

22nd International Workshop on

Weak Interactions and Neutrinos WIN'09

Electroweak Symmetry Breaking

Weak Decays, CP Violation and CKM

Neutrino Physics

Dark Matter

September 14-19, 2009 - Relais San Clemente, Perugia - ITALY

Direct neutrino mass measurements



Monica Sisti

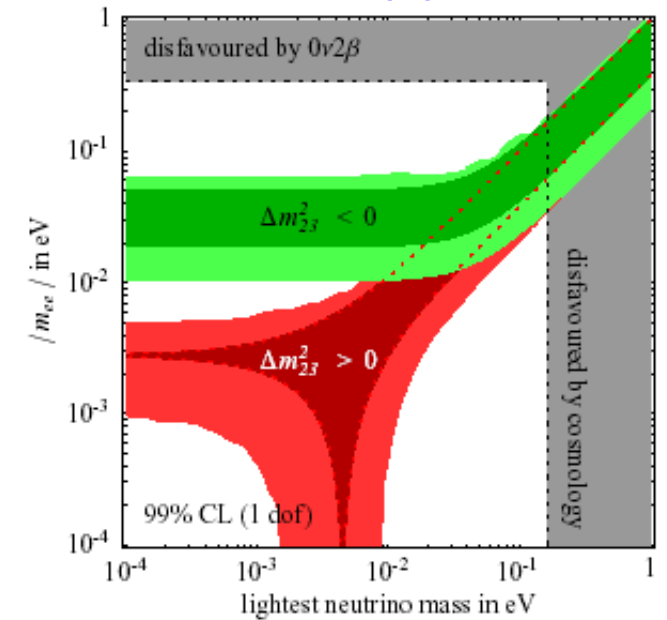
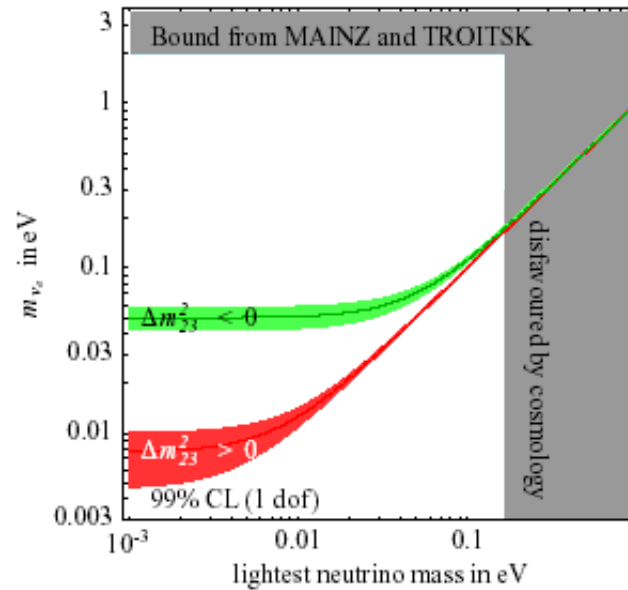
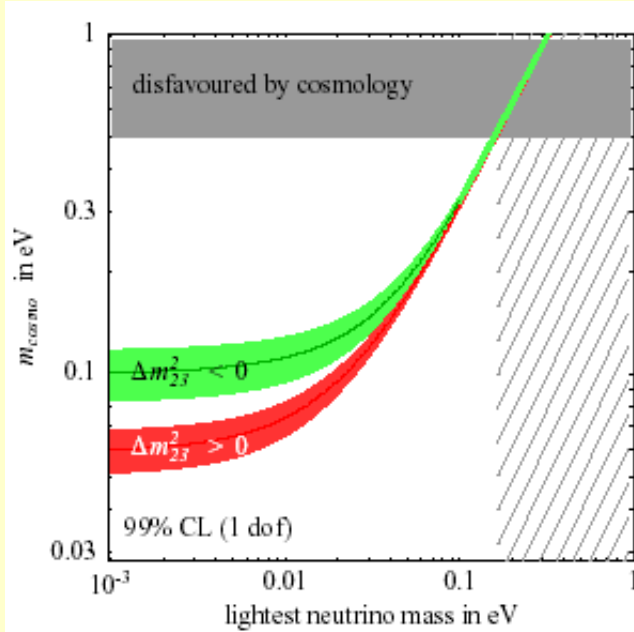
Università and INFN – Milano Bicocca



- **Crucial time for neutrino physics**
- **Importance of direct neutrino mass measurements**
- **Comparing the techniques: spectrometers vs. calorimeters and current state-of-the-art**
- **KATRIN: the future of spectrometer experiments**
- **Prospects for ^{187}Re experiments: the MARE project**

Our knowledge about neutrino properties

Strumia and Vissani: hep-ph/0606054



inverse hierarchy: $m_3 \ll m_1 \approx m_2$

normal hierarchy: $m_1 \approx m_2 \ll m_3$

degeneration: $m_1 \approx m_2 \approx m_3$

Tool

Measured quantity

Sensitivity (eV)

present future

Cosmology (CMB+LSS)

$$m_\Sigma \equiv \sum m_i$$

0.7 ÷ 1

0.07

yes

0ν double beta decay

$$m_{ee} \equiv |\sum m_i |U_{ei}|^2 e^{i\alpha_i}|$$

0.5

0.05

yes

Beta decay

$$m_\beta \equiv (\sum m_i^2 |U_{ei}|^2)^{1/2}$$

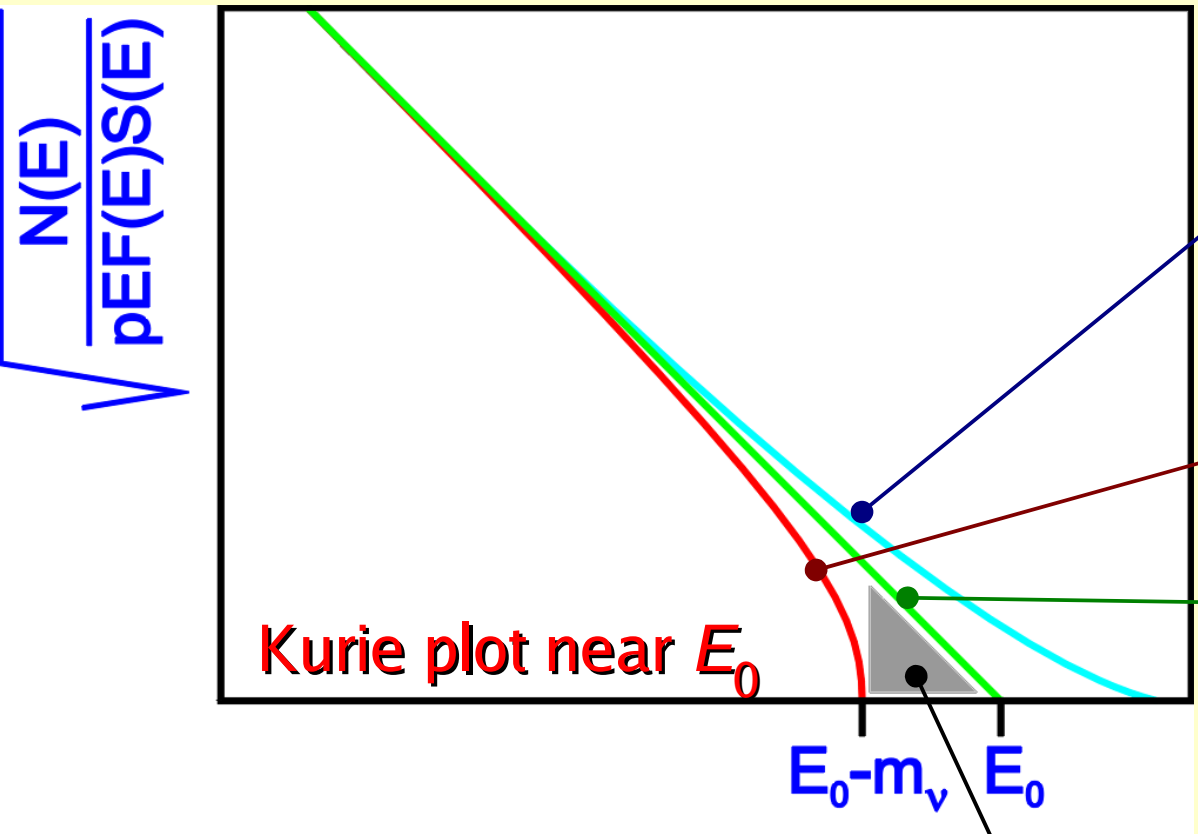
2

0.2

no

model dependency \uparrow

Assessing absolute neutrino mass scale



effect of:

- detector energy resolution
- background counts
- other systematics...

effect of $m_\nu \neq 0$

$N(E_\beta, m_{\bar{\nu}_e} = 0)$

fraction F of decays below the end-point

$$F(\delta E) = \int_{E_0 - \delta E}^{E_0} N(E_\beta, m_{\bar{\nu}_e} = 0) dE$$

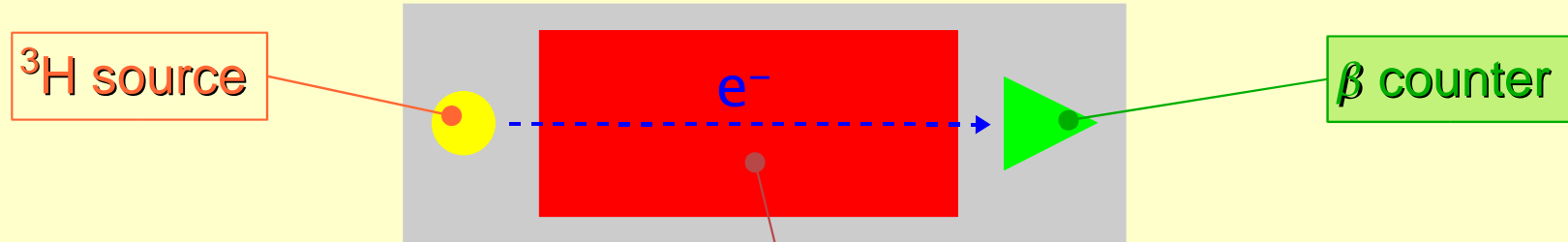
$$\approx \left(\frac{\delta E}{E_0} \right)^3$$

for ${}^3\text{H}$ β -decay $F(10 \text{ eV}) \approx 2 \times 10^{-10}$

- General experimental requirements**
- ◆ low endpoint energy E_0
 - ◆ high statistics at the β spectrum end-point
 - ◆ high energy resolution ΔE
 - ◆ high signal-to-background ratio at the end-point
 - ◆ small systematic effects

Experimental approaches to direct measurement

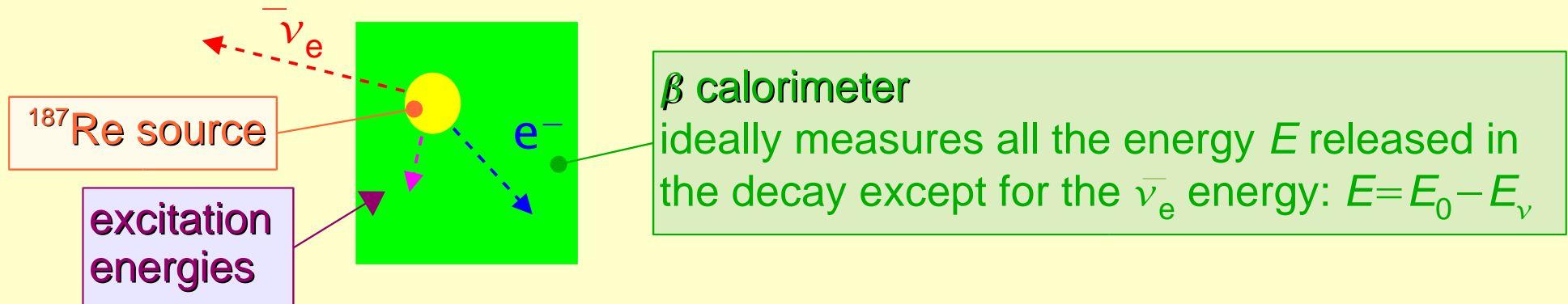
- Spectrometers: source \neq detector



β analyzer

- differential or integral spectrometer: β s from the ${}^3\text{H}$ spectrum δE are magnetically and/or electrostatically selected and transported to the counter

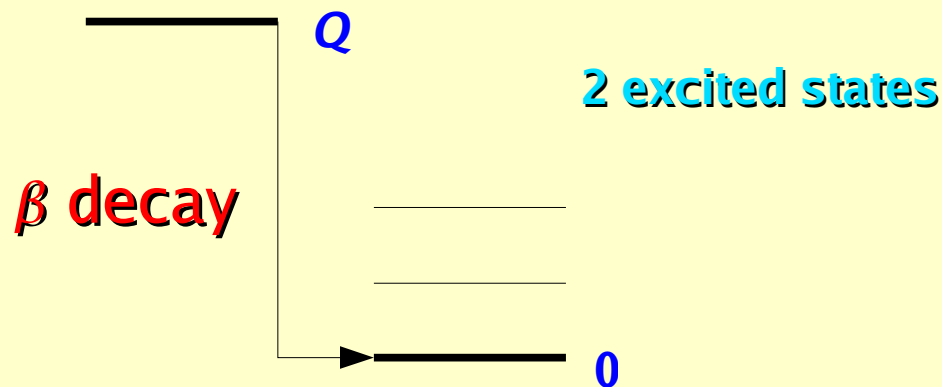
- Calorimeters: source \subseteq detector



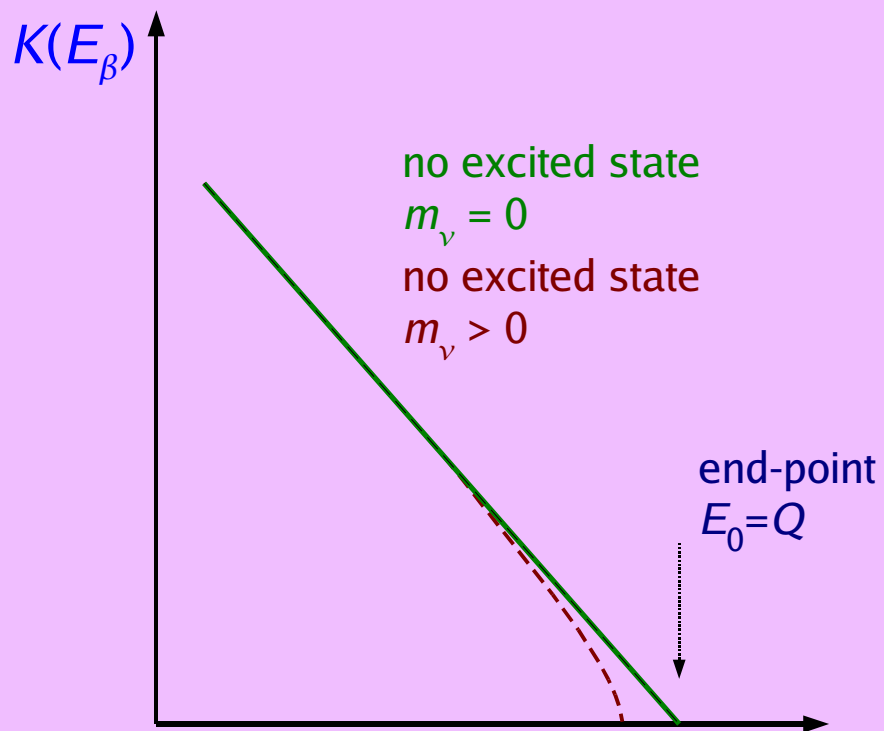
β calorimeter

ideally measures all the energy E released in the decay except for the $\bar{\nu}_e$ energy: $E = E_0 - E_{\bar{\nu}}$

