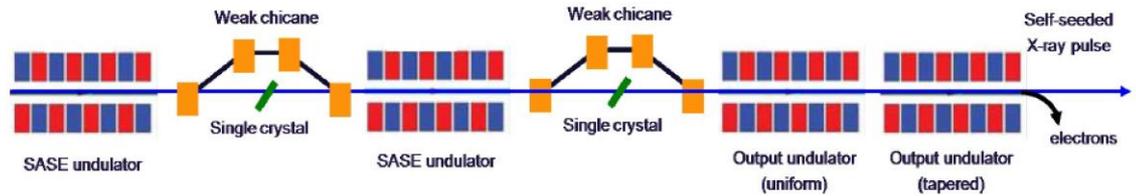
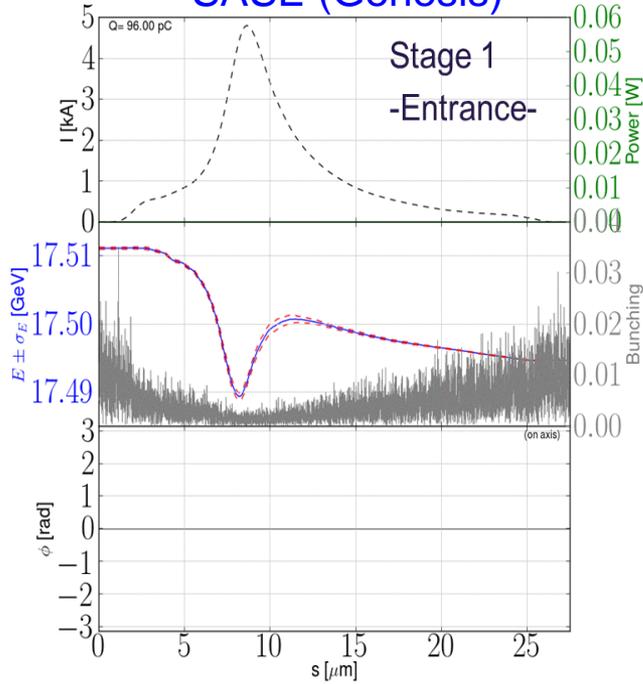
	Photon energy, [keV]	Integrated flux, [Ph/ pulse]	Photon pulse BW	Photon Flux, [Ph/s/meV]
w/o HXRSS	9	7e11	$\Delta\lambda/\lambda \sim 1.2e-3$ or $\sim 12eV$	1.5e12
w/ HXRSS (1)	9	7e12	$\Delta\lambda/\lambda \sim 1e-4$ or $\sim 940meV$	2.1e14
w/o HXRSS	14.4	1e11	$\Delta\lambda/\lambda \sim 1.6e-3$ or $\sim 20eV$	1e11
w/ HXRSS (2)	14.4	9e11	$\Delta\lambda/\lambda \sim 7e-5$ or $\sim 1eV$	4e13

(1) G. Geloni, V. Kocharyan and E. Saldin DESY 15-141; (2) O. Chubar, G. Geloni et al. J. Synchrotron Rad. 23 (2016).

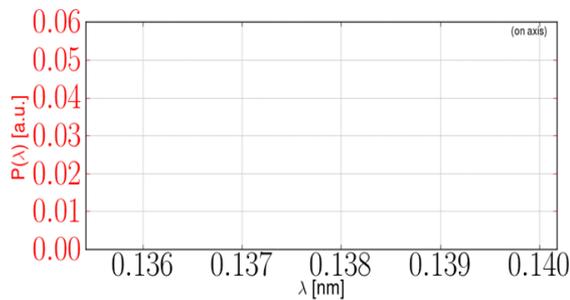
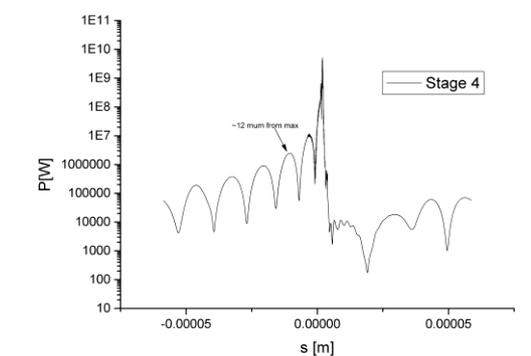
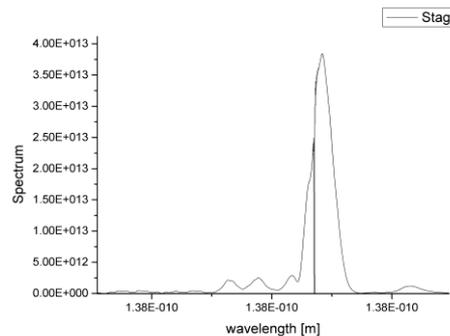
HXRSS Simulations



SASE (Genesis)



Seeding Stages (Ocelot)

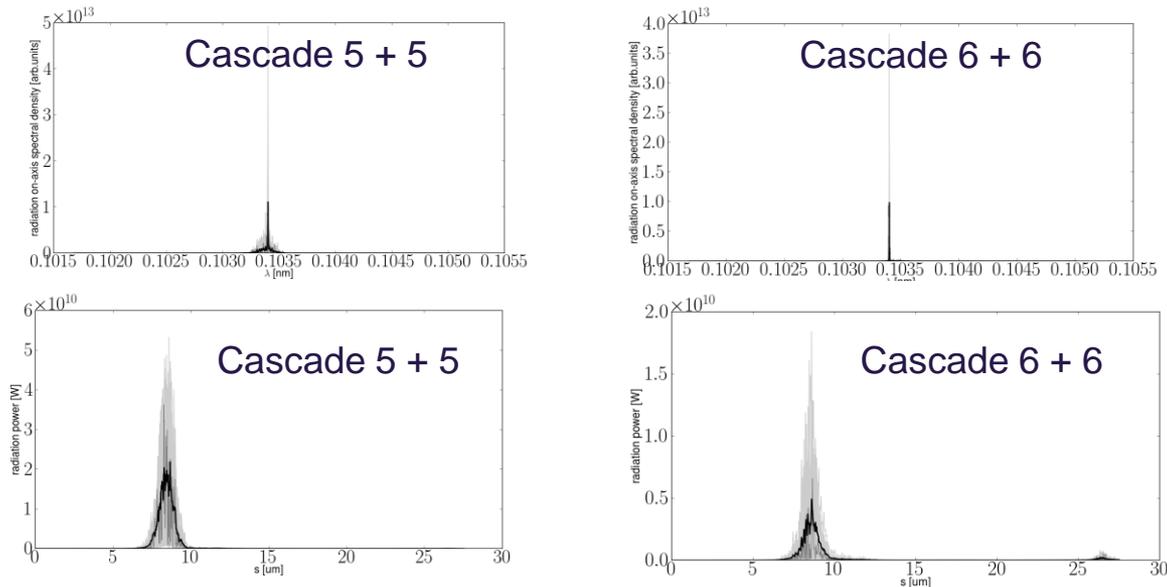


Courtesy of G. Geloni

HXRSS performance drops at **high energies** due to

- Spatiotemporal coupling for low Miller indexes
- Larger equivalent SASE shot noise to beat
- Longer gain length → **increase # modules before HXRSS** can improve the situation

100pC – 17.5 GeV – C004 – 12keV – 5th stage

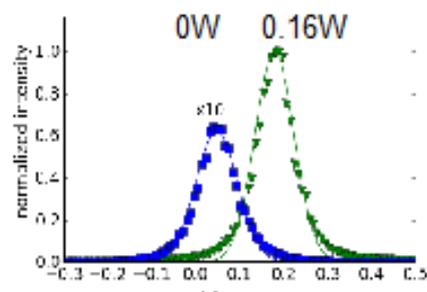


Courtesy of G. Geloni

Heat load on crystals by spontaneous radiation **also increases!**

High repetition rate: **27000 bunches/s** Heat load is an issue!

- **Spontaneous emission radiation** has a broad energy spectrum, it is considered as the main source of heat load.
- **Calculated total energy deposition** from spontaneous emission: $\sim 6 \mu\text{J}$.



X-ray reflection
energy shift
due to heat load
 $\Delta E/E = 3.1 \text{ e-4}$

Experiment by
L. Samoylova

Si(440) analyzer crystal scans

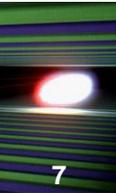
Courtesy of *H. Sinn*



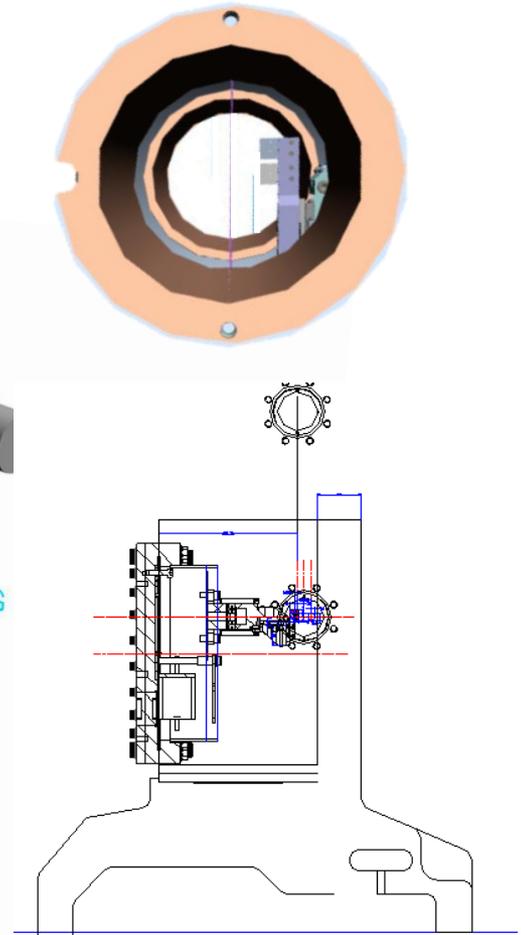
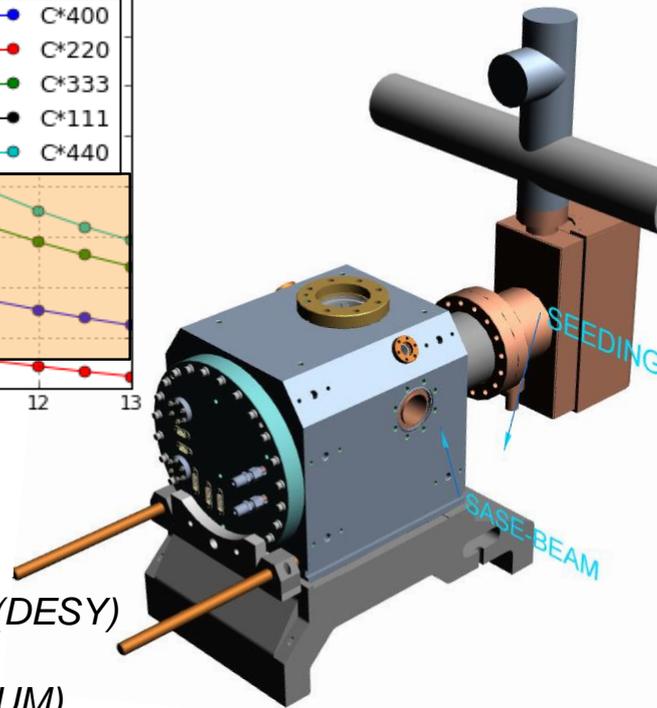
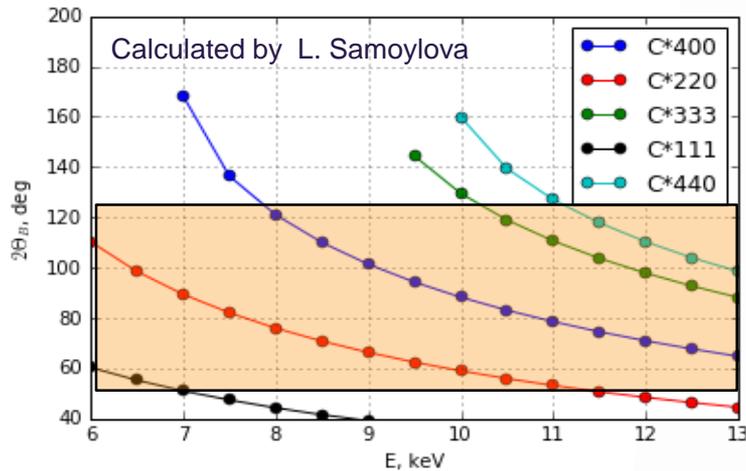
Frequency (Hz)	Angular Range (PV)
1200	70 urad
...	...
300	280 urad
...	...
10	0.48°

- **Pitch oscillator** will be treated as option (space foreseen and some development within design contract).
- **Oscillate bragg angle** can be used to compensate temperature cycle during pulse train.

