The SeaQuest experiment & displaced dark sectors

Stefania Gori University of Cincinnati

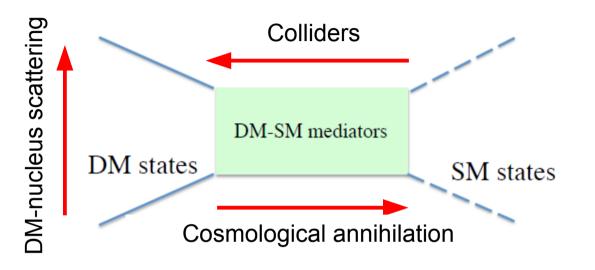
Aspen 2017 winter conference "From the LHC to dark matter and beyond"

> Aspen, March 21nd 2017

Light dark sectors

What is the Dark Matter energy scale?

 $\mathsf{D}\mathsf{M}$



EW mediators lead to the so called Lee-Weinberg window

 ${
m few}~{
m GeV} \lesssim m_{
m DM} \lesssim {
m few}~{
m TeV}$

Light WIMPs are facilitated by light mediators Dark sectors!

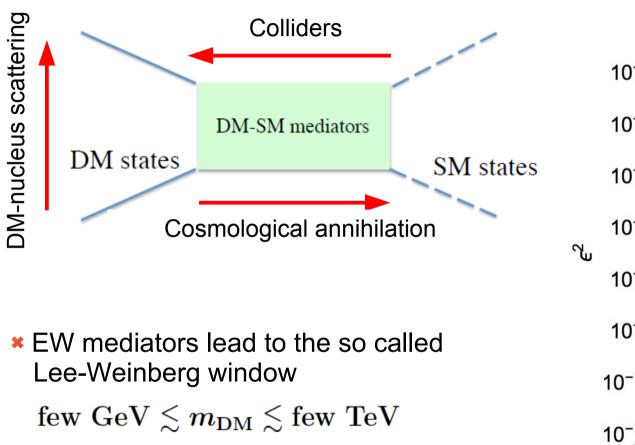
SM

mediator



Light dark sectors

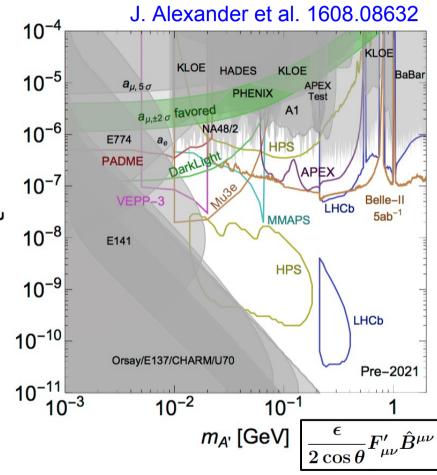
What is the Dark Matter energy scale?





Dark sectors!



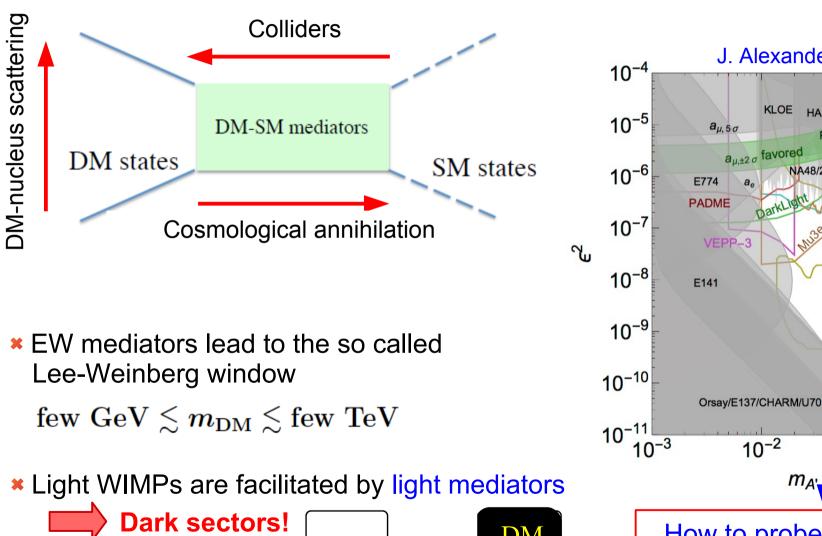


Light dark sectors

What is the Dark Matter energy scale?

DM

_р, S,



SM

mediator

How to probe the (100 MeV - few GeV) range?

m_{A'} [GeV]

