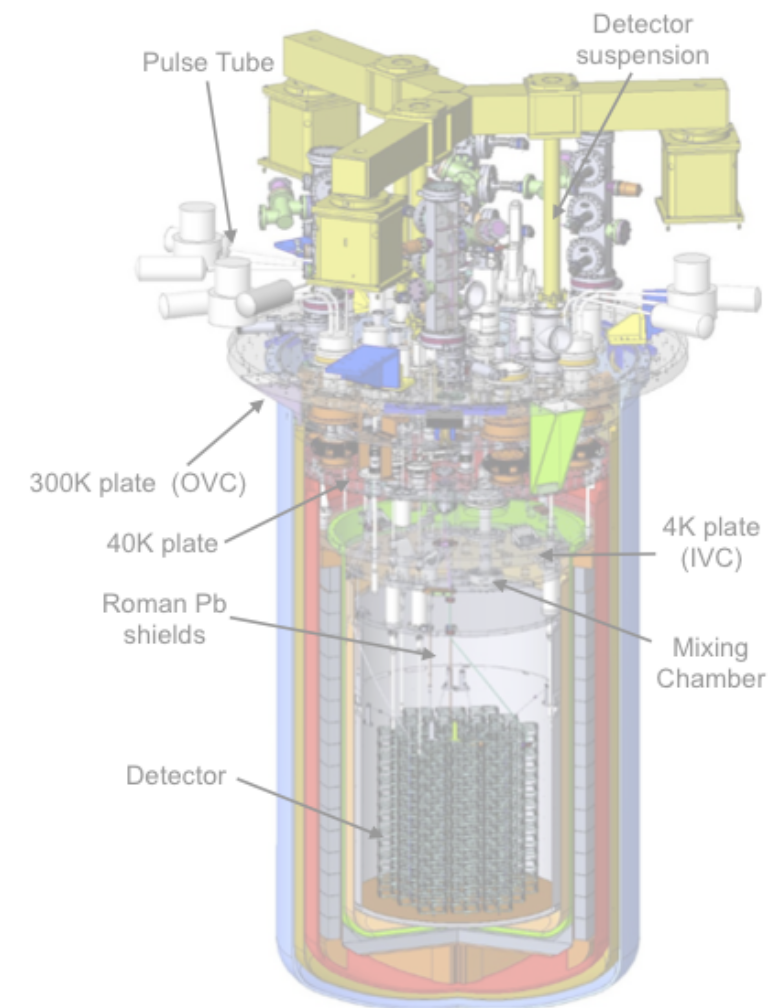


Neutrinoless $\beta\beta$ decay @ LNGS



99th Plenary ECFA meeting

June 30, 2016 - LNGS

Carlo Bucci

INFN - Laboratori Nazionali del Gran Sasso

Present knowledge about neutrinos

- neutrinos are massive fermions
- there are 3 active neutrino flavors (ν_α)
- neutrino flavor states are mixtures of mass states (ν_k)

$$|\nu_\alpha\rangle = \sum_k U_{\alpha k} |\nu_k\rangle$$

Pontecorvo–Maki–Nakagawa–Sakata matrix

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Atmospheric /
Accelerator

Reactor /
Accelerator

Solar /
Reactor

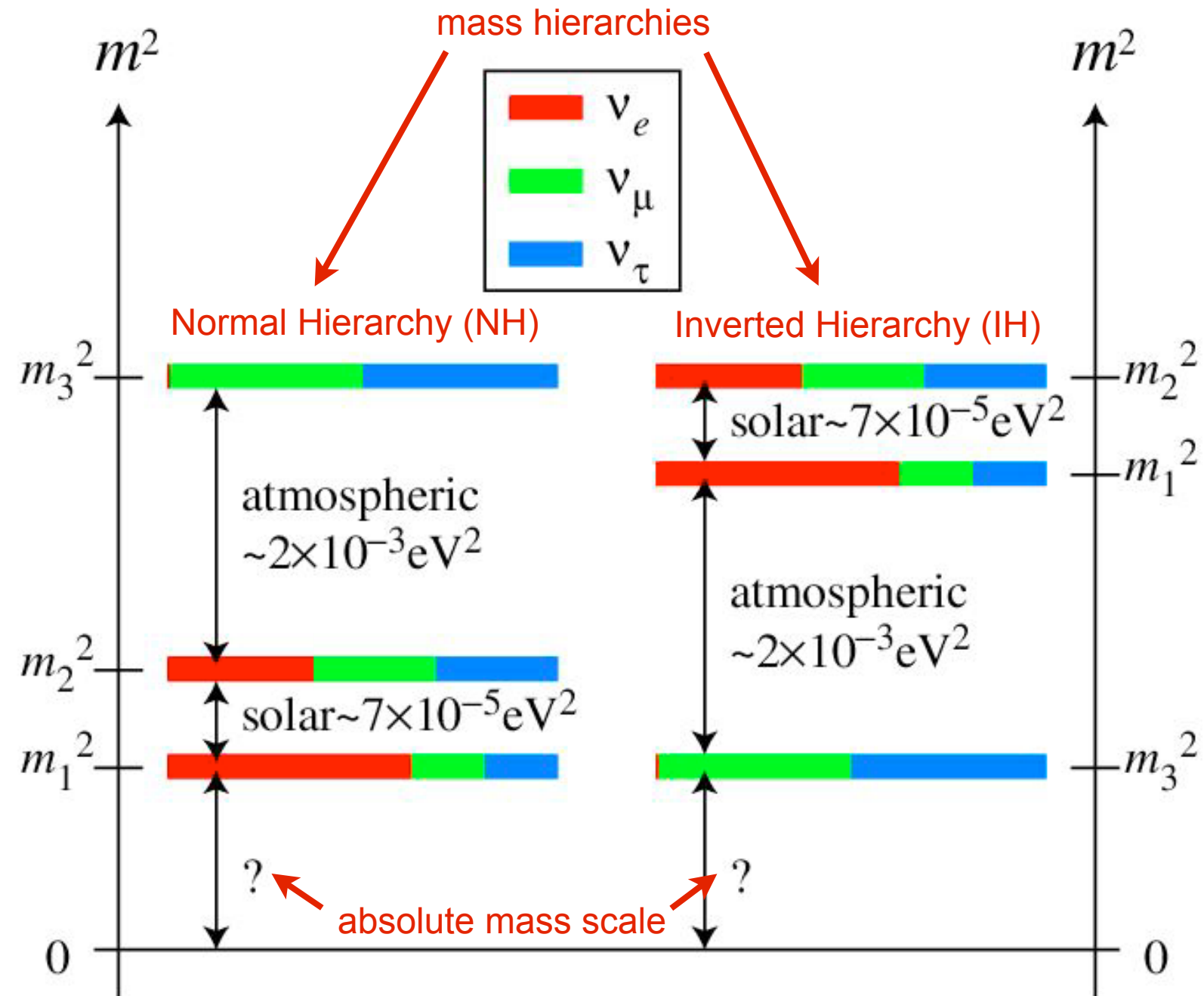
Measurements of neutrino parameters from:

- neutrino oscillations
- single beta decay
- cosmology
- neutrinoless double beta decay

Open questions

$0\nu\text{-}\beta\beta$ can give an answer to three fundamental questions:

- Dirac or Majorana nature
- absolute mass scale:
mass of the lightest ν
- hierarchy of masses

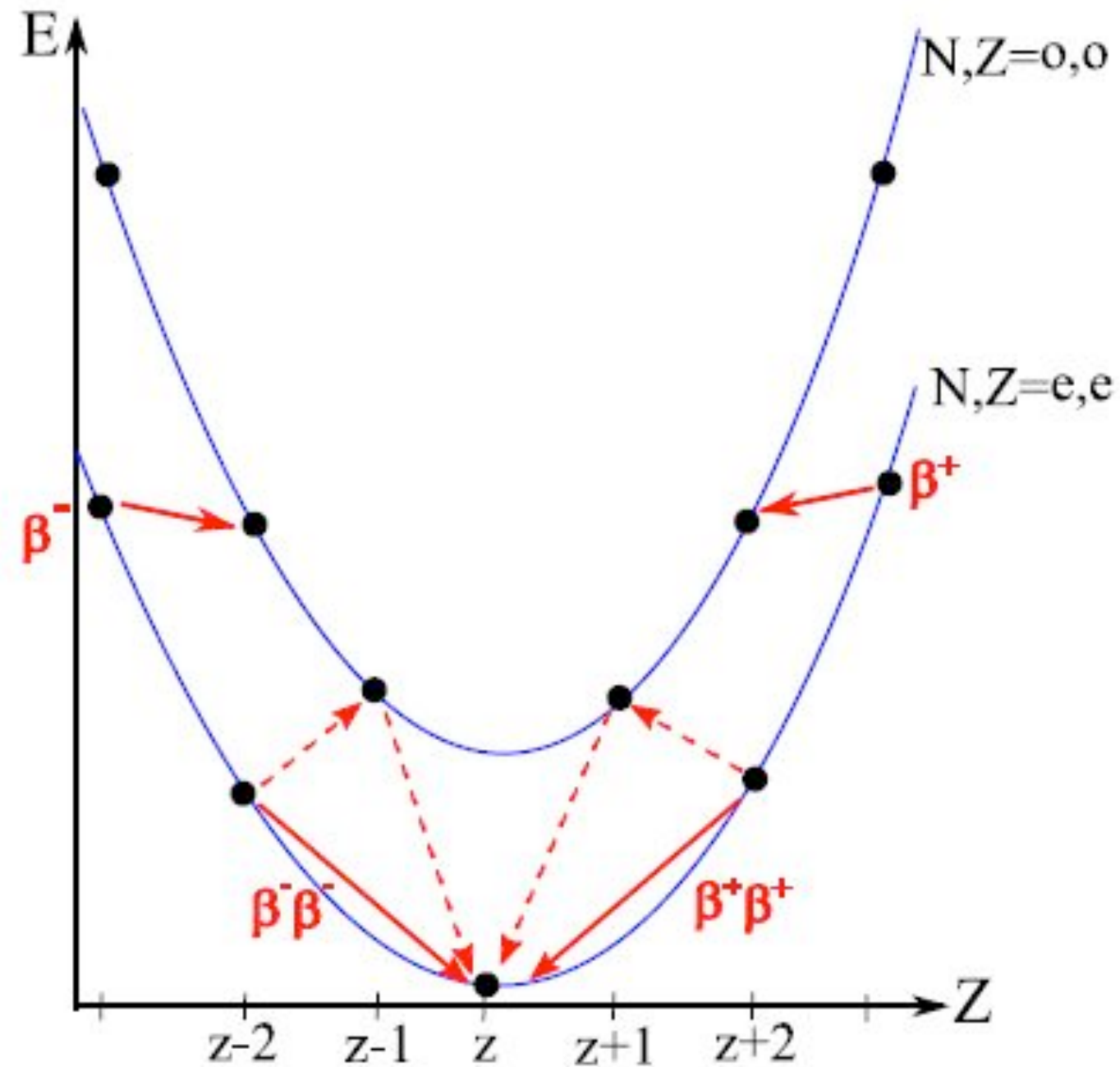


Oscillation experiments can determine the hierarchy but are blind to the other two questions

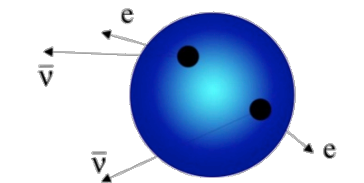
Double beta decay

Very rare nuclear decay $(A,Z) \rightarrow (A,Z+2) + 2e^- (+?)$

Isotope	Q-value [keV]	Isotopic abundance
^{48}Ca	4272	0,19
^{76}Ge	2039	7,8
^{82}Se	2996	9,2
^{96}Zr	3350	2,8
^{100}Mo	3034	9,6
^{116}Cd	2814	7,6
^{130}Te	2527	33,4
^{136}Xe	2459	8,9
^{150}Nd	3371	5,6



Double beta decay

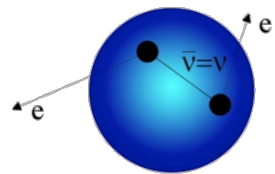
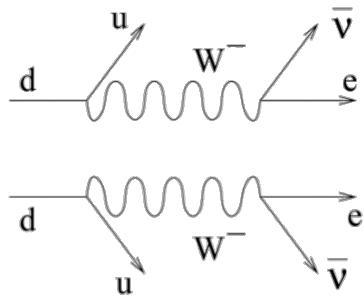


$$(A, Z) \rightarrow (A, Z + 2) + 2e^- + 2\bar{\nu}$$

2ν-DBD

2nd order process allowed in the SM

observed in several nuclei with $\tau^{2\nu} \sim 10^{19}-10^{21}$ y



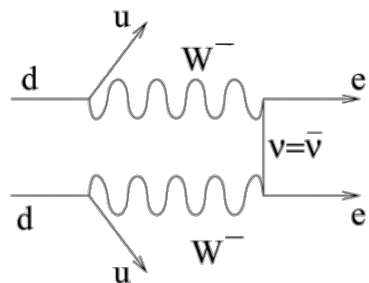
$$(A, Z) \rightarrow (A, Z + 2) + 2e^-$$

0ν-DBD (implies physics beyond SM)

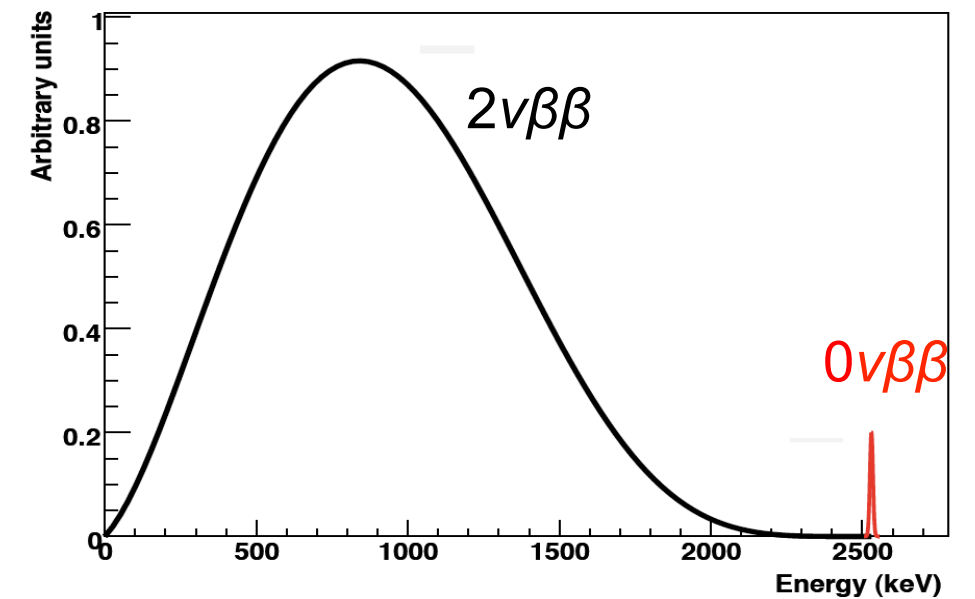
lepton number violating process

$\tau^{0\nu} > 10^{24}-10^{25}$ y

exists if neutrino is a Majorana particle and $m_\nu \neq 0$

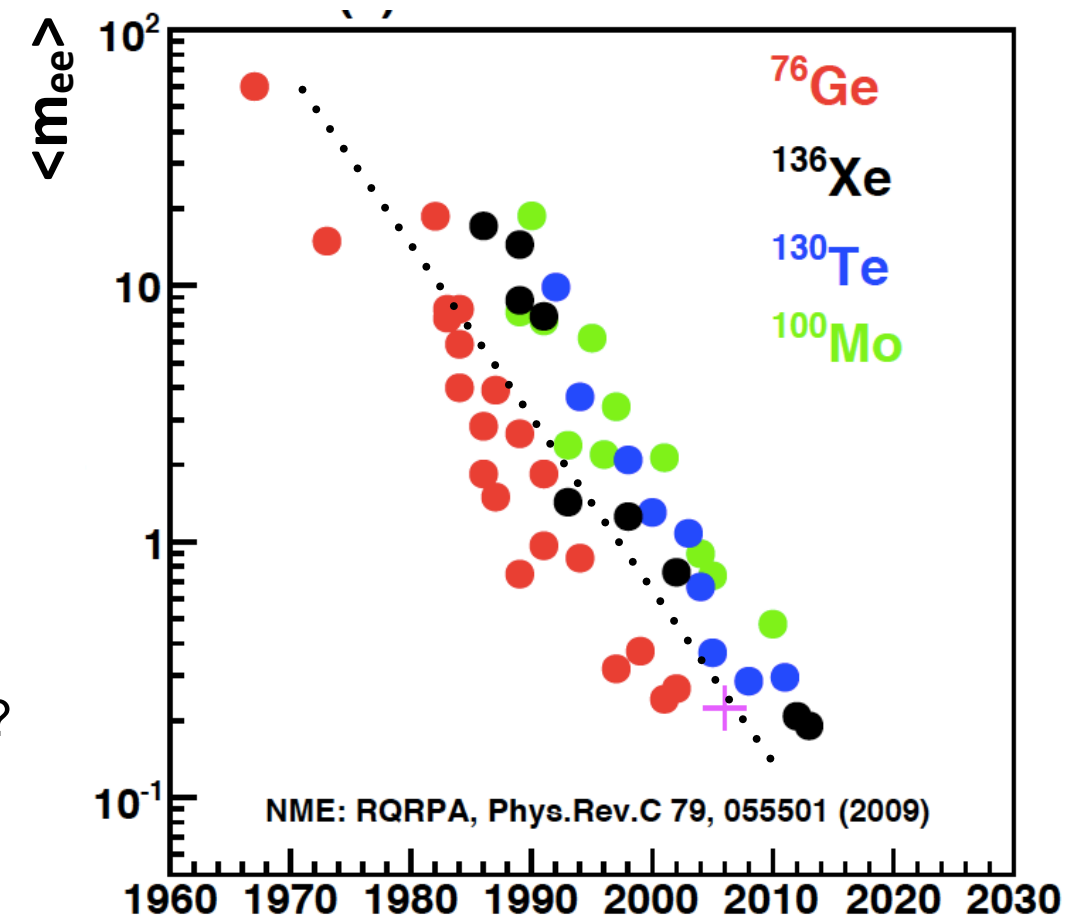


$\beta\beta$ summed e^- energy spectrum

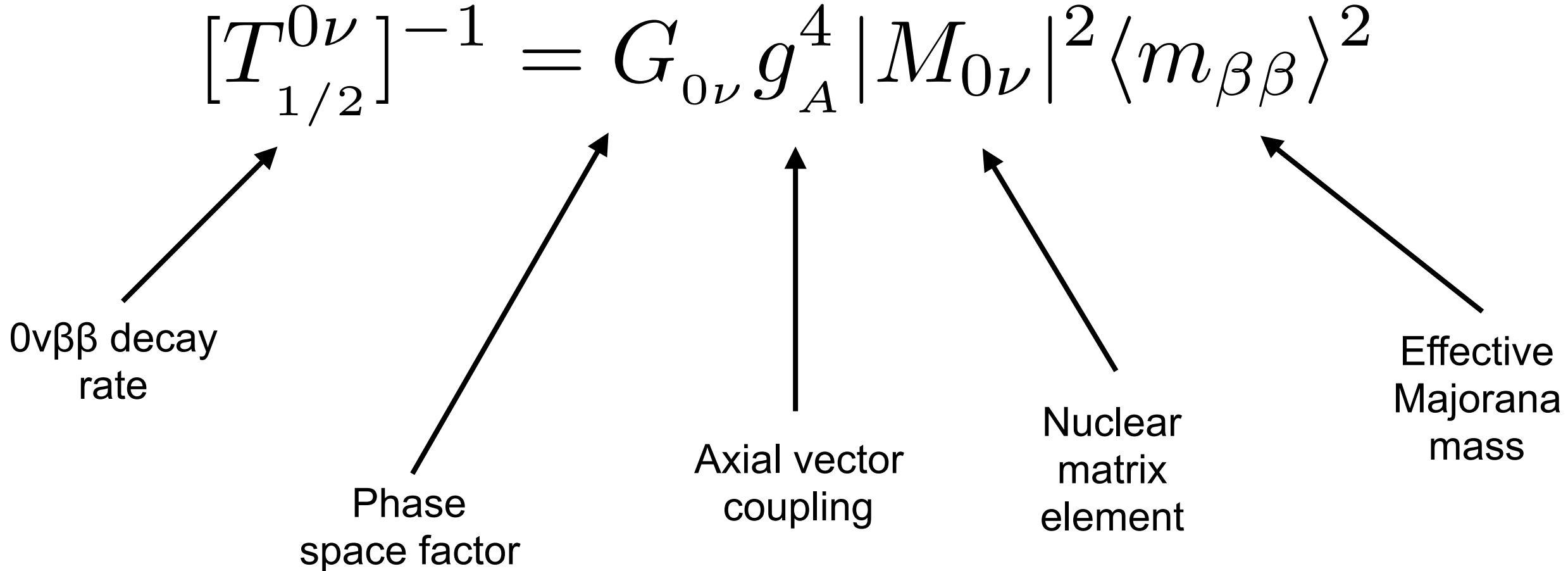


A long history of $0\nu\text{-}\beta\beta$ experiments

An order of magnitude on the effective Majorana mass every 15 years?



