# ΗH

#### report and plans

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LHCHXSWG meeting, October 2016

## HH in YR4

- recommendations for total rates in the SM (gluon fusion and rare channels)
- estimation of th. uncertainties in gluon fusion
- comparisons of available differential distributions
- recommendations for BSM benchmarks (EW singlet, 2HDM, effective Lagrangian)

### SM rates – a quick dictionary

from S. Jones

	$\sigma_{ m LO}~({ m fb})$	$\sigma_{ m NLO}~({ m fb})$	$\sigma_{ m NNLO}~({ m fb})$
HEFT	$17.07^{+30.9\%}_{-22.2\%}$	$\begin{array}{r} 31.93\substack{+17.6\%\\-15.2\%\\38.32\substack{+18.1\%\\-14.9\%}\end{array}$	$37.52^{+5.2\%}_{-7.6\%}$
B.I. HEFT	$19.85^{+27.6\%}_{-20.5\%}$	$38.32^{+18.1\%}_{-14.9\%}$	$\begin{array}{r} 37.52_{-7.6\%} \\ 43.63_{-7.6\%}^{+5.2\%} \ast \end{array}$
FTapprox	$19.85^{+27.6\%}_{-20.5\%}$	$34.26^{+14.7\%}_{-13.2\%}$	—
Full Theory	$19.85^{+27.6\%}_{-20.5\%}$	$32.91^{+13.6\%}_{-12.6\%}$	—
N.I. HEFT	_	$32.91^{+1\bar{3}.6\%}_{-12.6\%}$	$38.67^{+5.2\%}_{-7.6\%}*$

(Born/NLO Improved) Higgs Effective Field Theory -  $m_t \rightarrow \infty$  (finite at LO/NLO)

FTapprox. – complete NLO real emissions, approx. virtual contributions

\* reweighted to full NLO

Glover, van der Bij 88, Dawson, Dittmaier, Spira 98 Maltoni, Vryonidou, Zaro 14, de Florian, Grazzini, Hanga, Kallweit, Lindert, Maierhöfer, Mazzitelli, Rathlev 16 Borowka, Greiner, Heinrich, Kerner, Schlenk, Schubert, Zirke 16

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LHCHXSWG recommendations for  $\sqrt{s}$  = 14 TeV, m<sub>h</sub>=125 GeV,  $\mu_R = \mu_F = m_{hh}/2$ 

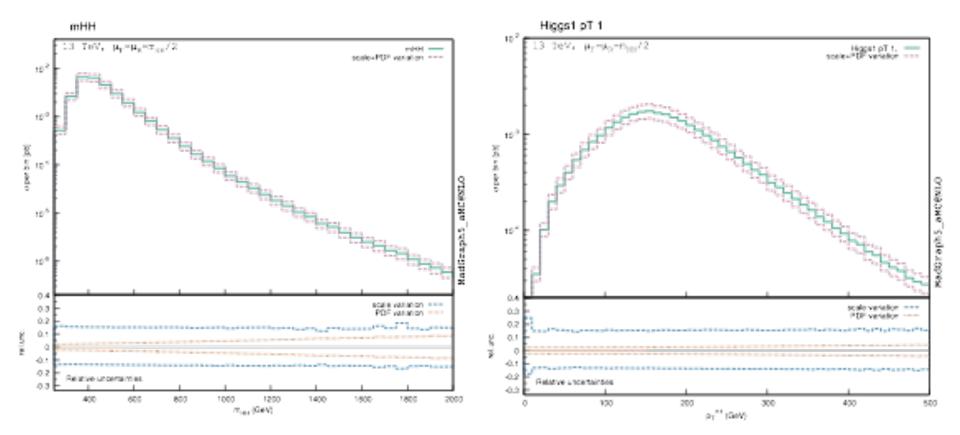
$$\sigma_{\text{NNLO+NNLL}} = 39.56^{+4.4\%}_{-6.0\%}$$
 fb