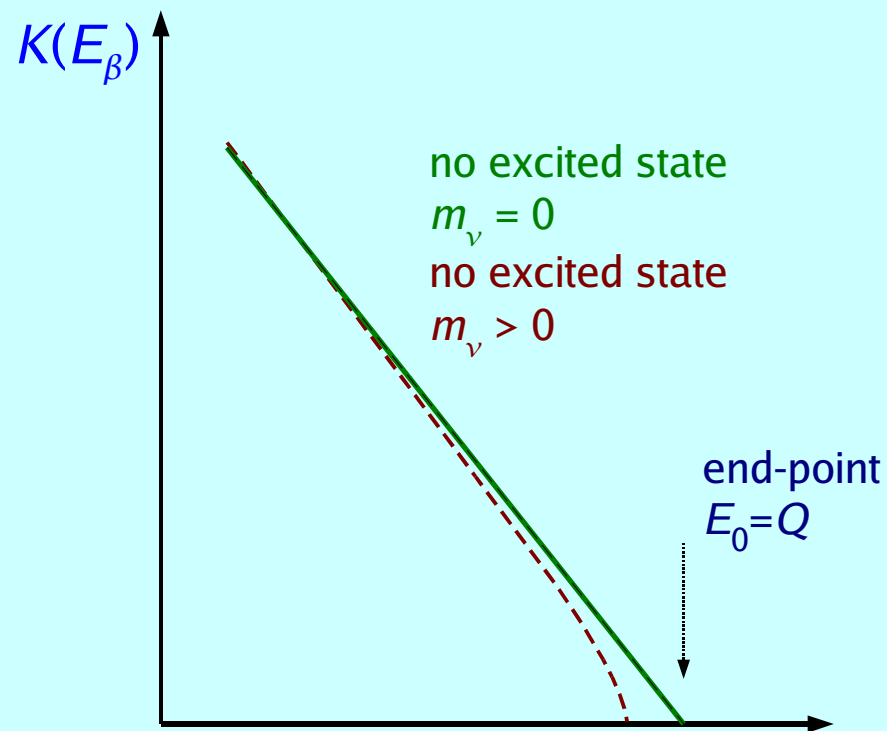
Spectrometry of beta sources



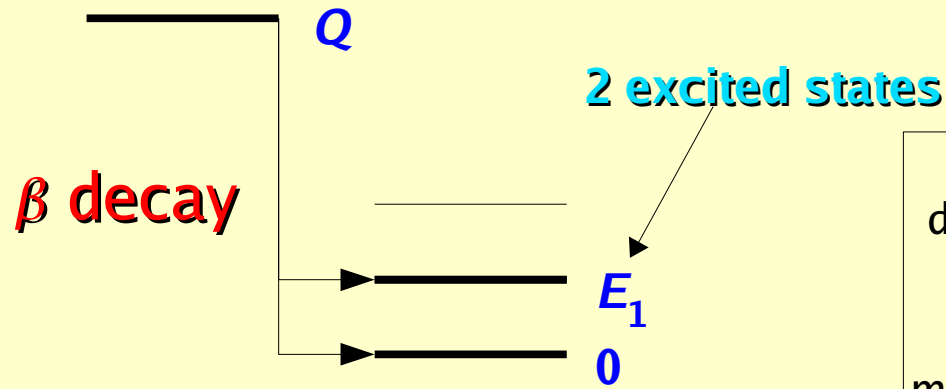
Spectrometers



Calorimeters



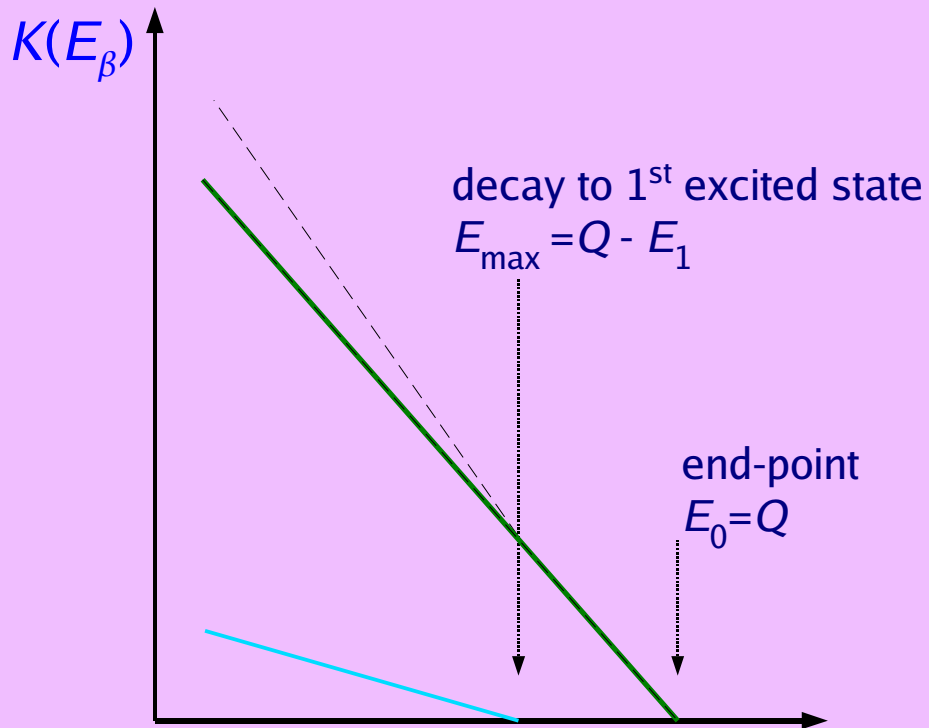
Spectrometry of beta sources



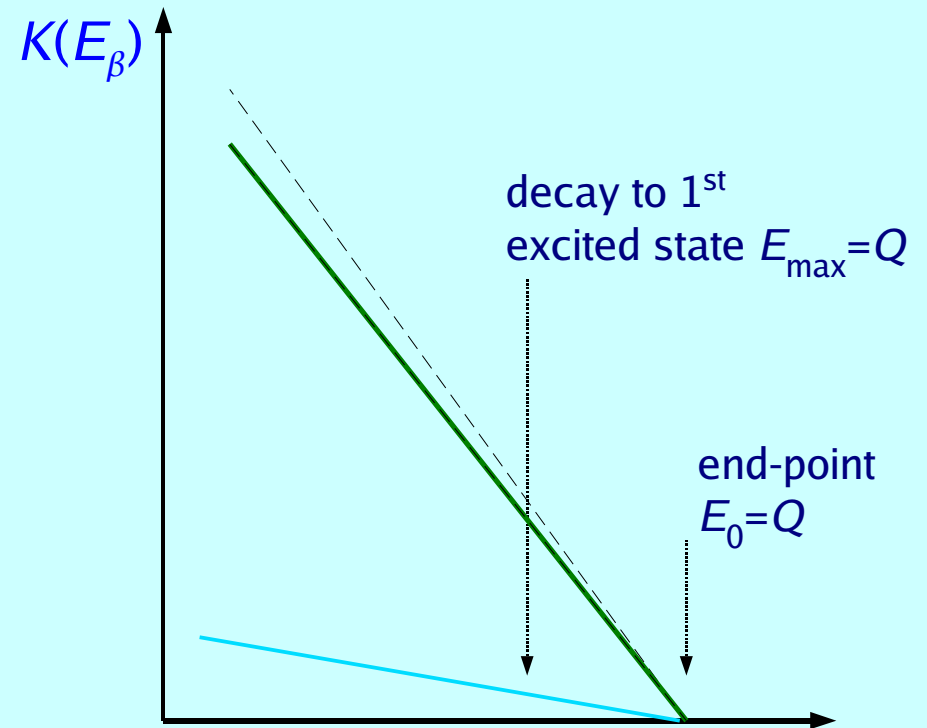
excitation energy is lost

de-excitation faster than
detector response time $t_d \sim 1 \mu s$
⇓
excitation energy is
measured together with β energy

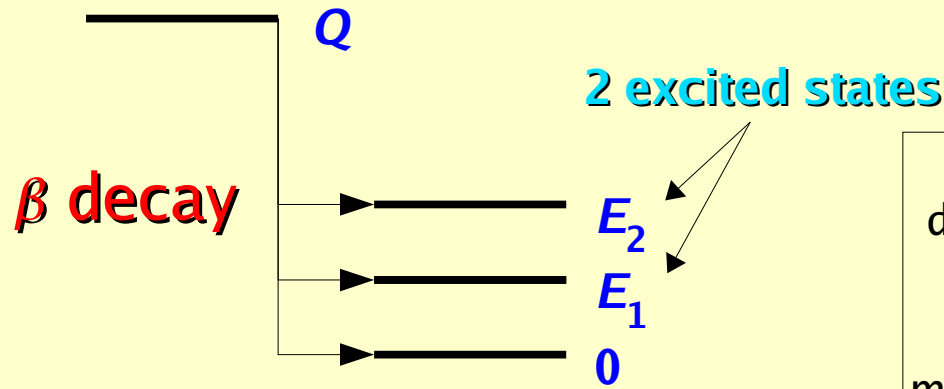
Spectrometers



Calorimeters



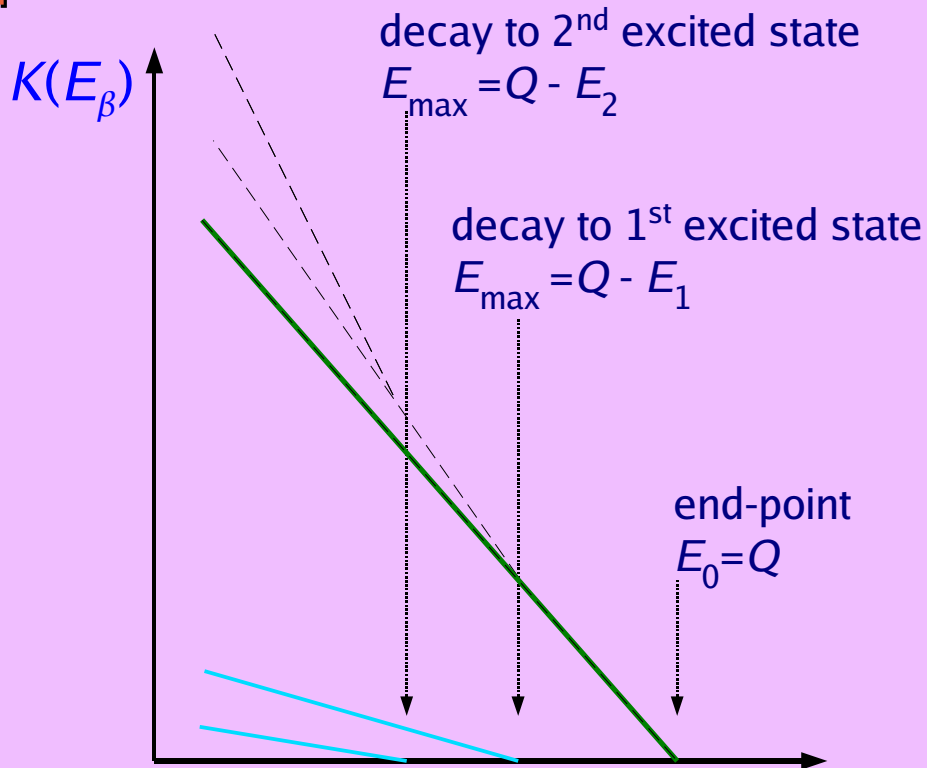
Spectrometry of beta sources



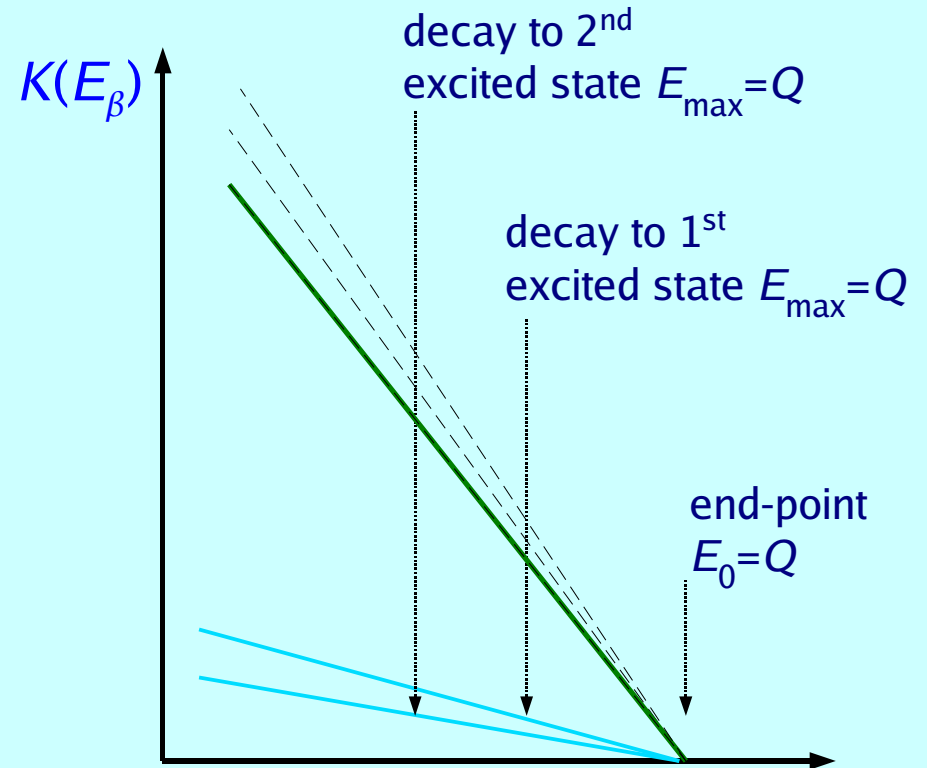
excitation energy is lost

de-excitation faster than
detector response time $t_d \sim 1 \mu\text{s}$
 \Downarrow
 excitation energy is
measured together with β energy

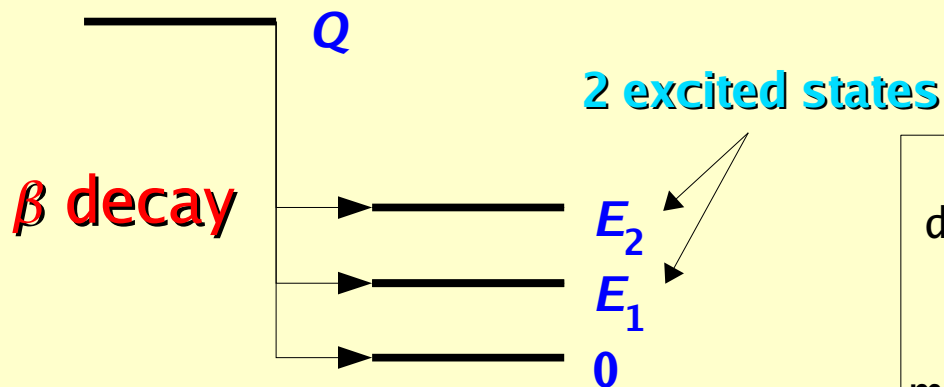
Spectrometers



Calorimeters



Spectrometry of beta sources



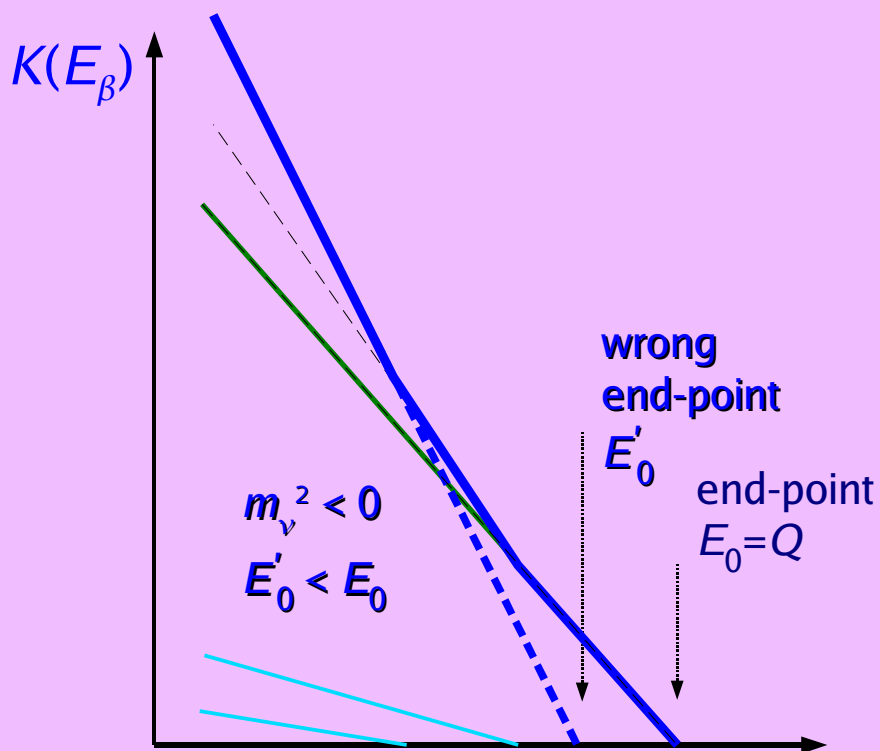
excitation energy is lost

de-excitation faster than detector response time $t_d \sim 1 \mu s$

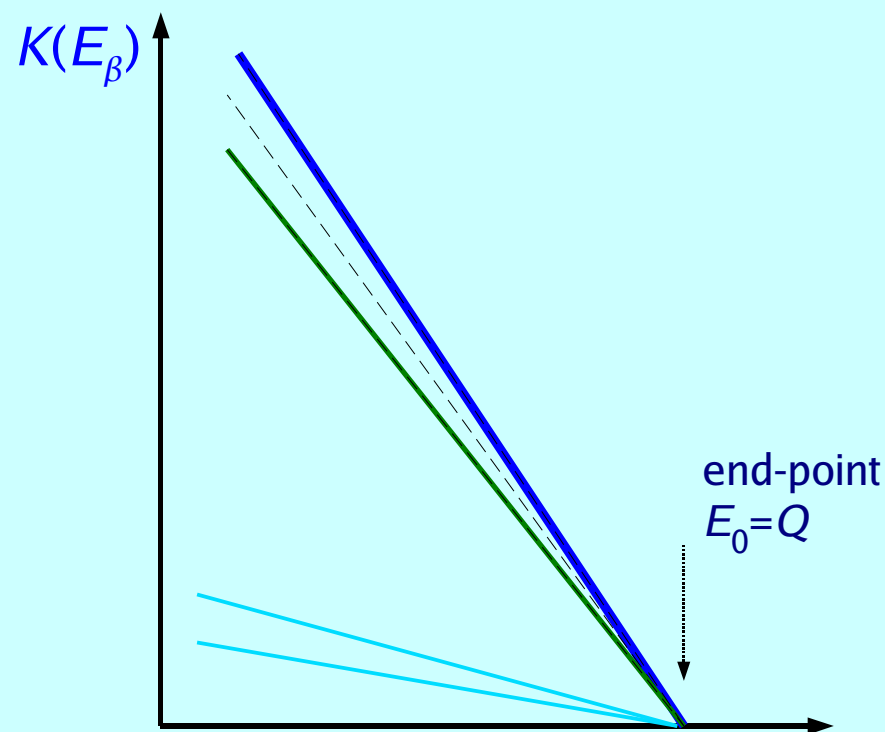
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excitation energy is measured together with β energy

