

KamLAND Results

Week Interactions and Neutrinos (WIN'09)

Sep. 15, 2008

I. Shimizu (Tohoku Univ.)

KamLAND Collaboration

S. Abe,¹ T. Ebihara,¹ S. Enomoto,¹ K. Furuno,¹ Y. Gando,¹ K. Ichimura,¹ H. Ikeda,¹ K. Inoue,¹ Y. Kibe,¹ Y. Kishimoto,¹ M. Koga,¹ A. Kozlov,¹ Y. Minekawa,¹ T. Mitsui,¹ K. Nakajima,¹, * K. Nakajima,¹ K. Nakamura,¹ M. Nakamura,¹ K. Owada,¹ I. Shimizu,¹ Y. Shimizu,¹ J. Shirai,¹ F. Suekane,¹ A. Suzuki,¹ Y. Takemoto,¹ K. Tamae,¹ A. Terashima,¹ H. Watanabe,¹ E. Yonezawa,¹ S. Yoshida,¹ J. Busenitz,² T. Classen,² C. Grant,² G. Keefer,² D.S. Leonard,² D. McKee,² A. Piepke,² M.P. Decowski,³ J.A. Detwiler,³ S.J. Freedman,³ B.K. Fujikawa,³ F. Gray,³, † E. Guardincerri,³ L. Hsu,³, ‡ R. Kadel,³ C. Lendvai,³ K.-B. Luk,³ H. Murayama,³ T. O'Donnell,³ H.M. Steiner,³ L.A. Winslow,³ D.A. Dwyer,⁴ C. Jillings,⁴, Å C. Mauger,⁴ R.D. McKeown,⁴ P. Vogel,⁴ C. Zhang,⁴ B.E. Berger,⁵ C.E. Lane,⁶ J. Maricic,⁶ T. Miletic,⁶ M. Batygov,⁷ J.G. Learned,⁷ S. Matsuno,⁷ S. Pakvasa,⁷ J. Foster,⁸ G.A. Horton-Smith,⁸ A. Tang,⁸ S. Dazeley,⁹, Å K.E. Downum,¹⁰ G. Gratta,¹⁰ K. Tolich,¹⁰ W. Bugg,¹¹ Y. Efremenko,¹¹ Y. Kamyshev,¹¹ O. Perevozchikov,¹¹ H.J. Karwowski,¹² D.M. Markoff,¹² W. Tornow,¹² K.M. Heeger,¹³ F. Piquemal,¹⁴ and J.-S. Ricol¹⁴

(The KamLAND Collaboration)



¹Research Center for Neutrino Science, Tohoku University, Sendai 980-8578, Japan

²Department of Physics and Astronomy, University of Alabama, Tuscaloosa, Alabama 35487, USA

³Physics Department, University of California, Berkeley and

Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA

⁴W. K. Kellogg Radiation Laboratory, California Institute of Technology, Pasadena, California 91125, USA

⁵Department of Physics, Colorado State University, Fort Collins, Colorado 80523, USA

⁶Physics Department, Drexel University, Philadelphia, Pennsylvania 19104, USA

⁷Department of Physics and Astronomy, University of Hawaii at Manoa, Honolulu, Hawaii 96822, USA

⁸Department of Physics, Kansas State University, Manhattan, Kansas 66506, USA

⁹Department of Physics and Astronomy, Louisiana State University, Baton Rouge, Louisiana 70803, USA

¹⁰Physics Department, Stanford University, Stanford, California 94305, USA

¹¹Department of Physics and Astronomy, University of Tennessee, Knoxville, Tennessee 37996, USA

¹²Triangle Universities Nuclear Laboratory, Durham, North Carolina 27708, USA and

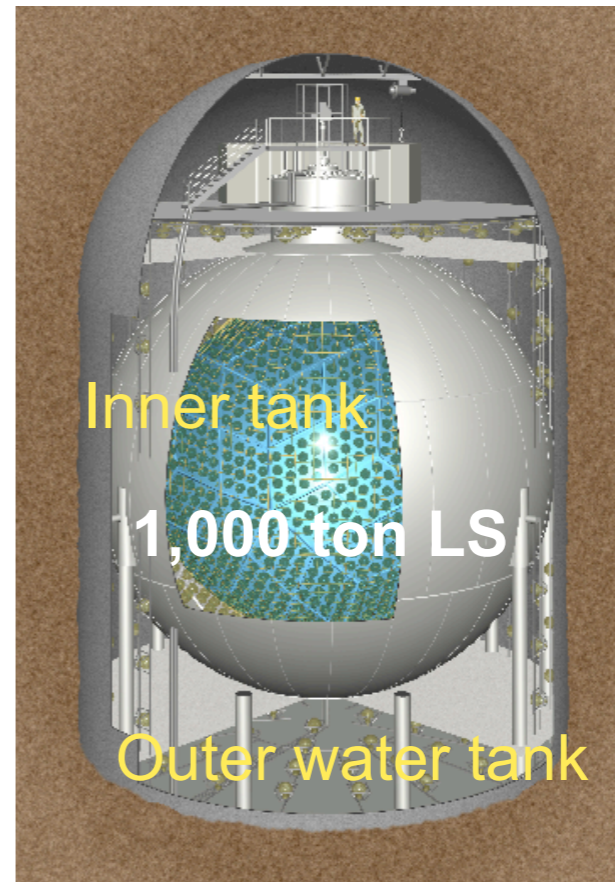
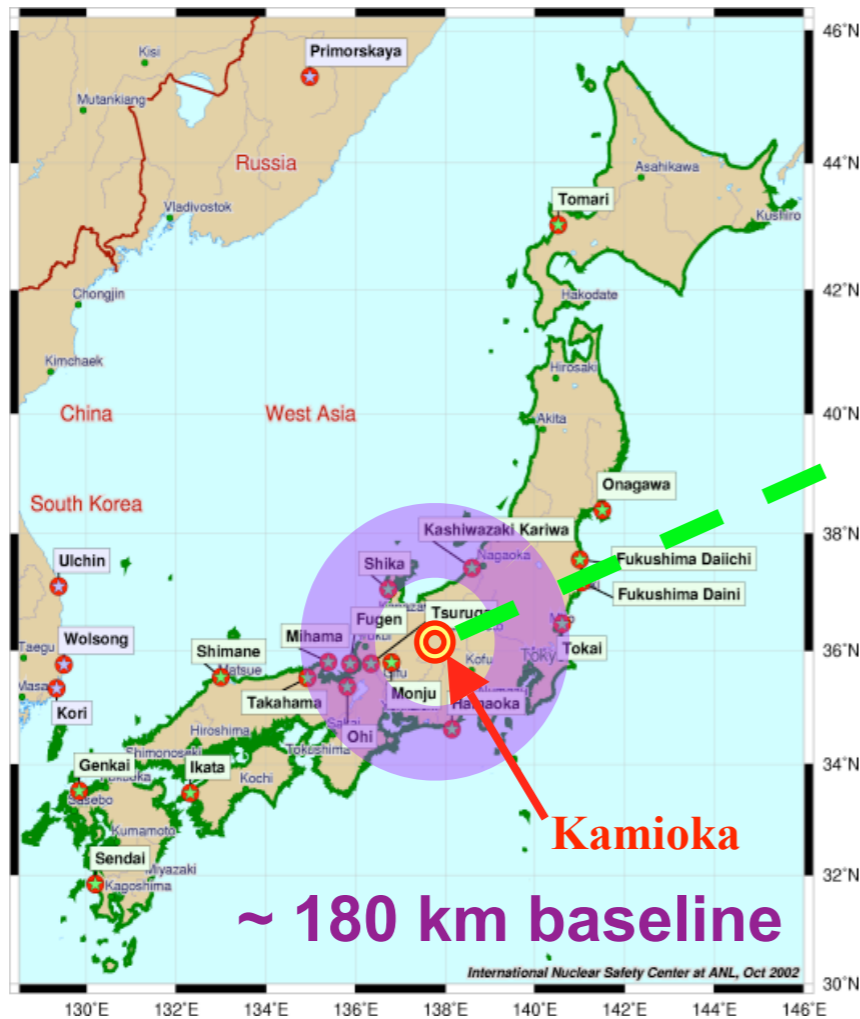
Physics Departments at Duke University, North Carolina Central University, and the University of North Carolina at Chapel Hill

¹³Department of Physics, University of Wisconsin, Madison, Wisconsin 53706, USA

¹⁴CEN Bordeaux-Gradignan, IN2P3-CNRS and University Bordeaux I, F-33175 Gradignan Cedex, France

KamLAND

Kamioka Liquid Scintillator Anti-Neutrino Detector



34% photo-coverage with
1325 17" and 554 20" PMTs

Reactor neutrino : sensitive to LMA solution

1st result (Dec. 2002)

2nd result (Jun. 2004)

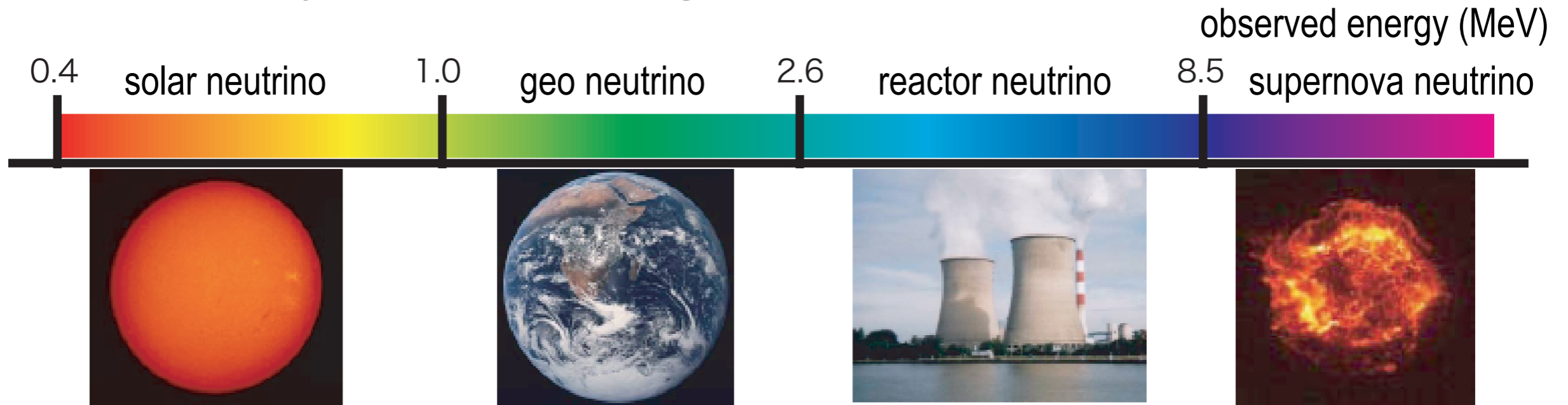
3rd result (Jan. 2008)