#### Gluon fusion cross-sections

NNI	NNLL+NNLO with top mass effects at NLO arXiv				
$m_h = 124.5 \text{ GeV}$	$\sigma'_{\rm NNLL}(fb)$	Scale Unc. (%)	PDF Unc. (%)	$\alpha_s$ Unc. (%)	
$\sqrt{s} = 7 \text{ TeV}$	7.132	+4.0 - 5.7	$\pm 3.4$	$\pm 2.8$	
$\sqrt{s} = 8 \text{ TeV}$	10.24	+4.1 - 5.7	$\pm 3.0$	$\pm 2.6$	
$\sqrt{s} = 13 \mathrm{TeV}$	33.70	+4.3 - 6.0	$\pm 2.1$	$\pm 2.3$	
$\sqrt{s} = 14  { m TeV}$	39.85	+4.4 - 6.0	$\pm 2.1$	$\pm 2.2$	
$m_h = 125~{ m GeV}$	$\sigma'_{\rm NNLL}(fb)$	Scale Unc. (%)	PDF Unc. (%)	$\alpha_s$ Unc. (%)	
$\sqrt{s} = 7 \text{ TeV}$		+4.0 - 5.7	$\pm 3.4$	$\pm 2.8$	
$\sqrt{s} = 8 \text{ TeV}$	10.16	+4.1 - 5.7	$\pm 3.1$	$\pm 2.6$	
$\sqrt{s}=13{ m TeV}$	33.45	+4.3 - 6.0	$\pm 2.1$	$\pm 2.3$	
$\sqrt{s} = 14  { m TeV}$	39.56	+4.4 - 6.0	$\pm 2.1$	$\pm 2.2$	
$m_h = 125.09~{ m GeV}$	$\sigma'_{ m NNLL}(fb)$	Scale Unc. (%)	PDF Unc. (%)	$\alpha_s$ Unc. (%)	
$\sqrt{s} = 7 \text{ TeV}$	7.068	+4.0 - 5.7	$\pm 3.4$	$\pm 2.8$	
$\sqrt{s} = 8 \text{ TeV}$	10.15	+4.1 - 5.7	$\pm 3.1$	$\pm 2.6$	
$\sqrt{s}=13{ m TeV}$	33.41	+4.3 - 6.0	$\pm 2.1$	$\pm 2.3$	
$\sqrt{s} = 14  { m TeV}$	39.51	+4.4 - 6.0	$\pm 2.1$	$\pm 2.2$	
$m_h = 125.5~{ m GeV}$	$\sigma'_{\rm NNLL}(fb)$	Scale Unc. (%)	PDF Unc. (%)	$\alpha_s$ Unc. (%)	
$\sqrt{s} = 7 \text{ TeV}$	7.023	+4.0 - 5.7	$\pm 3.4$	$\pm 2.8$	5% th.
$\sqrt{s} = 8 \text{ TeV}$	10.09	+4.1 - 5.7	$\pm 3.1$	$\pm 2.6$	
$\sqrt{s} = 13  { m TeV}$	33.21	+4.3 - 6.0	$\pm 2.1$	$\pm 2.3$	incertainties
$\sqrt{s} = 14  { m TeV}$	39.27	+4.4 - 5.9	$\pm 2.1$	$\pm 2.2$	5

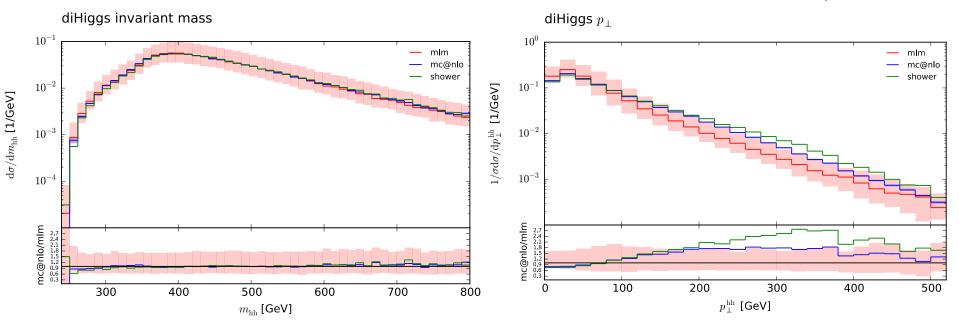
#### **Differential distributions**

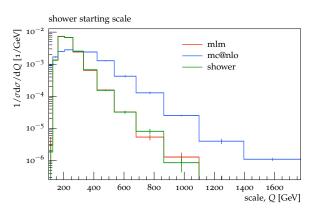
FTapprox from Madgraph5\_aMC@NLO matched to Pythia 8



