χ^2 and Goodness of Fit

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CERN Academic Training Course

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Least squares best fit What is σ ? Resume of straight line Correlated errors Goodness of fit with χ^2 Number of Degrees of Freedom Other G of F methods Errors of first and second kind **Combinations** THE paradox

Least Squares Straight Line Fitting

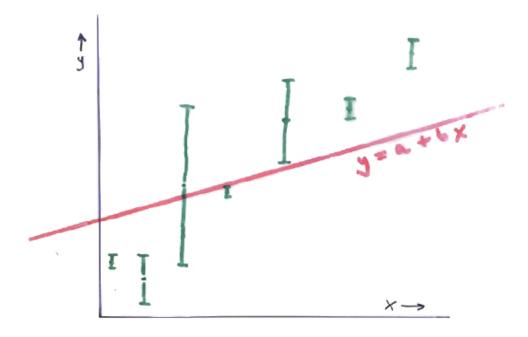
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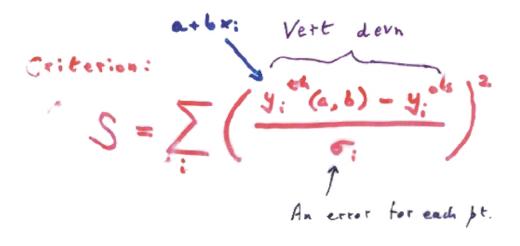
1) Does it fit straight line? (Goodness of Fit)

2) What are gradient and intercept? (Parameter Determination) Do 2) first

N.B.1 Can be used for non "a+bx" e.g. $a + b/x + c/x^2$ N.B.2 Least squares is not the only method

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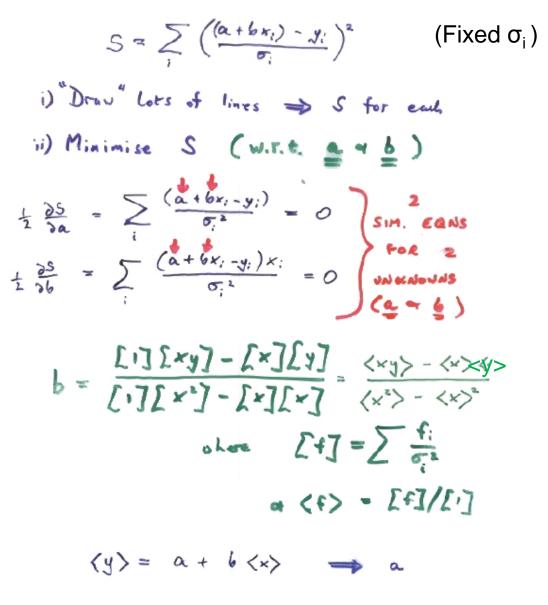
If theory and data OK: $y^{th} \sim y^{obs} \rightarrow S$ small Minimise $S \rightarrow$ best line Value of $S_{min} \rightarrow$ how good fit is

Which σ should we use?

Which o?	Exptl o	Theory σ
Name	Neyman	Pearson
Ease of algebra	Easier, so this version is used more	
If Th = 0.01, Exp = 1	Contributes 1 to S	Contributes 98 to S More plausible
S ~ χ^2 ?		More plausible
S = $(\hat{a} - a_1)^2 / \sigma^2 + (\hat{a} - a_2)^2 / \sigma^2$	Biassed down because smaller $a_i \rightarrow smaller \sigma$	Biassed up because larger $\hat{a} \rightarrow$ larger σ

(For $\hat{a} \sim a_i$, and both much larger than σ_i , 2 methods are very similar)

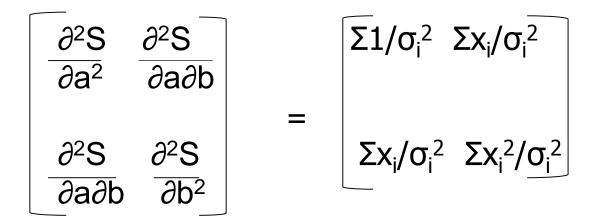
Straight Line Fit



N.B. L.S.B.F. passes through (<x>, <y>)

Correlated intercept and gradient?

