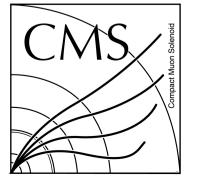
# Dark Matter @ 100 TeV





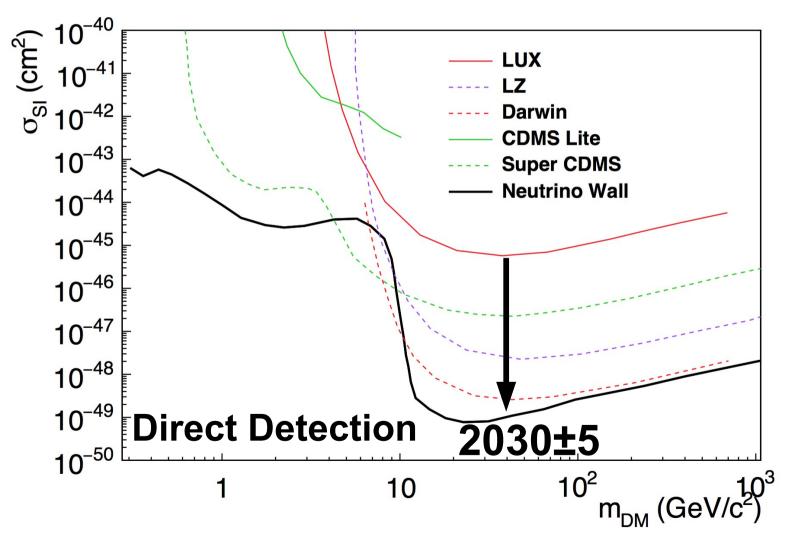
# Phil Harris (CERN)

w/help from K.Hahn(NWU) & MLM (CERN)



# Dark Matter searches not @ collider

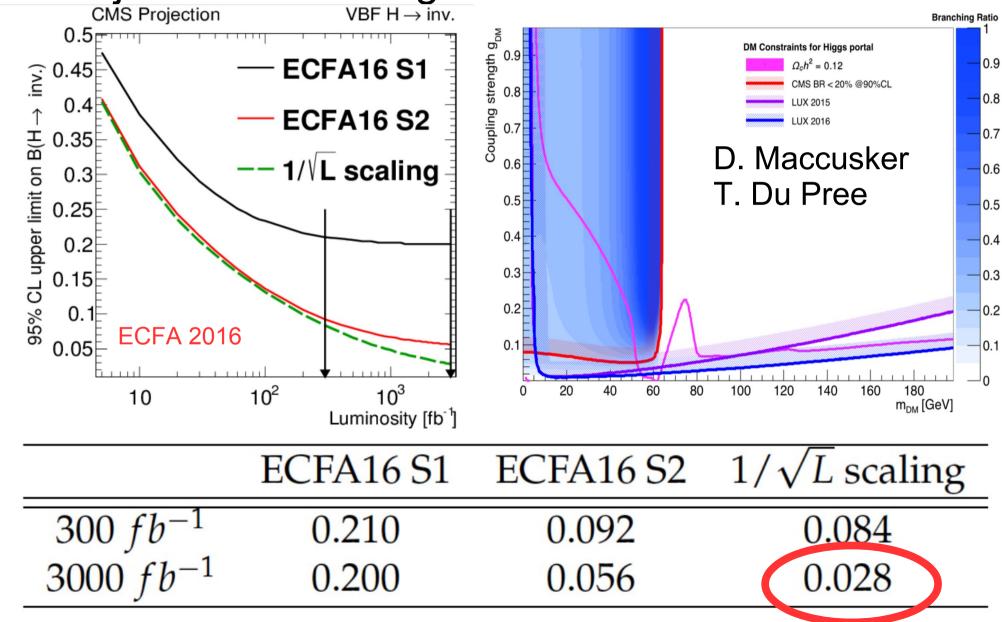
Dark matter searches not at colliders have clear benchmarks



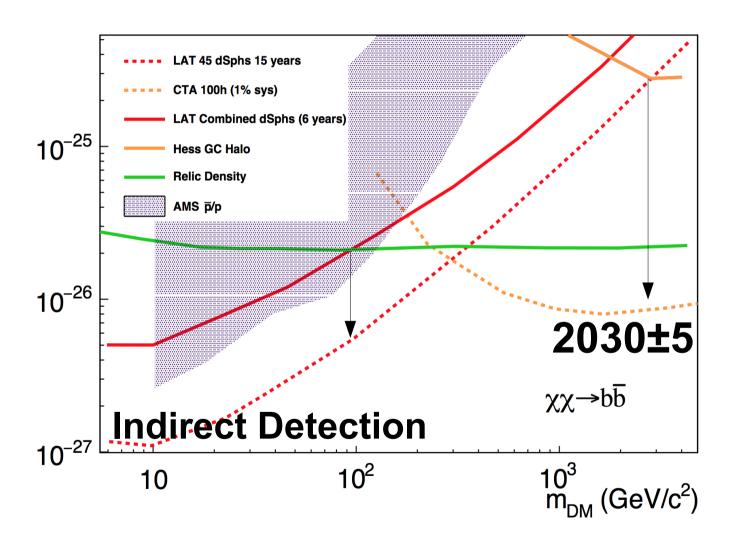
Goal: get to the Neutrino background wall

# Full Scaling expected Scaling

Projections at LHC go to <3%</li>



# Dark Matter searches not @ collider



Goal: get to the Relic density

### Question:

 Whats the simplest way to present LHC results in the context of Dark Matter?

### Question:

 Whats the simplest way to present LHC results in the context of Dark Matter?

- Answer:
  - $\mathsf{-}\ \mathsf{\sigma}_{\mathsf{Invisible}}$
- Assumes dark matter coupling to standard model

# Adding Dark Matter

- What drives dark matter interaction is production
  - Take the approach that this is defined by the mediator

### For a mediator search

• 
$$\mathcal{L}' = g_{DM} \chi \chi \Upsilon$$

Z'<sup>µ</sup> Spin 1

Uniform coupling to SM

$$\mathcal{L}' = \mathcal{L}' + g_{SM} Z'_{\mu} \overline{q} \gamma^{\mu} q$$

S Spin 0

Yukawa\* couplings to SM

$$\mathcal{L}' = \mathcal{L}' + g_{SM} S \overline{q} q$$

# Simplified Models 101

### Vector

$$g_{\rm DM} Z'_{\mu} \bar{\chi} \gamma^{\mu} \chi$$

EWK style coupling (equal to all quarks/leptons)

**Axial vector** 

$$g_{\rm DM} Z_{\mu}^{\prime\prime} \bar{\chi} \gamma^{\mu} \gamma^5 \chi$$

EWK style coupling (equal to all quarks/leptons)

### Scalar

$$g_{
m DM} S \, ar{\chi} \chi$$

Yukawa style coupling (Mass based coupling)

### Pseudoscalar

$$g_{\mathrm{DM}}P\,ar{\chi}\gamma^5\chi$$

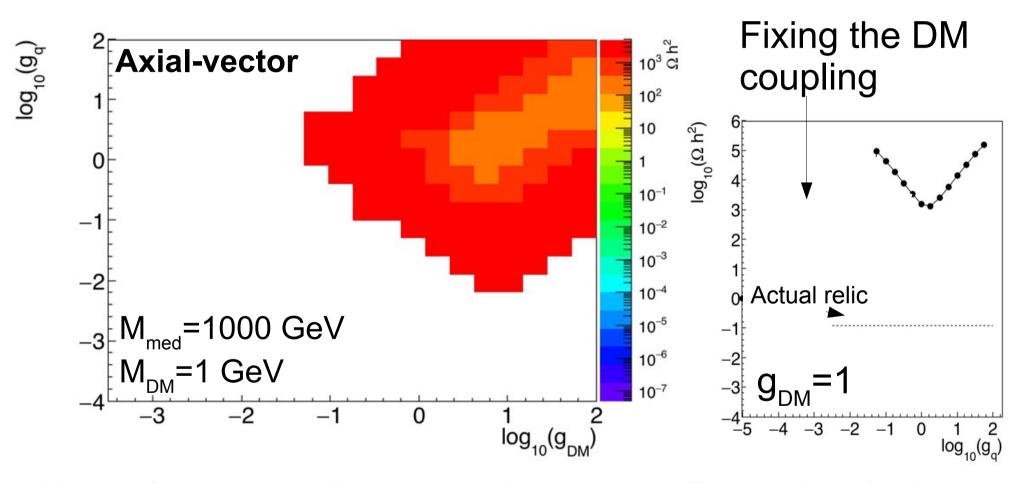
Yukawa style coupling (Mass based coupling)

# Establishing a collider benchmark

Relic Density??