Spectrometers



Calorimeters



Calorimetry of beta sources

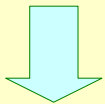
Calorimeters measure the entire spectrum at once

- use low E_0 β decaying isotopes to achieve enough statistics close to E_0
- best choice ^{187}Re : $E_0 = 2.47 \text{ keV} \Rightarrow F(\delta E=10 \text{ eV}) \approx (\delta E/E_0)^3 = 7 \times 10^{-8}$

Pile-up

- time unresolved superposition of β decays
- for a source activity A_β , a time resolution τ_R and an energy resolution function $R(E_\beta)$

$$N^{\text{exp}}(E_\beta) \approx (N(E_\beta) + \tau_R A_\beta \cdot N(E_\beta) \otimes N(E_\beta)) \otimes R(E_\beta)$$

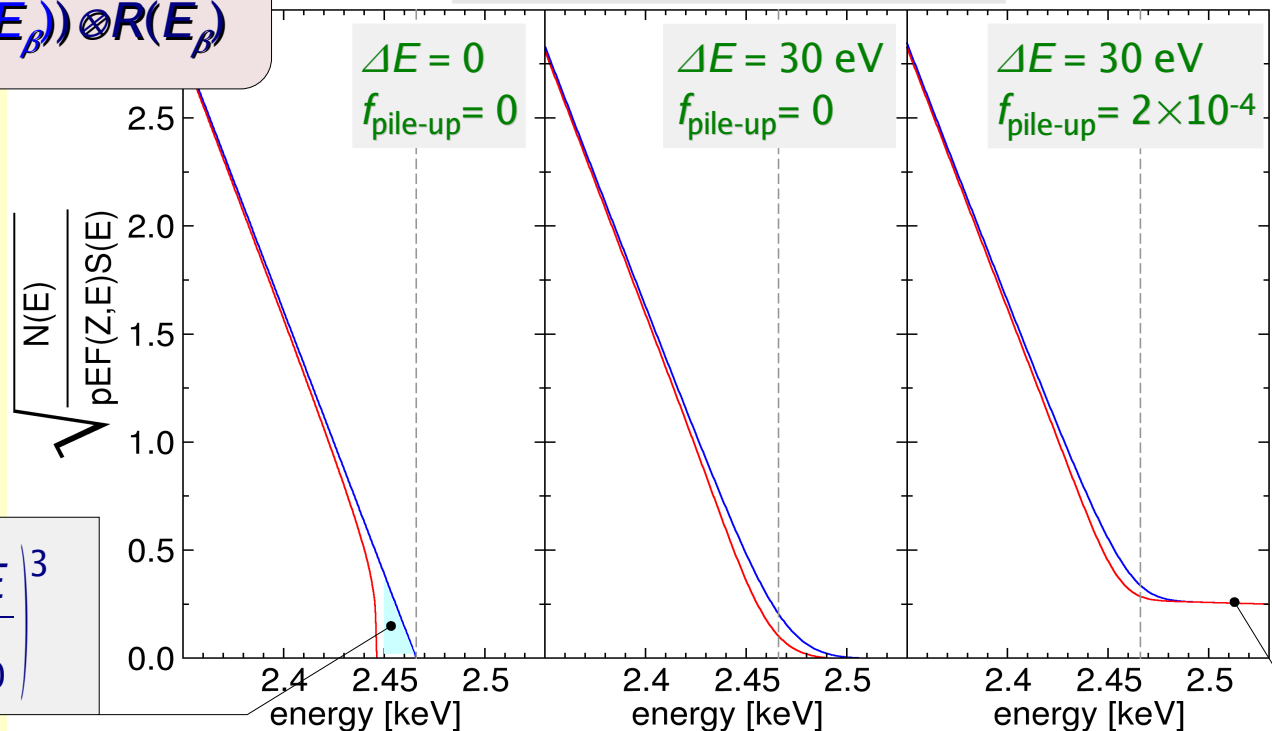


Arrays of microcalorimeters

$$F(\delta E) \approx \left(\frac{\delta E}{E_0} \right)^3$$

$^{187}_{75}\text{Re} \rightarrow ^{187}_{76}\text{Os} + e^- + \bar{\nu}_e$
 natural isotopic abundance: 63%
 $\rightarrow 1 \text{ mg metallic Re} \approx 1.0 \text{ dec/s}$

— $m_\nu = 0$ — $m_\nu = 20 \text{ eV}$



pile-up fraction: $f_{pp} = \tau_R A_\beta$

^3H beta decay in the last twenty years

ITEP

T_2 in complex molecule
magn. spectrometer (Tret'yakov)

m_ν

17-40 eV

Los Alamos

gaseous T_2 - source
magn. spectrometer (Tret'yakov)

< 9.3 eV

Tokio

T - source
magn. spectrometer (Tret'yakov)

< 13.1 eV

Livermore

gaseous T_2 - source
magn. spectrometer (Tret'yakov)

< 7.0 eV

Zürich

T_2 - source impl. on carrier
magn. spectrometer (Tret'yakov)

< 11.7 eV

Troitsk (1994-today)

gaseous T_2 - source
electrostat. spectrometer

< 2.5 eV

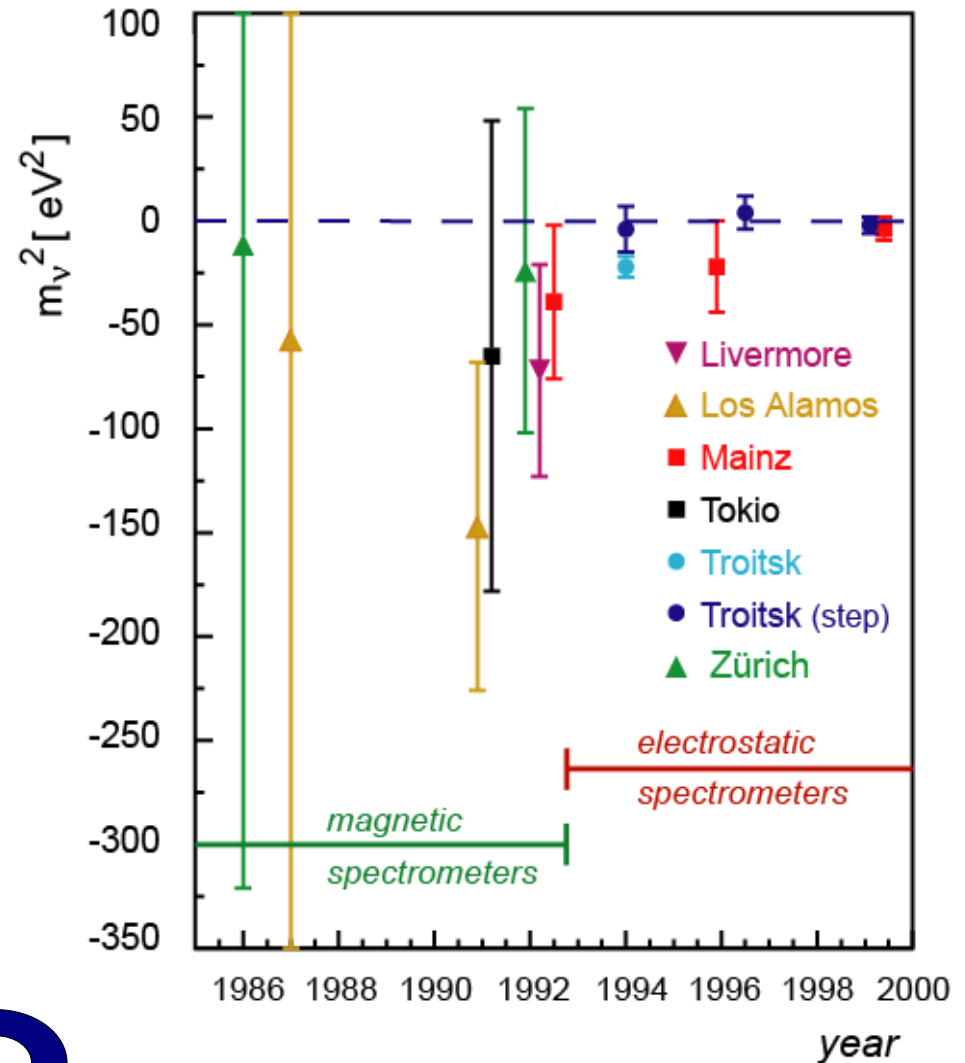
PDG 2008:

Mainz (1994-today)

frozen T_2 - source
electrostat. spectrometer

< 2.3 eV

experimental results



PDG evaluation: < 2 eV (95% C.L.)

Electrostatic filter with Magnetic Adiabatic Collimation

MAC-technique

adiabatic guiding of β -particles along the magnetic field lines

inhomogen. B-Feld:
stray field of 2 super-
conducting magnets

$B_{\max} = 3 - 6 \text{ T}$

$B_{\min} < 1 \text{ mT}$

very large solid angle !

$$\Delta\Omega \sim 2\pi$$

E-technique

energy analysis by
electro static retarding
field (electrodes)

integral particle
transmission $E > U_0$

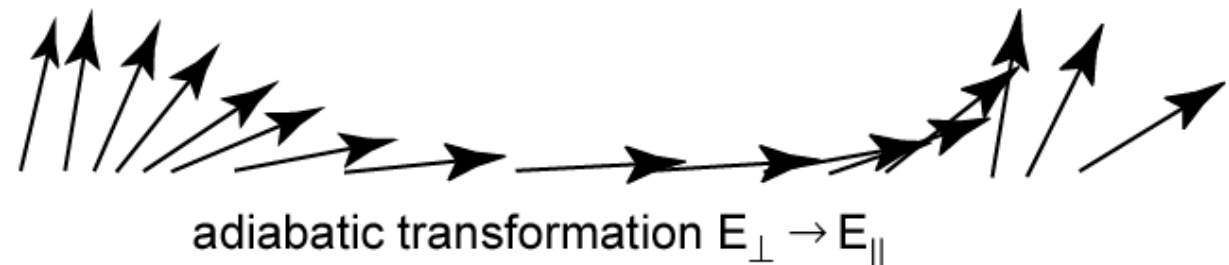
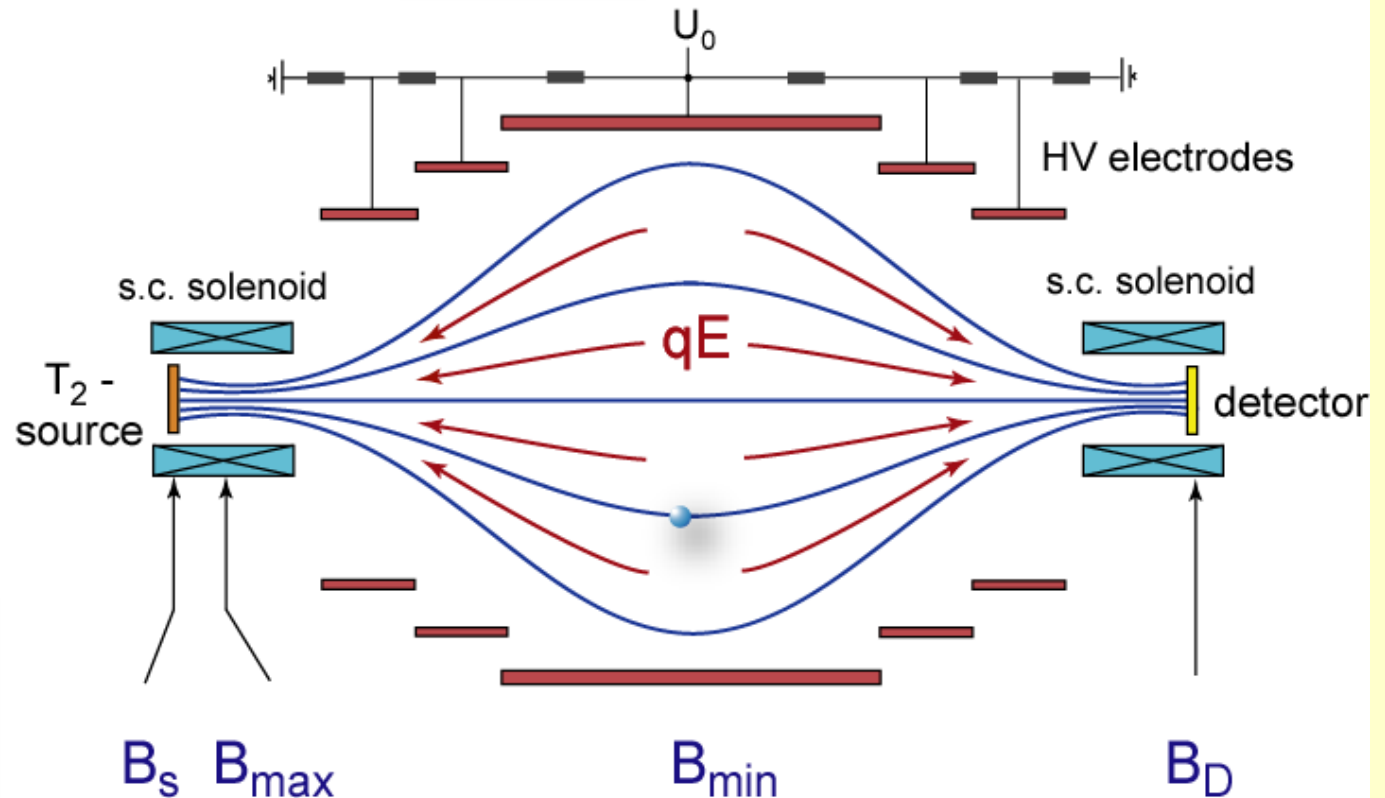
high pass filter !

MAC-E filter

$$U_0 < 30 \text{ kV}$$

$$\vec{F} = (\vec{\mu} \cdot \nabla) \vec{B} + q \vec{E}$$

$$\mu = E_{\perp} / B = \text{const}$$



The most sensitive neutrino mass experiments

Mainz & Troitsk have reached their intrinsic limit of sensitivity



Troitsk

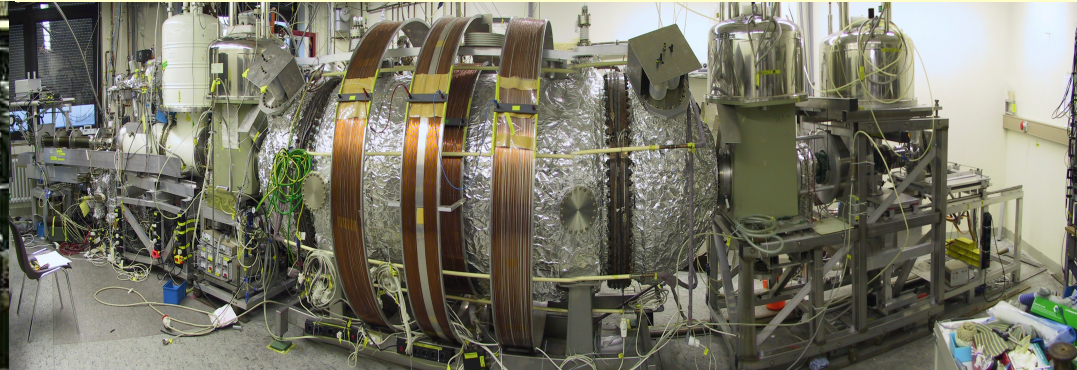
windowless gaseous T_2 source

$$m_\nu^2 = -2.3 \pm 2.5_{\text{stat}} \pm 2.0_{\text{sys}} \text{ eV}^2$$

$$m_\nu \leq 2.05 \text{ eV (95 \% C.L.)}$$

Lobashev, Nucl. Phys. A 719 (2003) 153

with a phenomenological correction for yet unexplained small anomaly in the spectrum



Mainz

quench condensed solid T_2 source

$$m_\nu^2 = -0.6 \pm 2.2_{\text{stat}} \pm 2.1_{\text{sys}} \text{ eV}^2$$

$$m_\nu \leq 2.3 \text{ eV (95 \% C.L.)}$$

Kraus et al., Eur. Phys. J. C 40 (2005) 447

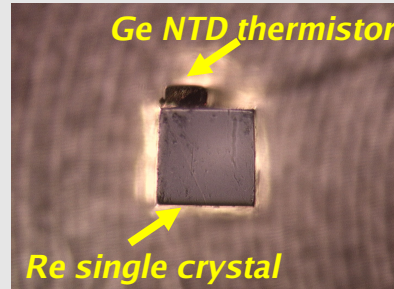
both experiments now used for systematic investigations

Precursor ^{187}Re experiments

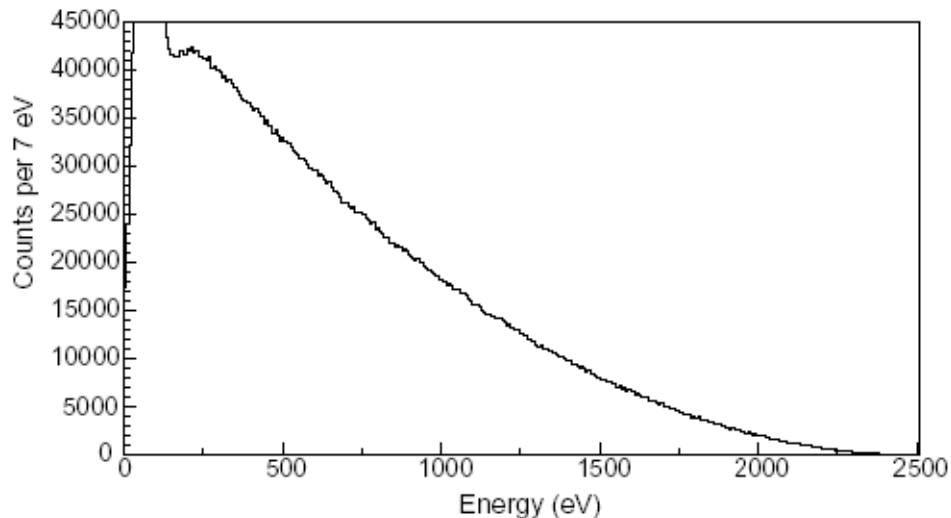
MANU (1999)

Genova

- 1 crystal of metallic Re: 1.6 mg
- ^{187}Re activity ≈ 1.6 Hz
- Ge-NTD thermistor
- $\Delta E = 96$ eV FWHM
- 0.5 years live-time
- $m_\nu^2 = -462^{+579}_{-679} \text{ eV}^2$
- $m_\nu < 26$ eV (95 % C.L.)



