

EFT in top physics

Facts, Remarks and Proposals

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Introductory remark

The facts, remarks and proposals discussed here are more general than the particular example discussed.

I will refer to *Wtb* (most studied top interaction) for definiteness. But you can imagine that the lessons to be learnt can be applied to other interactions as well.

Fact #1

Only a few among top physics measurements might qualify as *precision measurements*

$$F_0 = 0.681 \pm 0.012 \text{ (stat)} \pm 0.023 \text{ (syst)}$$

$$F_- = 0.323 \pm 0.008 \text{ (stat)} \pm 0.014 \text{ (syst)}$$

$$F_+ = -0.004 \pm 0.005 \text{ (stat)} \pm 0.014 \text{ (syst)}$$

CMS

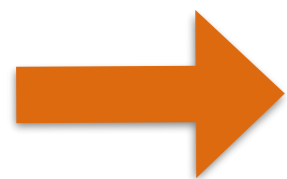
systematics dominated

while many others do not

$$\sigma_{t-\text{ch}} = 83.6 \pm 2.3 \text{ (stat)} \pm 7.4 \text{ (syst) pb}$$

$$\sigma_t / \sigma_{\bar{t}} = 1.95 \pm 0.10 \text{ (stat)} \pm 0.19 \text{ (syst)}$$

CMS



we are not dealing with precision physics when extracting dim6 top operator coefficients from data

Fact #2

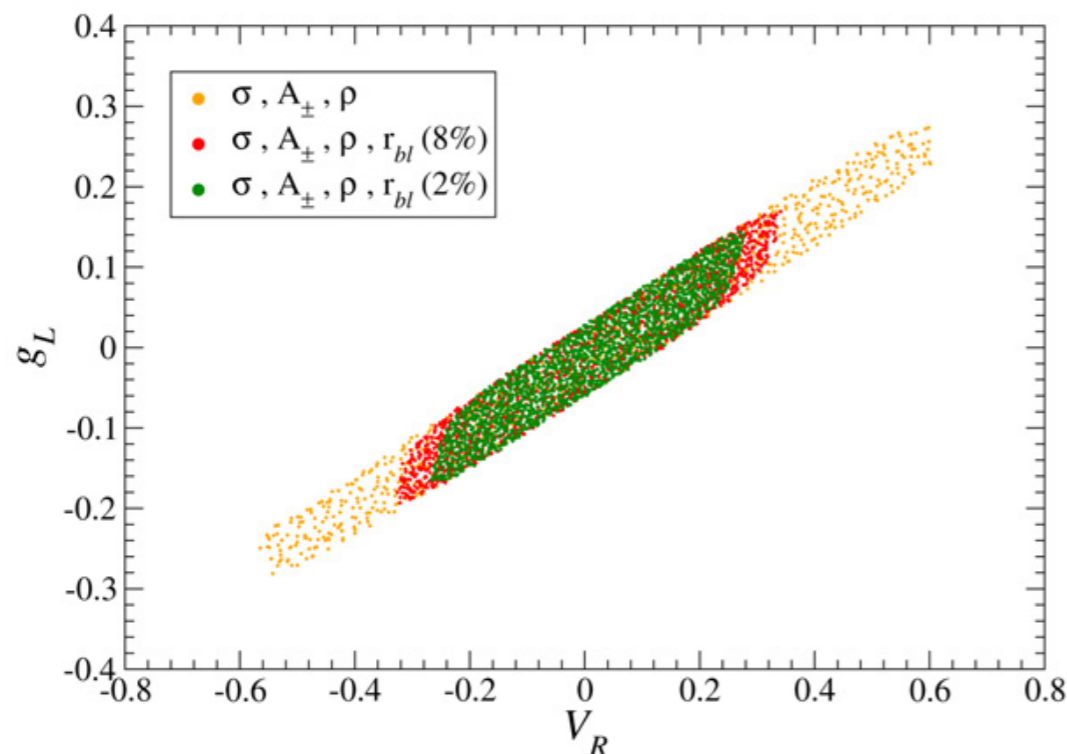
In observables there are often cancellations among anomalous contributions that further degrade the sensitivity

$$\bar{b}(p_2) [i\sigma^{\mu\nu} (p_1 - p_2)_\nu P_L + m_t \gamma^\mu P_R] t(p_1) = \bar{b}(p_2) (p_1 + p_2)^\mu P_L t(p_1)$$

for t, b on shell and $m_b=0$

→
$$\left[V_R = \frac{m_t}{M_W} g_L \right] \approx \text{SM}$$

no contribution to Γ_+ and Γ_- .