$0\nu\text{-}\beta\beta$ and Majorana mass

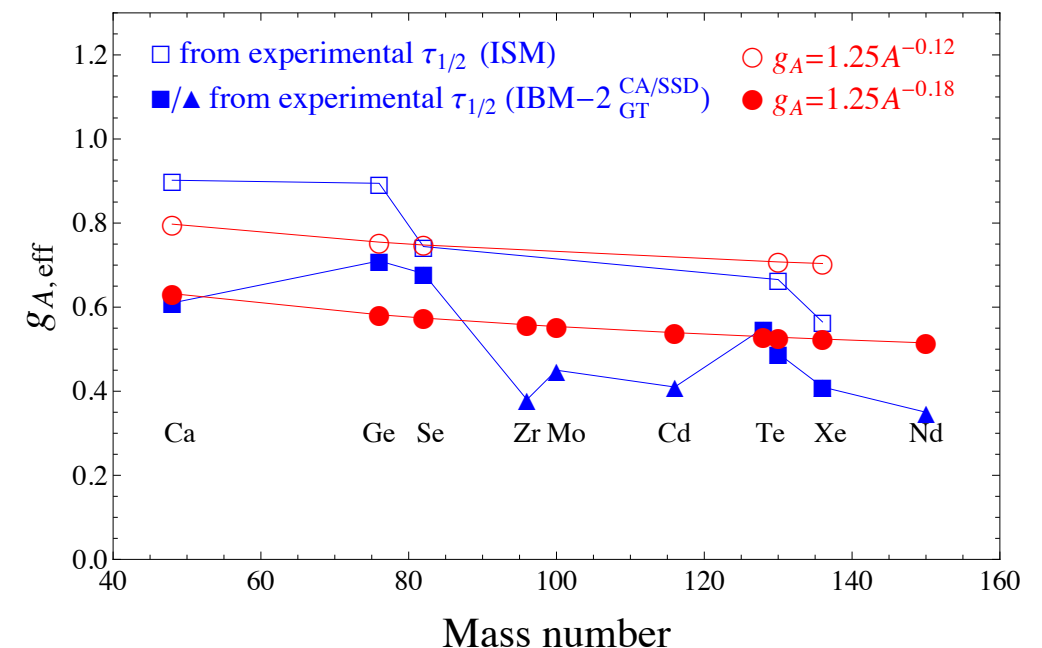
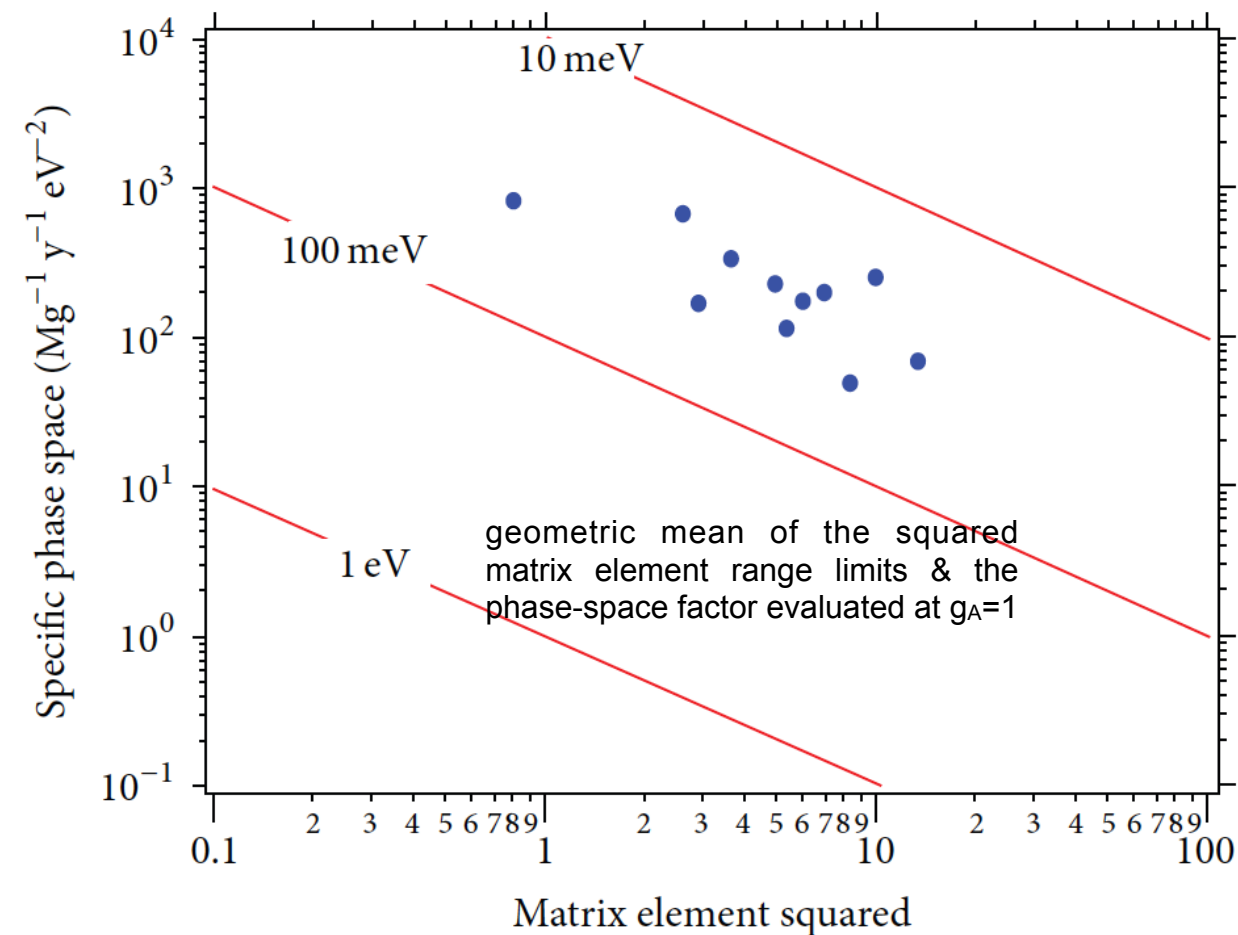
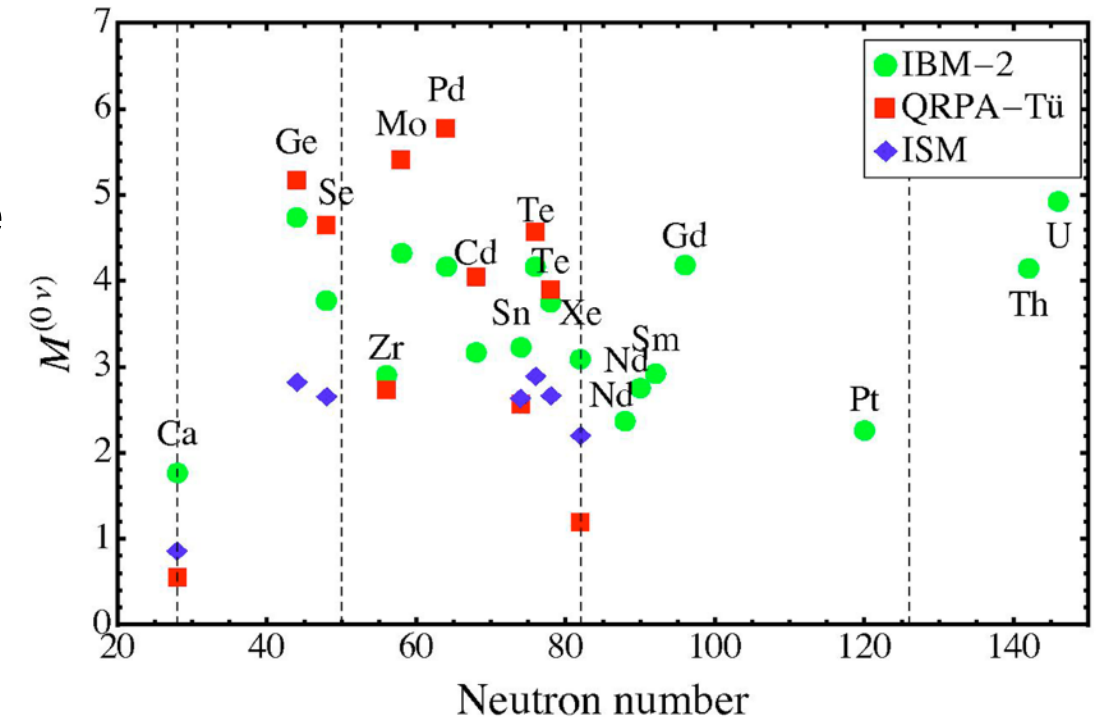


$$\langle m_{\beta\beta} \rangle = \sum_k U_{ek}^2 m_k = c_{12}^2 c_{13}^2 m_1 + s_{12}^2 c_{13}^2 e^{i\alpha} m_2 + s_{13}^2 e^{i\beta} m_3$$

NEUTRINO MASS EIGENVALUES
 NEUTRINO MIXING MATRIX

Is there a preferred isotope?

- Nuclear matrix elements calculations are rapidly improving and today the differences between different methods (IBM, QRPA, ISM) are much smaller than in the past
- Uncertainty on g_A plays a relevant role
factor 2 in g_A is a factor 16 in decay rate
- inverse correlation observed between phase space and square of the nuclear matrix element

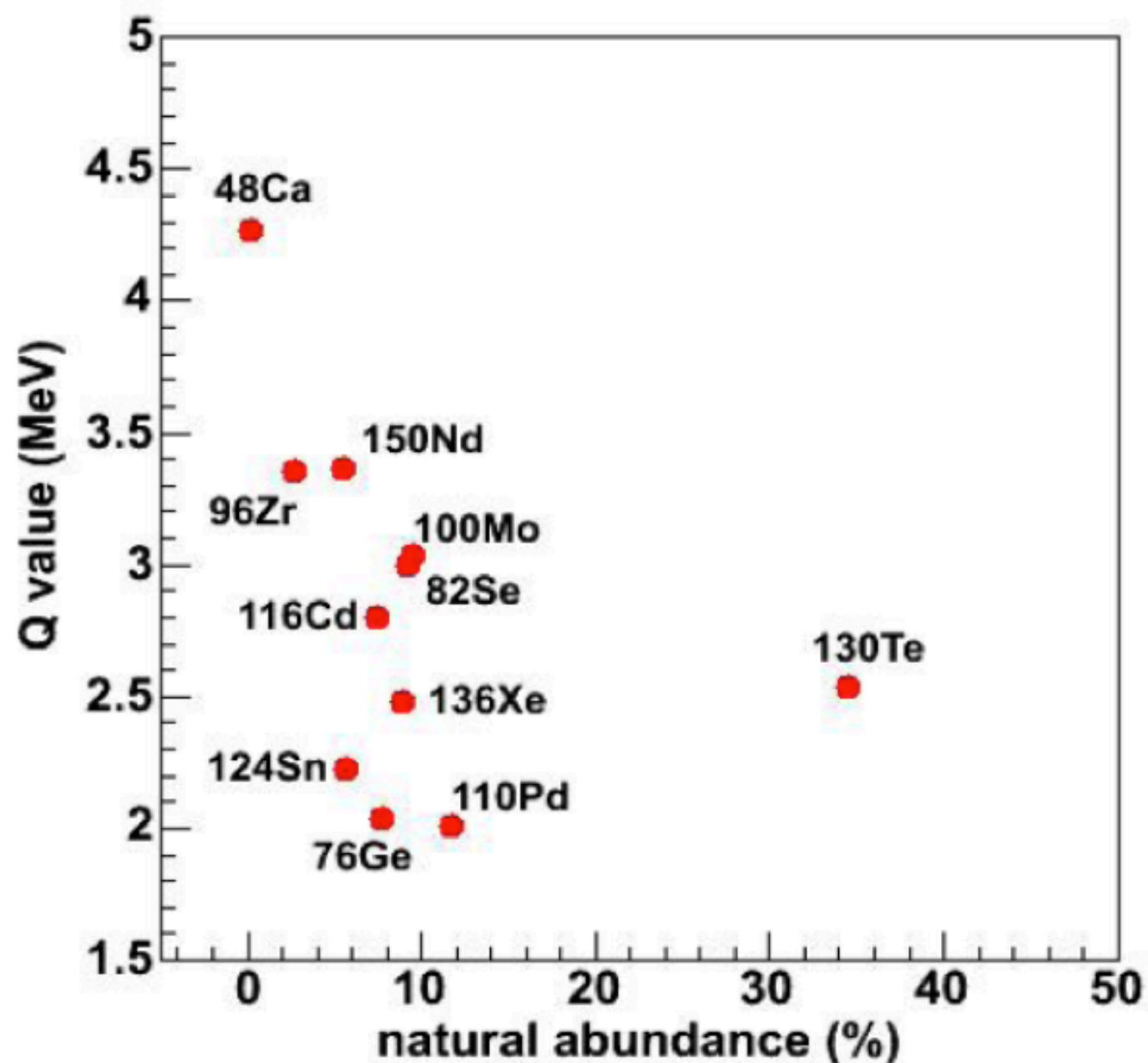


Isotope choice

In many cases driven by the detector characteristics.

- ^{76}Ge with Germanium diodes
- ^{136}Xe with Xenon TPCs
- bolometers and scintillators have multiple choices

- Isotopic abundance as high as possible
 - money issue
- Q-value as high as possible
 - lower environmental background
- 2ν -DBD half-life as high as possible
 - energy resolution



Sensitivity

Half-life corresponding to the minimum detectable number of events over background at a given confidence level

$$S_{0\nu} = \ln(2) N_A \frac{\eta \cdot \epsilon}{W} \sqrt{\frac{M \cdot T}{\Delta E \cdot B}}$$

finite background: $M \cdot T \cdot B \cdot \Delta E > 1$

M: active detector mass [kg]

T: measurement life time [anni]

B: background in the ROI [counts keV⁻¹ kg⁻¹ y⁻¹]

