

2016 progress in two-loop electroweak pseudoobservables and further prospects [S-matrix approach]

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Based on collaboration with:
levgen Dubovskyk, Ayres Freitas, Johann Usovitsch

- levgen Dubovskyk, Ayres Freitas, JG, Tord Riemann, Johann Usovitsch
"The two-loop electroweak bosonic corrections to $\sin^2 \theta_{\text{eff}}^b$ "
Phys.Lett. B762 (2016) 184
- TR LL16 talk, PoS LL2016 (2016) 075:
"30 years, some 700 integrals, and 1 dessert, or:
Electroweak two-loop corrections to the $Z\bar{b}b$ vertex",
arXiv:1610.07059;
- JG LL16 talk, PoS LL2016 (2016) 034:
"Numerical integration of massive two-loop Mellin-Barnes integrals in
Minkowskian regions",
arXiv:1607.07538

Outline

1 Introduction

- Electroweak Pseudo-observables (EWPOs)
- The effective weak mixing angle $\sin^2 \theta_{\text{eff}}^{\text{b}}$

2 Fresh rolls: 2-loop EW bosonic corrections to $\sin^2 \theta_{\text{eff}}^{\text{b}}$

3 Numerical 2-loop calculations

- Mellin-Barnes versus and sector decomposition methods

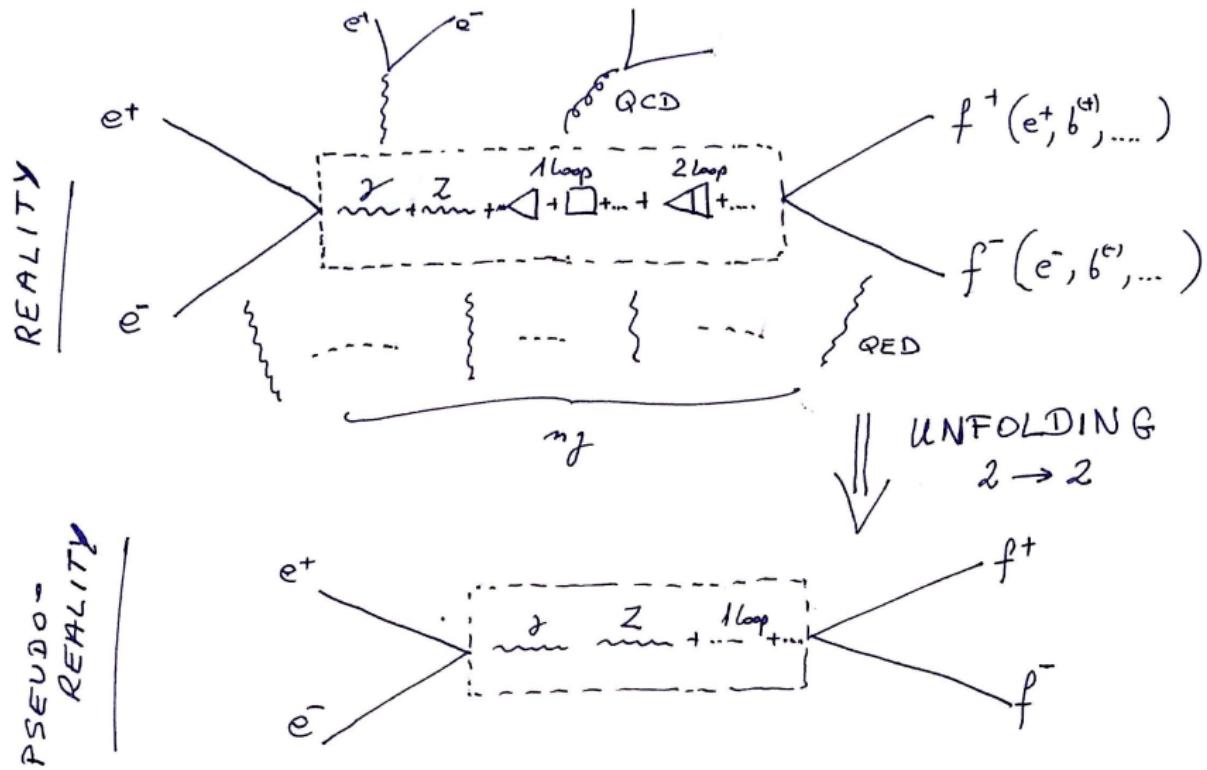
4 Pseudo-observables, S-matrix and $\gamma - Z$ interferences

5 Summary and Outlook: To be or not to be (optimistic)

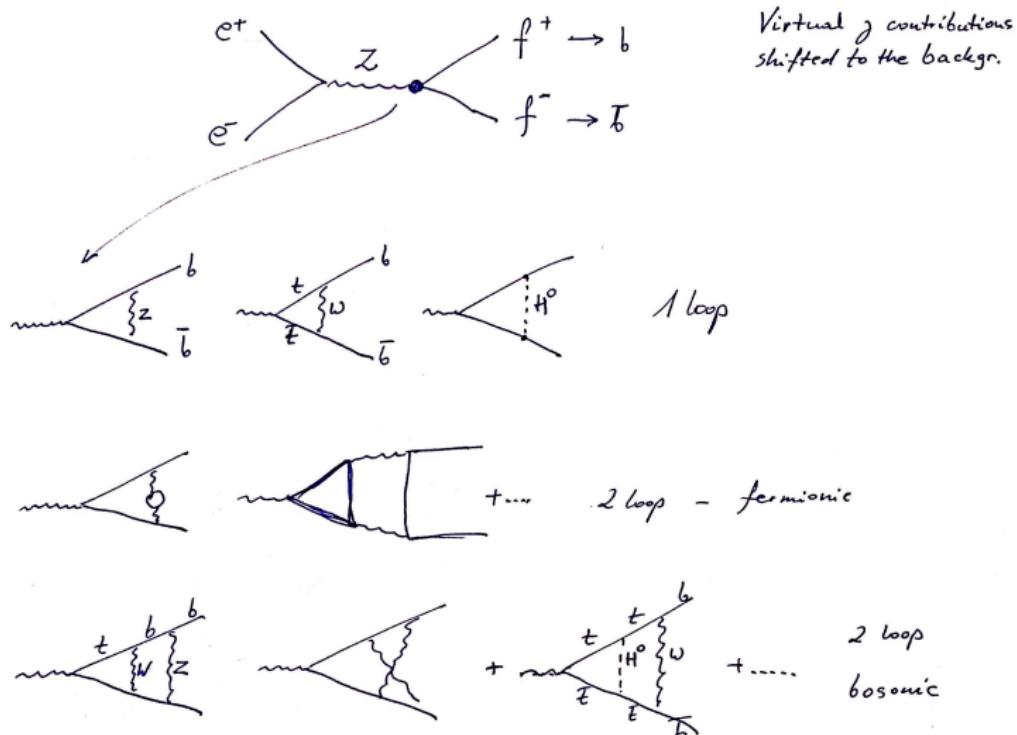
6 Backup slides: details of numerical methods