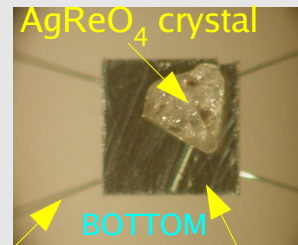
6.0×10^6 ^{187}Re decays above 420 eV



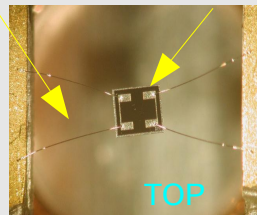
MIBETA (2002-2003)

Milano, Como, Trento

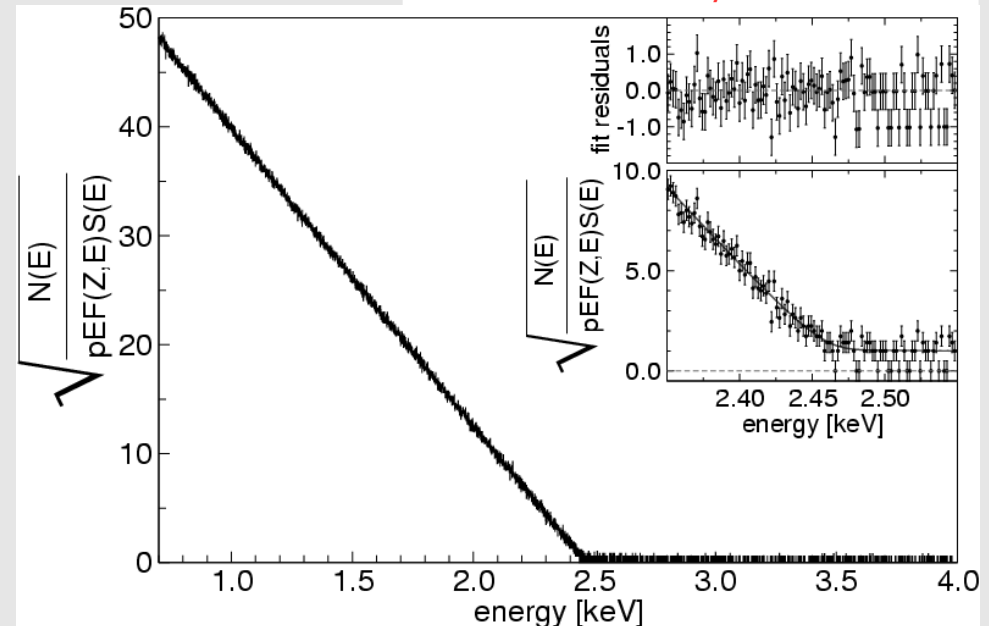
- 10 AgReO_4 crystals: 2.71 mg
- ^{187}Re activity = 0.54 Hz/mg
- Si thermistors (ITC-irst)
- $\Delta E = 28.5$ eV FWHM
- 0.6 years live time
- $m_\nu^2 = -112 \pm 207_{\text{stat}} \pm 90_{\text{sys}} \text{ eV}^2$
- $m_\nu < 15$ eV (90 % C.L.)



Al bonding wires
Si thermistor



6.2×10^6 ^{187}Re decays above 700 eV



Comparing techniques in β decay

Spectrometers

◆ Choice of β -emitter: ^3H

- $E_0 = 18.6$ keV
- $\tau_{1/2} = 12.3$ y
- $1/2^+ \rightarrow 1/2^+$ superallowed transition

Achieved sensitivity:
 ~ 2 eV

◆ Advantages

- ▲ high statistics
- ▲ high energy resolution
- ▲ simple atomic/molecular structure

◆ Drawbacks

- ▼ systematics due to source effects
- ▼ systematics due to excited final states
- ▼ background
- ▼ spectrometer stability
- ▼ ...

Future planned sensitivity:

KATRIN $\rightarrow 0.2$ eV (data taking: 2012)

Calorimeters

◆ Choice of β -emitter: ^{187}Re

- $E_0 = 2.5$ keV
- $\tau_{1/2} = 43.2$ Gy
- $5/2^+ \rightarrow 1/2^-$ unique 1^{st} forbidden

Achieved sensitivity:
 ~ 15 eV

◆ Advantages

- ▲ measure neutrino energy
- ▲ no backscattering/self-absorption
- ▲ no excited final state effects

◆ Drawbacks

- ▼ limited statistics
- ▼ systematics due to pile-up
- ▼ energy dependent background
- ▼ ^{187}Re spectral shape
- ▼ ...

Future planned sensitivity:

MARE-1 $\rightarrow 2 \div 4$ eV (data taking: 2009)

MARE-2 $\rightarrow 0.2$ eV (2015???)

Complementary techniques – Different systematics

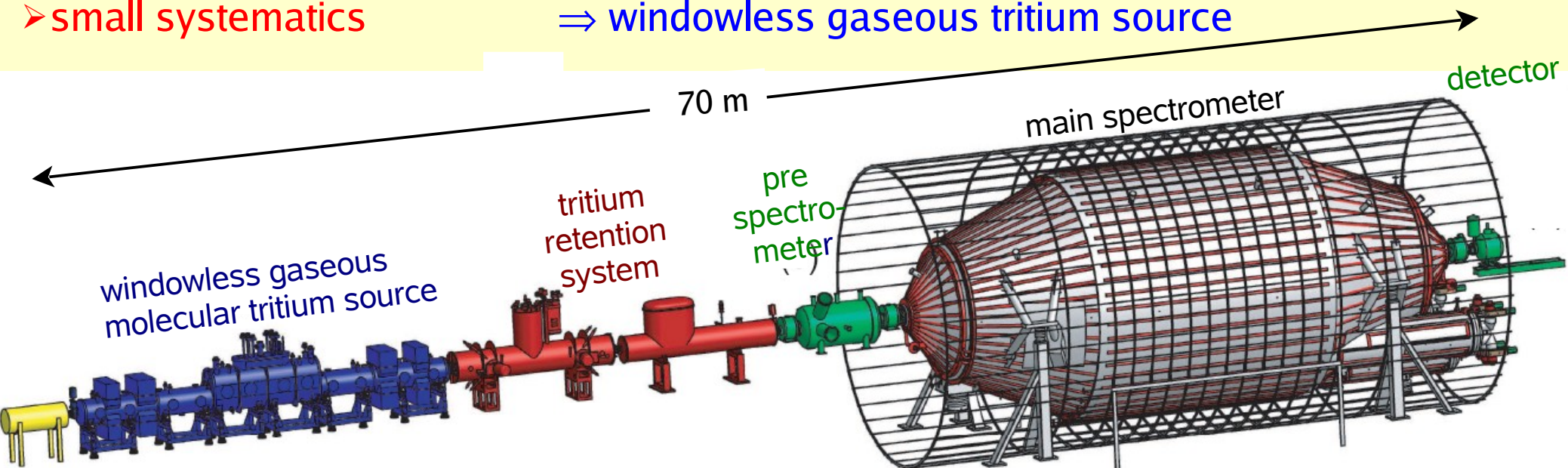
The Karlsruhe Tritium Neutrino experiment: KATRIN

Physics goal: one order of magnitude improvement on m_ν
 $2 \text{ eV} \rightarrow 0.2 \text{ eV}$

- higher energy resolution: $\Delta E \approx 1 \text{ eV}$
 since $E/\Delta E \sim A_{\text{spectrometer}} \Rightarrow$ larger spectrometer
- higher statistics
 $dN/dt \sim A_{\text{source}} \sim A_{\text{spectrometer}} \Rightarrow$ larger spectrometer
- longer measuring time: $\Rightarrow 100 \text{ d} \rightarrow 1000 \text{ d}$
- small systematics \Rightarrow windowless gaseous tritium source



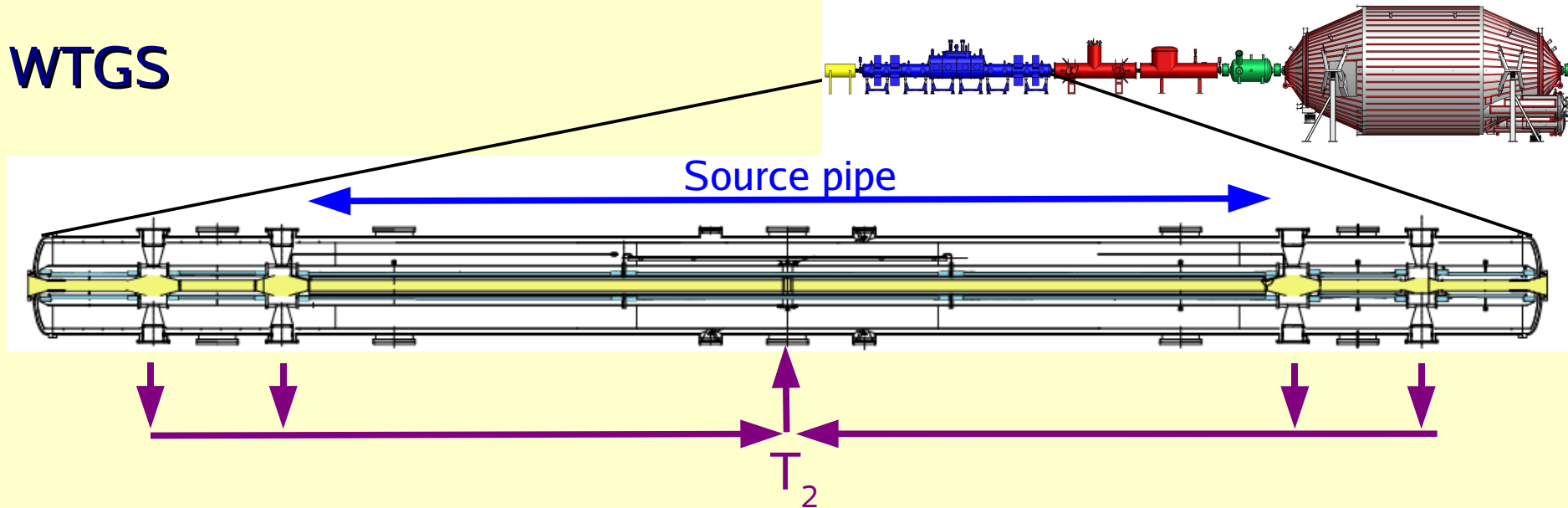
$\varnothing 10\text{m}$



(Scientific Report FZKA 7090)

Molecular Windowless Gaseous Tritium Source

WTGS



stainless steel tube in long superconducting solenoids

∅ 9 cm, length 10 m, T = 30 K (± 0.1%)

Magnetic field: 3.6 Tesla (± 2%)

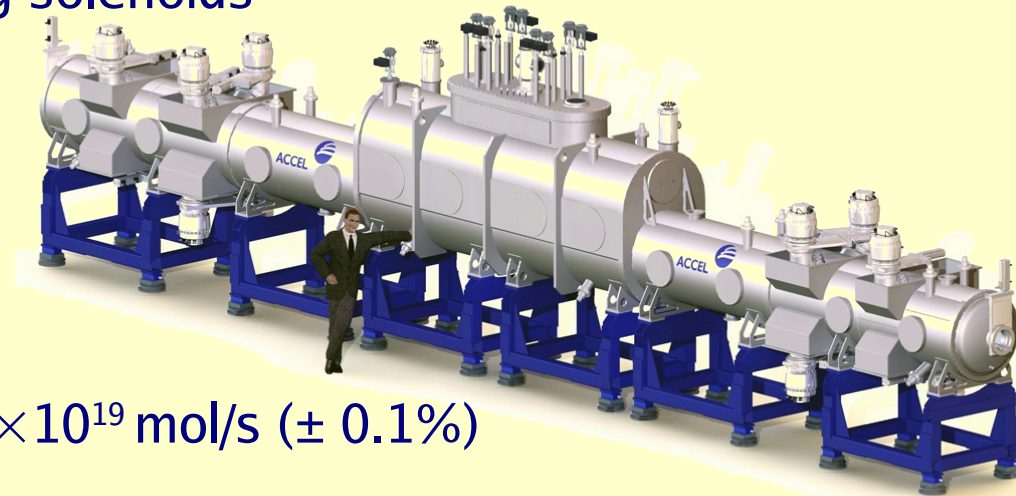
Tritium recirculation (and purification)

Tritium purity > 95%

Tritium injection: $p_{inj} = 0.003$ mbar, $rate_{inj} = 5 \times 10^{19}$ mol/s (± 0.1%)

Integral column density: $\rho d = 5 \times 10^{17}/\text{cm}^2$ (± 0.1%)

for high signal rate with small systematics
(maximum accepted solid angle: 51°)