W: molecular weight

$$S_{0\nu} = \ln(2) N_A \frac{\eta \cdot \epsilon}{W} M \cdot T$$

zero background: $M \cdot T \cdot B \cdot \Delta E \approx 1$

N_A: Avogadro number

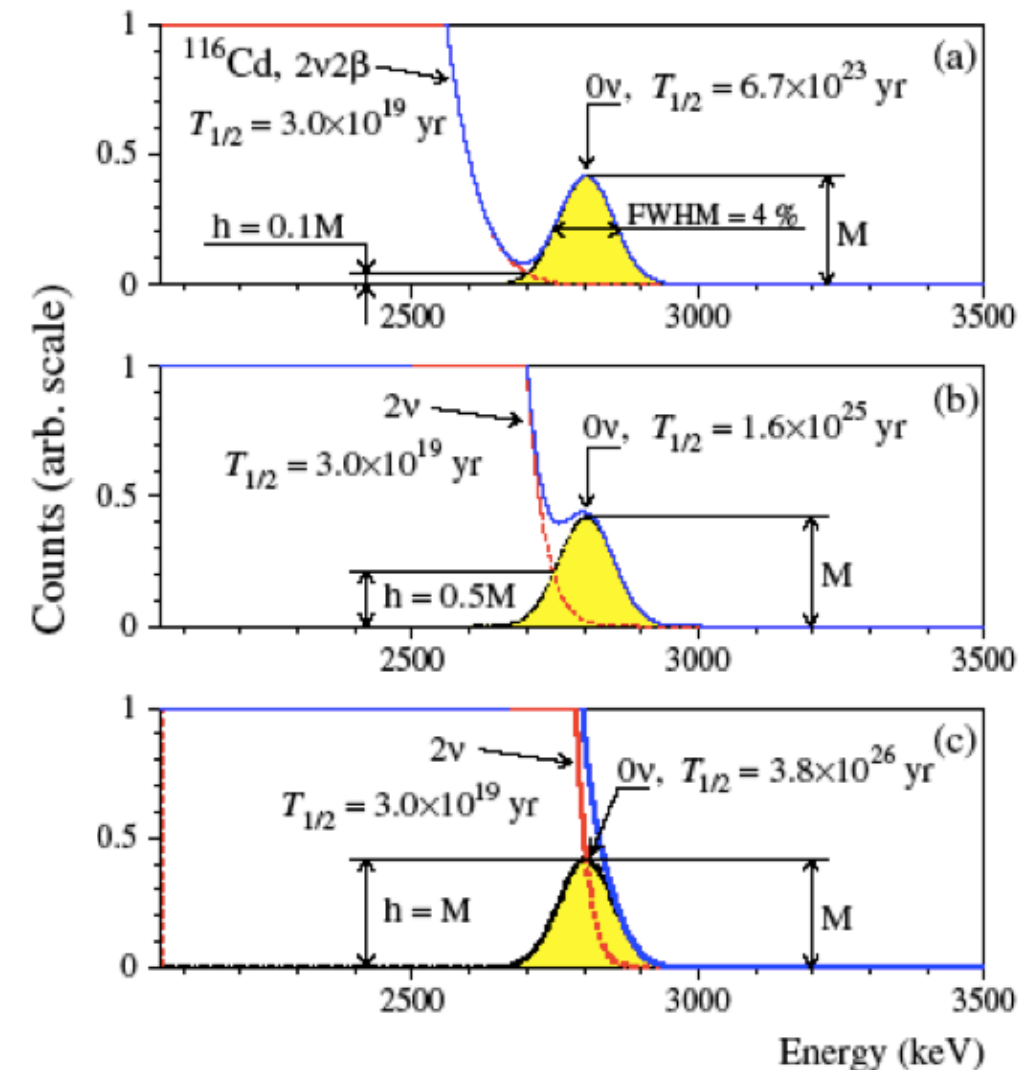
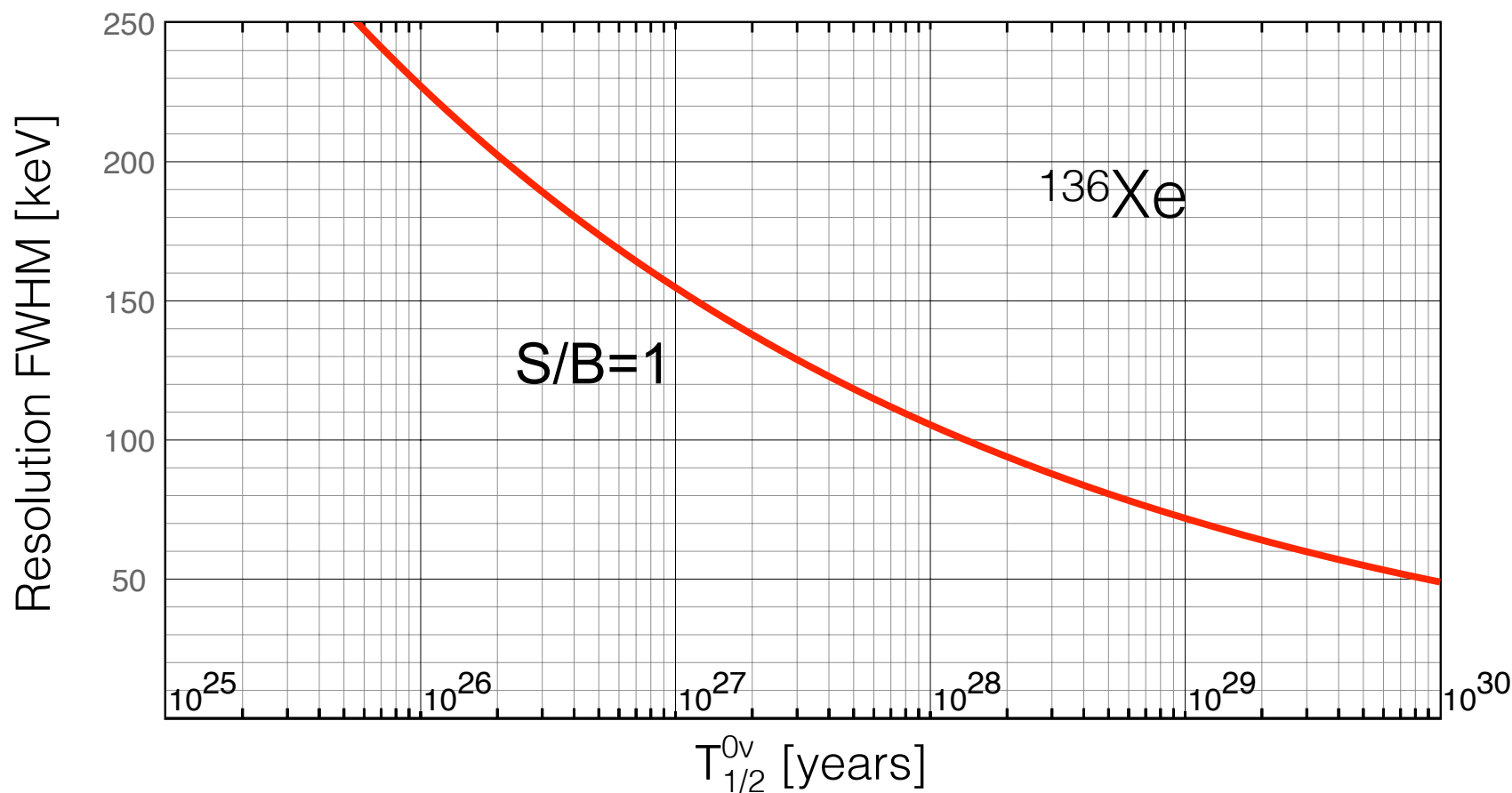
η: isotopic abundance

ε: detector efficiency

ΔE: FWHM energy resolution @ Q-value

Irreducible background from $2\nu\text{-}\beta\beta$

- The irreducible background induced by the $2\nu\text{-}\beta\beta$ could be mitigated just by the energy resolution
- The effect can be partially attenuated with an asymmetric ROI (but losing efficiency)



Zdesenko, Danevich, Tretyak, J. Phys. G 30 (2004) 971

$$\frac{S}{B} = \frac{m_e}{7Q\delta^6} \frac{\Gamma_{0\nu}}{\Gamma_{2\nu}} = \frac{m_e}{7Q\delta^6} \frac{T_{1/2}^{2\nu}}{T_{1/2}^{0\nu}}$$

Key ingredients

Number of nuclei of the chosen isotope

- Detector mass
- Isotopic abundance -> enrichment -> money

Background

- radio purity and shieldings
- discrimination capability

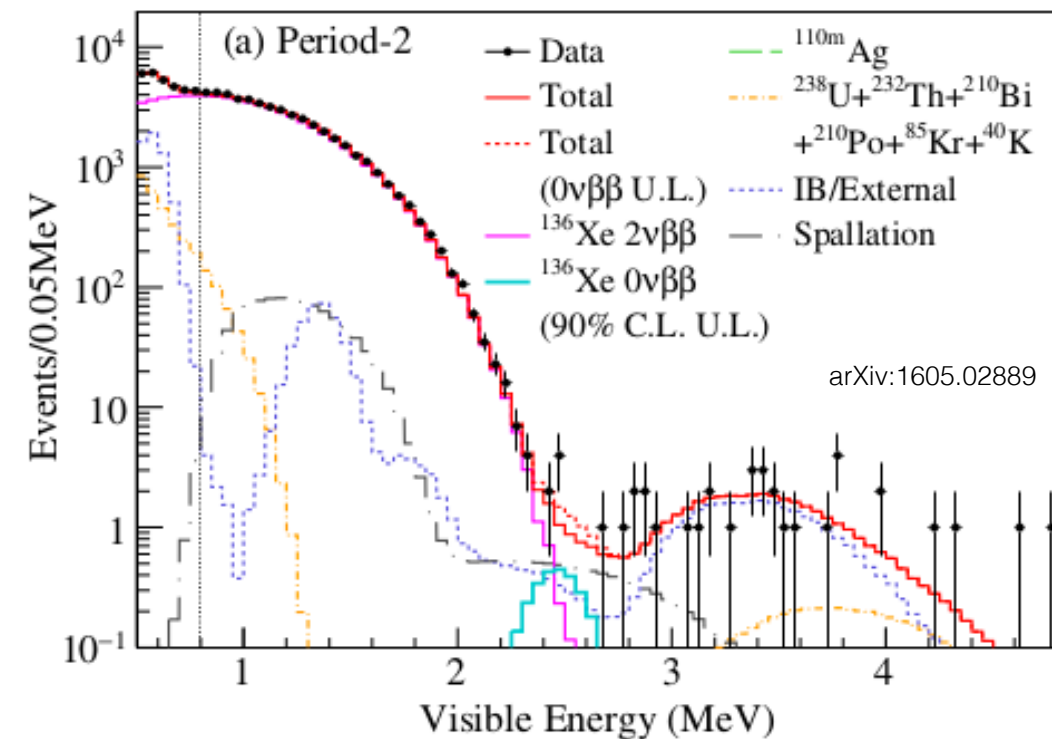
Energy resolution

- detector dependent
- easier identification of the background contributions
- less probability of accidental superposition with environmental radioactivity peaks (e.g. 2448 keV of ^{214}Bi at ~ 10 keV from ^{136}Xe Q-value
2505 keV sum of the ^{60}Co peaks at ~ 20 keV from ^{130}Te Q-value)
- mitigate the background induced by 2ν -DBD

Double beta decay experiments: limits vs discovery

- Solid state detectors: have better resolution but is more difficult to reach very low background and increase mass
 - ➔ discovery potential
- Liquid scintillators and TPCs have poor energy resolution but its easier to reach large masses and they can be purified
 - ➔ in case of a positive signal difficult to disentangle from possible background sources

Kamland-zen



LNGS has a leadership role in “discovery” experiments

Neutrinoless $\beta\beta$ decay @ LNGS

A long history in $0\nu\beta\beta$ search

- MiDBD (^{130}Te)
- Heidelberg-Moscow (^{76}Ge)
- Cuoricino (^{130}Te)
- GERDA-I (^{76}Ge)
- CUORE-0 (^{130}Te)

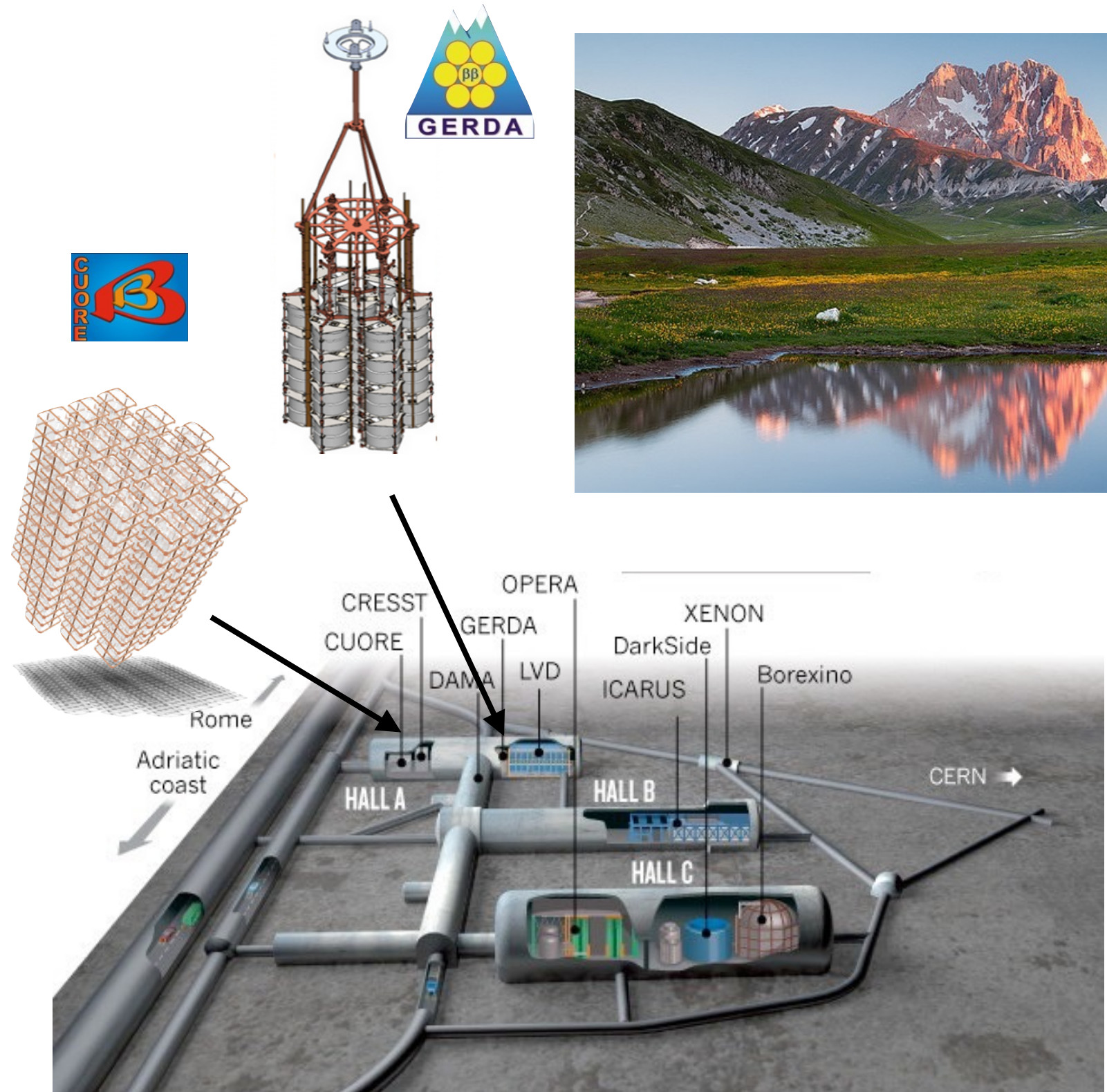
R&D projects

- Cobra (^{116}Cd)
- CUPID-0/LUCIFER (^{82}Se)
- DAMA R&D (^{116}Cd)

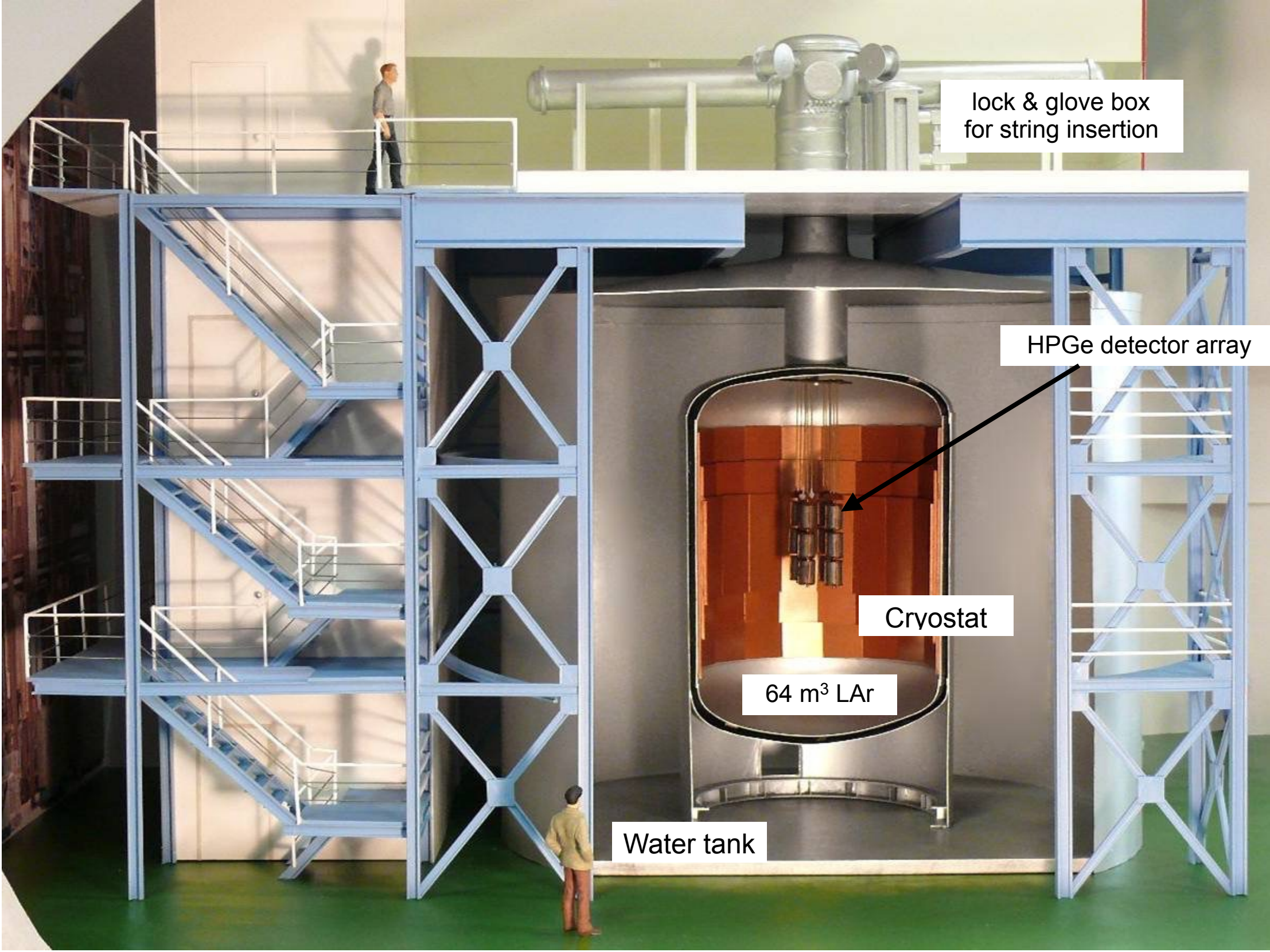
important role in the near future

- GERDA-II (^{76}Ge)
- CUORE (^{130}Te)

- ~ 3600 m.w.e. deep
- μs : $\sim 3 \times 10^{-8}/(\text{s cm}^2)$
- γs : $\sim 0.73/(\text{s cm}^2)$
- neutrons: $4 \times 10^{-6} \text{ n}/(\text{s cm}^2)$