- Construction of MB integrals
- Avoiding numerical instabilities

Road to the $Z\bar{b}b$ vertex and 2-loop EW corrections (1)



Road to the $Z\bar{b}b$ vertex and 2-loop EW corrections (2)



Pseudo-observables, an example: $d\sigma/d\cos\theta$ ($e^+e^- \rightarrow \bar{b}b$)

Close to the Z -boson peak and assuming Born-like v, a couplings:

$$\begin{aligned}\frac{d\sigma}{d\cos\theta} &\sim G_F^2 \left| \frac{s}{s - M_Z^2 + iM_Z\Gamma_Z} \right|^2 \\ &\times [(a_e^2 + v_e^2)(a_b^2 + v_b^2)(1 + \cos^2\theta) (2a_e v_e)(2a_b v_b)(2\cos\theta)]\end{aligned}$$

Factorizations:

Symmetric integration over $\cos\theta$

$$\sigma_T = \int_{-1}^1 d\cos\theta \frac{d\sigma}{d\cos\theta} \sim \left| \frac{s}{s - M_Z^2 + iM_Z\Gamma_Z} \right|^2 G_F(a_e^2 + v_e^2) \quad \mathbf{G_F}(\mathbf{a_b^2 + v_b^2})$$

Anti-symmetric integration over $\cos\theta$

$$A_{F-B} = \frac{\left[\int_0^1 d\cos\theta - \int_{-1}^0 d\cos\theta \right] \frac{d\sigma}{d\cos\theta}}{\sigma_T} \sim \overbrace{\frac{A_e}{2a_e v_e}}^{\sigma_T} \quad \overbrace{\frac{A_b}{2\mathbf{a_b v_b}}}_{\mathbf{a_b^2 + v_b^2}}$$

Pseudo-observable A_b

$$A_b = \frac{2\Re e \frac{v_b}{a_b}}{1 + \left(\Re e \frac{v_b}{a_b}\right)^2} = \frac{1 - 4|Q_b|\sin^2 \theta_{\text{eff}}^{\text{b}}}{1 - 4|Q_b|\sin^2 \theta_{\text{eff}}^{\text{b}} + 8Q_b^2(\sin^2 \theta_{\text{eff}}^{\text{b}})^2}$$

Definition of the effective weak mixing angle

$$\sin^2 \theta_{\text{eff}}^{\text{b}} = \frac{1}{4|Q_b|} \left(1 - \Re e \frac{v_b}{a_b} \right)$$

- Vertex form factor

$$V_\mu^{Zb\bar{b}} = \gamma_\mu [v_b(s) - a_b(s)\gamma_5] = \dots +$$

$$+$$

$$+$$

$$+\dots$$

$e^+e^- \rightarrow Z \rightarrow l\bar{l}, b\bar{b}$, status

- 1985 - 1-loop leptonic ($l\bar{l}$) EW and $b\bar{b}$ corrections (Akhundov, Bardin, Riemann)
- 2006 - 2-loop leptonic EW corrections (Awramik, Czakon, Freitas)
- 2008 - 2-loop $b\bar{b}$ EW corrections with fermionic sub-loops (Awramik, Czakon, Freitas, Kniehl)
- 2016 - Completion: 2-loop $b\bar{b}$ bosonic EW corrections - DFGRU

Our project started in 2012 (TR - AF meeting). Basis for success:

- ① Ayres Freitas: knowledge of the 2-loop renormalization scheme + experience in previous studies
- ② TR, JG, JU, ID: new numerical evaluations based on Mellin-Barnes (**MB**) approach to Feynman integrals — In 2012 we hoped to use known by that time versions of AMBRE/MB tools - completely naive assumption (!)

Our results: Effective weak mixing angle $\sin^2 \theta_{\text{eff}}^{\text{b}}$

- The standard model prediction for the effective weak mixing angle can be written as

$$\sin^2 \theta_{\text{eff}}^{\text{b}} = \left(1 - \frac{M_W^2}{M_Z^2}\right) (1 + \Delta\kappa_b)$$

- The bosonic electroweak two-loop corrections amount to

$$\Delta\kappa_b^{(\alpha^2, \text{bos})} = -0.9855 \times 10^{-4}$$

DFGRU, Phys.Lett. B762 (2016) 184

Collection of radiative corrections: full stabilization at $10^{-4}!$

Order	Value [10^{-4}]	Order	Value [10^{-4}]
α	468.945	$\alpha_t^2 \alpha_s$	1.362
$\alpha \alpha_s$	-42.655	α_t^3	0.123
$\alpha_t \alpha_s^2$	-7.074	α_{ferm}^2	3.866
$\alpha_t \alpha_s^3$	-1.196	α_{bos}^2	-0.986

Table: Comparison of different orders of radiative corrections to $\Delta \kappa_b$.