# Solving for the relic

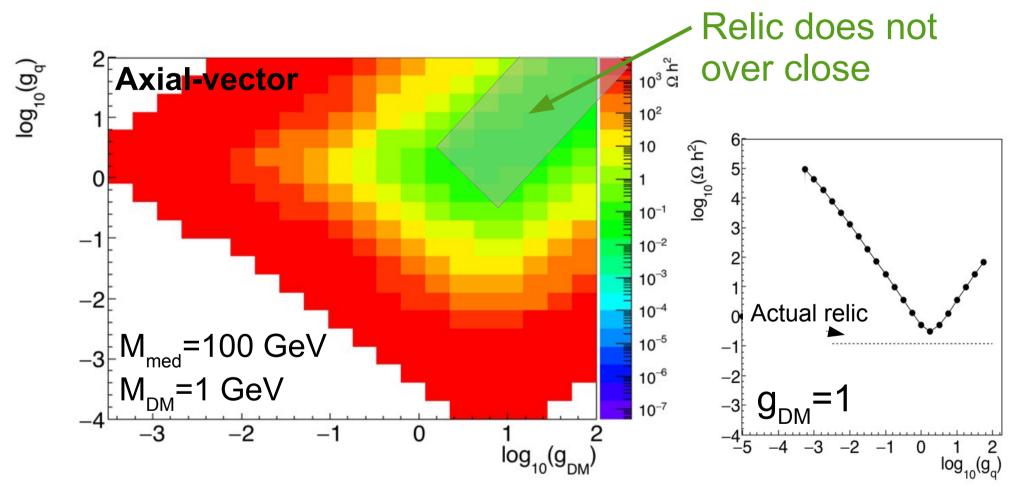
We can numerically solve for the solution



For a large mediator we have no allowed solution

# Solving for the relic

We can numerically solve for the solution

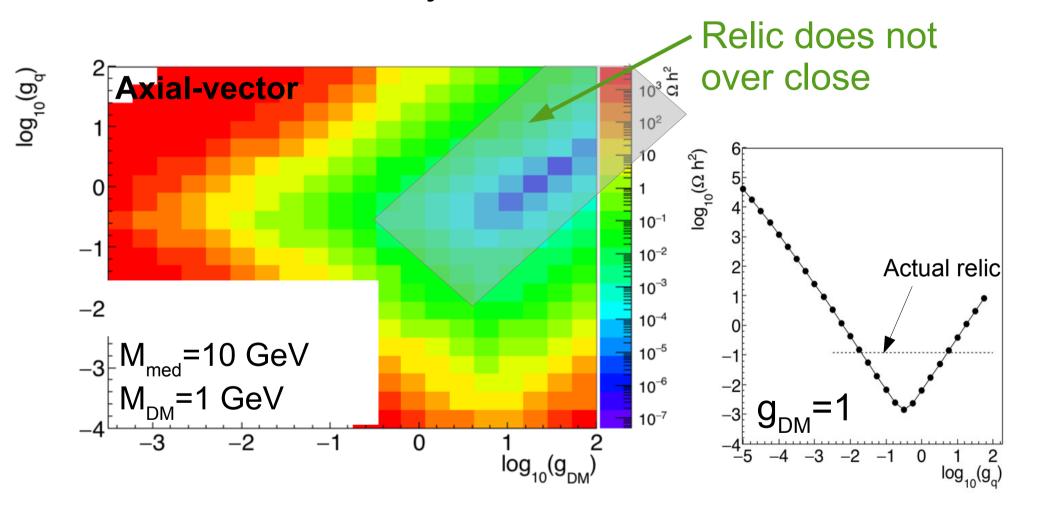


For smaller mediators allowed solution

Arxiv:1703.05703

# Solving for the relic

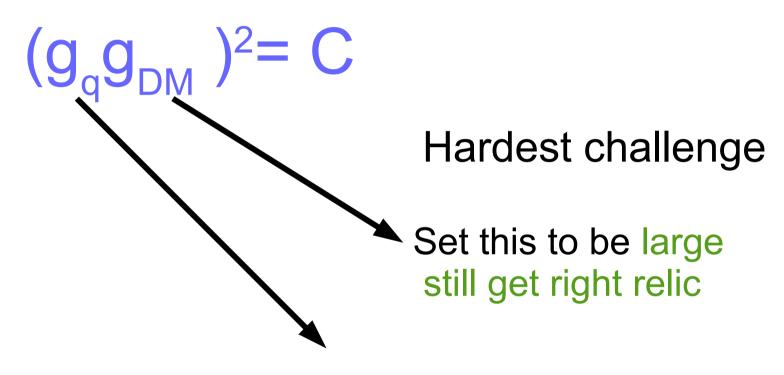
We can numerically solve for the solution



Arxiv:1703.05703

### A common theme of DM talks

Relic density is solved for a constant value of:

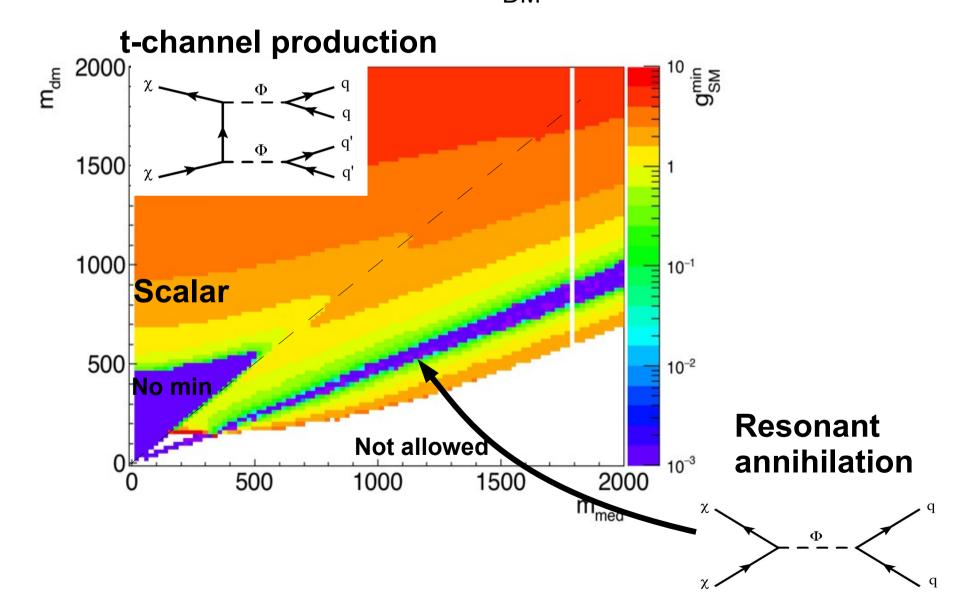


Set this to be small Weak coupling with the SM

Most challenging dark matter searches consist of : strong dark sector coupled weakly to the visible sector

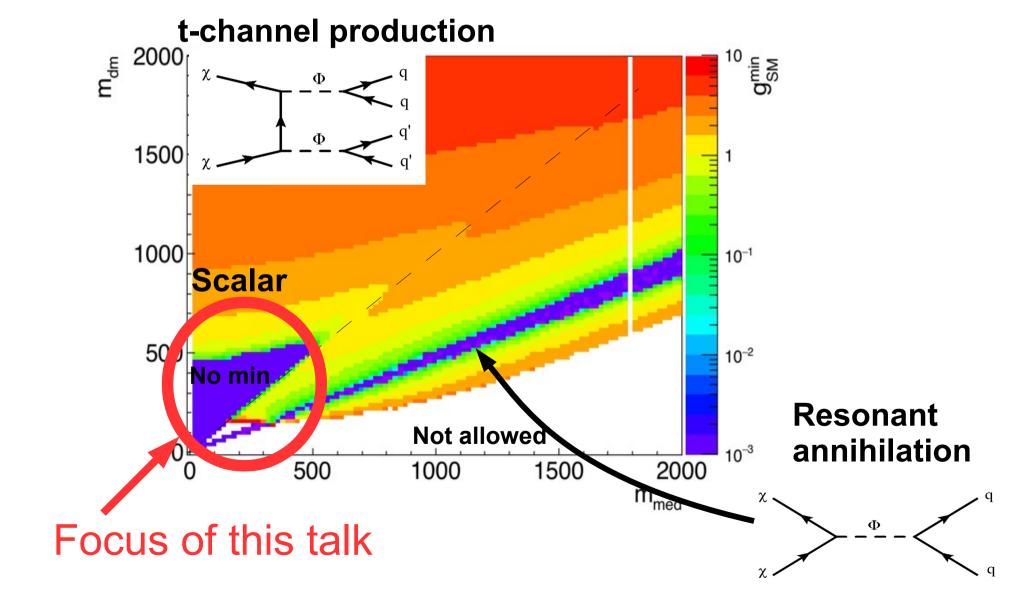
# What is the smallest coupling?