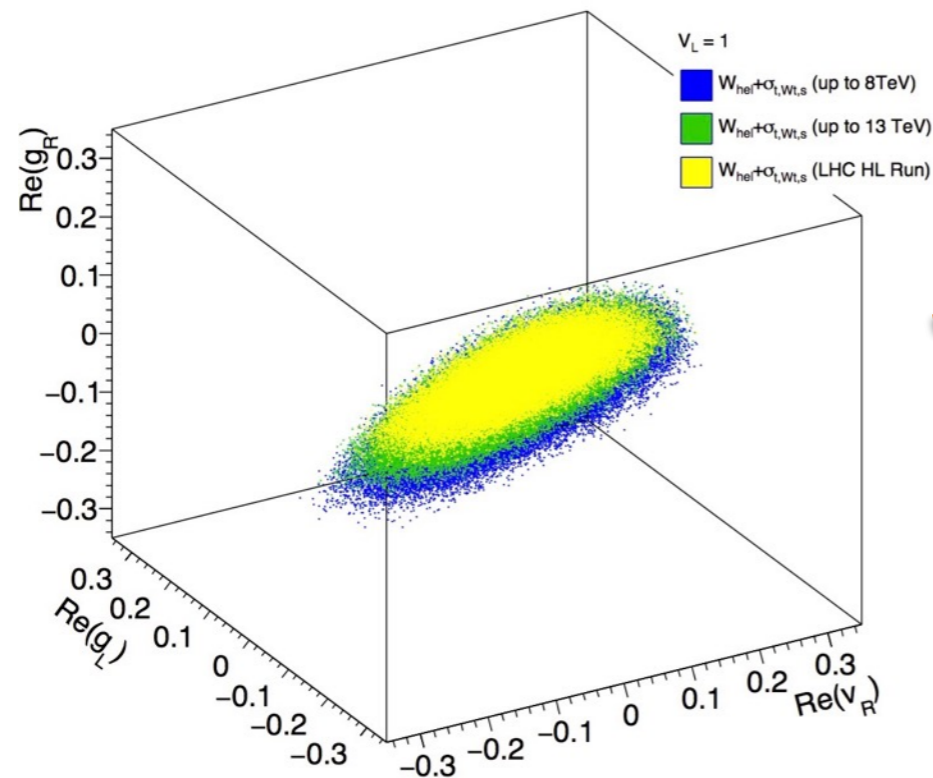


JAAS 0803.3810

$$\begin{aligned} \mathcal{L}_{Wtb} = & -\frac{g}{\sqrt{2}} \bar{b} \gamma^\mu (V_L P_L + V_R P_R) t W_\mu^- \\ & -\frac{g}{\sqrt{2}} \bar{b} \frac{i\sigma^{\mu\nu} q_\nu}{M_W} (g_L P_L + g_R P_R) t W_\mu^- + \text{h.c.} \end{aligned}$$

...and when all couplings are left arbitrary, limits are very loose

Birman et al. 1605.02679



$$\Lambda > 0.3 \text{ TeV} \quad (C = 1)$$



$$V_L = V_{tb} + C_{\phi q}^{(3,33)} \frac{v^2}{\Lambda^2}$$

$$g_L = \sqrt{2} C_{dW}^{33*} \frac{v^2}{\Lambda^2}$$

$$V_R = \frac{1}{2} C_{\phi\phi}^{33*} \frac{v^2}{\Lambda^2}$$

$$g_R = \sqrt{2} C_{uW}^{33} \frac{v^2}{\Lambda^2}$$

strong temptation:

ignore operators loosely constrained
as well as quadratic terms

Remark #1

Fifth Commandment of EFT: “*Thou shalt not kill quadratic terms*”

JAAS 1008.3225
JAAS et al. RMP

- often they are important, because of Facts #1 and #2
- in case they are not, they don't matter anyway

And in many cases it is not inconsistent to keep them while dropping dim-8

series converges

$$C_i^{(4)} \frac{E^2}{\Lambda^4} \ll C_i^{(2)} \frac{1}{\Lambda^2} \quad \longrightarrow \quad C_i^{(4)} \lesssim C_i^{(2)}$$

meaningful Λ
[Fact #2]

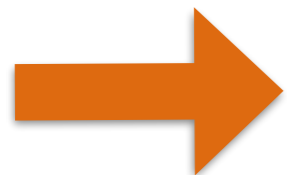
$$C_i^{(2)} \gg 1 \quad \longrightarrow \quad C_i^{(2)} \ll \left(C_i^{(2)}\right)^2$$

therefore

$$\frac{C_i^{(4)}}{\Lambda^4} \ll \left(\frac{C_i^{(2)}}{\Lambda^2}\right)^2$$

See [Contino et al. 1604.06444](#)
for an alternative argument

Quadratic terms are positive semidefinite and ensure that bounds on anomalous couplings exist even if measurements are insufficient or have little precision

