



Why could strangeon matter be “superluminal”? —sound speed in compact star

Jiguang Lu, Enping Zhou, Xiaoyu Lai &
Renxin Xu

Peking University

12, 05, 2017, Varadero, Cuba

maximum mass of pulsar

- The maximum mass of a neutron star can hardly surpass $2M_{\odot}$ due to the “hyperon puzzle”. (Schaffner-Bielich 2008)
- PSR J0715+1807, $1.97M_{\odot}$ (Demorest et al. 2010)
- PSR J0348+0432, $2.01M_{\odot}$ (Antoniadis et al. 2013)
- Stiffer matter state is needed.

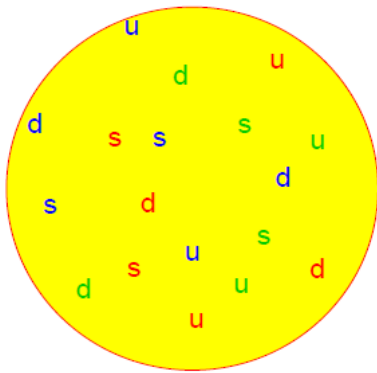
maximum mass of pulsar

- The maximum mass of a neutron star can hardly surpass $2M_{\odot}$ due to the “hyperon puzzle”. (Schaffner-Bielich 2008)
- PSR J0715+1807, $1.97M_{\odot}$ (Demorest et al. 2010)
- PSR J0348+0432, $2.01M_{\odot}$ (Antoniadis et al. 2013)
- Stiffer matter state is needed.

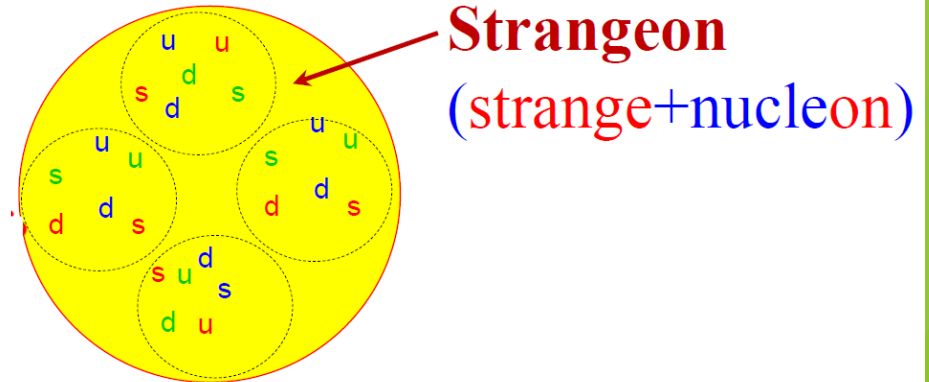
Strangeon matter!

What is a **strangeon**?

- ▶ speculated in generalized Witten's conjecture
- ▶ 3 flavor symmetry (u, d, s)
- ▶ Strangeons in bulk constitutes the true ground state of the strong-interaction at supra nuclear density.



Strange quark matter



Strangeon matter

What is the property of strangeon matter?

- ▶ A strangeon is much more massive than a nucleon. Thus the strangeon can be regarded as a traditional localized point particle rather than quantum wave.
- ▶ Strangeon matter should be much stiffer than neutron matter, which signifies a larger bulk modulus.
- ▶ $\sqrt{\frac{\partial P}{\partial \rho}} > c$ (superluminal sound speed??)

sound speed & causality

► There is no priori reason to believe the variational principle is physical, thus any result based on the Lagrangian may be invalid (Ellis, Maartens & MacCallum, 2007).

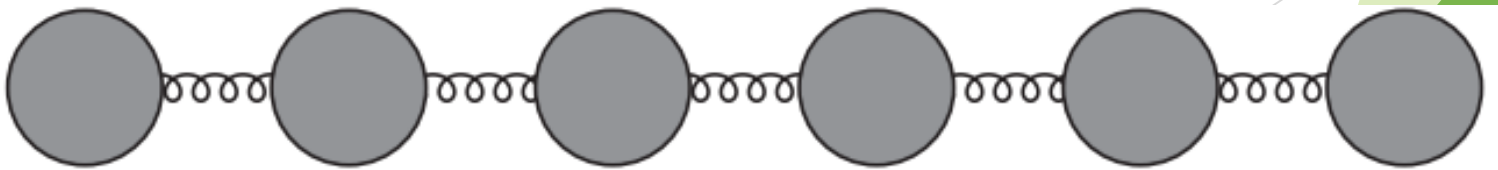
► Lorentz invariance per se does not prohibit macroscopic theories with superluminal sound, and superluminal propagation is not necessarily noncausal (Bruneton, 2007).