For a dark sector coupling g<sub>DM</sub>=1

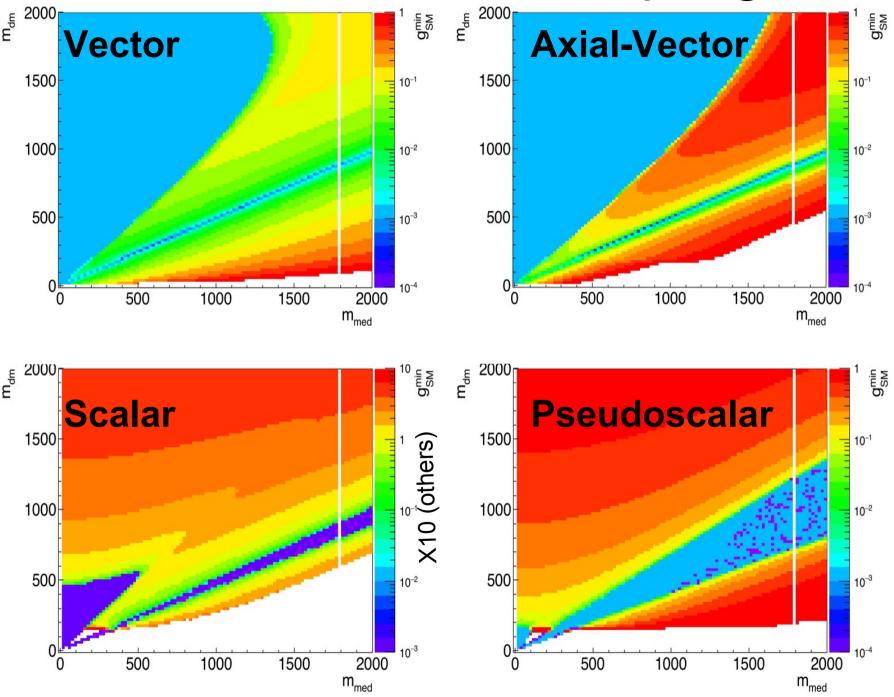


# What is the smallest coupling?

For a dark sector coupling g<sub>DM</sub>=1



# Min Couplings for all



# Summary Benchmarks

### • Spin 1:

Aim to probe couplings down 0.01 for m<sub>Med</sub> > 100 GeV

### • Spin 0 :

- Aim to probe couplings down 0.1 for  $m_{Med} > 300 \text{ GeV}$
- Try to cover m<sub>Med</sub> < 300 by any means possible</li>

### In previous talks

- Have shown for SM-like couplings @ 100 TeV can probe most/if not all phase space
- A few tough to reach places exist
  - This talk will focus on how far we can push these

# Using the Luminosity

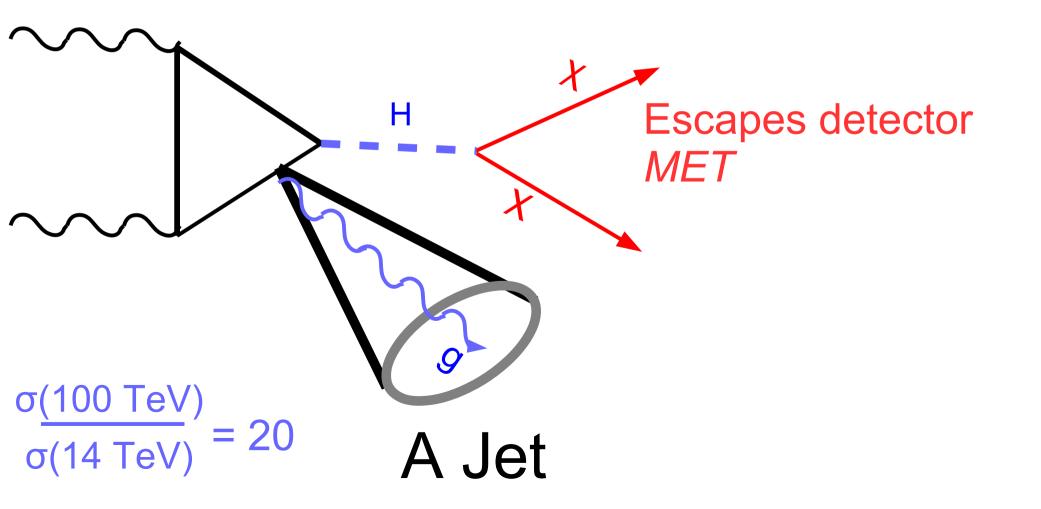
# FCC-hh as Higgs Production tool

- Rate of Higgs production at 100 TeV is very large
  - 800 Higgs events per pb

- Focus of this talk :
  - Whats our sensitivity to H→Inv?
- H→Inv probes a large variety of models
  - Benchmark for exotic Higgs sensitivity
  - Benchmark for low mass scalars
- Fundamental question:
  - What are the advantages of such high rates

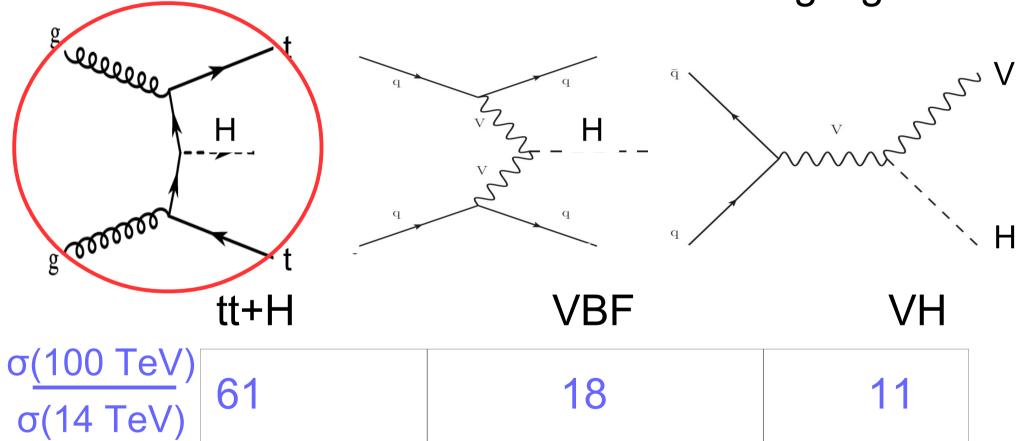
# The Basic Monojet Search

Escaping detector gives us signatures of *MET* 



### **Additional Probes**

Higgs production has additional interesting signatures



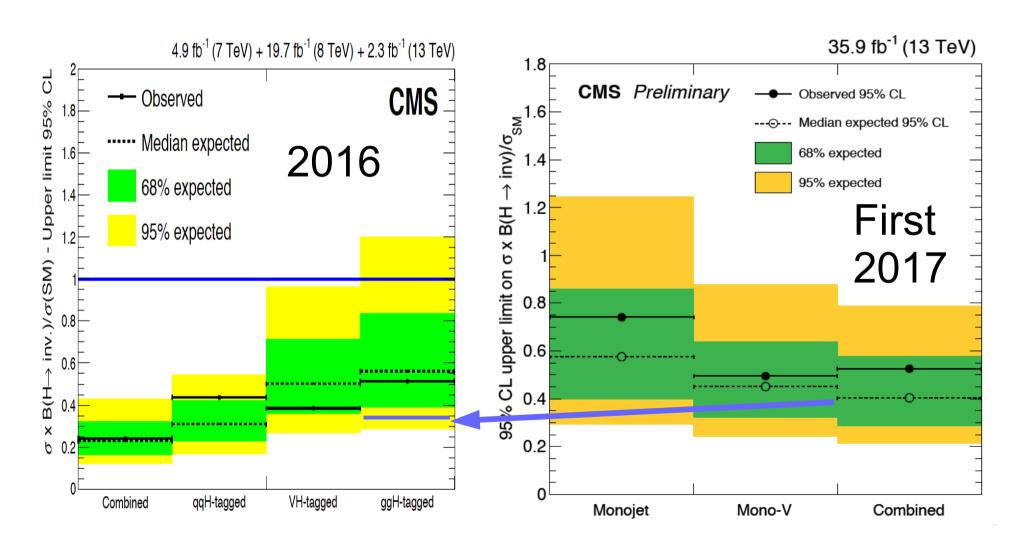
tt+H has a very distinct initial state

Large cross section increase makes:

tt+H→Invisible the golden invisible channel

# Current Higgs Invisible Search

This model is the same as Higgs invisible search

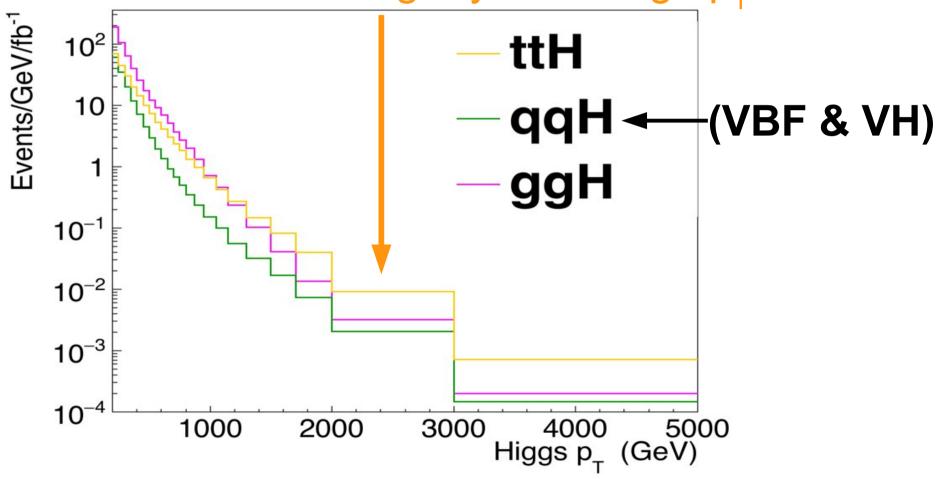


BR(H→Inv) < 24% (CMS) 25% (ATLAS)