Monochromator Design

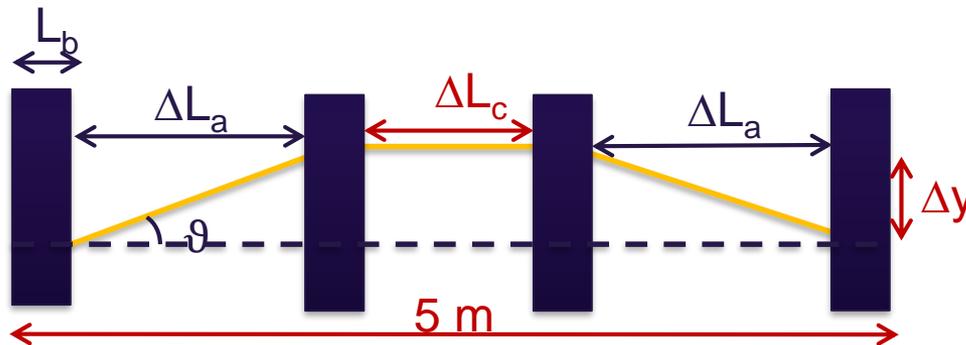
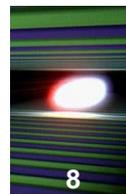


- View window for reflected signal detected by YAG screen: window angle at $90^\circ \pm 33.5^\circ$ view covers all the reflections
- Two crystals with different orientations on one holder



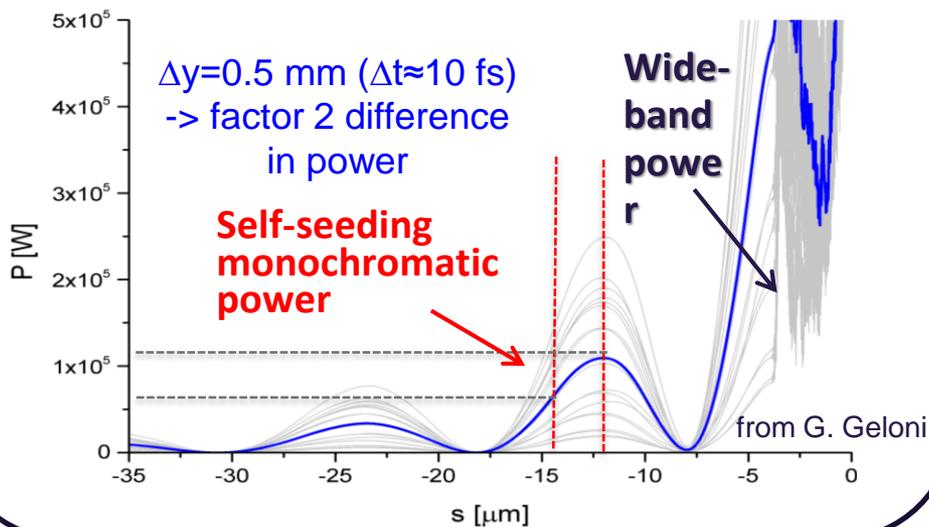
- Chamber design: *T. Wohlenberg (DESY)*
- Diamond and holder design:
D. Shu (ANL), S. Terentiev (TISNUM)

HXRSS Chicane design



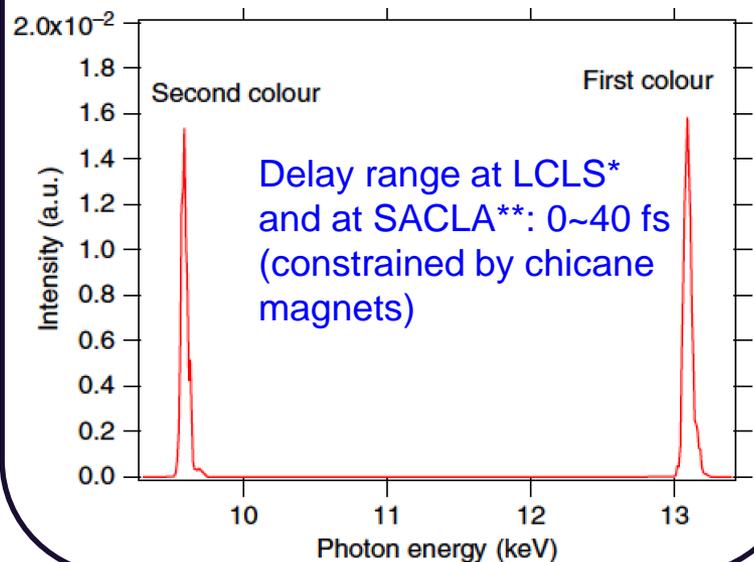
Minimum delay

(<40 fs preferred, constrained by min. e- beam offset)



Maximum delay

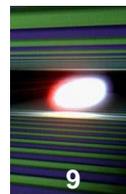
(> 400 fs preferred for **2-colour SASE**)



*A. A. Lutman et al., PRL 110, 134801 (2013)

** Toru Hara et al., Nature communications 4 (2013)

Chicane Design: maximum delay

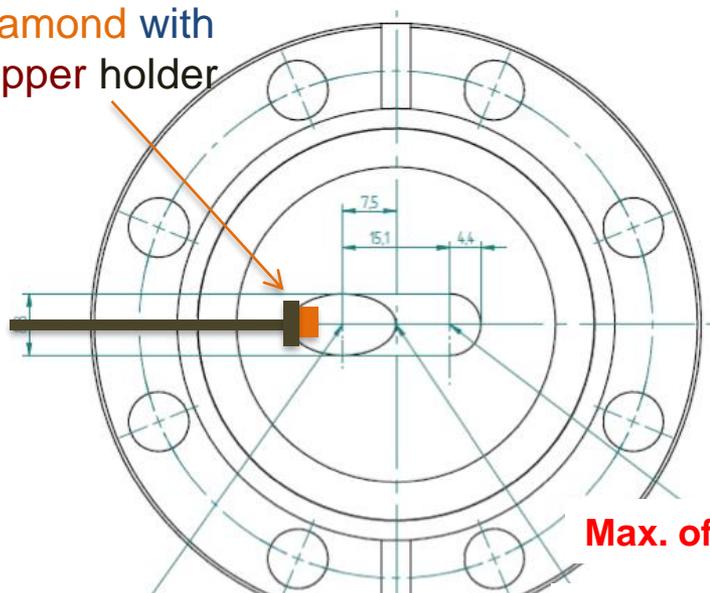


9

Chicane chamber center is shifted to insure a maximum offset of 15 mm

436 fs max. delay can be achieved at European XFEL for <11.6 GeV beam

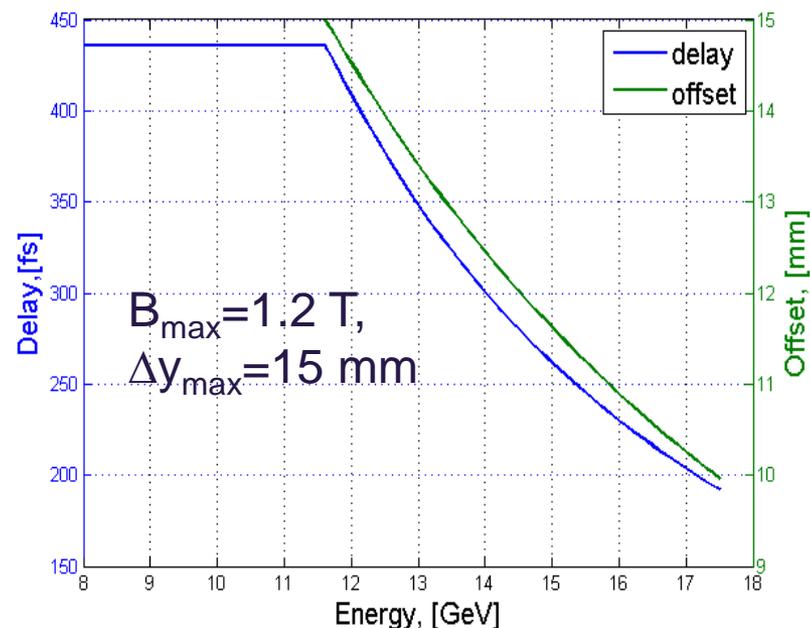
Diamond with copper holder



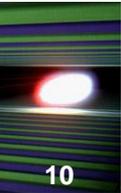
Max. offset of 15 mm

Center of undulator system
(beam w/o seeding)

Center of chicane
(dipole magnets, BPM
and Monochromator)

