The LUX approach to the photon PDF

P. Nason in collaboration with A. Manohar, G. Salam and G. Zanderighi

INFN, sez. di Milano Bicocca

CERN, September 13, 2016

- ► The Master Equation
- ► The LUX PDF set
- Structure functions data
- Elastic data
- Uncertainties
- Some applications
- LUX and Hoppet resources

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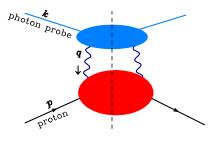
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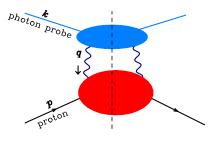
$$\begin{aligned} \sigma &= \int \frac{\mathrm{d}^4 q}{(2\pi)^4} \frac{e_{\mathrm{phys}}^4(q^2)}{q^4} \\ &\times \langle k | \tilde{J}_p^{\ \mu}(-q) J_p^{\nu}(0) | k \rangle \\ &\times \langle \rho | \tilde{J}_{\mu}(q) J_{\nu}(0) | \rho \rangle \end{aligned}$$

Kinematics constraints:

$$egin{aligned} Q^2 &= -q^2 > 0, \ 0 &< x_{
m bj} = Q^2/(2p \cdot q) \ \leq 1. \end{aligned}$$

Same kinematic restrictions as in DIS.

- $\stackrel{1}{=} \frac{1}{4\pi} \langle p | \tilde{J}_{\mu}(q) J_{\nu}(0) | p \rangle = -g_{\mu\nu} F_1(Q^2, x_{\rm bj}) + \frac{p^{\mu} p^{\nu}}{p \cdot q} F_2(Q^2, x_{\rm bj}) + \dots$ (Notice: full F_1 and F_2 , not only inelastic)
- > Photon induced process can be given in terms of F_1 , F_2
- Hence: the photon PDF must be calculable in terms of F_1 , F_2 .



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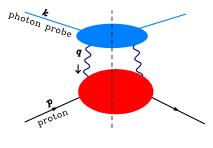
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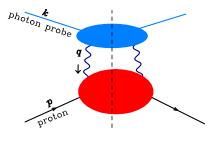
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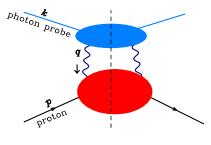
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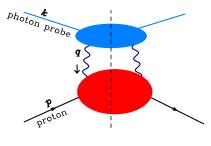
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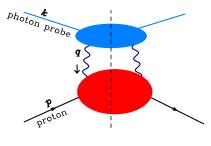


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- Compute the cross section with the Master Formula
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$$xf_{\gamma/\rho}(x,\mu^2) = \frac{1}{2\pi} \int_x^1 \frac{dz}{z} \left\{ -\alpha(\mu^2) z^2 F_2\left(\frac{x}{z},\mu^2\right) \right\}$$

$$+\underbrace{\int_{\frac{x^{2}m_{p}^{2}}{1-z}}^{\frac{\mu^{2}}{1-z}} \frac{dQ^{2}}{Q^{2}} \frac{\alpha^{2}(Q^{2})}{\alpha(\mu^{2})} \left[\left((1+(1-z)^{2}) + \frac{2x^{2}m_{p}^{2}}{Q^{2}} \right) F_{2}(x/z,Q^{2}) - z^{2}F_{L}\left(\frac{x}{z},Q^{2}\right) \right] \right\}$$

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- $Q^2 \approx m_p^2$ region formally of order α , i.e. NLO (as $\overline{\mathrm{MS}}$ term).
- Straightforward to improve at NNLO in α_s (Master Equation is exact, compute the parton model process at NNLO)
- Also accurate at (α/α_s)², provided that α(Q²) and F₂ include leading log electromagnetic evolution.
- Valid at all μ's: MUST match evolution accuracy with one extra α_s. Agrees with De FLorian, Sborlini, Rodrigo αα_s splitting functions, arXiv:1512.00612.