disappearance of $\bar{\nu}_e$

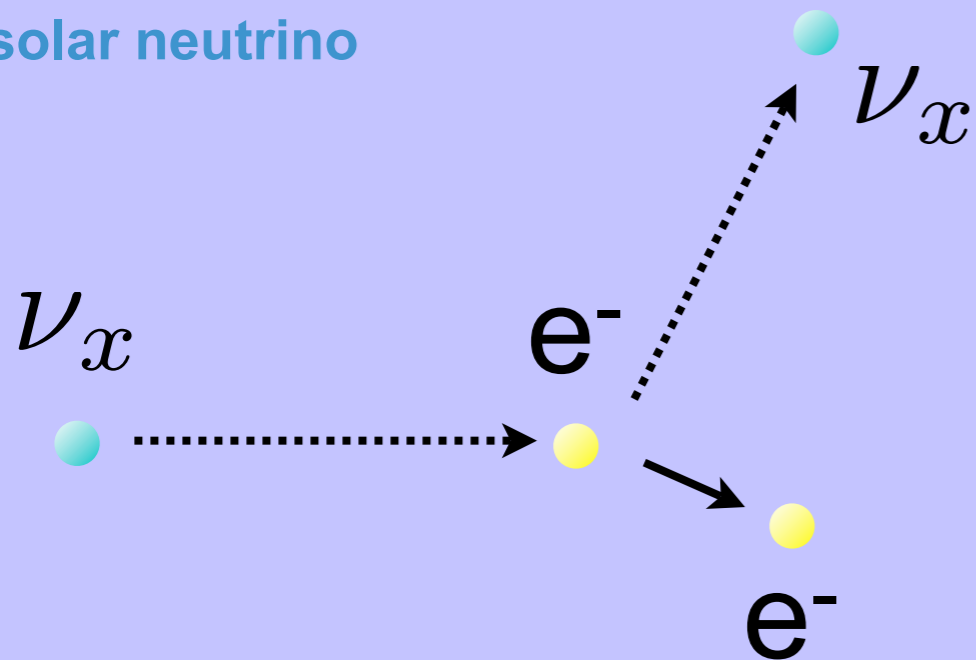
spectral distortion of $\bar{\nu}_e$

precise measurement of oscillation parameters

Physics Target in KamLAND

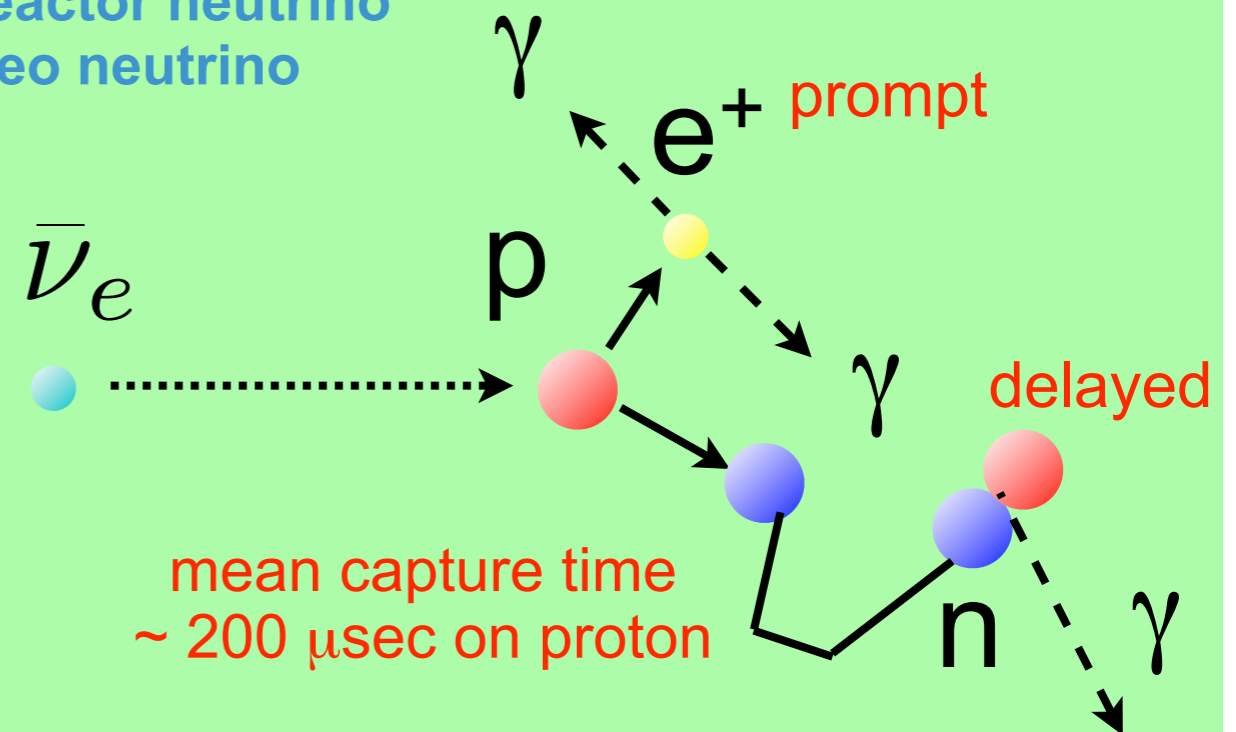


solar neutrino



neutrino detection by electron scattering

reactor neutrino geo neutrino



anti-neutrino detection by inverse beta-decay

Reactor and Geo Neutrino Analysis

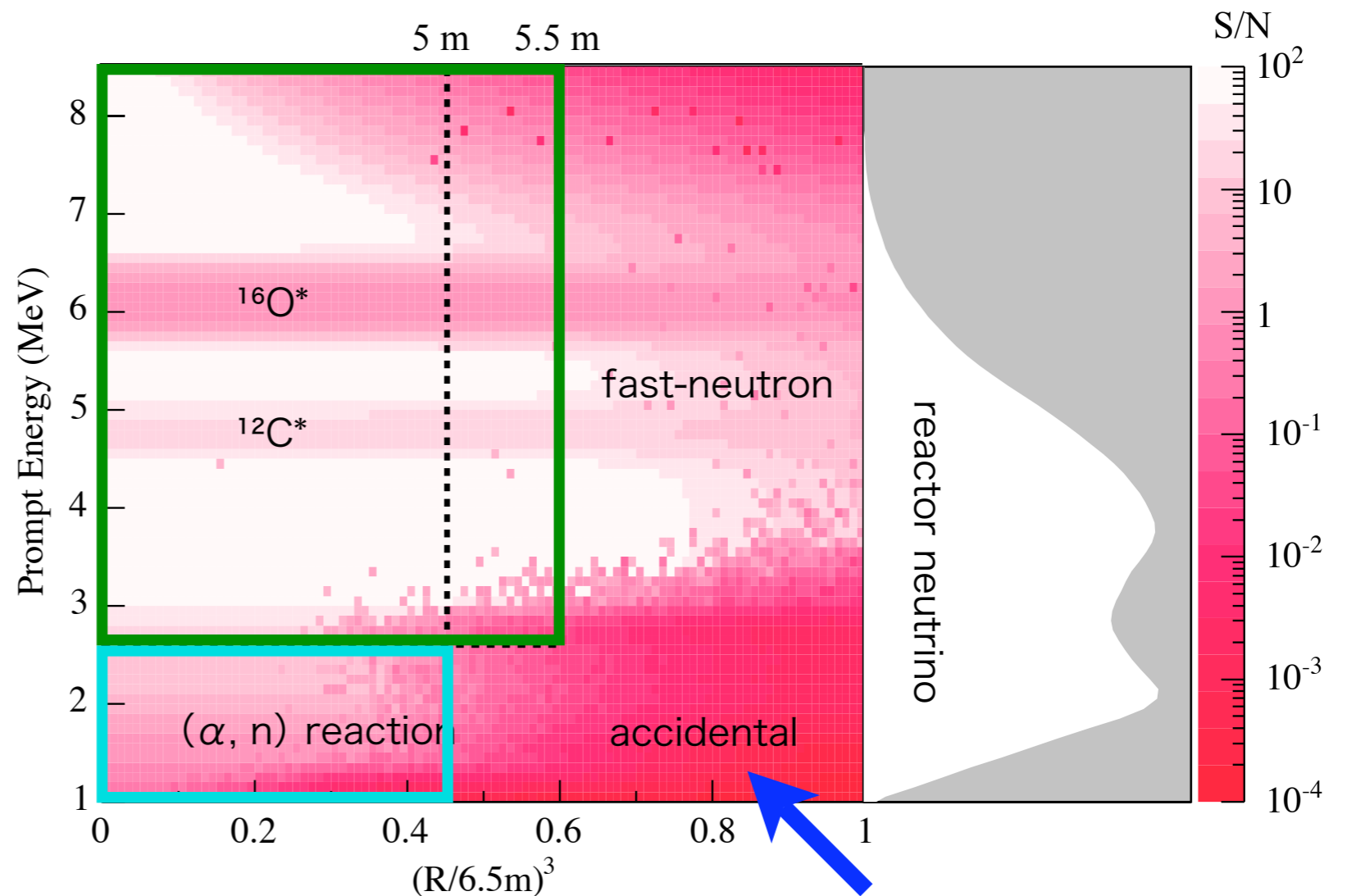
previous result

S / B ratio map (energy v.s. radius)

separated analysis window for reactor and geo neutrinos

reactor neutrino
(2.6 - 8.5 MeV, R 5.5 m)

geo neutrino
(0.9 - 2.6 MeV, R 5.0 m)



large accidental B.G.
caused by external γ -rays

Analysis improvement

- (1) efficient **accidental** background rejection
- (2) combined analysis of **reactor** and **geo** neutrinos

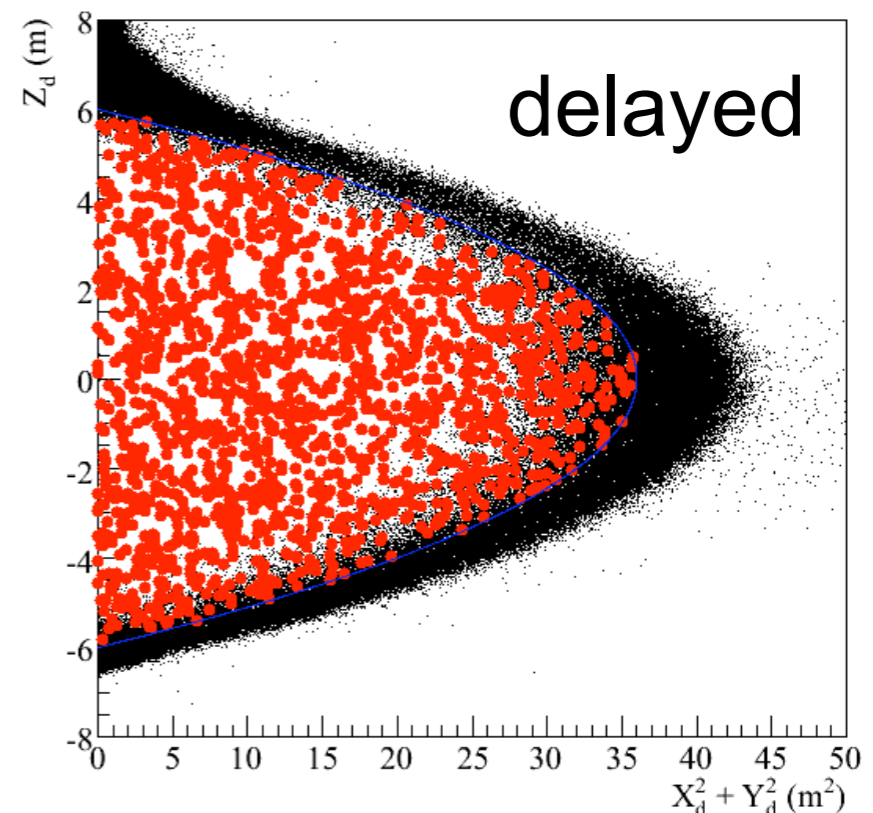
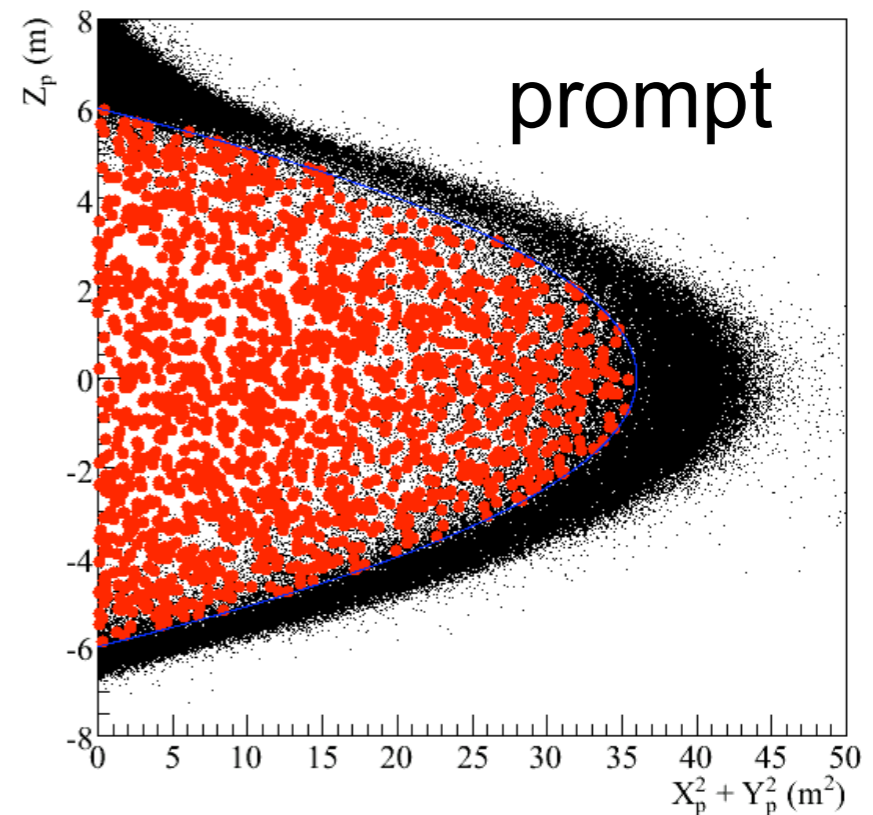
Anti-Neutrino Event Selection

(a) accidental B.G. discrimination

- $0.5 < \Delta T < 1000 \mu\text{s}$
- $\Delta R < 2 \text{ m}$
- $1.8 \text{ MeV} < E_{\text{delayed}} < 2.6 \text{ MeV}$ or
 $4.0 \text{ MeV} < E_{\text{delayed}} < 5.8 \text{ MeV}$
- $0.9 \text{ MeV} < E_{\text{prompt}} < 8.5 \text{ MeV}$
- $R_{\text{prompt}}, R_{\text{delayed}} < 6.0 \text{ m}$
- L-selection from 6 parameters

(b) μ spallation cut

- $\Delta T_{\mu} > 2 \text{ s}$ after showing μ
- $\Delta T_{\mu} > 2 \text{ s}$ or $\Delta L > 3 \text{ m}$ after
non-showering μ ($\Delta Q < 10^6 \text{ p.e.}$)