► Sound speed $c_s = \sqrt{\frac{\partial P}{\partial \rho}}$ Caporaso & Brecher (1979) **??**

pointed out that it is possible to construct lattice models with $P > \rho c^2$ and $dP > c^2 d\rho$, but sound speed remains subluminal.

Signal in strangeon matter

- ▶ The particle (strangeon) is so massive that it has small wave-packet, thus, the strangeon can be regarded as a traditional localized point particle rather than quantum wave.
- ▶ **two-body short-range repulsive conservative** interaction between particles
- ▶ For simplification, the oscillation propagation in a 1-D discrete chain is considered.



frequency domain result of strangeon matter

- ▶ Assume a sine wave propagate in the chain, and the wave form is stable.

frequency domain result of strangeon matter

- ▶ Assume a sine wave propagate in the chain, and the wave form is stable.

- ▶ Sound speed $c_s = \frac{1}{\sqrt{\frac{1}{2} \left[\frac{1}{\frac{\partial P}{\partial \rho}} + \sqrt{\frac{1}{\left(\frac{\partial P}{\partial \rho}\right)^2 + l^2 \omega^2 c^2}} \right]}}$

- ▶ **It could be superluminal!!**

frequency domain result of strangeon matter

- ▶ Assume a sine wave propagate in the main, and the wave form is stable.

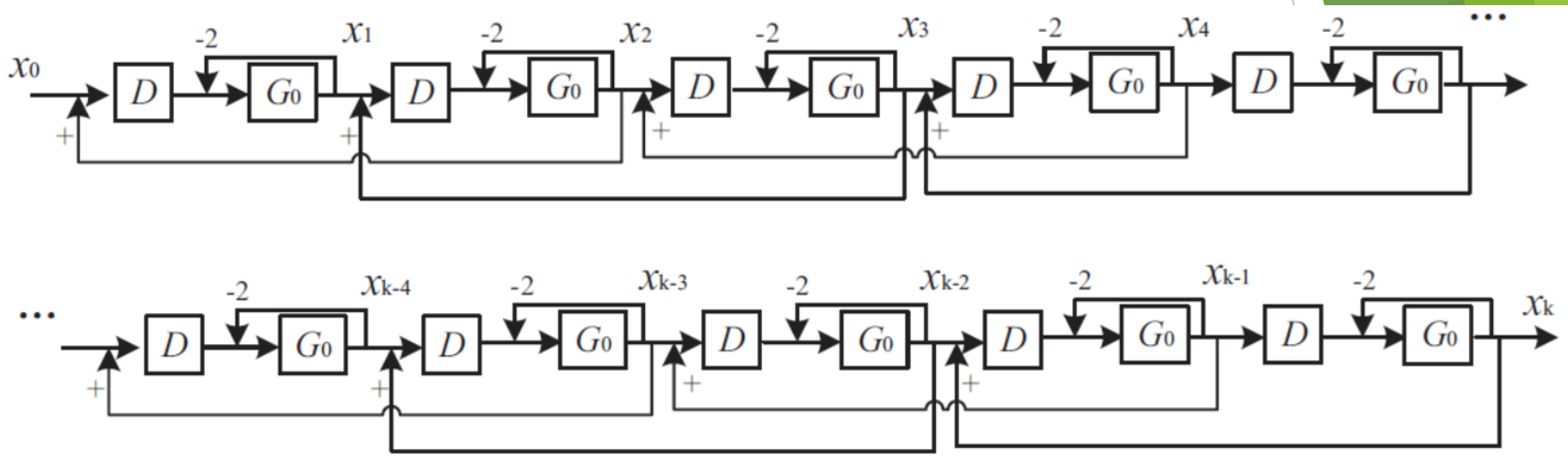
- ▶ Sound speed $c_s = \frac{1}{\sqrt{\frac{1}{2} \left[\frac{1}{\frac{\partial P}{\partial \rho}} + \frac{1}{(\frac{\partial P}{\partial \rho})^2 + \frac{1}{\omega^2 c^2}} \right]}}$