Input Parameters: M_Z , Γ_Z , M_W , Γ_W , M_H , m_t , α_s and $\Delta \alpha$

1-loop contributions	Akhundov:1985
fermionic EW 2-loop corrections	Awramik:2008
$\mathcal{O}(\alpha \alpha_s)$ QCD corrections	Djouadi:1987,Djouadi:1987,Kniehl:1989,Kniehl: 1991, Fleischer:1992,Buchalla:1992,Czarnecki:1996
partial higher-order corrections of orders $\mathcal{O}(\alpha_t \alpha_s^2)$	Avdeev:1994,Chetyrkin:1995
$\mathcal{O}(\alpha_t \alpha_s^3)$	Schroder:2005,Chetyrkin:2006,Boughezal:2006
$\mathcal{O}(\alpha^2 \alpha_t)$ and $\mathcal{O}(\alpha_t^3)$	vanderBij:2000,Faisst:2003

Simple fitting formula

$$\Delta\kappa_b^{(\alpha^2, \text{bos})} = k_0 + k_1 c_H + k_2 c_t + k_3 c_t^2 + k_4 c_H c_t + k_5 c_W \quad (1)$$

$$c_H = \log \left(\frac{M_H}{M_Z} \times \frac{91.1876 \text{GeV}}{125.1 \text{GeV}} \right)$$

$$c_t = \left(\frac{m_t}{M_Z} \times \frac{91.1876 \text{GeV}}{173.2 \text{GeV}} \right)^2 - 1 \quad (2)$$

$$c_W = \left(\frac{M_W}{M_Z} \times \frac{91.1876 \text{GeV}}{80.385 \text{GeV}} \right)^2 - 1$$

$$\begin{aligned} k_0 &= -0.98605 \times 10^{-4}, & k_1 &= 0.3342 \times 10^{-4}, & k_2 &= 1.3882 \times 10^{-4}, \\ k_3 &= -1.7497 \times 10^{-4}, & k_4 &= -0.4934 \times 10^{-4}, & k_5 &= -9.930 \times 10^{-4} \end{aligned} \quad (3)$$

The deviations to the full calculation amount to average (maximal) 5×10^{-8} (1.2×10^{-7}), in the input parameter ranges.

Currently most precise prediction for $\sin^2 \theta_{\text{eff}}^{\text{b}}$

$$\begin{aligned} \sin^2 \theta_{\text{eff}}^{\text{b}} = & s_0 + d_1 L_H + d_2 L_H^2 + d_3 \Delta_\alpha + d_4 \Delta_t + d_5 \Delta_t^2 \\ & + d_6 \Delta_t L_H + d_7 \Delta_{\alpha_s} + d_8 \Delta_t \Delta_{\alpha_s} + d_9 \Delta_Z \end{aligned} \quad (4)$$

$$\begin{aligned} L_H &= \log \left(\frac{M_H}{125.7 \text{GeV}} \right), \quad \Delta_t = \left(\frac{m_t}{173.2 \text{GeV}} \right)^2 - 1, \quad \Delta_Z = \frac{M_Z}{91.1876 \text{GeV}} - 1, \\ \Delta_\alpha &= \frac{\Delta \alpha}{0.0059} - 1, \quad \Delta_{\alpha_s} = \frac{\alpha_s}{0.1184} - 1. \end{aligned} \quad (5)$$

$$\begin{aligned} s_0 &= 0.232704, \quad d_1 = 4.723 \times 10^{-4}, \quad d_2 = 1.97 \times 10^{-4}, \quad d_3 = 2.07 \times 10^{-2}, \\ d_4 &= -9.733 \times 10^{-4}, \quad d_5 = 3.93 \times 10^{-4}, \quad d_6 = -1.38 \times 10^{-4}, \\ d_7 &= 2.42 \times 10^{-4}, \quad d_8 = -8.10 \times 10^{-4}, \quad d_9 = -0.664. \end{aligned} \quad (6)$$

- M_W is calculated from the Fermi constant G_μ [Awramik, et al., 2004]
- The deviations to the full calculation amount to average (maximal) 2×10^{-7} (1.3×10^{-6}), in the input parameter ranges.

Numerical 2-loop calculations

Our approach:

**Direct numerical calculations
in Minkowskian kinematics**

Direct numerical integrations in Minkowskian regions >NLO

Sector decomposition (SD)

Fiesta 3 [A.V.Smirnov, 2014]

SecDec 2 [S. Borowka, G. Heinrich, 2012]

SecDec 3 [S. Borowka, G. Heinrich, P. Jones, M. Kerner, J. Schlenk, T. Zirke, 2013]

- Nicodemos, ver 2.0 [A. Freitas]

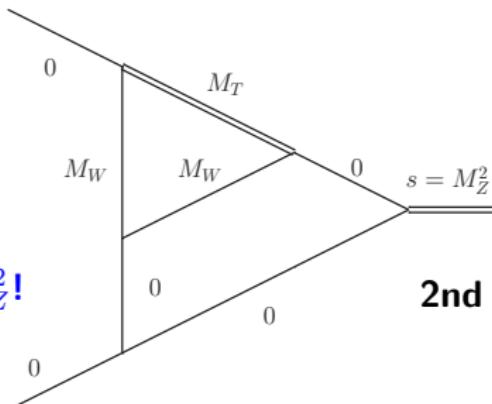
Now: the Mellin-Barnes (MB) method.

Toolbox: AMBRE/MB/MBnumerics/CUBA

Two steps (automatic construction and numerical evaluation):

- ① **AMBRE** [K.Kajda (planar, ver.2.2), E.Dubovik (non-planar,ver3.0)] - **PlanarityTest** [K.Bielas, E.Dubovik]
- ② **MBnumerics** [J. Usovitsch, E. Dubovik] - a completely new software !

One of the most difficult IR-divergent integrals with 2 scales



Huge oscillations for $s = M_Z^2$!