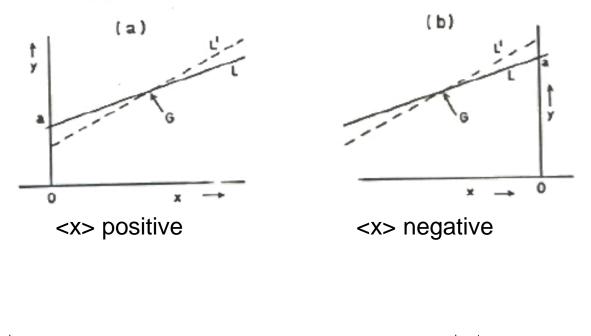
2 * Inverse covariance matrix =

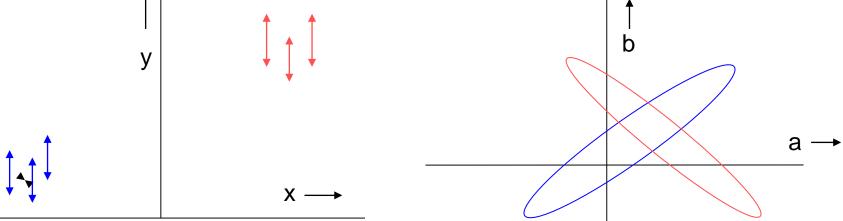


Invert \rightarrow Covariance matrix Covariance $\sim -\Sigma x_i / \sigma_i^2 = [x]$ If measure intercept at weighted c. of g. of x for data points, cov = 0 i.e. gradient and intercept there are uncorrelated

So track params are usually specified at centre of track.

Covariance(a,b) ~ -<x>





Measurements with correlated errors e.g. systematics?

In general :
$$S = \sum_{ij} \widetilde{\Delta}_i H_{ij} \Delta_j$$

where $\Delta_j = (observe - pred.)_j$

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Comments on Least Squares method

1) Need to bin

Beware of too few events/bin

2) Extends to n dimensions

but needs lots of events for n larger than 2 or 3

3) No problem with correlated uncertainties

4) Can calculate S_{min} "on line" i.e. single pass through data

$$\Sigma (y_i - a - bx_i)^2 / \sigma^2 = [y_i^2] - b [x_iy_i] - a [y_i]$$

5) For theory linear in params, analytic solution

6) Goodness of Fit

$$\bigstar \bigstar \bigstar$$

	Individual events (e.g. in cos θ)	y _i ±σ _i ∨ x _i (e.g. stars)	
1) Need to bin?	Yes	No need	
4) χ^2 on line	First histogram	Yes	10

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2.		•		•		

	Moments	Max Like	Least squares
Easy?	Yes, if	Normalisation, maximisation messy	Minimisation
Efficient?	Not very	Usually best	Sometimes = Max Like
Input	Separate events	Separate events	Histogram
Goodness of fit	Messy	No (unbinned)	Easy
Constraints	No	Yes	Yes
N dimensions	Easy if	Norm, max messier	Easy
Weighted events	Easy	Errors difficult	Easy
Bgd subtraction	Easy	Troublesome	Easy
Inverse covariance matrix	Observed spread, or analytic	$ \left\{ -\frac{\partial^2 \ln \mathcal{L}}{\partial p_i \partial p_j} \right\} $	$\left\{\frac{\partial^2 S}{2\partial p_i \partial p_j}\right\}$
Main feature	Easy	Best	Goodness of Fit

Goodness of Fit: χ^2 test

- 1) Construct S and minimise wrt free parameters
- 2) Determine v = no. of degrees of freedom

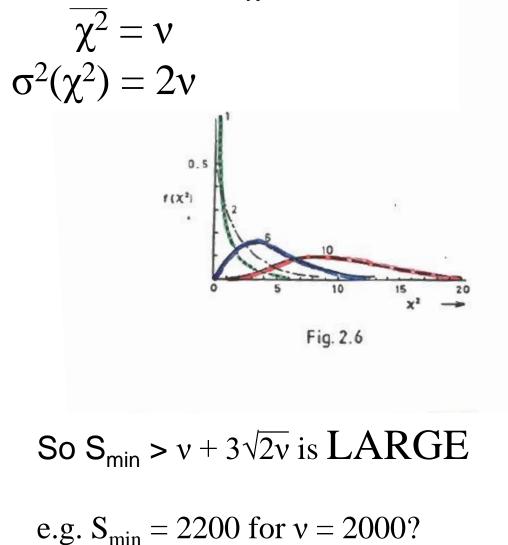
v = n - p n = no. of data points p = no. of FREE parameters