WGTS-demonstrator
will arrive early 2010

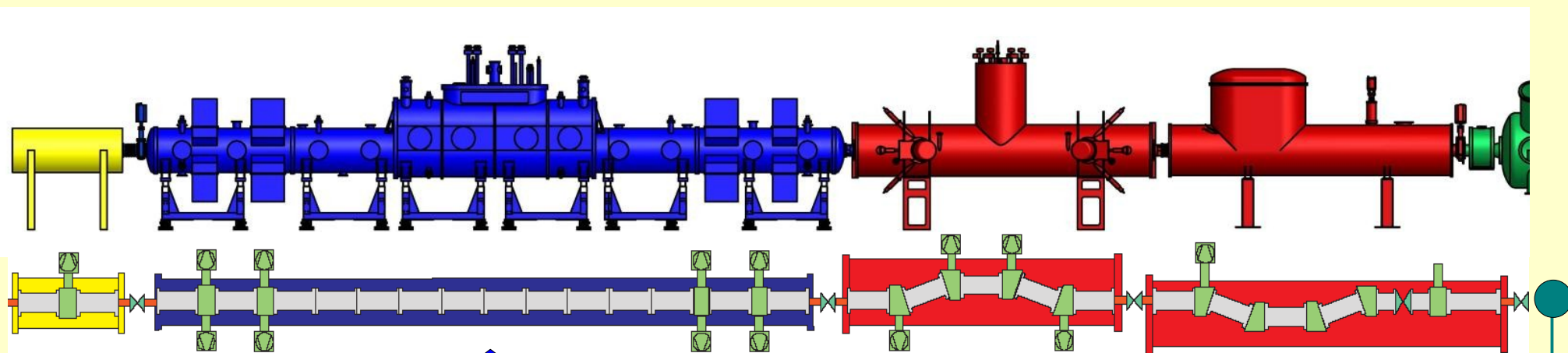
Transport and pumping sections

adiabatic electron guiding & T_2 flow reduction factor of $\sim 10^{14}$

Molecular windowless
gaseous tritium source

Differential
pumping

Cryogenic
pumping



T_2 -injection 1.8 mbar l/s (STP)
= 1.7×10^{11} Bq/s = 40 g/d

$R > 10^7$
 $\approx 10^{-7}$ mbar l/s

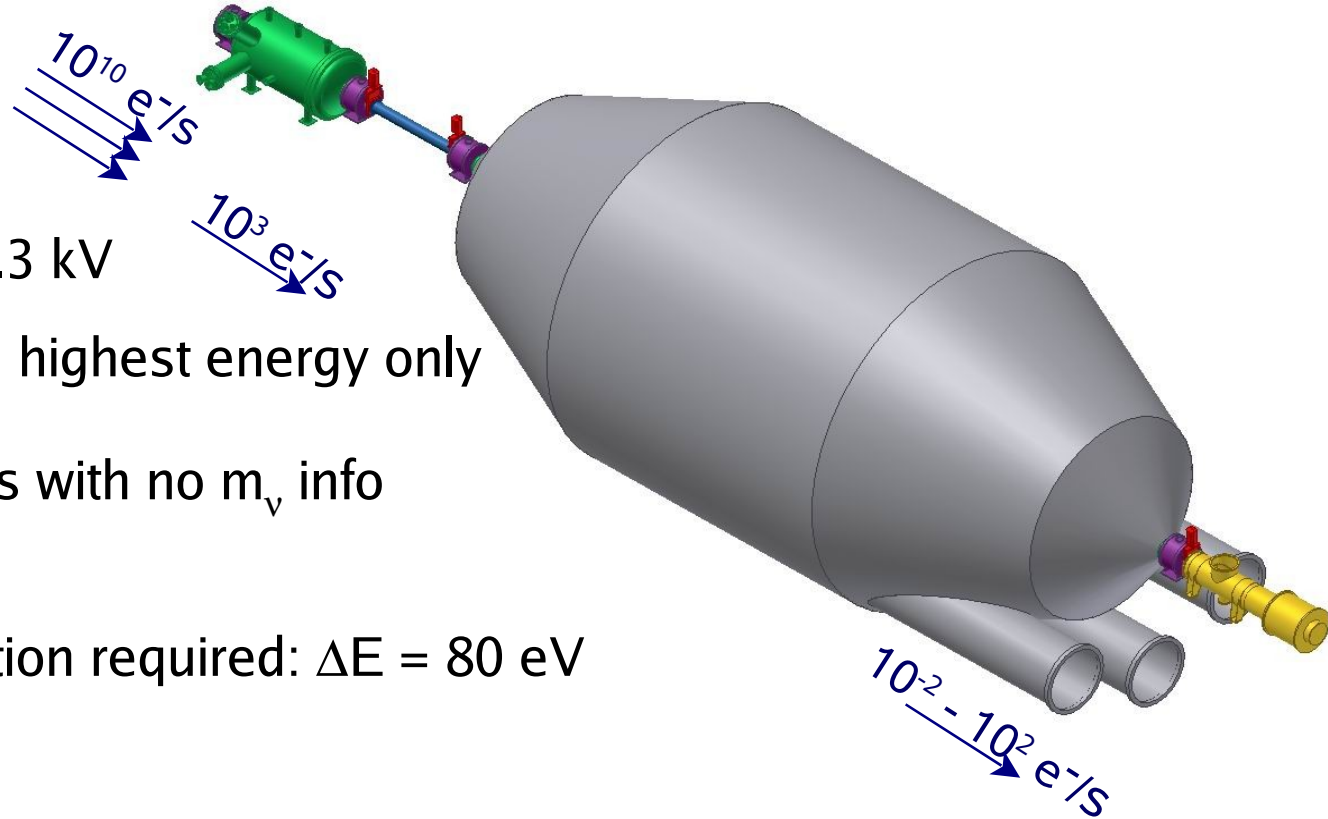
$R > 10^7$
 $< 2.5 \cdot 10^{-14}$ mbar l/s

$p(T_2) < 10^{-20}$ mbar

Pre and main electrostatic spectrometers

Pre spectrometer

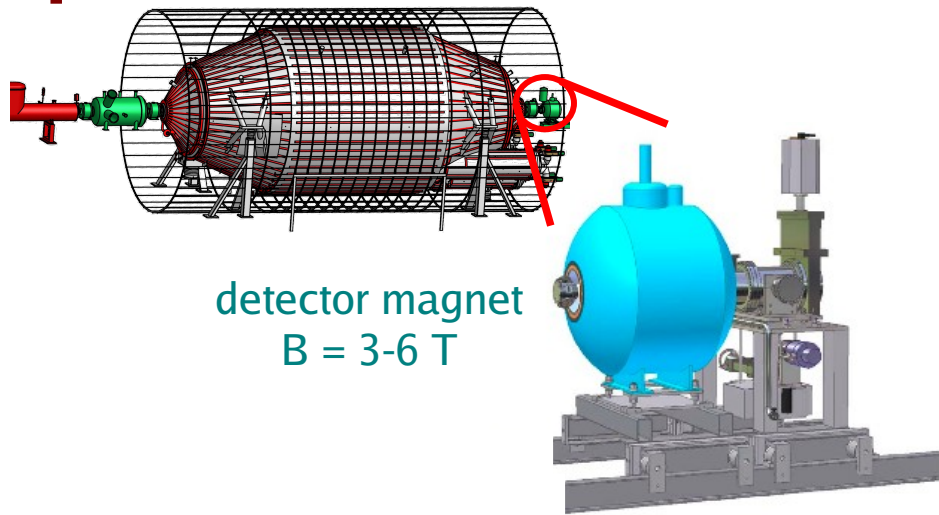
- \varnothing 1.7 m, length 3.5 m
- Fixed retarding potential: -18.3 kV
- Transmission of electron with highest energy only (10^{-7} part in last 100 eV)
m \rightarrow Filter all β -decay electrons with no m_ν info
 \rightarrow Reduction of background
- only moderate energy resolution required: $\Delta E = 80$ eV



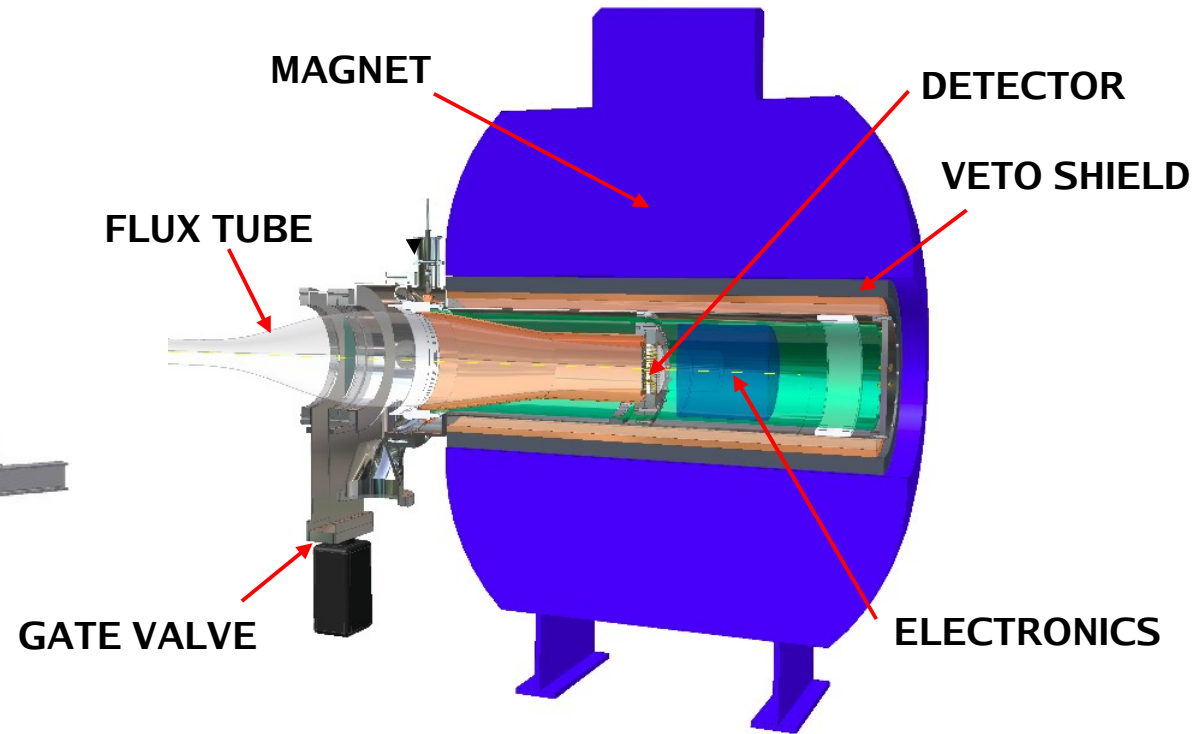
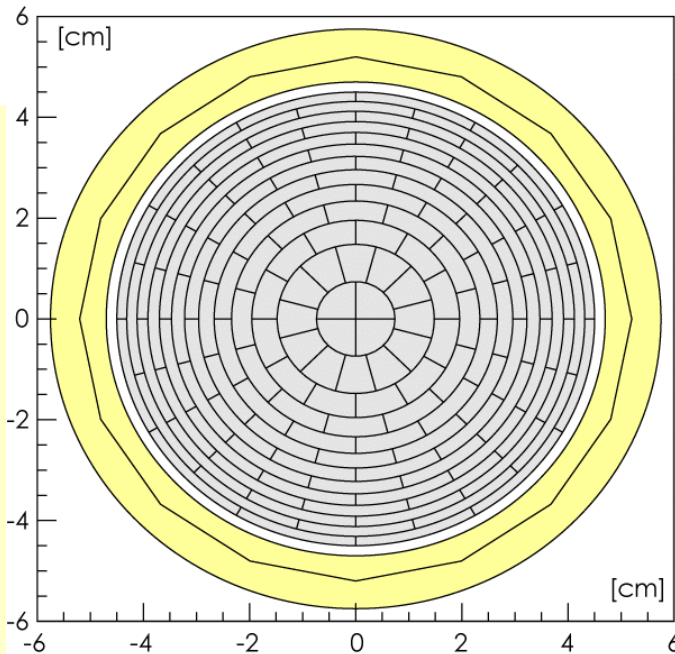
Main spectrometer:

- \varnothing 10 m, length 24 m
 \Rightarrow large energy resolution: $\Delta E = 0.93$ eV
 \Rightarrow high luminosity: $L = A_{\text{Seff}} \frac{\Delta\Omega}{4\pi} = A_{\text{analyse}} \frac{\Delta E}{(2E)} = 20 \text{ cm}^2$
- Variable retarding potential: -18.4–18.6 kV
- ultrahigh vacuum requirements (background) $p < 10^{-11}$ mbar (EHV)
- „simple“ construction: vacuum vessel at HV + „massless“ screening electrode

Detector setup



$A = 63 \text{ cm}^2$

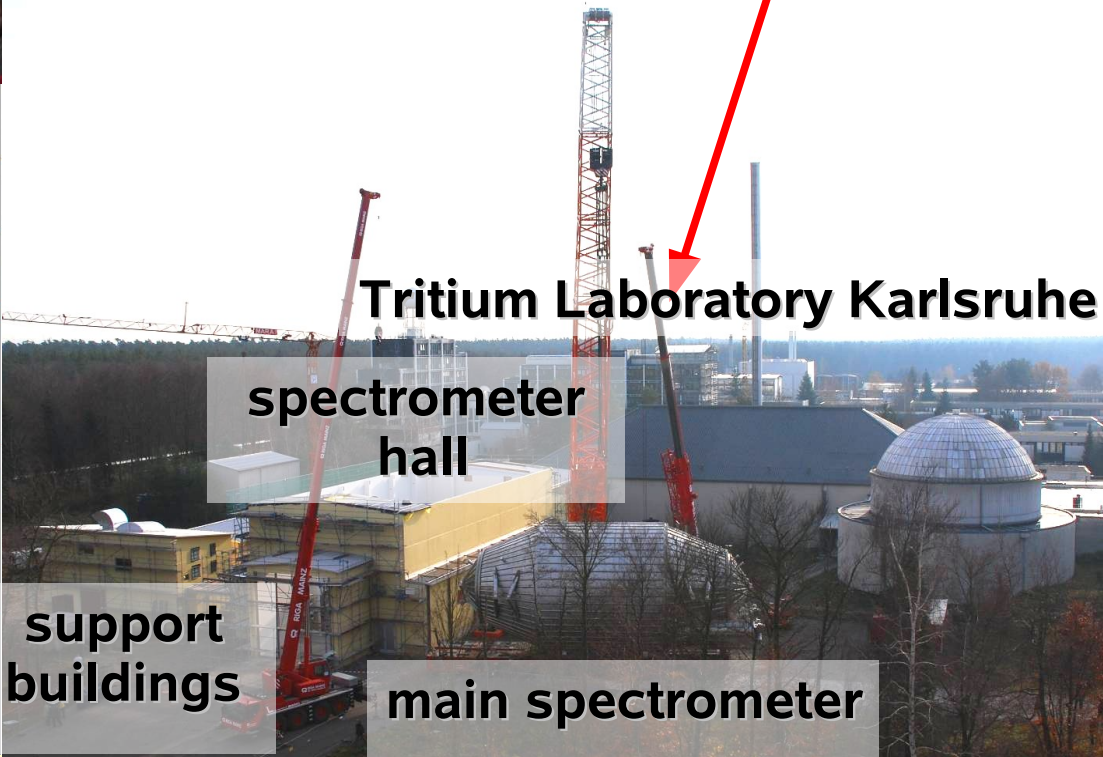
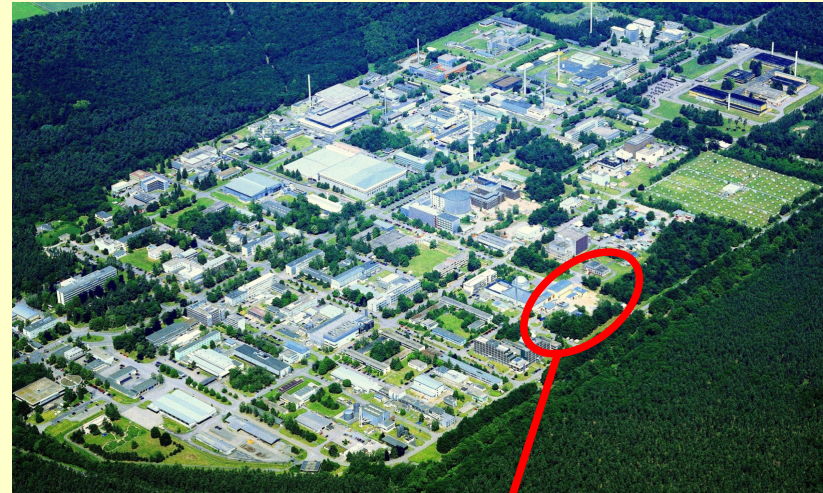


- Si-Pin diode
- Detection of transmitted β -decay electrons (mHz to kHz)
- **Low background for endpoint investigation**
- High energy resolution $\Delta E < 1 \text{ keV}$
- 12 rings with 30° segmentation + 4 fold center = **148 pixels**
→ record azimuthal and radial profile of flux tube

KATRIN @ Forschungszentrum Karlsruhe



Leopoldshafen, 25.11.06



Tritium Laboratory Karlsruhe

spectrometer hall

support buildings

main spectrometer

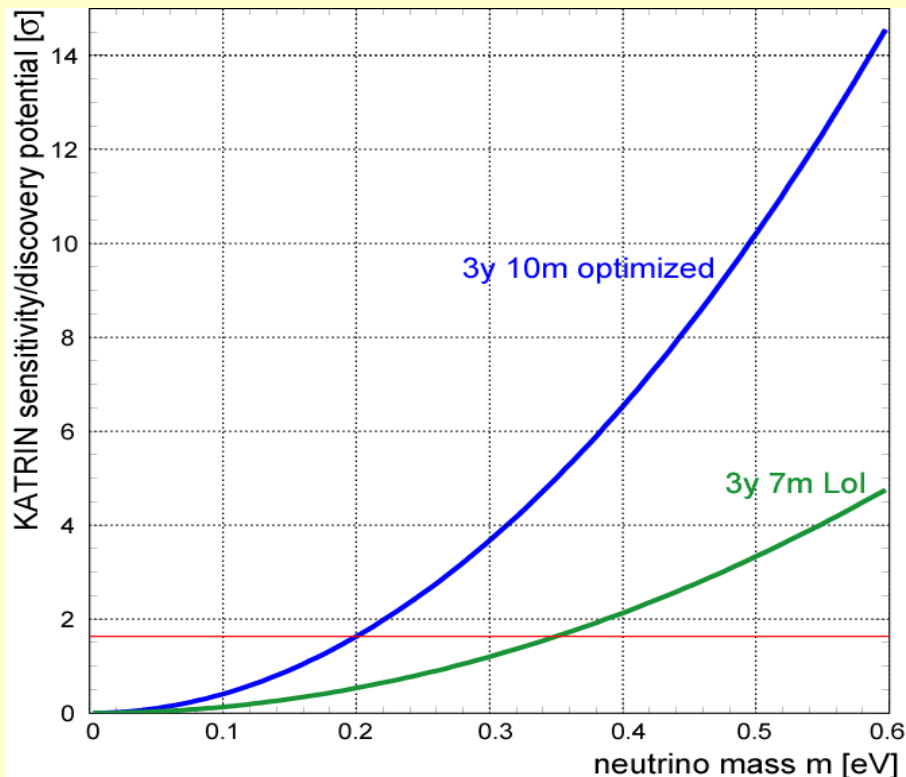


8800 km

KATRIN time schedule

KATRIN: 0.2 eV sensitivity:

- 2009-11 commissioning of main spectrometer and detector
- 2009-12 commissioning of tritium source and tritium elimination lines
- 2012- regular data taking for 5-6 years (3 full-beam-years)



Expectation for 3 full beam years: $\sigma_{\text{syst}} \sim \sigma_{\text{stat}}$

sensitivity:
 $m_\nu < 0.2\text{eV}$ (90%CL)