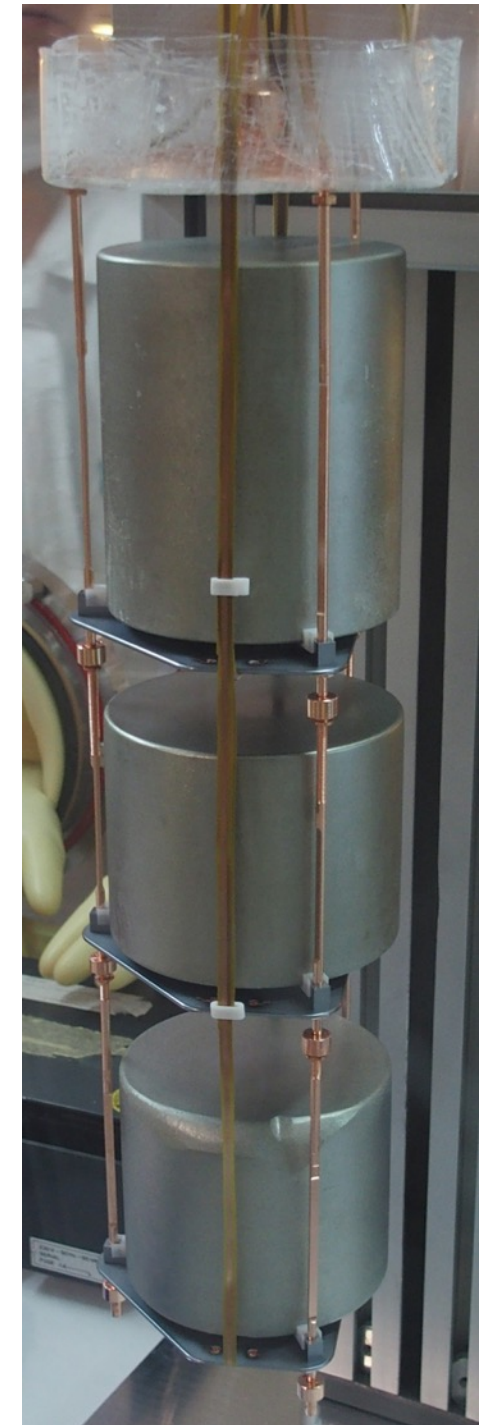
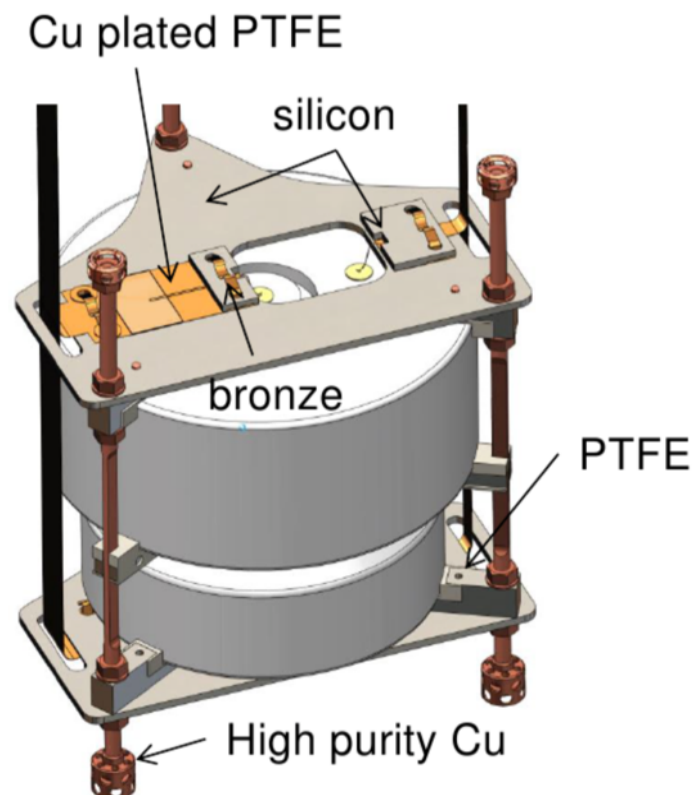
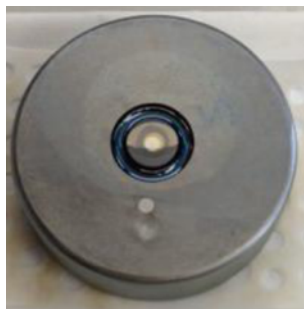


GERmanium Detector Array

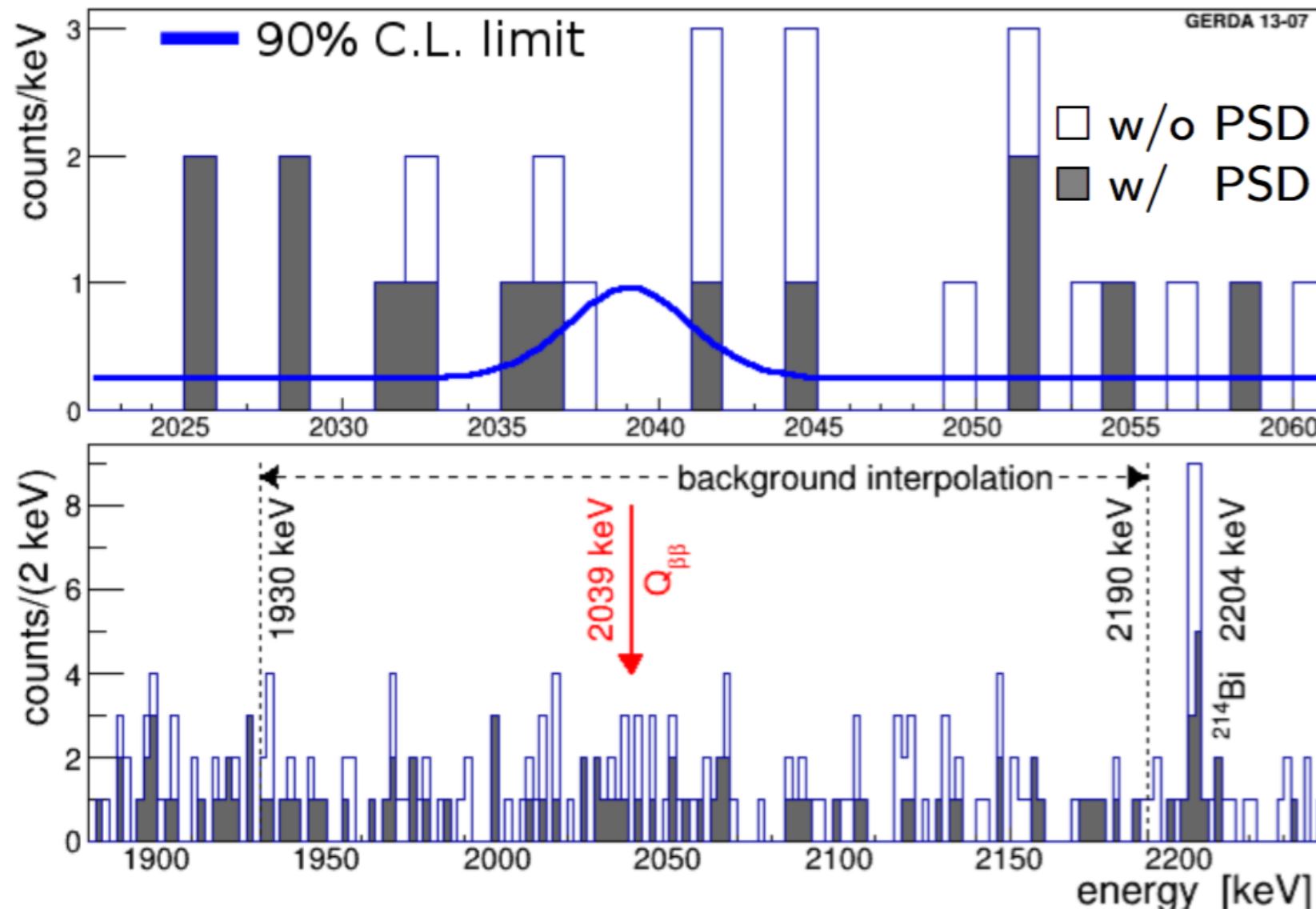


GERDA

- High purity semiconductor
- Enrichment: $\sim 86\%$ of ^{76}Ge
- $Q_{\beta\beta}$ value 2039 keV
- Energy resolution @ $Q_{\beta\beta} \sim 0.2\%$ (FWHM)
- High detection efficiency
- Modular
- Point-like energy deposition (PSD)



GERDA Phase I results



- Exposure 21.6 kg · yr
- Blind analysis
- Detection efficiency:
 - 62% for Coax
 - 66% for BEGe
- Sensitivity $2.4 \cdot 10^{25}$ yr
- No counts in $2039 \text{ keV} \pm 1\sigma$

$$T_{1/2} > 2.1 \times 10^{25} \text{ yr}$$

$$\langle m_{\beta\beta} \rangle < 0.2\text{--}0.4 \text{ eV (90\% C.L.)}$$

GERDA Phase II

Goals:

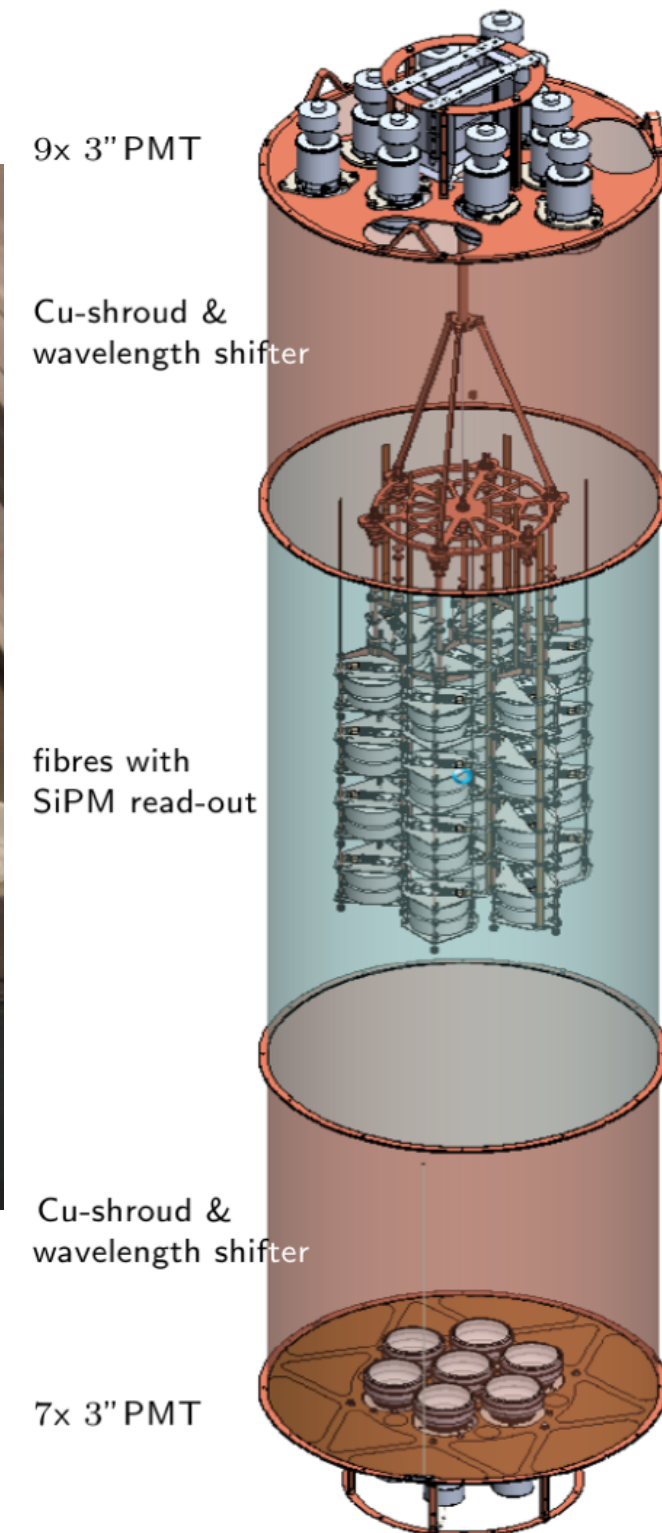
- Sensitivity above $1 \cdot 10^{26}$ yr
- Exposure ~ 100 kg \cdot yr
- Background index 10^{-3} c/(keV \cdot kg \cdot yr)

What's new?

- 30 custom BEGe detectors (~ 20 kg)
- energy resolution ~ 3 keV FWHM @ 2MeV
- pulse shape discrimination of bulk SSE against surface events (α & β) and MSE

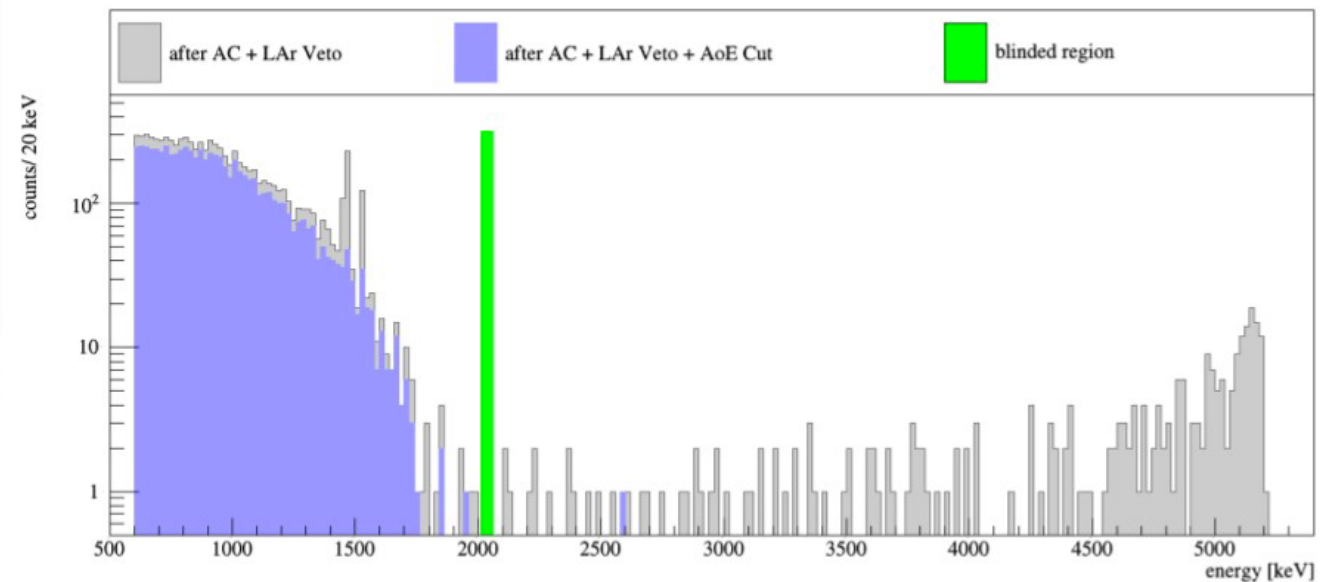
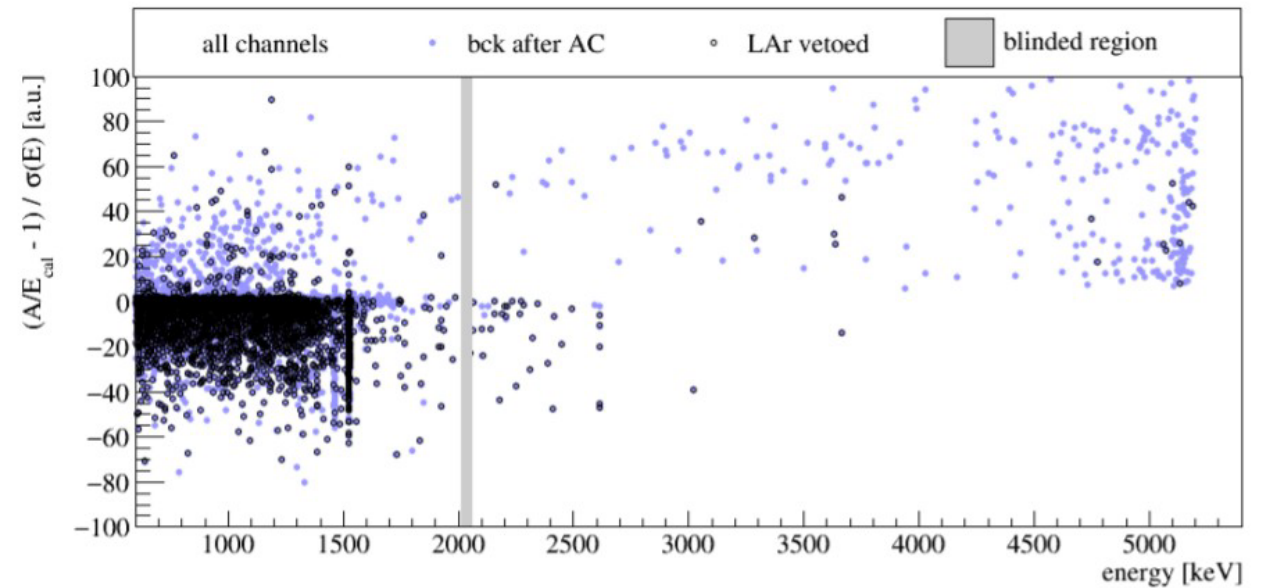
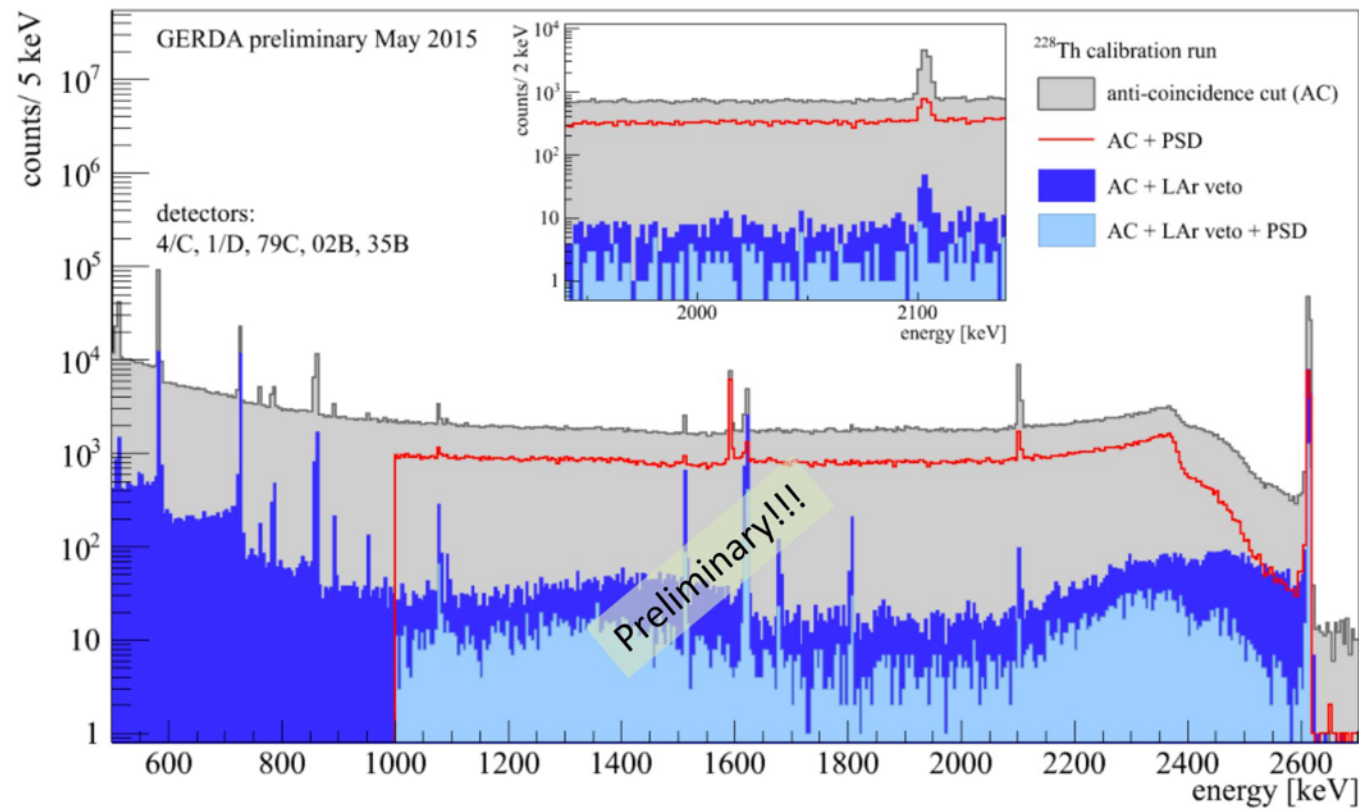
Background veto