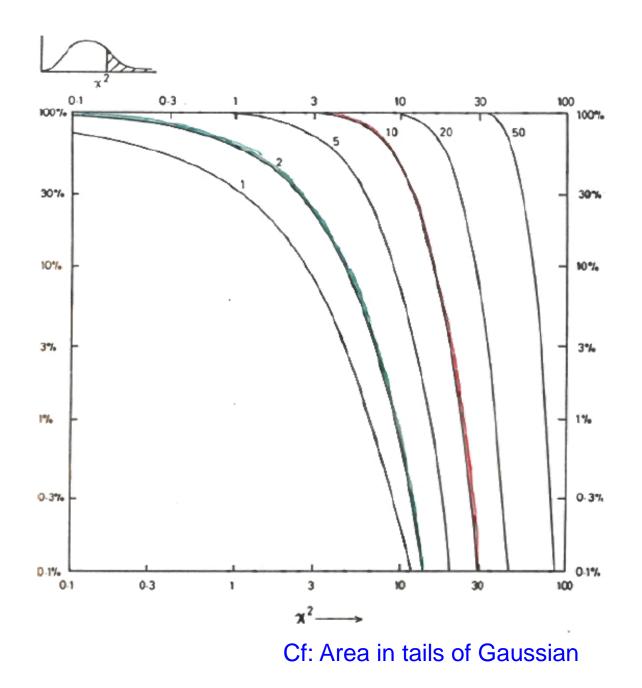
3) Look up probability that, for v degrees of freedom, $\chi^2 \ge S_{min}$

Uses i) Poisson ~ Gaussian if expected number not too small ii) For N y_i distributed as Gaussian N(0,1), $\Sigma y_i^2 \sim \chi^2$ with ndf = N

So works ASYMPTOTICALLY. Otherwise use MC for dist of S (or binned \mathcal{L})

Properties of mathematical χ^2 distribution:





χ^2 with v degrees of freedom?

v = data - free parameters ?

Why asymptotic (apart from Poisson \rightarrow Gaussian)? a) Fit flatish histogram with $y = N \{1 + 10^{-6} \cos(x - x_0)\}$ $x_0 = \text{free param}$

b) Neutrino oscillations: almost degenerate parameters $y \sim 1 - A \sin^2(1.27 \Delta m^2 L/E)$ 2 parameters $\xrightarrow{} 1 - A (1.27 \Delta m^2 L/E)^2$ 1 parameter Small Δm^2

Goodness of Fit

 χ2 Very general Needs binning Not sensitive to sign of deviation

Run Test



Kolmogorov-Smirnov



Aslan and Zech `Energy Test' Durham IPPP Stats Conf (2002)

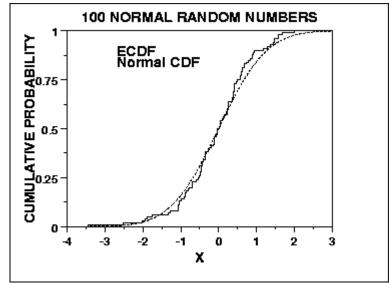
Binned *L*ikelihood (= Baker-Cousins}

etc

Goodness of Fit: Kolmogorov-Smirnov

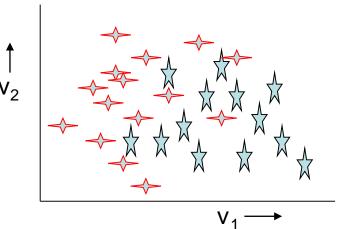
Compares data and model cumulative plots (or 2 sets of data) Uses largest discrepancy between dists. Model can be analytic or MC sample

Uses individual data points Not so sensitive to deviations in tails (so variants of K-S exist) Not readily extendible to more dimensions Distribution-free conversion to p; depends on n (but not when free parameters involved – needs MC)



Goodness of fit: 'Energy' test

Assign +ve charge to data \rightarrow ; -ve charge to M.C. Calculate 'electrostatic energy E' of charges If distributions agree, E ~ 0 If distributions don't overlap, E is positive Assess significance of magnitude of E by MC



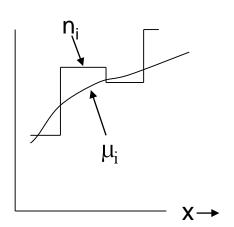
N.B.