- ▶ It could be superluminal!!

because of the reflected wave

time domain result of strangeon matter

- ▶ time domain response of an impulse input



$$D(s) = \exp\left(-\frac{ls}{c}\right) \quad G_0(s) = -\frac{1}{ms^2} \frac{\partial F}{\partial x} \Big|_{x=l}$$

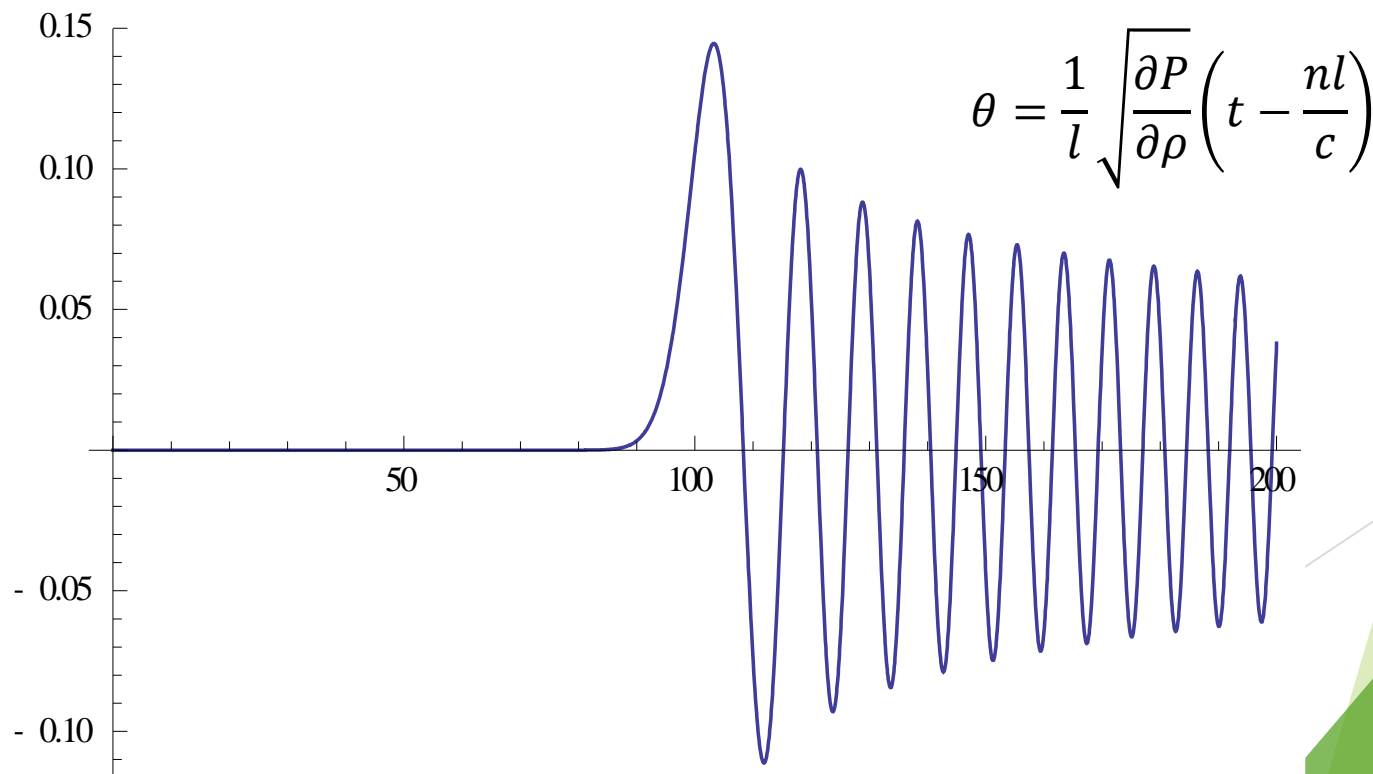
- ▶ With Meson's gain formula

$$G(s) = \frac{2^k G_1^k}{k+1} = \frac{1}{k+1} \left(\frac{\frac{\partial P}{\partial \rho}}{\frac{l^2 s^2}{2} + \frac{\partial P}{\partial \rho}} \right)^k \exp\left(-\frac{kls}{c}\right)$$