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- $Q^2 \approx m_p^2$ region formally of order α , i.e. NLO (as $\overline{\mathrm{MS}}$ term).
- Straightforward to improve at NNLO in α_s (Master Equation is exact, compute the parton model process at NNLO)
- ► Also accurate at (α/α_s)², provided that α(Q²) and F₂ include leading log electromagnetic evolution.
- Valid at all μ's: MUST match evolution accuracy with one extra α_s. Agrees with De FLorian, Sborlini, Rodrigo αα_s splitting functions, arXiv:1512.00612.

Ideal use:

- Get $F_{2/L}$ at low Q^2 from available data.
- ▶ PDF global fit, including EM evolution, with the photon density constrained by the previous equation, F_{2/L} taken from data at low Q² and computed from the PDF's at high Q²

- Low Q^2 region cannot be neglected.
- (α/α_s)² terms arising from the evolution of QED coupling cannot be neglected (α(m²_μ))/α(M²_Z) ≈ 0.94)
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- Start from a standard set (e.g. PDF4LHC15_nnlo_100);
- Compute the photon PDF at µ = 100 GeV, with the low Q² component determined from A1, CLAS and Hermes GD11-P fits, and the high Q² part determined from the input PDF with standard NNLO coefficient functions.
- ► Evolve down to 10 GeV, including QED evolution only for splitting processes that affect the photon: $P_{\gamma q}$, $P_{\gamma g}$, $P_{\gamma \gamma}$ (with $\alpha \alpha_s$ terms included).
- Fix the momentum sum rule by rescaling the gluon (a factor of 0.99299 is needed).
- Evolve up including full QED evolution (with αα_s terms included).

This procedure is such that the structure functions at a scale of 10 GeV, where they are strongly data constrained, remain consistent with the new pdf set, while the $(\alpha/\alpha_s)^2$ due to photon radiation are included in the quark distributions at high scale

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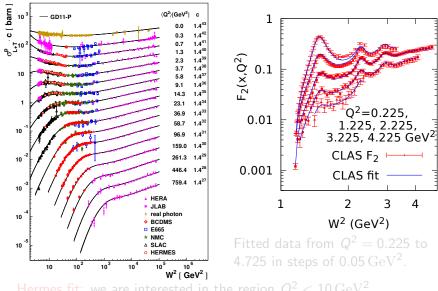
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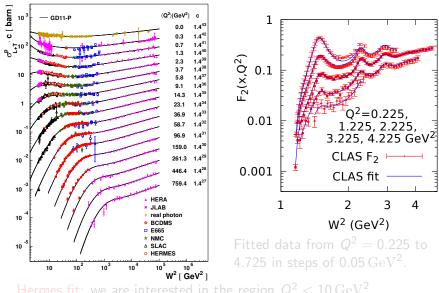
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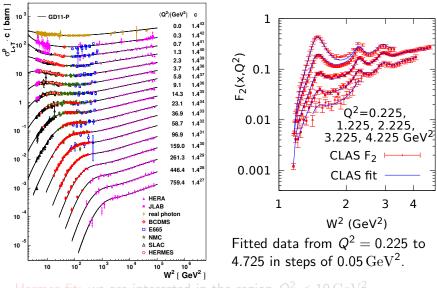
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Continuum data region: 4 $\text{GeV}^2 < W^2 \lesssim 10^5 \text{GeV}^2 \ (x \to 10^{-4}).$

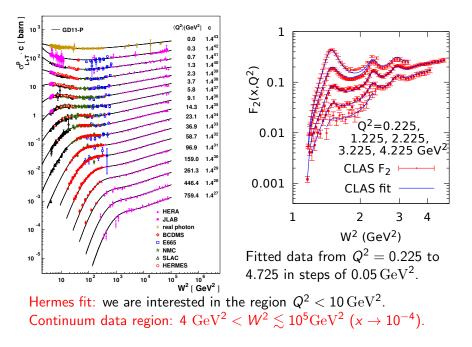


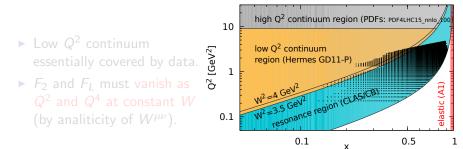
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Hermes fit: we are interested in the region $Q^2 < 10 \text{ GeV}^2$. Continuum data region: $4 \text{ GeV}^2 < W^2 \lesssim 10^5 \text{GeV}^2 (x \rightarrow 10^{-4})$.