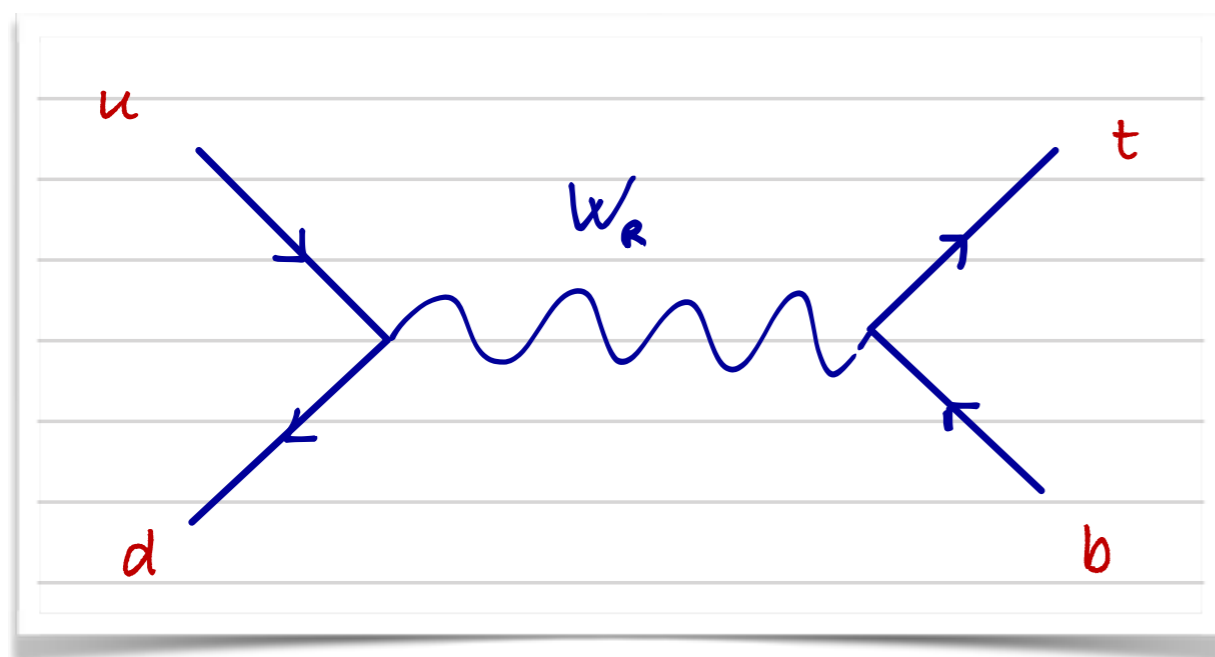


evaluating *theory uncertainty* by switching on/off quadratic terms leads to an absurd uncertainty when they dominate, which is often the case

Remark #2

Second Commandment of EFT: “*You shalt not drop operators in vain*”

JAAS 1008.3225

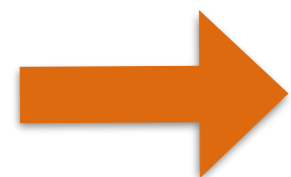


$$(\bar{d}_R \gamma^\mu u_R)(\bar{t}_R \gamma_\mu b_R)$$

New physics may only generate operators that do not interfere with SM

When possible, one should consider all contributing operators

It is not absolutely necessary, however: operators are gauge invariant



you lose generality but not consistency

Remark #3

EFT [in top physics] is a consistent framework to parameterise unknown heavy new physics

it is not an *extended SM*

Then, why performing global fits to C_i ?

○ get precise constraints on new physics?



○ identify *directions* where new physics contributions may cancel



○ identify new observables more sensitive to new physics!



more important than the C_i are the measurements

Fact #3

These *cancellations* are related to the fact that we have dropped operators from our list

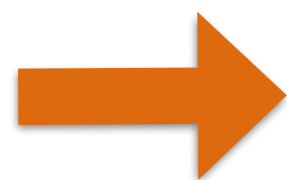
$$\left. \begin{aligned} O_{Dd}^{33} &\supset \frac{gv}{2} \bar{b} p_2^\mu P_L t W_\mu^- \\ O_{\bar{D}d}^{33} &\supset -\frac{gv}{2} \bar{b} p_1^\mu P_L t W_\mu^- \end{aligned} \right\} \longrightarrow O_{Dd}^{33} - O_{\bar{D}d}^{33} \supset \frac{gv}{2} \bar{b} (p_1 + p_2)^\mu P_L t W_\mu^-$$

does not contribute for helicity ± 1

$$O_{Dd}^{33} + O_{\bar{D}d}^{33} \supset \frac{gv}{2} \bar{b} (p_1 - p_2)^\mu P_L t W_\mu^-$$

does not contribute for any helicity

Fact #2 \longleftrightarrow we are almost insensitive to some operators



“barring cancellations” must be barred.

fine tuning is this



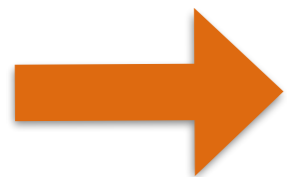
Interlude: genesis of dim-6 operator list

Original list	80 operators	Buchmuller & Wyler NPB 268 621, 1986
Added missing four-fermion operator	81 operators	Artz, Einhorn, Wudka hep-ph/9405214
Removed 7: O_{qW} , O_{qB} , O_{uB} , O_{dB} , O_{IW} , O_{IB} , O_{eB}	74 operators	Grzadkowski, Hioki, Ohkuma, Wudka hep-ph/0310159
Removed 9: O_{qG} , O_{uG} , O_{dG} , O_{Du} , $O_{\bar{D}u}$, O_{Dd} , $O_{\bar{D}d}$, O_{De} , $O_{\bar{D}e}$	65 operators	JAAS 0811.3842
Removed 1 four-fermion operator	64 operators	Nomura 0911.1941
Removed 4 four-fermion operators	60 operators	JAAS 1008.3562
Removed 1 four-scalar operator	59 operators	Grzadkowski, Iszkrzynski, Misiak, Rosiek 1008.4884

Fact #4

The origin of the *insensitivity* is that

$$\bar{b}(p_2)(p_1 + p_2)^\mu t(p_1) \cdot \varepsilon_\mu^\pm(p_1 - p_2)^* = 0$$



$$\rho_W = \rho_W^{\text{SM}} + \begin{pmatrix} 0 & 0 & 0 \\ 0 & \times & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