- LAr scintillating read-out
- better anti-coincidence



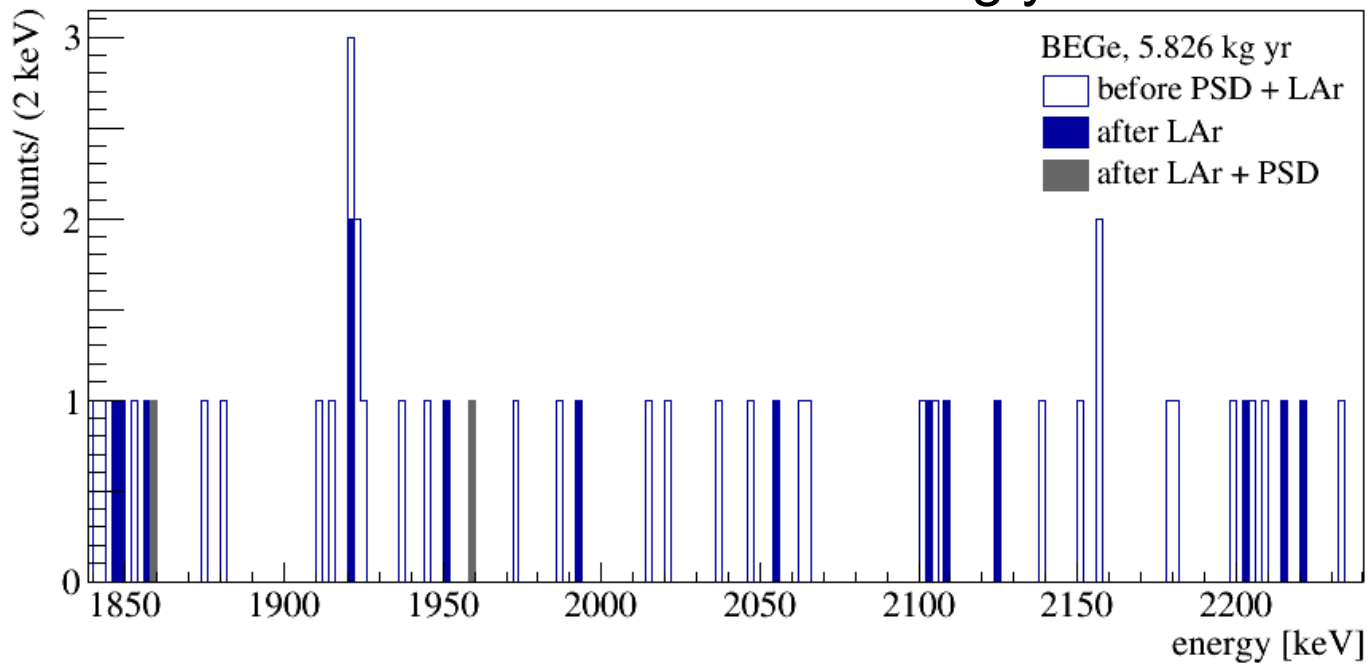
Background discrimination

- Anti-coincidence
- LAr veto
- PSD cuts

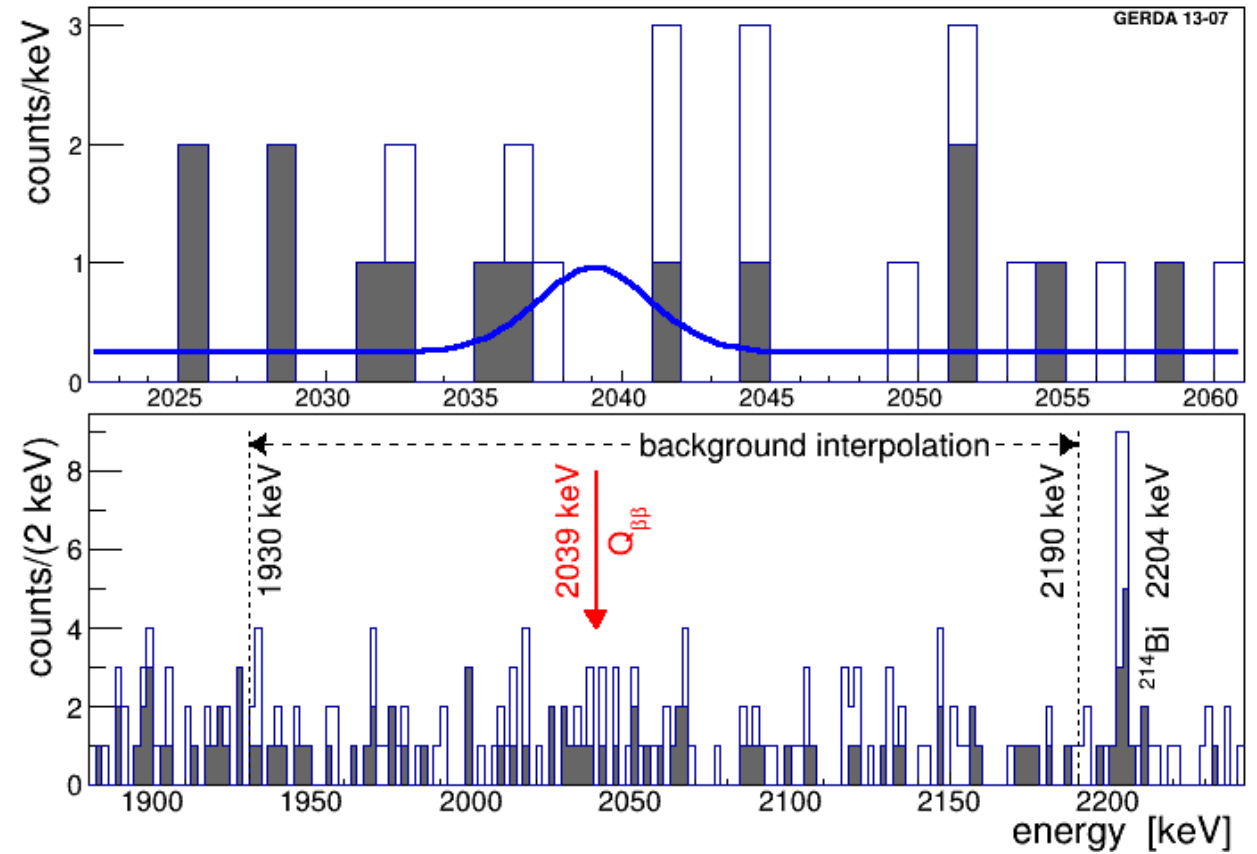


GERDA Phase II preliminary results

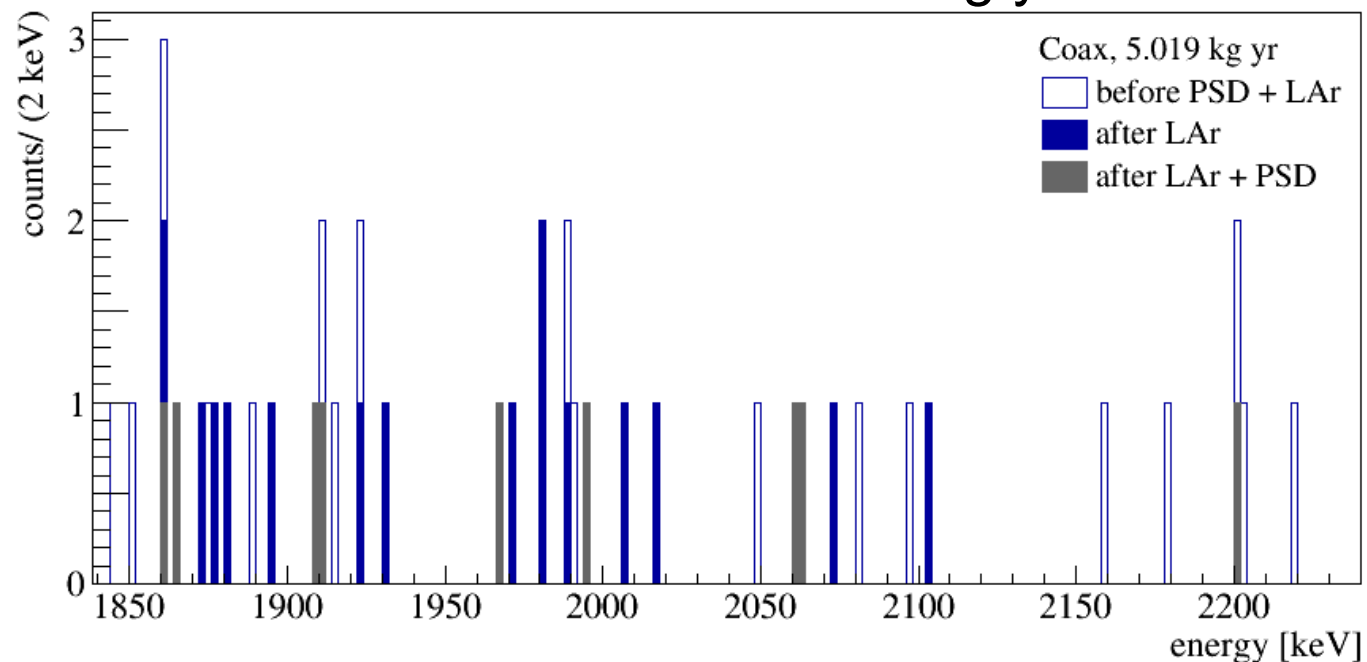
Phase II BEGe: 5.8 kg yr



Phase I: 21.7 kg yr



Phase II Coax: 5.0 kg yr



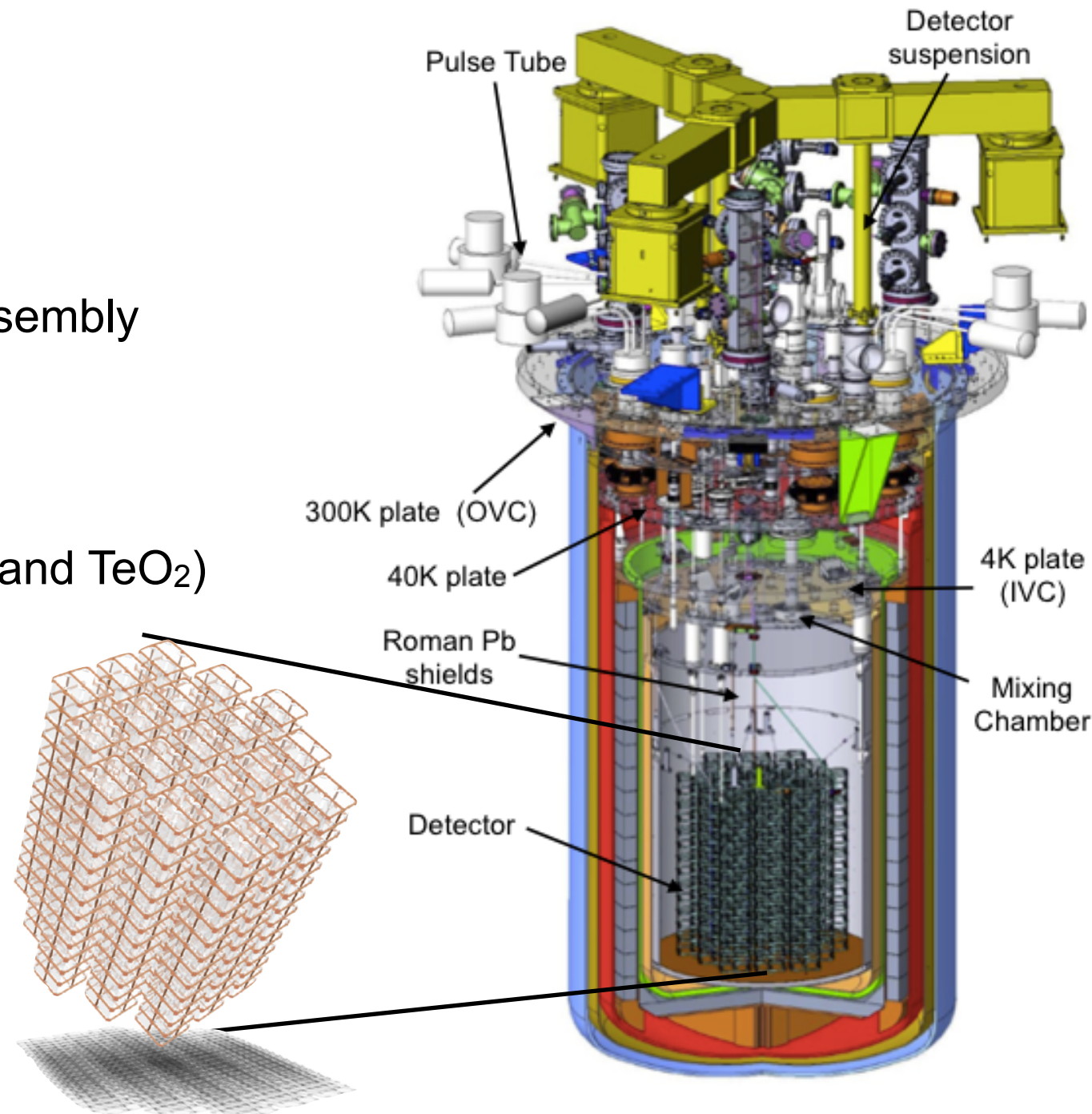
Background (BEGe) $0.7^{+1.2}_{-0.5} \times 10^{-3} \text{ c/keV/kg/yr}$
 $T_{1/2} > 5.3 \cdot 10^{25} \text{ yr}$ (90% CL, frequentist)
 $T_{1/2} > 3.5 \cdot 10^{25} \text{ yr}$ (90% credible, Bayesian)

This result suggests future Ge experiments with 200 kg and beyond

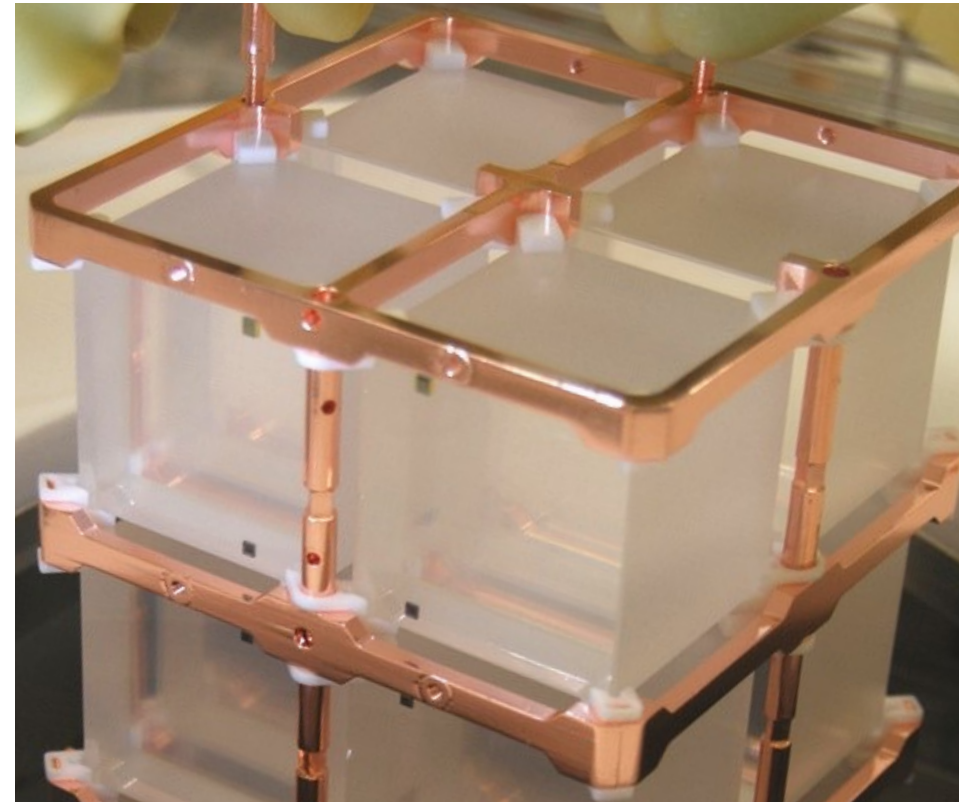
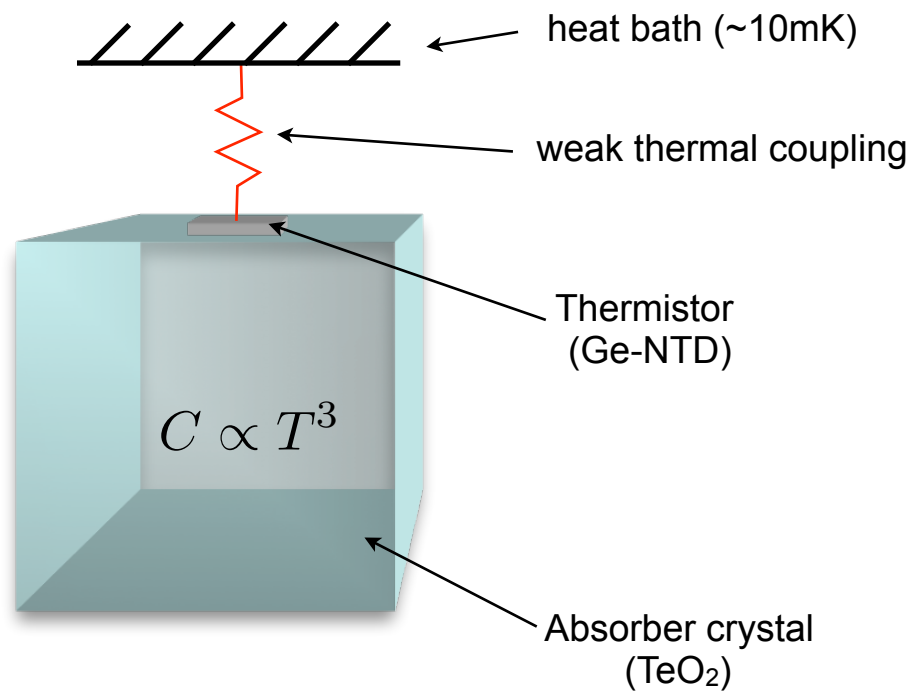
The CUORE challenge

Operate a huge thermal detector array in a extremely low radioactivity and low vibrations environment