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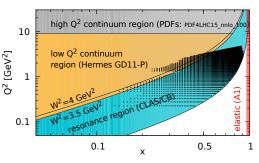


Also:

$$F_2(x, Q^2) = \frac{1}{4\pi^2 \alpha} \frac{Q^2(1-x)}{1 + \frac{4x^2 m_p^2}{Q^2}} (\sigma_T(x, Q^2) + \sigma_L(x, Q^2)) \Longrightarrow_{Q^2 \to 0} \frac{Q^2 \sigma_{\gamma p}(W)}{4\pi^2 \alpha^2}.$$

At small Q^2 , $\sigma_T \Longrightarrow \sigma_{\gamma p}(W)$, becoming a function of W only (the *CM* energy in photoproduction), and σ_L vanishes. Photoproduction data included in Hermes and Christy-Bosted parametrizations.

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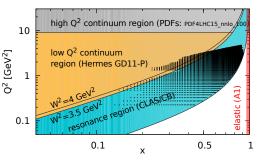
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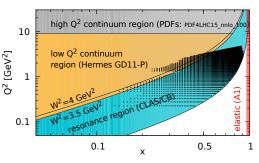


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Elastic Contribution

 F_2 and F_L receive an elastic contribution that we must include:

$$\begin{split} F_2^{\rm el} &= \frac{G_E^2(Q^2) + G_M^2(Q^2)\tau}{1+\tau} \,\delta(1-x), \\ F_L^{\rm el} &= \frac{G_E^2(Q^2)}{\tau} \,\delta(1-x), \end{split}$$

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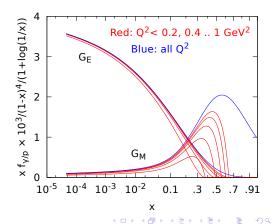
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so that the elastic contribution falls rapidly with Q^2 .

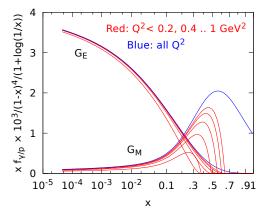
$$\begin{aligned} x f_{\gamma}^{\text{el}}(x,\mu^2) &= \frac{1}{2\pi} \int_{\frac{x^2 m_p^2}{1-x}}^{\frac{\mu^2}{1-x}} \frac{\mathrm{d}Q^2}{Q^2} \frac{\alpha^2(Q^2)}{\alpha(\mu^2)} \Biggl\{ \left(1 - \frac{x^2 m_p^2}{Q^2(1-x)} \right) \frac{2(1-x)G_E^2(Q^2)}{1+\tau} \\ &+ \left(2 - 2x + x^2 + \frac{2x^2 m_p^2}{Q^2} \right) \frac{G_M^2(Q^2)\tau}{1+\tau} \Biggr\}. \end{aligned}$$

