

Depleted Argon for Future Direct Dark Matter Searches

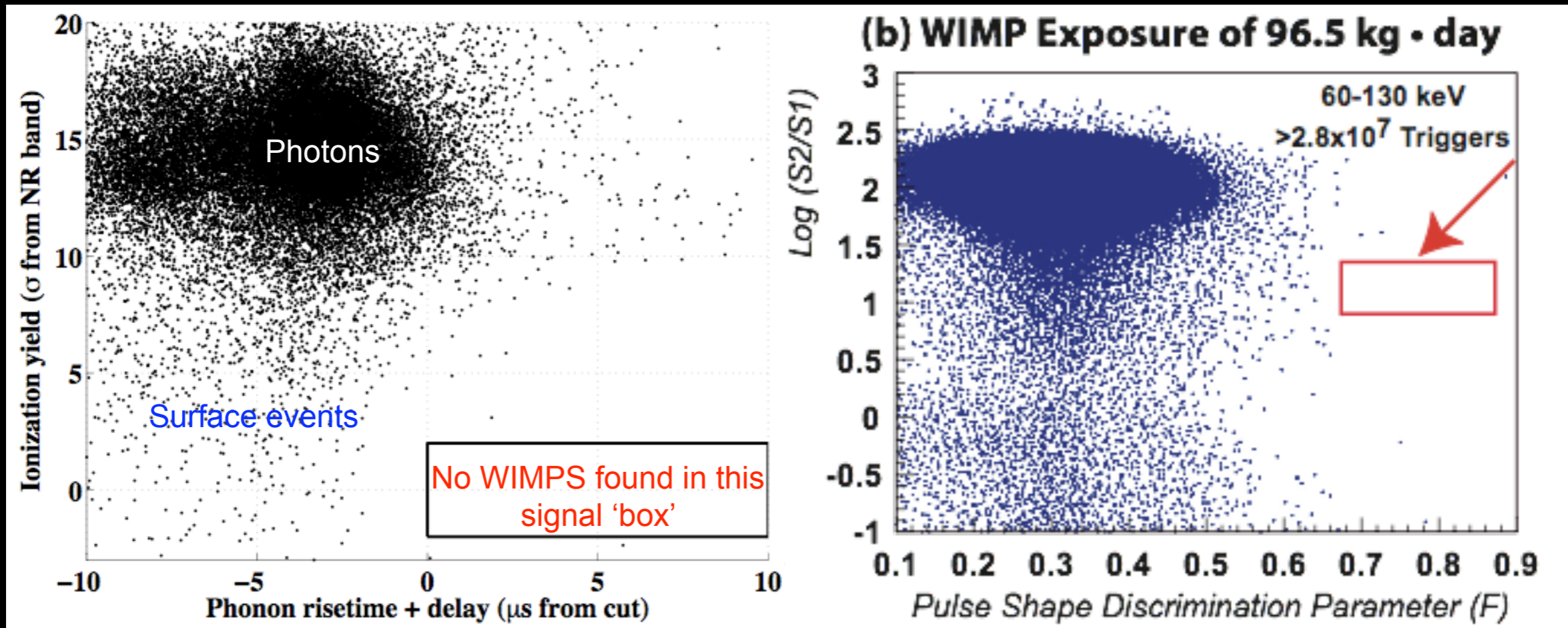
Cristiano Galbiati
Princeton University

WIN'09
Perugia, Italy

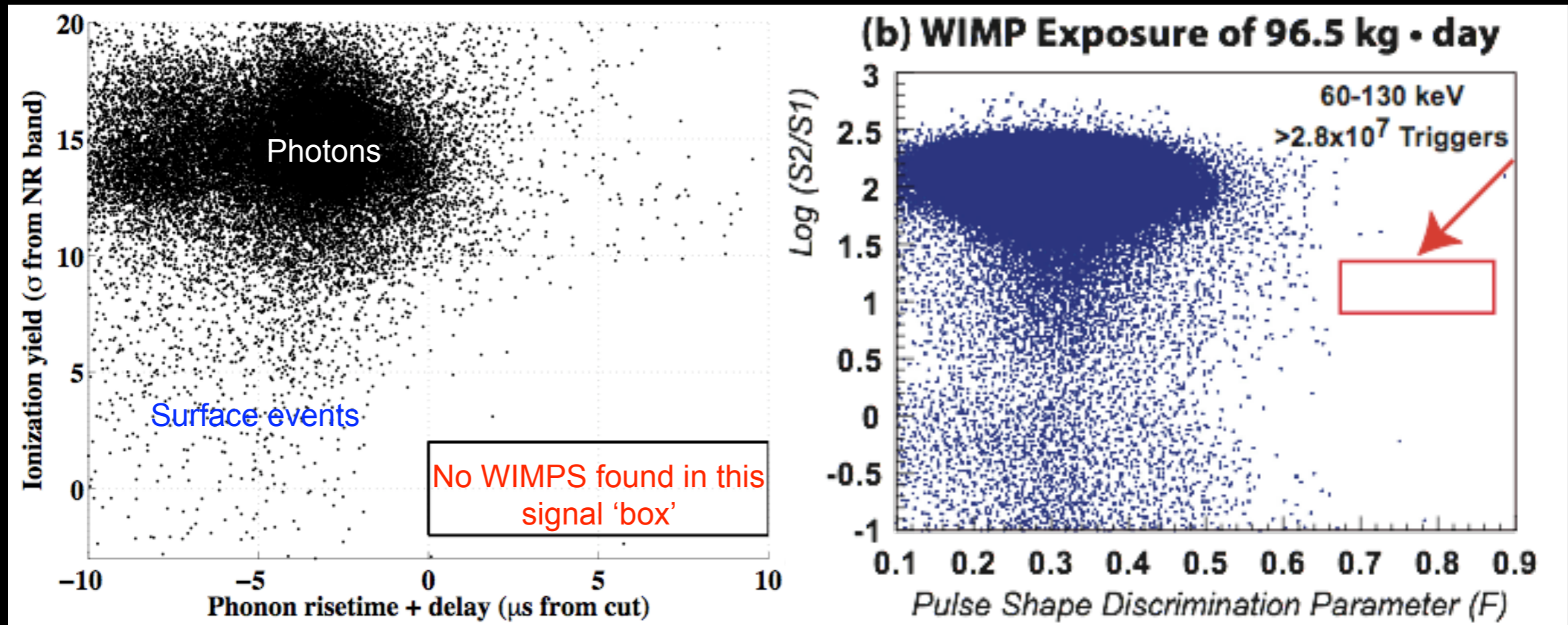
Sep 17, 2009



Two is Better than One!



Two is Better than One!