KATRIN pictures and plots: thanks to C. Weinheimer

A project for a new ^{187}Re experiment: MARE

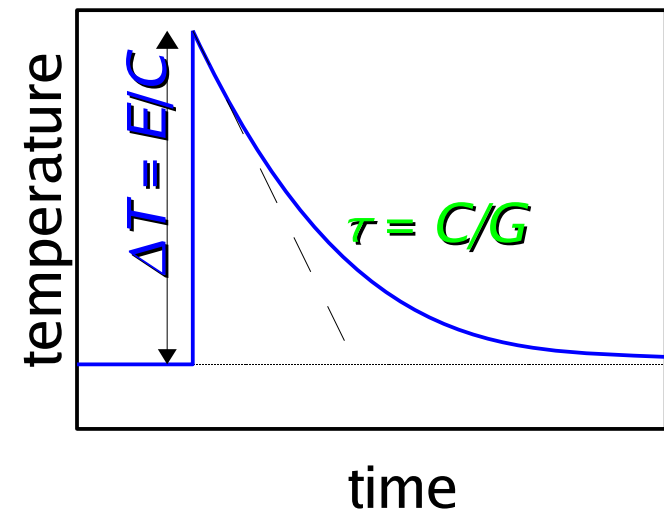
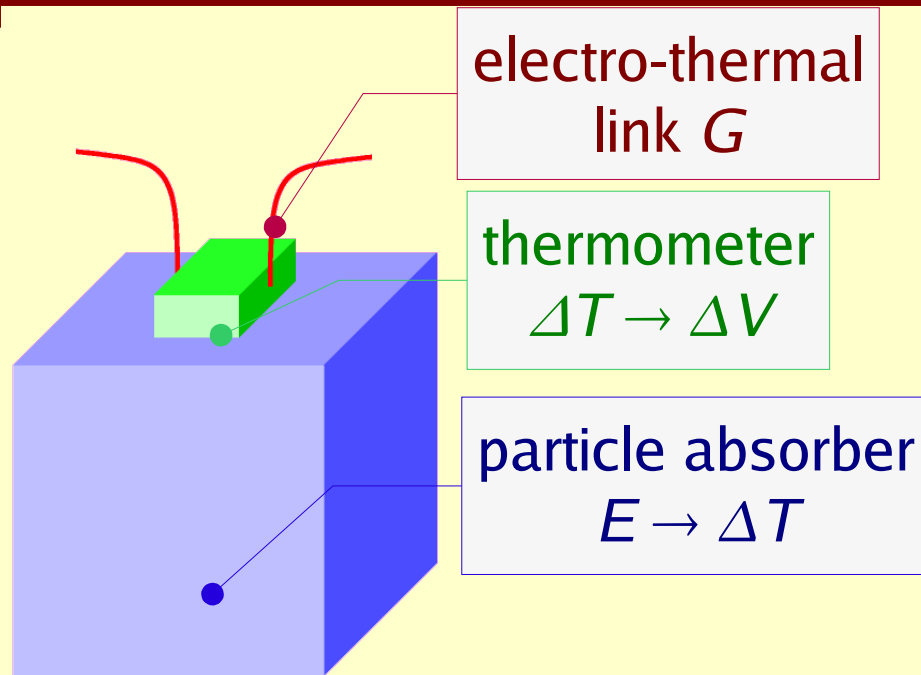
<http://crio.mib.infn.it/wig/silicini/proposal/>

MARE: Microcalorimeter Arrays for a Rhenium Experiment

Università di Genova, and INFN-Genova, Italy
Goddard Space Flight Center, NASA, Maryland, USA
Kirkhhof-Institute Physik, Universität Heidelberg, Germany
Università dell'Insubria, and INFN-Milano-Bicocca, Italy
Università di Milano-Bicocca, and INFN-Milano-Bicocca, Italy
NIST, Boulder, Colorado, USA
ITC-irst, Trento, and INFN-Padova, Italy
PTB, Berlin, Germany
University of Miami, Florida, USA
Università di Roma 'La Sapienza', and INFN-Roma1, Italy
SISSA, Trieste, Italy
Wisconsin University, Madison, Wisconsin, USA



Cryogenic detectors



Detection principle:

- $\Delta T = E/C$ where C is the total thermal capacity

⇒ low C : $C \sim T^3$ in dielectrics or superconductors below T_C

→ low T (i.e. $T \ll 1K$)

- time response: $\tau = C/G$

- ultimate limit to energy resolution:

statistical fluctuation of internal energy U : $\langle \Delta U^2 \rangle = k_B T^2 C$

- example: 1 mg of Re @ 100 mK
 $C \sim T^3$ (Debye) $\Rightarrow C \sim 10^{-13}$ J/K
6 keV X-ray $\Rightarrow \Delta T \sim 10$ mK
 $\Rightarrow \Delta U \sim 1$ eV
 $G \sim 10^{-11}$ W/K $\Rightarrow \tau \sim 10$ ms

- All the deposited energy is measured
- The detector is fully sensitive

¹⁸⁷Re experiment statistical sensitivity

$$\Sigma(m_\nu) \approx 20 \text{ eV}$$

1/10

$$\Sigma(m_\nu) = 2 \text{ eV}$$

1/10

$$\Sigma(m_\nu) = 0.2 \text{ eV}$$

- MIBETA detectors with $\Delta E_{\text{FWHM}} = 30 \text{ eV}$, $\tau_R = 1.5 \text{ ms}$

- ▷ $A_\beta = 0.15 \text{ Hz} \rightarrow f_{\text{pp}} = 2 \times 10^{-4}$
- ▷ $t = 3.6 \text{ y} \times \text{det} \rightarrow 1.6 \times 10^7 \text{ events}$
- ▷ $\Sigma_{\text{exp}}(m_\nu) = 15 \text{ eV}$

- detectors with $\Delta E_{\text{FWHM}} = 10 \text{ eV}$, $\tau_R = 100 \mu\text{s}$

- ▷ for $A_\beta = 0.3 \text{ Hz} \rightarrow f_{\text{pp}} = 3 \times 10^{-5}$
- ▷ $\Sigma_{\text{MC}}(m_\nu) = 2 \text{ eV}$ with $2 \times 10^{10} \text{ events}$
- ▷ $t = 2000 \text{ y} \times \text{det}$

- detectors with $\Delta E_{\text{FWHM}} = 1 \text{ eV}$, $\tau_R = 1 \mu\text{s}$

- ▷ for $A_\beta = 1 \text{ Hz} \rightarrow f_{\text{pp}} = 10^{-6}$
- ▷ $\Sigma_{\text{MC}}(m_\nu) = 0.2 \text{ eV}$ with $2.5 \times 10^{13} \text{ events}$
- ▷ $t = 8 \times 10^5 \text{ y} \times \text{det}$

The MARE project

Goal: a **sub-eV** direct neutrino mass measurement complementary to the KATRIN experiment

MARE is divided in two phases:

MARE-1

MARE-2

new experiments with large arrays using available technology and ready to start as soon as possible

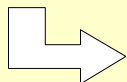
2÷4 eV m_ν sensitivity

very large experiment with a m_ν statistical sensitivity close to KATRIN but still improvable: 5 years from now for further detector R&D

0.2 eV m_ν sensitivity

phase I is needed:

- because it's the only possible one with present technology
- to investigate systematics in thermal calorimeters



very important to cross-check spectrometer results

MARE-1: TES vs. Silicon thermistors

- **aim: high statistics measurement with a ready-to-use technology**
 - ▷ few eV statistical sensitivity in few years
 - ▷ investigate systematics in thermal calorimeters with $10^9 \div 10^{10}$ events
 - ▷ cross-check spectrometer results

MARE-1 SEMICON (MIBETA2)

U. Milano-Bicocca / INFN Sez. Mi-Bicocca
U. Insubria / INFN Sez. Mi-Bicocca
ITC-Irst / INFN Sez. Padova
U. Wisconsin, Madison
NASA/Goddard

- about 300 element arrays
- well known Si implanted thermistors
- AgReO_4 crystals

MARE-1 TES (MANU2)

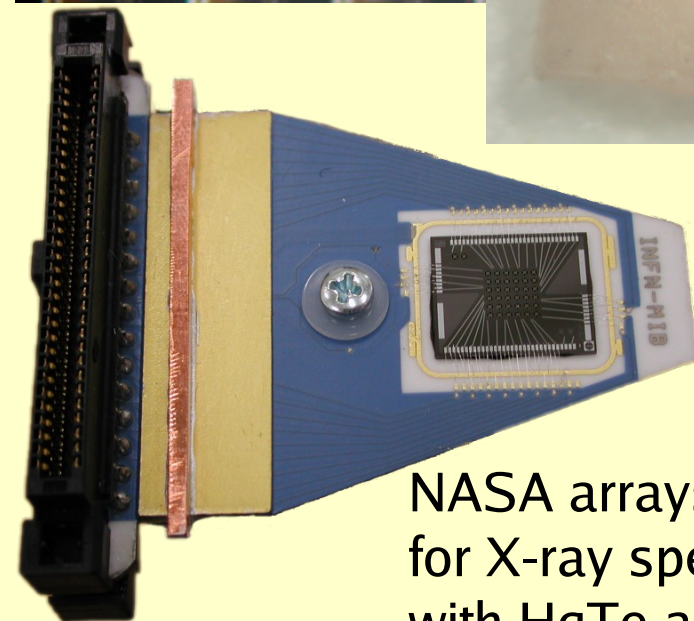
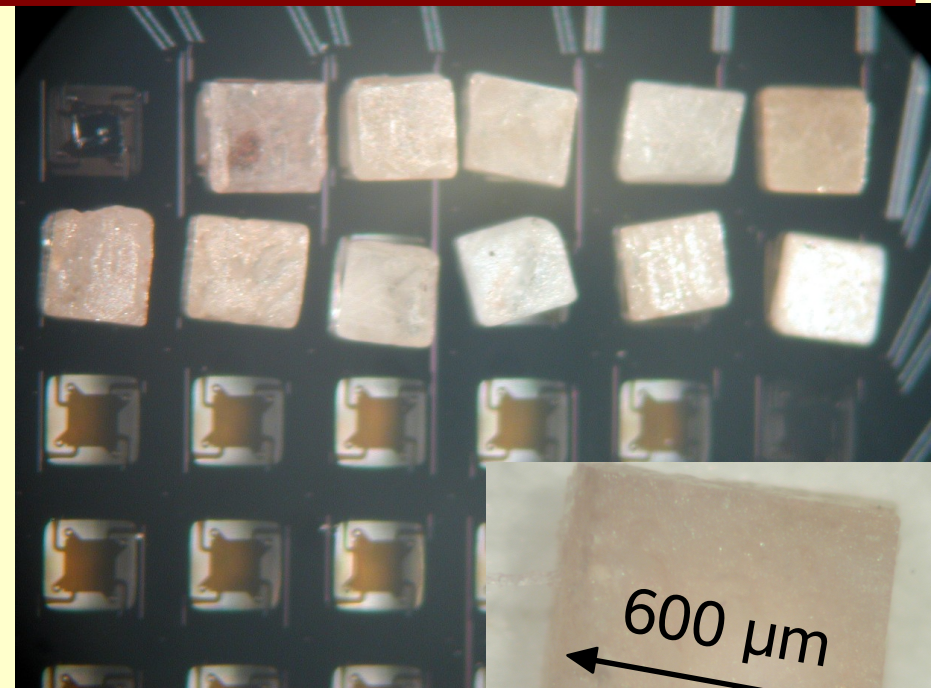
U. Genova / INFN Sez. Genova
U. Miami, Florida
PTB Berlin, Germany

- about 300 element arrays
- newly developed transition edge sensors
- Re crystals

- ▷ cross check
- ▷ common effort on systematics
- ▷ joint analysis to improve limit

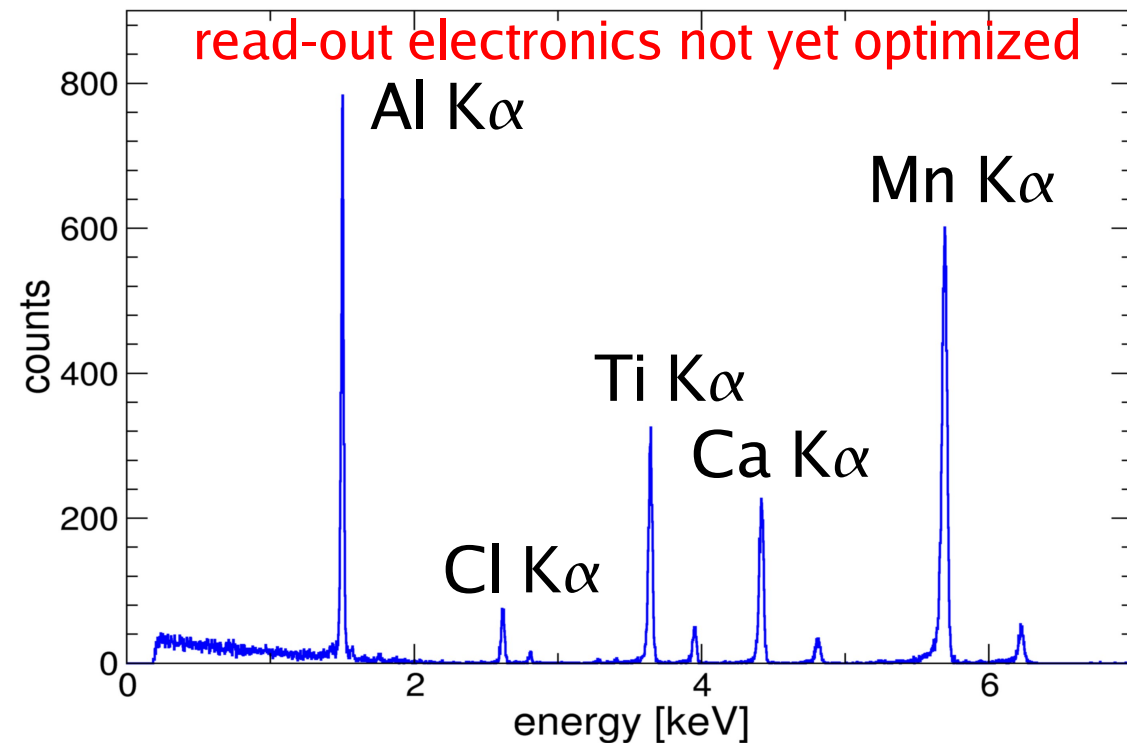
MARE-1 SEMICON

- NASA/GSFC XRS2-2 arrays
 - ▷ 6x6 pixels
- flat AgReO_4 single crystals
 - ▷ $m \approx 0.5$ mg
- detector R&D phase results
 - ▷ best operating $T \approx 90\text{mK}$
 - ▷ $\Delta E \approx 30$ eV, $\tau_R \approx 250$ μs



NASA arrays optimized for X-ray spectroscopy with HgTe absorber → ASTRO-E2 mission

read-out electronics not yet optimized



MARE-1 SEMICON: MC statistical sensitivity

year	1	2	3	4
new detectors	72	216	0	0
total detectors	72	288	288	288
statistics [det*y]	72	360	648	936
activity [c/s]	0.27	$m_{\text{AgReO}_4} = 500 \mu\text{g}$		
statistics [events]	6.10E+08	3.05E+09	5.49E+09	7.94E+09