#### Towards estimating shower uncertainties

from A. Papaefstathiou





- "mc@nlo" (blue): MG5\_aMC@NLO samples,
- "mlm" (red): MLM-merged HERWIG samples,
- "shower" (green): LO+HERWIG parton shower samples.

#### BSM resonant benchmarks

• Singlet model

 $V = -m^2 \Phi^{\dagger} \Phi - \mu^2 S^2 + \lambda_1 (\Phi^{\dagger} \Phi)^2 + \lambda_2 S^4 + \lambda_3 \Phi^{\dagger} \Phi S^2$ 

• input parameters:  $m_h = 125 \text{ GeV}, M_H, \cos \alpha, v, \tan \beta = v/\langle s \rangle$ 

cross-section tables included in the Report

• 2HDM

	$\tan \beta$					$m_{12}^2$
B1	1.75	-0.5881	300	441	442	38300
B2	1.50	-0.6792	700	701	670	180000
B7	10.00	-0.5881 -0.6792 0.1015	500	500	500	24746

#### EFT benchmarks

Effective Lagrangian for a CP even Higgs

#### Many thanks for contributions to YR4!

J. Adelman, A. Apyan, J. Baglio, A. Carvalho, D. de Florian, M. Dall'Osso, T. Dorigo, F. Goertz, C. Gottardo, J. Grigo, R. Gröber, P. Hebda, G. Heinrich, B. Hespel, N. Konstantinidis, I. Lewis, J. Mazzitelli, M. Mühlleitner, N. Styles, A. Papaefstathiou, J. Rojo, M. Spannowsky, M. Spira, M. Tosi, C. Vernieri, E. Vryonidou, M. Zaro, T. Zirke.

#### Now we are before YR5...

### Future plans

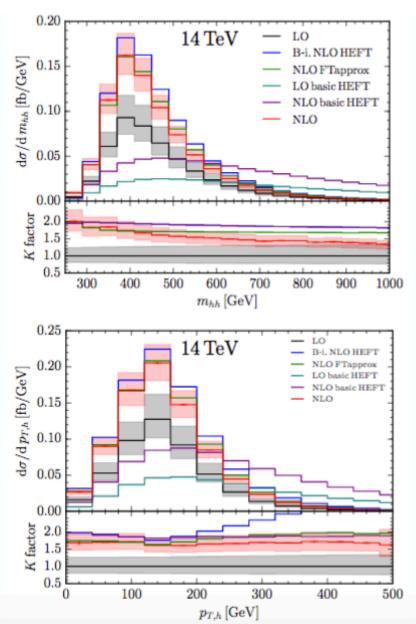
- 1. distributions at NLO (finite top mass effects)
- 2. distributions at NNLO (HEFT)
- 3. bridging parton level to hadron level
- 4. general directions for BSM
- 5. additional processes? GF+2j, more WBF, VVHH?

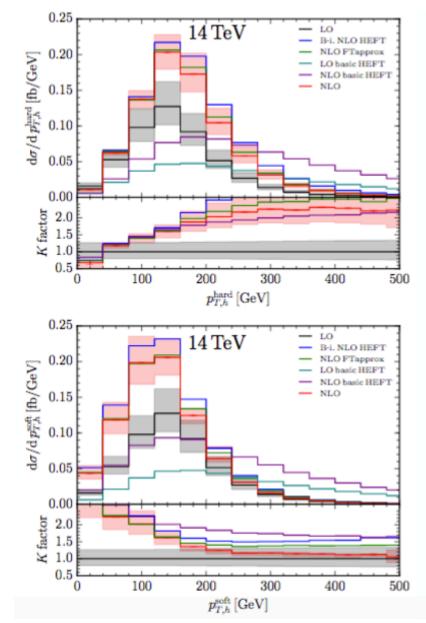
based on contributions and discussions during the meeting https://indico.cern.ch/event/573916/