2nd problem: Γ 's.

$$\begin{aligned}
 & \int dz_1 \int dz_2 (-s)^{-2-2\epsilon} \left(-\frac{s}{MT^2}\right)^{-z_2} \left(-\frac{s}{MW^2}\right)^{-z_1} \Gamma[-z_1] \Gamma[-z_2] \Gamma[-z_3] \\
 \times & \frac{\Gamma[-1 - 2\epsilon - z_1 - z_2] \Gamma[-\epsilon - z_1 - z_2] \Gamma[2 + 2\epsilon + z_1 + z_2]}{\Gamma[-3\epsilon - z_1 - z_2] \Gamma[1 - 2\epsilon - z_1 - z_2] \Gamma[1 - z_3] \Gamma[-2\epsilon - 2z_1 - 2z_2 - z_3]} \\
 \times & \Gamma[-2\epsilon - z_1 - 2z_2 - z_3] \Gamma[-1 - 2\epsilon - z_1 - z_2 - z_3] \\
 \times & \Gamma[-\epsilon - z_1 - z_2 - z_3] \Gamma[1 + z_2 + z_3] \Gamma[1 + \epsilon + z_1 + z_2 + z_3]
 \end{aligned}$$

Solutions: see Backup Slides and MBnumerics.m

One of the most difficult IR-divergent integrals with 2 scales, cont'd

MBnumerics.m

2016-04-21

Johann Usovitsch

=

$$1.541402128186602 + 0.248804198197504*I$$

+

$$0.12361459942846659 - 1.0610332704387688 *I * \text{eps}^{-1}$$

+

$$-0.33773737955057970 + 3.6*10^{-17}*I * \text{eps}^{-2}$$

Time needed **43 min.**

SecDec

=

$$1.541 + 0.2487*I$$

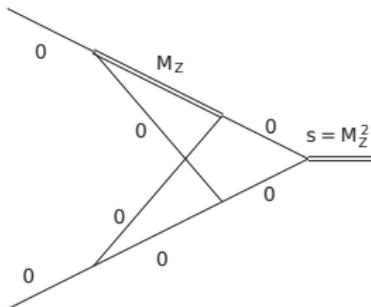
+

$$0.123615 - 1.06103*I * \text{eps}^{-1}$$

+ $-0.3377373796 - 5*10^{-10}*I*\text{eps}^{-2}$

Time needed **24 hours**

The worst case for SD, fine with MBnumerics



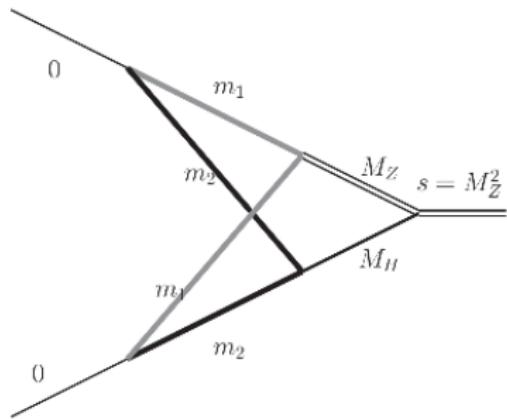
Euclidean results (constant part):

Analytical :	-0.4966198306057021
MB(Vegas) :	-0.4969417442183914
MB(Cuhre) :	-0.4966198313219404
FIESA :	-0.4966184488196595
SecDec :	-0.4966192150541896

Minkowskian results (constant part):

Analytical :	$-0.778599608979684 - 4.123512593396311 \cdot i$
MBnumerics :	$-0.778599608324769 - 4.123512600516016 \cdot i$
MB(Vegas) :	big error
MB(Cuhre) :	NaN
FIESA :	big error
SecDec :	big error

8-dim MB integral (less accurate) for the $Z\bar{b}b$ vertex



$$m_1 = M_t, m_2 = M_W$$

The integrals contain up to three dimensionless parameters:

$$\left\{ \frac{M_H^2}{M_Z^2}, \frac{M_W^2}{M_Z^2}, \frac{m_t^2}{M_Z^2}, \frac{(M_Z + i\varepsilon)^2}{M_Z^2} \right\}$$

Important

MB and SD methods are very much complementary!

- MB works well for hard threshold, on-shell cases, not many internal masses, SD is powerful for integrals with internal masses.
- see e.g.: J. Gluza, K. Kajda, T. Riemann and V. Yundin
"Numerical Evaluation of Tensor Feynman Integrals in Euclidean Kinematics"
Eur. Phys. J. C **71** (2011) 1516; [arXiv:1010.1667 [hep-ph]]

10^{-8} accuracy achieved for **any** self-energy and vertex Feynman integral
with one of the methods.

How to unfold - rough scheme

We have to describe

$$e^+ e^- \rightarrow (\gamma, Z) \rightarrow f^+ f^-(\gamma), \quad (7)$$

S-matrix Ansatz in the complex energy plane

$$\begin{aligned} \mathcal{A}_{e^+ e^- \rightarrow b\bar{b}} &= \underbrace{\frac{R_Z}{s - s_Z} + \frac{R_\gamma}{s}}_{\gamma-Z \text{ interference}} + \overbrace{S + (s - s_Z)S' + \dots}^{\text{Background}}, \\ s_Z &= \overline{M}_Z^2 - i\overline{M}_Z\overline{\Gamma}_Z \end{aligned}$$

- R, S, S', \dots are individually gauge-invariant and UV-finite - **unitarity and analyticity of the S-matrix**. IR-finite, when soft and collinear real photon emission is added. [Willenbrock, Valencia, 1991] [Sirlin, 1991] [Stuart, 1991]

[Riemann, 1991, 1992] [H. Veltman, 1994] [Passera, Sirlin, 1998] [Gambino, Grassi, 2000]

[Awramik, Czakon, Freitas, 2006].