### **Additional Observation**

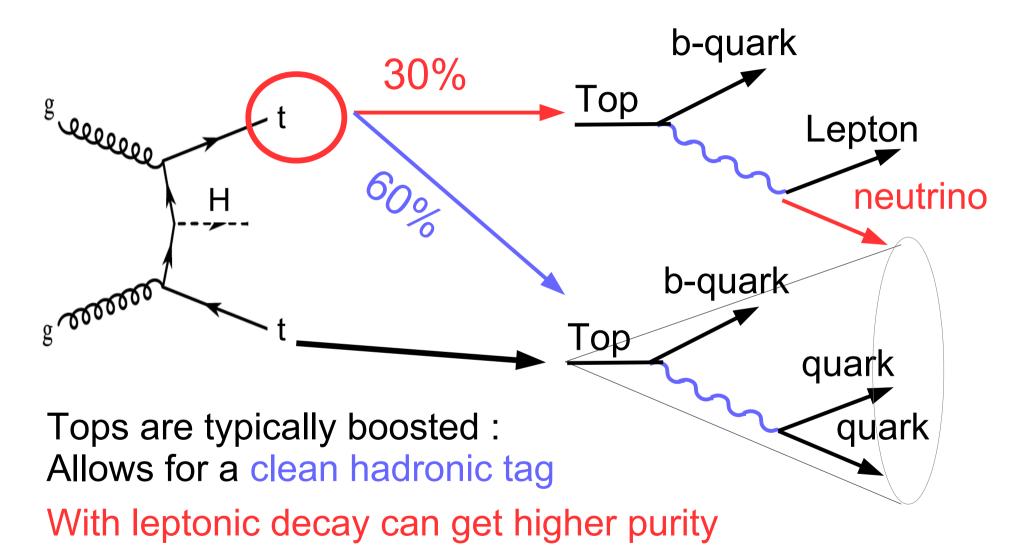
One key feature MLM pointed out

tt+H has a larger yield at high p<sub>⊤</sub>



p<sub>T</sub> spectra of Z & W bosons drops rapidly at high p<sub>T</sub> Inclusive ttH can be made relatively pure

# Designing the tt+H Analysis

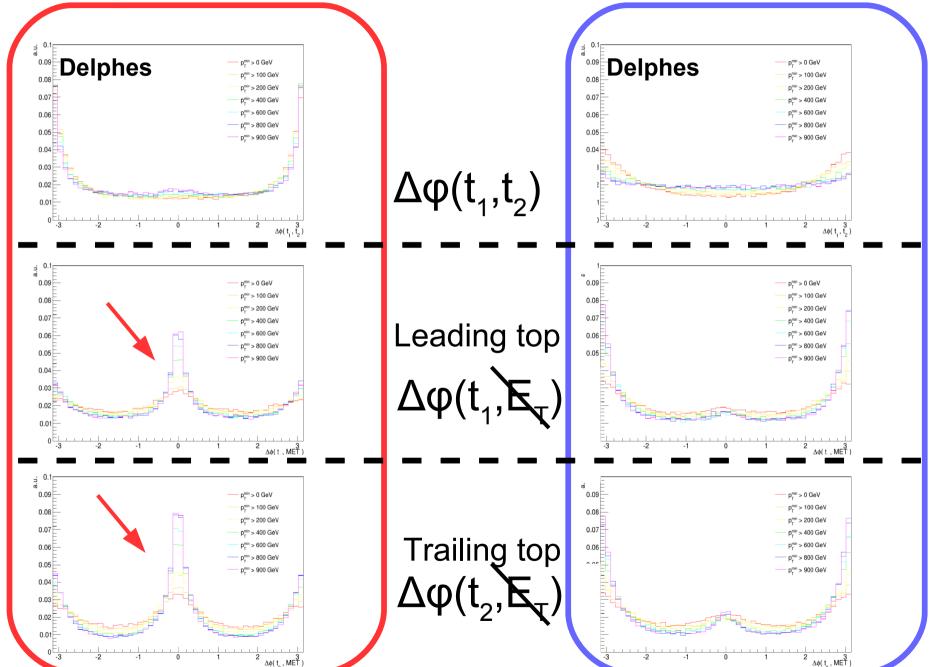


### Here:

consider lepton w/another hadronic top jet

### Background: SM tt

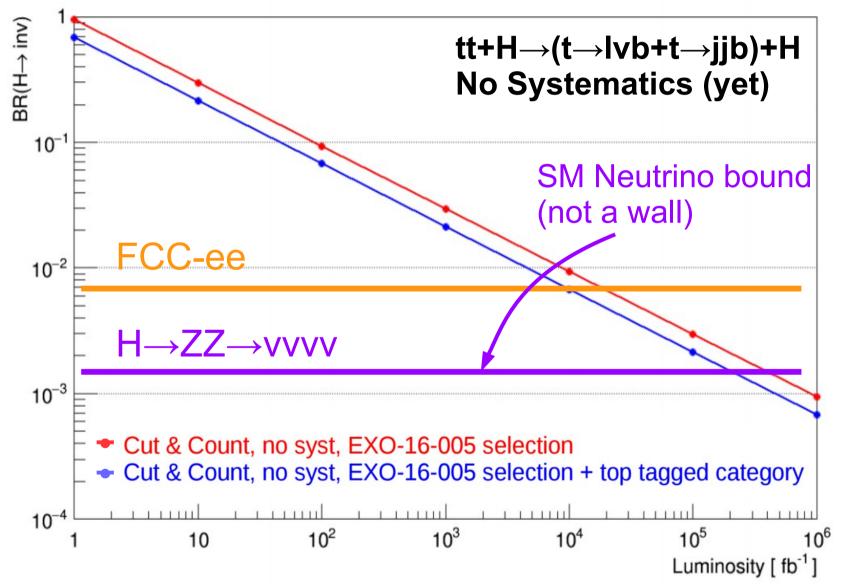
### Signal : tt+H→Inv



Boosted tt+H analysis has an S/B of 1 to 1 (very pure

# Implications with a Pure category

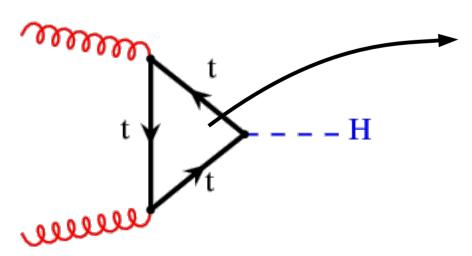
Currently considering semi-leptonic channel without systematics



Crosses both FCC bounds and SM H Invisible bound

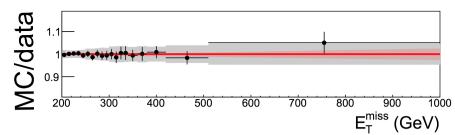
# Monojet(s) analysis

- Consider an analysis:
  - Veto leptons for  $|\eta| < 4.0$
  - Fit the MET spectrum
    - Predict the MET spectrum with the highest level of precision
- In MET tail S/B is 2-5%
  - Aim to just exploit low purity with very large yields