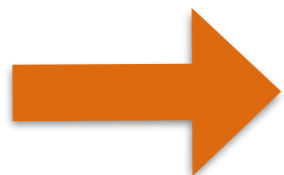
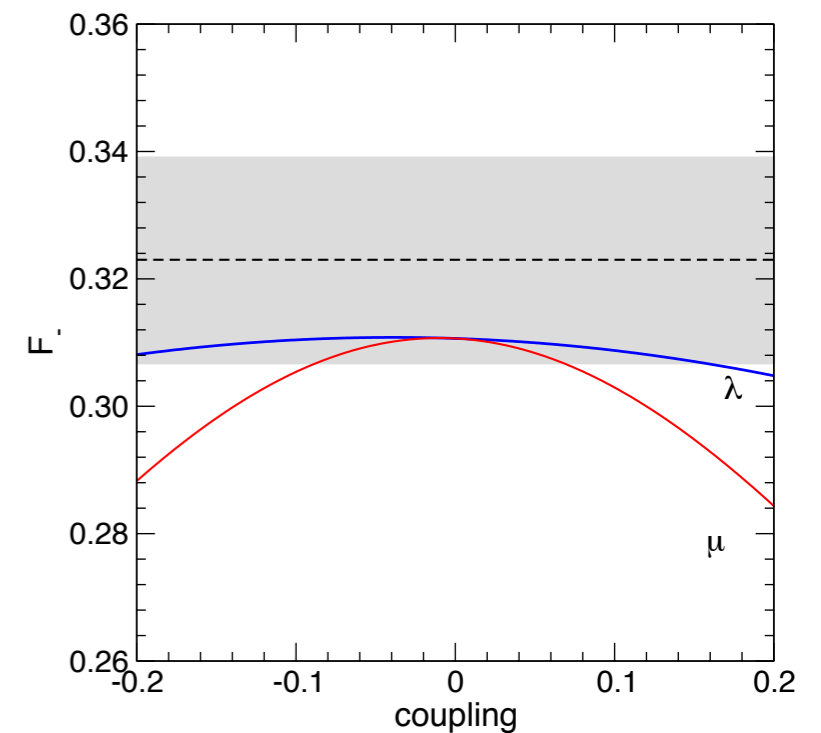
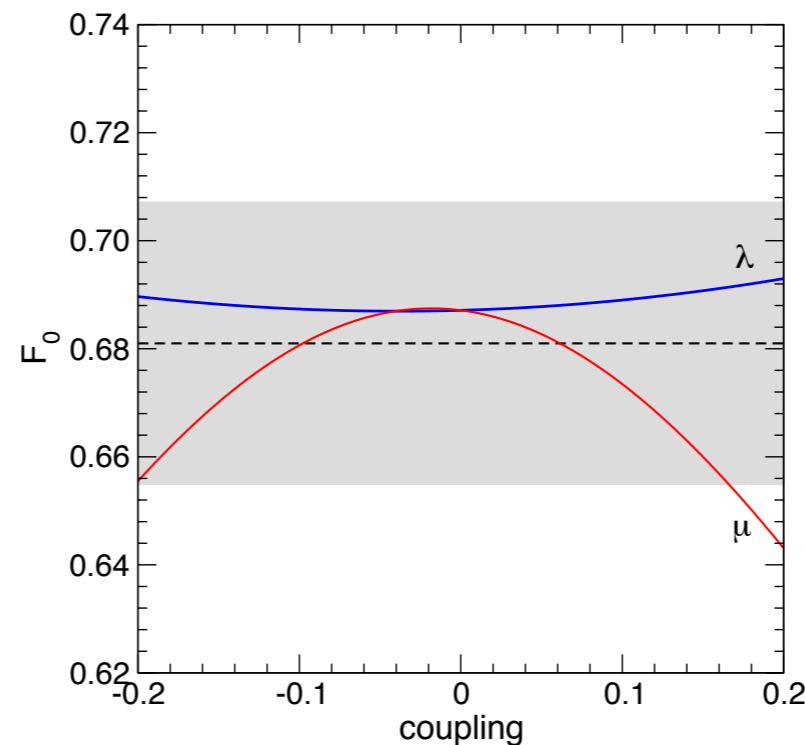
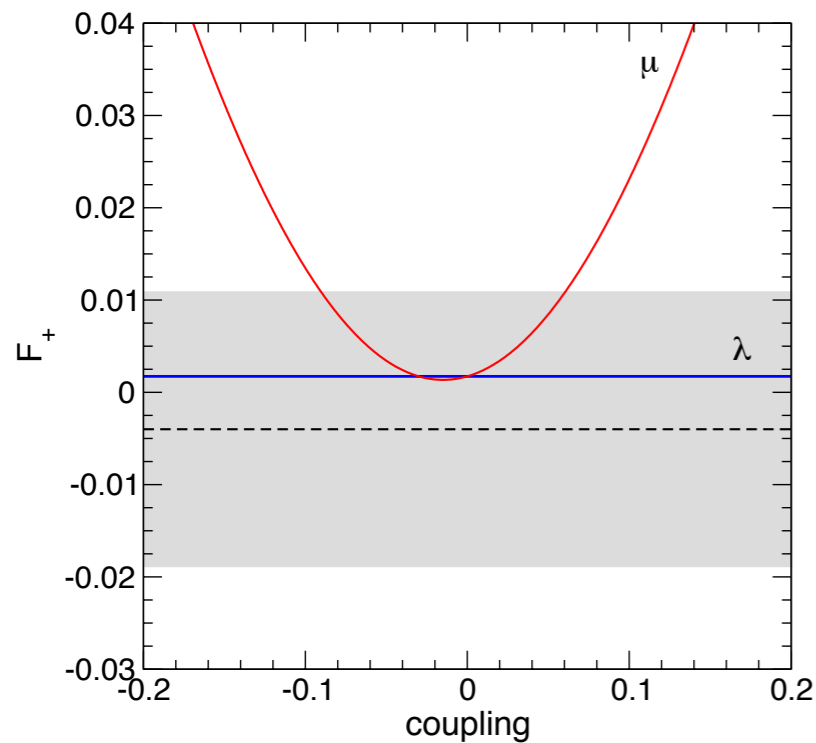


Therefore, the dependence of F_+ , F_- (and all W polarisations) on these operators is residual, stemming from Γ_0 in the denominator

Replacing V_R and g_L by two orthogonal combinations

$$\begin{pmatrix} \lambda \\ \mu \end{pmatrix} = \frac{1}{[m_t^2 + M_W^2]^{\frac{1}{2}}} \begin{pmatrix} m_t & M_W \\ -M_W & m_t \end{pmatrix} \begin{pmatrix} V_R \\ g_L \end{pmatrix}$$

the (in)sensitivity of helicity fractions to λ is apparent



The λ problem

But there is life beyond W helicity fractions!