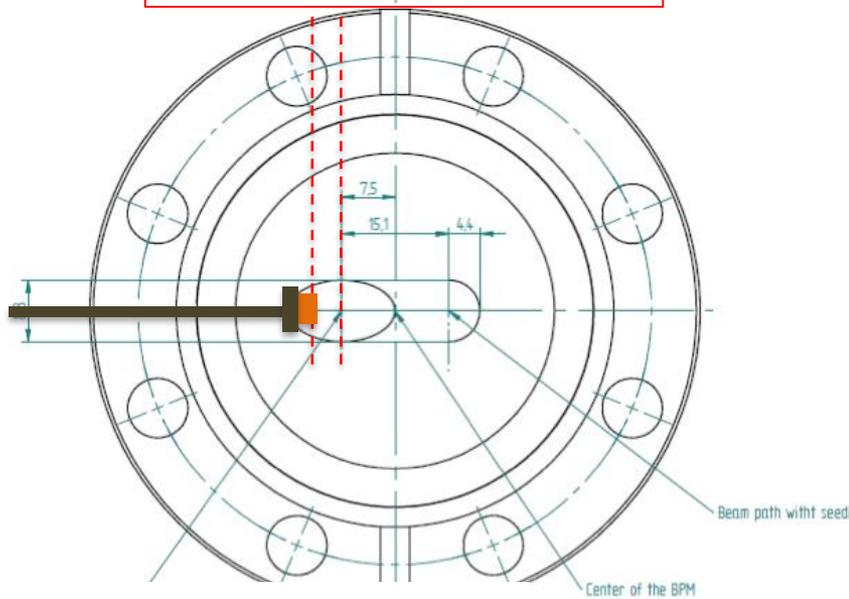


Chicane Design: minimum delay



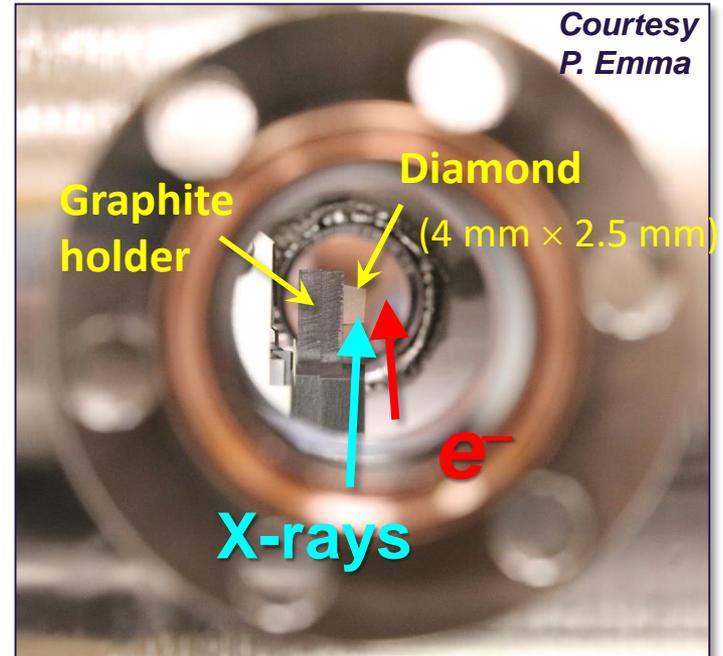
European XFEL HXRSS with **450 times** higher duty cycle than LCLS

minimum offset: ? mm



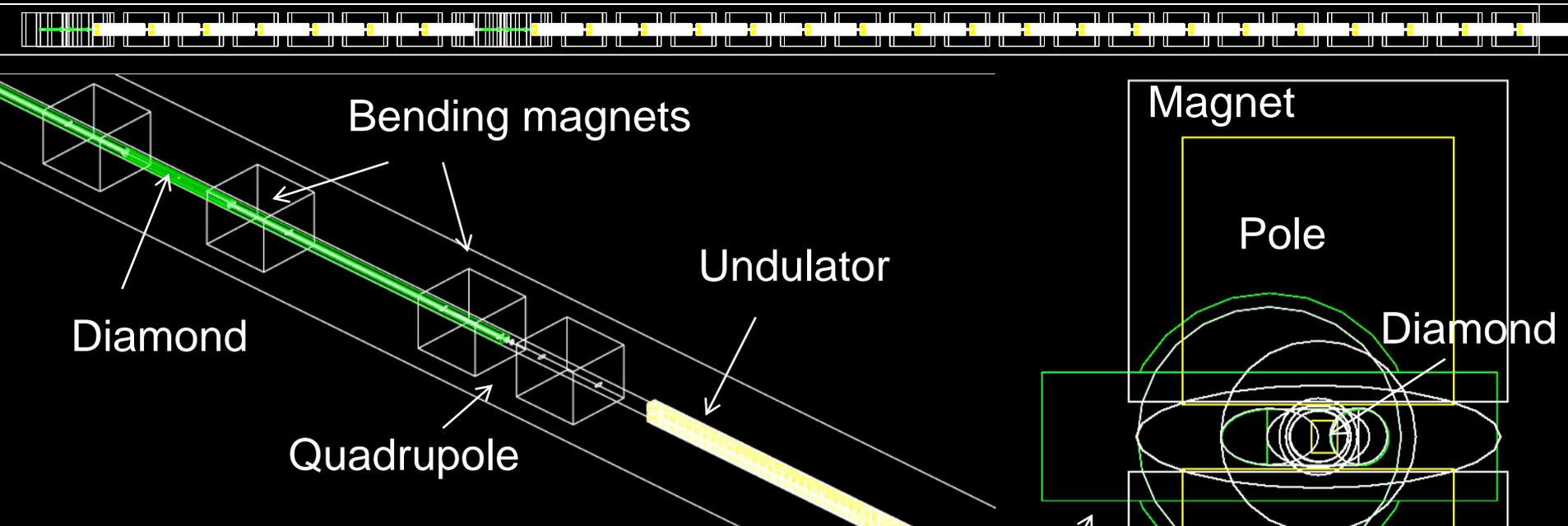
Center of undulator system (beam w/o seeding)

LCLS HXRSS: diamond is **2.5 mm** from beam center



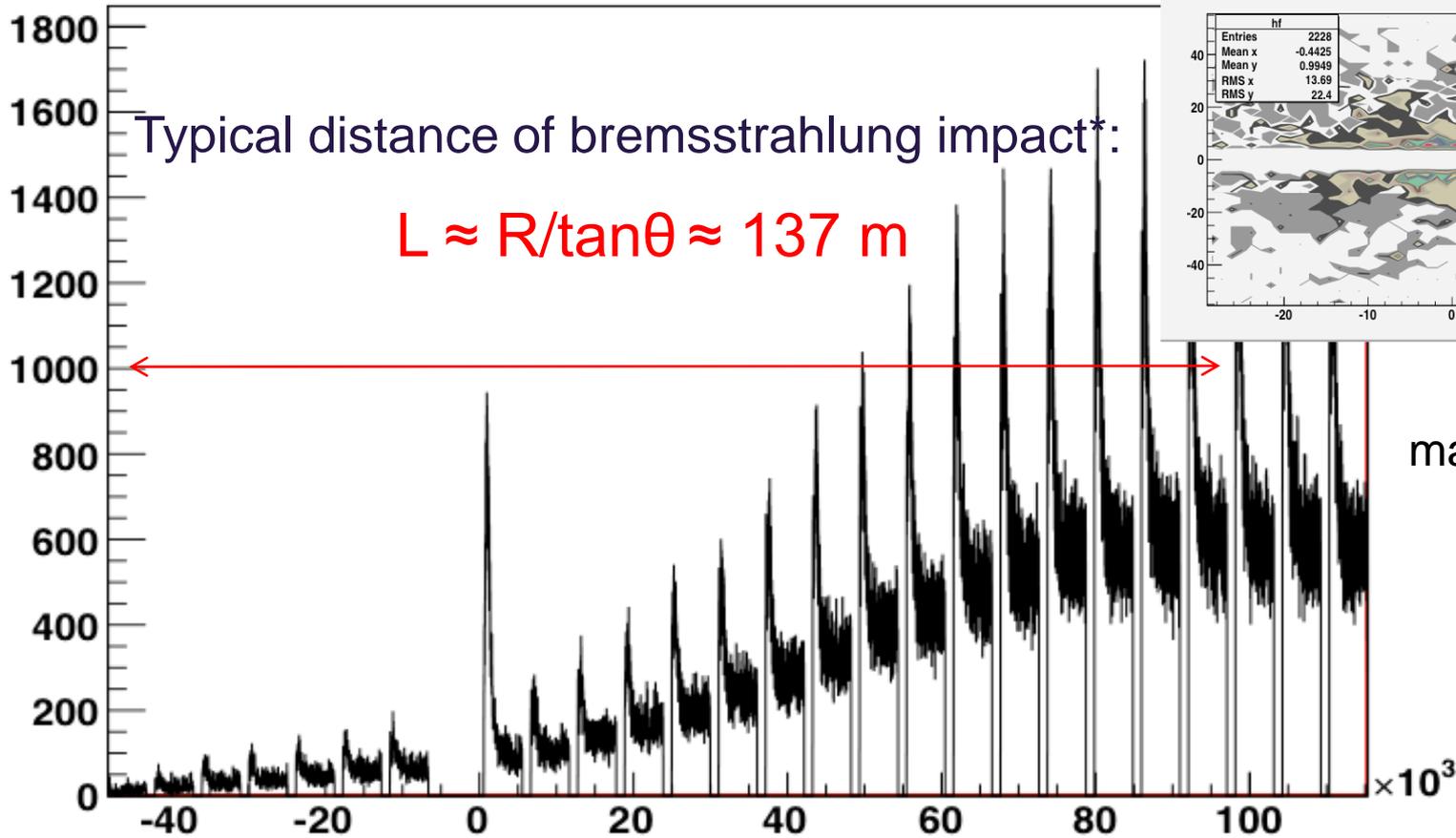
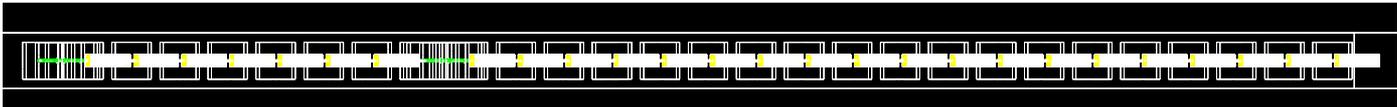
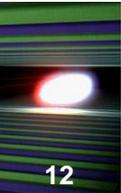
Undulator damage: GEANT4 simulations

- HXRSS+7 undulators+HXRSS+18 undulators



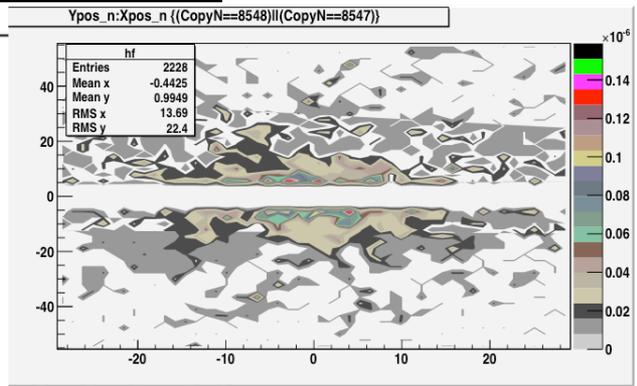
- Dipoles and quadrupoles are simplified as boxes with elliptical apertures
- Input beam: 10^6 e^- with 17.5 GeV energy hitting the crystal with 90° angle

Neutron fluence along undulator line



Typical distance of bremsstrahlung impact*:

$$L \approx R/\tan\theta \approx 137 \text{ m}$$



Normalized maximum neutron fluence :
 $\sim 1 \times 10^{-7} \text{ n/cm}^2/\text{e}^-$

* A. Fasso, Dose Absorbed in LCLS Undulator Magnets, I. Effect of a 100 μm Diamond Profile Monitor, RP-05-05, May 2005.