- 1) Works in many dimensions
- 2) Needs metric for each variable (make variances similar?)
- 3) $E \sim \Sigma q_i q_j f(\Delta r = |r_i r_j|)$, $f = 1/(\Delta r + \varepsilon)$ or $-\ln(\Delta r + \varepsilon)$

Performance insensitive to choice of small ϵ

See Aslan and Zech's paper at: http://www.ippp.dur.ac.uk/Workshops/02/statistics/program.shtml

Binned data and Goodness of Fit using *L*-ratio



For histogram, uses Poisson prob $P(n;\mu)$ for n observed events when expect μ .

Construct \mathcal{L} -ratio = Product{P(n_i; μ_i)/P(n_i; μ =n_i)} P(n_i; μ =n_i) is best possible μ for that n_i Need denoms because P(100;100.0) very different from P(1;1.0)

-2* \mathcal{L} ratio ~ χ^2 when μ_i large and $n_i \sim \mu_i$ Better than Neyman or Pearson χ^2 when μ_i small

Baker and Cousins, NIM 221 (1984) 437

Wrong Decisions

Error of First Kind

Reject H0 when true Should happen x% of tests

Errors of Second Kind

Accept H0 when something else is true Frequency depends on i) How similar other hypotheses are e.g. $H0 = \mu$ Alternatives are: $e \quad \pi \quad K \quad p$ ii) Relative frequencies: $10^{-4} \ 10^{-4} \ 1 \quad 0.1 \quad 0.1$

 Aim for maximum efficiency ← Low error of 1st kind maximum purity ← Low error of 2nd kind
 As χ² cut tightens, efficiency ↑ and purity↓
 Choose compromise

How serious are errors of 1st and 2nd kind?

Result of experiment

 e.g Is spin of resonance = 2?
 Get answer WRONG

 Where to set cut?

 Small cut ⇒ Reject when correct
 Large cut ⇒ Never reject anything

 Depends on nature of H0 e.g.

 Does answer agree with previous expt?
 Is expt consistent with special relativity?

2) Class selector e.g. b-quark / galaxy type / γ-induced cosmic shower Error of 1st kind: Loss of efficiency Error of 2nd kind: More background Usually easier to allow for 1st than for 2nd

3) Track finding

Combining: Uncorrelated exptl results

Simple Example of Minimising S

Measurements a, ±5; Treasurements $a_1 \pm 6$, $a_2 \pm 6_2$ $a_1 \pm 6$; $a_2 \pm 6_2$ $a_2 \pm 6_2$ $a_3 \pm 6$; $a_4 \pm 6$; $a_1 \pm 6$; $a_2 \pm 6$; $a_1 \pm 6$; $a_2 \pm 6$; $a_3 \pm 6$; $a_4 \pm 6$; $a_5 \pm 6$; $a_1 \pm 6$; $a_2 \pm 6$; $a_3 \pm 6$; $a_4 \pm 6$; $a_1 \pm 6$; $a_2 \pm 6$; $a_3 \pm 6$; $a_4 \pm 6$; $a_5 \pm 6$; Minimise S w.r.t. à $\frac{1}{2}\frac{\partial S}{\partial x} = \sum \frac{\hat{a} - A}{\sigma^2} = 0$ à Zon = Zaiz ₩ Error on à quien by $\sigma = \left(\frac{1}{2}\frac{\partial S}{\partial a^2}\right)^{-k_1}$ IN PARABOLIC AH EQUIN TO 23 = 2 2 52 S -> S ... + 1 $\frac{1}{\sigma^2} = \sum \frac{1}{\sigma^2} \Phi$ ing parame $\frac{1}{2} \frac{3^2 S}{3 k_i \partial k_j} = 10 VERSE ERAOR S$ Many parame

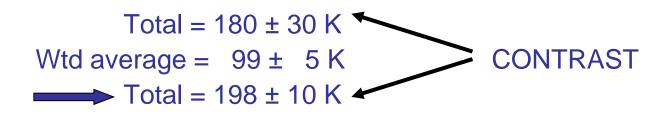
N.B. Better to combine data rather than results

So $\hat{a} = \sum w_i a_i / \sum w_i$, where $w_i = 1/\sigma_i^2$

Difference between weighted and simple averaging

Isolated island with conservative inhabitants How many married people ?