Likelihood Selection

L-selection for accidental B.G. discrimination

Accidentals PDF $f_{acc}(E_p, E_d, \Delta R, \Delta T, R_p, R_d)$

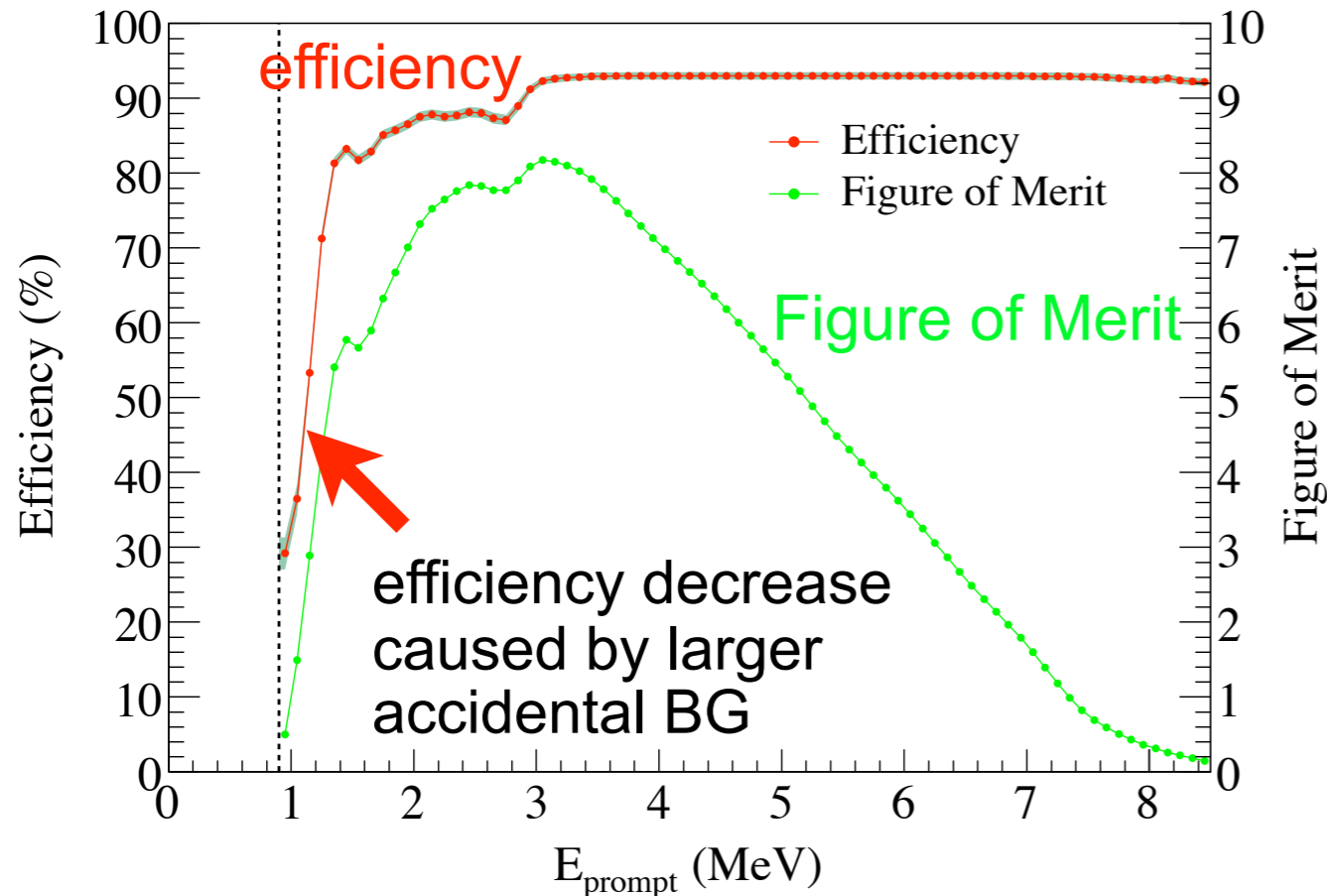
Signal PDF $f_{\bar{\nu}_e}(E_p, E_d, \Delta R, \Delta T, R_p, R_d)$

L-selector $L = \frac{f_{\bar{\nu}_e}}{f_{\bar{\nu}_e} + f_{acc}}$

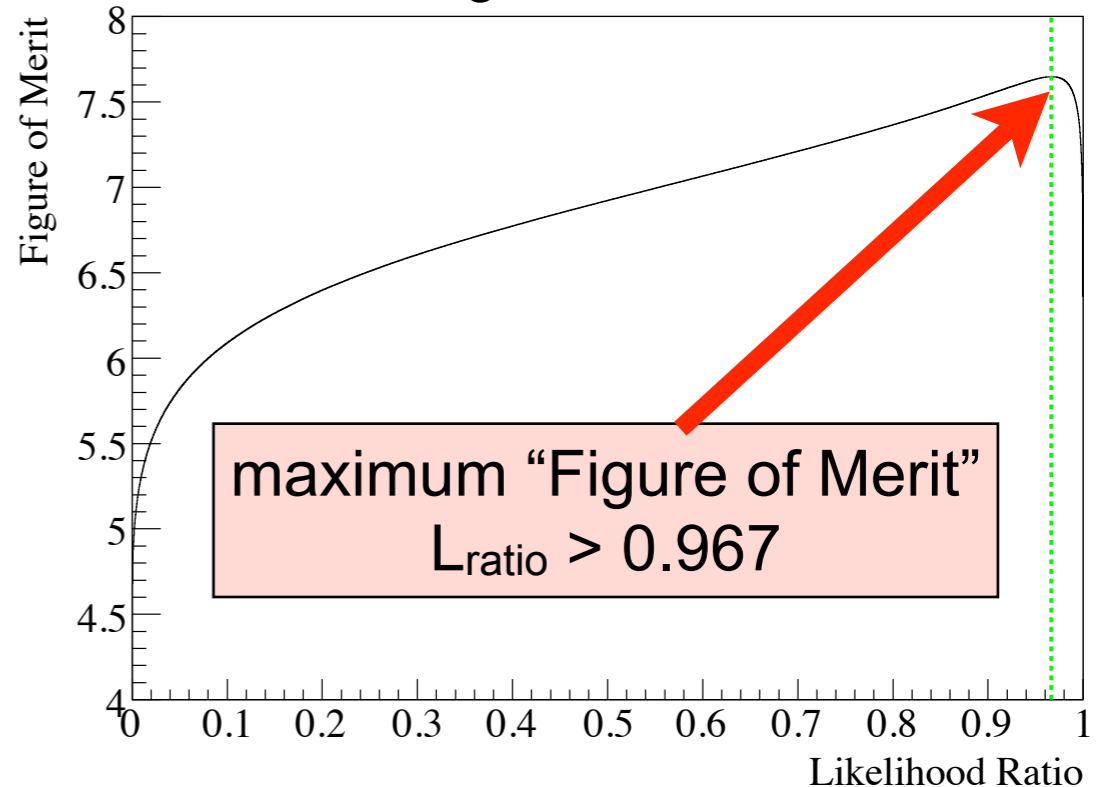
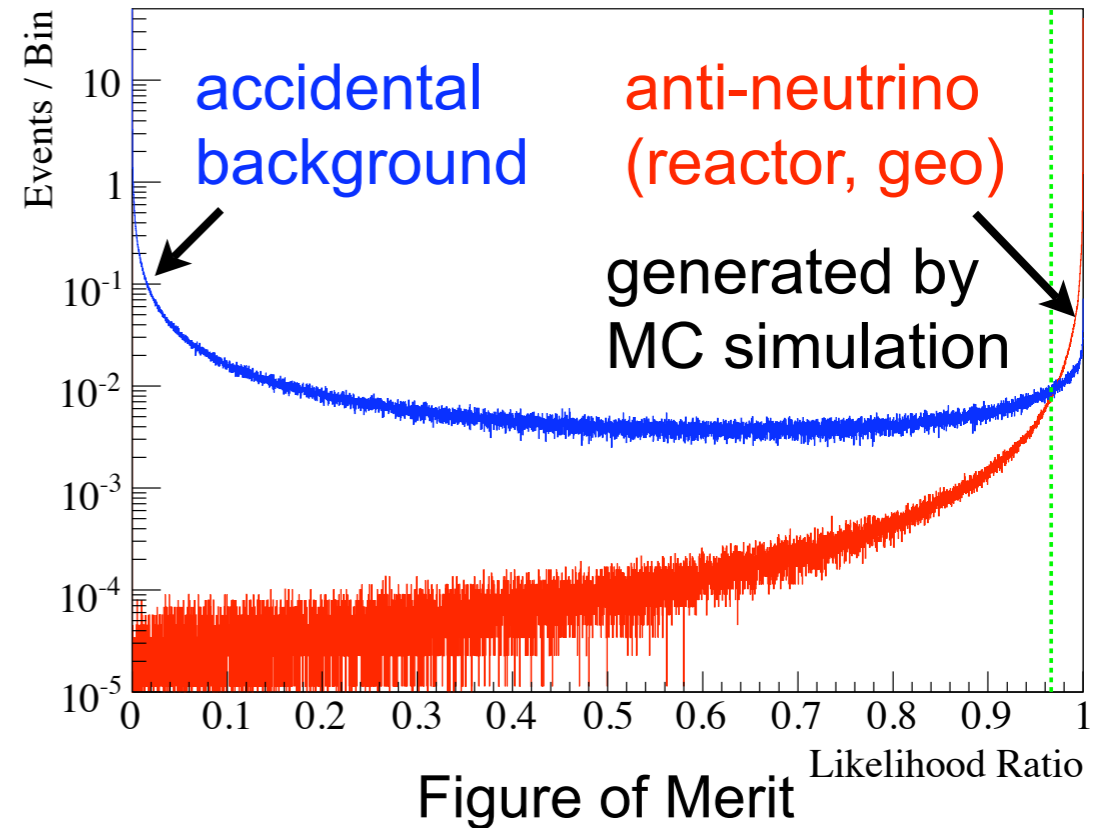
Maximize
"Figure of Merit"
for each E_p bin

$$FOM = \frac{S}{\sqrt{S + B_{acc}}}$$

Detection efficiency



$2.2 < E_{prompt} < 2.3$ MeV



Systematic Uncertainty

“full volume” calibration lowered the fiducial volume error

(4.7% in previous analysis)

Detector related

Reactor related

Fiducial volume

1.8%

$\bar{\nu}_e$ spectra

2.4%

Energy scale

1.5%

Reactor power

2.1%

L-selection eff.