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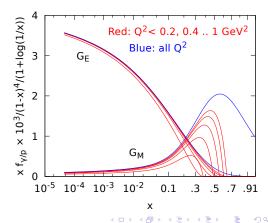
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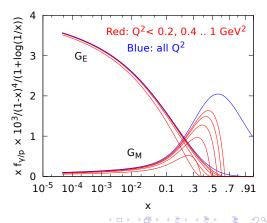
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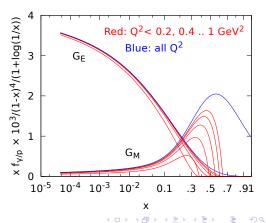
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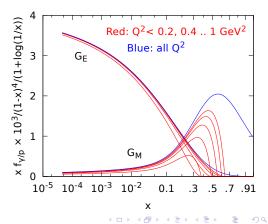
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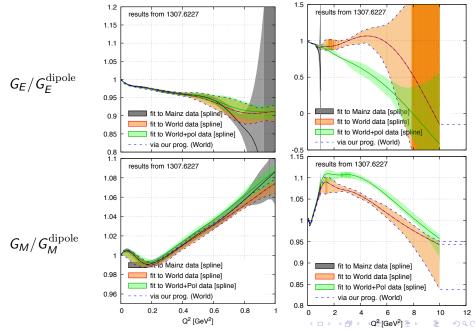


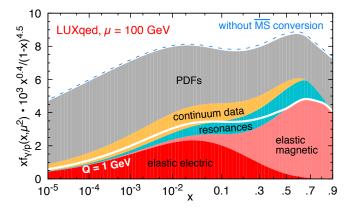
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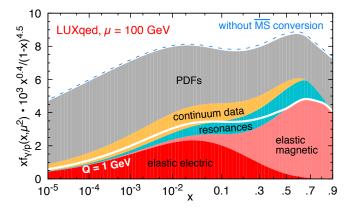


Elastic Data, A1 experiment and World data

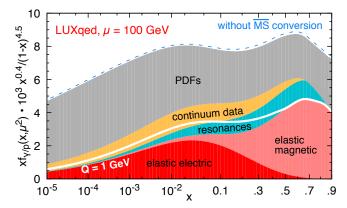




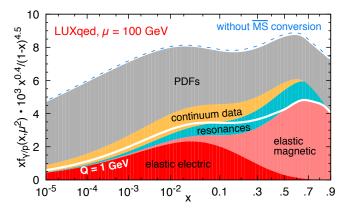
- Important elastic component. Magnetic prevails for x > 0.2.
- Continuum and resonance contributions not negligible
- Very important contribution from $Q^2 < 1 \text{ GeV}^2$.



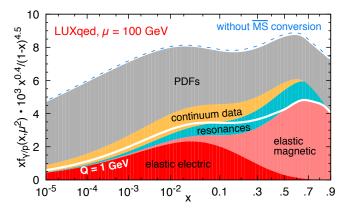
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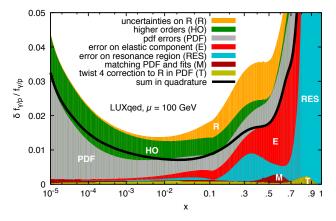
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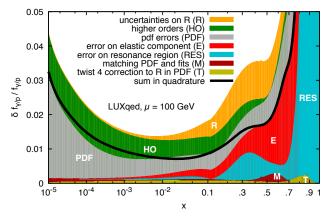


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At small x, higher order effects and PDF's dominate the error.

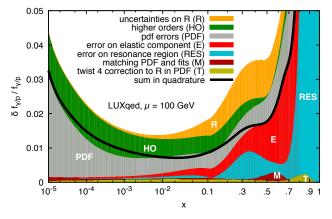
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- Total uncertainty at the percent level. Further improvements possible!



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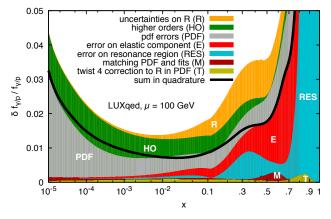
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Uncertainties included in LUX

Added members with variations in photon PDF calculation:

- 0-100: original PDF members (PDF4LHC15_nnlo_100)
- ▶ 101: Replace CLAS parametrization of resonance region with Christy-Bosted one. (Becomes particuarly crazy al large *x*).
- ▶ 102: rescale R in low Q^2 region by 1.5.
- ▶ 103: rescale *R* in high-*Q*² region with a higher-twist component.
- 104: Use 'World" elastic fit from A1: no polarization data, no fit to Two Photon Exchange effects.
- ▶ 105: Use lower edge of elastic fit error band.
- ▶ 106: Start using PDF's from $Q^2 = 5$ rather than 9 GeV^2 .
- ► 107: Upper limit of integration in f_γ formula changed to μ² instead of μ²/(1 − z), with suitable correction of MS term.