J. Alexander et al. 1608.08632

KLOE

A1

HPS

MMAPS

HPS

 10^{-1}

APEX

Test

APEX

KLOE

DarkLight

 $a_{\mu,\pm 2\sigma}$ favored

 10^{-2}

HADES

NA48/2

PHENIX

KLOE

Belle-II

LHCb

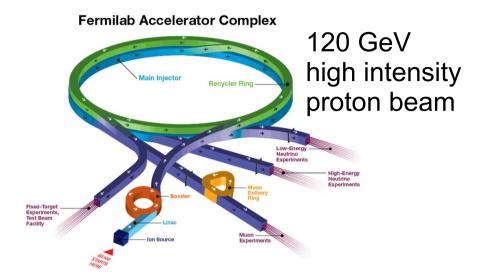
Pre-2021

BaBar





Fermilab intensity frontier



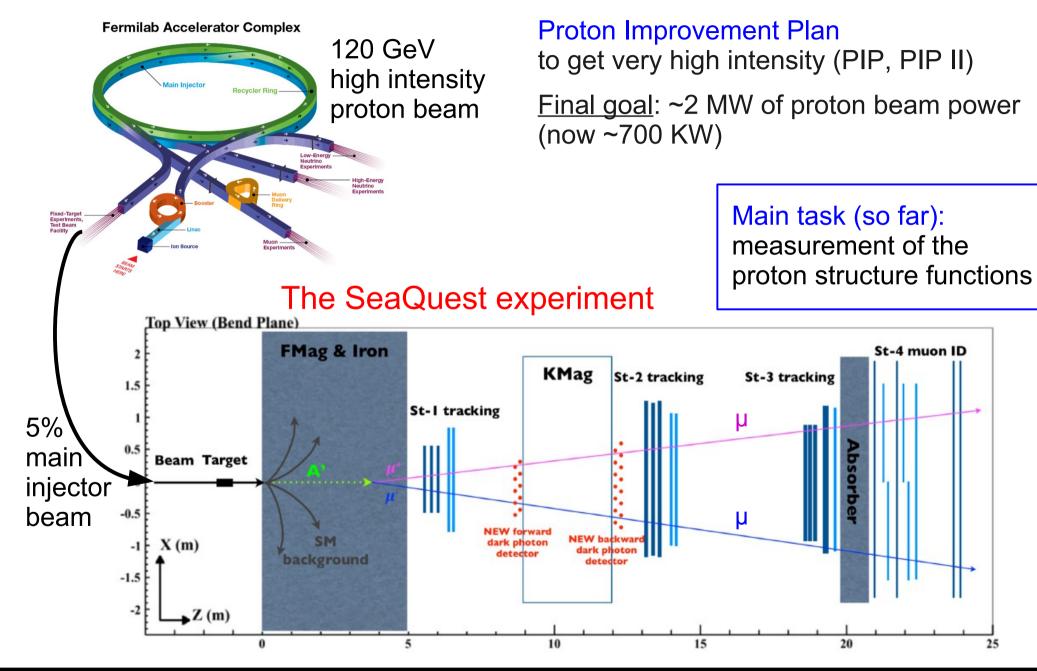
Proton Improvement Plan

to get very high intensity (PIP, PIP II)

<u>Final goal</u>: ~2 MW of proton beam power (now ~700 KW)

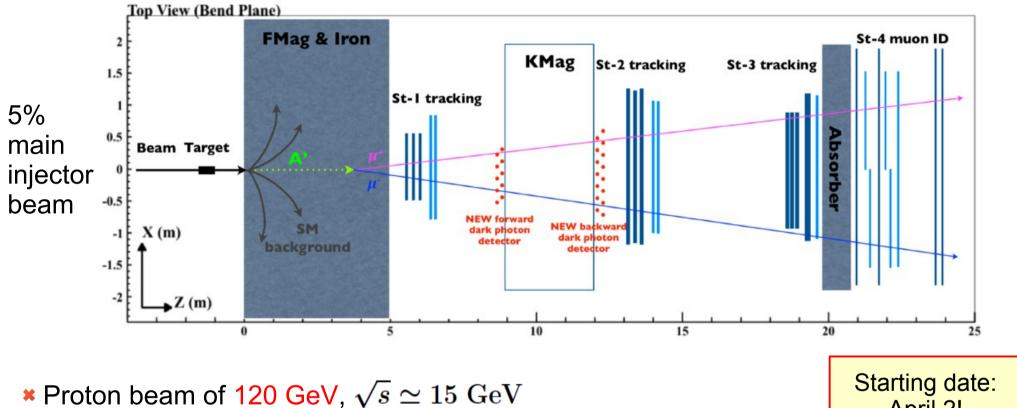


Fermilab intensity frontier





The SeaQuest fact sheet

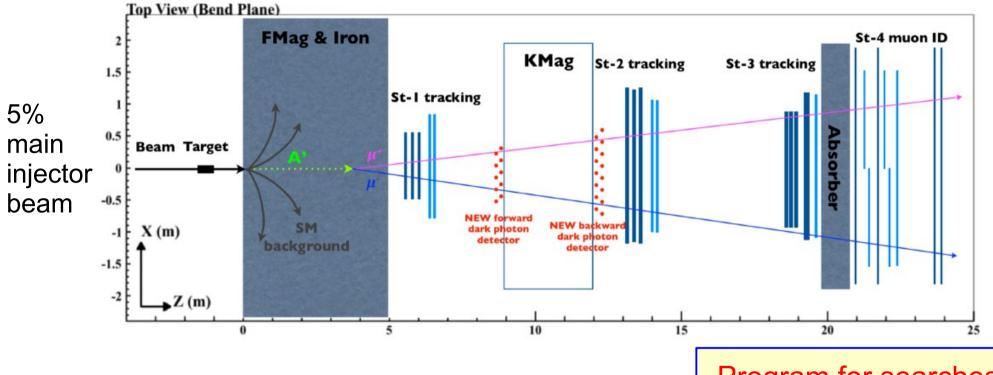


Recent installation of a displaced trigger

April 2!

- Phase I parassitic run: 1.44 × 10¹⁸ POT in 200 days
- * <u>Possible Phase II</u>: upgrade of the detector + larger luminosity (~ 10^{20} POT?)

The SeaQuest fact sheet



- ***** Proton beam of 120 GeV, $\sqrt{s} \simeq 15~{
 m GeV}$
- Recent installation of a displaced trigger

Program for searches for New Physics?

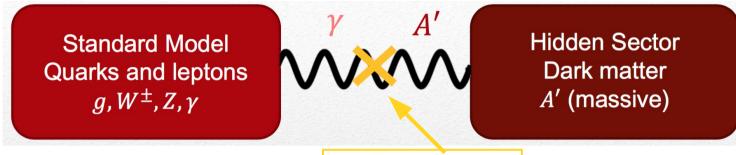
- Phase I parassitic run: 1.44 × 10¹⁸ POT in 200 days
- * Possible Phase II: upgrade of the detector + larger luminosity (~ 10²⁰ POT?)

Main features:

- Higher beam energy (if compared to LSND, MiniBooNE, ...)
- Smaller detector-target distance (if compared to CHARM, NOMAD, ...)