Few Events in \sim Tons per Yr at few tens of KeV
Crucial the presence of dual, semi-independent
discrimination

Discrimination in Depleted Argon TPC

- **Pulse shape discrimination of primary scintillation (S1)** based on the difference in decay times between singlet (≈ 7 ns) and triplet ($1.6 \mu\text{s}$) components of the emitted UV light
 - Minimum ionizing: triplet/singlet $\sim 3/1$
 - Nuclear recoils: triplet/singlet $\sim 1/3$
 - Theoretical Identification Power exceeds 10^8 for > 60 photoelectrons (Boulay & Hime 2004)
- **Difference in ratio of the prompt scintillation (S1) to the drift time-delayed ionization (S2)**
- **Precise determination of events location in 3D**
- **Up-front reduction of large ^{39}Ar cosmogenic rate**

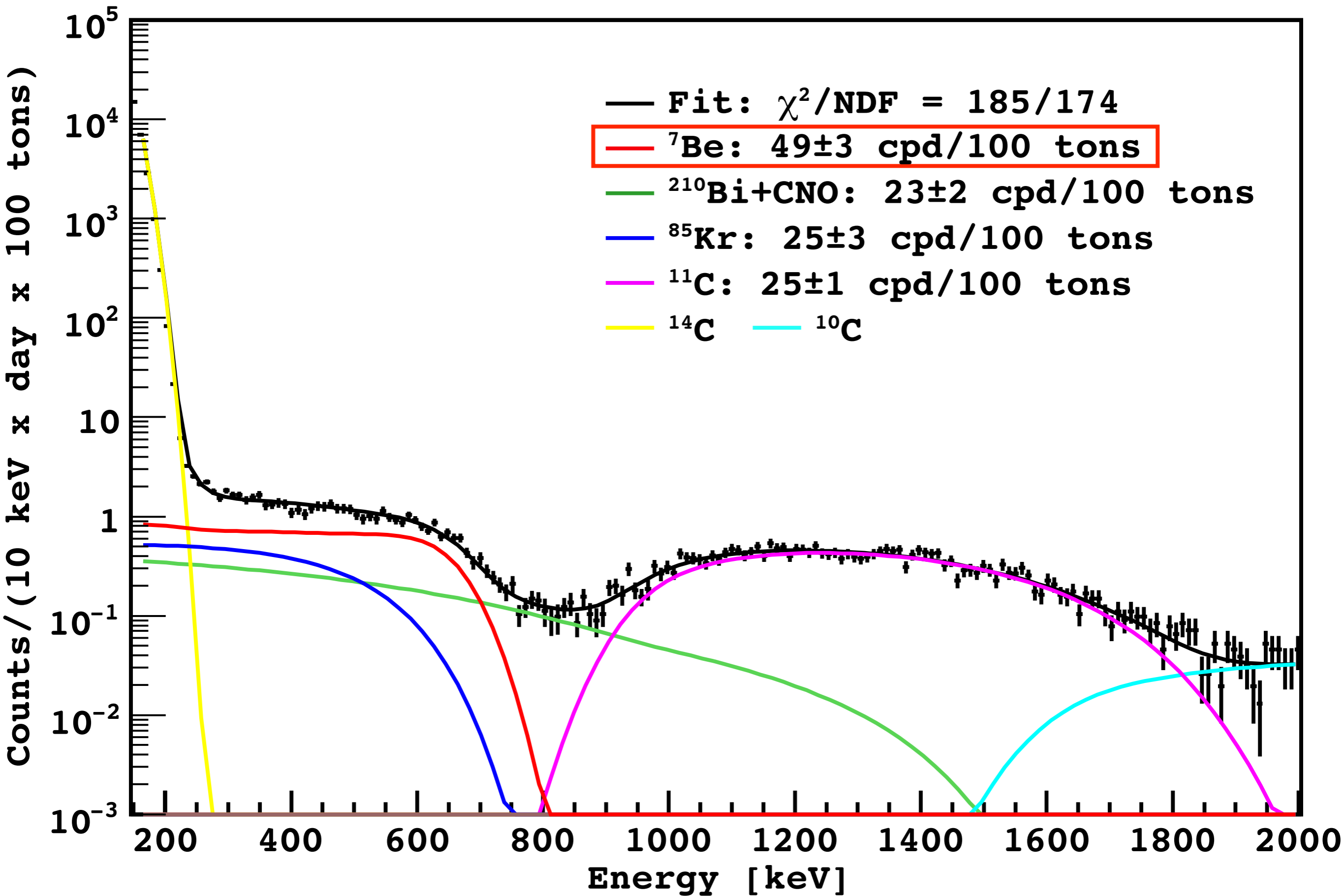
Why is underground argon desirable?

- Radioactive ^{39}Ar produced by cosmic rays in atmosphere
 - beta decays, $Q = 565 \text{ keV}$, $t_{1/2} = 269 \text{ years}$
- In atmospheric argon:
 - $^{39}\text{Ar}/\text{Ar}$ ratio 8×10^{-16}
 - specific activity 1 Bq/kg
- Limits size (and sensitivity) of argon detectors to 500-1000 kg due to ^{39}Ar events pile-up

Why is underground argon desirable?