Number of married men $= 100 \pm 5 \text{ K}$ Number of married women $= 80 \pm 30 \text{ K}$



GENERAL POINT: Adding (uncontroversial) theoretical input can improve precision of answer Compare "kinematic fitting"

BLUE

Best Linear Unbiassed Estimate

Combine several possibly correlated estimates of same quantity e.g. V₁, V₂, V₃ σ_1^2 COV_{12} **Covariance matrix** COV_{13} cov_{12} σ_2^2 cov_{13} cov_{23} COV₂₃ Uncorrelated Positive correlation Negative correlation $cov_{ii} = \rho_{ii} \sigma_i \sigma_i$ with $-1 \le \rho \le 1$ Lyons, Gibault + Clifford

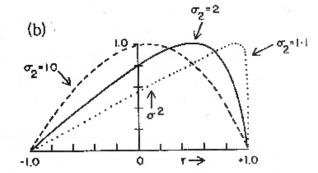
NIM A270 (1988) 42

$$\begin{split} v_{best} &= w_1 v_1 + w_2 v_2 + w_3 v_3 & \text{Linear} \\ \text{with} \quad w_1 + w_2 + w_3 = 1 & \text{Unbiassed} \\ \text{to give } \sigma_{best} &= \min \left(\text{wrt } w_1, w_2, w_3 \right) & \text{Best} \\ \text{For uncorrelated case, } w_i \sim 1/\sigma_i^2 & \text{For correlated pair of measurements with } \sigma_1 < \sigma_2 \\ v_{best} &= \alpha v_1 + \beta v_2 & \beta = 1 - \alpha \\ \beta &= 0 \text{ for } \rho = \sigma_1/\sigma_2 \\ \beta &< 0 \text{ for } \rho > \sigma_1/\sigma_2 & \text{i.e. extrapolation!} & \text{e.g. } v_{best} = 2v_1 - v_2 \end{split}$$

Extrapolation is sensible:

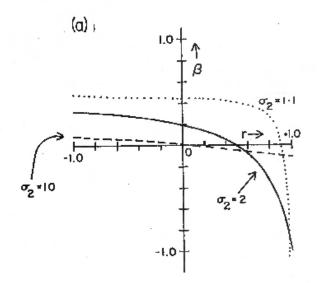
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V _{true}	V ₁	V_2	



Beware extrapolations because

[b] σ_{best} tends to zero, for $\rho = +1$ or -1



[a] v_{best} sensitive to ρ and σ_1/σ_2

N.B. For different analyses of ~ same data, ρ ~ 1, so choose 'better' analysis, rather than combining

N.B. σ_{best} depends on σ_1 , σ_2 and ρ , but not on $v_1 - v_2$ e.g. Combining 0±3 and x±3 gives x/2 ± 2

$$\mathsf{BLUE} = \chi^2$$

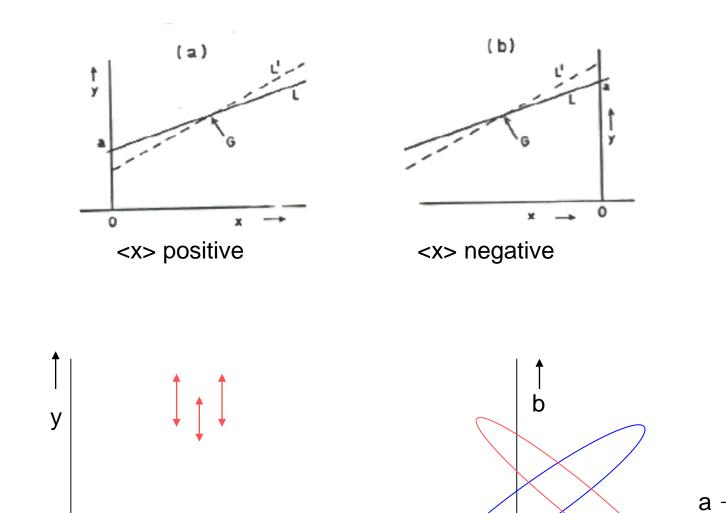
 $S(v_{best}) = \Sigma (v_i - v_{best}) E^{-1}_{ij} (v_j - v_{best})$, and minimise S wrt v_{best} S_{min} distributed like χ^2 , so measures Goodness of Fit But BLUE gives weights for each v_i

Can be used to see contributions to σ_{best} from each source of uncertainties e.g. statistical and systematics

different systematics

For combining two or more possibly correlated measured quantities {e.g. intercepts and gradients of a straight line), use χ^2 approach. Alternatively. Valassi has extended BLUE approach Covariance(a,b) ~ -<x>