JAAS & Bernabéu I508.04592

Being a spin-1 particle, the W boson has no less than

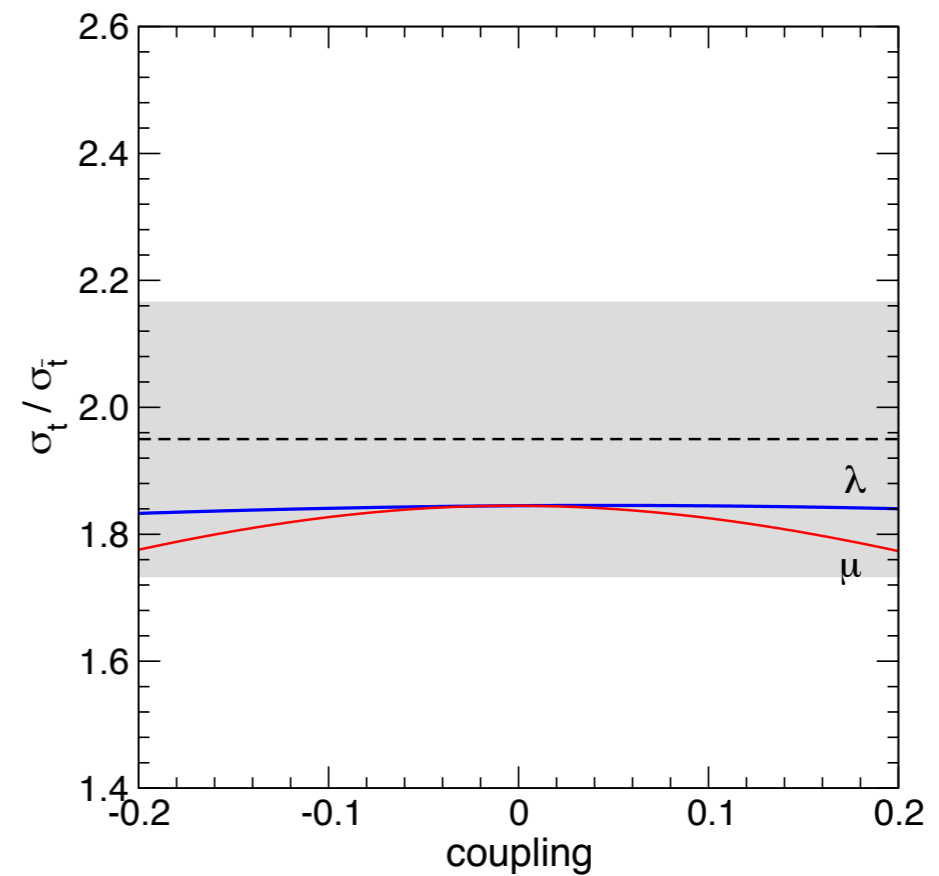
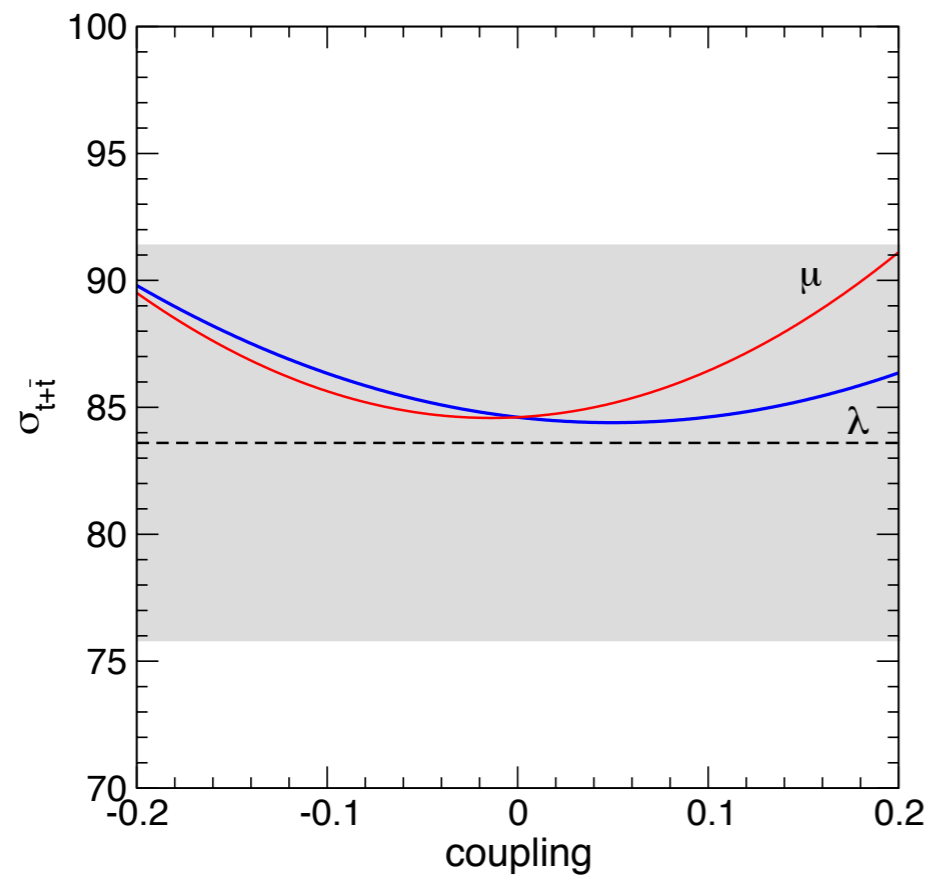
8 (eight!) polarisation observables