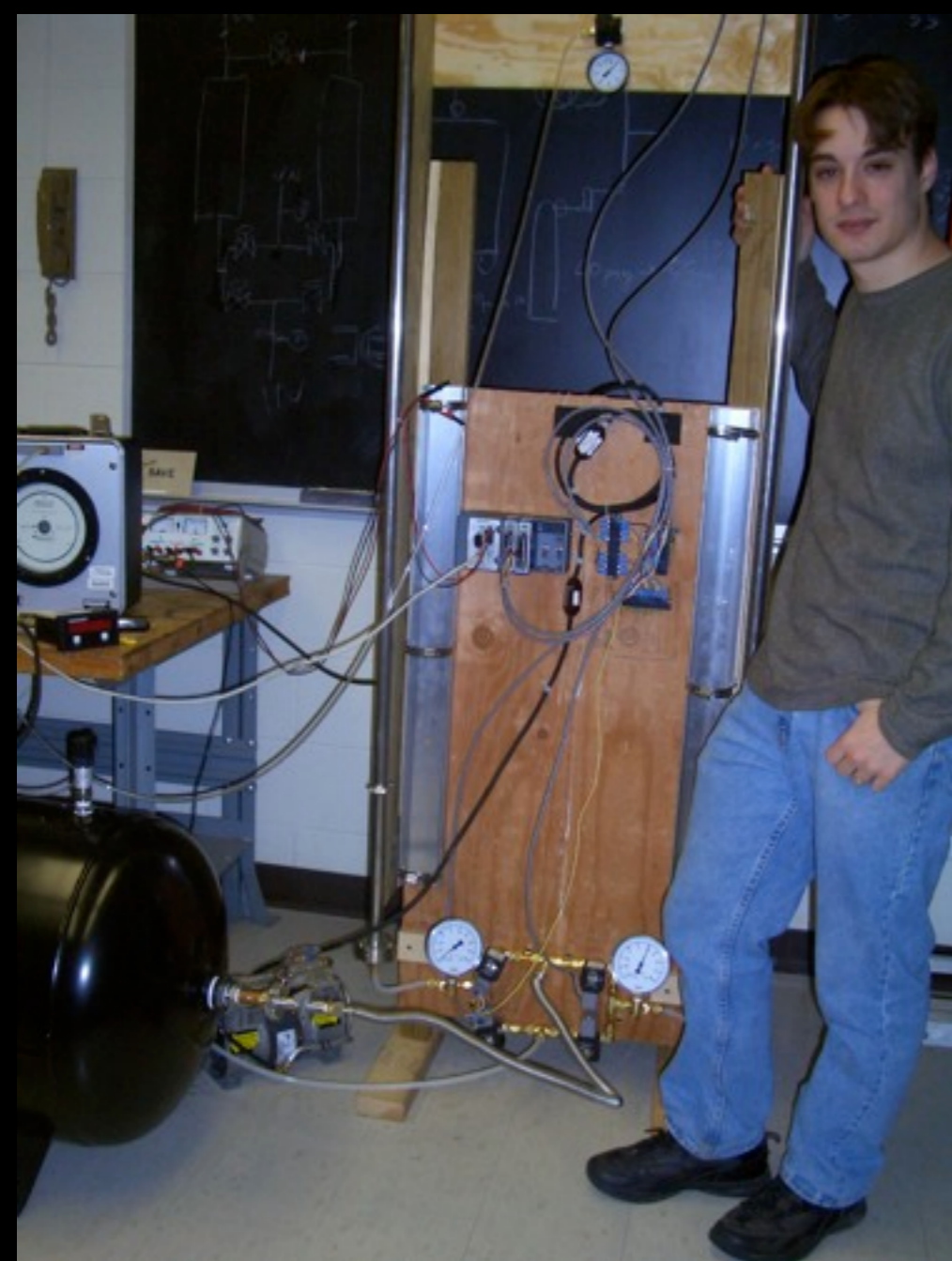
- ^{39}Ar -depleted argon available via centrifugation or thermal diffusion, but expensive at the ton scale!
- ^{39}Ar production by cosmic rays strongly suppressed underground
- Motivated by success in Borexino
 - Low background from ^{14}C crucial for observation of low energy neutrinos with organic liquid scintillators.
 - Hydrocarbons in deep underground reservoirs results in low cosmogenic ^{14}C

New Results: 192 Days



Necessary to pre-scan sources of interest for ^{39}Ar

- ^{39}Ar also produced underground by neutron activation, from fission and (α, n) neutrons
 - $^{39}\text{K}(n, p)^{39}\text{Ar}$
- ^{39}Ar content depends on local content of U, Th, and K, and on rock porosity
- In some groundwater samples $^{39}\text{Ar}/\text{Ar}$ ratio measured up to a factor $20\times$ (2000%) of the atmospheric ratio
- Cannot rely on ^{39}Ar simply being low. Pre-scan of ^{39}Ar activity on small samples necessary for program.



← Prototype Purification Plant
at Princeton

Sampling gas field in the West



All across the West:
Texas, New Mexico,
Colorado, Kansas

Sources Viable for Large Scale Production

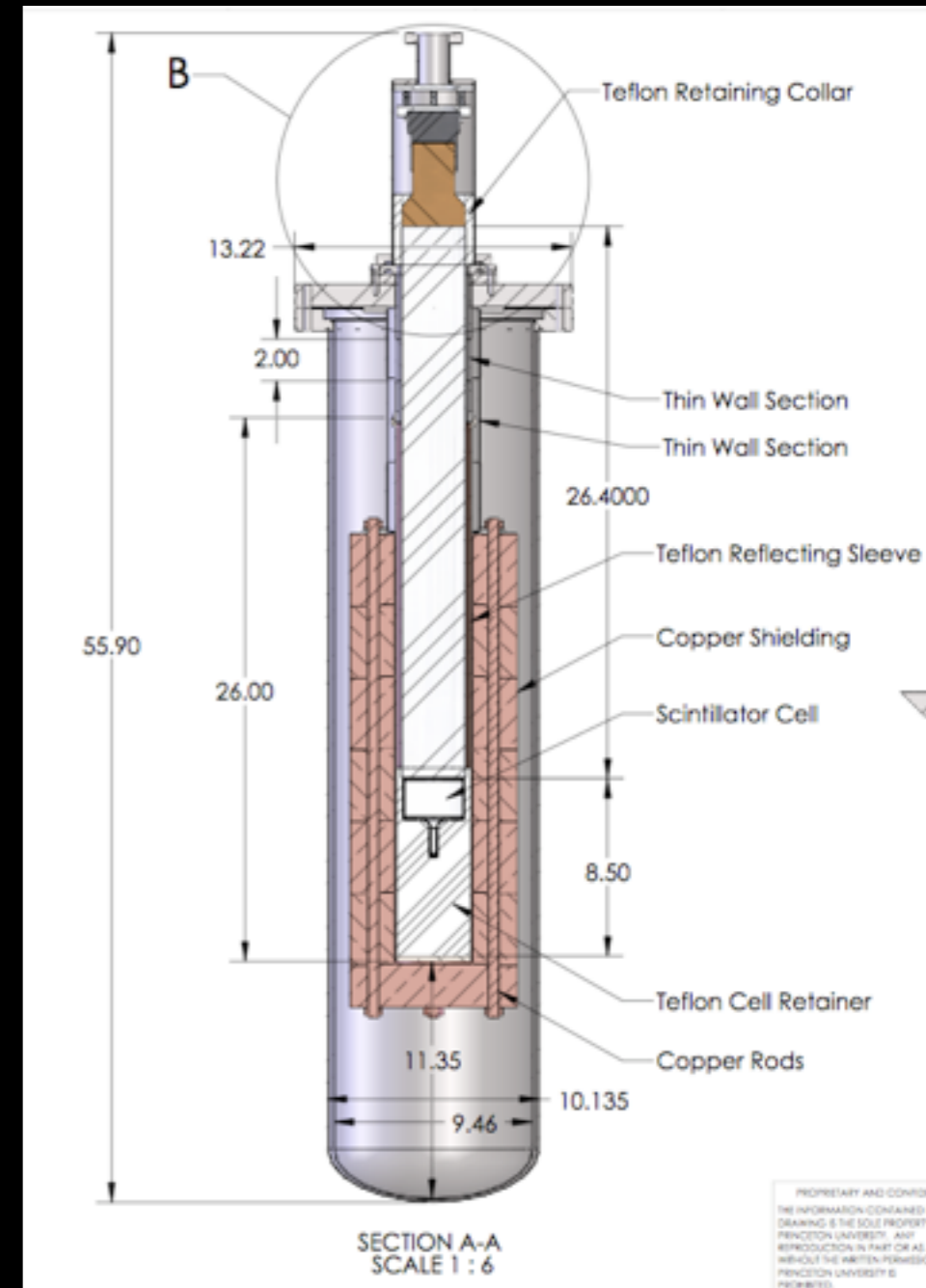
	National Helium Reserve	Doe Canyon Complex
^{39}Ar	<50 mBq/kg	<40 mBq/kg
Depletion	>20	>25
Production capacity	30 tons/yr	>10 tons/yr

Measurements of low levels of ^{39}Ar

- Single phase ~ 1 kg Ar liquid scintillator
- Long light pipe to Room Temp PMT
- Internal Cu shield. External Pb shielding not shown
- Measure ^{39}Ar spectrum.
- Goal: < 1 mB/kg (0.1% atm.)