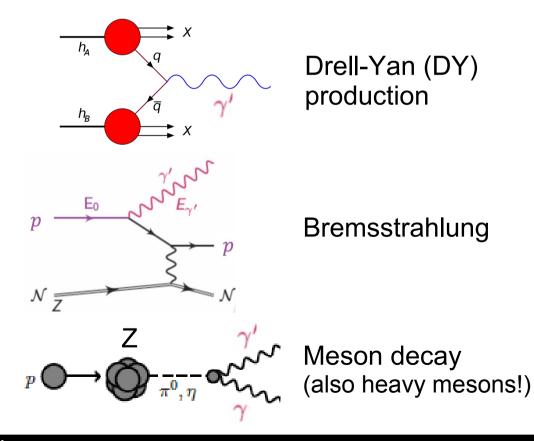


A minimal dark photon model



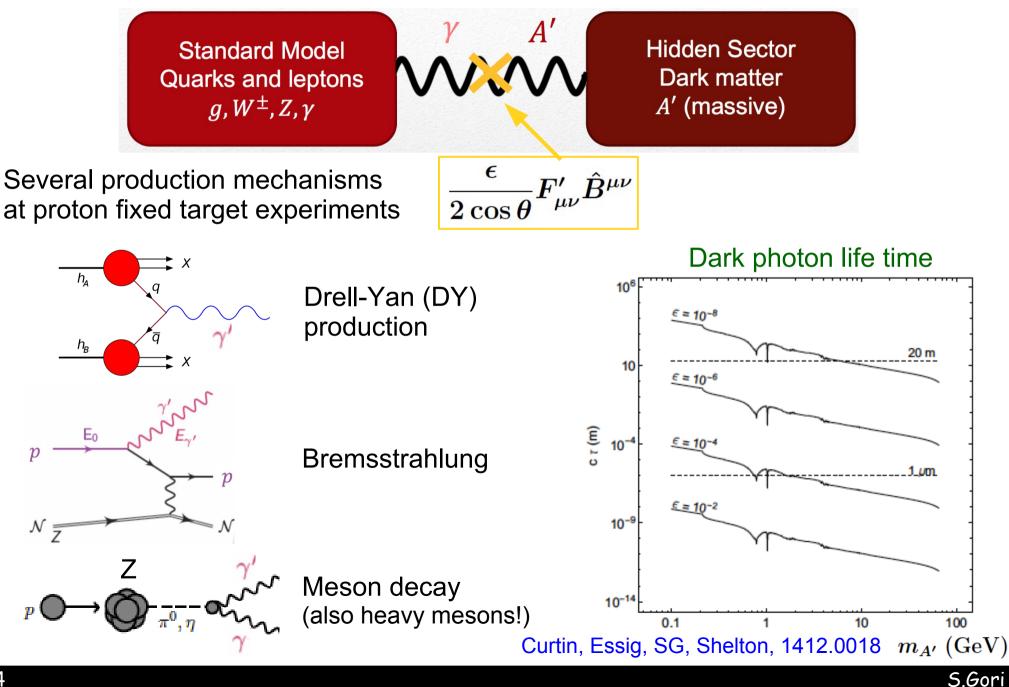
Several production mechanisms at proton fixed target experiments