All errors are taken as symmetric.

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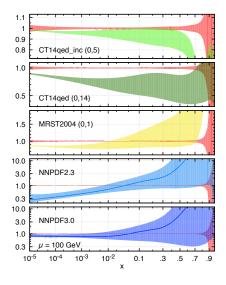
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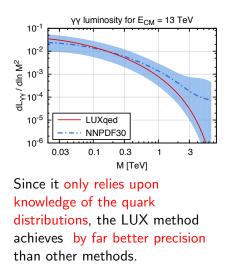
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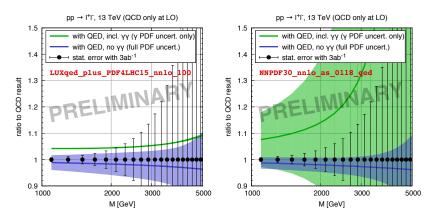
Approaches that use some lepton scattering information (in particular CT14qed_inc) do achieve better precision (note different y axis in panel).

APPLICATION TO HIGGS PHYSICS

pp \rightarrow H W ⁺ (\rightarrow l ⁺ v) + X at 13 TeV	
non-photon induced contributions	91.2 ± 1.8 fb
photon-induced contribs (NNPDF23)	6.0 ^{+4.4} –2.9 fb
photon-induced contribs (LUXqed)	4.4 ± 0.1 fb

non-photon numbers from LHCHXSWG (YR4)

di-lepton spectrum



LUXQED photon has few % effect on di-lepton spectrum and negligible uncertainties

- LUXqed_plus_PDF4LHC15_nnlo_100 set available from LHAPDF
- Additional plots and validation info available from <u>http://cern.ch/luxqed</u>
- Preliminary version of HOPPET DGLAP evolution code with QED (order α and αα_s) corrections available from hepforge:

svn checkout http://hoppet.hepforge.org/svn/branches/qed hoppet-qed

(look at tests/with-lhapdf/test_qed_evol_lhapdf.f90 for an example; interface may change, documentation missing)

- Photon PDF can be extracted with great precision from available knowledge of proton structure function and form factors.
- ▶ The needed low Q² data is available thanks to extensive low and intermediate energy Nuclear Physics studies.
- Our study aimed at NLO precision including terms suppressed by one power of α_s or by a power of α/α_s relative to the leading term. This leads to precisions at the percent level.
- The study of structure functions and form factors at low energy is still ongoing in the Nuclear Physics Community (further progress will come).
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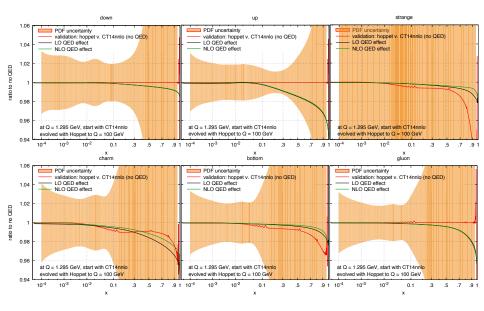
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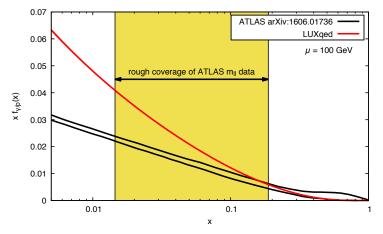
EXTRA SLIDES

Impact of QED evolution



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ratio of ATLAS photon (1606.01736) to LUXqed



ATLAS result based on reweighting of NNPDF23 with highmass ($M_{\rm ll} > 116$ GeV) data

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