Inner Tube

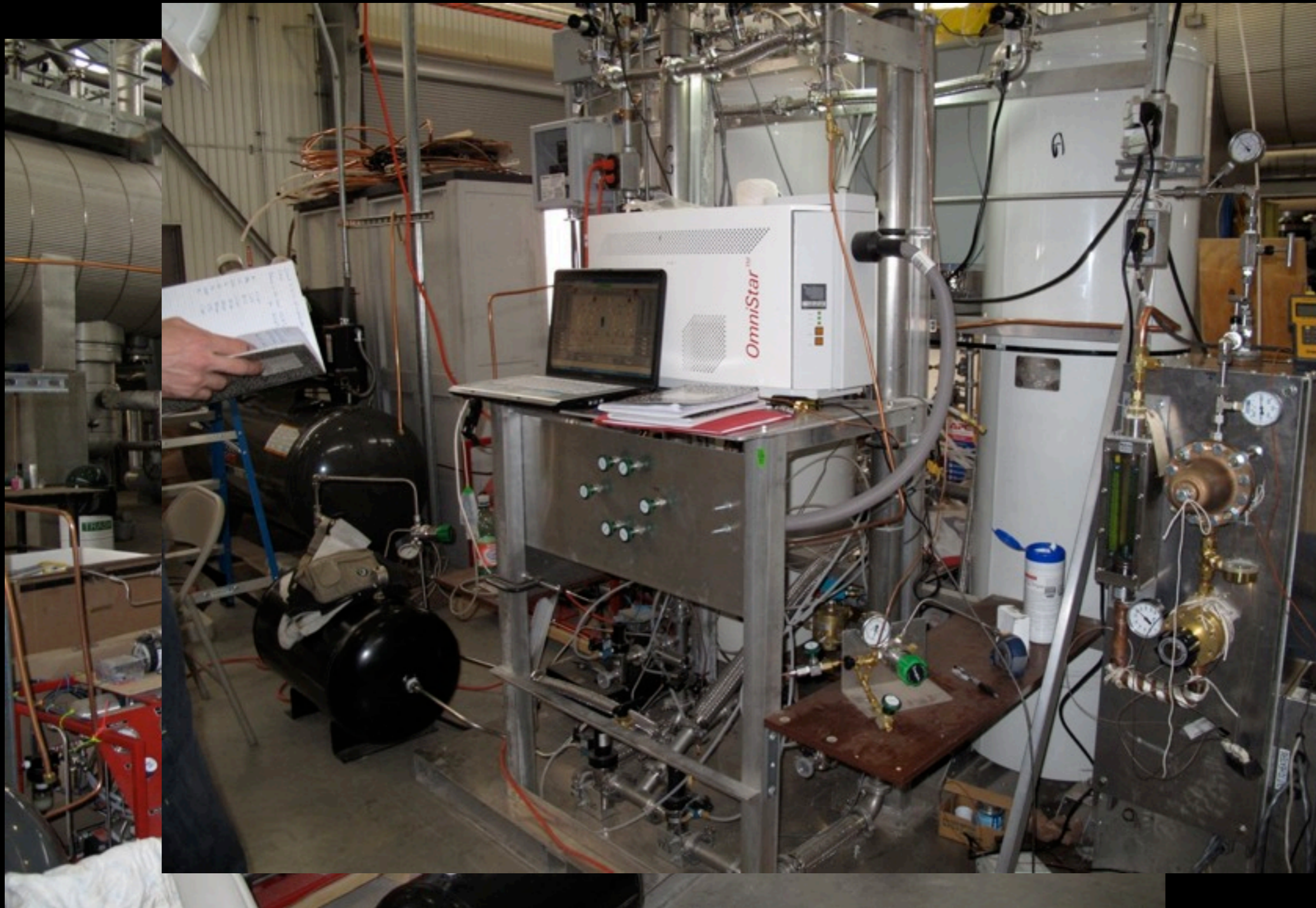


Cryostat with Inner Tube and internal Cu shield

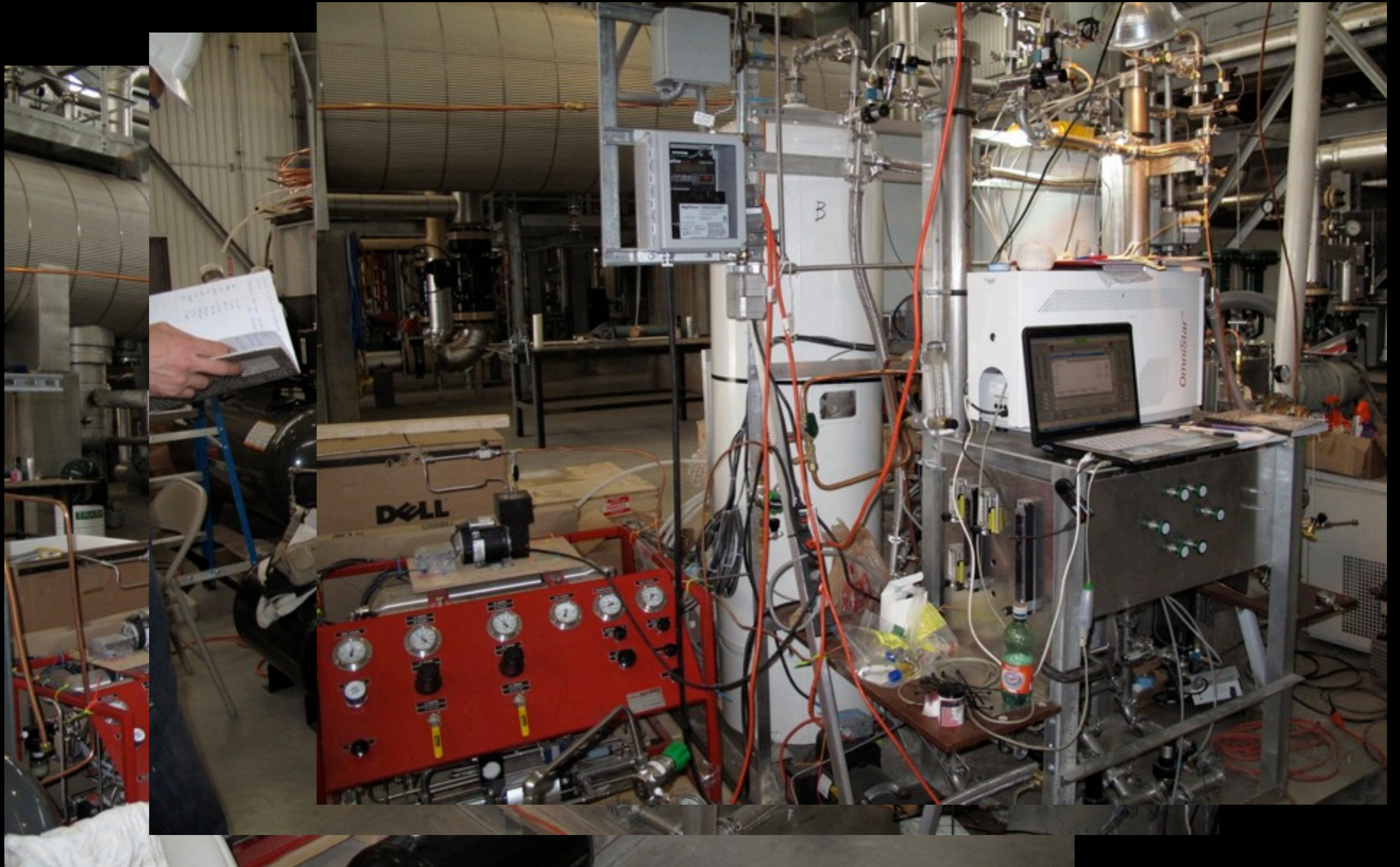
Princeton Prototype Plant for Industrial Scale Production: Achieved 0.5 kg/day