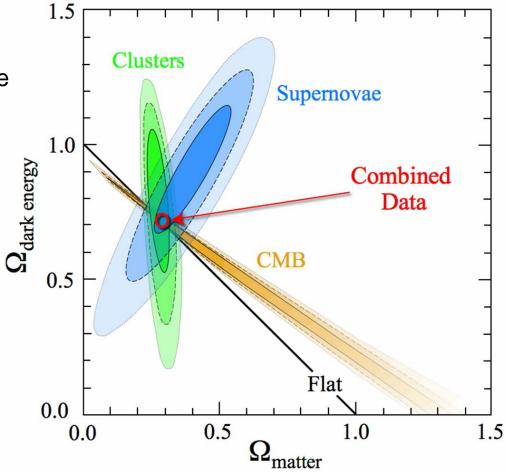
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Uncertainty on $\Omega_{dark energy}$

When combining pairs of variables, the uncertainties on the combined parameters can be much smaller than any of the individual uncertainties e.g. $\Omega_{dark energy}$



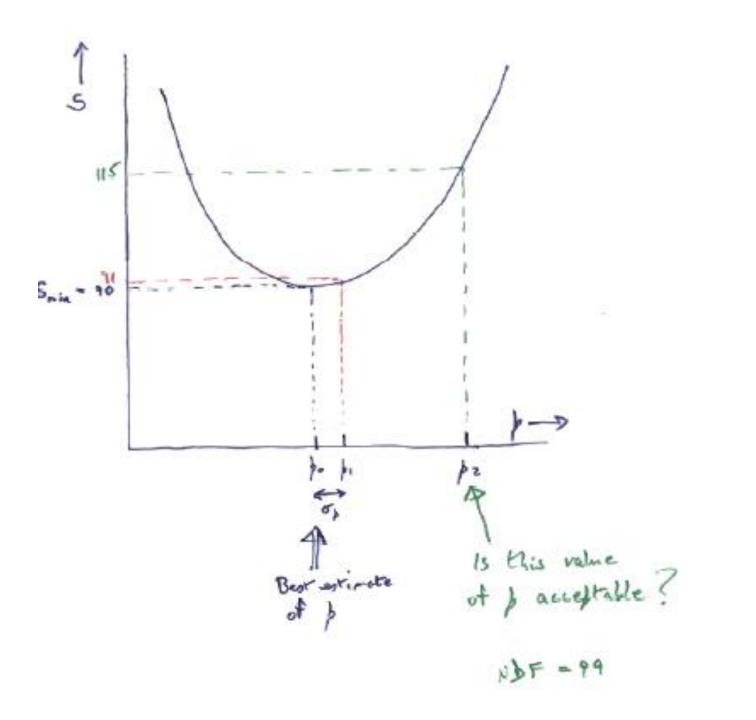
THE PARADOX

Histogram with 100 bins Fit with 1 parameter S_{min} : χ^2 with NDF = 99 (Expected $\chi^2 = 99 \pm 14$)

For our data, $S_{min}(p_0) = 90$ Is p_2 acceptable if $S(p_2) = 115$?