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#### **NLO** distributions

#### from S. Jones





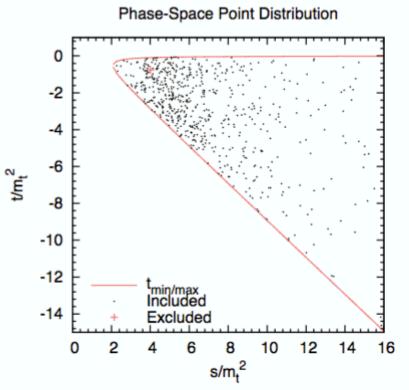
# Phase space sampling for NLO calculations

Events for virtual:

1) VEGAS algorithm applied to LO matrix element  $\mathcal{O}(100k)$  events computed

2) Using LO events unweighted events generated using accept/reject method  $\mathcal{O}(30k)$  events remain

3) Randomly select 665 Events, compute at NLO



Median GPU time per PS point: 2 hours Total compute time used: 4680 GPU Hours Wall time: **6 days**  from S. Jones

### Possible future improvements

- Differential distributions are a bottle neck since much CPU needed for phase space sampling
- Cannot be matched to PS and used out-of-thebox
- Investigating the use of the Grid infrastructure
- PS matching (in progress, may need grid)
- Improving N.I. HEFT for fully differential distributions (see later)

# Towards differential distributions at NNLO HEFT

NNLO + threshold resummation at NNLL available — Large increase (~20%) w.r.t. NLO

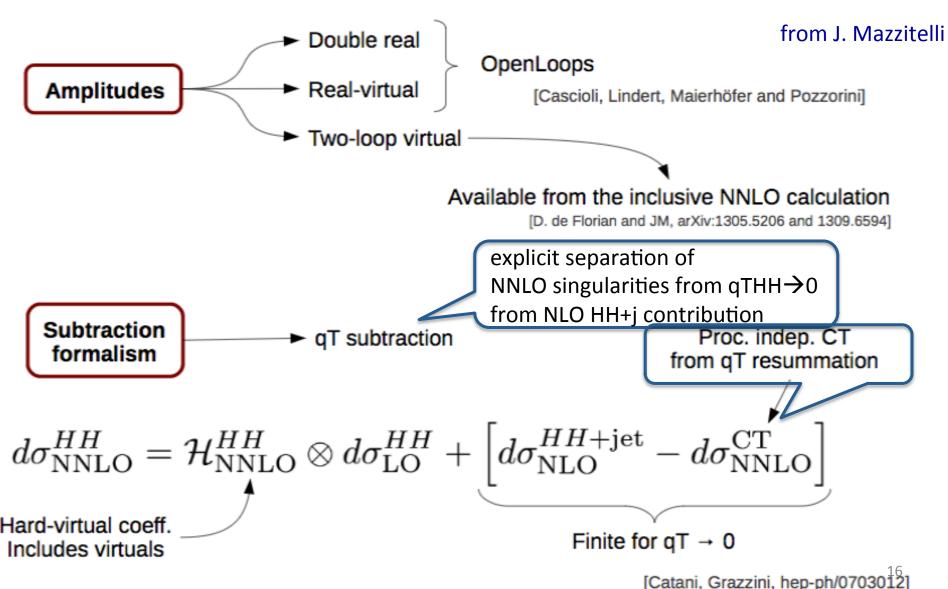
Substantial reduction of scale uncertainty (~±5%)