The term $R_\gamma(s)/s$ is part of the background

- The poles of \mathcal{A} have complex residua R_Z and R_γ .
- There is only ONE pole in mathematics, while in physics we observe two of them: photon exchange at $s = 0$, Z exchange at $s_0 = s_Z$. Mathematically, the appearance of the photon pole is result of summing of part of background around Z pole, $s_0 = s_Z$

[T. Riemann, APPB 2015]

$$\begin{aligned}
 \frac{R_\gamma(s)}{s} &= \frac{\sum_{n=0}^{\infty} R_n (s - s_0)^n}{s} \\
 &= \frac{\sum_{n=0}^{\infty} R_n (s - s_0)^n}{s_0 - (s_0 - s)} \\
 &= \sum_{n=0}^{\infty} R_n (s - s_0)^n \frac{1}{s_0} \frac{1}{1 - \frac{s_0 - s}{s_0}} \\
 &= \sum_{n=0}^{\infty} R_n (s - s_0)^n \frac{1}{s_0} \left[1 + \frac{s_0 - s}{s_0} + \left(\frac{s_0 - s}{s_0} \right)^2 \dots \right];
 \end{aligned}$$

Conclusions

- We used numerical approach to $Z \rightarrow bb$ based on MB and SD methods.
- The main challenge was the calculation of massive two-loop vertex diagrams, now AMBRE/MB/MBnumerics/CUBA works with 8 digits and for MB integrals of $\text{dim} < 5$.
- New automatized tools **AMBRE 3** and **MBnumerics** for the evaluation of the Mellin-Barnes integrals in Minkowskian kinematics together with sector decomposition programs **SecDec 3** and **Fiesta 3** are sufficient to calculate all needed integrals for Z resonance physics.
- Continuum physics (cross sections) needs also 2-loop boxes, **this has to be studied**.
- Final calculation at two-loop order to the electroweak effective weak mixing angle $\sin^2 \theta_{\text{eff}}^b$ is presented as a simple fitting formula

Outlook

Further plans connected also with FCC:

- ① Evaluation of other pseudoobservables, Γ_{Zbb} , $\Gamma_{Z_{\text{tot}}}$, ...
- ② S-matrix theory: exact two-loops description of the Z-physics resonance,
e.g. A. Leike, T. Riemann, and J. Rose, "S-matrix approach to the Z line shape",
Phys. Lett. B273 (1991) 513, [hep-ph/9508390];
T. Riemann, "S-matrix Approach to the Z Resonance", APPB46 (2015) 11, 2235;
DFGRU, PoS LL2016 (2016) 075: "30 years, some 700 integrals, and 1 dessert, or:
Electroweak two-loop corrections to the $Z\bar{b}b$ vertex", arXiv:1610.07059;
- ③ Further development of MB tools;
- ④ Further applications: e.g. including box diagrams (cross sections).
- ⑤ Open. Finally, two-loop EW enough for FCC? N^3LO with
AMBRE/MB/MBnumerics? (Really exciting prospect!)

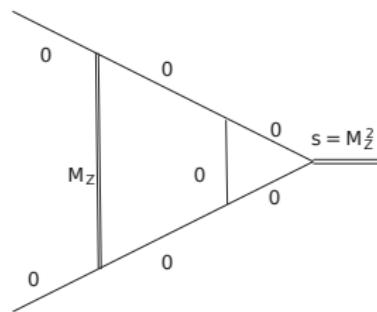
Note recent progress in differential equation method of solving Feynman (master) integrals: Henn 2013 (concept), Prausa 2017 (algorithm)

Backup slides

Basic problem: Steping up from Euclidean to direct calculation in Minkowskian kinematics

$$\frac{1}{(-)p^2 - m^2} \longrightarrow \text{singularities} \longrightarrow \frac{1}{(-)p^2 - m^2 + i\delta}$$

Resonance: $s = M_Z^2$, $s = -M_Z^2$



Step 1

Construction of MB integrals

<http://us.edu.pl/~gluza/ambre/>

Mellin-Barnes representations in HEP - method

- "Om definita integraler", R. H. Mellin, Acta Soc. Sci. Fenn. 20(7), 1 (1895),
 "The theory of the gamma function", E. W. Barnes Messenger Math. 29(2), 64 (1900).

$mathematics \rightarrow \frac{1}{(A+B)^\lambda} = \frac{1}{\Gamma(\lambda)} \frac{1}{2\pi i} \int_{-i\infty}^{+i\infty} dz \Gamma(\lambda+z) \Gamma(-z) \frac{B^z}{A^{\lambda+z}}$
$physics \rightarrow \frac{1}{(p^2 - m^2)^a} = \frac{1}{\Gamma(a)} \frac{1}{2\pi i} \int_{-i\infty}^{+i\infty} dz \Gamma(a+z) \Gamma(-z) \frac{(m^2)^z}{(p^2)^{a+z}}$

It is recursive \implies multidimensional complex integrals.

$$\int_{-\frac{1}{3}-i\infty}^{-\frac{1}{3}+i\infty} dz_1 \int_{-\frac{2}{3}-i\infty}^{-\frac{2}{3}+i\infty} dz_2 \left(\frac{-s}{M_Z^2} \right)^{-z_1} \frac{\Gamma[-z_1]^3 \Gamma[1+z_1] \Gamma[z_1-z_2] \Gamma[-z_2]^3 \Gamma[1+z_2] \Gamma[1-z_1+z_2]}{s \Gamma[1-z_1]^2 \Gamma[-z_1-z_2] \Gamma[1+z_1-z_2]}$$

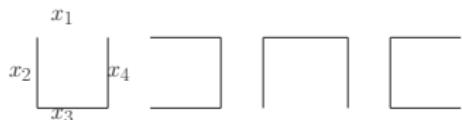
Overlaped integrals

Multiloop Feynman diagrams, general MB integrals

$$\frac{1}{D_1^{n_1} D_2^{n_2} \dots D_N^{n_N}} \rightarrow \int \prod_{j=1}^N dx_j x_j^{n_j-1} \delta(1 - \sum_{i=1}^N x_i) \frac{U(x)^{N_\nu - d(L+1)/2}}{F(x)^{N_\nu - dL/2}}$$

$N_\nu = n_1 + \dots + n_N$

The functions U and F are called graph or Symanzik polynomials.