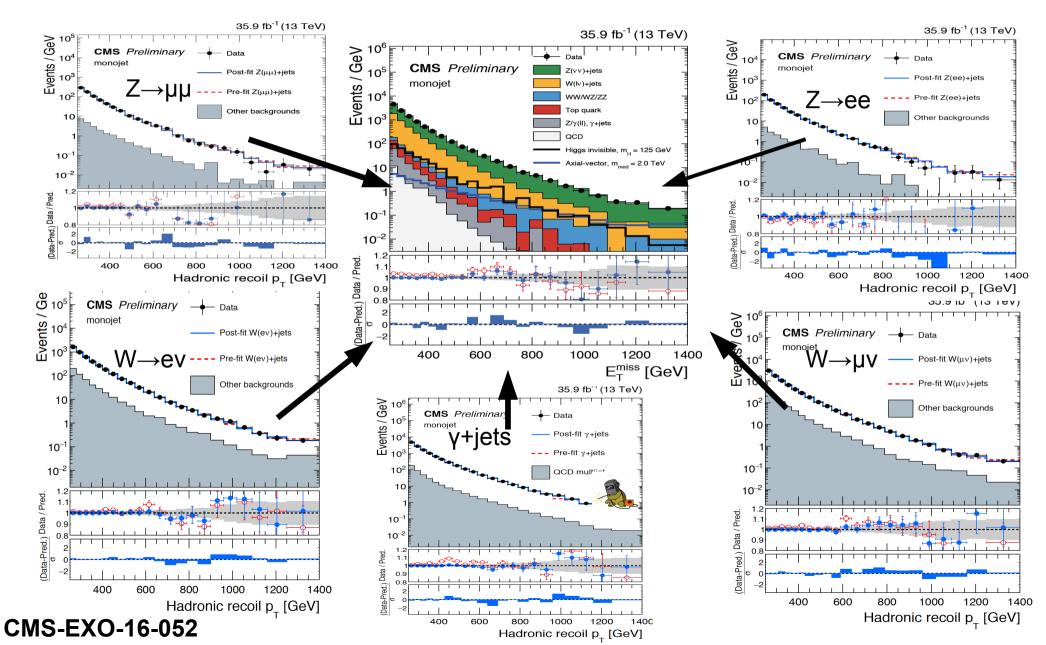


Finite top quark mass contribution crucial

Approximately known to NLO

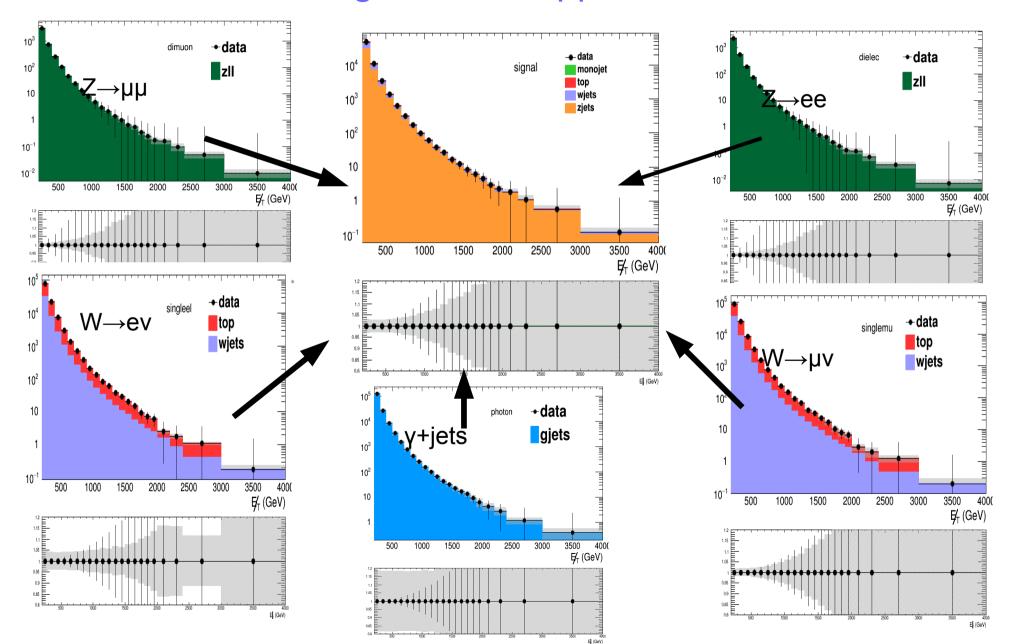


### 5 Control regions 15% uncertainty @ 1 TeV



# Monojet analysis @ CMS

### The same fitting scheme applies to 100 TeV

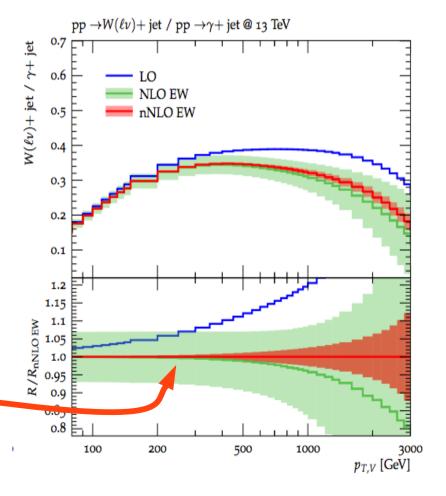


# The foundation of this analysis

Going from  $\gamma$  or  $W \rightarrow Z$ 

Unc. 
$$\frac{d\sigma^{Y(W)}}{dp_T} = \frac{d\sigma^Z}{dp_T}$$

- Key to this analysis ratios
  - Require best theoretical calculations
  - Current (N)NLO theoretical prescription brought additional ~40% on 36/fb analysis



# The foundation of this analysis

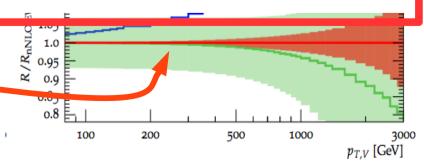
Going from  $\gamma$  or  $V \rightarrow Z$ 

Unc. 
$$\frac{d\sigma^{Y(W)}}{dp_{T}} / \frac{d\sigma^{Z}}{dp_{T}} / \frac{d\sigma^{Z}}{dp_{T}} = \frac{d\sigma^{V(W)}}{dp_{T}} / \frac{d\sigma^{V(W)}}{dp_{T}} = \frac{d$$

Precise predictions for V+jets dark matter backgrounds

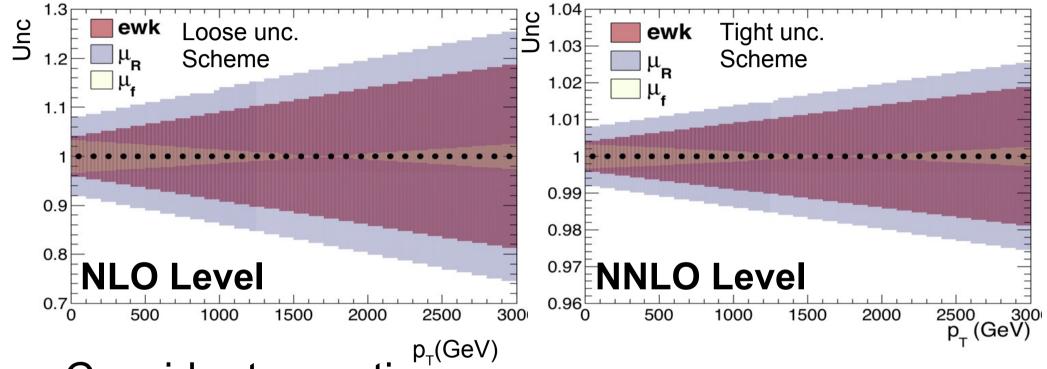
J. M. Lindert<sup>1</sup>, S. Pozzorini<sup>2</sup>, R. Boughezal<sup>3</sup>, J. M. Campbell<sup>4</sup>, A. Denner<sup>5</sup>, S. Dittmaier<sup>6</sup>, A. Gehrmann-De Ridder<sup>2,7</sup>, T. Gehrmann<sup>2</sup>, N. Glover<sup>1</sup>, A. Huss<sup>7</sup>, S. Kallweit<sup>8</sup>, P. Maierhöfer<sup>6</sup>, M. L. Mangano<sup>8</sup>, T.A. Morgan<sup>1</sup>, A. Mück<sup>9</sup>, F. Petriello<sup>3,10</sup>, G. P. Salam\*<sup>8</sup>, M. Schönherr<sup>2</sup>, and C. Williams<sup>11</sup>

prescription brought additional ~40% on 36/fb analysis



# Benchmarks for this study

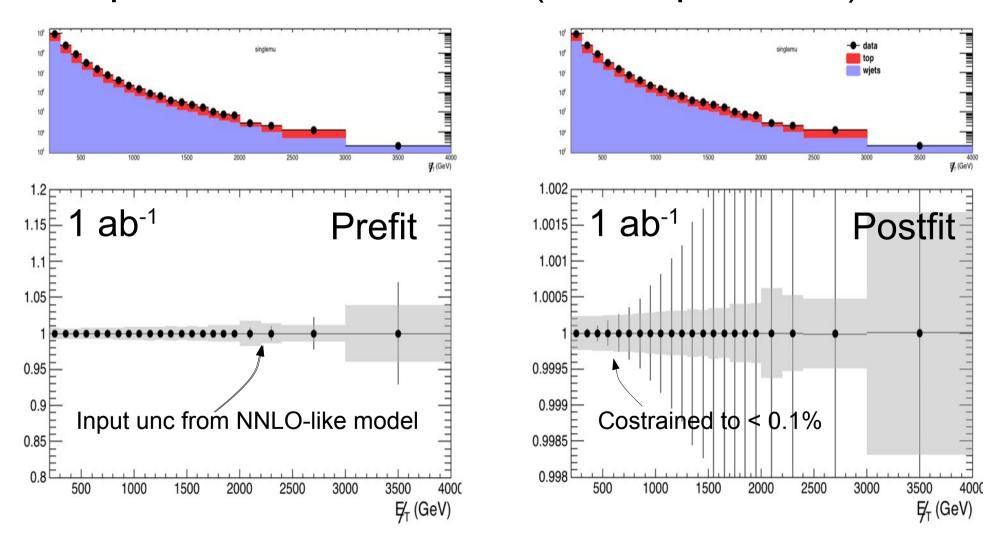
What are reasonable uncertainty choices



- Consider two options :
  - A Loose uncertainty →Comparable to NLO
  - A Tight uncertainty →Comparable to NLO
- Using: 0.5%/0.25%/5% e/μ/τ efficiency & 1% lumi