Princeton Prototype Plant for Industrial Scale Production: Achieved 0.5 kg/day



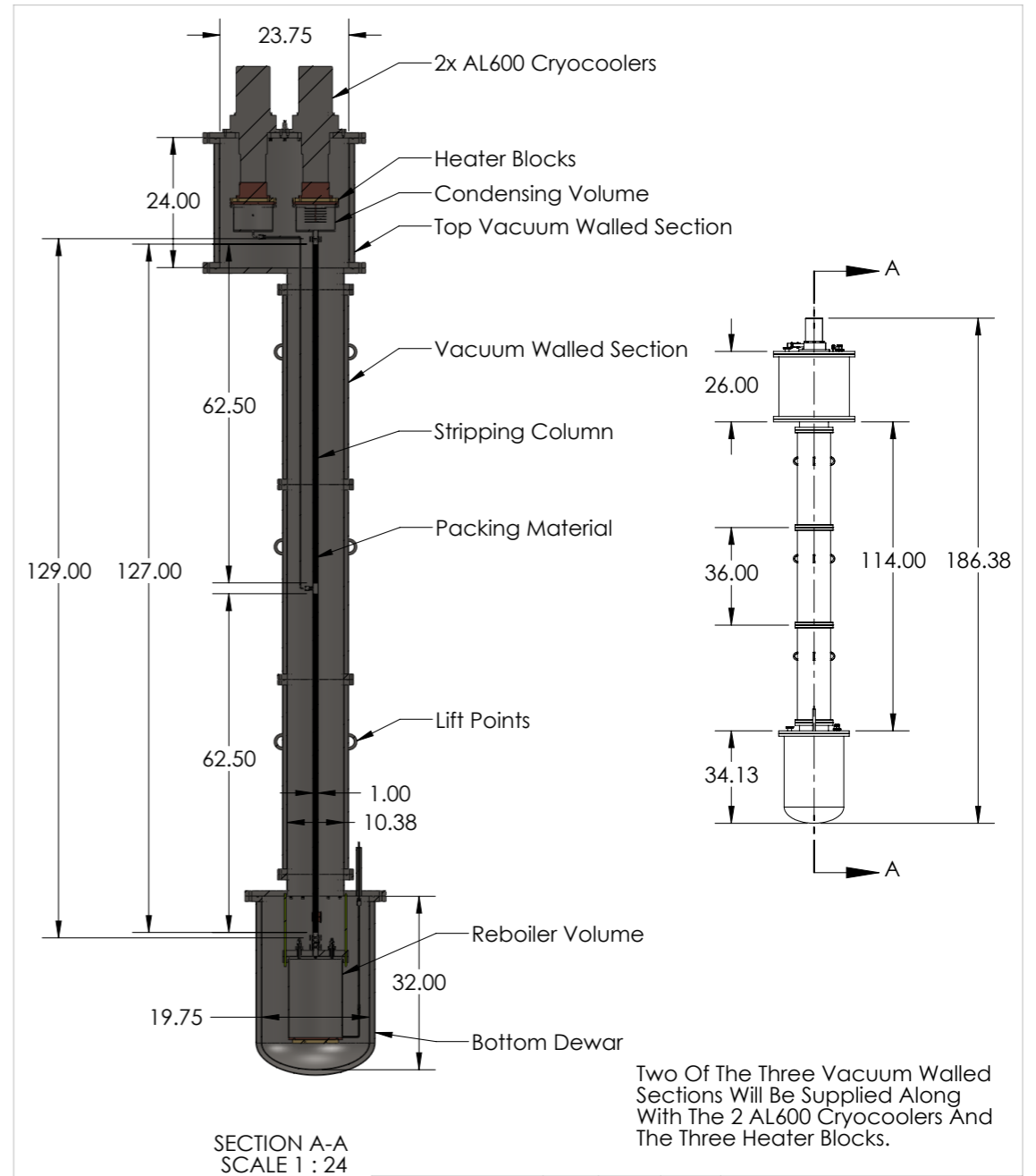
Princeton Prototype Plant for Industrial Scale Production: Achieved 0.5 kg/day



Purification by cryogenic distillation

Cryogenic distillation column designed at Princeton, to be commissioned Summer-Fall 2009