$$\begin{aligned} \frac{1}{\Gamma} \frac{d\Gamma}{d\cos\theta^* d\phi^*} = & \frac{3}{8\pi} \left\{ \frac{1}{2} (1 + \cos^2\theta^*) + \langle S_3 \rangle \cos\theta^* \right. \\ & + \left[\frac{1}{6} - \frac{1}{\sqrt{6}} \langle T_0 \rangle \right] (1 - 3\cos^2\theta^*) \\ & + \langle S_1 \rangle \cos\phi^* \sin\theta^* + \langle S_2 \rangle \sin\phi^* \sin\theta^* \\ & - \langle A_1 \rangle \cos\phi^* \sin 2\theta^* - \langle A_2 \rangle \sin\phi^* \sin 2\theta^* \\ & \left. + \langle B_1 \rangle \cos 2\phi^* \sin^2\theta^* + \langle B_2 \rangle \sin 2\phi^* \sin^2\theta^* \right\} \end{aligned}$$

first measured
by ATLAS

Their measurement will improve the global limits when the precision is better, but does not solve the λ problem

t-channel single top cross sections depend on λ and μ but have too large uncertainties to constrain λ effectively

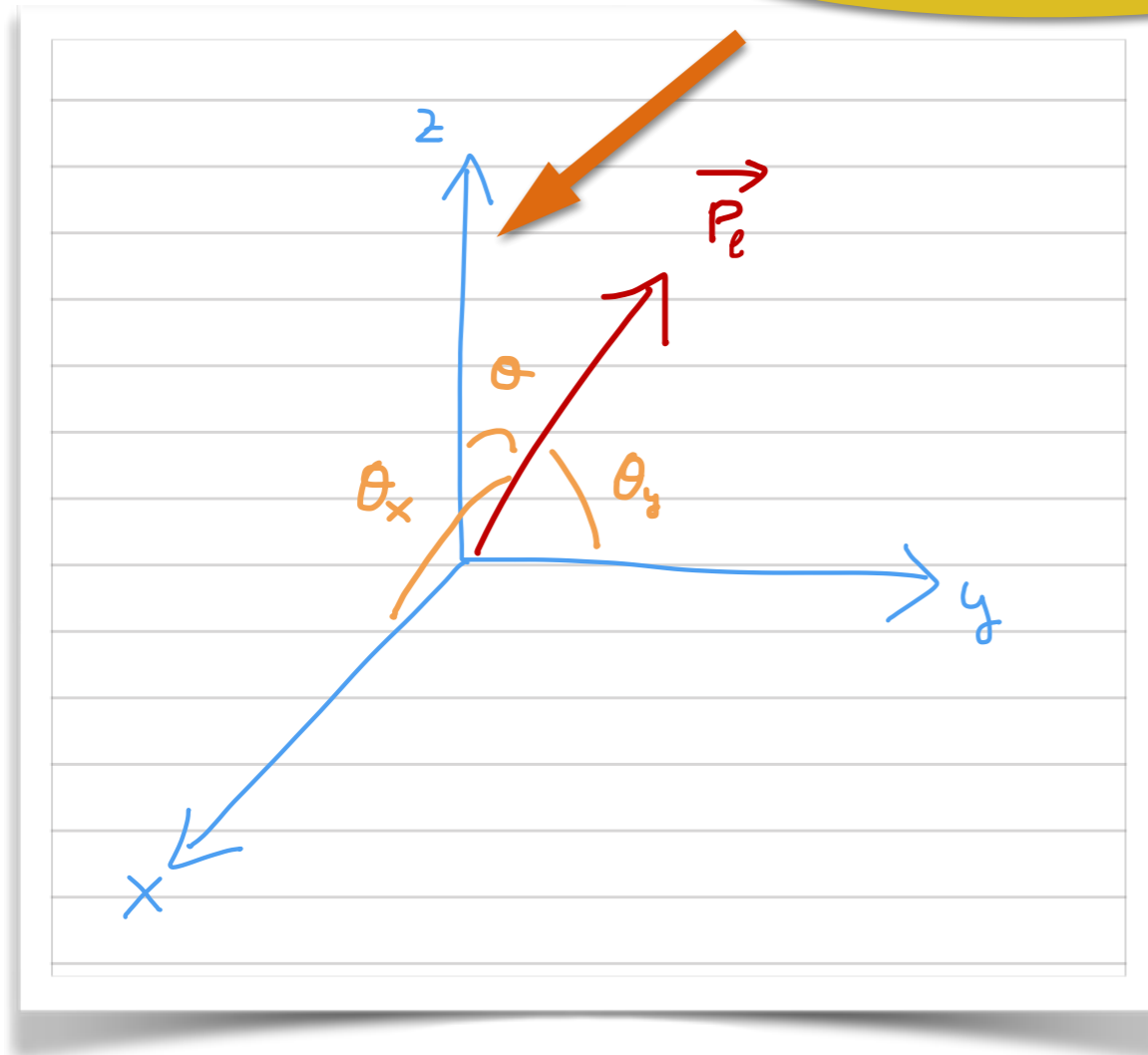


Proposal #1

More polarisation measurements and with higher precision

JAAS & Santos I404.I585

spectator jet direction



$$\frac{1}{\sigma} \frac{d\sigma}{d \cos \theta} = \frac{1}{2} (1 + \alpha P_z \cos \theta)$$