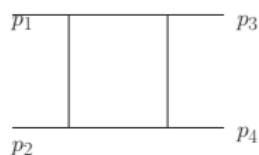
$$\mathbf{U} = \mathbf{x}_1 + \mathbf{x}_2 + \mathbf{x}_3 + \mathbf{x}_4 \quad ! \text{ 1-loop} \longrightarrow 1$$

Trees contributing to the polynomial U for the square diagram



$$\mathbf{F} = \mathbf{t} \cdot \mathbf{x}_1 \mathbf{x}_3 + \mathbf{s} \cdot \mathbf{x}_2 \mathbf{x}_4$$

2 – trees contributing to the polynomial F for the square diagram



Dimension of MB integrals depends on factorizations of F and U !

Cuts of internal lines such that:

- U : (i) every vertex is still connected to every other vertex by a sequence of uncut lines; (ii) no further cuts without violating (i)
- F : (iii) divide the graph into two disjoint parts such that within each part (i) and (ii) are obeyed and such that at least one external momentum line is connected to each part;

Step 2

Numerics of MB integrals

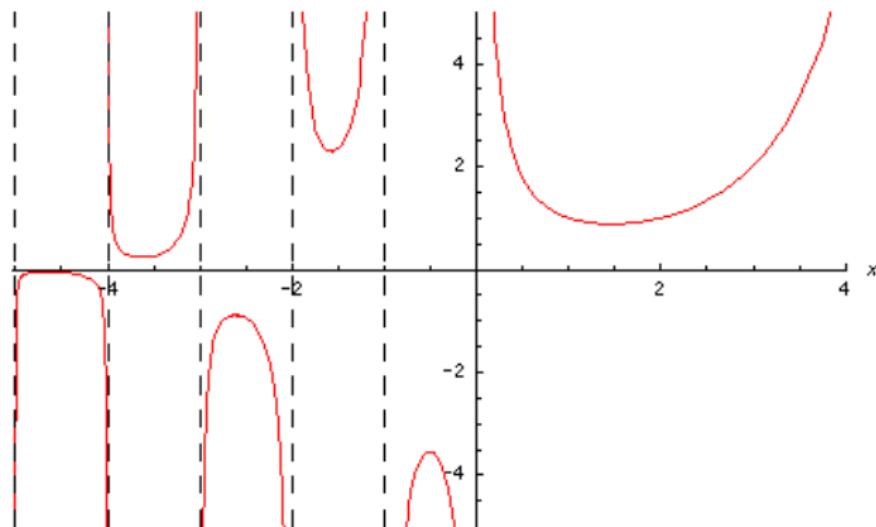
<http://mbtools.hepforge.org/>

Gamma function: Singularities in the complex plane

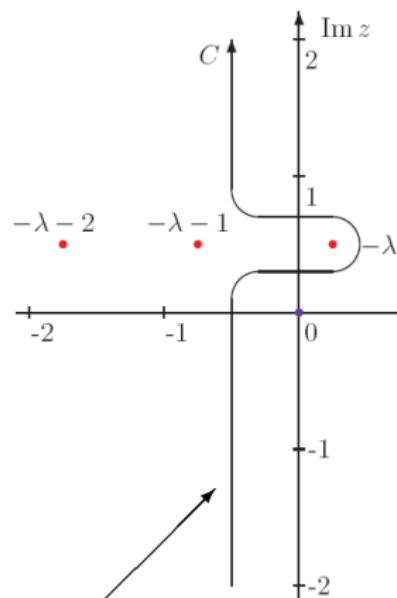
$$\Gamma(z) = \int_0^\infty t^{z-1} e^{-t} dt$$

$$\int dz \Gamma[z + \lambda]$$

SINGULARITIES



REGULAR



Contours: shifts, deformations

(* shifting contours *)

In[203]:=

```
sim = Gamma[-z]
```

Out[203]=

```
Gamma[-z]
```

In[227]:=

```
Sum[-Residue[Gamma[-z], {z, n}], {n, 0, 100}] // N
```

Out[227]=

```
0.367879
```

In[226]:=

```
n1 = NIntegrate[
  1 / (2 Pi) sim /. z → -1 / 20 + I y, {y, -10, 10}]
```

Out[226]=

```
0.367879 + 0. i
```

In[230]:=

```
n2 = NIntegrate[
  1 / (2 Pi) sim /. z → 1 / 20 + I y, {y, -10, 10}]
```

Out[230]=

```
-0.632121 + 0. i
```

In[231]:=

```
n2 - n1
```

Out[231]=

```
-1. + 0. i
```

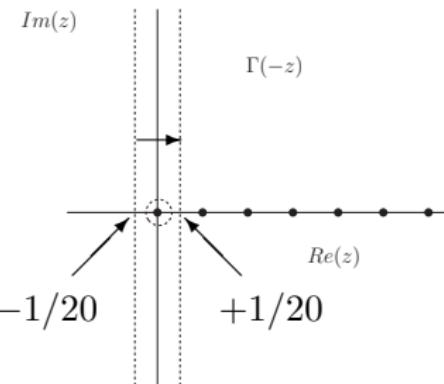
In[232]:=

```
Residue[sim, {z, 0}]
```

Out[232]=

```
-1
```

(* B512m2 *)



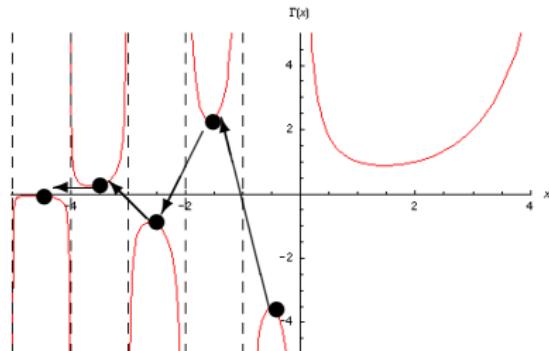
Two basic observations for shifting z follows

$$\int dz_1 \dots dz_k \dots I(\dots, \text{Re}[z_k] + n + \text{Im}[z_k], \dots) \quad I_{\text{orig}}$$

$$= \text{Residue} \left[\int dz_1 \dots dz_k \dots I \right]_{\text{Re}[z_k] + n} \quad I_{\text{Res}}$$