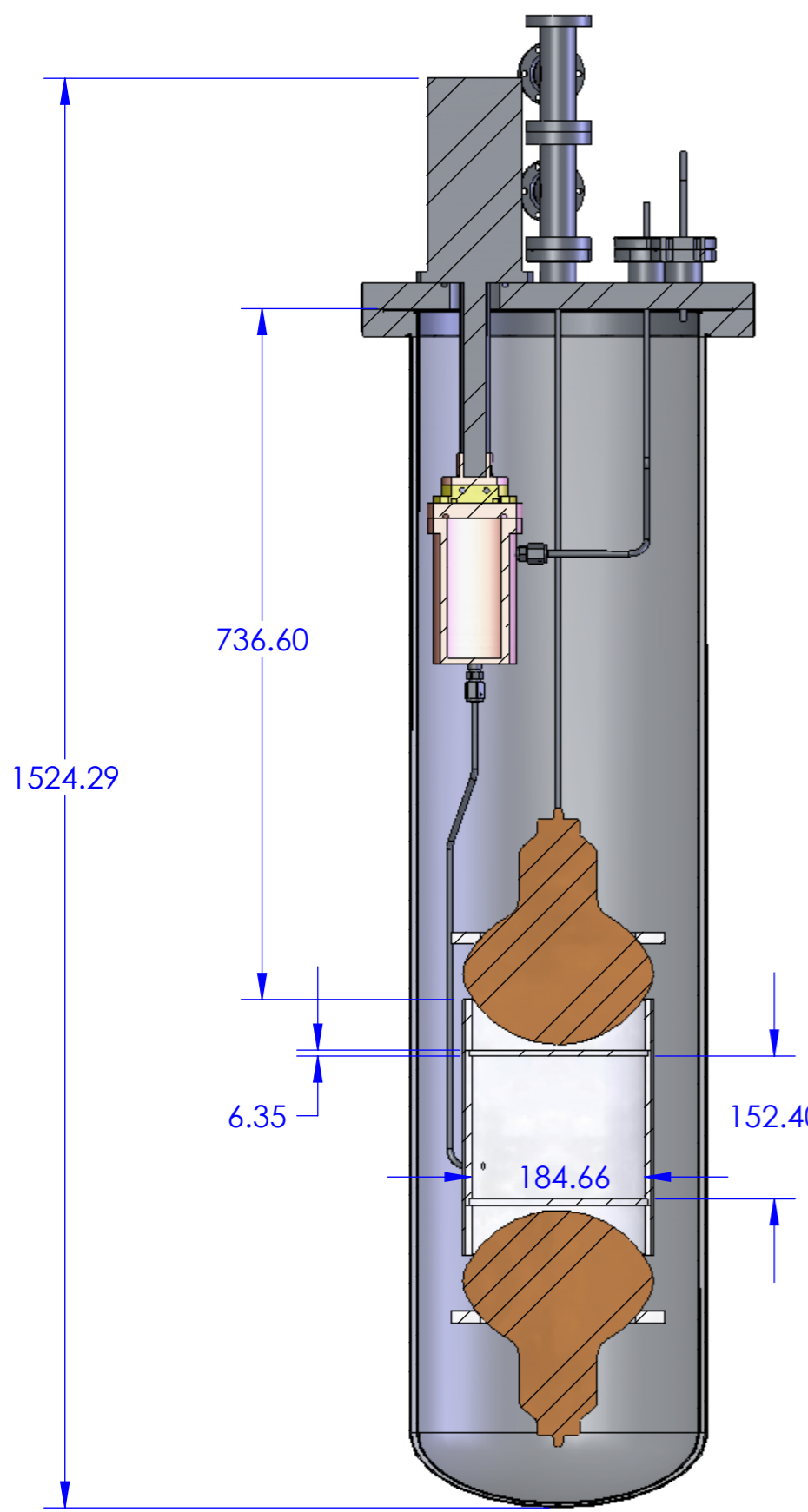
Performance:
purification of 20 kg per day of argon to 4-5 nines of purity



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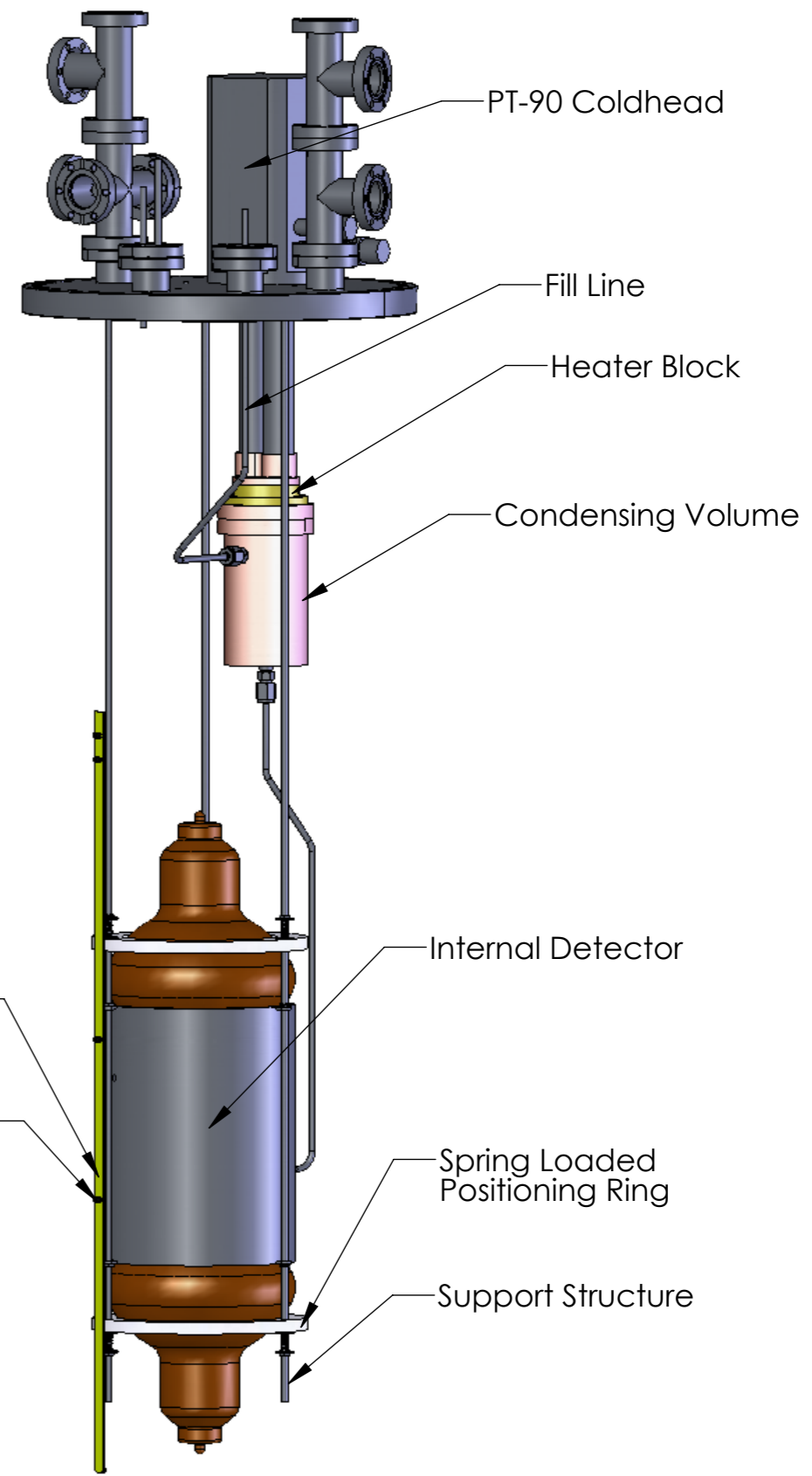
DIMENSIONS ARE IN INCHES		NAME	DATE
TOLERANCES:		TH	4/13/09
FRACTIONAL ±			
ANGULAR; MACH ± BEND ±			
TWO PLACE DECIMAL ±			
THREE PLACE DECIMAL ±			
MATERIAL	Mostly 304 Stainless Steel	Q.A.	
FINISH		COMMENTS:	
DO NOT SCALE DRAWING		Any Questions Call Tristen Hohman: Office # - 609-258-6761 Cell # - 908-399-0783	