$$rac{\epsilon}{2\cos heta}F'_{\mu
u}\hat{B}^{\mu
u}$$

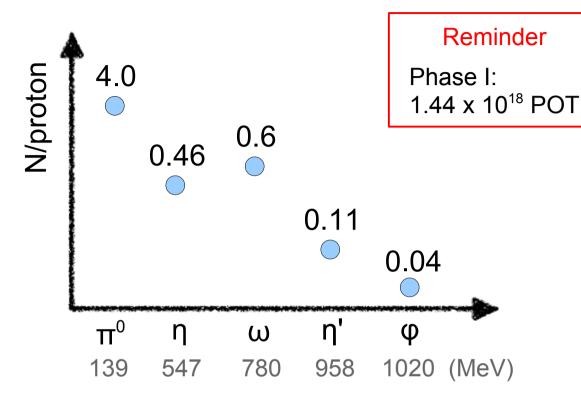




A minimal dark photon model



Huge meson production at SeaQuest

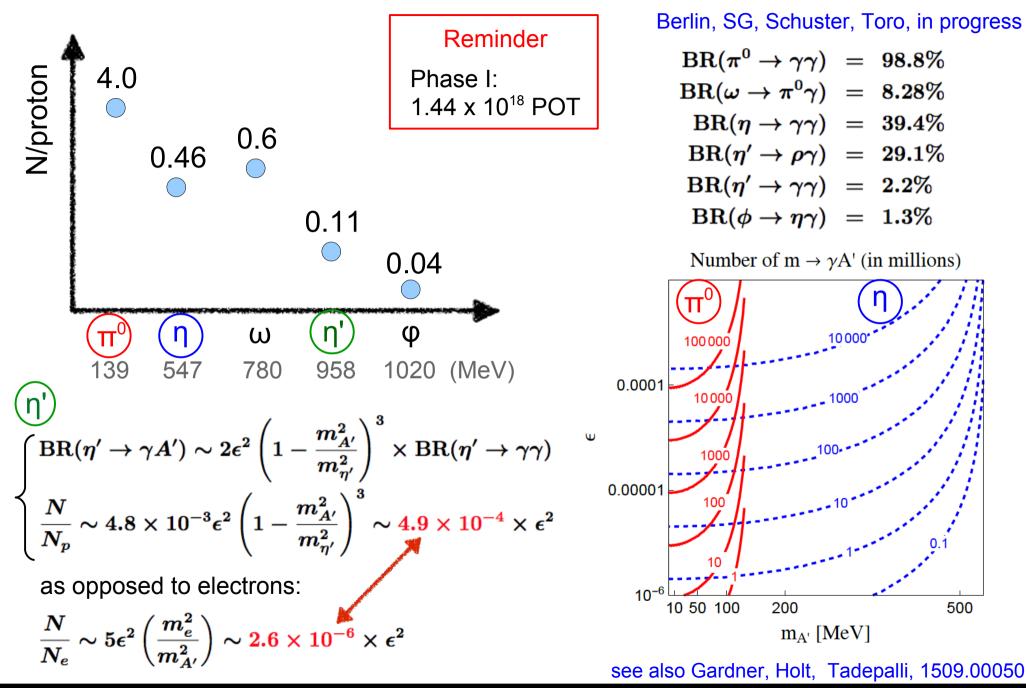


Berlin, SG, Schuster, Toro, in progress

 $\begin{array}{rcl} \mathrm{BR}(\pi^0 \rightarrow \gamma \gamma) &=& 98.8\% \\ \mathrm{BR}(\omega \rightarrow \pi^0 \gamma) &=& 8.28\% \\ \mathrm{BR}(\eta \rightarrow \gamma \gamma) &=& 39.4\% \\ \mathrm{BR}(\eta' \rightarrow \rho \gamma) &=& 29.1\% \\ \mathrm{BR}(\eta' \rightarrow \gamma \gamma) &=& 2.2\% \\ \mathrm{BR}(\phi \rightarrow \eta \gamma) &=& 1.3\% \end{array}$



Huge meson production at SeaQuest



Prompt vs. displaced dark photons

Before 2017, SeaQuest could only see prompt di-muons

In collaboration with the SeaQuest collaboration:

Expected sensitivities for Phase I (solid) & Phase II (dashed)

 10^{-2} onstraint SeaQues Prompt Muon a-2 Constraint DY-like A 10⁻³ Ψ HPS DarkLight 10^{-4} APEX Full Belle II BaBar KLOE 0.3 0.1 10 1 3 DY-dark $m_{A'}$ [GeV] photon

Systematics limited (S/B ~ few 10^{-5})

Improvement can be achieved with a better mass resolution (6% assumed)

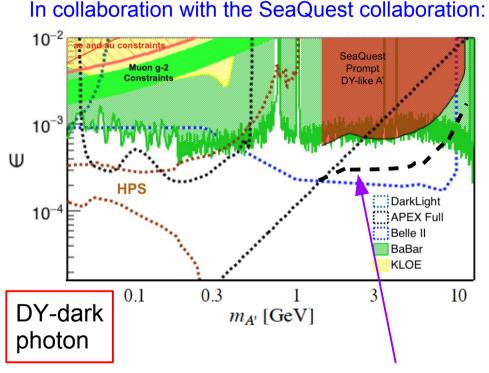


Prompt vs. displaced dark photons

Before 2017, SeaQuest could only see prompt di-muons

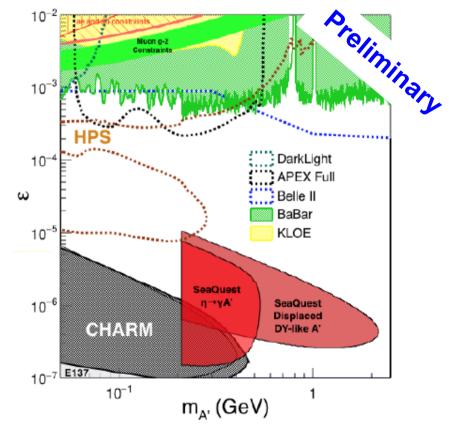
Very recent installation of a <u>displaced trigger</u>

Expected sensitivities for Phase I (solid) & Phase II (dashed)



Systematics limited (S/B ~ few 10⁻⁵)

Improvement can be achieved with a better mass resolution (6% assumed)



We ask for displaced muons originating after the bump almost background 0 experiment

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Beyond resonance searches: the IDM case

With the installation of the <u>displaced trigger</u>, the SeaQuest program broadens considerably!

Any dark sector coupled to protons and muons (even very weakly)

What is the physics case we can establish?

Work in progress on SIMP models, inelastic DM models, and right handed neutrinos

Berlin, SG, Schuster, Toro, in preparation

Beyond resonance searches: the IDM case

With the installation of the <u>displaced trigger</u>, the SeaQuest program broadens considerably!

Any dark sector coupled to protons and muons (even very weakly)

Inelastic Dark Matter

In a nutshell:

Tucker-Smith, Weiner 0101138, 0402065

***** Dirac spinor charged under a U(1)_D symmetry $\psi = (\eta, \xi^{\dagger})$

$$-\mathcal{L} \supset m_D\eta \xi + rac{m_\eta}{2}\eta\eta + rac{m_\xi}{2}\xi \xi + h.c.$$

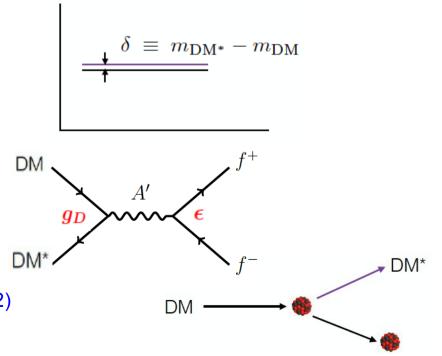
It is natural that m_n , $m_{\epsilon} << m_{D}$

Main interactions are inelastic

- leading annihilation process DM $\text{DM}^* \to \text{SM}$

- very suppressed direct detection signals

(see however Bramante et al., 1608.02662)