0.6%

Fuel composition

1.0%

OD veto

0.2%

Long-lived nuclei

0.3%

Cross section

0.2%

Time lag

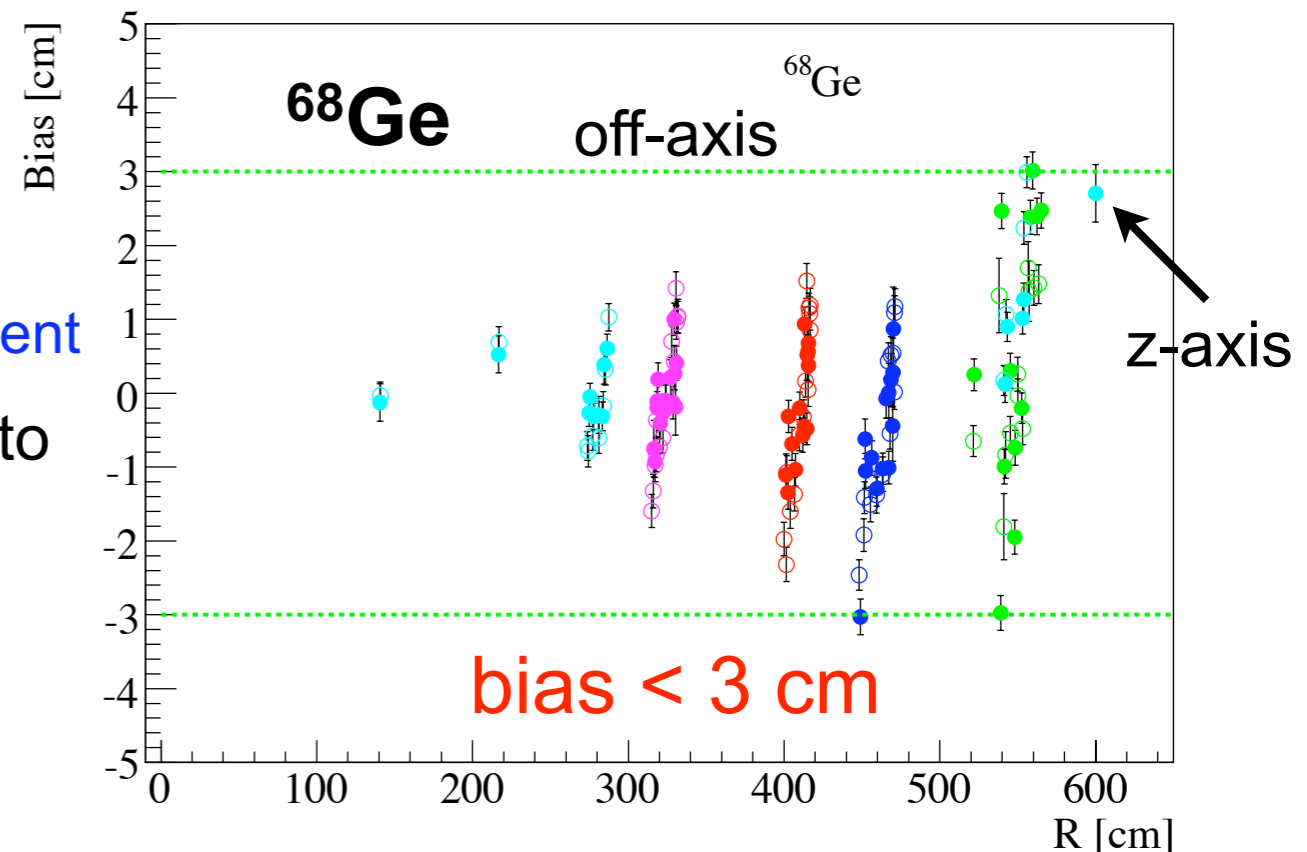
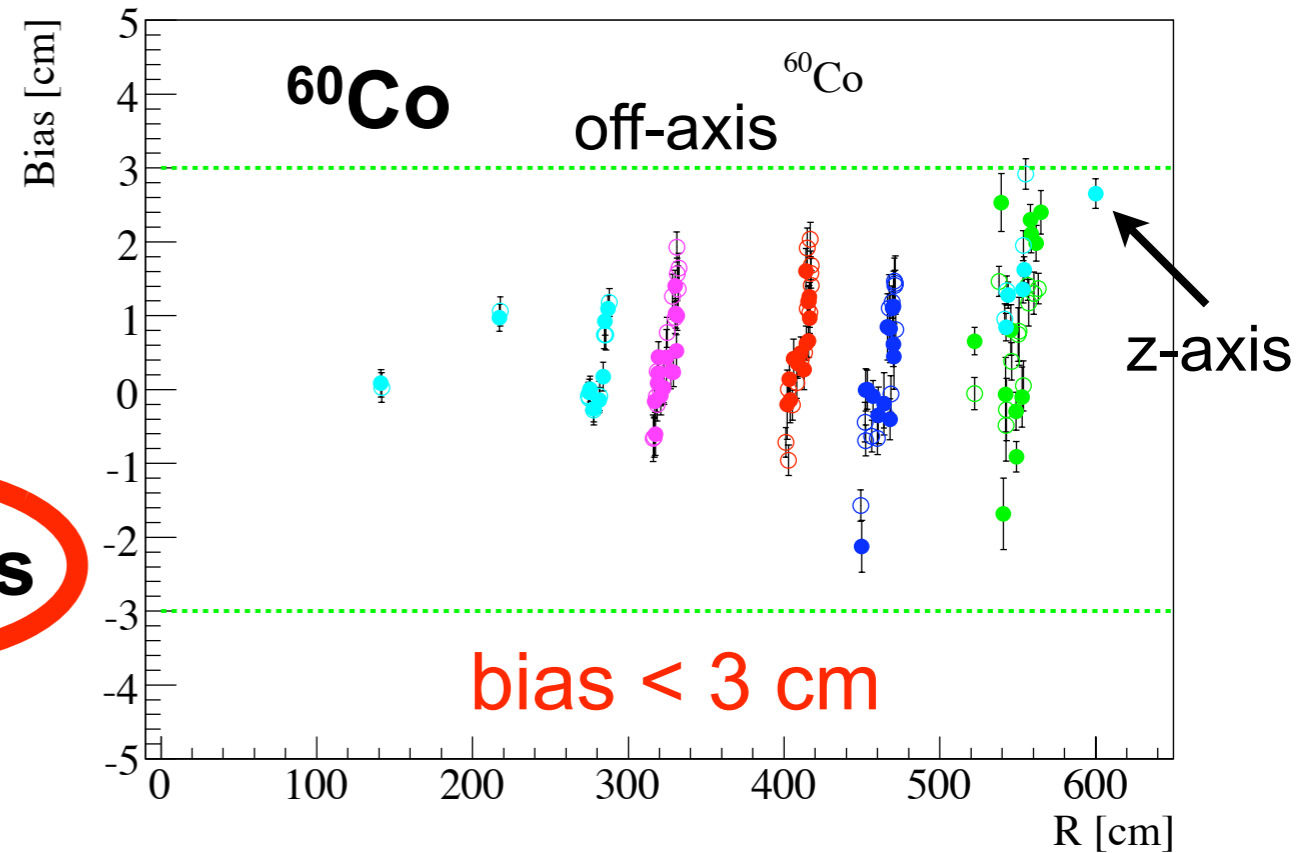
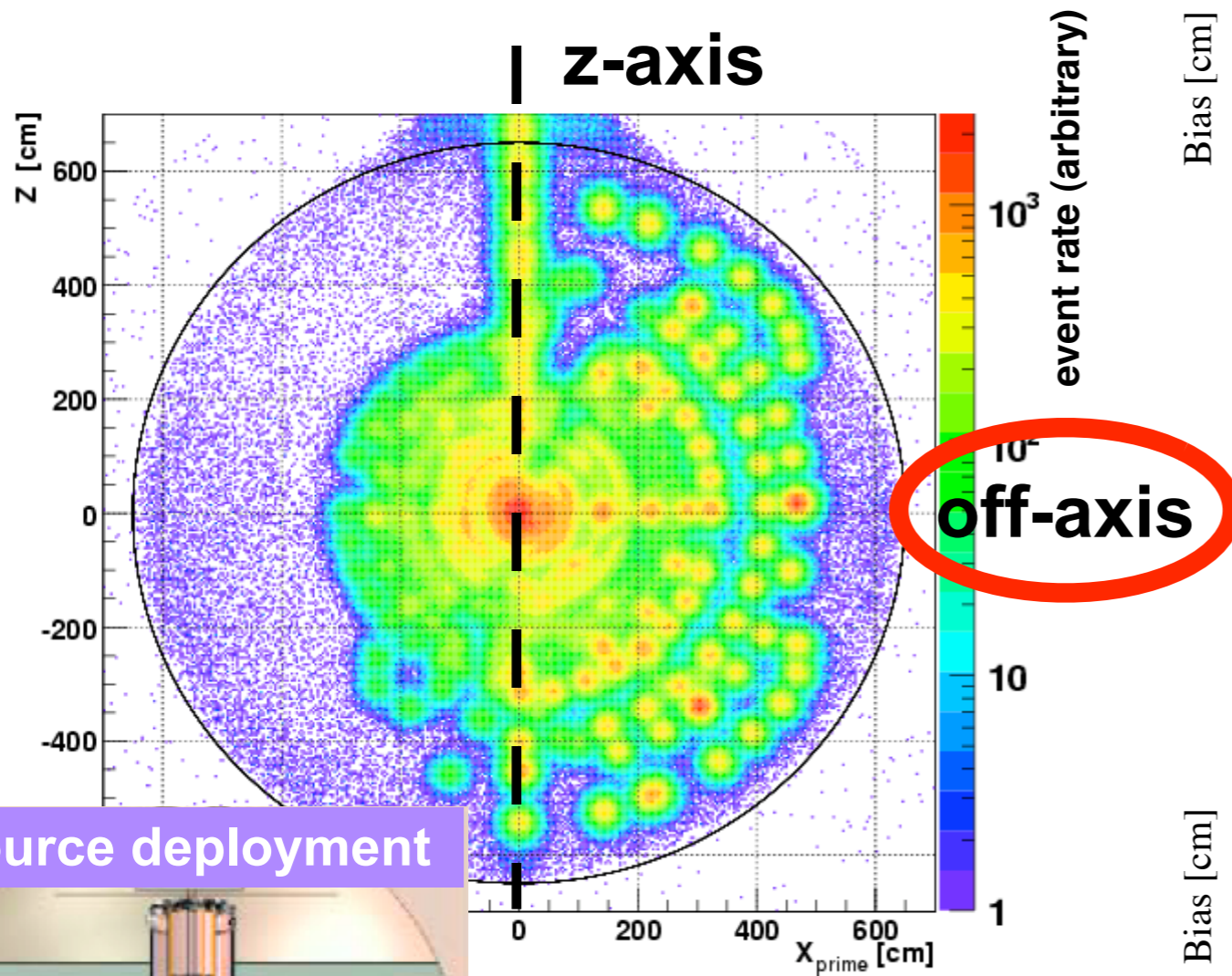
0.01%

2.4%

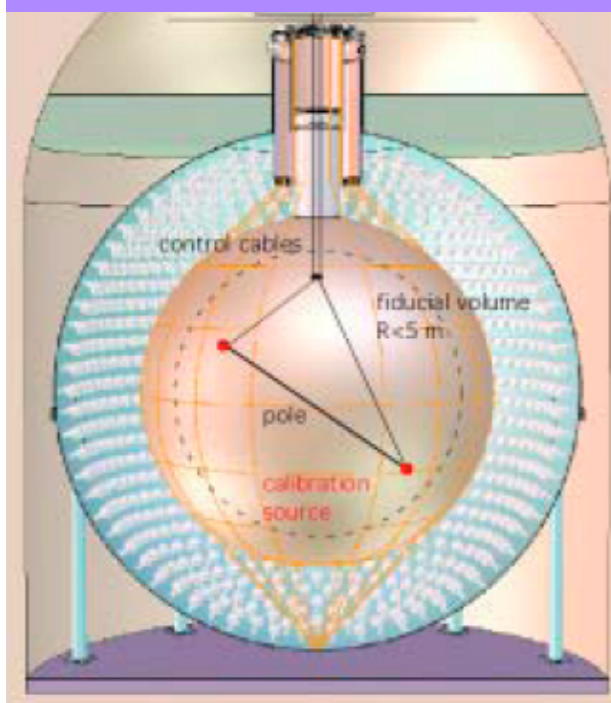
3.4%

Total systematic uncertainty : 4.1%

Full Volume Calibration



source deployment

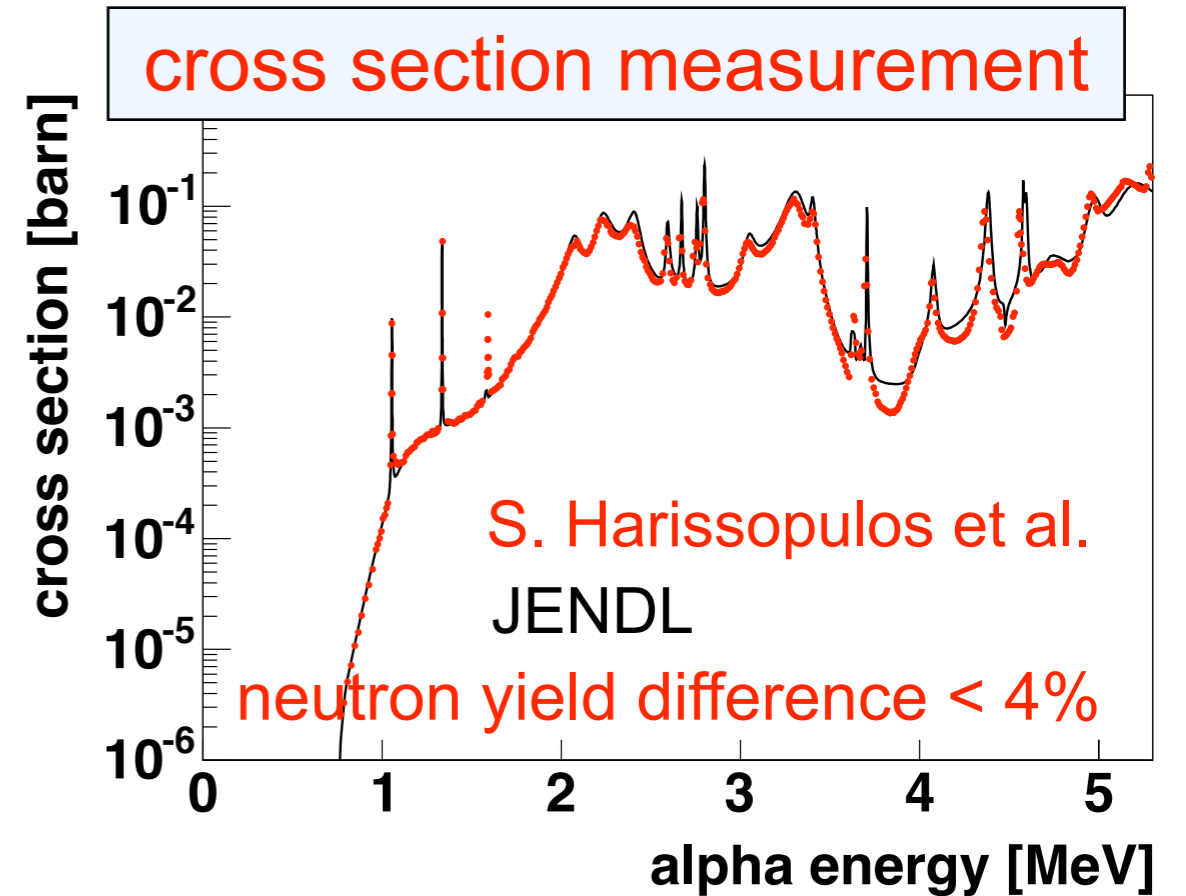
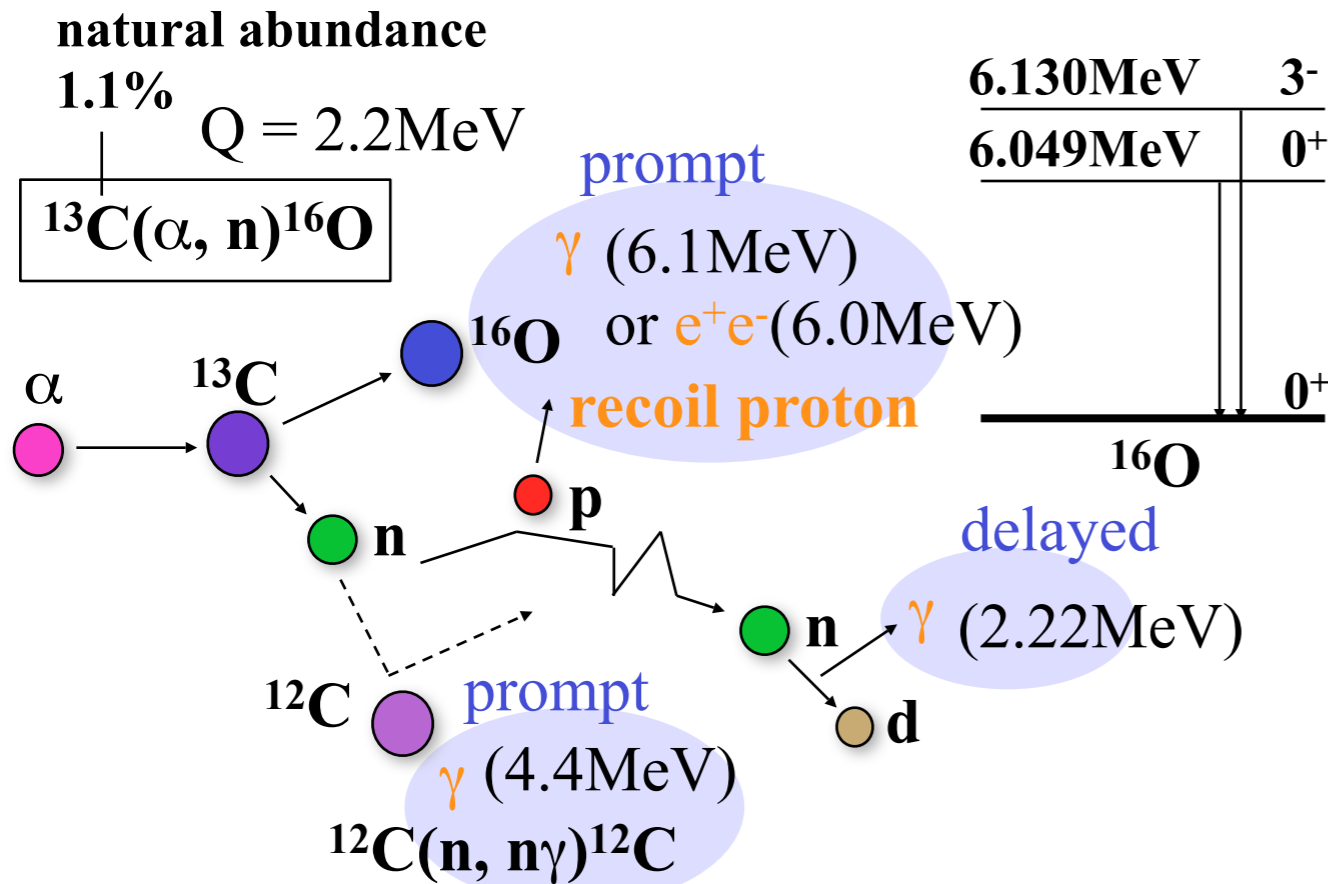