$$\frac{1}{\sigma} \frac{d\sigma}{d \cos \theta_x} = \frac{1}{2} (1 + \alpha P_x \cos \theta_x)$$

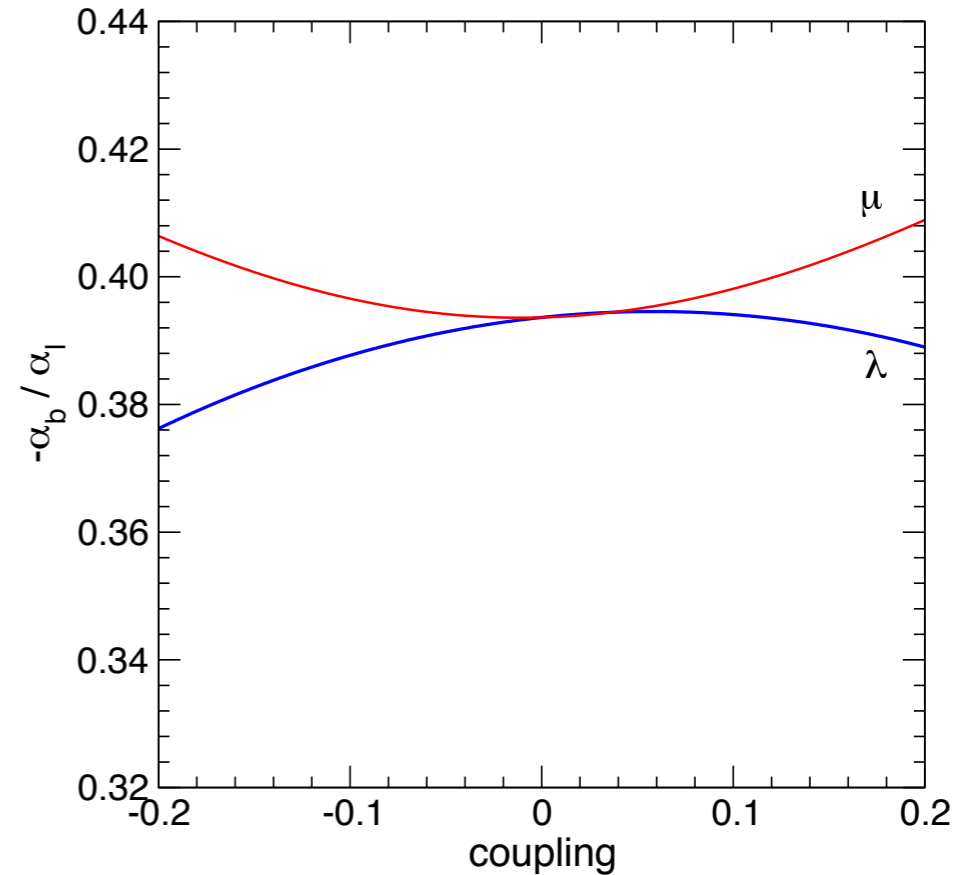
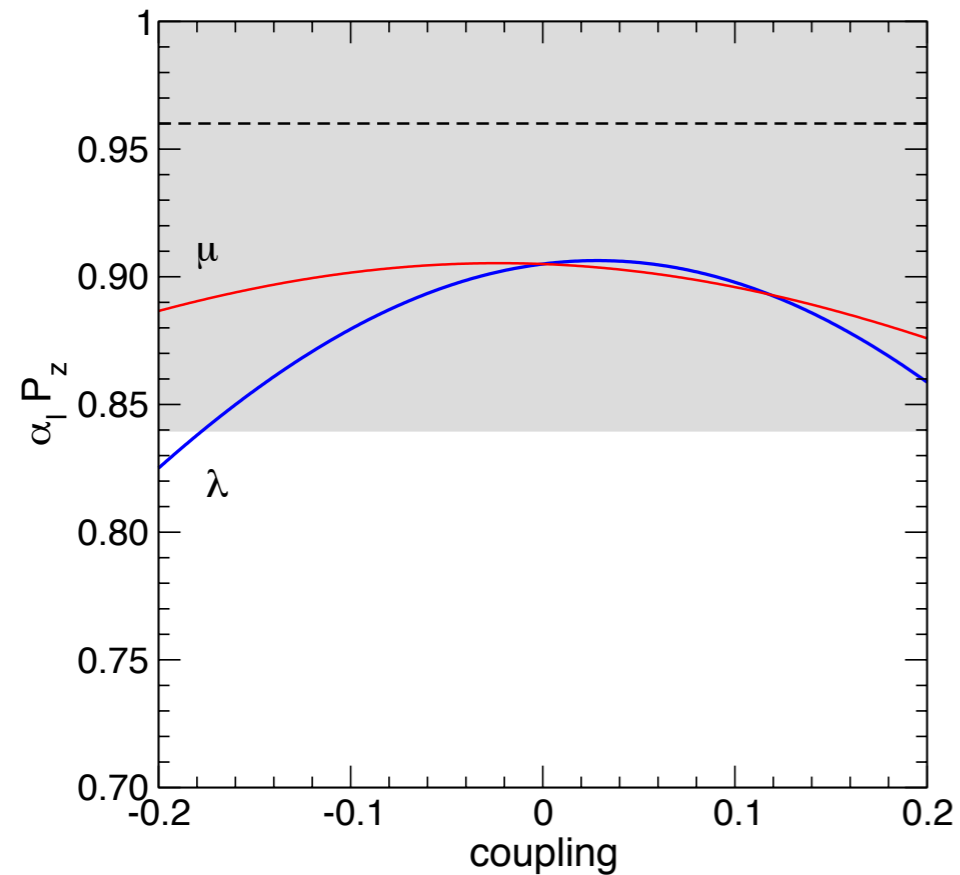
$$\frac{1}{\sigma} \frac{d\sigma}{d \cos \theta_y} = \frac{1}{2} (1 + \alpha P_y \cos \theta_y)$$