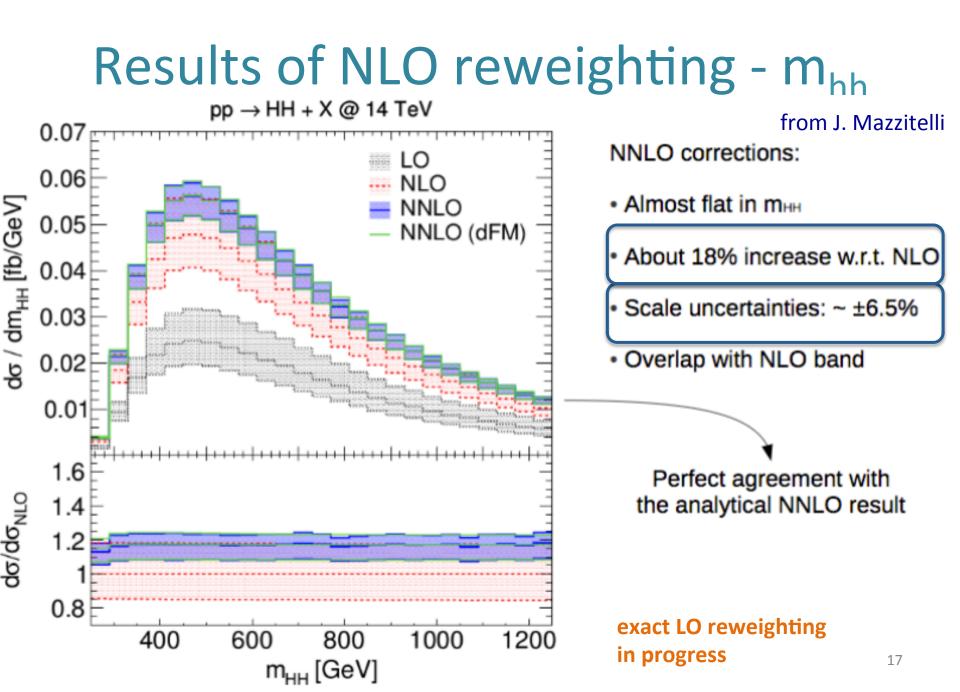
BSM: NNLO including dim 6 operators in progress [de Florian, Fabre, Mazzitelli] [D. de Florian and JM, arXiv:1505.07122]

Fully differential calculation at NNLO in QCD is desirable

[de Florian, Grazzini, Hanga, Kallweit, Lindert, Maierhoefer, JM and Rathlev, arXiv:1606.09519]

#### **NNLO HEFT - technicalities**



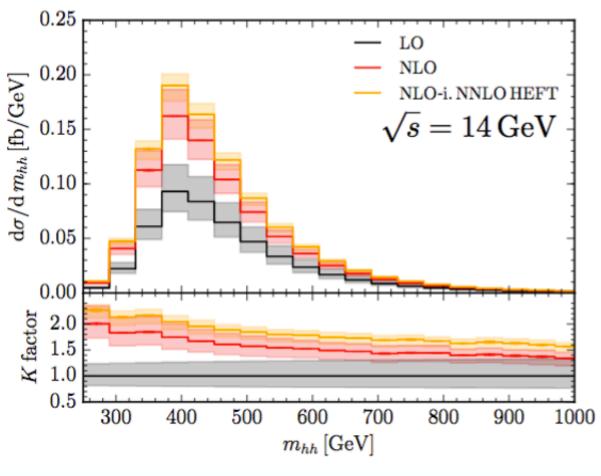


### Results of NLO reweighting - p<sub>h</sub><sup>T</sup>

from J. Mazzitelli

 $pp \rightarrow HH + X @ 14 TeV$  $pp \rightarrow HH + X @ 14 TeV$ LO О. NLO NLO dơ / dp<sub>T, нн</sub> [fb/GeV] tb/GeV NNLO NNLO 10 10<sup>-1</sup> dσ / dp\_ 10-2  $10^{-2}$ 10<sup>-3</sup> 10<sup>-3</sup>  $p_i^{\rm T} > 30 \,\,{\rm GeV}, \, |\eta_j| < 4.4, \, R = 0.4$ dσ/dσ<sub>NLO</sub> dσ/dσ<sub>NLO</sub> 2 2 100 200 300 400 500 600 700 800 100 200 300 400 500 600 700 800 ۲Ô р<sub>т. нн</sub> [GeV] p<sub>T,i</sub> [GeV] NNLO corrections up to 80-60%, scale uncertainties ~30-40% 18