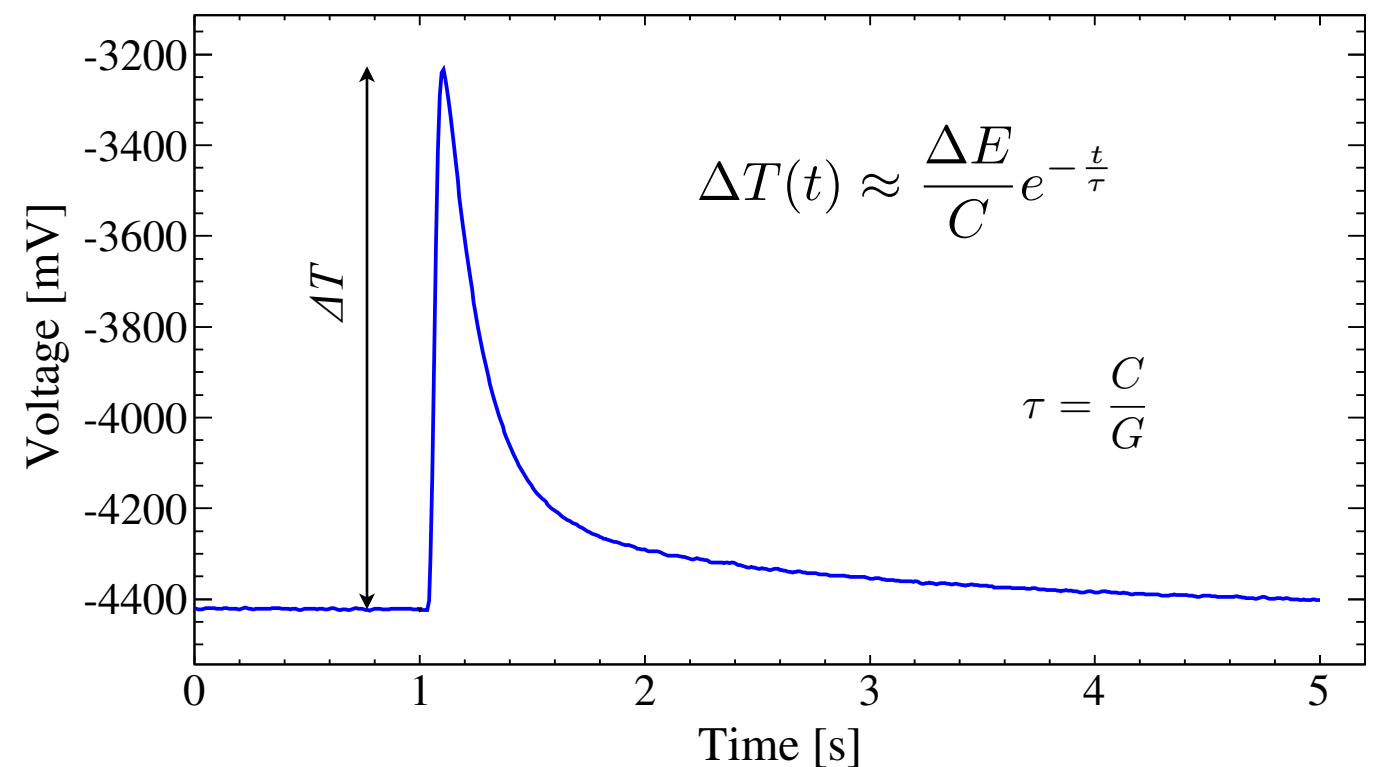
- Closely packed array of 988 TeO_2 crystals (19 towers of 52 crystals $5 \times 5 \times 5 \text{ cm}^3$, 0.75 kg each)
- Mass of TeO_2 : 741 kg ($\sim 206 \text{ kg}$ of ^{130}Te)
- Energy resolution: 5 keV @ 2615 keV (FWHM)
- Stringent radiopurity controls on materials and assembly
- Operating temperature: $\sim 10 \text{ mK}$
- Mass to be cooled down: $\sim 15 \text{ tons}$ (lead, copper and TeO_2)
- Background aim: $10^{-2} \text{ c/keV/kg/year}$
- $T_{1/2}$ sensitivity (90% C.L.): $0.95 \times 10^{26} \text{ yr}$



Thermal detectors



- wide choice of detector materials
 - low heat capacity @ T_{work}
- excellent energy resolution ($\sim 1 \text{‰}$ FWHM)
 - huge number of energy carriers (phonons)
- equal detector response for different particles
 - true calorimeters
- slowness
 - in rare event search doesn't matter



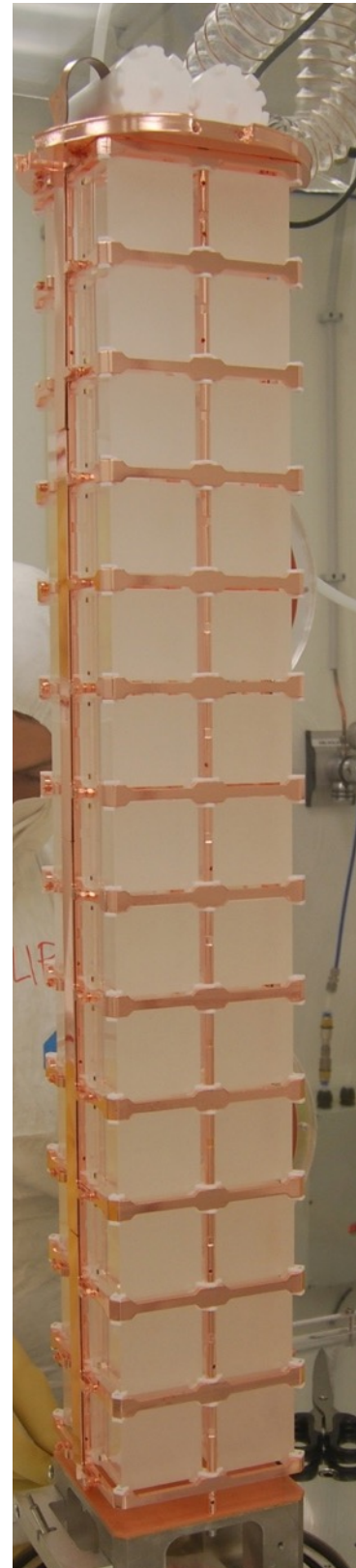
CUORE-0

CUORE-0 is the **first tower** produced out of the CUORE assembly line.

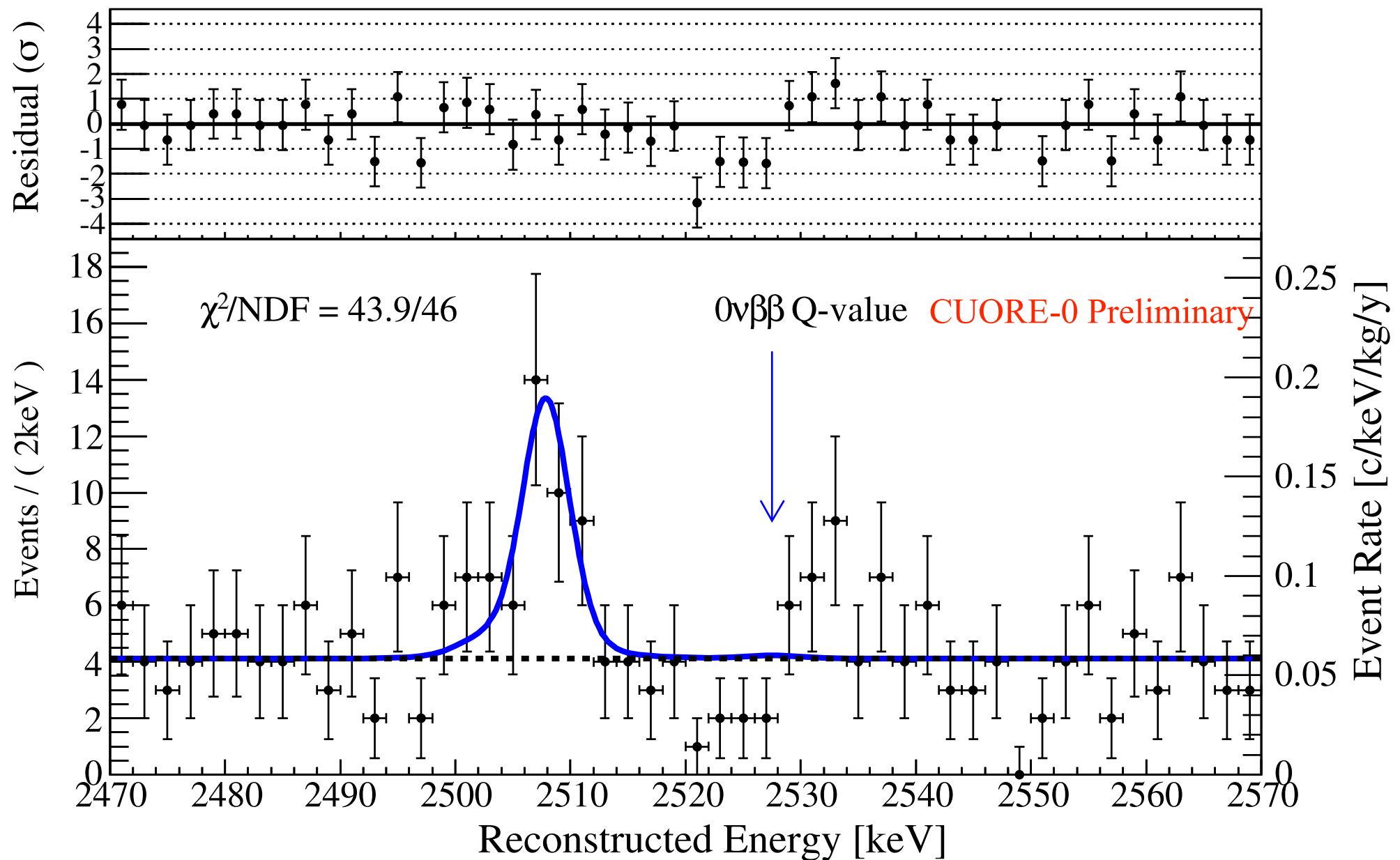
- 52 TeO_2 5x5x5 cm^3 crystals (~ 750 g each)
- 13 floors of 4 crystals each
- total detector mass: 39 kg TeO_2 (10.9 kg of ^{130}Te)

CUORE-0 has been taking data since March 2013 in the 25 year old Cuoricino cryostat.

- **Proof of concept** of CUORE detector in all stages
- Test and debug of the CUORE **tower assembly line**
- Test of the CUORE **DAQ and analysis framework**
- Check of the radioactive **background reduction**
- Extend the physics reach beyond Cuoricino while CUORE is being assembled
- Sensitive $0\nu\text{-}\beta\beta$ experiment



CUORE-0 results

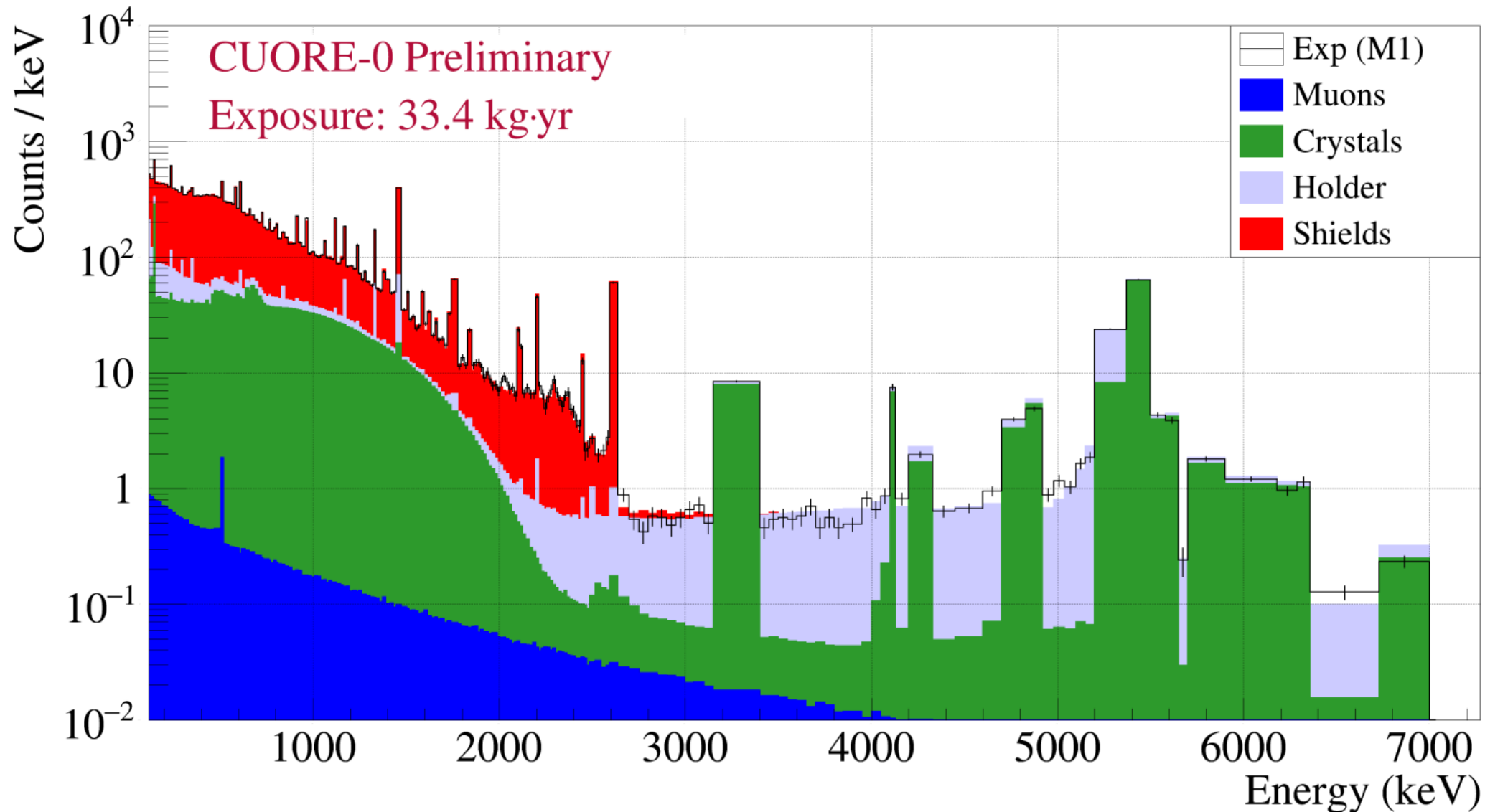


Background index: 0.058 ± 0.004 (stat.) ± 0.002 (syst.) c keV⁻¹ kg⁻¹ yr⁻¹

0 ν - $\beta\beta$ ¹³⁰Te Bayesian 90% C.L. limit: $T_{1/2} > 2.7 \times 10^{24}$ yr

$\langle m_{\beta\beta} \rangle < (270-650)$ meV

CUORE-0 background reconstruction



$0\nu\beta\beta$ ROI background:

- Shields: ~72%
- Holder: ~21%
- Crystals: ~5%
- Muons: ~2%

- In CUORE the contribution from the cryostat shields will be strongly reduced
- Main background will come from degraded alpha particles

CUORE Towers Assembly

- Assembly of all the 19 CUORE towers completed in 2014



Assembly line improved
after CUORE-0

CUORE-0

51/52 NTD connected

51/52 heaters connected

CUORE

988/988 NTD connected

988/988 heaters connected

- Also a mockup tower for the Detector installation phase and a minitower to be used during the cryostat commissioning runs were produced

Cryogenic system commissioning: phase 1

Phased commissioning adding complexity at each step

- Phase I: individual systems test

- Outer/Inner vacuum chamber

- Cryostat

- Final temperatures:

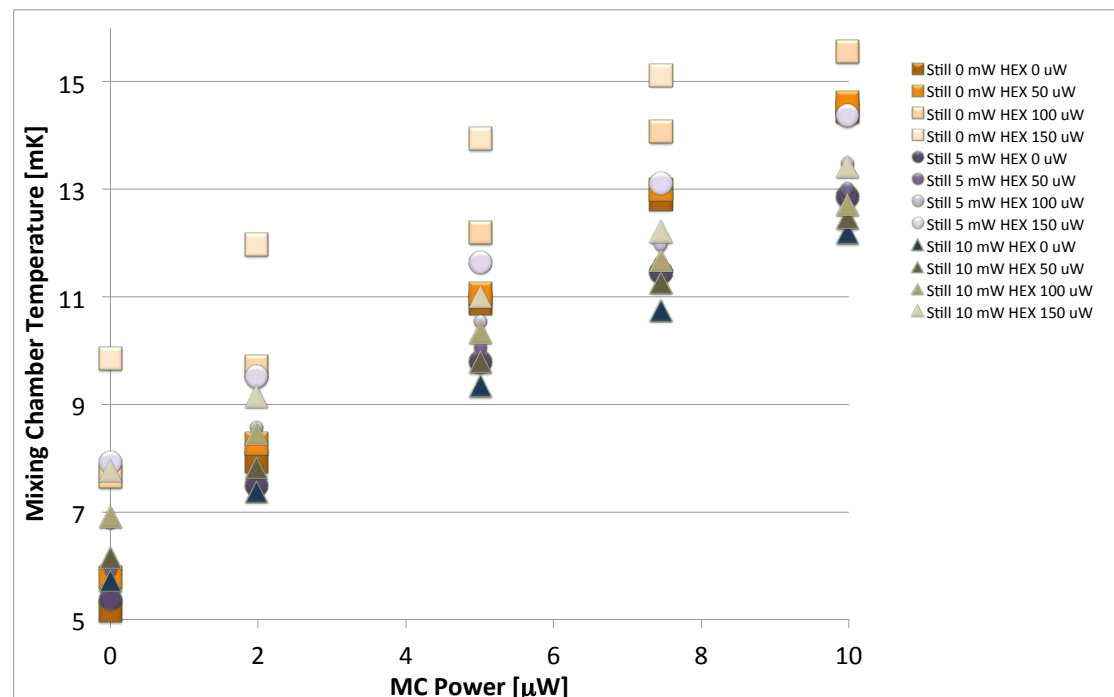
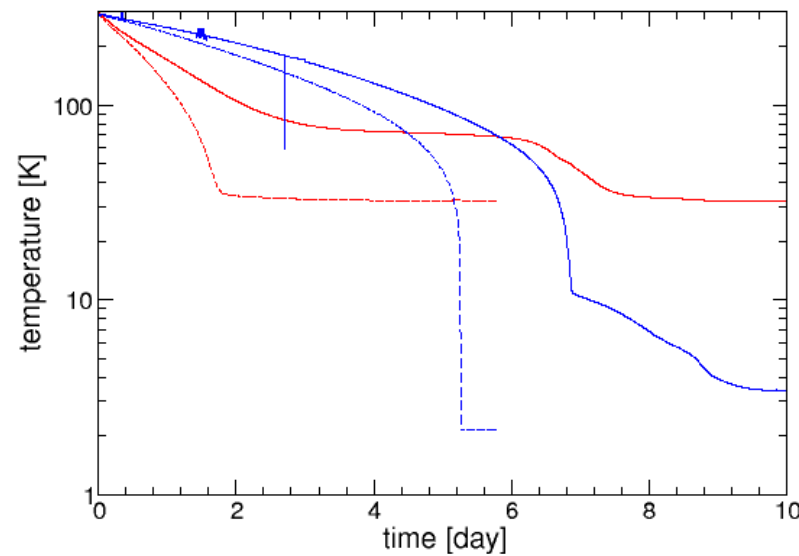
- 32 K at the 40K stage