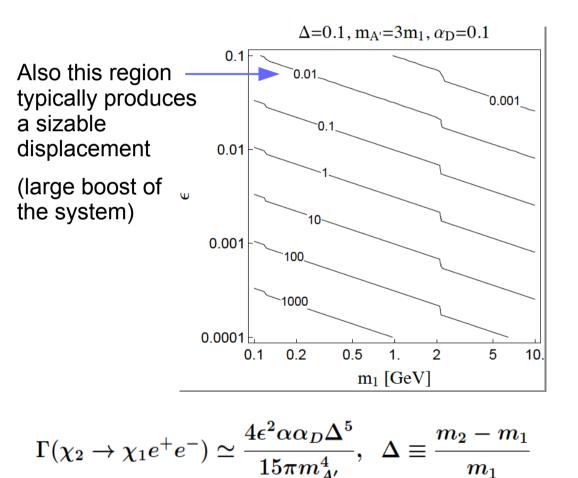


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A long lived DM excited state

Berlin, SG, Schuster, Toro, in preparation

Life time of the excited state (in meters)

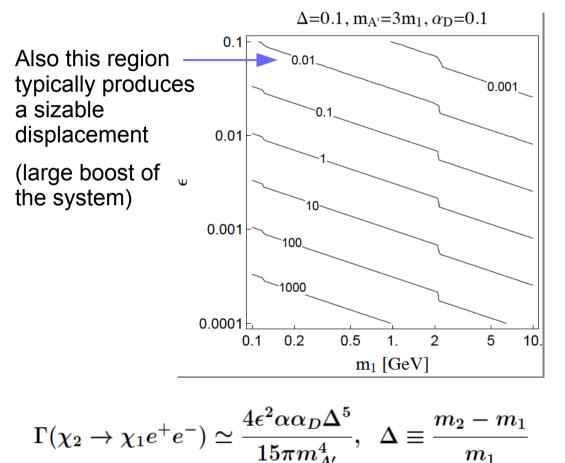




A long lived DM excited state

Berlin, SG, Schuster, Toro, in preparation

Life time of the excited state (in meters)



A variety of search strategies:

Production topologies:

 $(\gamma, j)A' \rightarrow \chi_2 \chi_1 \rightarrow (f \bar{f} \chi_1) \chi_1$ $(\gamma, j)A' \rightarrow f \bar{f}$ (suppressed by ϵ^2)

Babar (and future B-factories)

- visible (monophoton + 2 electrons/muons)
- invisible (monophoton + missing)

× LHC

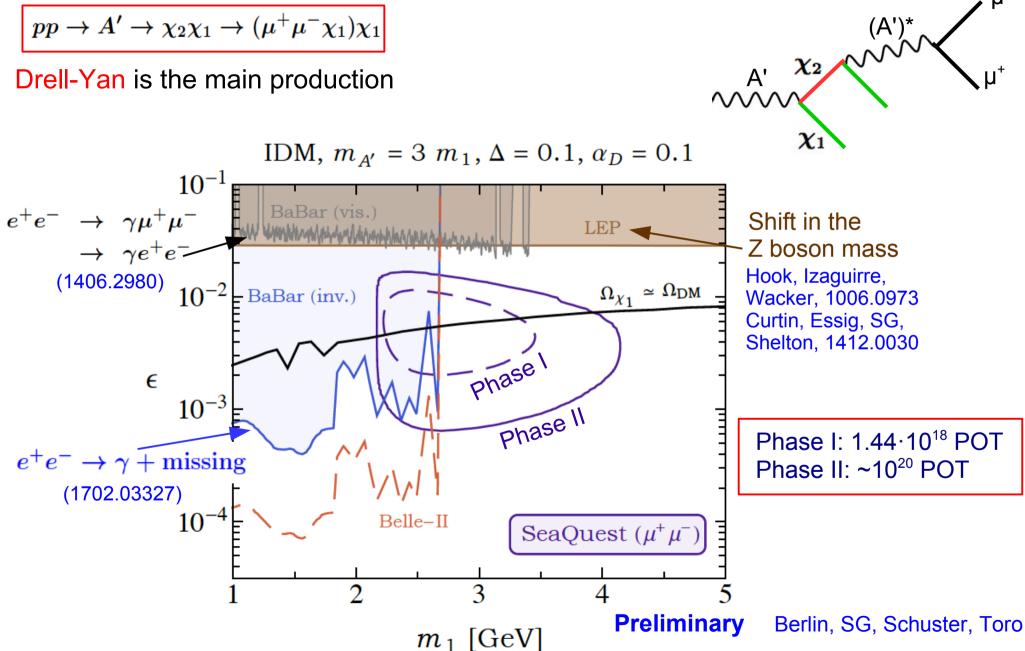
- monojet
- proposed search for displaced (soft) muons plus a jet (Izaguirre, Krnjaic, Shuve, 1508.03050)
- (low energy) electron and proton fixed target experiments (BDX, LSND, E137...)
 - electrons
 - missing mass

What is the role of SeaQuest?



The SeaQuest reach

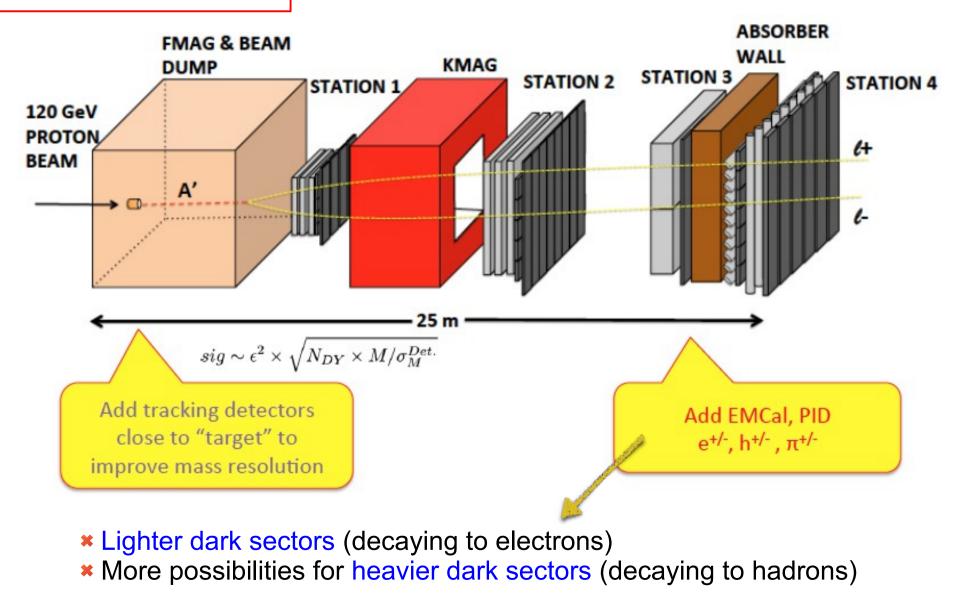
$$pp
ightarrow A'
ightarrow \chi_2 \chi_1
ightarrow (\mu^+ \mu^- \chi_1) \chi_1$$