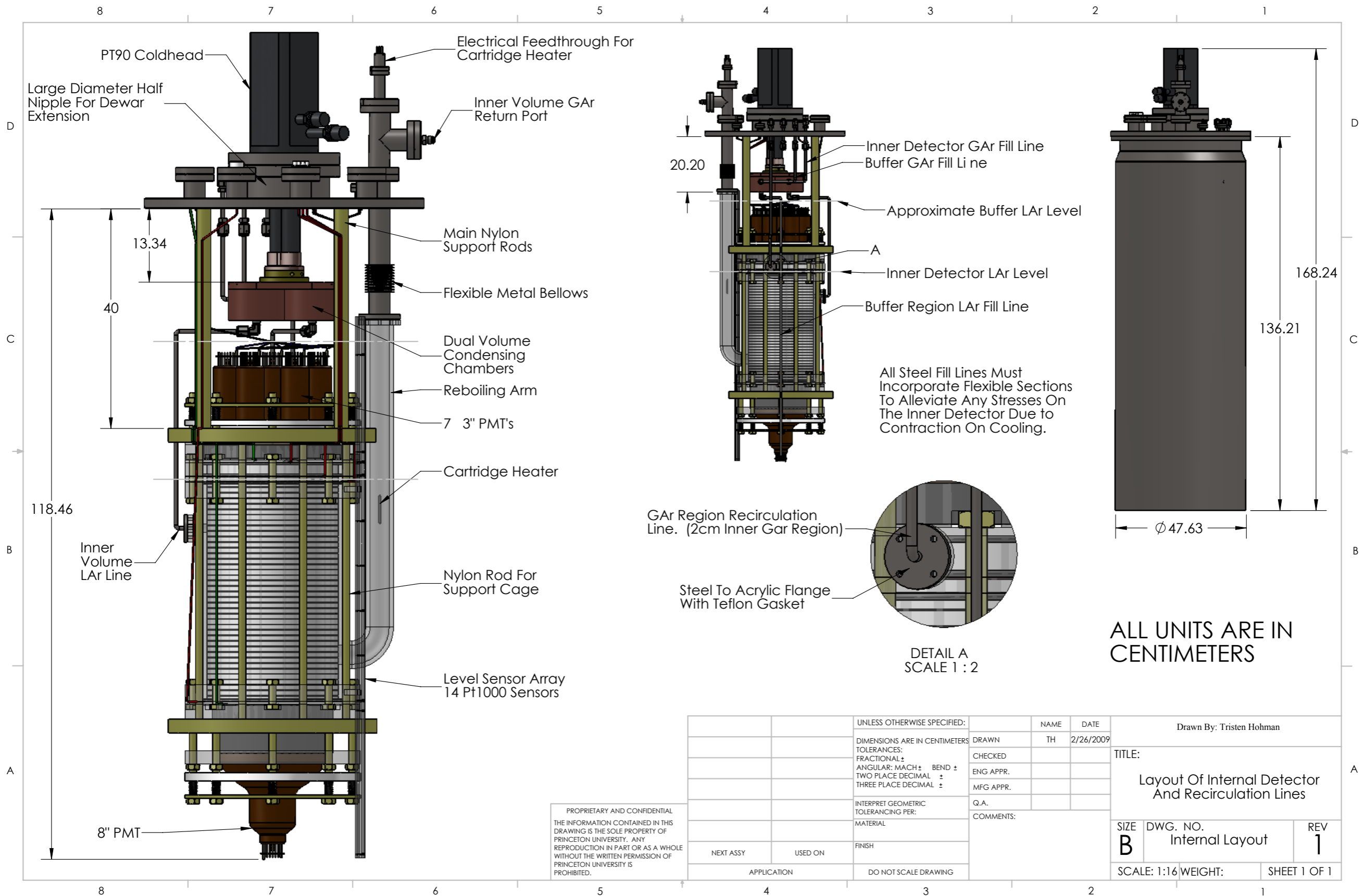
Drawn By: Tristen Hohman			
Main View of Cryostripping Column			
SIZE	DWG. NO.	REV.	
A	cryostripping column revised 3	3	
SCALE: 1:48	WEIGHT:	SHEET 1 OF 1	



- Level Probe Positions
(From top flange)
- 1 - 454.12 mm
 - 2 - 477.52 mm
 - 3 - 774.70 mm
 - 4 - 944.88 mm
 - 5 - 1231.9 mm

64 kg LAr
6.7 kg active mass
Light Yield: 5 pe/keV





PT90 Coldhead

Large Diameter Half Nipple For Dewar Extension

13.34

40

118.46

Inner Volume LAr Line

8" PMT

Electrical Feedthrough For Cartridge Heater

Inner Volume GAR Return Port

Main Nylon Support Rods

Flexible Metal Bellows

Dual Volume Condensing Chambers

Reboiling Arm

7 3" PMT's

Cartridge Heater

Nylon Rod For Support Cage

Level Sensor Array 14 Pt1000 Sensors

20.20

Inner Detector GAR Fill Line

Buffer GAR Fill Line

Approximate Buffer LAr Level

A

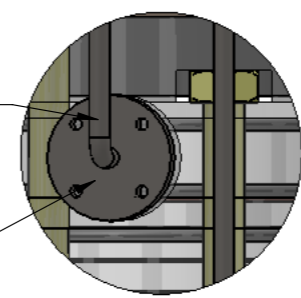
Inner Detector LAr Level

Buffer Region LAr Fill Line

All Steel Fill Lines Must Incorporate Flexible Sections To Alleviate Any Stresses On The Inner Detector Due to Contraction On Cooling.