“4pi calibration” system for the off-axis source deployment

bias < 3 cm corresponds to 1.8% volume uncertainty

cross-checked by ¹²B/¹²N uniformity

(α , n) Background Estimation



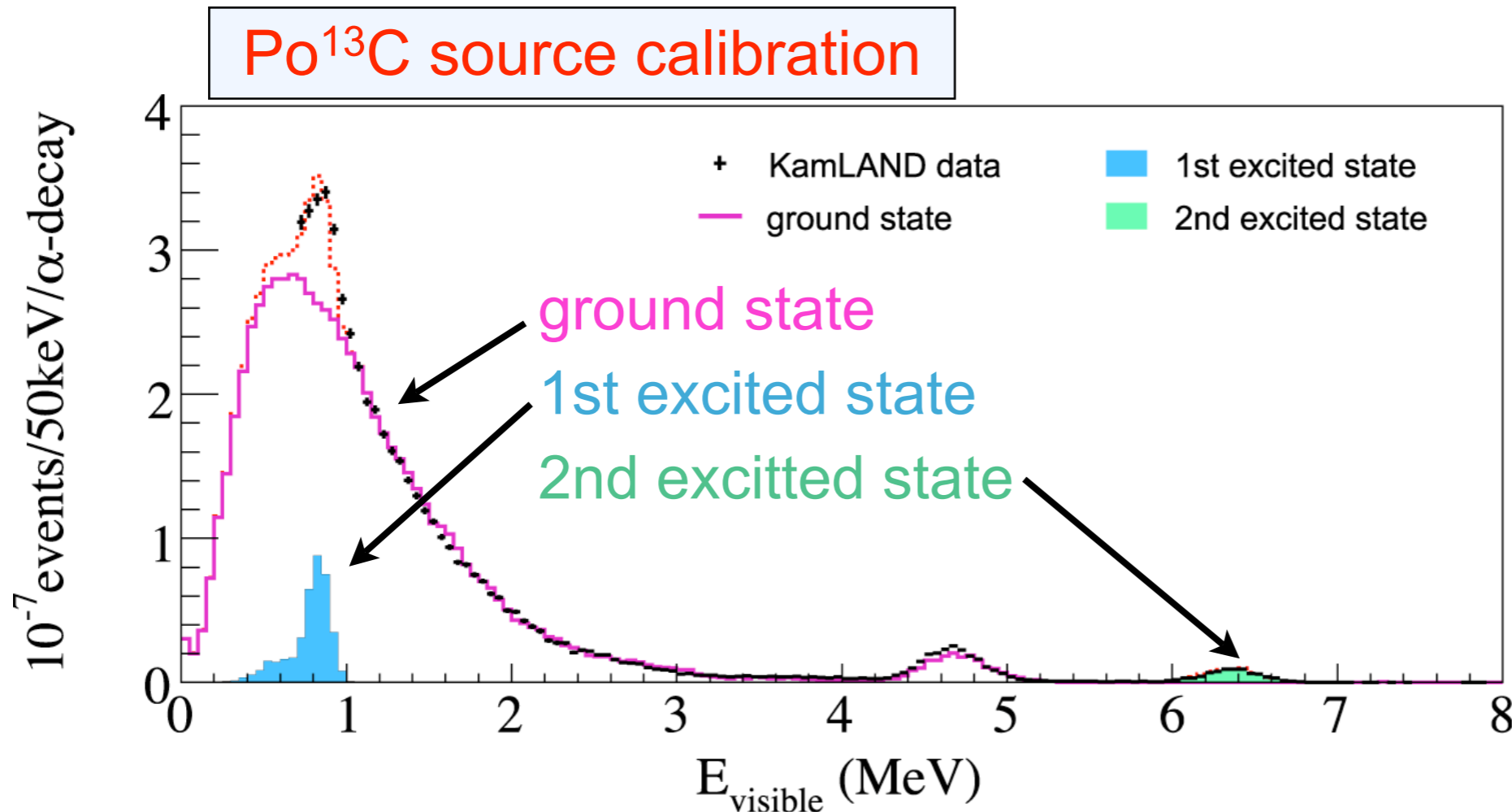
total cross section was determined precisely

1. low energy	$^{13}\text{C}(\alpha, n)^{16}\text{O}$ (g.s.)	n
2. 4.4MeV	$^{13}\text{C}(\alpha, n)^{16}\text{O}$ (g.s.) $\rightarrow ^{12}\text{C}(\text{n}, \text{n}\gamma)^{12}\text{C}$	n $\gamma + \text{n}$
3. 6MeV	$^{13}\text{C}(\alpha, n)^{16}\text{O}^*$ (1st e.s. 6.049MeV)	e^+e^-
	$^{13}\text{C}(\alpha, n)^{16}\text{O}^*$ (2nd e.s. 6.130MeV)	$\gamma + \text{n}$

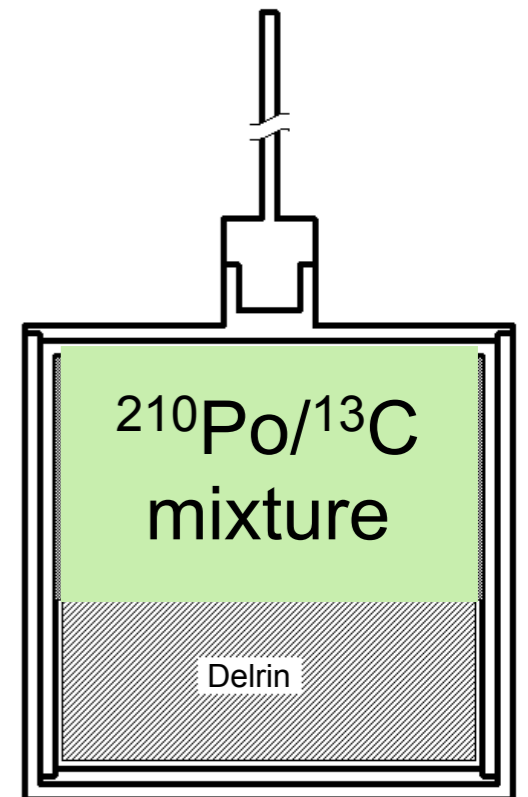
Cross section for each branch should be measured

Cross Section Measurement

direct measurement of $^{13}\text{C}(\alpha, n)^{16}\text{O}$ reaction in KamLAND



Po¹³C source



(α, n) background estimation

163.3 ± 18.0 events for ground state

18.7 ± 3.7 events for excited state

Estimation uncertainty

11% for ground state

20% for excited state

Rate Analysis above 2.6 MeV

“Reactor” rate analysis
(2.6 MeV threshold)

No osci. expected	1554
Background	63
Observed events	985

Ratio = (obs. - B.G.) / No osci.
 $0.593 \pm 0.020(\text{stat}) \pm 0.026(\text{syst})$

8.5 σ disappearance significance

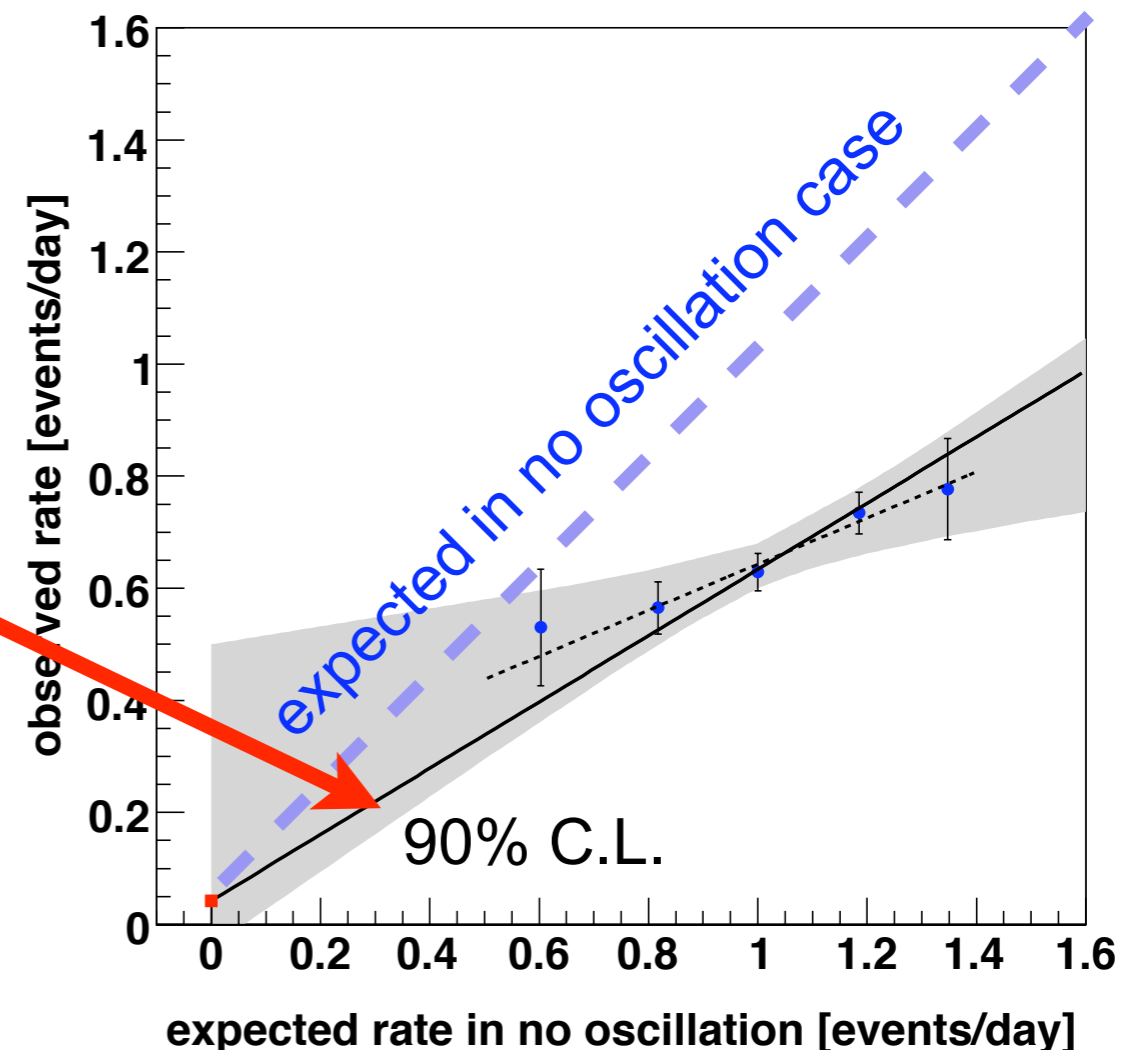
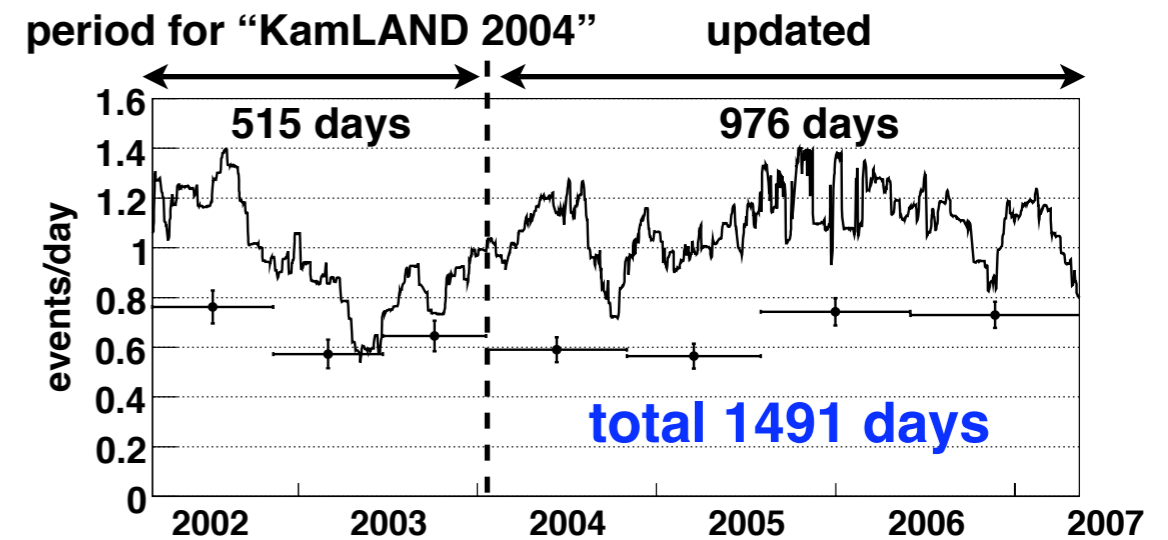
Fit constrained through B.G. expected