- Maximum neutron flux allowed for 0.01 % demagnetization of Nd-Fe-B magnets: $1 \times 10^{11} \text{ n/cm}^2$ *.
- Assuming 0.01% demagnetization in 20 years with 10 shifts (8 hours each) /month for HXRSS operation, the **maximum allowed number of e⁻/bunch** (with 27000 bunches/s):

$$N_{\max} = 1 \times 10^{11} / 20 \times 12 \times 10 \times 8 \times 60 \times 60 \times 27000 \times 10^{-7} \approx 5 \times 10^5 \text{ e}^-/\text{bunch}$$

- Maximum charge per bunch: $1 \text{ nC} \approx 6.25 \times 10^9 \text{ e}^-/\text{bunch}$
- Maximum horizontal beam size in undulator section: $\sigma_x \approx 50 \text{ } \mu\text{m}$

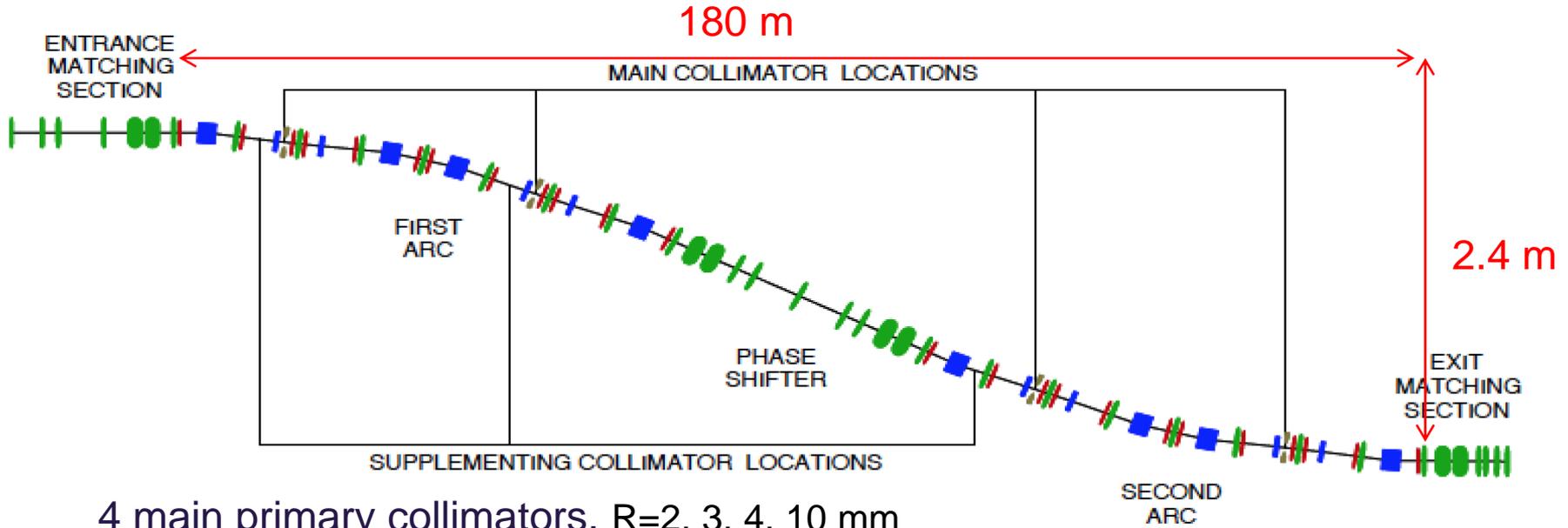
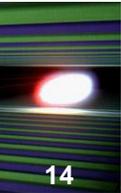
If there is **no halo**, the crystal can enter **as close as $\sim 4\sigma_x \approx 200 \text{ } \mu\text{m}$** to a gaussian beam core with above mentioned conditions!

What if there is $100 \sigma_x$ of halo?

*M. Santana Leitner, et al, Radiation Protection Studies for LCLS Tune Up Dump, SLAC-PUB-14020, 2010.

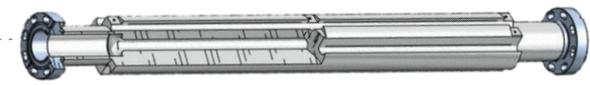
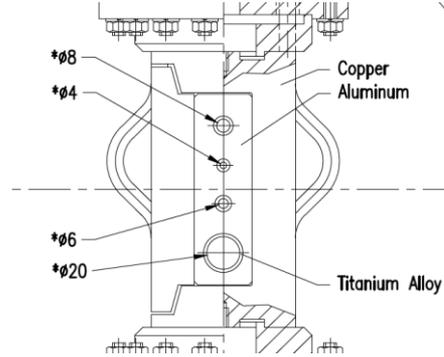
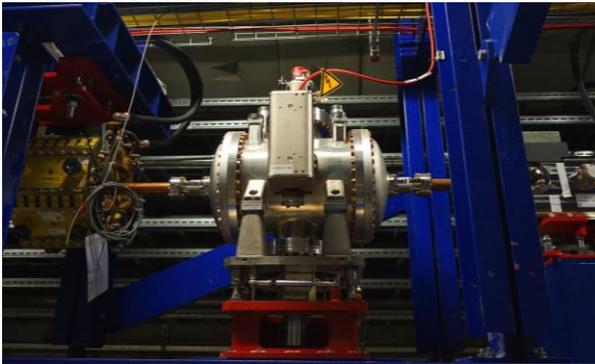
*J. Alderman, P.K. Job, R.C. Martin, C.M. Simmons, G.D. Owen, and O J. Puhl, Radiation-Induced Demagnetization of Nd-Fe-B Permanent Magnets, Advanced Photon Source Report LS-290 (2000).

Collimation System Design

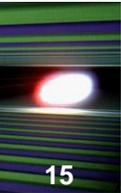


4 main primary collimators, R=2, 3, 4, 10 mm
(Titanium alloy+Aluminum+Copper)

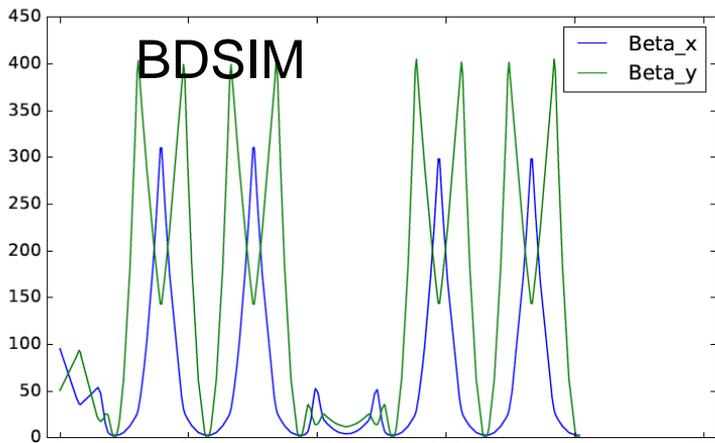
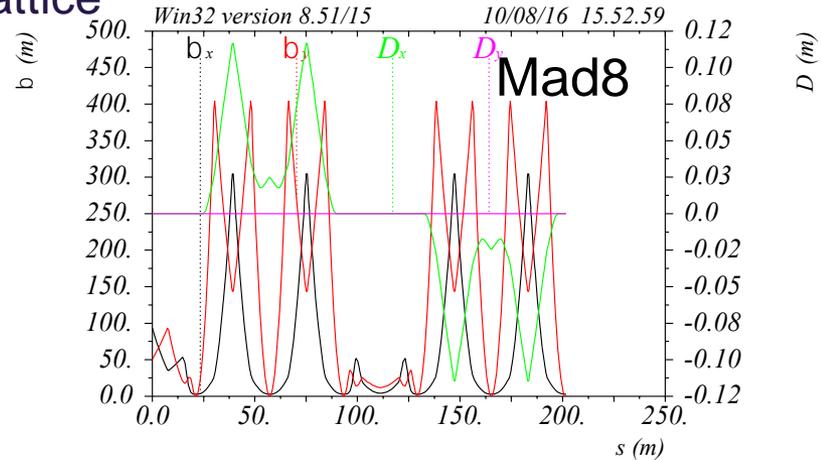
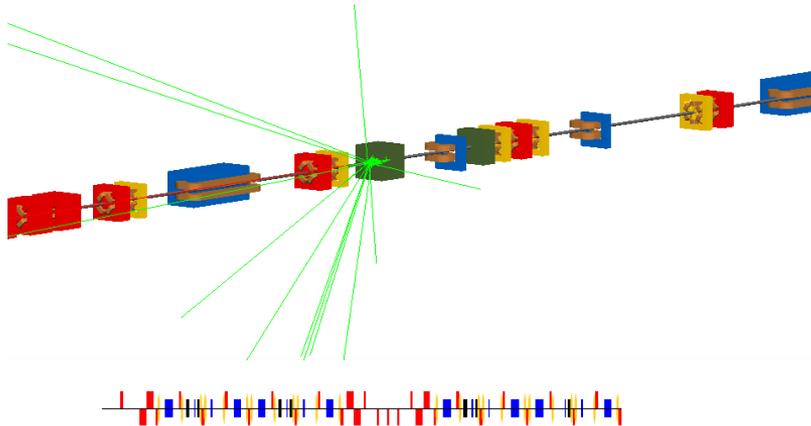
Supplementary collimator aperture: R=10 mm (Aluminum)



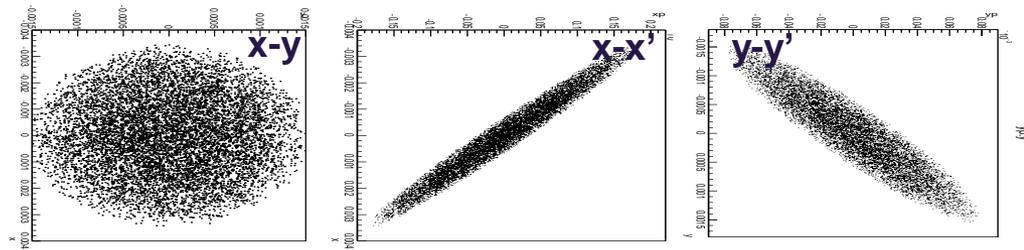
* V.Balandin, R.Brinkmann, W.Decking, N.Golubeva, TESLA-FEL 2007-05



BDSIM*: GEANT4 simulation with accelerator lattice

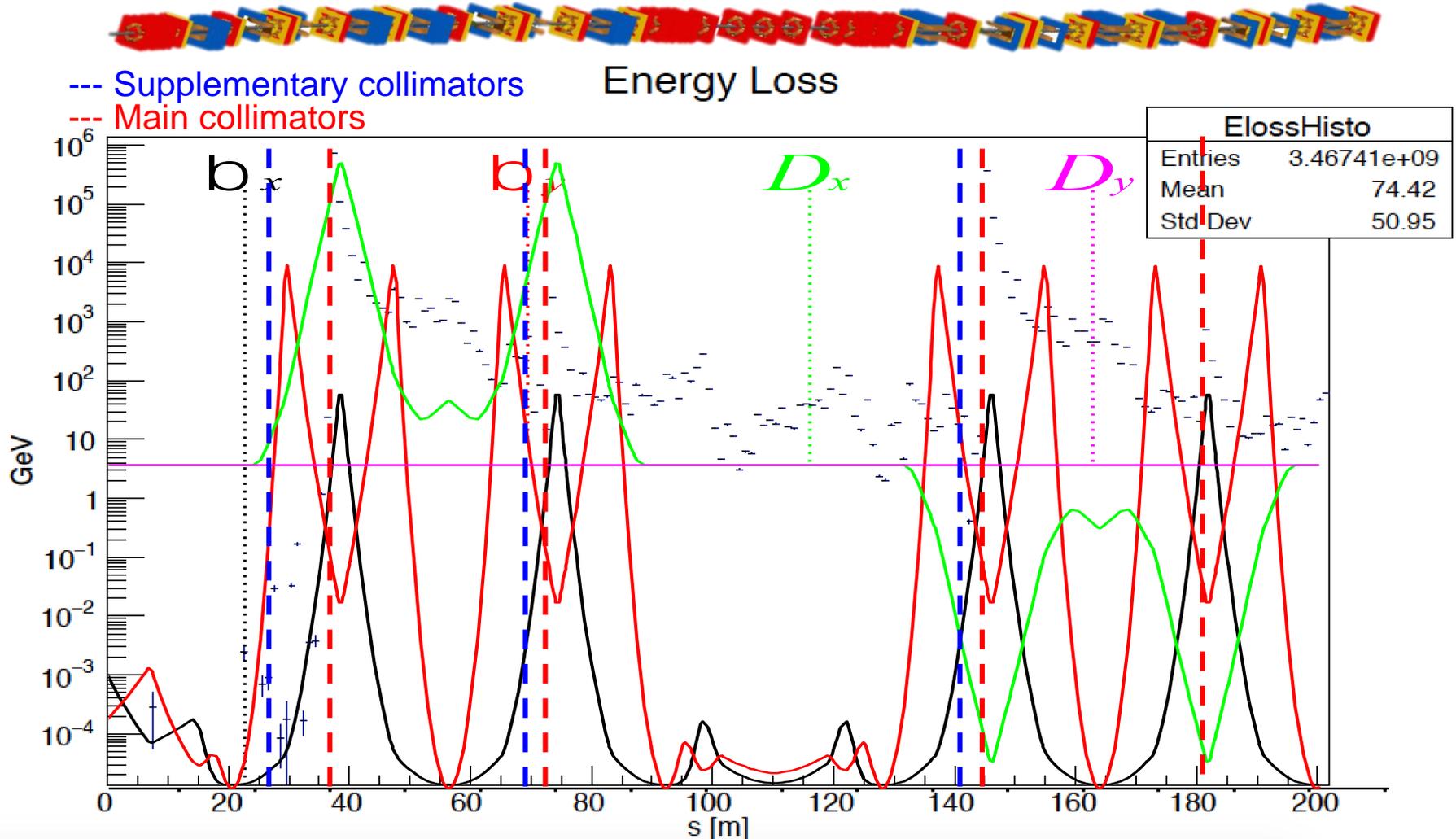


First simulations are performed with 10^5 e- with $\pm 50\sigma$ uniform distribution in (x, x', y, y') phase space and with $R=2$ mm collimator aperture