GAR Region Recirculation Line. (2cm Inner Gar Region)

Steel To Acrylic Flange With Teflon Gasket



DETAIL A
SCALE 1 : 2

168.24

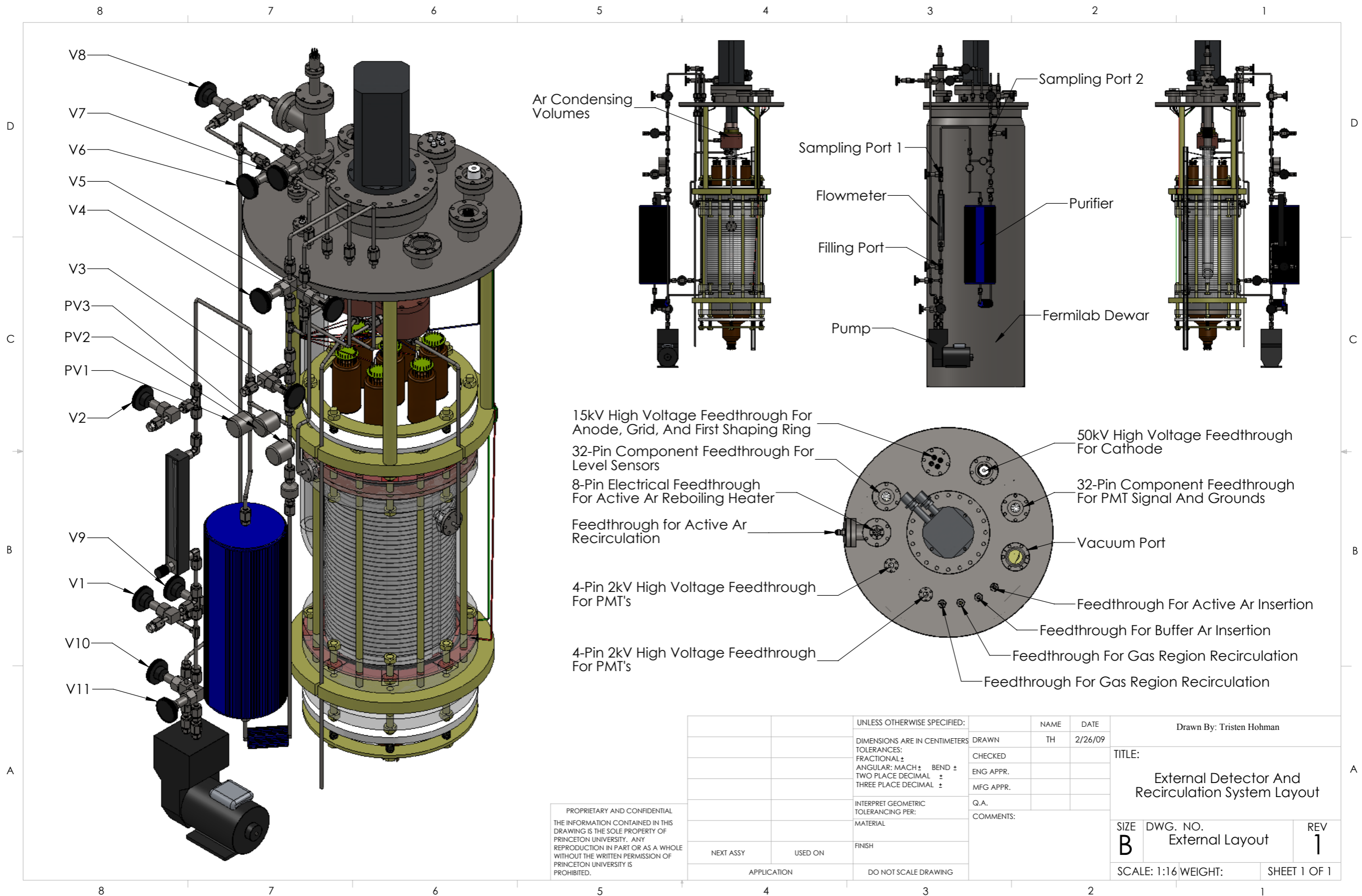
136.21

Ø 47.63

ALL UNITS ARE IN CENTIMETERS

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		UNLESS OTHERWISE SPECIFIED:		NAME	DATE	Drawn By: Tristen Hohman	
		DIMENSIONS ARE IN CENTIMETERS		DRAWN	TH	2/26/2009	
		TOLERANCES:		CHECKED		TITLE:	
		FRACTIONAL ±		ENG APPR.		Layout Of Internal Detector	
		ANGULAR: MACH ± BEND ±		MFG APPR.		And Recirculation Lines	
		TWO PLACE DECIMAL ±		Q.A.		SIZE	DWG. NO.
		THREE PLACE DECIMAL ±		COMMENTS:		B	Internal Layout
		INTERPRET GEOMETRIC TOLERANCING PER:				SCALE: 1:16	WEIGHT:
		MATERIAL					SHEET 1 OF 1
NEXT ASSY	USED ON	FINISH				REV	1
APPLICATION		DO NOT SCALE DRAWING					



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		DIMENSIONS ARE IN CENTIMETERS		DRAWN	TH	2/26/09	
		TOLERANCES:		CHECKED		TITLE:	
		FRACTIONAL ±		ENG APPR.		External Detector And Recirculation System Layout	
		ANGULAR: MACH ± BEND ±		MFG APPR.		SIZE DWG. NO. REV	
		TWO PLACE DECIMAL ±		Q.A.		B External Layout 1	
		THREE PLACE DECIMAL ±		COMMENTS:		SCALE: 1:16 WEIGHT: SHEET 1 OF 1	
		INTERPRET GEOMETRIC TOLERANCING PER:					
		MATERIAL					
NEXT ASSY	USED ON	FINISH					
APPLICATION		DO NOT SCALE DRAWING					

Past and Current Efforts

- **WARP 3.2-kg**
 - funded by INFN, construction at Pavia, operation at LNGS with participation from Princeton
 - completed 2006 - first and so far only dark matter result with argon - *Astropart. Phys.* **28**, 495 (2008)
- **WARP 140-kg [NSF PHY-0603376]**
 - funding from INFN and NSF (through Princeton)
 - installation completed; commissioning ongoing
- **Investigation of depleted argon from underground wells [NSF PHY-0704220] [Princeton U. Seed Fund]**
 - demonstrated availability of large quantity of depleted argon, ^{39}Ar reduced by >25

Short- and Long-Term Future

- **DUSEL (2013 and beyond)** [[NSF PHY-0919363](#)]
 - MAX - Multi-ton Argon and Xenon
 - 20 institutions, 5-t DAr and 2.5-t Xe TPCs
 - NSF announced intention to fund S4 proposal
- **Depleted Argon Production** [[NSF PHY-0811186](#)]
 - funded from NSF through Princeton
 - 2 kg/day, depletion >25 , by end of 2009
- **DarkSide Detector (2009-2012)**
 - first depleted argon TPC
 - demonstration of key features
 - competitive dark matter search

DarkSide



UMass Amherst
Arizona State University
Augustana College
Black Hills State University
Fermilab
University of Houston
University of Notre Dame
Princeton University
Temple University
UCLA

Like the jelly beans in this jar, the Universe is mostly dark: 96 percent consists of dark energy (about 70%) and dark matter (about 26%). Only about four percent (the same proportion as the lightly colored jelly beans) of the Universe - including the stars, planets and us - is made of familiar atomic matter.

The End