, μ⁻

SeaQuest phase II: more lumi & upgrade

Plan for 2020 - 2025

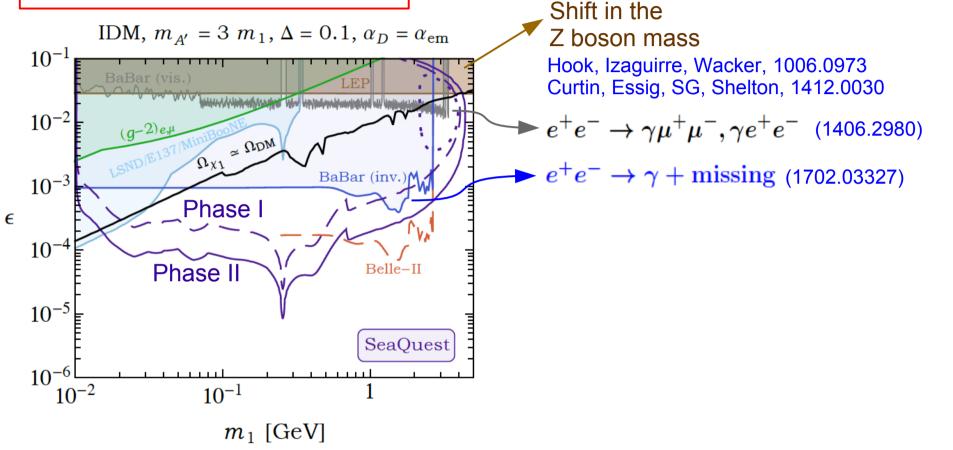






IDM SeaQuest reach, electron signature

$pp ightarrow A' ightarrow \chi_2 \chi_1 ightarrow (e^+e^-\chi_1) \chi_1$



Preliminary Berlin, SG, Schuster, Toro



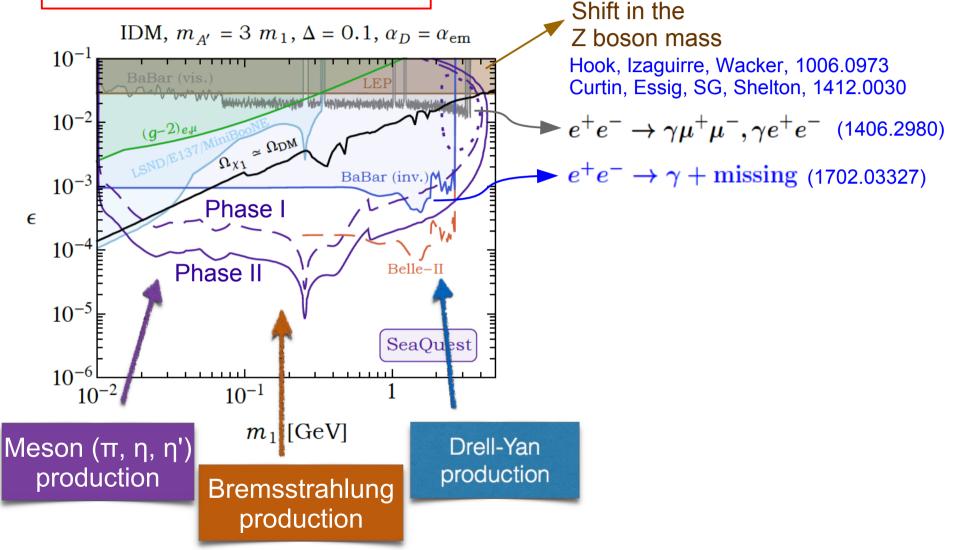
Berlin, SG, Schuster, Toro, in preparation

IDM SeaQuest reach, electron signature

Berlin, SG, Schuster, Toro, in preparation

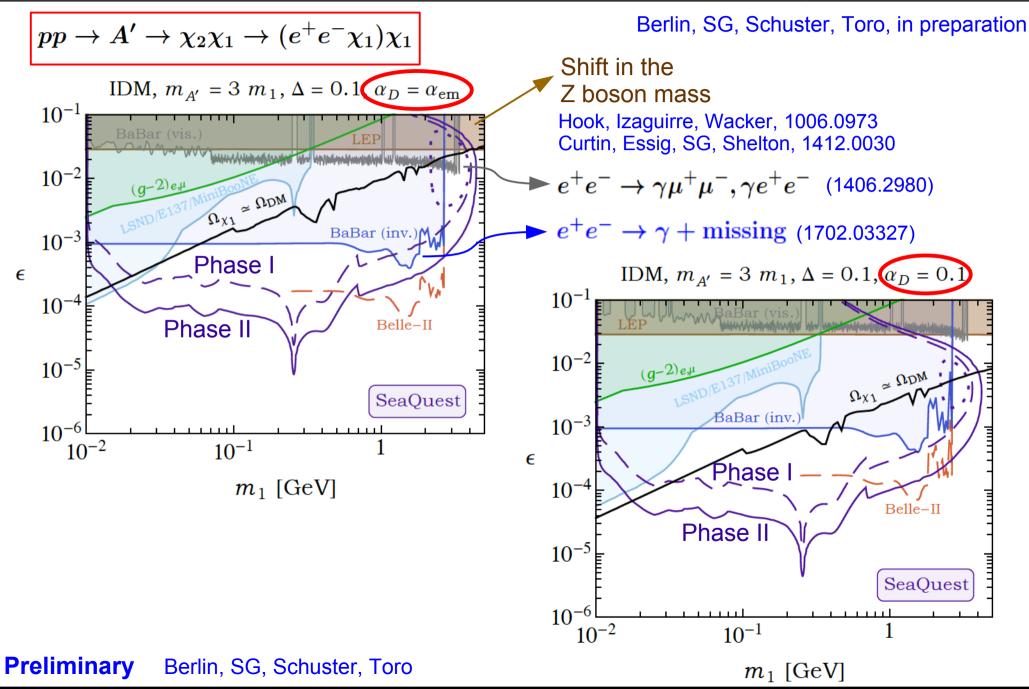
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$pp ightarrow A' ightarrow \chi_2 \chi_1 ightarrow (e^+e^-\chi_1) \chi_1$



