Dark Matter in Vector Boson Fusion

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Outline

I. General VBF Overview

- II. EFTs and the Higgs Portal
- III. Electroweakinos

I. General VBF Overview



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 VBF kinematics



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- Without Higgs, $V_L V_L \rightarrow V_L V_L$ violates unitarity at ~ TeV.
- *SM-ish* Higgs gives E² behavior in amplitudes.
- 2nd largest production channel for Higgs.

II. EFTs and the Higgs Portal

Effective Field Theories

Effective Field Theories



$$\begin{split} \mathcal{L}_{\mathrm{D5a}} &\supseteq \ \frac{1}{\Lambda} \left[\bar{\chi} \chi \right] \left[\frac{Z_{\mu} Z^{\mu}}{2} + W_{\mu}^{+} W^{-\mu} \right], \\ \mathcal{L}_{\mathrm{D5b}} &\supseteq \ \frac{1}{\Lambda} \left[\bar{\chi} \gamma^{5} \chi \right] \left[\frac{Z_{\mu} Z^{\mu}}{2} + W_{\mu}^{+} W^{-\mu} \right], \\ \mathcal{L}_{\mathrm{D5c}} &\supseteq \ \frac{g}{\Lambda} \left[\bar{\chi} \sigma^{\mu\nu} \chi \right] \left[\frac{\partial_{\mu} Z_{\nu} - \partial_{\nu} Z_{\mu}}{\cos \theta_{W}} - ig \left(W_{\mu}^{+} W_{\nu}^{-} - W_{\nu}^{+} W_{\mu}^{-} \right) \right], \\ \mathcal{L}_{\mathrm{D5d}} &\supseteq \ \frac{g}{\Lambda} \left[\bar{\chi} \sigma_{\mu\nu} \chi \right] \epsilon^{\mu\nu\sigma\rho} \left[\frac{\partial_{\sigma} Z_{\rho} - \partial_{\rho} Z_{\sigma}}{\cos \theta_{W}} - ig \left(W_{\sigma}^{+} W_{\rho}^{-} - W_{\rho}^{+} W_{\sigma}^{-} \right) \right], \\ \mathcal{L}_{\mathrm{D6a}} &\supseteq \ \frac{g}{\Lambda^{2}} \partial^{\nu} \left[\bar{\chi} \gamma^{\mu} \chi \right] \left[\frac{\partial_{\mu} Z_{\nu} - \partial_{\nu} Z_{\mu}}{\cos \theta_{W}} - ig \left(W_{\mu}^{+} W_{\nu}^{-} - W_{\nu}^{+} W_{\mu}^{-} \right) \right], \\ \mathcal{L}_{\mathrm{D6b}} &\supseteq \ \frac{g}{\Lambda^{2}} \partial_{\nu} \left[\bar{\chi} \gamma_{\mu} \chi \right] \epsilon^{\mu\nu\rho\sigma} \left[\frac{\partial_{\sigma} Z_{\rho} - \partial_{\rho} Z_{\sigma}}{\cos \theta_{W}} - ig \left(W_{\sigma}^{+} W_{\rho}^{-} - W_{\rho}^{+} W_{\sigma}^{-} \right) \right], \\ \mathcal{L}_{\mathrm{D7a}} &\supseteq \ \frac{1}{\Lambda^{3}} \left[\bar{\chi} \chi \right] W^{i,\mu\nu} W_{\mu\nu}^{i}, \\ \mathcal{L}_{\mathrm{D7b}} &\supseteq \ \frac{1}{\Lambda^{3}} \left[\bar{\chi} \chi^{5} \chi \right] W^{i,\mu\nu} W_{\mu\nu}^{i}, \\ \mathcal{L}_{\mathrm{D7c}} &\supseteq \ \frac{1}{\Lambda^{3}} \left[\bar{\chi} \chi^{5} \chi \right] \epsilon^{\mu\nu\rho\sigma} W_{\mu\nu}^{i} W_{\rho\sigma}^{i}, \\ \mathcal{L}_{\mathrm{D7d}} &\supseteq \ \frac{1}{\Lambda^{3}} \left[\bar{\chi} \gamma^{5} \chi \right] \epsilon^{\mu\nu\rho\sigma} W_{\mu\nu}^{i} W_{\rho\sigma}^{i}. \end{split}$$