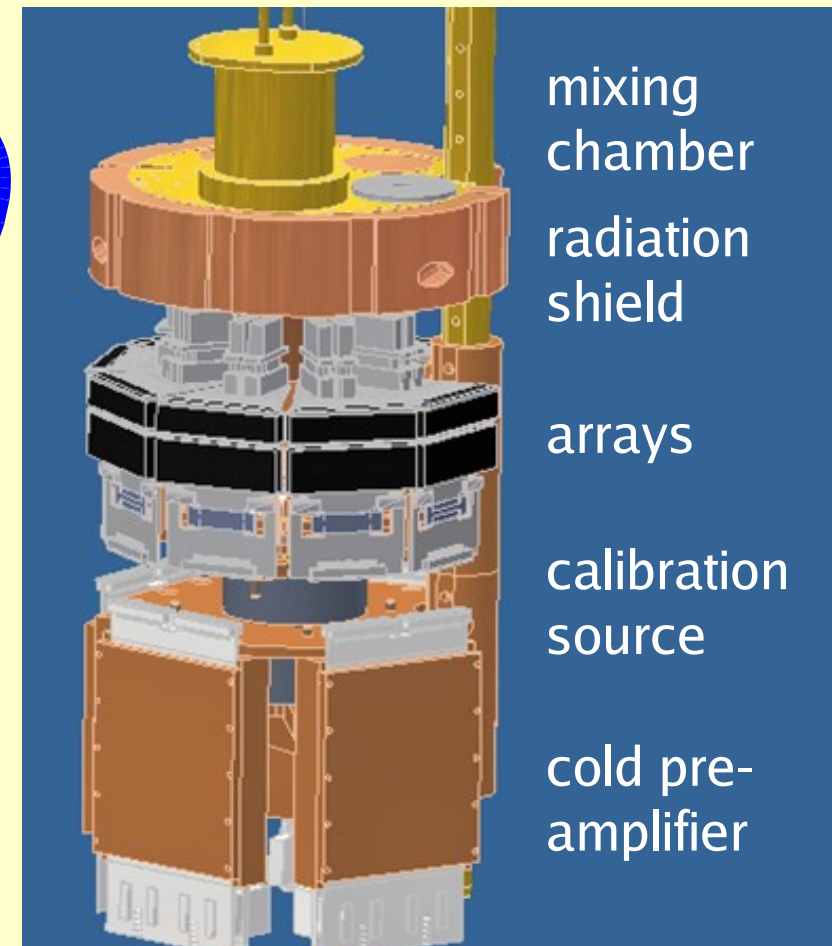
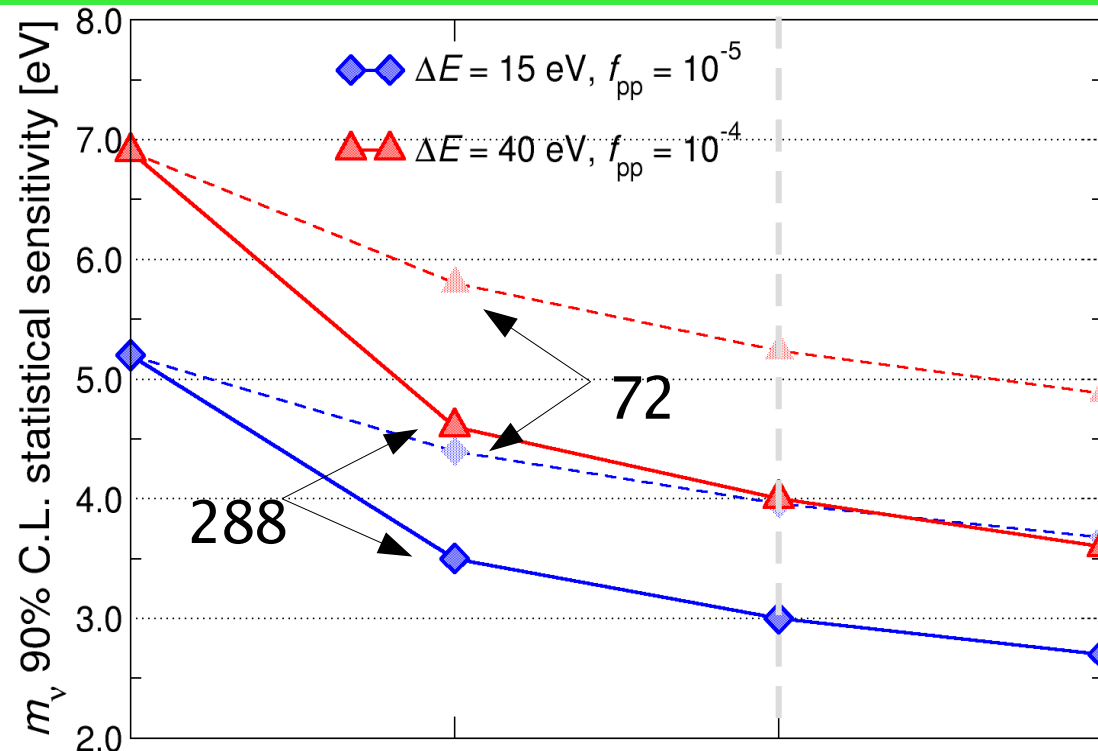
- setup ready for 8 arrays
- 288 AgReO_4 crystals
- now starting with 2 arrays (72 ch.)
- gradual deployment
- ▷ further detector optimization

$\Delta E = 40 \text{ eV}$ $\tau = 400 \mu\text{s}$ $f_{\text{pp}} = 1.0\text{E-}4$

m_ν sensitivity (90%)	6.9	4.6	4.0	3.6
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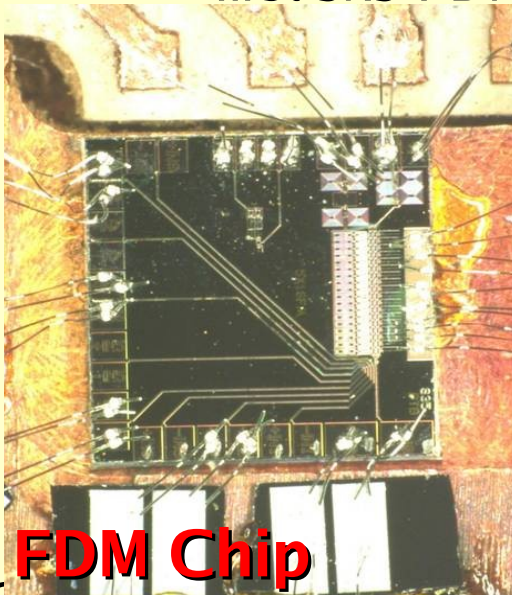
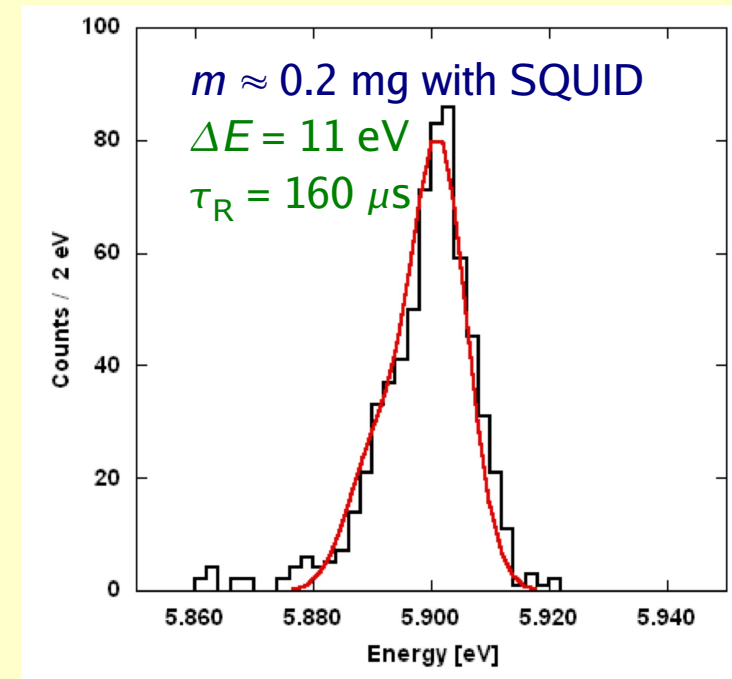
$\Delta E = 15 \text{ eV}$ $\tau = 50 \mu\text{s}$ $f_{\text{pp}} = 1.0\text{E-}5$

m_ν sensitivity (90%)	5.2	3.5	3.0	2.7
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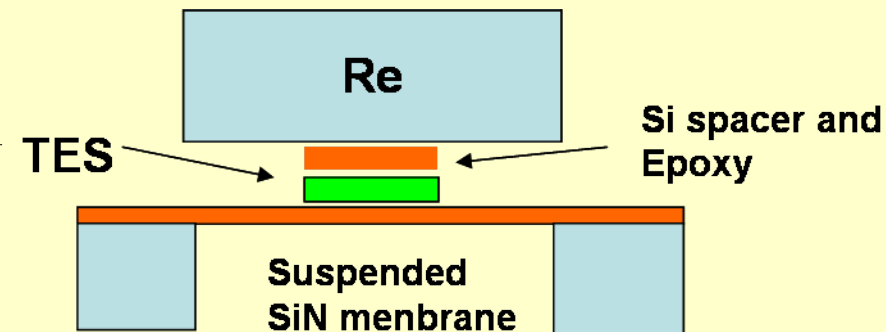
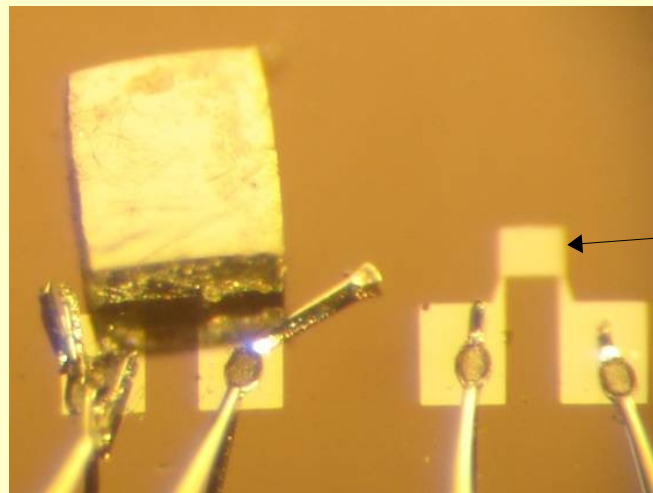


MARE-1 TES

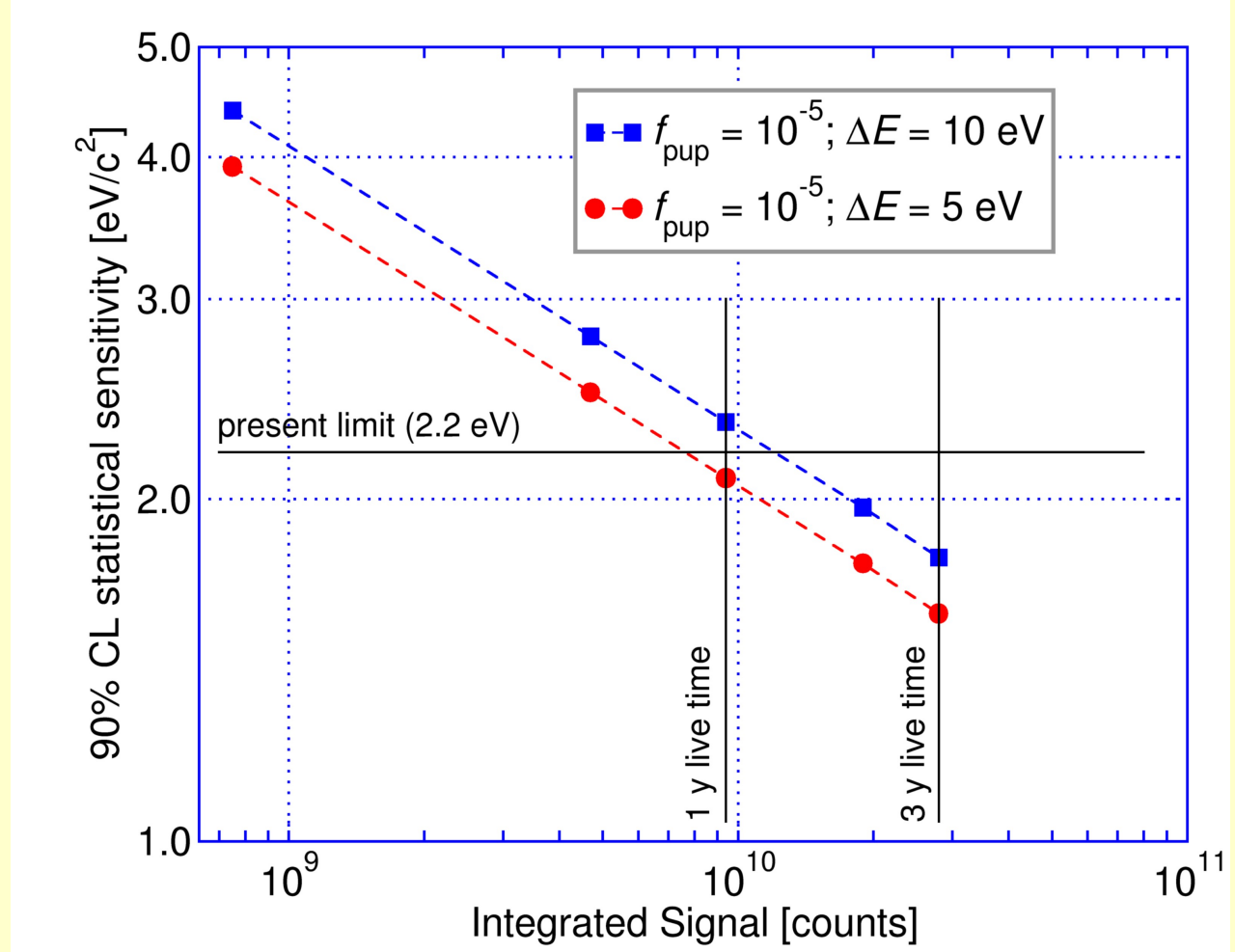
- Pulsed Laser Deposition of thin films: pure **Ir** or **Ir bilayers**
- detectors with **metallic rhenium** absorbers
- **300 channel array**
- detector R&D goal:
 - ▶ **1 mg Re crystals with: $\Delta E = 5$ eV, $\tau_R = 10$ μ s**
 - ▶ a further step towards MARE-2
- two read-out options
 - ▶ JFETs with **cold impedance transformer**
 - ▶ frequency multiplexed SQUIDs (FDM)
 - **first 3x3 FDM chip is under test now**



FDM Chip



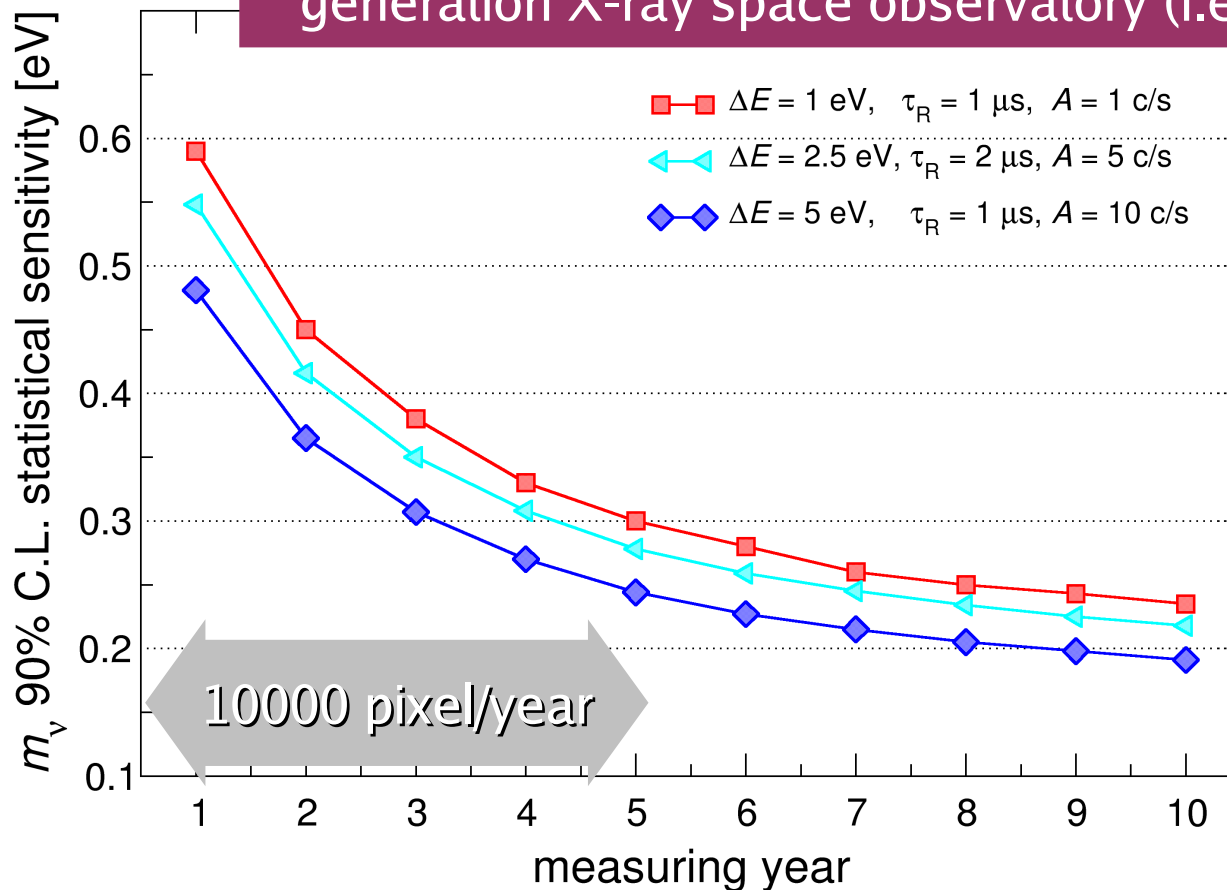
MARE-1 TES statistical sensitivity



- 300 rhenium crystals in 2 refrigerators
 - ▷ $m \approx 1 \text{ mg}$
- Ir/Au or Al/Ag TES at 100 mK
 - ▷ $\Delta E = 10 \text{ eV}, \tau_R = 10 \mu\text{s}, f_{\text{pp}} = 10^{-5}$
 - ▶ about 3×10^{10} events in 3 years $\Rightarrow m_\nu < 1.8 \text{ eV}$