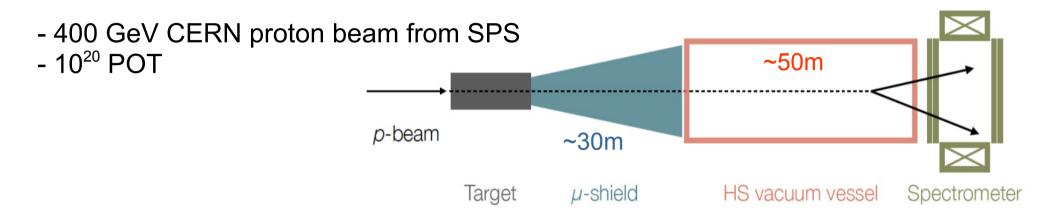
Preliminary Berlin, SG, Schuster, Toro

IDM SeaQuest reach, electron signature



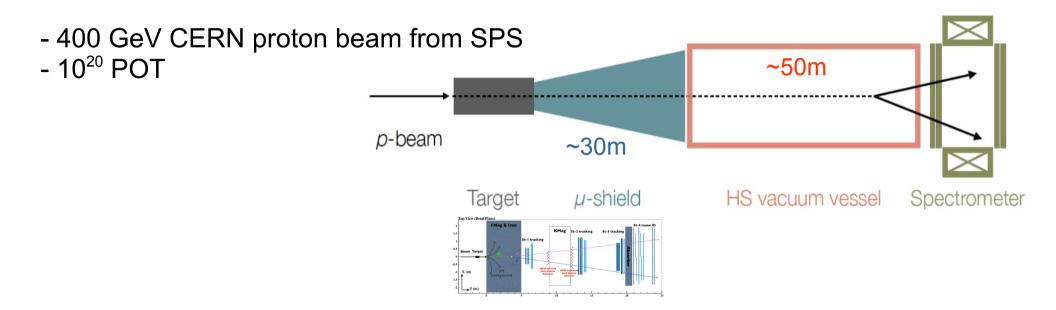
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Ongoing proposal for the ShiP experiment