time domain result of strangeon matter

- ▶ time domain response of an impulse input

$$x_k(t) = \mathcal{L}^{-1}[\mathcal{L}[x_0(t)]G(s)] = \frac{2^{k-\frac{1}{2}}\sqrt{\pi}A}{(k+1)\Gamma(k)} \sqrt{\frac{2}{l^2} \frac{\partial P}{\partial \rho}} \theta^{k-\frac{1}{2}} J_{k-\frac{1}{2}}(\theta)$$



sound speed in strangeon matter

- ▶ signal propagation time

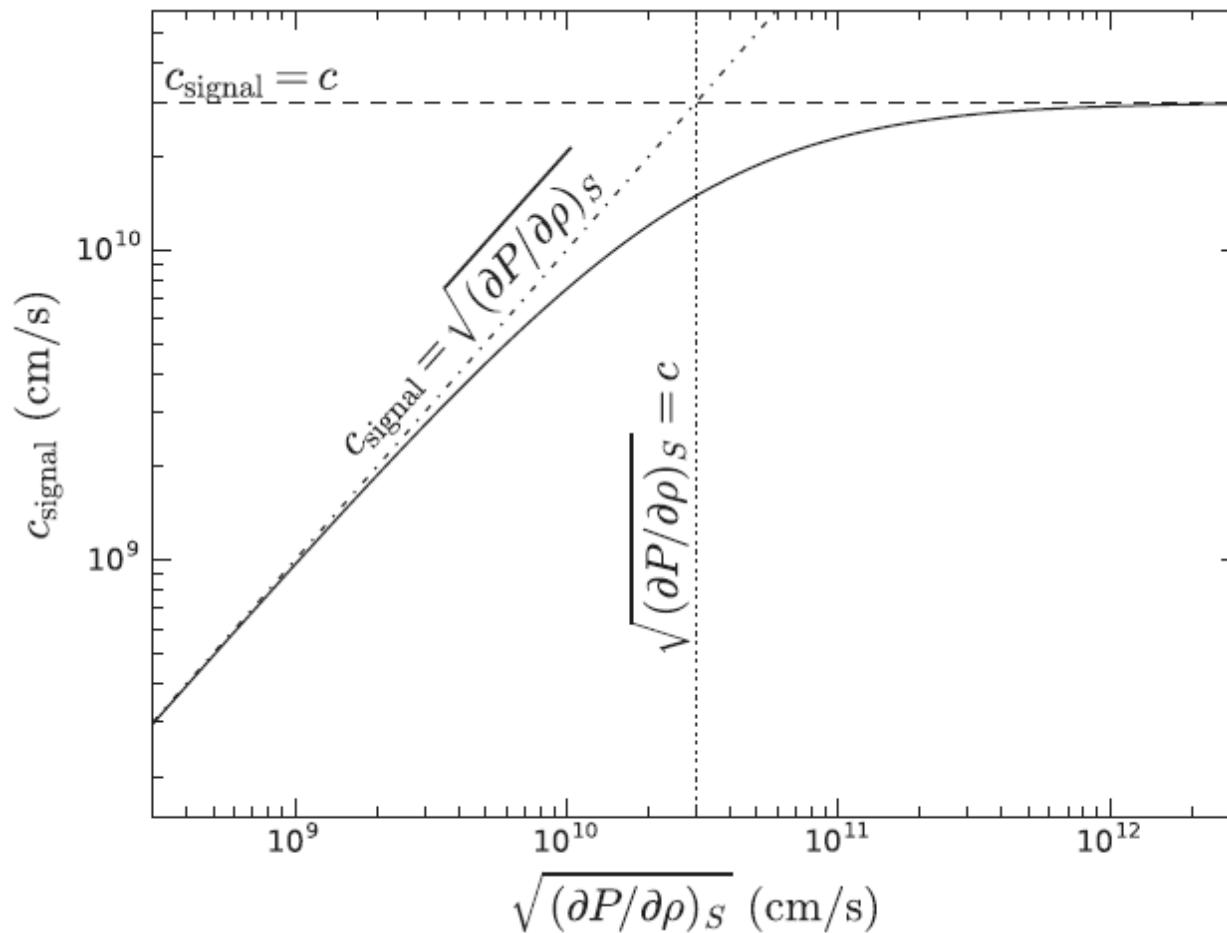
$$t_{\text{signal}} = \left[\frac{n - \frac{3}{2} + 1.855757 \left(n - \frac{3}{2}\right)^{\frac{1}{3}} + O[1]}{\sqrt{\frac{\partial P}{\partial \rho}}} + \frac{n}{c} \right] l$$