$$\chi^2 / \text{ndf} = 2.8 / 4$$

Fit with a horizontal line

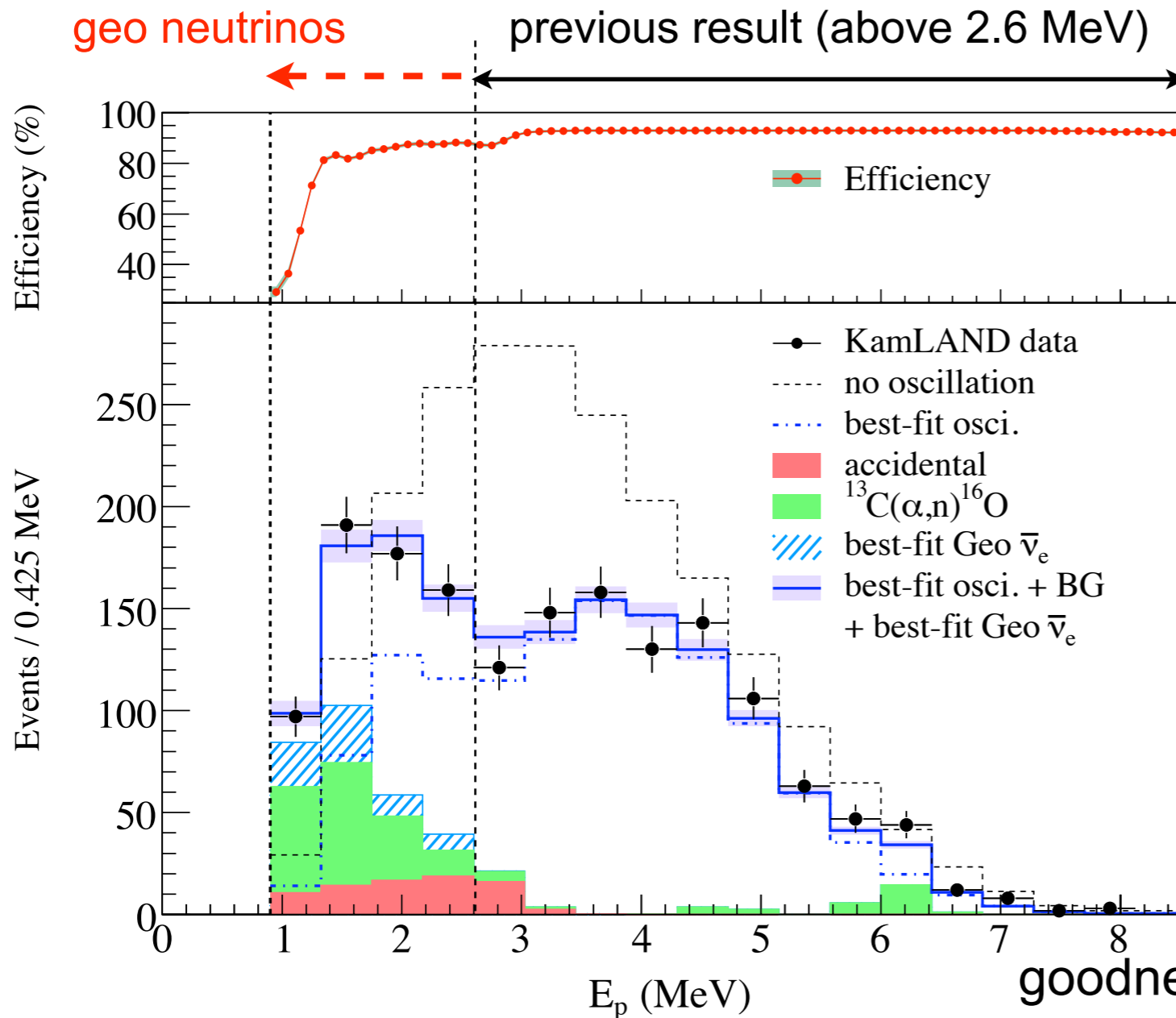
$$\chi^2 / \text{ndf} = 12.0 / 4$$

(1.7% C.L.)



Energy Spectrum above 0.9 MeV

exposure : 2881 ton-year (3.8 × 766 ton-year for “KamLAND 2004”)



“Geo + Reactor”
combined analysis

No osci. expected 2179

Background
(w/o geo neutrino) 276

Observed events 1609

best-fit

$(\tan^2\theta, \Delta m^2)$
 $= (0.56, 7.58 \times 10^{-5} \text{ eV}^2)$

free parameter : geo neutrinos
 (U, Th) = (37.1, 30.2) events

goodness of fit using equal probability bins

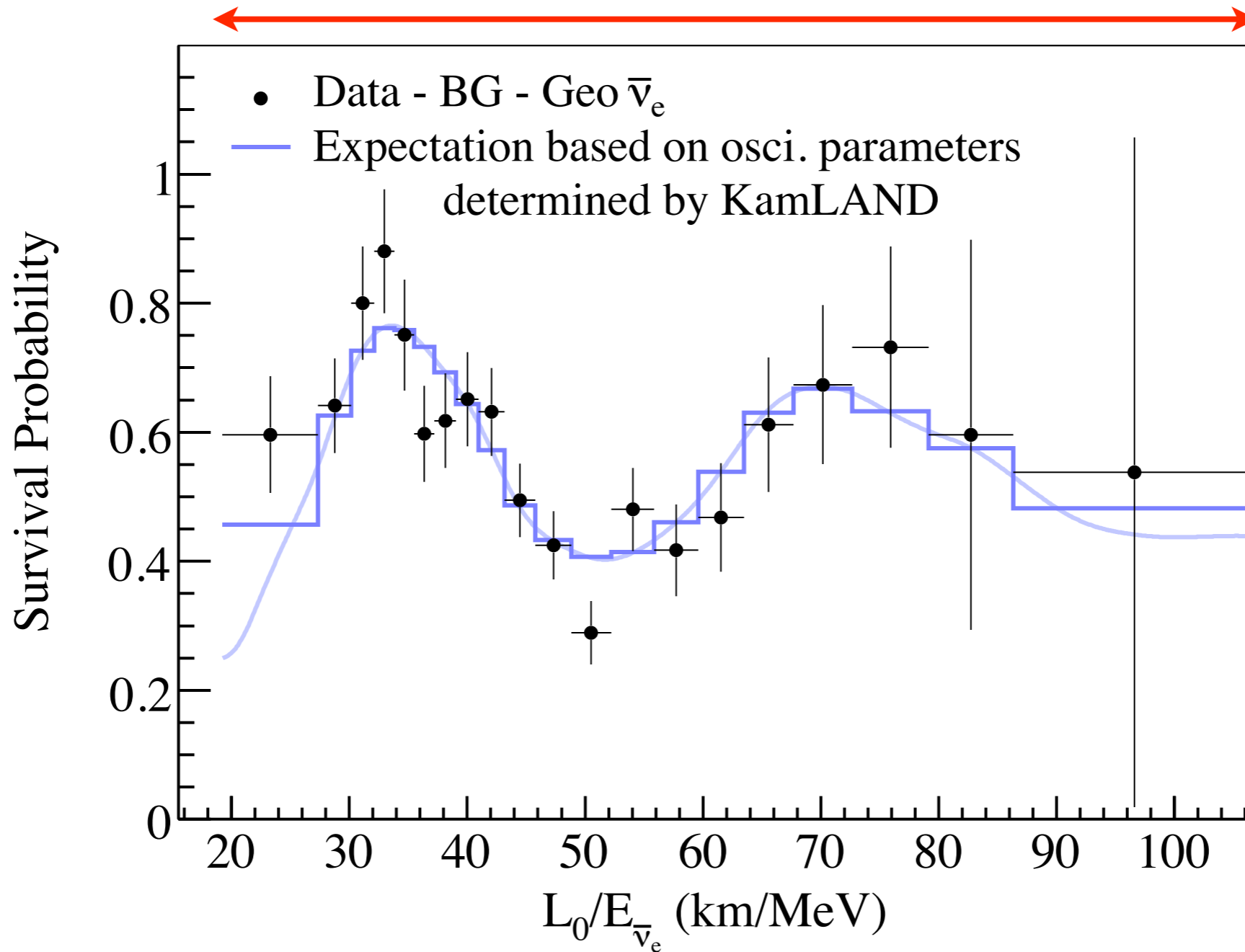
best-fit $\chi^2 / \text{ndf} = 20.9 / 16$ (18.4% C.L.)

no osci. $\chi^2 / \text{ndf} = 63.6 / 17$

Scaled no oscillation spectrum is excluded at 5.1 σ

L/E Plot

this result (above 0.9 MeV)



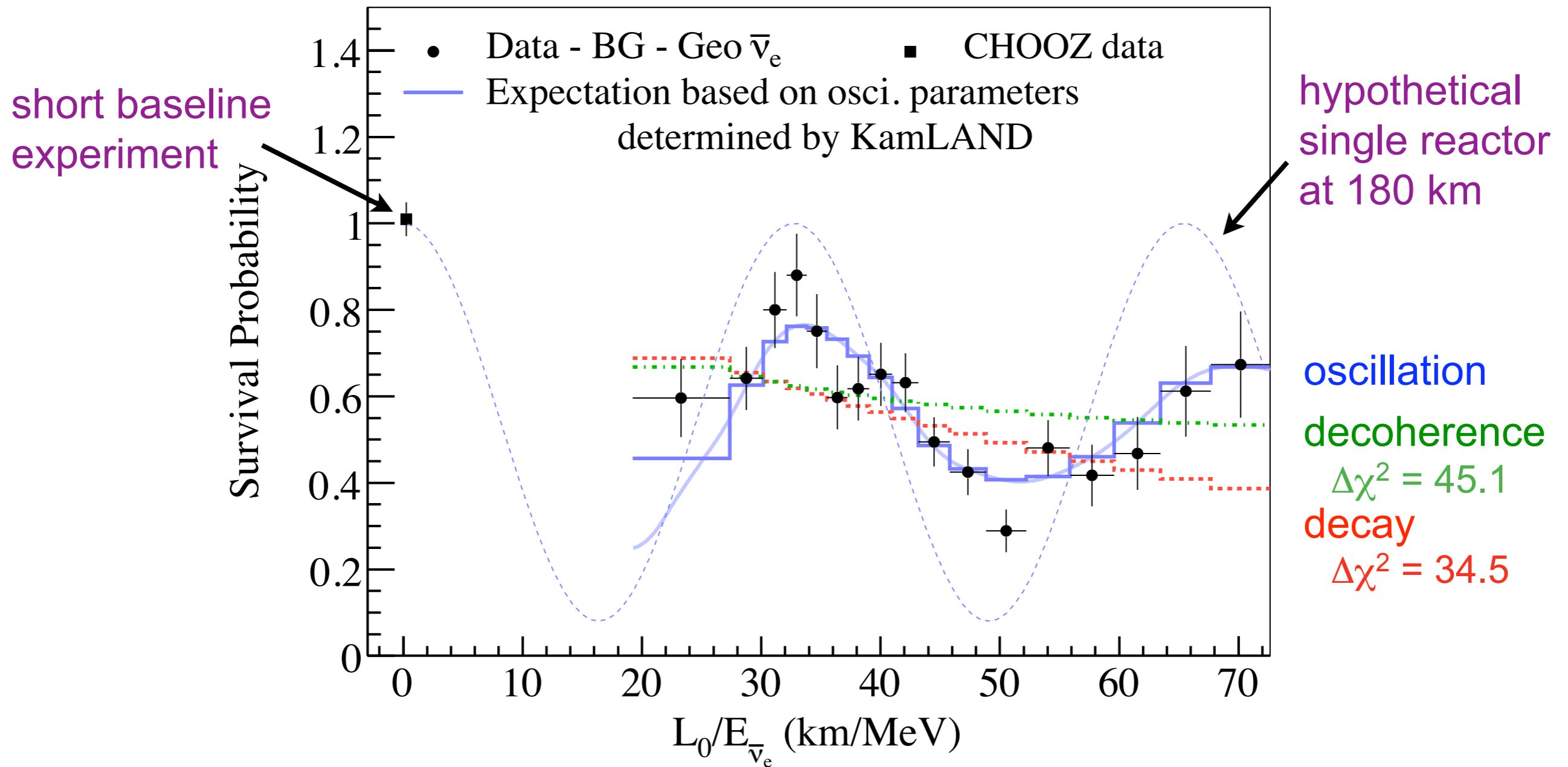
~ 2 cycle of oscillation

strong evidence of neutrino oscillation

Alternate Hypothesis

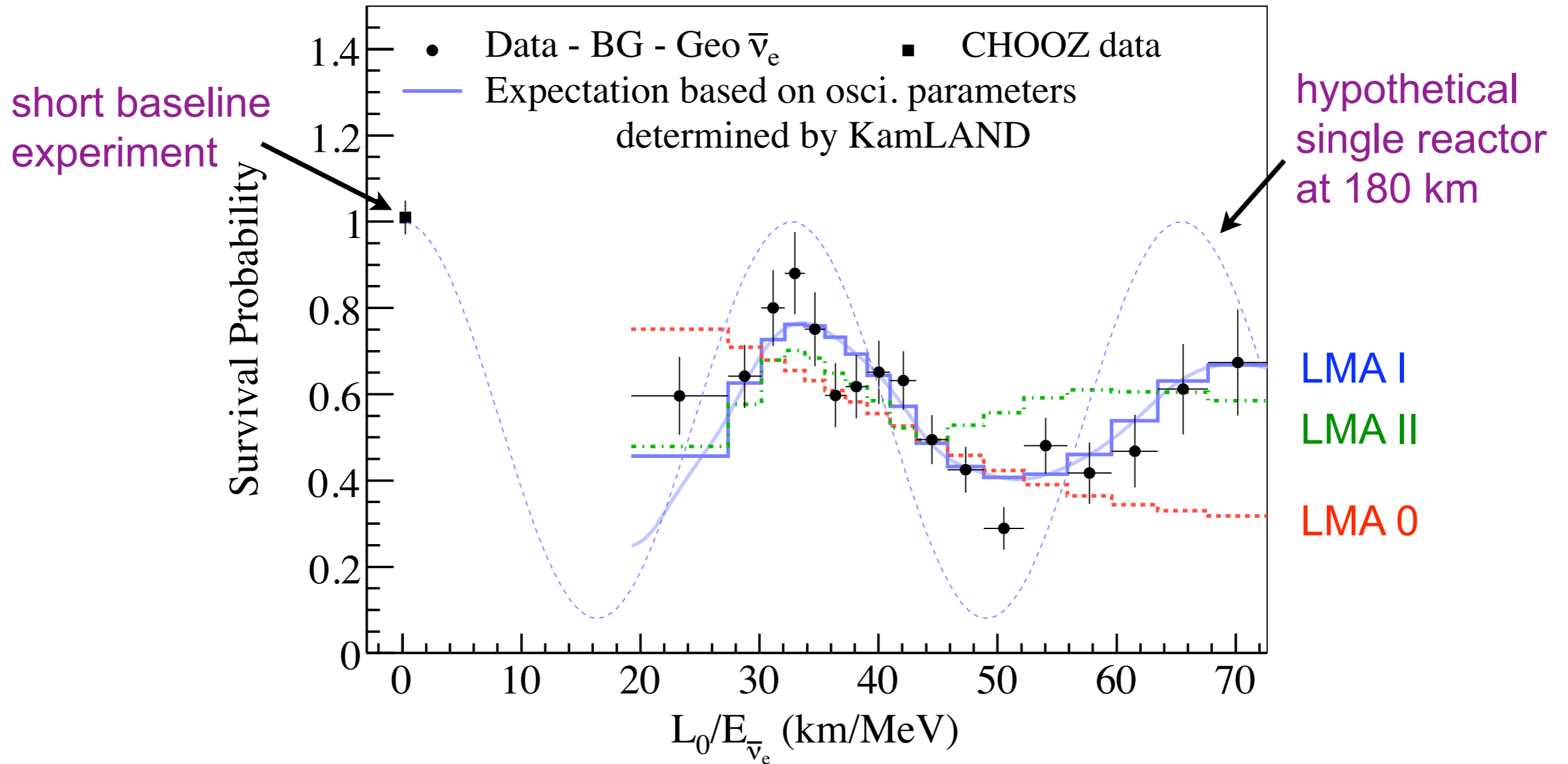
V. D. Barger et al., Phys. Rev. Lett. 82, 2640 (1999)