#### Combining the two improvements



# First attempt to combine full NLO

Borowka, Greiner, Heinrich, SPJ, Kerner, Schlenk, Zirke 16 +

#### NNLO HEFT (Differential)

de Florian, Grazzini, Hanga, Kallweit, Lindert, Maierhöfer, Mazzitelli, Rathlev 16

 $\frac{\mathrm{d}\sigma^{\mathrm{approx.}}}{\mathrm{d}m_{hh}} \equiv \frac{\mathrm{d}\sigma_{\mathrm{NLO}}}{\mathrm{d}m_{hh}} \times \frac{\mathrm{d}\sigma^{\mathrm{HEFT}}_{\mathrm{NLO}}/\mathrm{d}m_{hh}}{\mathrm{d}\sigma^{\mathrm{HEFT}}_{\mathrm{NLO}}/\mathrm{d}m_{hh}}$ 

**Bin-by-bin** rescaling of NLO by NNLO HEFT K-factor

#### Plans for improving the combination

- Combining NNLO<sub>mt $\rightarrow \infty$ </sub> with NLO<sub>finite mt</sub> :
  - Reweighting NNLO with the exact LO
  - With 1D NLO K-factors reweight m<sub>hh</sub> distribution and apply the correction to others

- Different procedures → estimation of the HEFT uncertainty for NNLO distributions
- Inclusion of Higgs decays

### Parton to hadron level

from A. Papaefstathiou

- Current state-of-the art of HH Monte Carlo simulations relies on full LO + full hard radiation + virtual corrections from HEFT.
- LO+PS vs. PS+full real radiation efficiencies can vary up to 20% in an LHC analysis.
- Differences between merged and matched samples due to uncertainties related to the shower starting scale.

### MC outlook

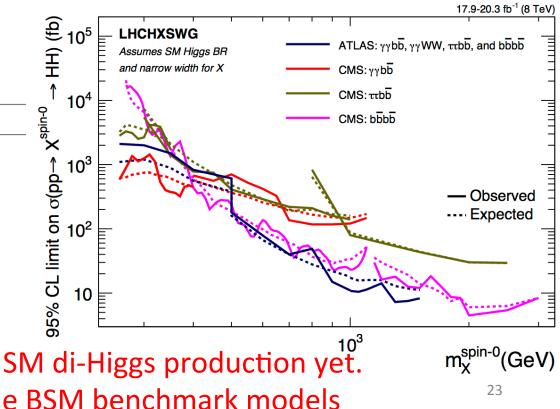
- Use the same inputs
- Discuss with experiments which MC generators and settings are used to include them in the comparison
- Can we gain in efficiency of signal selection with better MC modeling?

#### **Experimental searches**

#### Run 1 limits on non resonant SM-like cross-sections

Channel	Experiment	Observed (pb)	Expected (pb)	
$\gamma\gamma bar{b}$	ATLAS	2.2	1.0	
$\gamma\gamma bar{b}$	CMS	0.71	0.60	
$ au au bar{b}$	ATLAS	1.6	1.3	
au au bar b	CMS	0.59	0.94	
$bar{b}bar{b}$	ATLAS	0.62	0.62	
$\gamma\gamma WW^*$	ATLAS	11.0	6.7	
Combination	ATLAS	0.69	0.47	





Searches not sensitive to the SM di-Higgs production yet.  $\rightarrow$  need to develop reasonable BSM benchmark models

#### BSM models

• Inherit SM improvements in BSM benchmarks

- open questions about the progress:
  - new resonant models? need to work with other groups on 2HDM, MSSM, NMSSM benchmarks relevant for HH
  - + more EFT operators? (CP violation?)
  - + EFT operators in other production channels? (VVhh, tthh,...)

# Thoughts for general consideration before YR5

 harmonise BSM activities with WG2 and WG3 to make sure our benchmarks are not excluded already by single Higgs processes

 the necessity for providing recommendations for 4 mass points for rare processes (noninteger values a bottleneck for NLO computations)

## Conclusions

- Many theory developments already in YR4!
- Ongoing efforts in improving predictions for SM gluon fusion production
- Works towards estimating shower systematic
- Plans to improve BSM benchmarks with SM developments
- More discussions needed for BSM directions