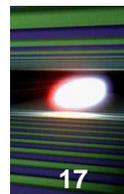


- Effects from sextuples are not considered for the first simulations.
- X and Y coordinate are exchanged in the simulations. *I. Agapov, et al, NIMA:606.3 (2009): 708-712.

Energy loss along the collimation line

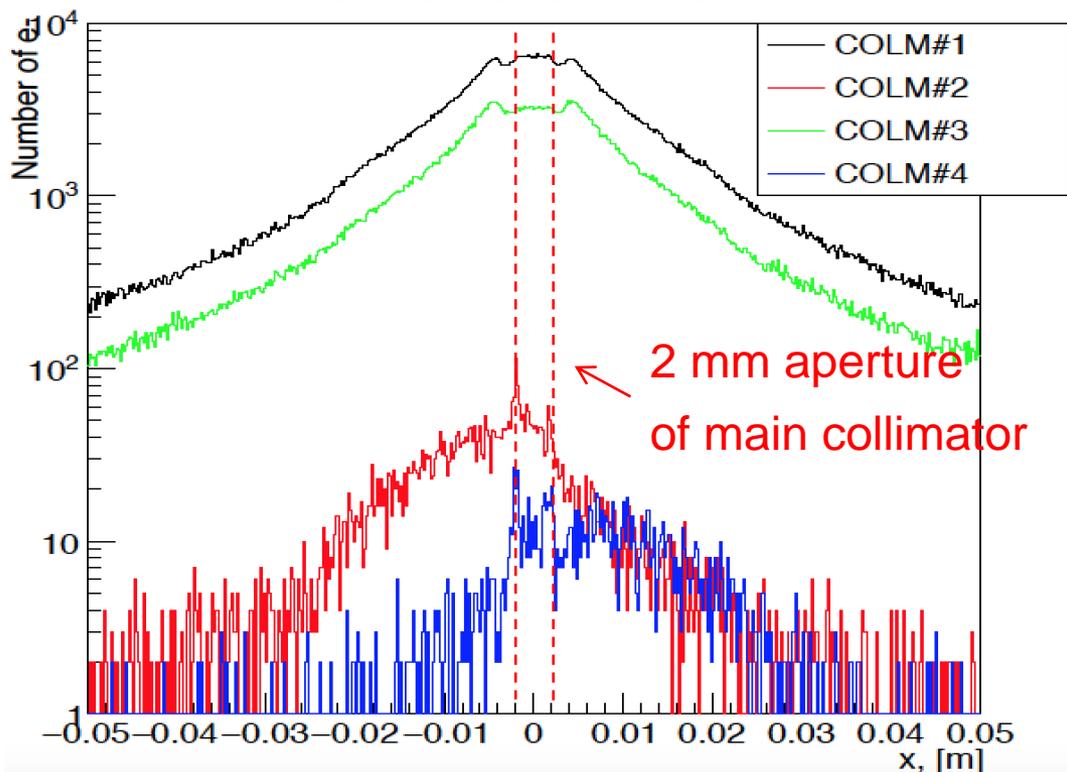


-> can be compared with the BLM measured results in the upcoming commissioning.

Secondary e⁻ @ main collimators

17

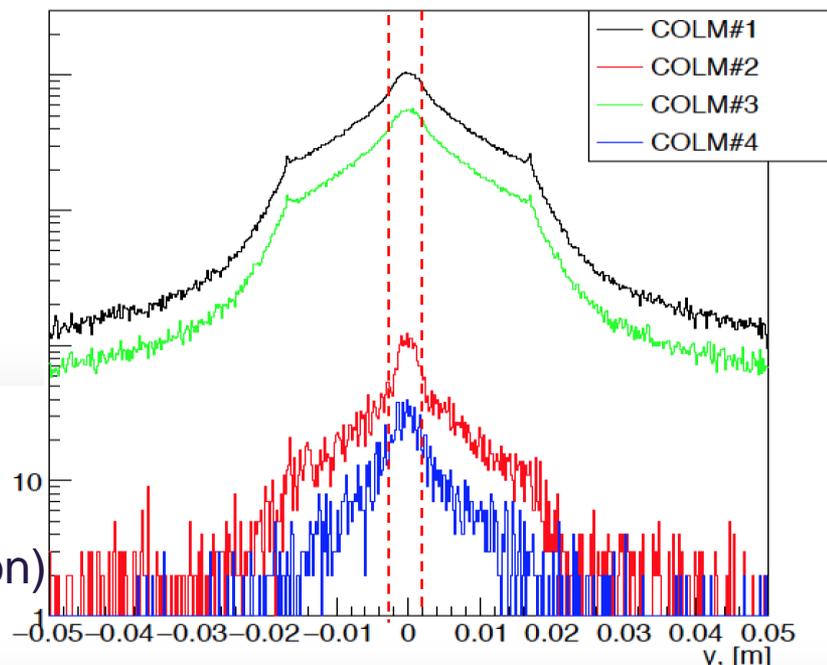
Horizontal distribution



- Secondary particles and residual primary particles beyond 2 mm are collimated by COLM#2 and COLM#4 (the asymmetric distributions in x indicate the bending direction)

- Primary particles beyond 2 mm aperture are collimated mainly by the first main collimator in the two arcs: COLM#1 (black) and COLM#3 (green)

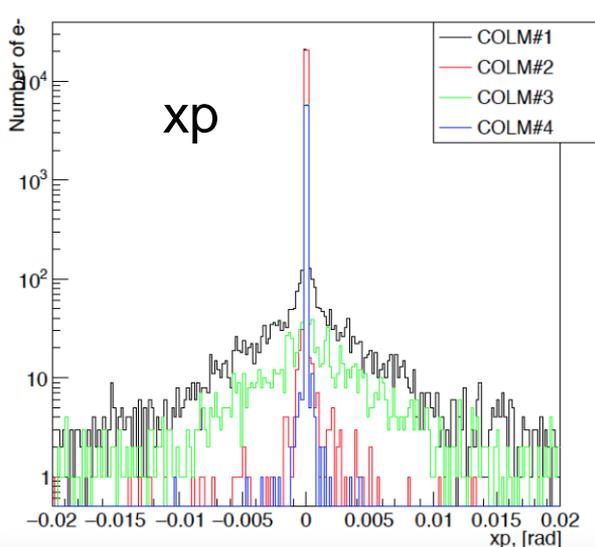
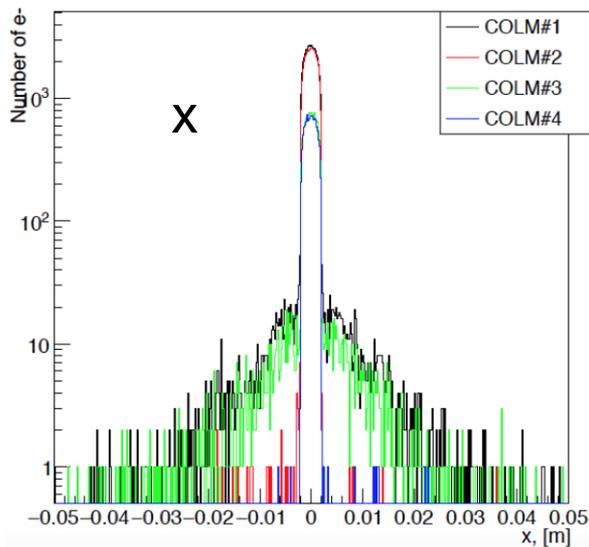
Vertical distribution



Primary e- distributions @ main collimators

18

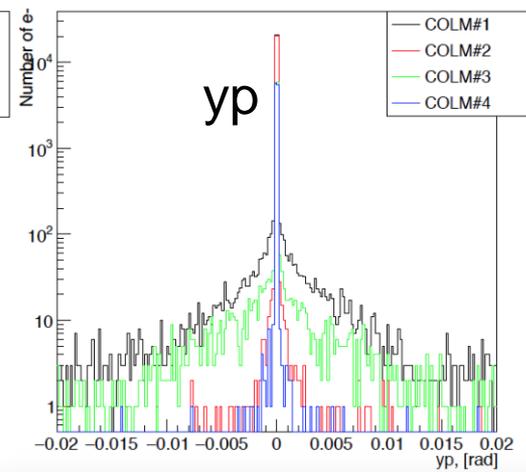
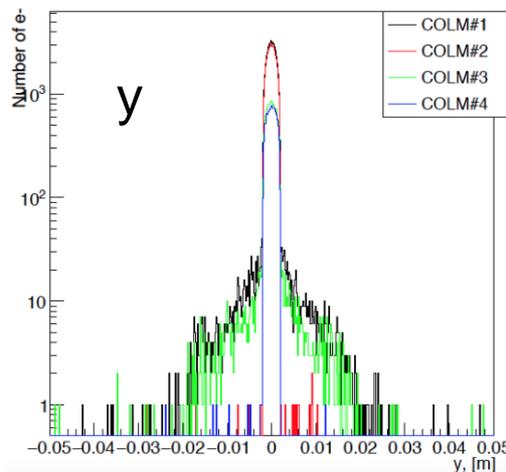
Horizontal distribution



- Primary particles within $\sim \pm 20\sigma_{x,xp}$ & $\sim \pm 35\sigma_{y,yp}$ can pass the collimators with 2 mm apertures freely;

Vertical distribution

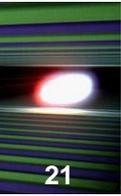
- Almost no primary halo particles left beyond the 2 mm aperture after the 4th collimator -> in this case, **the crystal** can be inserted up to a distance of **~ 2 mm** to the beam core (**~ 13 fs of minimum delay**) !