- 3.3 K at the 4K stage

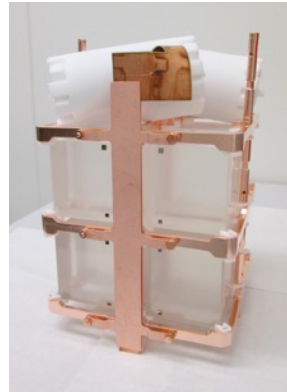
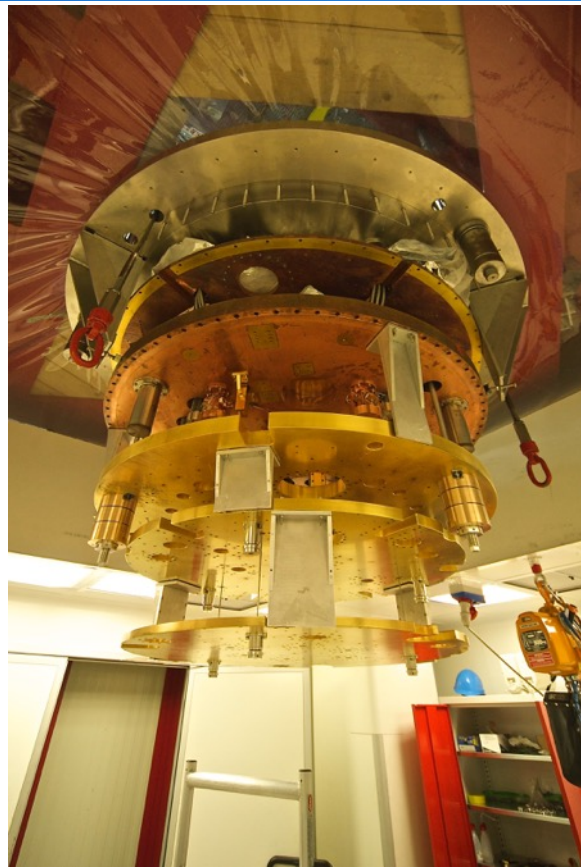
- Dilution Unit

- Lowest temperature: 4.95 mK

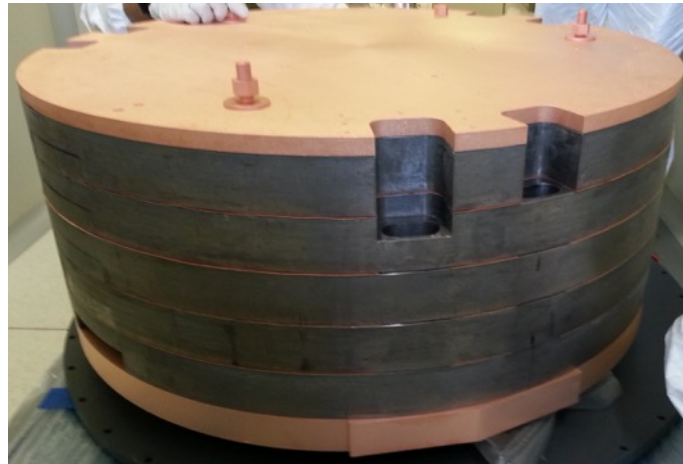
- Phase II: system integration



Cryogenic system commissioning: phase 2



Insertion of few
TeO₂ detectors



Cryostat
+
Dilution Unit

Wiring

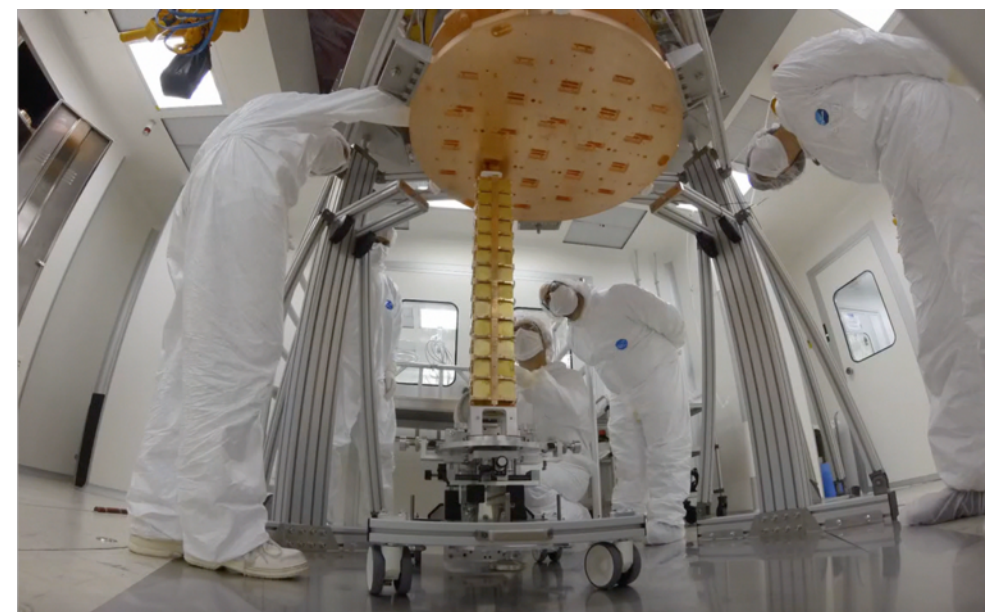
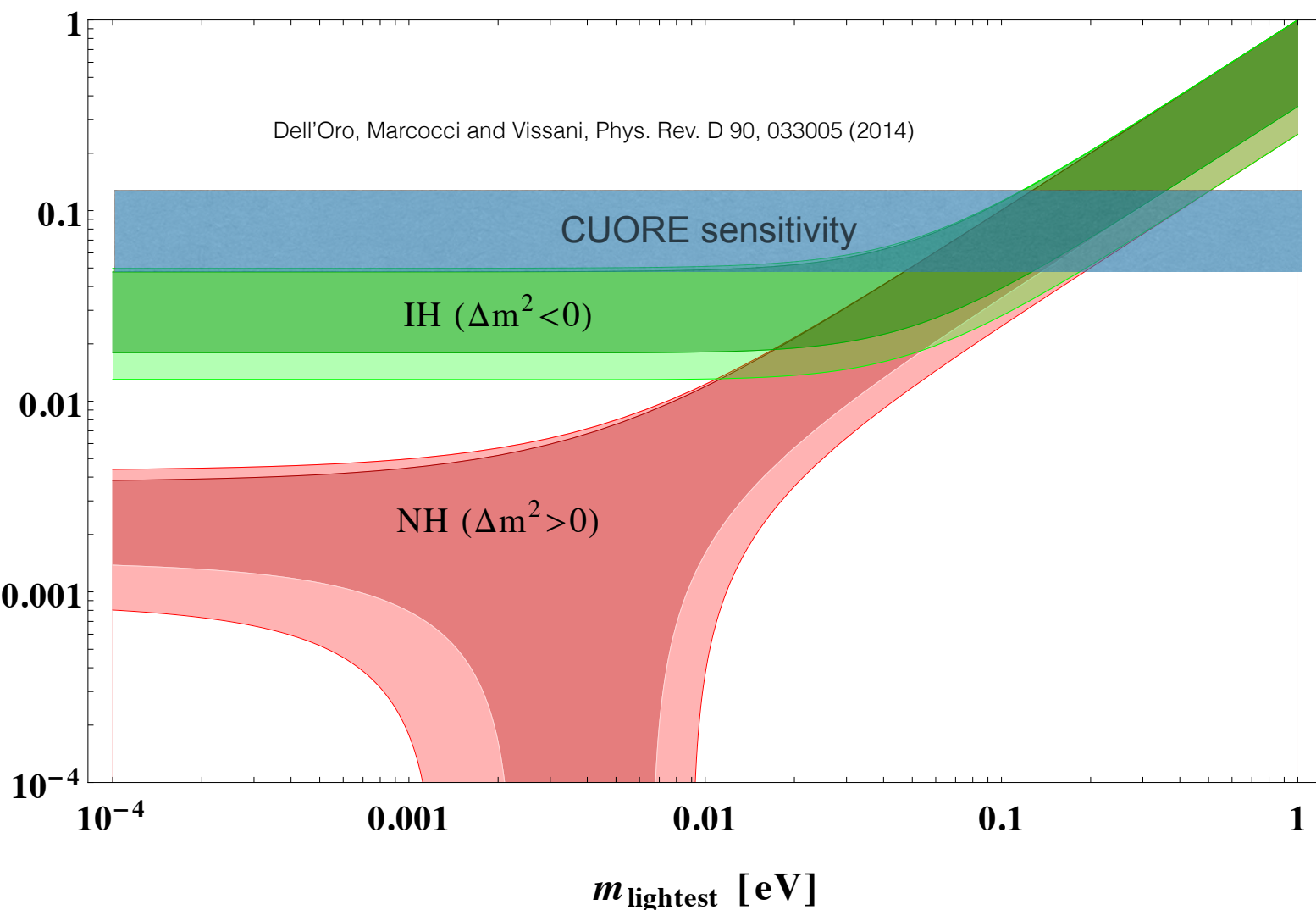
Top Pb shield
Detector Calibration System
Towers support plate
Fast Cooling System

side roman Pb shield



Present Status

- Cryogenic commissioning is completed
- Installation of all the 19 CUORE towers in the cryostat will start in 3 weeks
- CUORE cooldown foreseen in October



- $T_{1/2}$ sensitivity 0.95×10^{26} years @ 90% C.L. (5 years)
- $m_{\beta\beta}$ sensitivity 50-130 meV @ 90% C.L. (5 years)

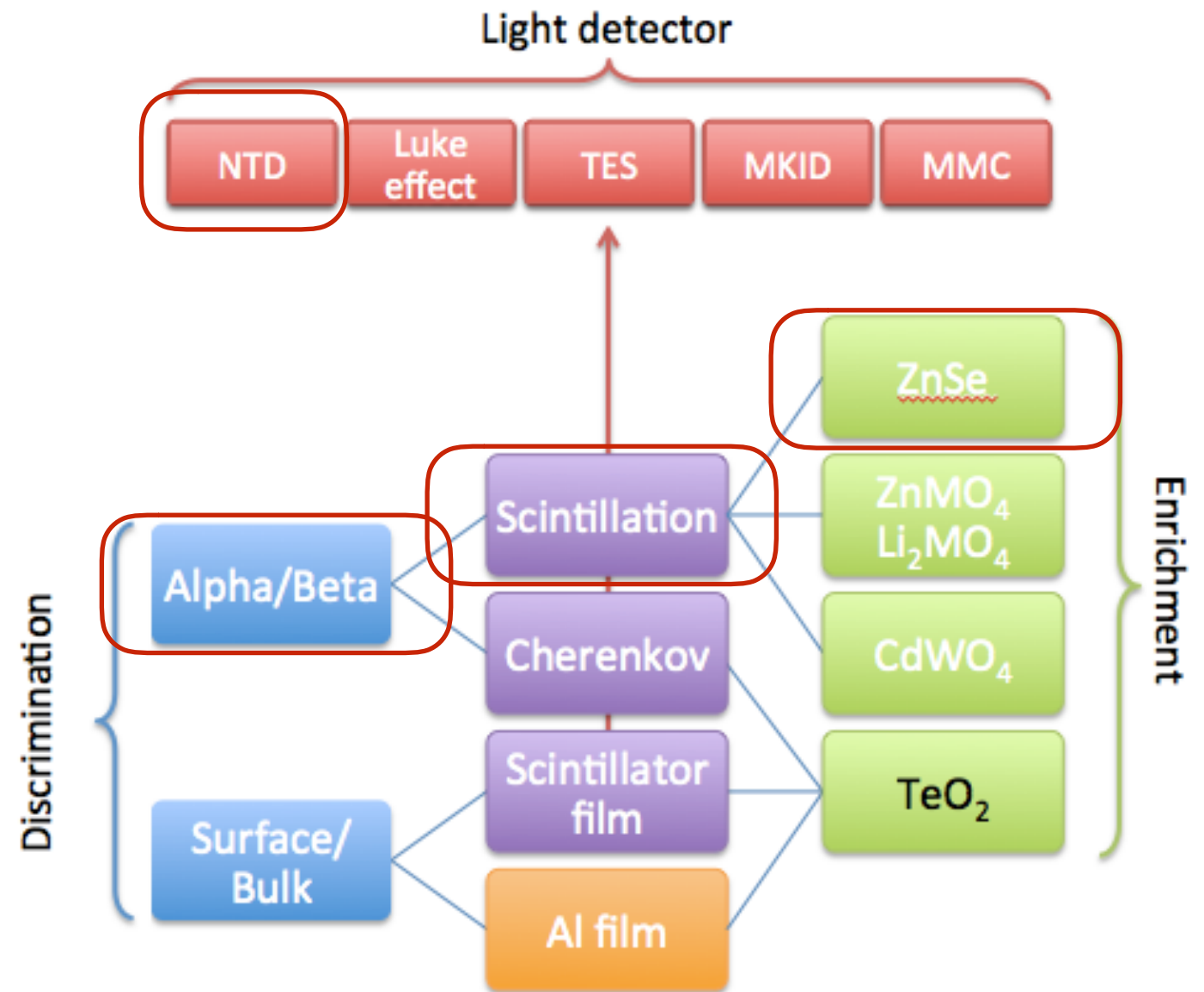
CUPID

- CUPID (CUORE Upgrade with Particle IDentification) aims at the realization of a ton-scale bolometric $0\nu\text{-}\beta\beta$ experiment, based on the experience learned in CUORE.
- CUPID plans to use the CUORE infrastructure @ LNGS, once CUORE completes operation
- Several R&D in progress

Required upgrades

- background discrimination
- isotopic enrichment
- stricter material selection
- possibly new shielding concepts

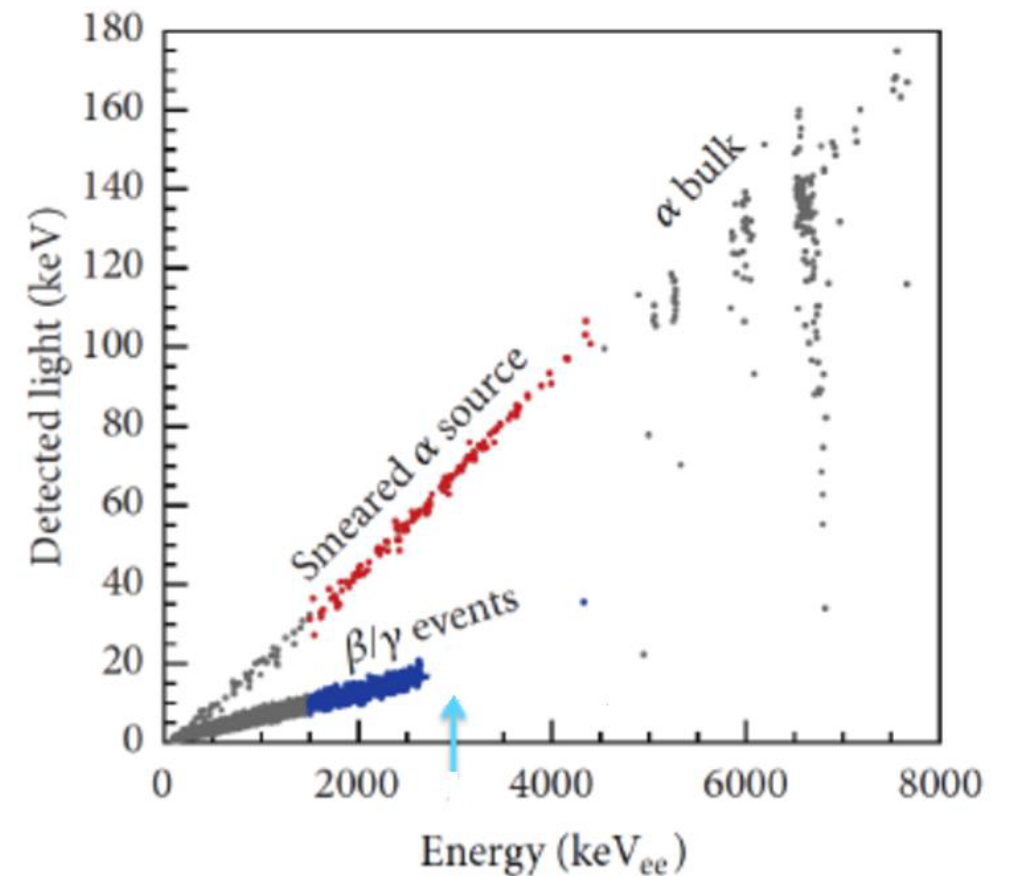
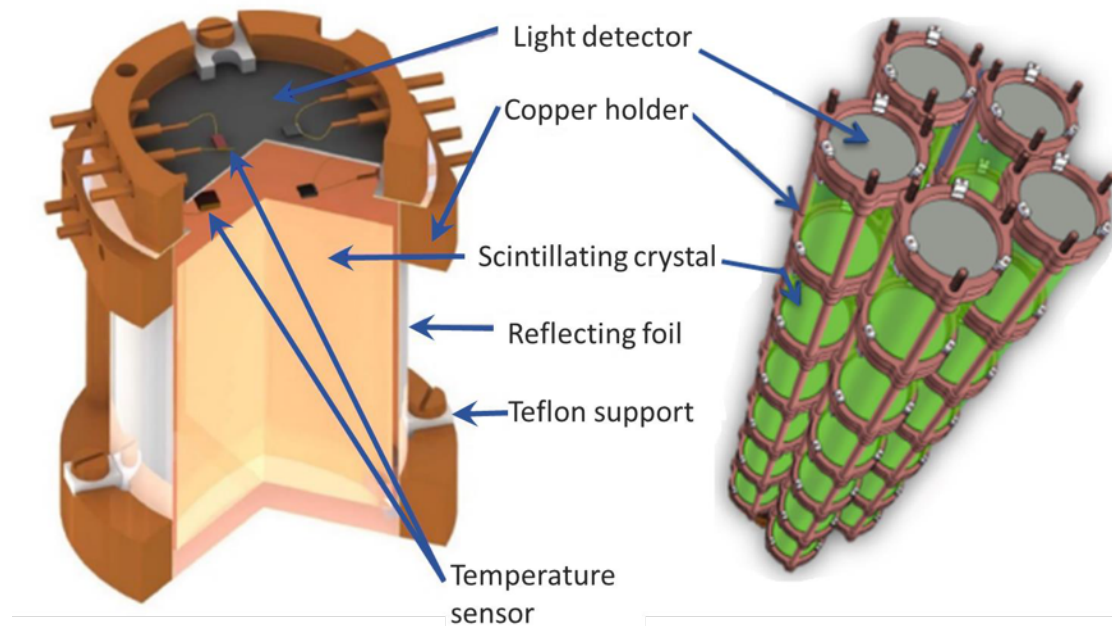
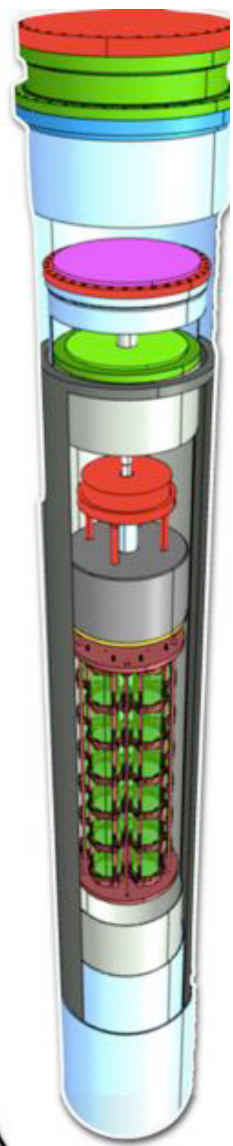
Phased program to test different techniques in small scale demonstrators