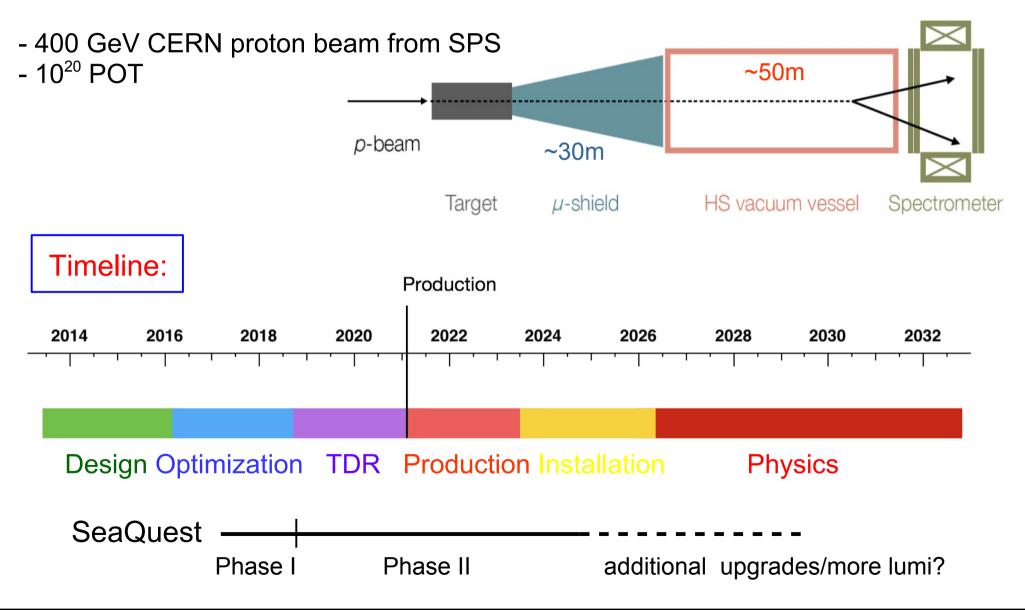


Ongoing proposal for the ShiP experiment





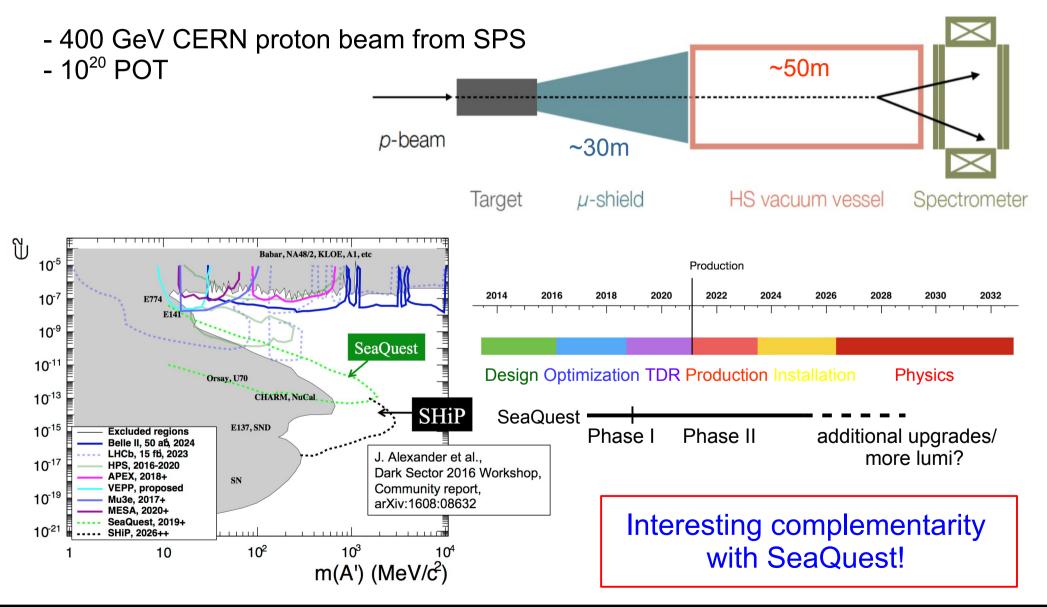
Ongoing proposal for the ShiP experiment



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Ongoing proposal for the ShiP experiment



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Conclusions

High energy proton beam dump experiments

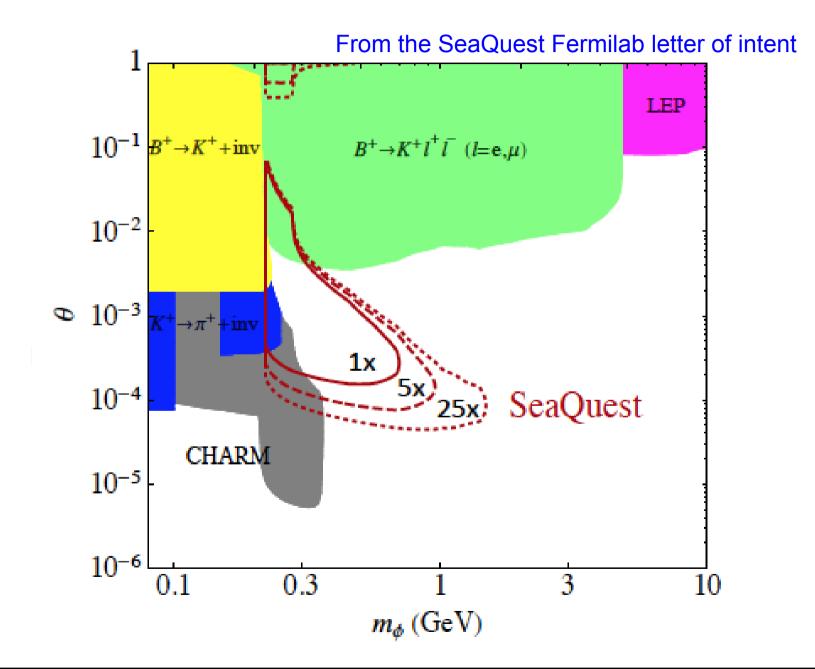
Interesting setup for looking for Dark particles. Huge rates! Feasible high intensity program for Fermilab

- * Beyond di-muon bump-hunt, Seaquest
- can efficiently look for displaced signatures
- has a unique opportunity to test large part of the remaining parameter space of Inelastic Dark Matter (IDM) models
- can be upgraded with a EMCal to see electron signatures
- can test additional models: SIMPs, right handed neutrinos, ... (in progress)
- * Additional "dark signatures" to look for?
- Ultimate program at the ShiP experiment at CERN (400 GeV proton beam, ~10²⁰ POT, starting in ~2026?)





SeaQuest reach for dark scalars

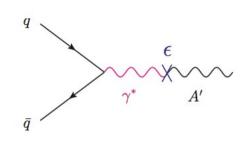




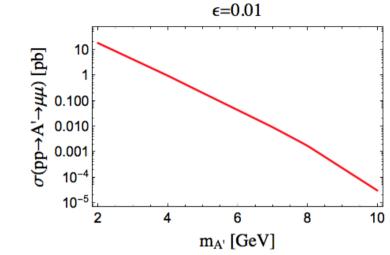


Drell-Yan, prompt

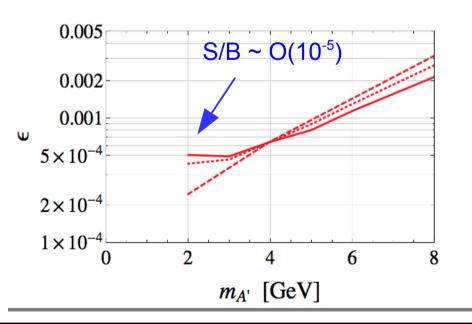
The cross sections for dark photon production are rather sizable



Expected lumi: 35 ab⁻¹



Our preliminary reach:



The exact reach depends sizably on the mass resolution and on the momentum of the two muons

 $= E_{\mu} > 10 \text{ GeV}, p_{x}/p_{z} > 0.05 \text{ and } dM = 225 \text{ MeV}$ $= E_{\mu} > 1.5 \text{ GeV} \text{ and } dM = 225 \text{ MeV}$ $= E_{\mu} > 10 \text{ GeV}, p_{x}/p_{z} > 0.05 \text{ and } dM = 6\% \text{ m}_{A'}$ $= p_{v}/p_{z} < 0.1$