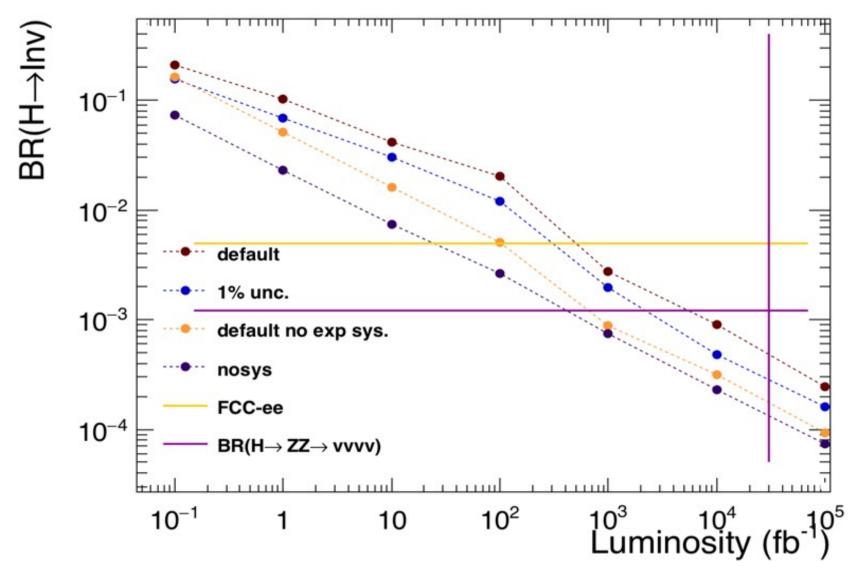
# What is the precision?

Can probe a few % effects (NNLO precision)



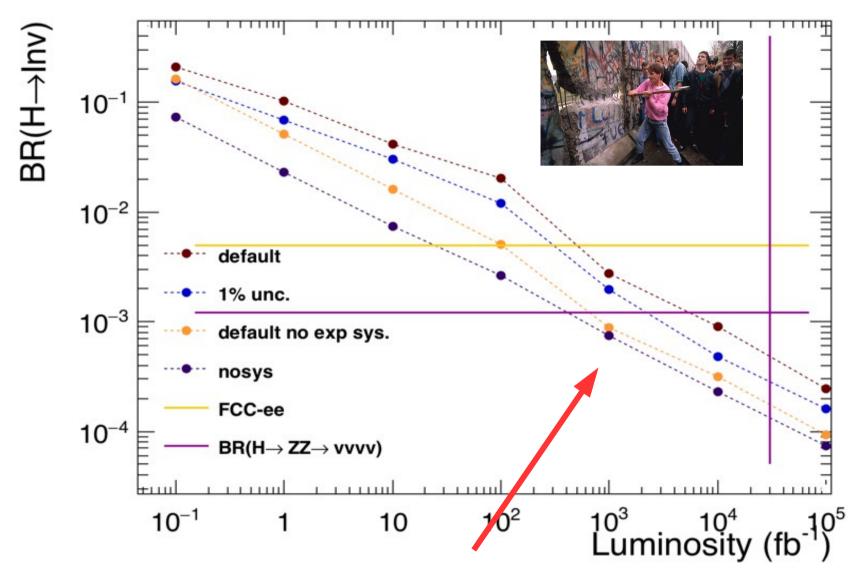
Through this scheme we can probe boson pT to 10<sup>-4</sup> level

## How do things scale?



Cross the SM neutrino wall at FCC with < 1 ab<sup>-1</sup>

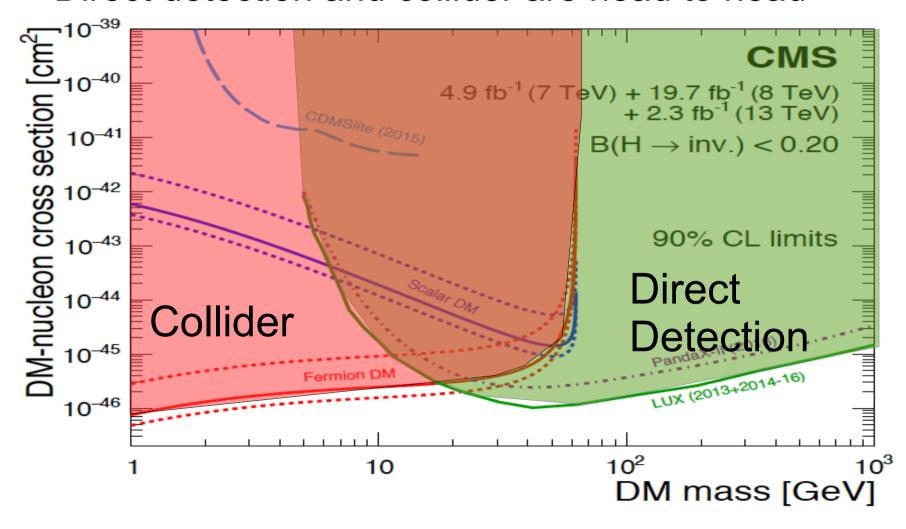
## How do things scale?



There is no systematics wall

### **Current Bounds**

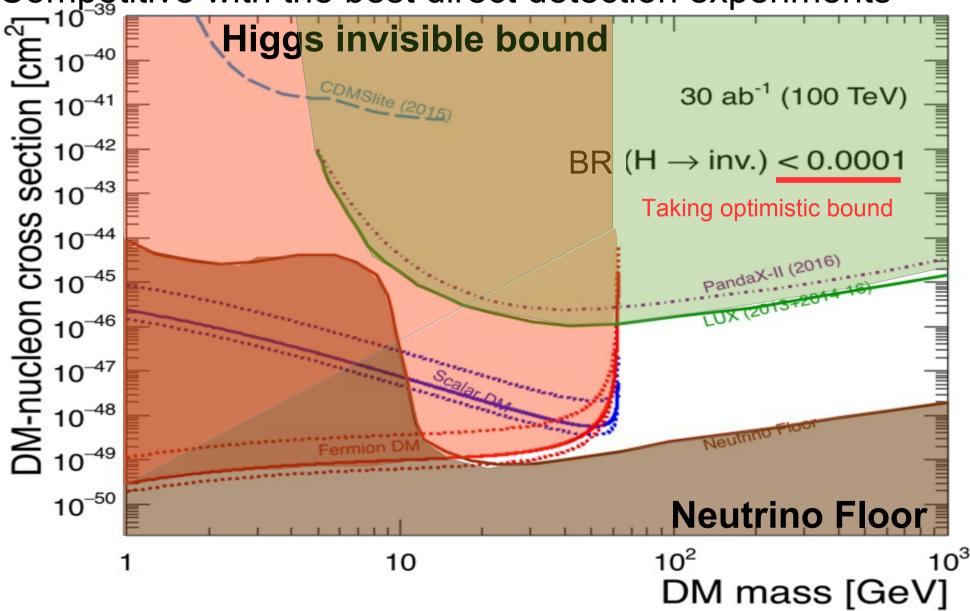
- Higgs to invisible :
  - Direct detection and collider are head to head



Competitive with the best direct detection experiments

#### **Future Bounds**

Competitive with the best direct detection experiments

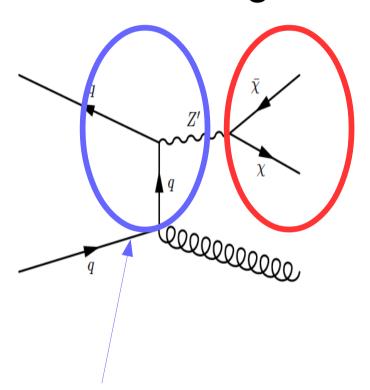


Higgs invisible of 10<sup>-4</sup> corresponds to g<sub>sm</sub> from 10<sup>-3</sup> to 10<sup>-2</sup>