- **HXRSS + tapering enable specific high-brightness applications** -> simulations are on-going to finalize the location of the 2 HXRSS stages;
- Energy shift due to heat load on diamond -> can be compensated by pitch oscillator
- **Issue of beam halo and radiation damage**-> can be studied by simulations with different beam halo distributions and compared with beam loss/halo monitors measured results in the future -> **important for HXRSS, corrugated structure** with very small gaps (~1.4 mm planed at European XFEL) similar for LCLS-II.
- HXRSS project Schedule:
 - Present status: chicane (magnets, vacuum chamber, girder) and monochromator chamber design completed, components ordered
 - End 2016: final design of monochromator
 - End 2017: Two HXRSS mono ready for installation
 - 2018- Installation

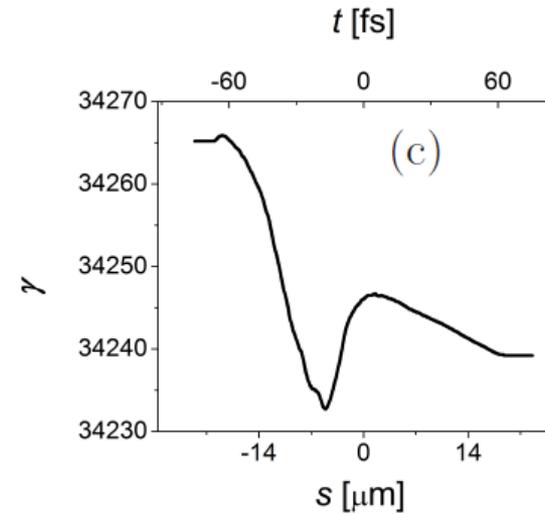
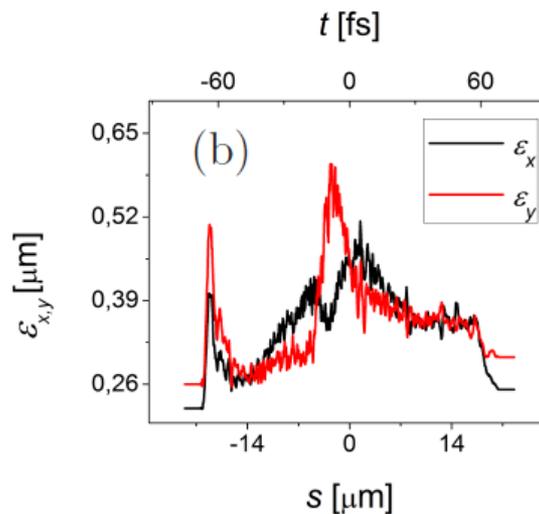
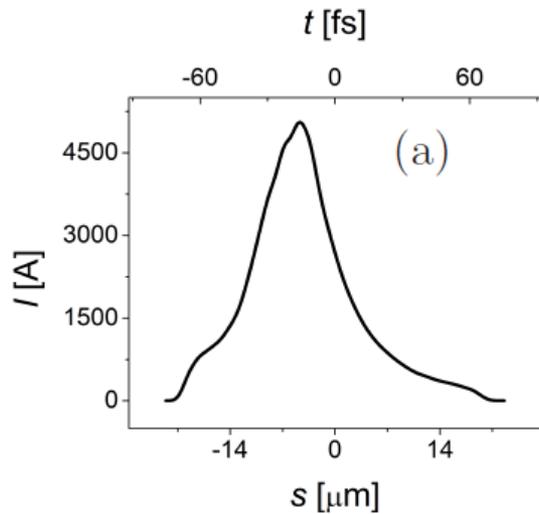
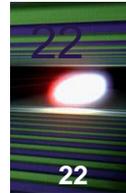
Thank You!

- Thanks to I. Agapov, V. Balandin, W. Decking, G. Feng, L. Froehlich, G. Geloni, N. Golubeva et al. for helpful information and discussions!
- Thanks to colleagues from LCLS for sharing the HXRSS design and operation experiences!
- Thanks to S. Boogert and L. Nevay (RHUL) for the support on BDSIM!



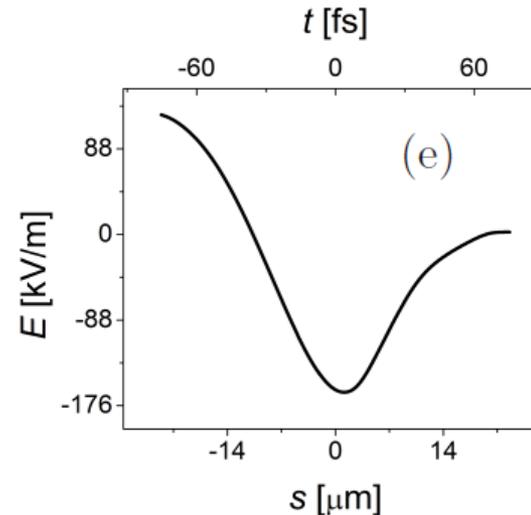
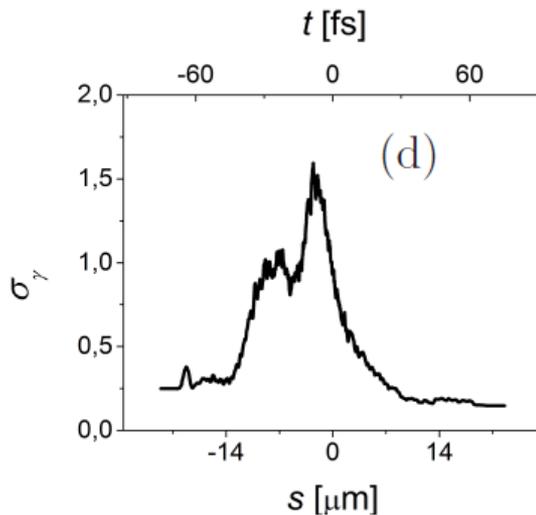
Back Up...

Ultra-High Resolution Inelastic X-ray Scattering UHRIX at 9 keV

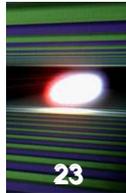


Start-to-end
simulations by
the DESY/MPY
Group

(Standard mode
of operation of
European XFEL
at 250pC)

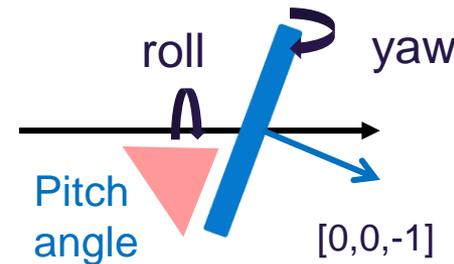
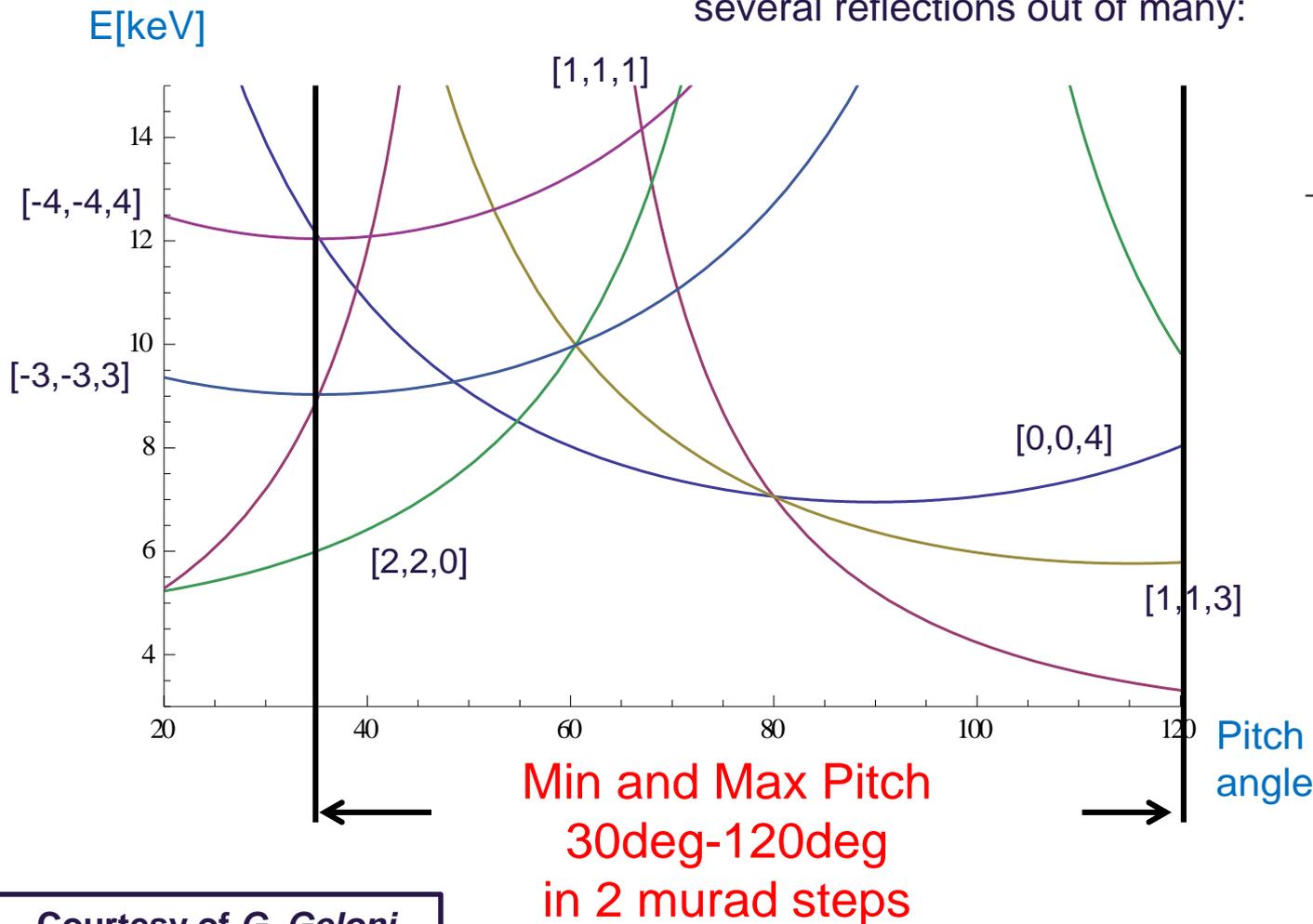


Energy Range and Reflections



We can cover the range from 3keV up to >12keV using only a several reflections out of many:

hkl	Min. Energy (keV)
111	3.01034
311	5.76401
400	6.95161
333	9.03035
444	12.0404