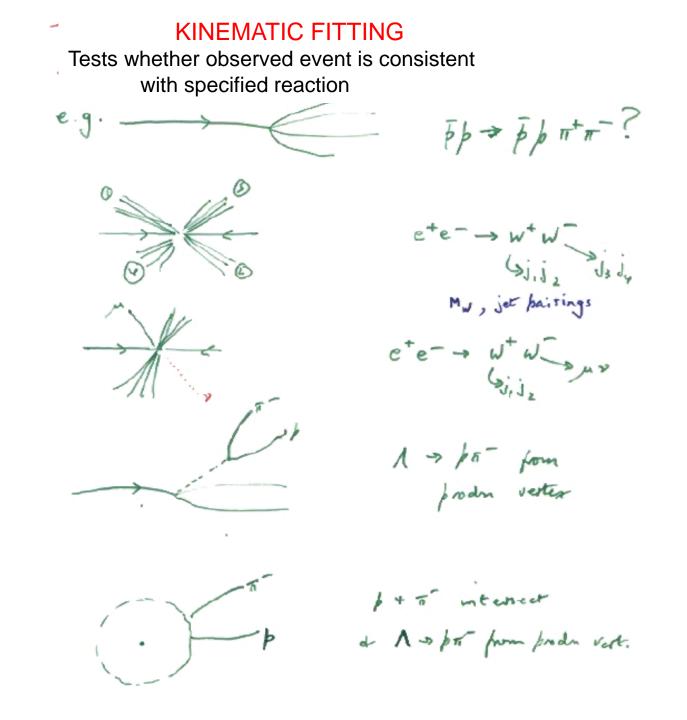
1) YES. Very acceptable χ^2 probability

2) NO. $\sigma_p \text{ from } S(p_0 + \sigma_p) = S_{\min} + 1 = 91$ But $S(p_2) - S(p_0) = 25$ So p_2 is 5 σ away from best value



Next time: Bayes and Frequentism: the return of an old controversy

> The ideologies, with examples Upper limits Feldman and Cousins Summary



Kinematic Fitting: Why do it?

1) Check whether event consistent with hypothesis [Goodness of Fit]

2) Can calculate missing quantities

3) Good to have tracks conserving E-P [Param detn.]

4) Reduces uncertainties

[Param detn.]

[Param detn.]

Kinematic Fitting: Why do it?

1) Check whether event consistent with hypothesis [Goodness of Fit] Use S_{min} and ndf

2) Can calculate missing quantities [Param detn.] e.g. Can obtain |P| for short/straight track, neutral beam; p_x , p_y , p_z of outgoing v, n, K⁰

3) Good to have tracks conserving E-P [Param detn.] e.g. identical values for resonance mass from prodn or decay

4) Reduces uncertainties [Param detn.] Example of "Including theoretical input reduces uncertainties"

How we perform Kinematic Fitting ?

Observed event: 4 outgoing charged tracks Assumed reaction: $pp \rightarrow pp\pi^+\pi^-$

Measured variables: 4-momenta of each track, vimeas

(i.e. 3-momenta & assumed mass)

Then test hypothesis:

Observed event = example of assumed reaction

i.e. Can tracks be wiggled "a bit" to do so? Tested by:

 $S_{min} = \sum (v_i^{fitted} - v_i^{meas})^2 / \sigma^2$

where v_i^{fitted} conserve 4-momenta

(Σ over 4 components of each track)

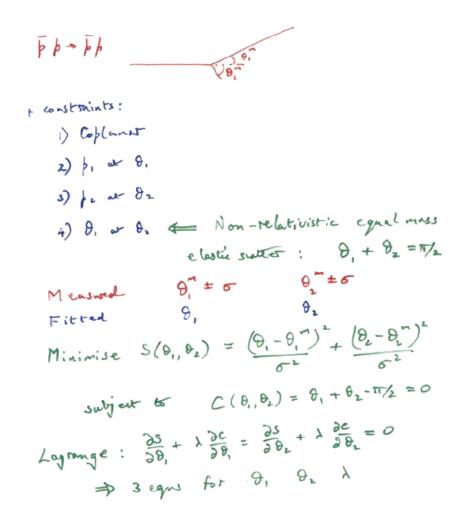
N.B. Really need to take correlations into account

i.e. Minimisation subject to constraints (involves Lagrange Multipliers)

'KINEMATIC' FITTING

Angles of triangle: $\theta_1 + \theta_2 + \theta_3 = 180$ $\theta_1 \quad \theta_2 \quad \theta_3$ Measured 50 60 73 ± 1 Sum = 183 Fitted 49 59 72 180 $\chi^2 = (50-49)^2/1^2 + 1 + 1 = 3$ Prob { $\chi^2_1 > 3$ } = 8.3% **ALTERNATIVELY:** Sum = 183 ± 1.7 , while expect 180 Prob{Gaussian 2-tail area beyond 1.73σ } = 8.3%

Toy example of Kinematic Fit



Eque simple to solve because C(Q, Q2) linear in Q, , D2 $\Rightarrow \theta_{i} = \theta_{i}^{m} + \frac{1}{2}(\pi_{i} - \theta_{i}^{m} - \theta_{i}^{m})$ $\theta_{1} = \theta_{1}^{m} + \frac{1}{2} \left(\frac{\pi}{2} - \theta_{1}^{m} - \theta_{2}^{m} \right)$ $\sigma(\theta_1) = \sigma(\theta_2) = \sigma/\sqrt{2}$

i.e. KINEMATIC FIT → REDUCED UNCERTAINTIES