# Beyond Invisible Searches

#### What else?

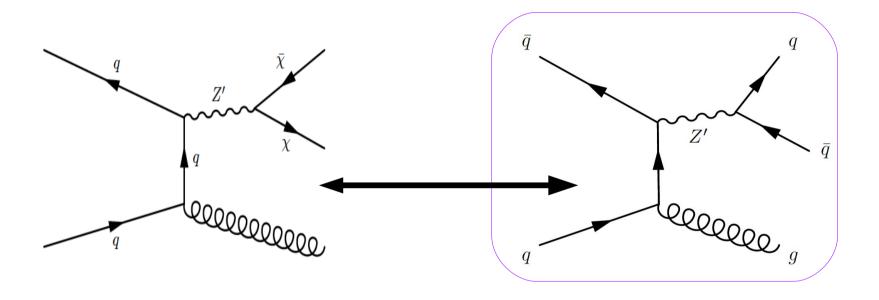
Without loss of generality we also have dijets



Mediator is coupling to quarks and to Dark matter

#### What else?

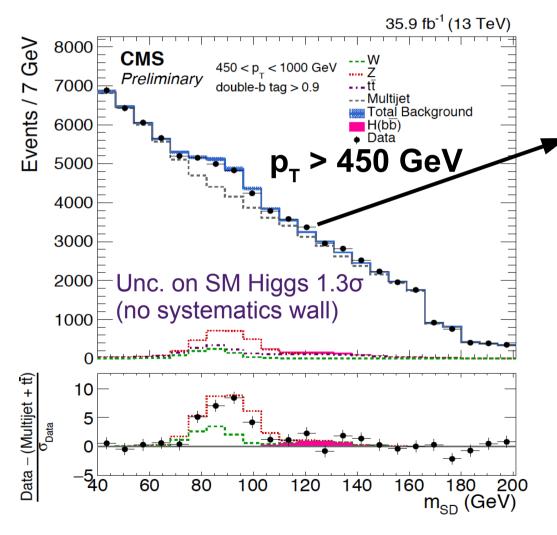
Without loss of generality we also have dijets



This is a dijet+ISR search Mediator is coupling to quarks and to Dark matter Mediator can decay to quarks

# Just beginning to search directly

Technique for tagging high p<sub>⊤</sub> objects in its infancy



First result for high p<sub>T</sub>
"Visible Higgs"
Came out this week

Can exploit same ideas From invisible to probe

Higgs production w/small couplings

To full FCC-hh gives BR(H→bb) sensitivity 10<sup>-3</sup>

# Other Dark Matter

#### Global View of Dark Matter

Key
Benefit @ 100 TeV
Detector Demand

**Dark Matter** 

Low mass mediators

Rate

(theoretical) precision

High mass mediators

Increased cross section Basic capabilities

Compressed Spectra

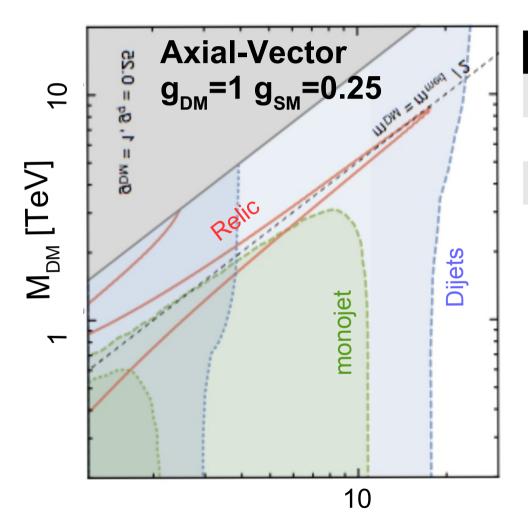
More Boost High res detector (low p<sub>T</sub>/displaced vtx)

unexpected

New phase space Maximal flexibility

#### High Mass Reach

Probing High mass reach

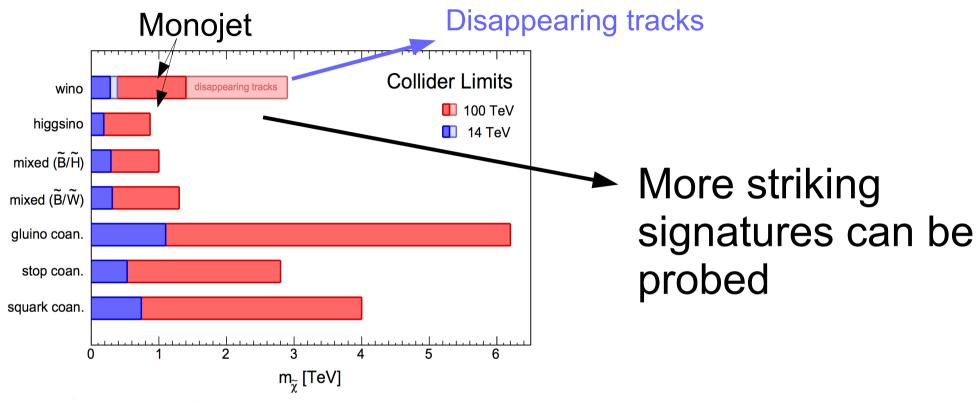


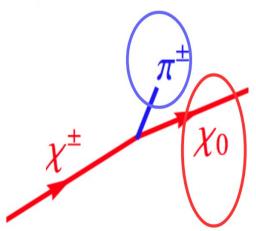
M<sub>med</sub> [TeV]

Model	Collider	(In)Direct	Relic	
Vector	15 TeV	>100 TeV	70 TeV	
Axial	15 TeV	6 TeV	8 TeV	
Scalar	3.5 TeV	3 TeV	6 TeV	
Pseudo 100 T	4 TeV V eV bound	1 TeV?	40 TeV	
	Bounds from other methods			
Relic density				

w/high mass mediators searches
Cover large most(if not all) allowed space)

#### More striking signatures





With compressed scenarios can have:

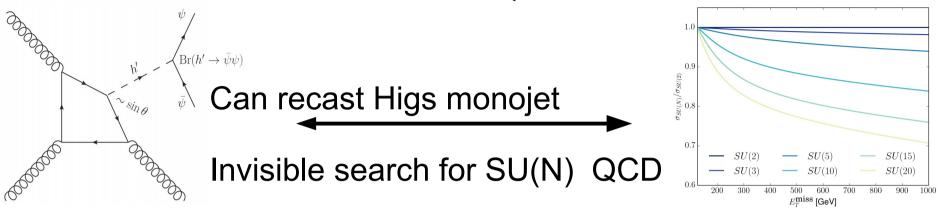
**Disappearing Tracks** 

Low  $p_{\tau}$  leptons (< 10 GeV)

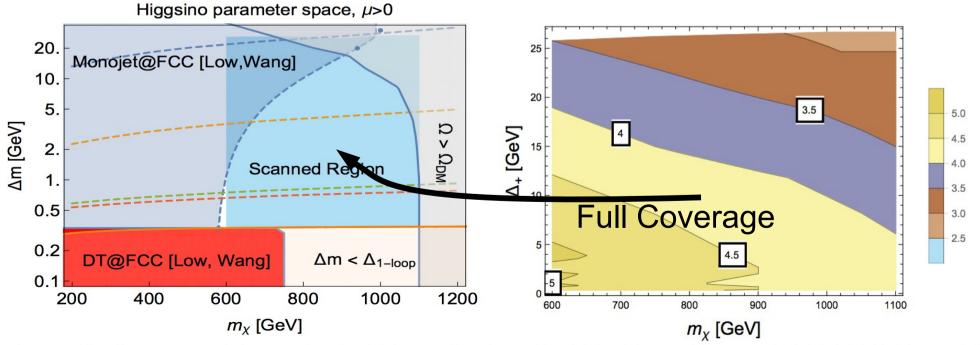
Displaced vertices

#### More exotic searches

Can consider extended QCD sectors



Exending compressed spectra to mono-Z



https://indico.cern.ch/event/550509/contributions/2413850/attachments/1399186/2134448/Zurita\_F