also: model-independent measurements

JAAS & Herrero-Hahn I208.6006

$$\alpha_\ell P_z = 0.96 \pm 0.05 \text{ (stat)} \pm 0.10 \text{ (syst)}$$

ATLAS



Improve $\alpha_l \cdot P_z$ measurement (?)

and measure P_x, P_y

Measure α_b / α_l [maybe at ILC?]

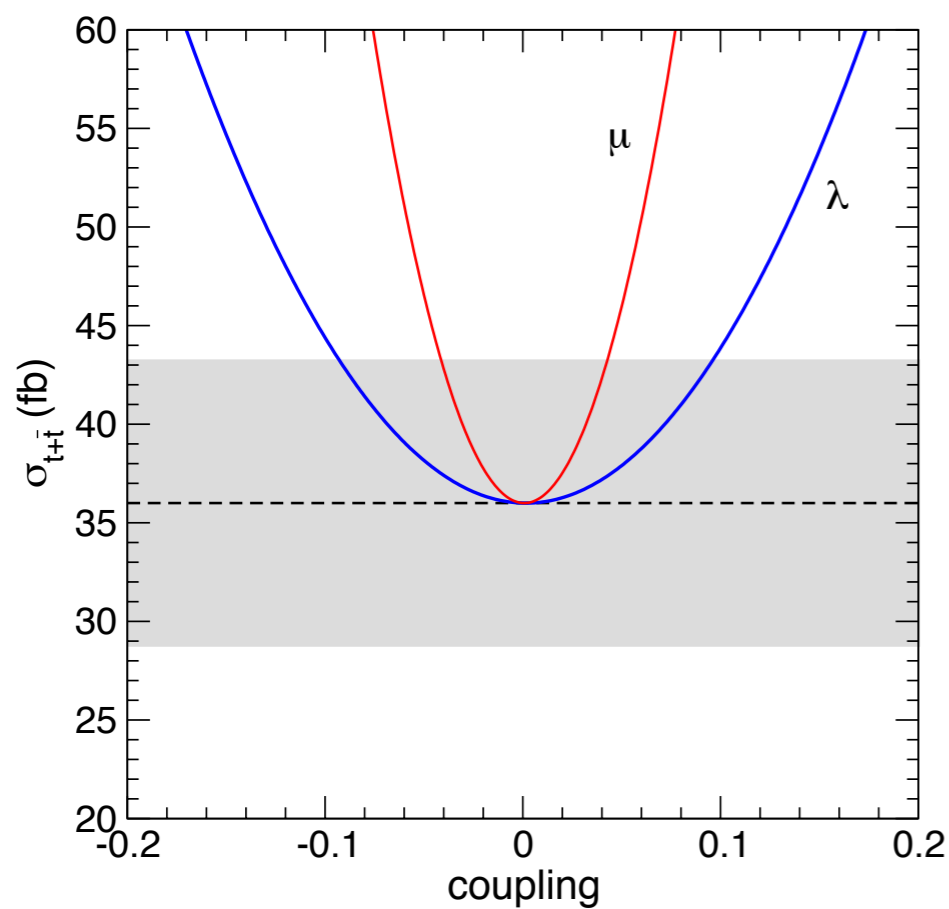
$$\alpha_b / \alpha_\ell = ? \pm ?$$

Proposal #2

Measurements at high Q , even with low precision, can be very constraining

JAAS & MLM '14

$$\sigma_{t\bar{b}+\bar{t}b}(m_{tb} > 1 \text{ TeV})$$



20% uncertainty

Final remarks

Besides measuring W helicity fractions 1205.2484, again CMS-PAS-TOP-12-015, and again CMS-PAS-TOP-12-020, and again 1308.3879, and again 1410.1154, and again CMS-PAS-TOP-14-017, and again 1605.09047,

one should consider other observables. Fortunately, new polarisation measurements are becoming available.

Many things still to be done at the pheno side [my incomplete to-do list]

- Can we possibly get limits on all 4f operators? 90 of them contribute to single top
- Global fit to top ttV operators: $O_{\varphi q}^{(1)}$, $O_{\varphi q}^{(3)}$, $O_{\varphi u}$, $O_{\varphi \varphi}$, O_{uW} , O_{dW} , $O_{uB\varphi}$, $O_{uG\varphi}$

This will trigger proposals for new measurements, stay tuned.