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VBF requires lower MET thresholds compared to Mono-X ⇒ EFT formalism is valid



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"Moore's Law" of Direct Detection



arXiv:1310.8327

Off-Shell Production



• Add a Z₂ symmetric singlet scalar: $\mathcal{L} \supset -c_{\phi} \phi^2 |H|^2$

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- $c_{\phi} \sim 1$ motivated from EW baryogenesis, naturalness.
- $m_{\phi} > m_h / 2$ and no mass-mixing makes this difficult to probe.
- VBF has good balance of signal rate and background separation.

Craig et al., arXiv:1412.0258



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III. Electroweakinos

Neutralinos

$$M_{\tilde{\chi}} = \begin{pmatrix} \tilde{H}_u & , & \tilde{H}_d & , & \tilde{W} & , & \tilde{B} \\ 0 & \mu & -g_2 v_u / \sqrt{2} & g_1 v_u / \sqrt{2} \\ \mu & 0 & g_2 v_d / \sqrt{2} & -g_1 v_d / \sqrt{2} \\ -g_2 v_u / \sqrt{2} & g_2 v_d / \sqrt{2} & M_2 & 0 \\ g_1 v_u / \sqrt{2} & -g_1 v_d / \sqrt{2} & 0 & M_1 \end{pmatrix}$$

1. Winos and Higgsinos

• Thermal production favores 1 TeV Higgsinos and 3 TeV Winos.

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- Prospects for direct detection solely dependent on known parameters.



Solon, Hill, arXiv:1309.4092



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VBF Signal



VBF Signal



VBF Cuts

• Backgrounds: Z+jets, W+jets, tt, QCD multijet

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Cut	14 TeV		$100 { m TeV}$	
	Wino	Higgsino	Wino	Higgsino
$n_{ m jet}$	2	2	2	2
$ \eta(j) $	5	5	7	7
$p_T(j_{\text{tag}}) \text{ (GeV)}$	45	45	75	50
$\Delta\eta(j_1,j_2)$	3.75	3.75	4.25	4.25
$\Delta \phi(j_1,j_2)$	2	2	2	3
$M(j_1, j_2)$ (TeV)	2	1	10	5
	400-700		1100 - 2500	
$p_T(j_{\text{veto}}) \text{ (GeV)}$	45	45	50	50
$p_T(e,\mu)$ (GeV)	20	20	20	20
$p_T(\tau) \; (\text{GeV})$	30	30	40	40
$\eta(e)$	2.5	2.5	2.5	2.5
$\eta(\mu)$	2.1	2.1	2.1	2.1
$\eta(au)$	2.3	2.3	2.3	2.3

Results



High Luminosity



2. Binos

Cosmology

Motivates small mass splittings.





Ibarra, Pierce, Shah, Vogl, arXiv:1501.03164



AB, D. Robertson, M. Solon, K. Zurek, arXiv:1511.05964

VBF Signal

- Compressed bino-slepton spectra difficult to probe with direct searches.
- VBF jets + MET + central leptons
- Mass reach of O(100) GeV binos



Dutta et al. arXiv:1411.6043

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- Most powerful direct probe for off-shell production.
- Consistent EFT models of dark matter at Run 2.
- Complementary (but not most sensitive) channel for electroweakinos.
- Tagging additional central activity is an interesting alternative.

Backup Slides

On-Shell Production



• Simplified Model: $\mathcal{L} \supset g_{\chi} \ h \ \overline{\chi} \chi$

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Brooke et al., arXiv:1603.07739

VBF Production On-Shell

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