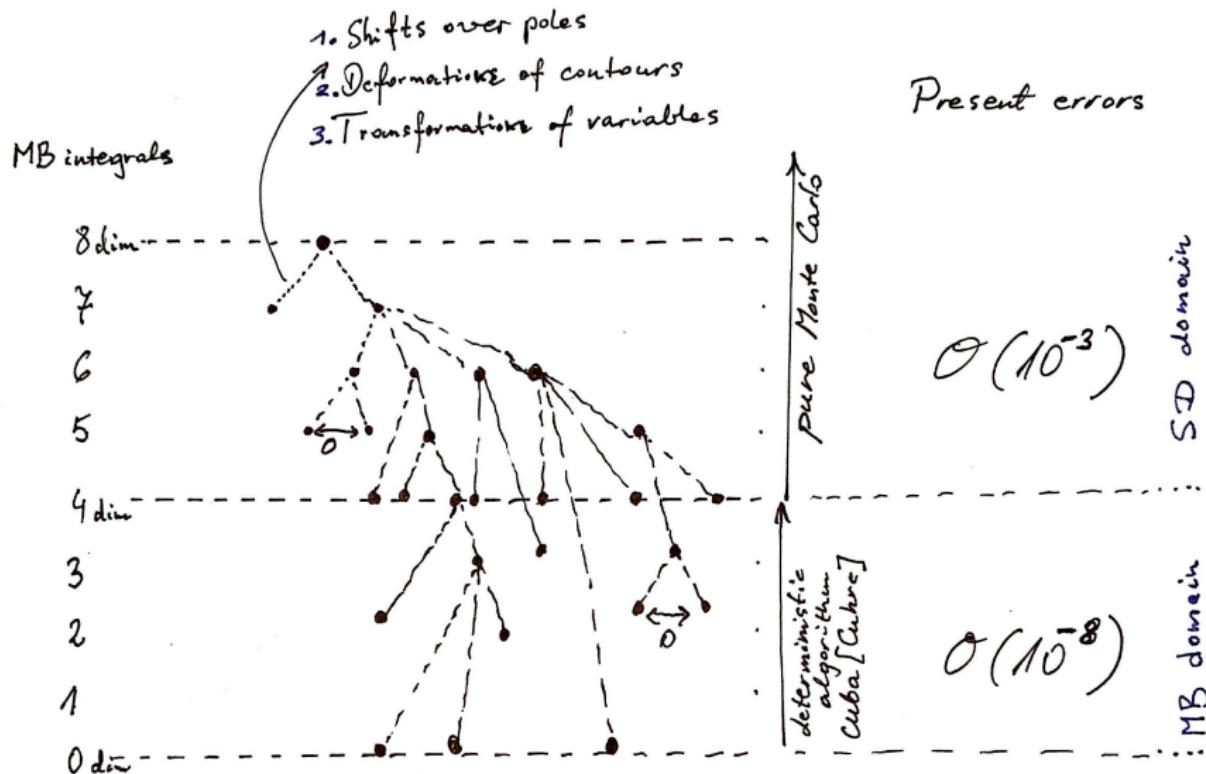
$$+ \int dz_1 \dots dz_k \dots I(\dots, \text{Re}[z_k] + (n+1) + \text{Im}[z_k], \dots) \quad I_{\text{new}}$$

- ① Residues **lower** dimensionality of original MB integrals.
- ② Integral after passing a pole (proper shifts) **can be made smaller**.



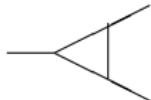
Top-bottom approach to evaluation of multidimensional MB integrals

MBnumerics.m - I. Dubovsky, J. Usovitsch, T. Riemann



BASIC PROBLEMS in Minkowski kinematics

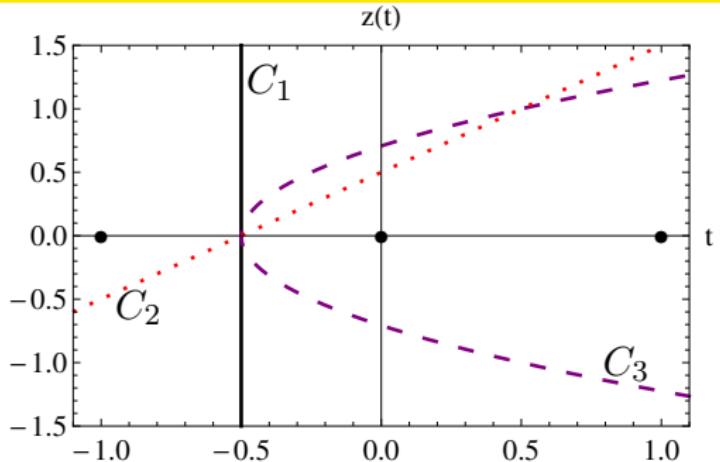
- I. Bad oscillatory behavior of integrands;
- II. Fragile stability for integrations over products and ratios of Γ functions.



$$\begin{aligned}
 V(s) &= \frac{e^{\epsilon\gamma_E}}{i\pi^{(4-2\epsilon)/2}} \int \frac{d^{(4-2\epsilon)}k}{[(k+p_1)^2 - m^2][k^2][(k-p_2)^2 - m^2]} \\
 &= \frac{V_{-1}(s)}{\epsilon} + V_0(s) + \dots,
 \end{aligned}$$

$$\begin{aligned}
 V_{-1}(s)|_{m=1} &= -\frac{1}{2s} \int_{-\frac{1}{2}-i\infty}^{-\frac{1}{2}+i\infty} \frac{dz}{2\pi i} \underbrace{(-s)^{-z}}_{\text{Problem I}} \overbrace{\frac{\Gamma^3(-z)\Gamma(1+z)}{\Gamma(-2z)}}^{\text{Problem II}} \\
 &= \frac{1}{2} \sum_{n=0}^{\infty} \frac{s^n}{\binom{2n}{n} (2n+1)} = \frac{2 \arcsin(\sqrt{s}/2)}{\sqrt{4-s}\sqrt{s}},
 \end{aligned}$$