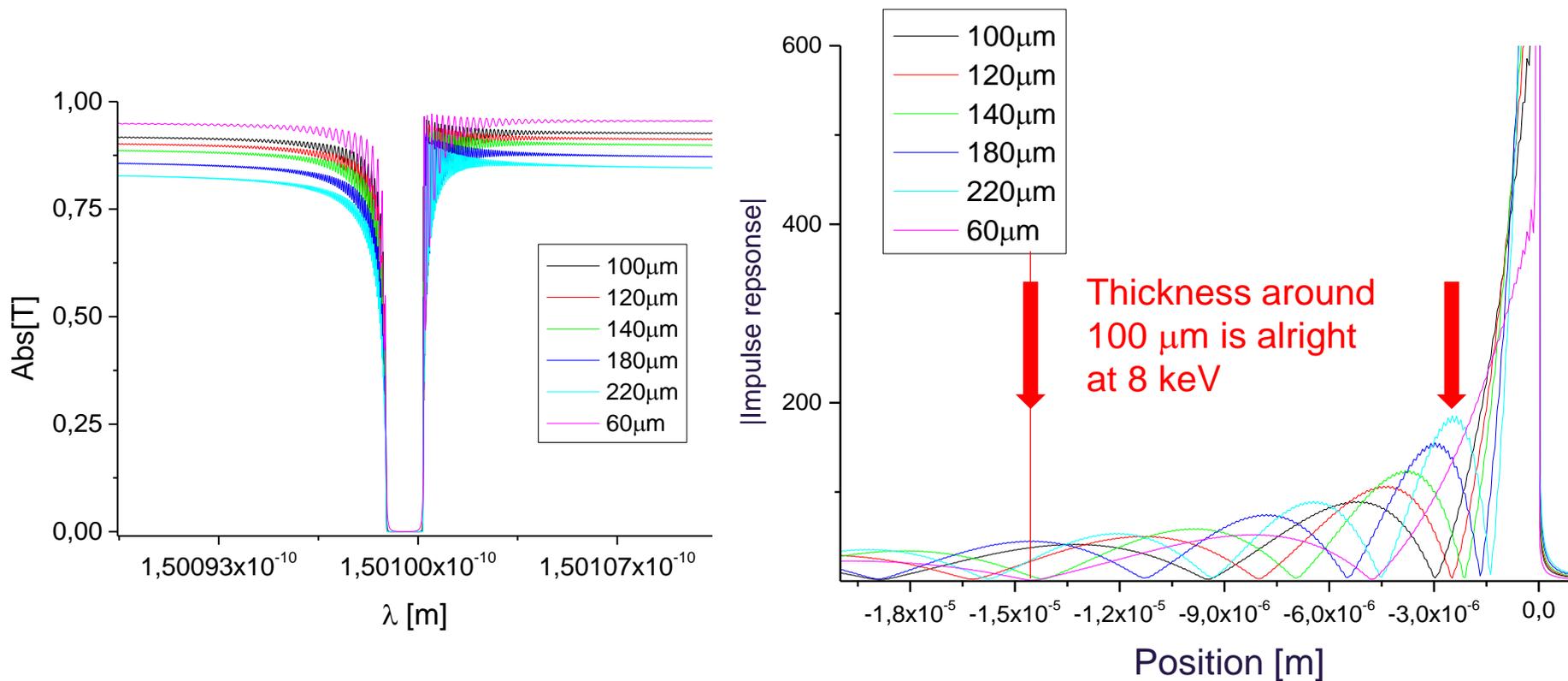
We should cover the FEL BW with +/- 3 degrees in yaw (and roll) with angular steps of 20murad

Courtesy of G. Geloni

Courtesy of G. Geloni

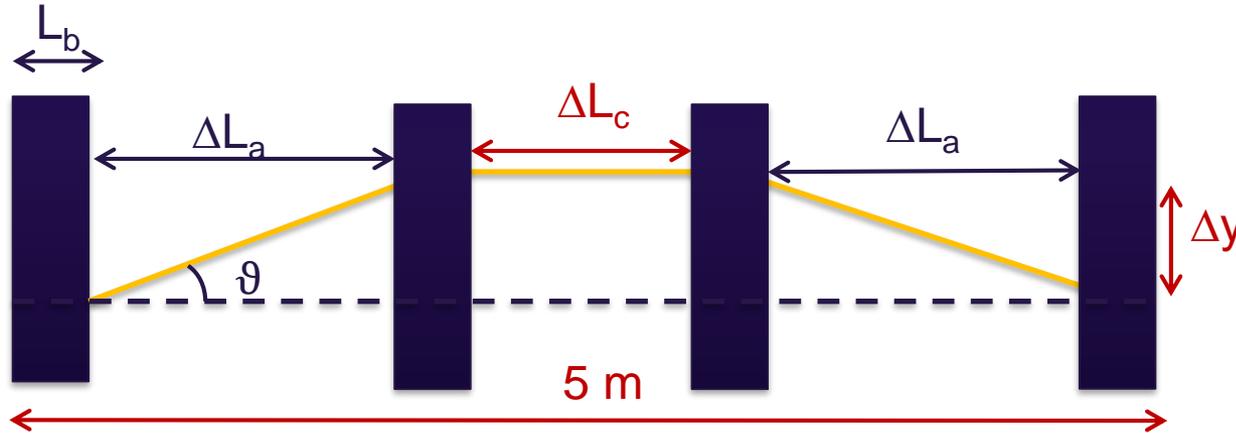
Thickness optimization relates to

- 1) Crystal impulse response
- 2) Minimal delay achievable
- 3) Heat loading issues (due to absorption)



C400 at 8.26keV

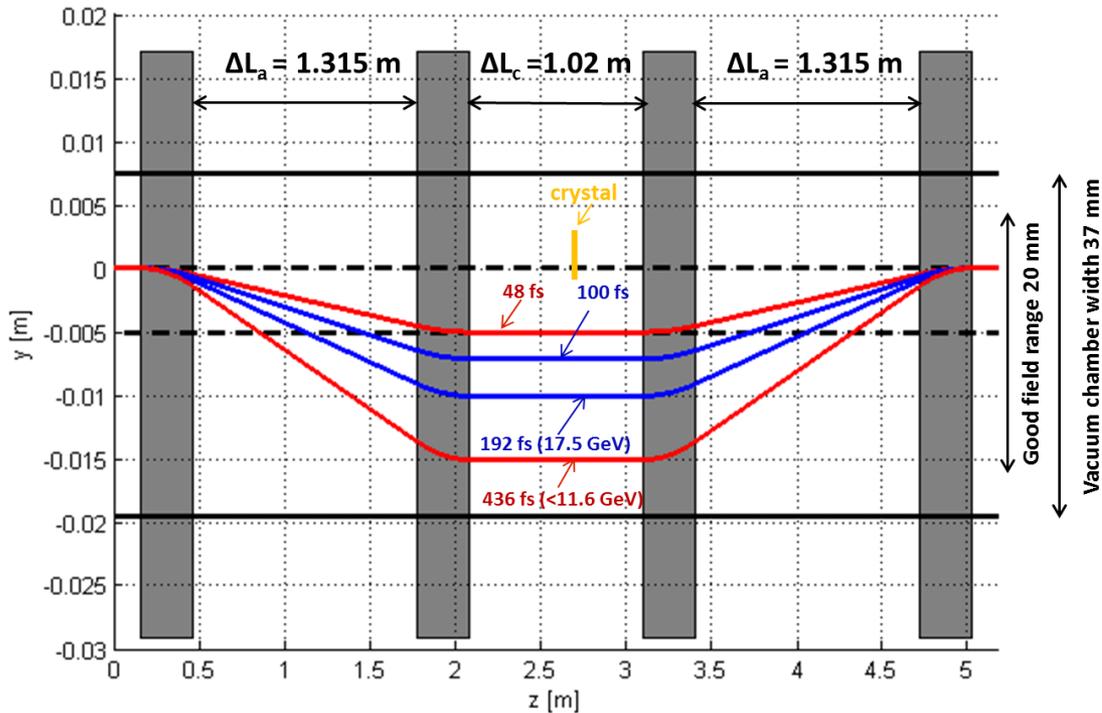
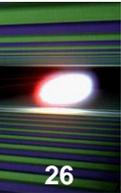
Chicane Design: Functions



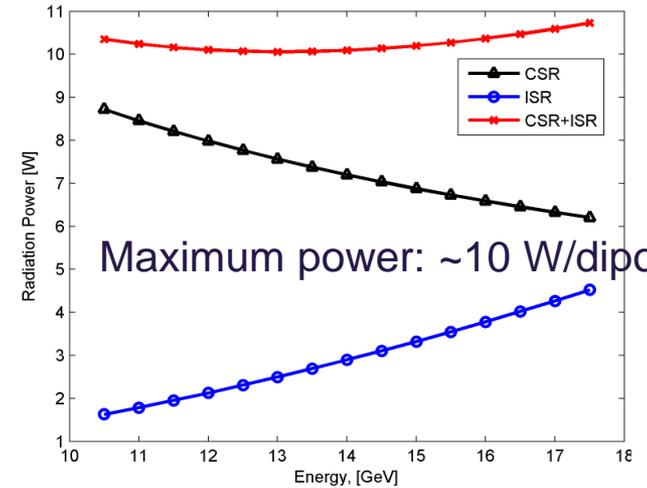
Functions of chicane:

- Create an offset between the electron bunch and the monochromator
- Wash out the SASE micro-bunching produced by the upstream undulators
- Produce an adjustable delay for the electron bunch to match the seed x-ray
- Correcting the phase error introduced by removing one undulator segment
- Scan the FEL photon pulse to obtain the length of the pulse
- Produce two color SASE pulses

Chicane Layout and Radiation Issues

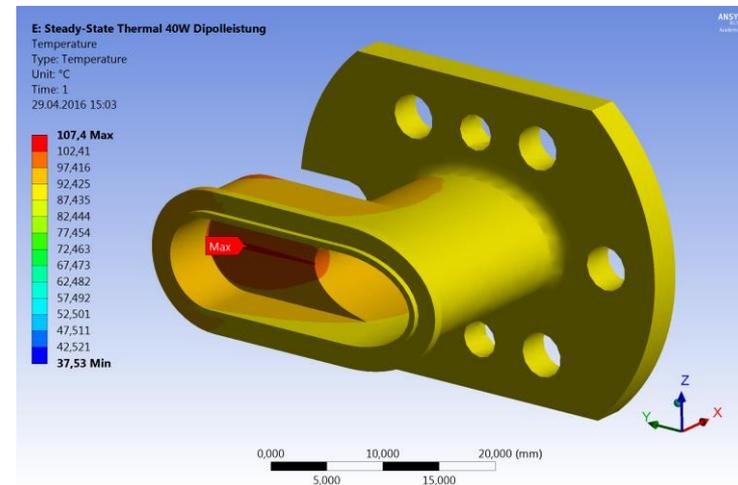


Radiation Power

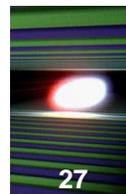


Max. Temperature: 107 degree

- Worst case scenario with 27000 bunches with 1 nC at 17.5 GeV
- Planned operation far below the dump power threshold (300 kW) : max. 13500 bunches with 1 nC at 17.5 GeV.
- Water cooling may be added



Chicane Main Parameters

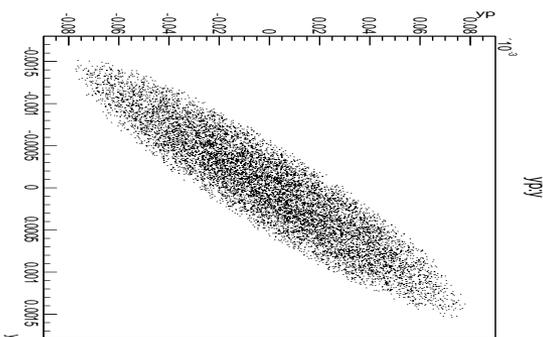
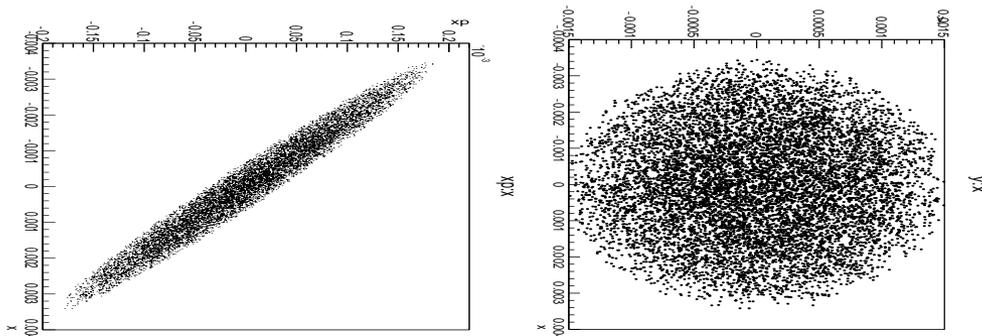


Parameters*	Symbol	Normal	Minimum	Maximum	Unit
Electron beam delay	Δt	48.4 – 100	0	192.2	fs
Electron beam offset	Δy	5 – 7.2	0	10	mm
Width of good field range	w_g	20	10	-	mm
Deflection angle	ϑ	3.1 – 4.4	0	6.2	mrad
R56	R_{56}	29 – 60	0	115.3	μm
Dipole field	B_0	0.60 – 0.87	0	1.2	T
Integrated field strength	$\int B dl$	0.18 – 0.26	0	0.36	T·m
Magnet yoke length	L_b	0.3	-	-	m
Magnet vertical gap	g	16	-	-	mm
Vacuum chamber width	w	27	-	-	mm
Distance between outer and inner magnets	ΔL_a	1.315	-	-	m
Distance between inner magnets	ΔL_c	1.02	-	-	m
Nominal current	I	175	-	-	A

*Parameters for 17.5 GeV beam

Input electron beam distribution

uniform distributions



betx=93.554;
bety=50.482;
alfx=4.857;
alfy=-2.389;
ex=2*Nsig*5.31e-11;
ey=2*Nsig*1.89e-11;
deltaE/E=1.5%.