- ▶ sound speed

$$c_{\text{signal}} = \frac{nl}{t_{\text{signal}}} \approx \frac{1}{\frac{1}{\sqrt{\frac{\partial P}{\partial \rho}}} + \frac{1}{c}} < c$$

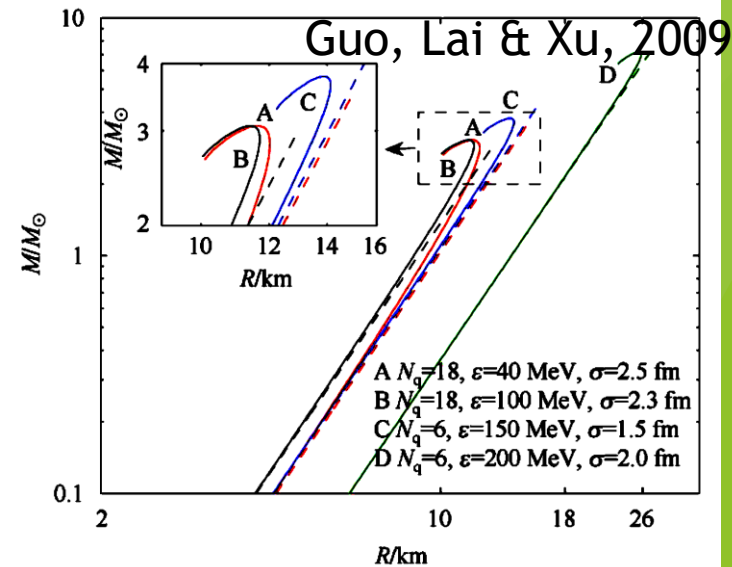
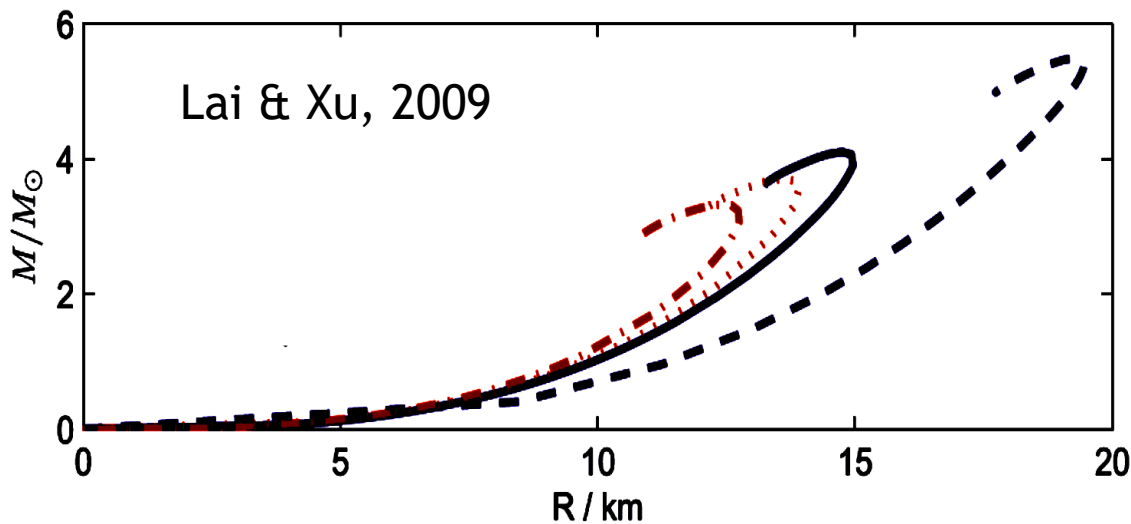
The signal speed
would never be
superluminal!!

sound speed in strangeon matter



sound speed and the maximum mass of pulsar

- $\frac{\partial P}{\partial \rho} < c^2 \Rightarrow M_{\max} \approx 3.2M_{\odot}$ (Rhoades & Ruffini, 1972), is often given the status of a rigorous upper limit in astrophysical literature.



SUMMARY

- ▶ The sound speed in strangeon matter is still subluminal.
- ▶ The large mass compact star could be the evidence for strangeon star.

Thanks!