Contour deformations



$$z(t) = x_0 + it : \quad V_{-1}^{C_1}(s) = \int_{-\infty}^{+\infty} (i) dt J[z(t)];$$

$$z(t) = x_0 + \theta t + it : \quad V_{-1}^{C_2}(s) = \int_{-\infty}^{+\infty} (\theta + i) dt J[z(t)]$$

$$z(t) = x_0 + at^2 + it : \quad V_{-1}^{C_3}(s) = \int_{-\infty}^{+\infty} (2at + i) dt J[z(t)]; .$$

$$s = 2, z(t) = \Re[-1/2] + i y, \quad y \in (-a, +a)$$

$$V_{-1}(2)|_{\text{analyt.}} = \mathbf{0.78539816339744830962} = \frac{\pi}{4}$$

$$V_{-1}(2)|_{\text{Pantis}}^{MB.m} = 0.7925 - 0.0225 i$$

$$V_{-1}(2)|_{C_1, a=15} = 0.7548660085063523 - \underline{0.229985258820015} i$$

$$V_{-1}(2)|_{C_1, a=10^2} = 0.73479313088852537844 + \underline{0.074901423602937676597} i$$

$$V_{-1}(2)|_{C_1, a=10^3} = 0.84718185073531076915 - \underline{0.094865760649354977853} i$$

$$V_{-1}(2)|_{C_1, a=10^4} = 4.4574554985139977188 + \underline{4.5139812364645122275} i$$

✓ $V_{-1}(2)|_{C_2} = \mathbf{0.7853981633859819} - 5.420140575251864 \cdot 10^{-15} \checkmark i$

✓ $V_{-1}(2)|_{C_3} = \mathbf{0.7853981632958756} + 2.435551760271437 \cdot 10^{-15} \checkmark i$

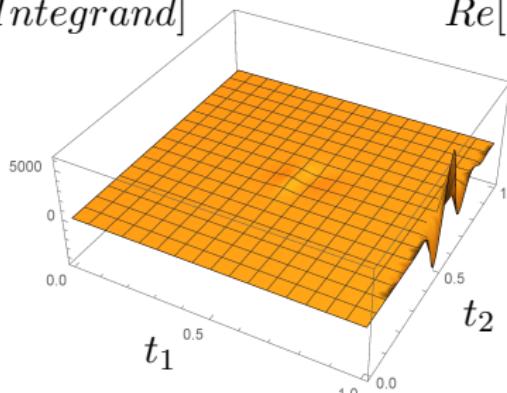
Transformations of integration variables (Mappings)

$$\int_{-\frac{1}{3}-i\infty}^{-\frac{1}{3}+i\infty} dz_1 \int_{-\frac{2}{3}-i\infty}^{-\frac{2}{3}+i\infty} dz_2 \left(\frac{-s}{M_Z^2} \right)^{-z_1} \frac{\Gamma[-z_1]^3 \Gamma[1+z_1] \Gamma[z_1-z_2] \Gamma[-z_2]^3 \Gamma[1+z_2] \Gamma[1-z_1+z_2]}{s \Gamma[1-z_1]^2 \Gamma[-z_1-z_2] \Gamma[1+z_1-z_2]}$$

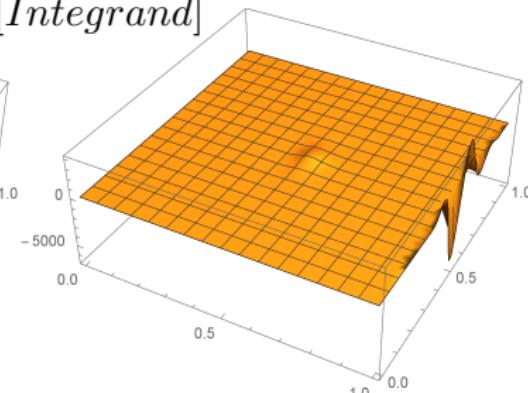
Logarithmic (in MB.m, M. Czakon, CPC 2006):

$$z_k = x_k + i \ln \left(\frac{t_k}{1-t_k} \right), \quad t_k \in (0, 1), \quad \text{the Jacobians: } J_k(t_k) = \frac{1}{t_k(1-t_k)}.$$

Im[Integrand]



Re[Integrand]

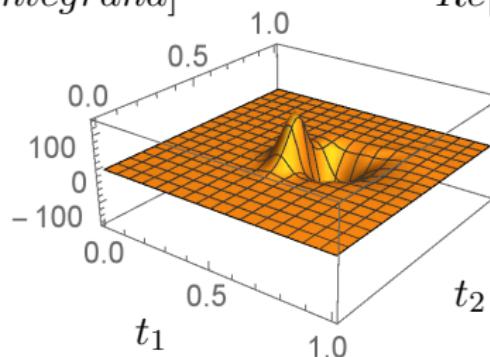


Transformations of variables (Mappings)

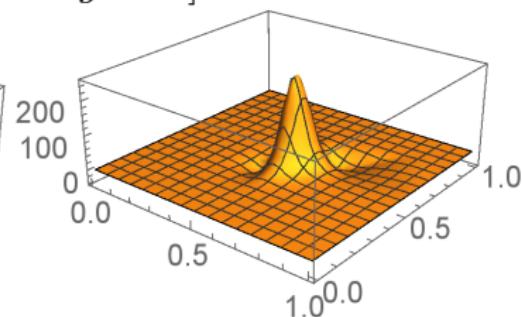
Tangent (in MBnumerics.m, ID, JU, TR, 2016):

$$z_k = x_k + i \frac{1}{\tan(-\pi t_k)}, \quad t_k \in (0, 1), \quad \text{the Jacobians : } J_k = \frac{\pi}{\sin^2[(\pi t_k)]}$$

Im[Integrand]



Re[Integrand]



In addition, $\Gamma \rightarrow e^{\ln \Gamma}$ improves numerical stability considerable, either.

The most difficult cases (for SD)