R&D towards CUPID: [arXiv:1504.03612](https://arxiv.org/abs/1504.03612)

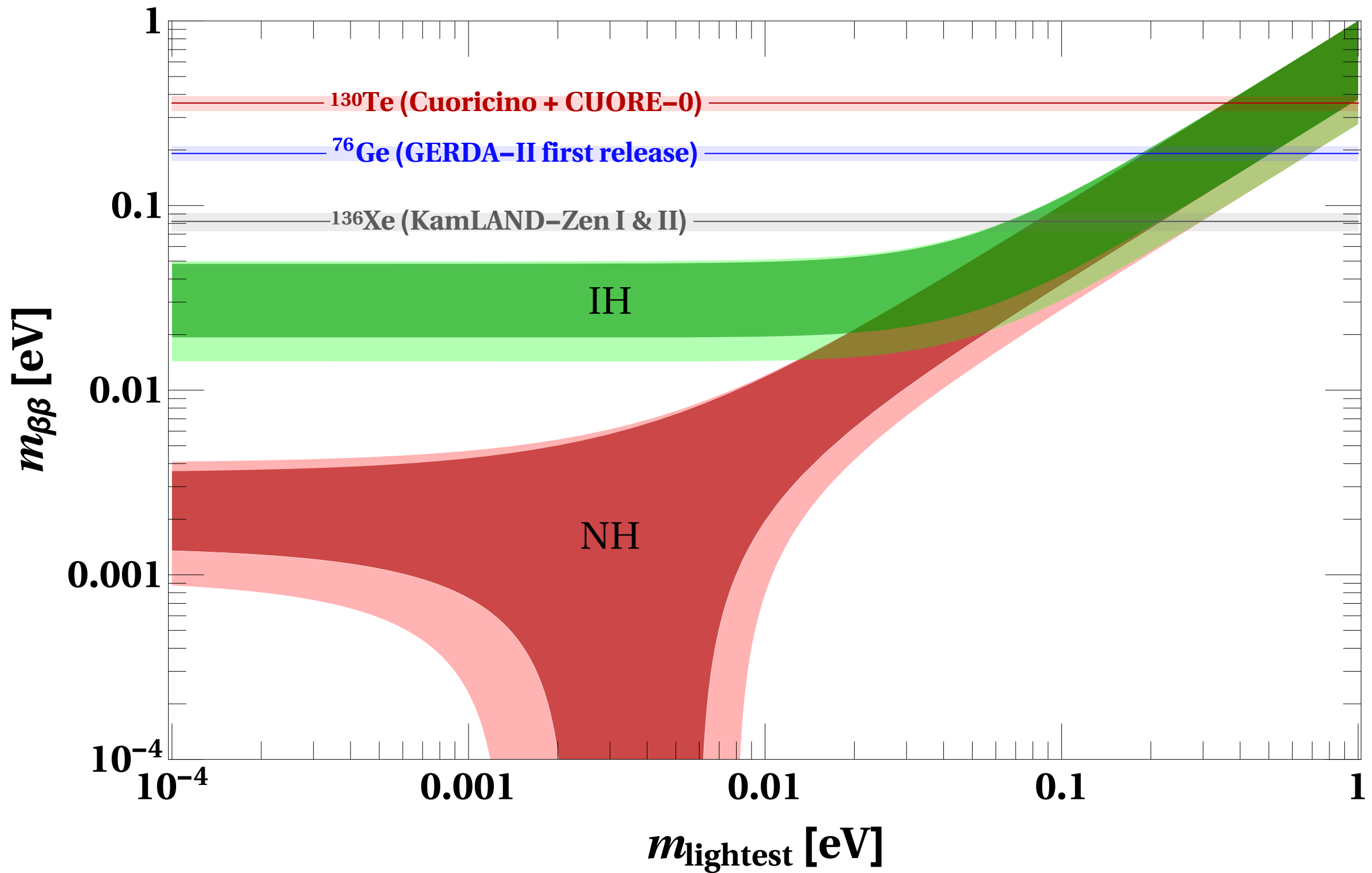
CUPID : [arXiv:1504.03599](https://arxiv.org/abs/1504.03599)

First demonstrator: CUPID-0/LUCIFER

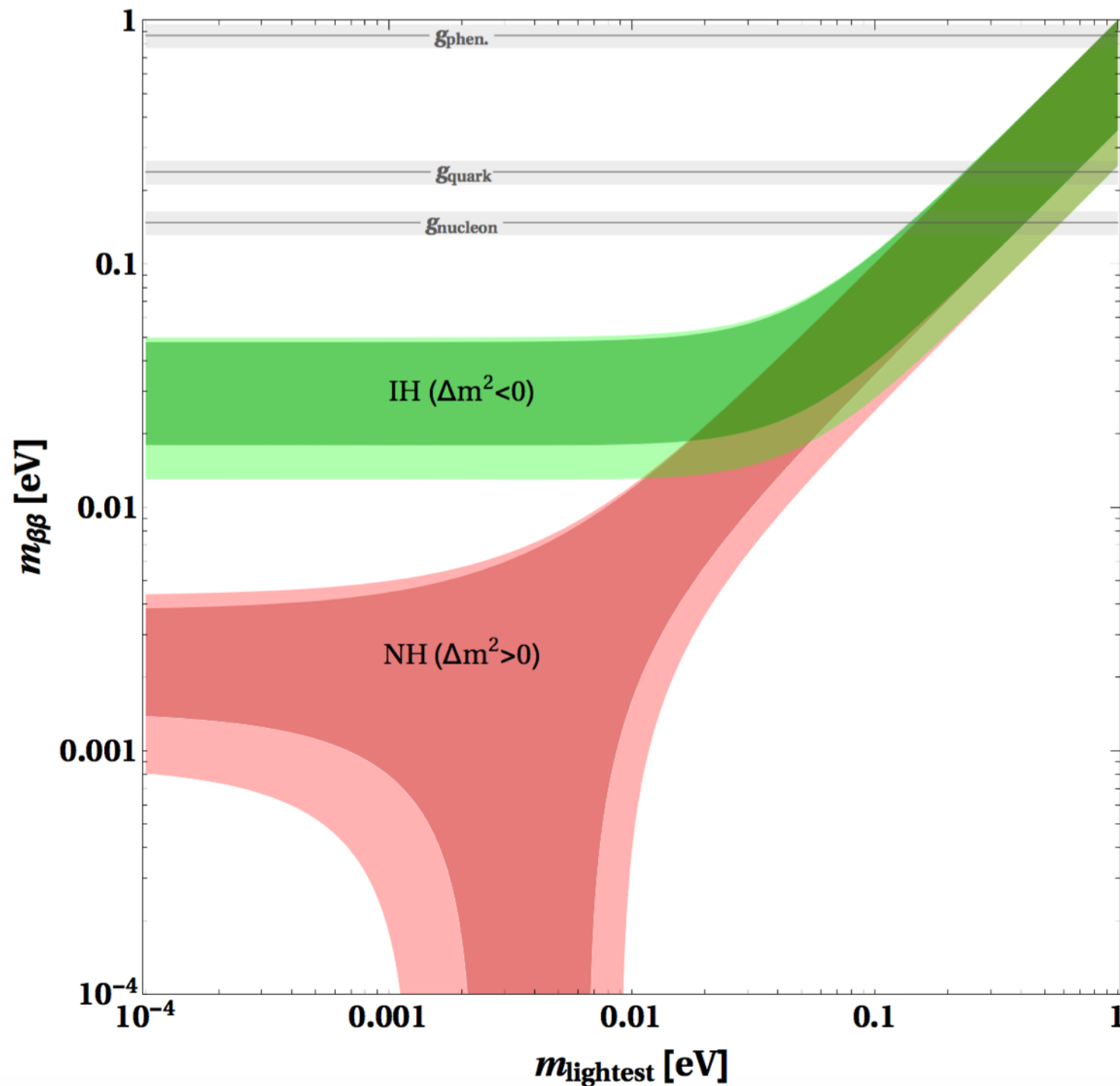


- 30 Zn^{82}Se crystals ~ 440 g each @ 95% enrichment operated as scintillating bolometers
- Bolometers arranged in 5 towers and faced to Ge light detectors
- Total mass: 13.2 kg (7 kg ^{82}Se)
- Expected bkg @ ROI 10^{-3} c/keV/kg/y
- Expected energy resolution @ ROI: 10 keV FWHM
- Data taking: Autumn 2016

Present limits on $m_{\beta\beta}$



Importance of the g_A quenching



$$t_{0\nu}^{1/2} \propto \mathcal{M} = g_A^{-4} \mathcal{M}_{0\nu}^{-2}$$

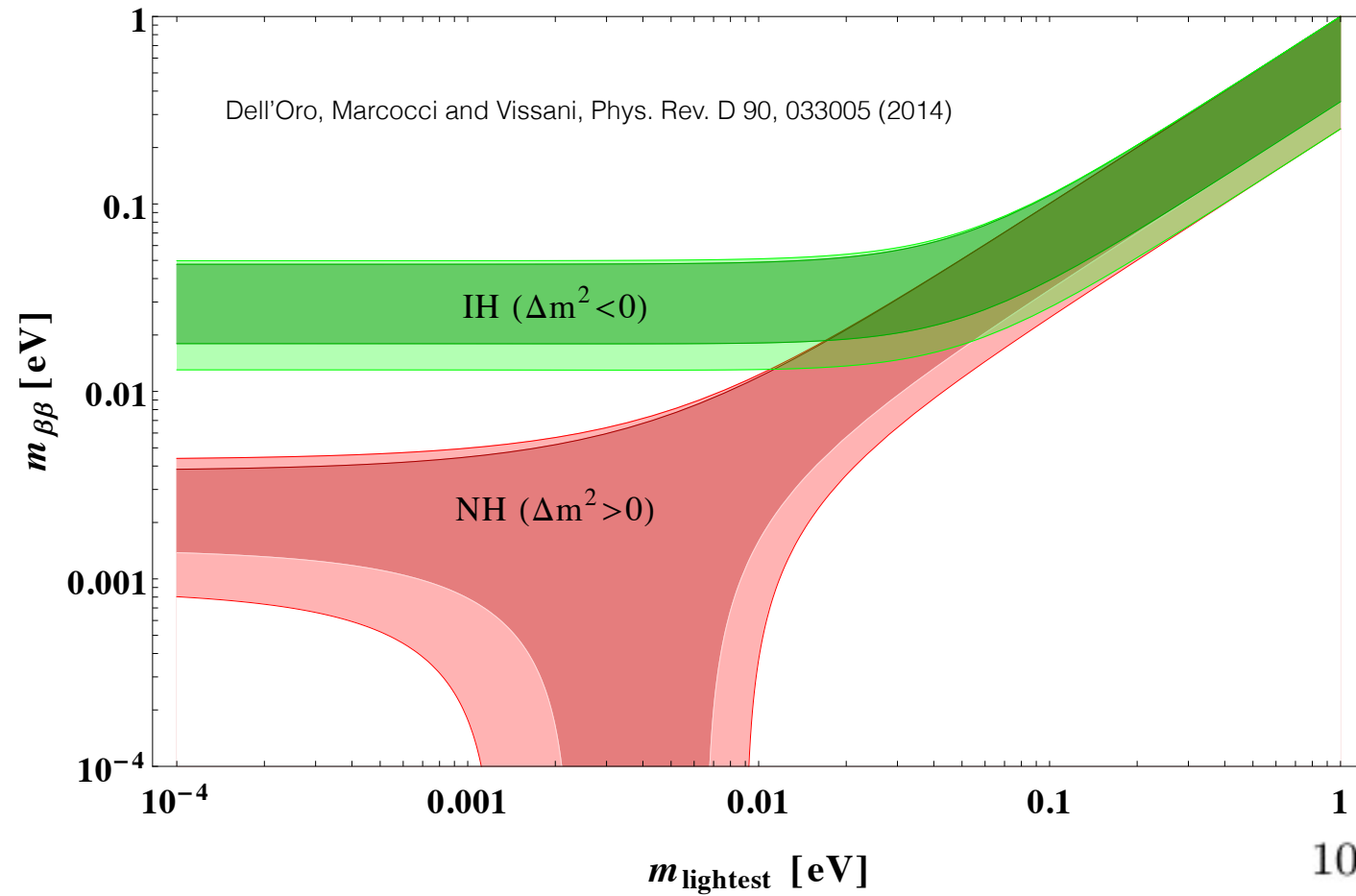
$$g_A = \begin{cases} g_{A, \text{nucleon}} & = 1.269 \\ g_{A, \text{quark}} & = 1 \\ g_{A, \text{phen.}} & = 1.269 \cdot A^{-0.18} \end{cases}$$

g_A	$m_{\beta\beta}^{\text{min}} [\text{eV}]$
$g_{A, \text{nucleon}}$	0.15 ± 0.03
$g_{A, \text{quark}}$	0.24 ± 0.05
$g_{A, \text{phen.}}$	0.87 ± 0.17

NMEs (IBM-2): J. Barea *et al.*, *Phys. Rev. C* 91, 034304 (2015)
 PSFs: J. Kotila & F. Iachello, *Phys. Rev. C* 85, 034316 (2012)
 ^{136}Xe : A. Gando *et al.*, *Phys. Rev. Lett.* 110, 062502, (2013)

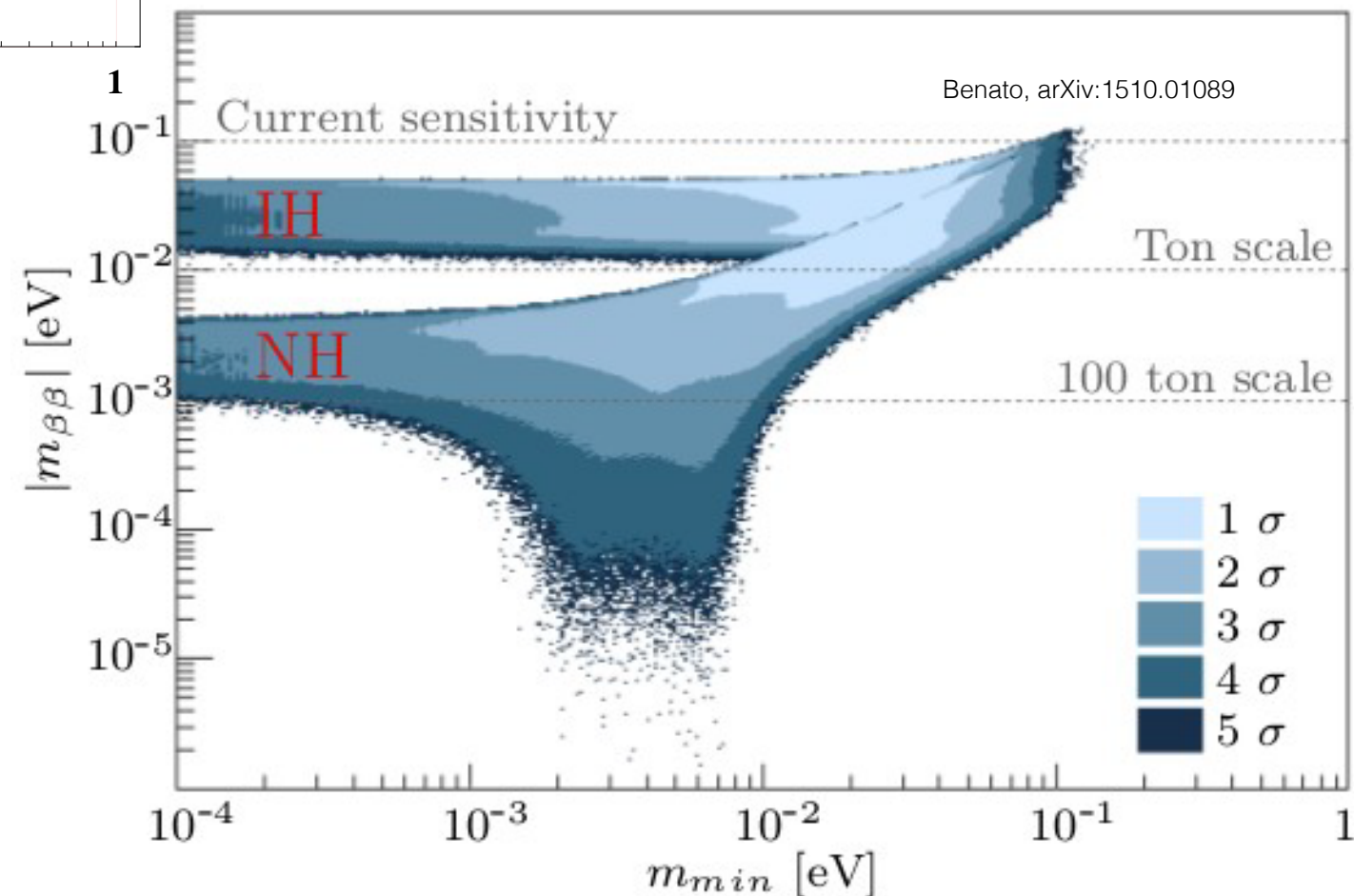
Dell'Oro, Marcocci and Vissani, *Phys. Rev. D* 90, 033005 (2014)

Normal hierarchy is unreachable?



- Data from oscillations experiments
- Main uncertainty given by the Majorana phases

- Toy Montecarlo
- Bound on the sum of neutrino masses from cosmology
- Flat probability distribution for Majorana phases



Conclusions

- Very strong worldwide competition, many results in the last months (and others to come)
- $0\nu\text{-}\beta\beta$ discovery could be within reach
 - If nothing is found we have to go to larger target mass
 - more money (about 50-100 M€)
 - isotopic enrichment will be the dominant cost
 - IMHO worthwhile
- improvement in NME calculation and especially better knowledge on g_A is advisable
- LNGS has a relevant role in $0\nu\text{-}\beta\beta$