MARE-2

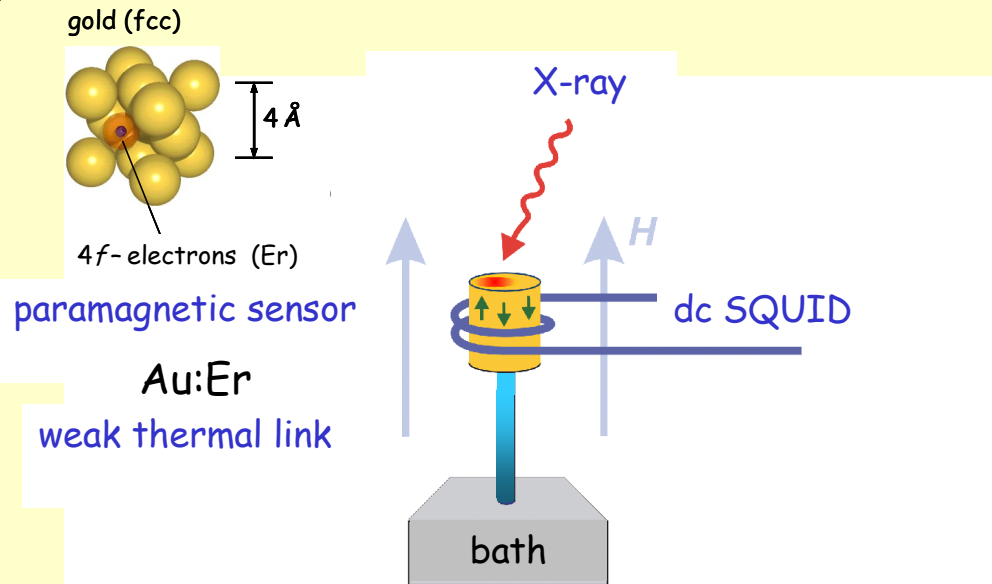
- only statistical analysis
- 50000+ detectors gradually deployed
 - ▷ 5 arrays with 10000 detectors each
 - ▷ one array deployed per year for the first 5 years
 - ▷ arrays distributed in many laboratories around the world
 - ▷ about $10^{13} \div 10^{14}$ events after 5 years
- technical requirements not far from that for next generation X-ray space observatory (i.e. XEUS, Con-X)



10000 pixel *kits*
 $\Delta E \approx 1 \text{ eV}$
 $\tau_R \approx 1 \mu\text{s}$
 $A_\beta \approx 1 \div 10 \text{ Hz}$

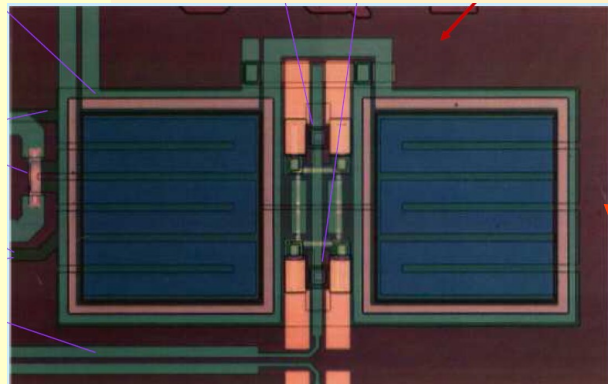
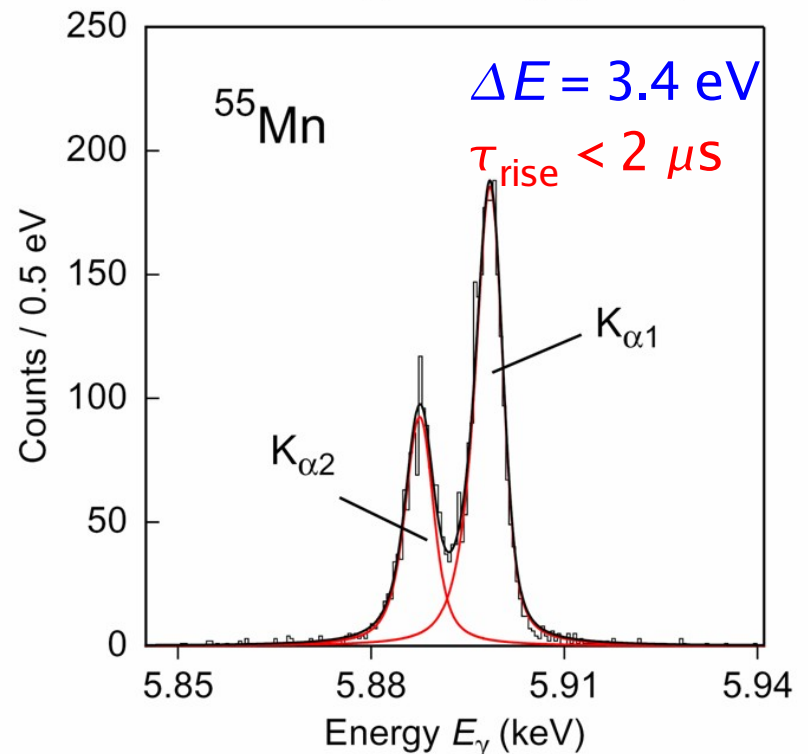
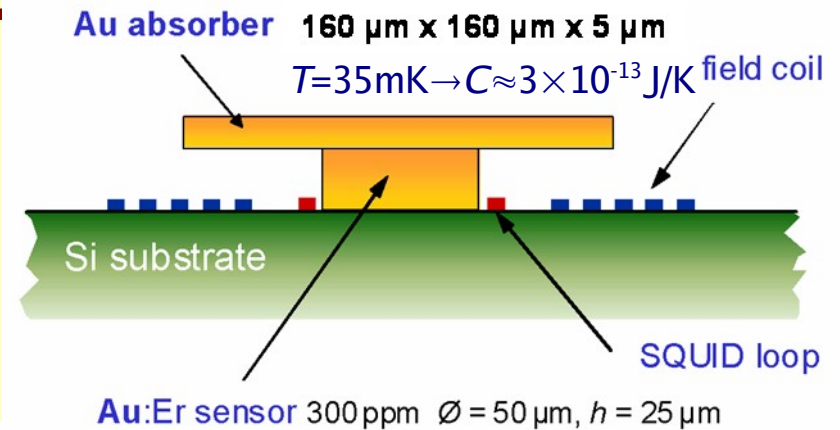
need for
new sensor R&D
and
new read-out techniques

MMC - Magnetic Micro Calorimeters (Heidelberg)



$$\delta M = \frac{\partial M}{\partial T} \delta T = \frac{\partial M}{\partial T} \frac{E_\gamma}{C_{ges}}$$

- ▶ suitable for large capacity absorbers
- ▶ very fast $\sim \mu s$
- ▶ high energy resolution $\sim eV$

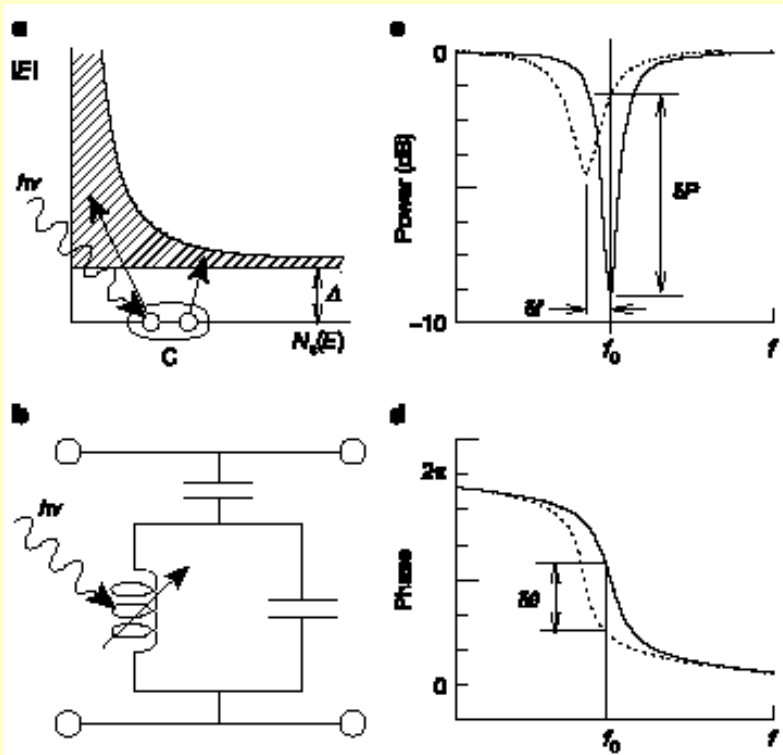


sensor design optimization for MARE-2
rhenium absorbers in progress
⇒ meander pick-up coils without external B field

MKIDs – Multiplexed Kinetic Inductance Detectors

- resonator exploiting the T dependence of inductance in a superconducting film

- ▶ **qp detectors** suitable for large absorbers
- ▶ **fast** devices for high single pixel activity A_β and low pile-up f_{pp}
- ▶ **high energy resolution**
- ▶ **multiplexing** for very large number of pixel



Sensitivity

$$\Delta E = 5 \text{ eV}$$

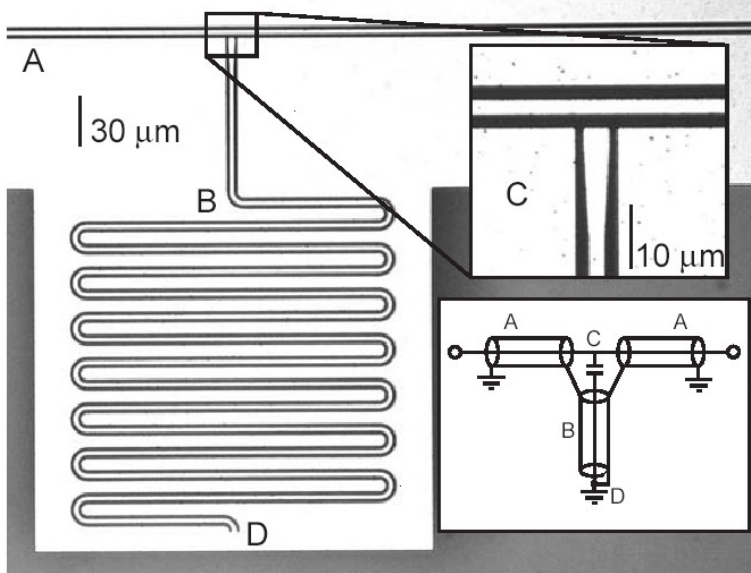
$$t_M = 36000 \text{ detectors} \times 3 \text{ years}$$

$$A_\beta = 20 \text{ c/s/det}$$

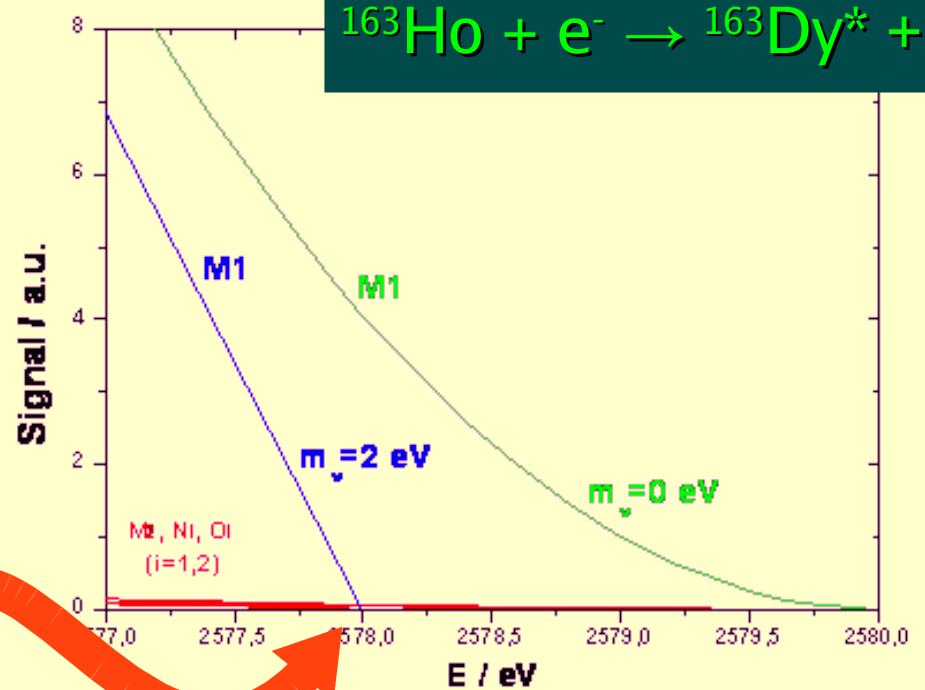
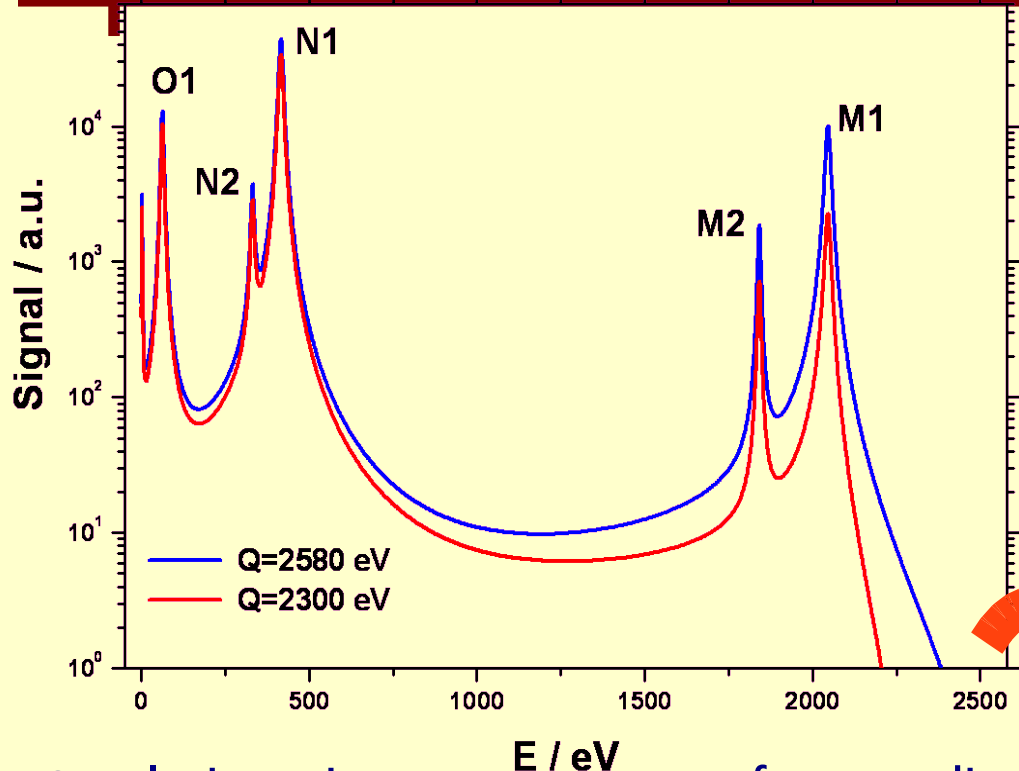
$$\bullet \tau_{\text{rise}} = 1 \mu\text{s} \Rightarrow m_\nu < 0.2 \text{ eV}$$

$$\bullet \tau_{\text{rise}} = 100 \mu\text{s} \Rightarrow m_\nu < 0.4 \text{ eV}$$

- KIDs developed for astrophysics
- application to bulky absorber still requires further efforts



^{163}Ho electron capture measurement



- calorimetric measurement of non-radiative Dy atomic de-excitations (Coster-Kronig, Auger...)
- fraction of events at end-point may be as high as for ^{187}Re : depends on Q_{EC} (≈ 2.5 keV)

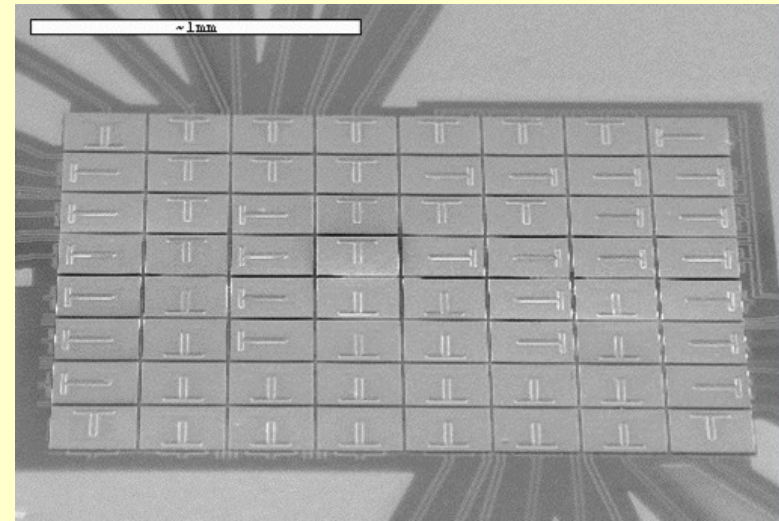
► $Q_{\text{EC}}?$

- fewer active nuclei are needed ($\tau \approx 4000$ y)

► can be implanted in any suitable absorber

► first implantation tests at ISOLDE are encouraging

- new NASA/Goddard TES arrays ($\Delta E = 2\text{eV}$) can be implanted with ^{163}Ho



Conclusions

Investigation of β -decay kinematics is the only model-independent approach to the absolute neutrino mass scale.

KATRIN is the ultimate tritium β -decay experiment: it will reach a sensitivity of 0.2 eV on m_ν . Expected data taking in 2012.

^{187}Re calorimetry is complementary to tritium experiments and can give sub-eV sensitivity to m_ν .

The **MARE** project 1st phase is just starting. R&D improvements on the detector technology are crucial for the 2nd phase.