- The Courant-Snyder invariant can be expressed as:

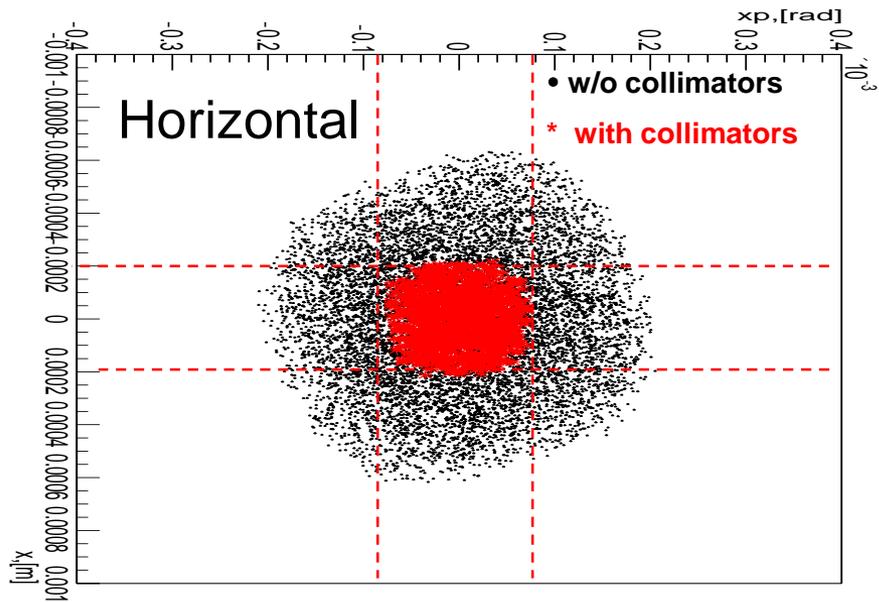
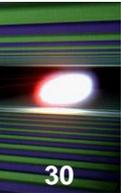
$$\tilde{x}^2 + \tilde{x}'^2 = \gamma x^2 + 2\alpha x x' + \beta x'^2 = \frac{\epsilon}{\pi}$$

with the normalized variables:

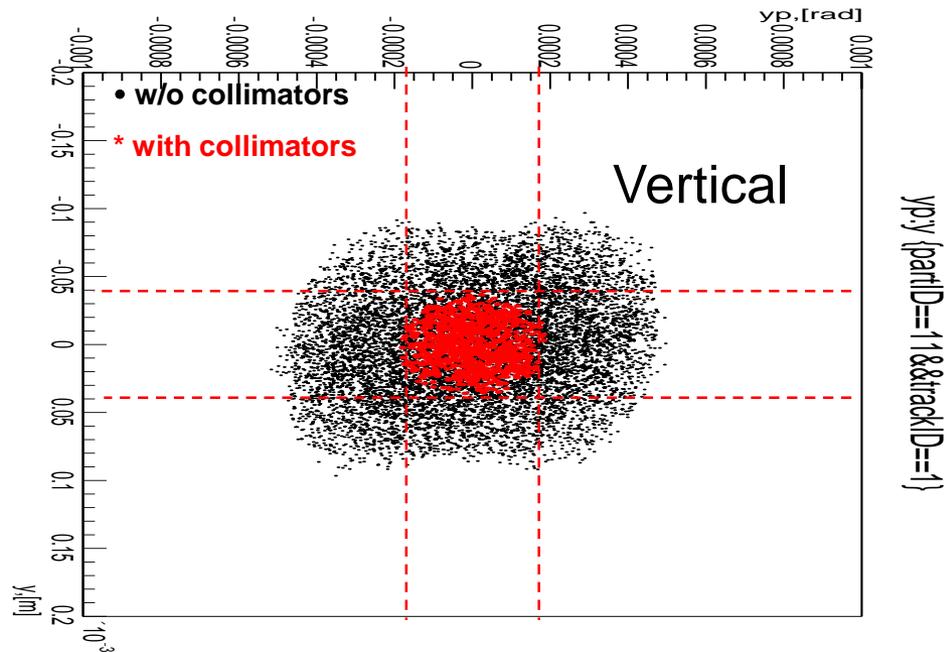
$$\tilde{x} = \frac{x}{\sqrt{\beta}} \quad \text{and} \quad \tilde{x}' = \frac{d\tilde{x}}{d\phi} = \sqrt{\beta} x' + x \frac{\alpha}{\sqrt{\beta}},$$

(\tilde{X}, \tilde{X}') and (\tilde{Y}, \tilde{Y}') are enlarged by N times with uniform distributions.

Phase space @ end of collimation section



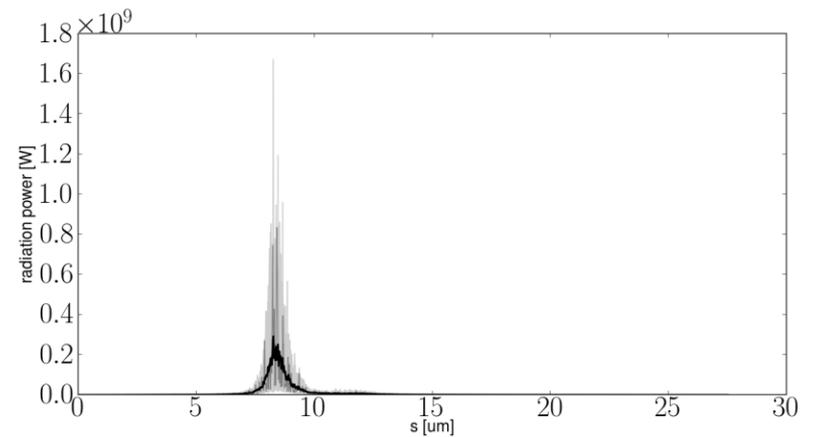
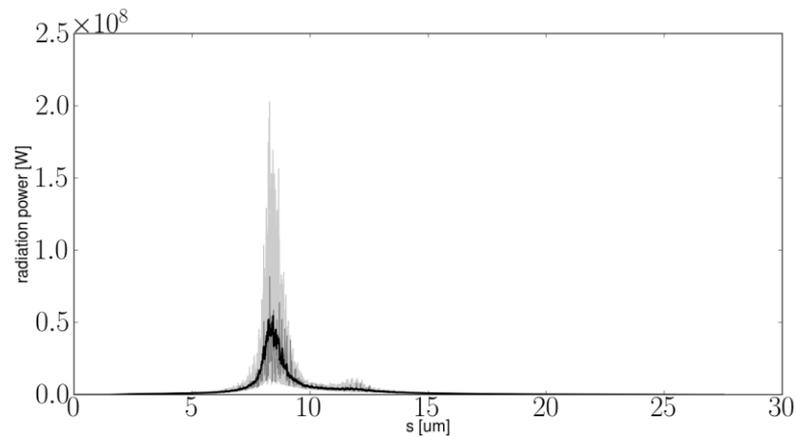
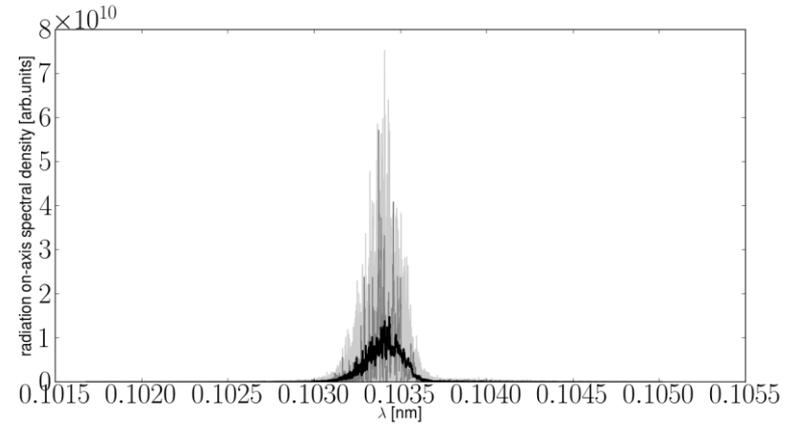
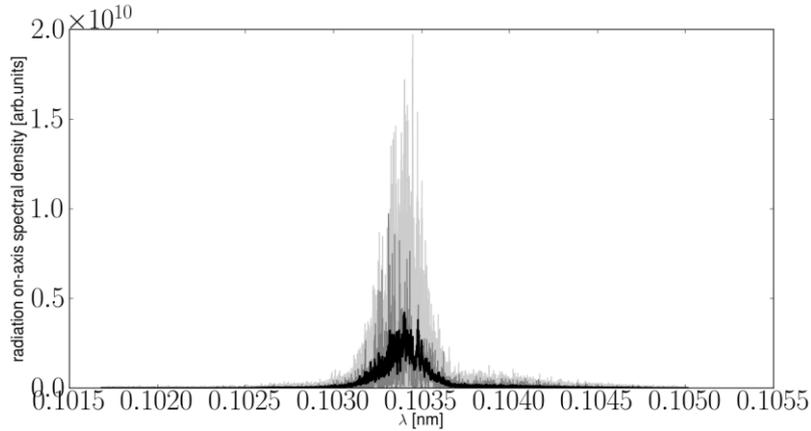
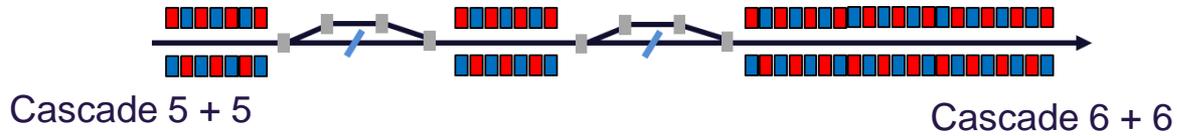
$xp_x \{partID==1 \&\&trackID==1\}$



$yp_y \{partID==1 \&\&trackID==1\}$

* Keep in mind that horizontal and vertical planes are exchanged in the simulation

100pC – 17.5 GeV – C004 – 12keV – 1st stage



100pC – 17.5 GeV – C004 – 12keV – 3rd stage