E. Lisi et al, Phys. Rev. Lett. 85, 1166 (2000)



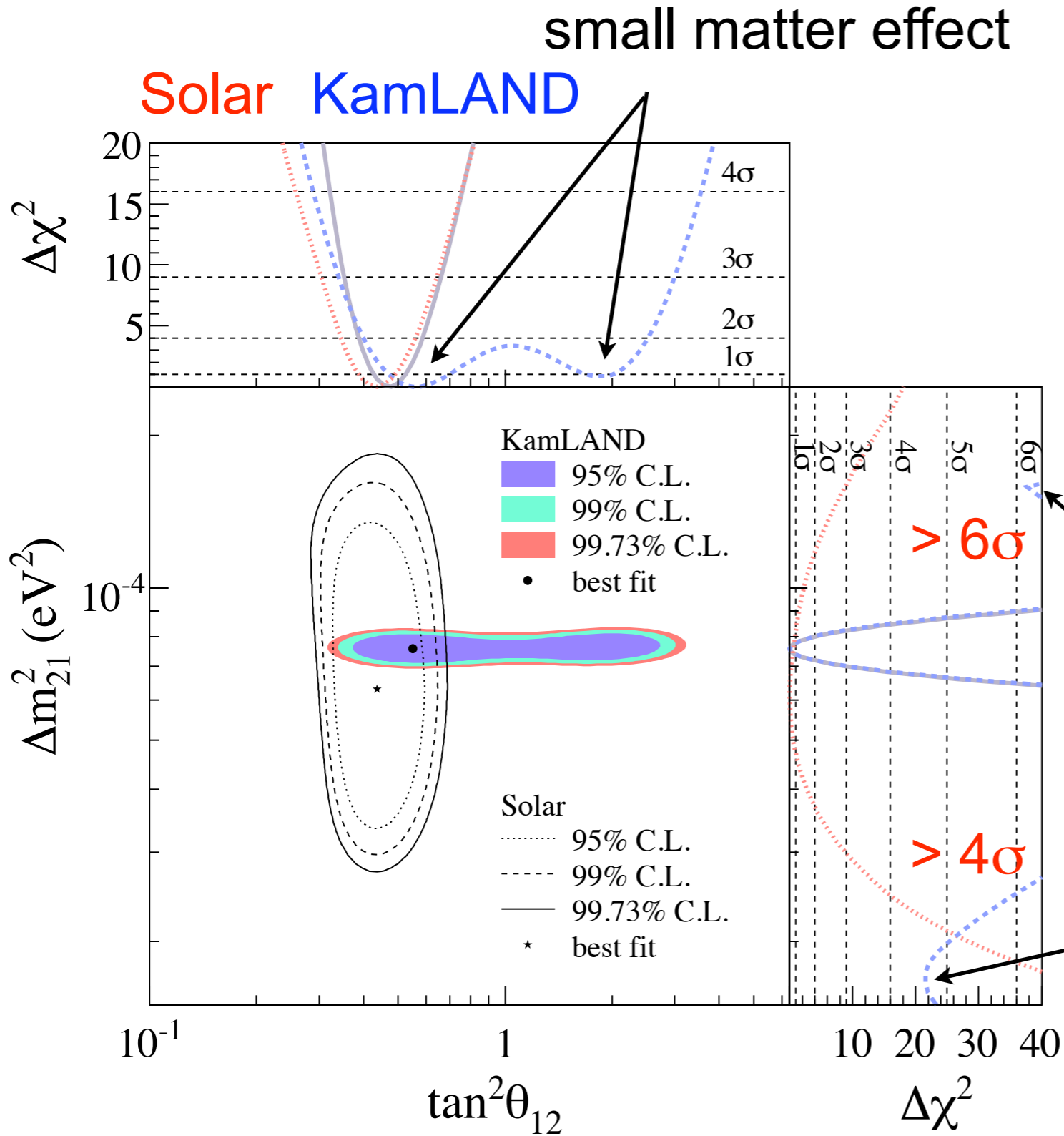
best model is neutrino oscillation

Alternate Wavelength



LMA 0 and LMA II are disfavored at more than 4σ

Oscillation Parameters

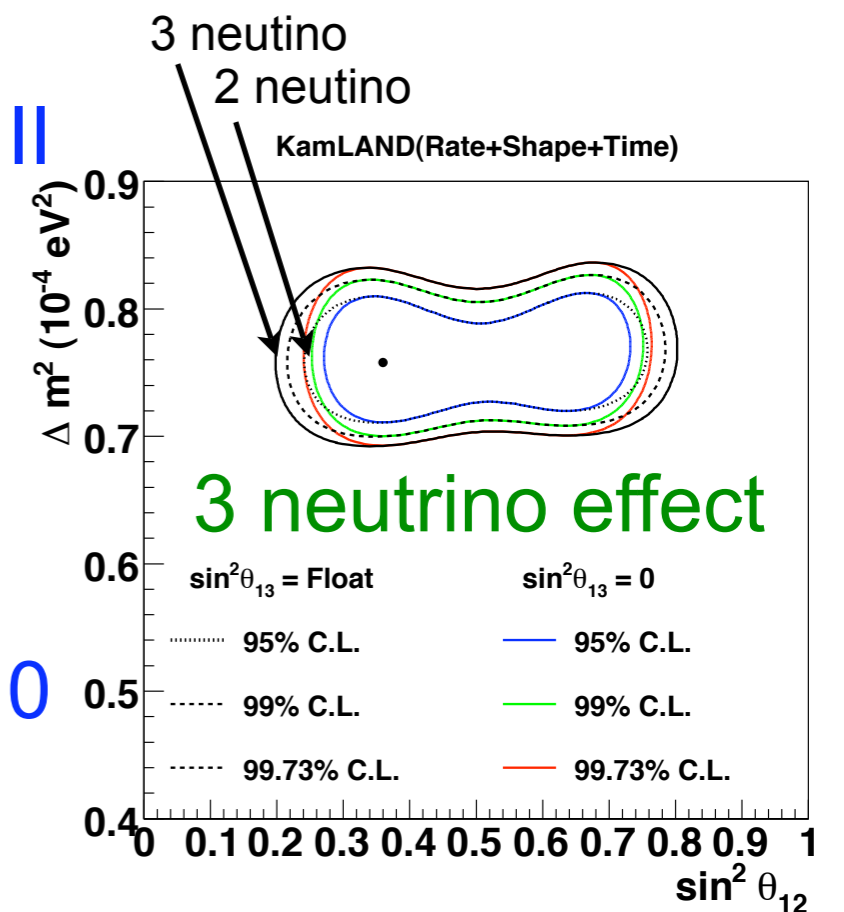


KamLAND only

$$\tan^2\theta = 0.56^{+0.14}_{-0.09}$$

$$\Delta m^2 = 7.58^{+0.21}_{-0.20} \times 10^{-5} \text{ eV}^2$$

(marginalized error)



same result for Δm^2

LMA II

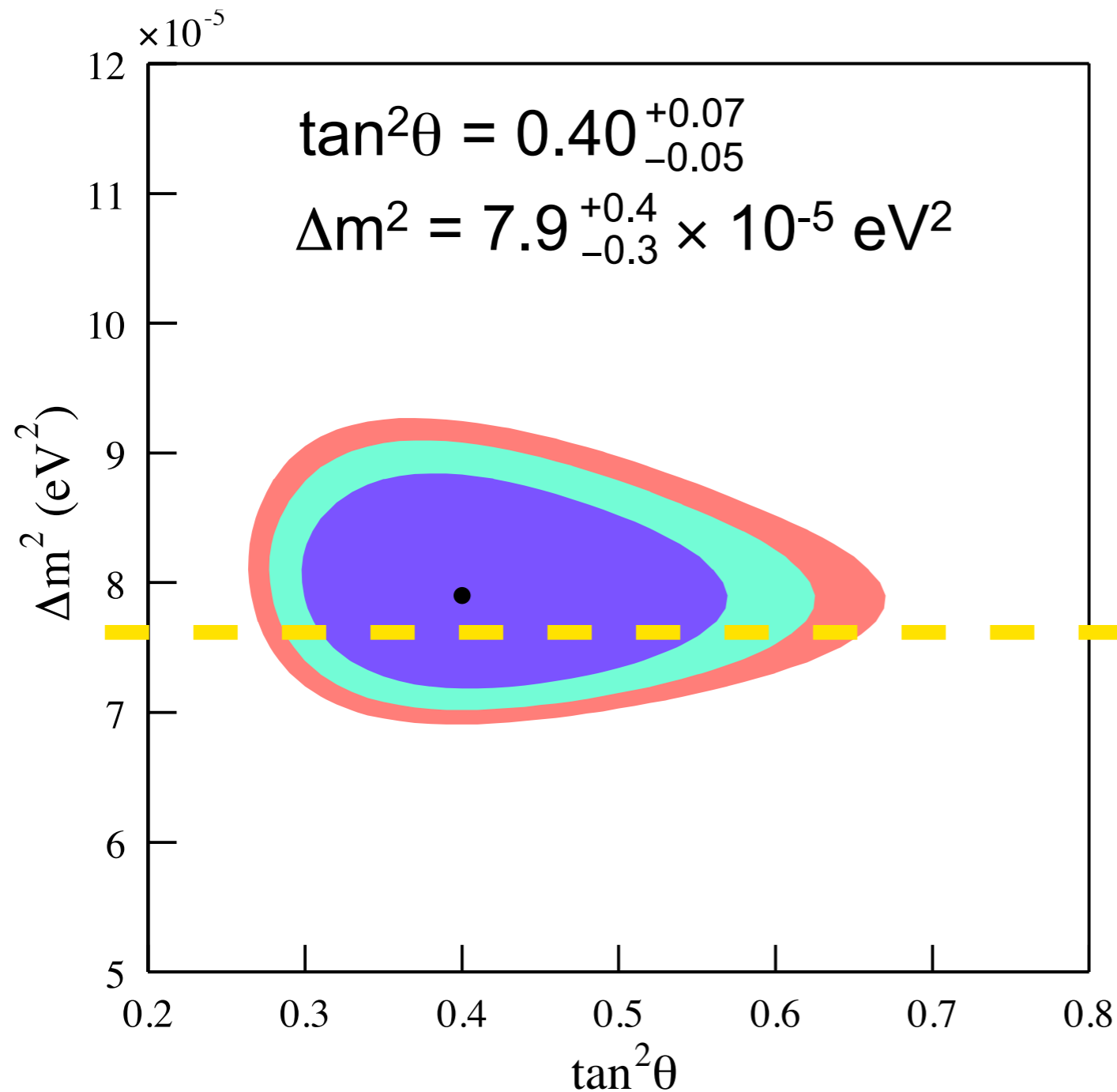
LMA 0

> 6σ

> 4σ

Precise Measurement of Δm^2

KamLAND 2004

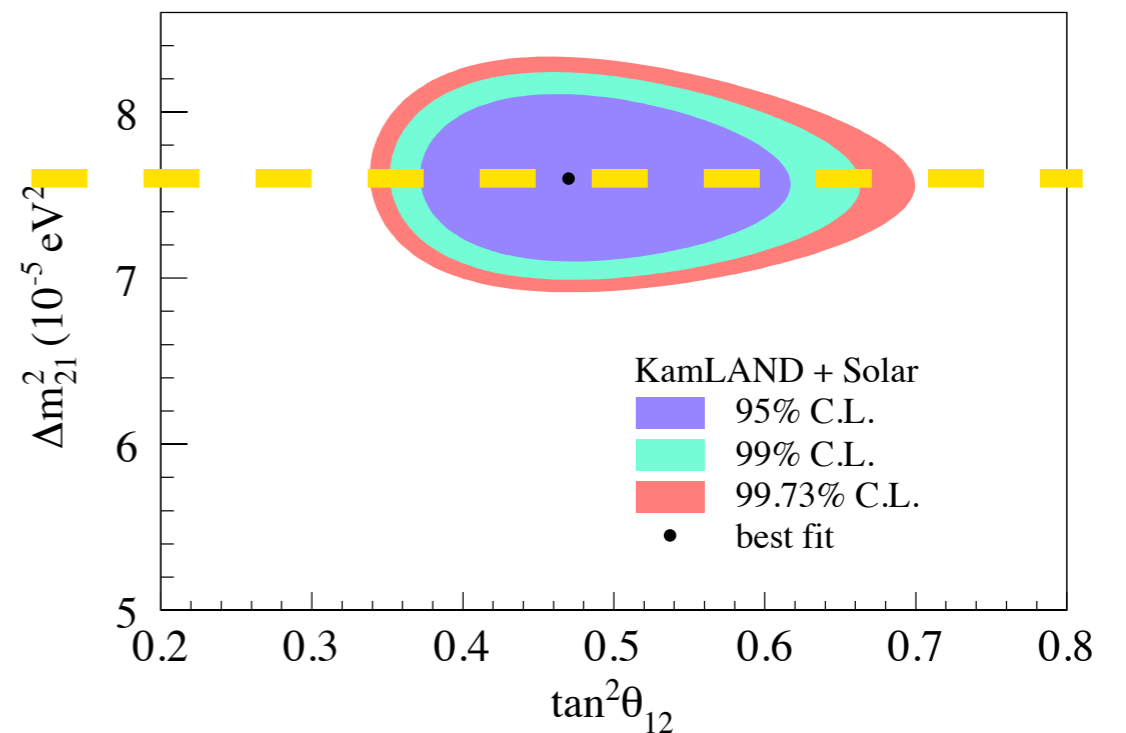


This result

KamLAND + Solar

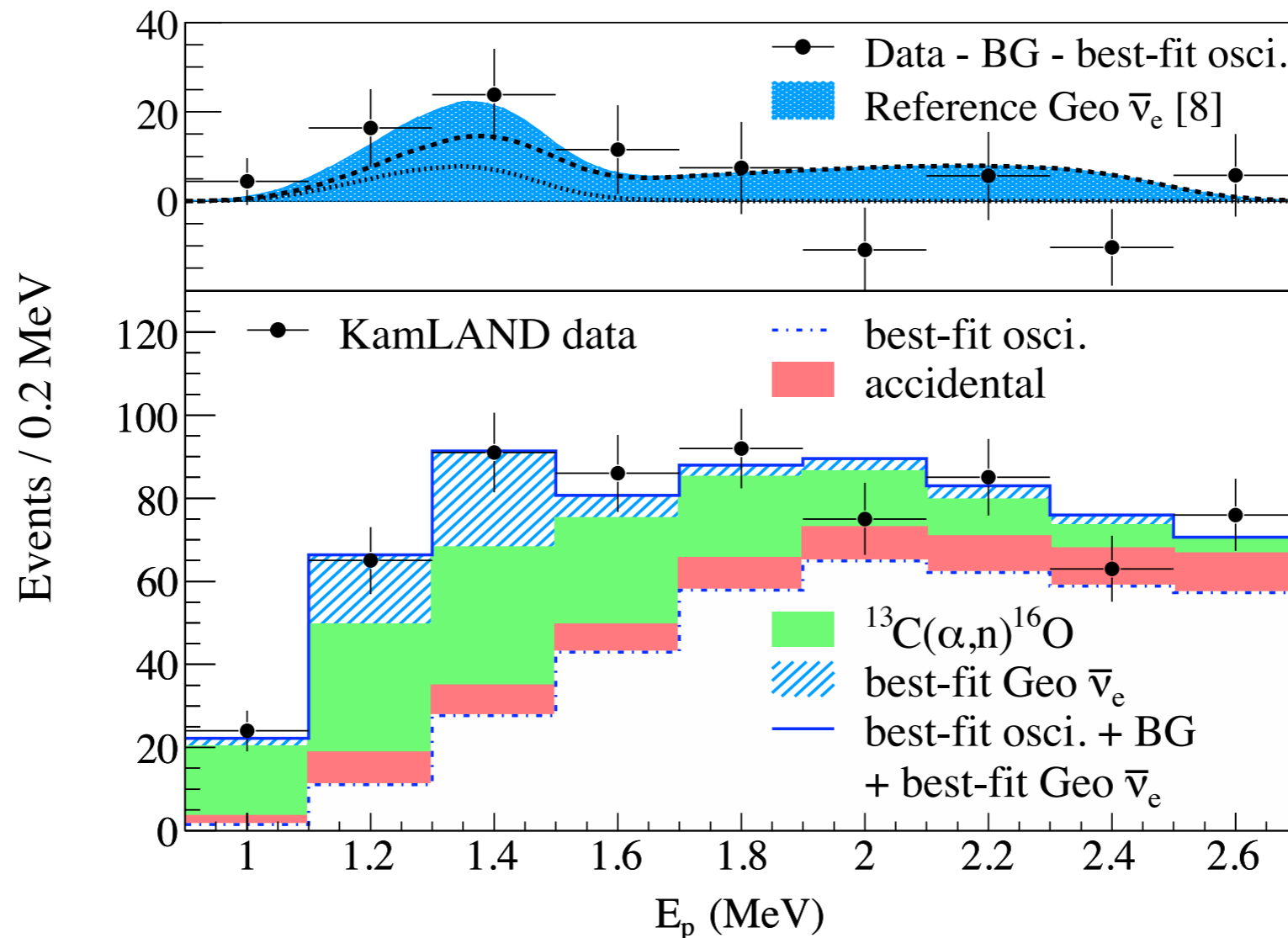
$$\tan^2\theta = 0.47^{+0.06}_{-0.05}$$
$$\Delta m^2 = 7.59^{+0.21}_{-0.21} \times 10^{-5} \text{ eV}^2$$

Δm^2 : systematic uncertainty 2.0%
dominated by linear energy scale uncertainty



Δm^2 is measured at 2.8% precision by KamLAND

Geo Neutrino Estimation



Reference model (16 TW)

U : 56.6 event (29.2 TNU)
 Th : 13.1 event (7.7 TNU)

total : 36.9 TNU

U+Th = $74.9^{+27.3}_{-27.2}$ event

$38.9^{+14.4}_{-14.2}$ TNU

(previous result : $57.4^{+32.0}_{-30.0}$ TNU)

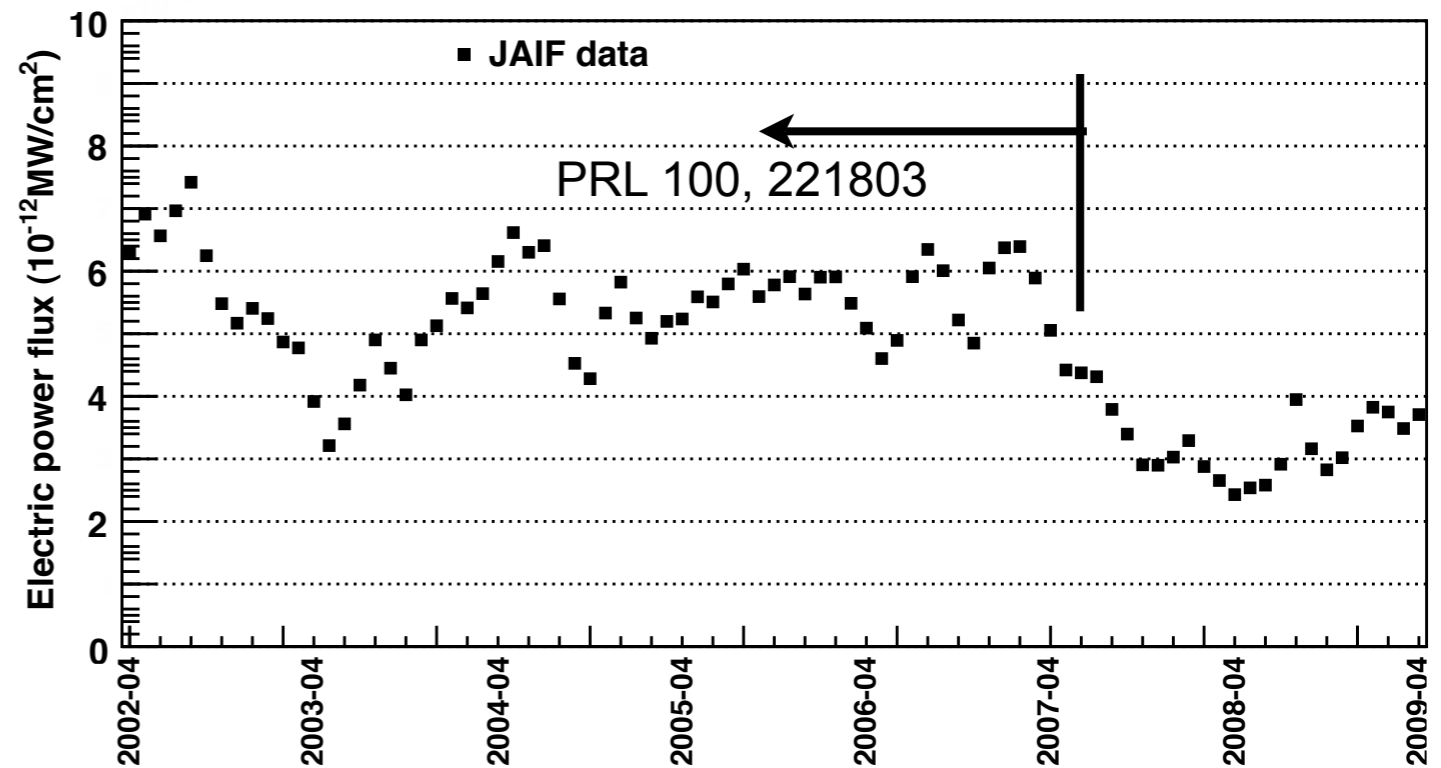
TNU (Terrestrial Neutrino Unit) = events/ 10^{32} target-proton/year

Future Prospects

Status after PRL data-set

(1) Japanese nuclear reactor is operated on low-power recently

(2) After the purification of the LS, the geo neutrino observation will be improved by the (α, n) background reduction

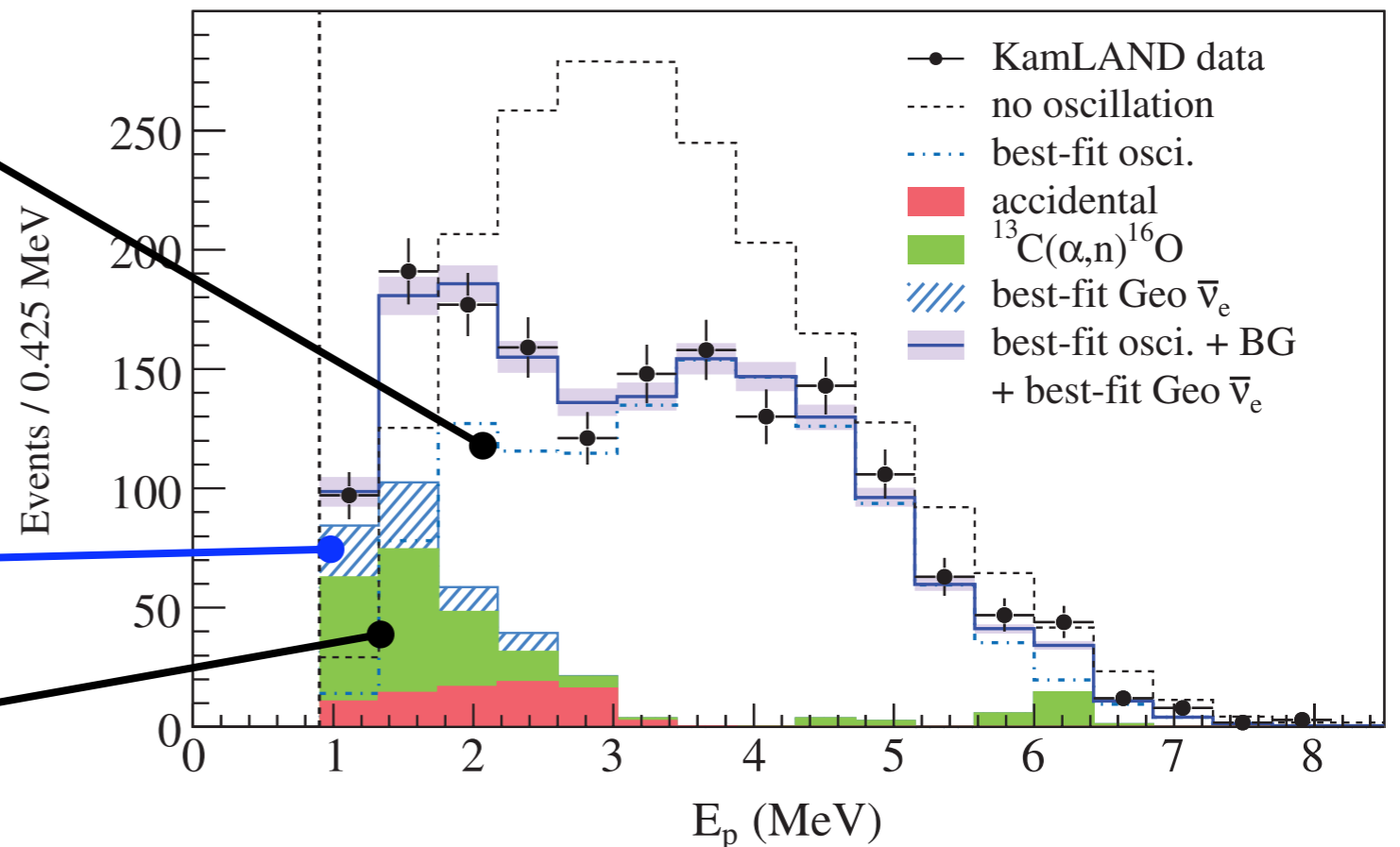


good data for geo neutrino

reactor neutrino is low

geo-neutrinos are being highlighted

(α, n) BG is eliminated



Summary

- KamLAND improved sensitivity to $\bar{\nu}_e$ observation.
data-set : 766 ton-yr \rightarrow 2881 ton-yr (α, n) B.G. uncertainty :
E threshold : 2.6 MeV \rightarrow 0.9 MeV 32% \rightarrow 10% (ground state)
syst. uncertainty : 6.5% \rightarrow 4.1% 100% \rightarrow 20% (excited state)
- In the reactor neutrino analyses, we showed
 - Oscillatory shape \sim 2 cycle of neutrino oscillation.
 - Exclusion of LMA II and 0 at more than 4σ C.L.
 - Precise measurement of oscillation parameters.
- We will continue the $\bar{\nu}_e$ observation to further reduce the uncertainties on the oscillation parameters and geo neutrino fluxes.