# Spectrum of Dark matter

 Deep understanding of monojet extends to many models

 Additional searches extend to

- Disappearing track
- Mono-Z
- Displaced jets
- ...More exotic

Final State	Analysis	section		
jet+MET	Wino, Higgsino DM	4.3.1 - 4.3.4		
jet+MET	Higgs Portal	4.3.5		
jet+MET	Simplified Vector/Axial	4.4.1 - 4.4.3		
jet+MET	Simplified Scalar/Pseudo	4.4.1 - 4.4.3		
jet+MET	Gluion/stop coannihalation	4.5.1		
VBF jets +MET	Wino, Higgsino DM	4.3.1 - 4.3.2		
VBF jets +MET	Higgs Portal	4.3.5		
photon+MET	Wino	4.3.2		
Disappearing tracks	Wino, Higgsino	4.3.1 - 4.3.2		
Disappearing tracks	Fiveplet DM	4.3.3		
Disappearing tracks	Relic-Neutralino	4.3.4		
lepton+ $\gamma$ +MET	Relic-Neutralino	4.3.4		
$Z_D  o ll + (Z_D  o ll)$	Dark Photons	4.4.4, 4.6.3		
displaced jets	Dark QCD/Hidden Valley	4.6.2		
long lived charged particle	Super-WIMPS/Gravitino	4.6.4		
dijet	Simplified Vector/Axial	4.4.1 - 4.4.3		
0 ' 64 6 1 44 14 '4 1 11 '4 11 4 4				

**Table 5:** Overview of the final states and the associated model, with a link to the respective section.

#### Conclusions

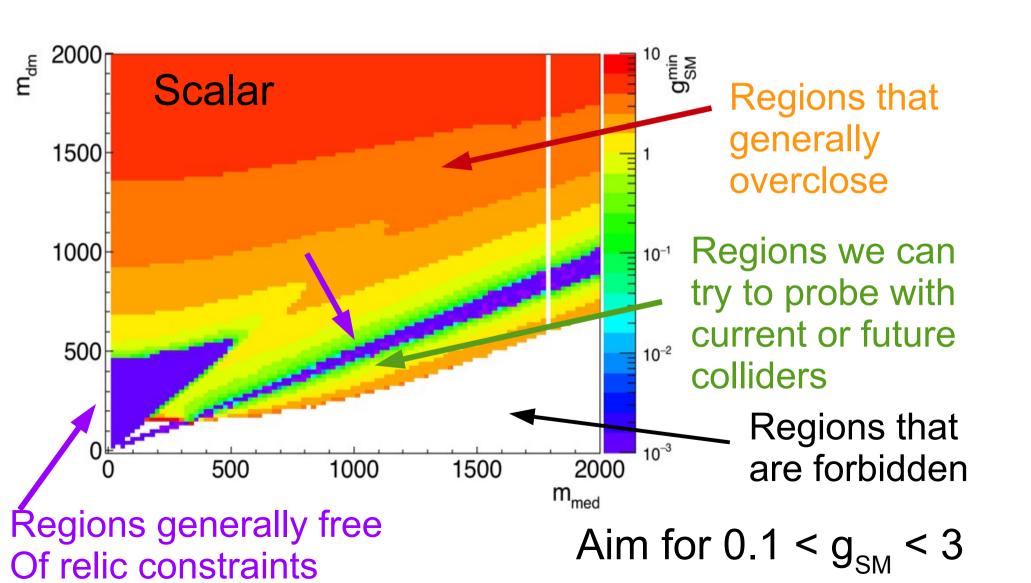
- A key aspect to FCC-hh is incredible rate
  - Allows us to probe Higgs invisible beyond neutrino wall
  - Extends Higgs invisible search well beyond FCC-ee
  - Extends to SM Higgs invisible
    - Gives us a signal we can calibrate
  - Higgs invisible bound translated to low mass scalar
    - Probes most of the allowed minimal coupling phase space

- Dark matter at FCC-hh
  - Four part study in High rate/High Mass/Exotics
  - In all cases: capability to exceed or match all other exp.

# Thanks!

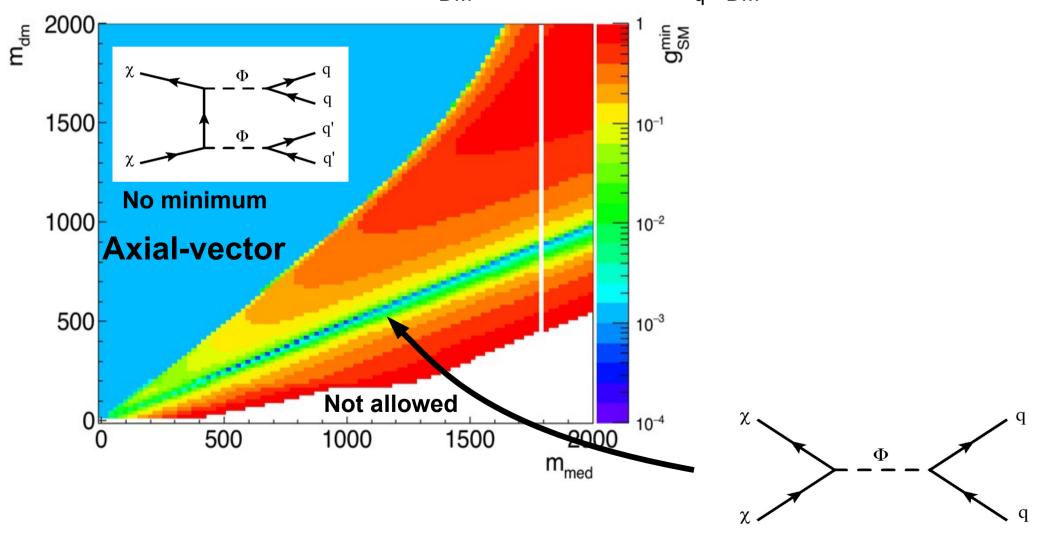
#### What do we conclude?

What is driving the results is the coupling



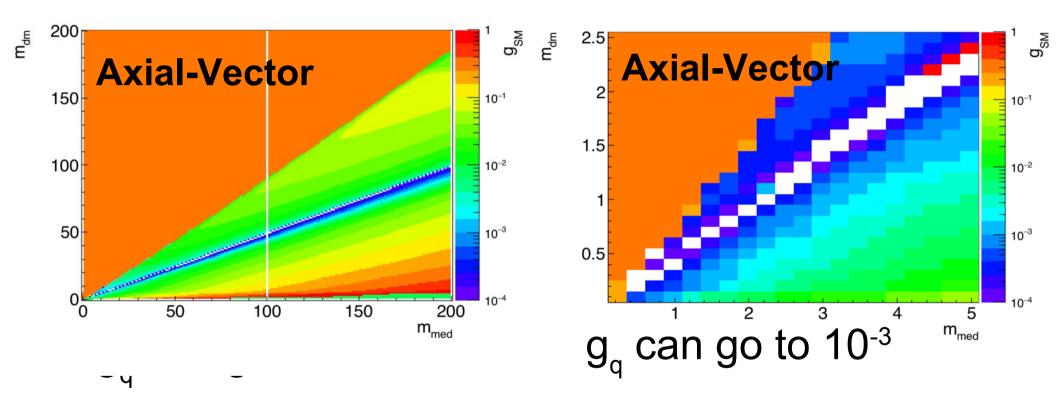
#### Min couplings

- Can split the solution to the max and min coupling
  - In this case we fix  $g_{DM} = 1$  (product  $g_{QDM}$  defines bound)



#### What about at low mass?

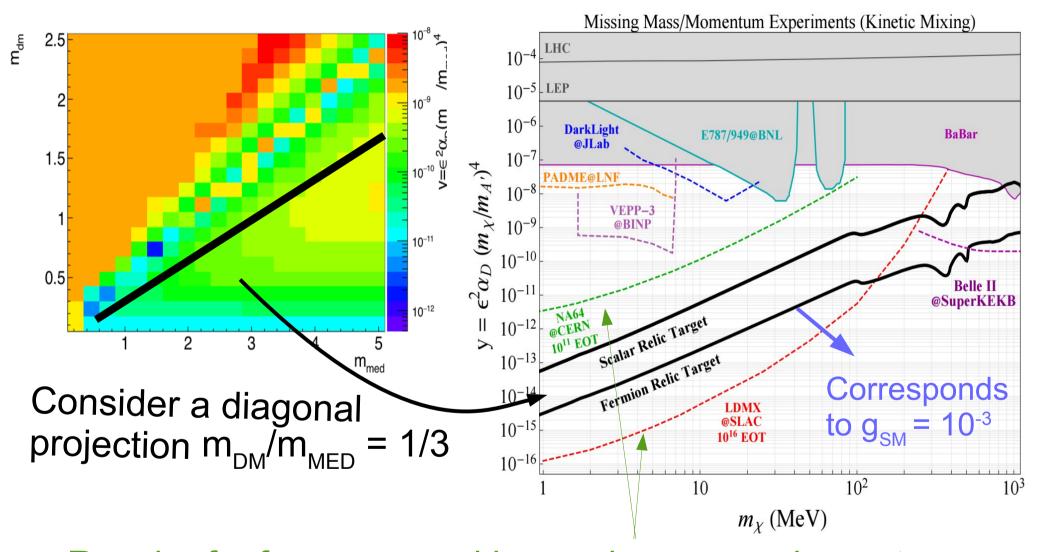
Coupling becomes a real challenge



At low masses we can have very small couplings

However we have more strategies

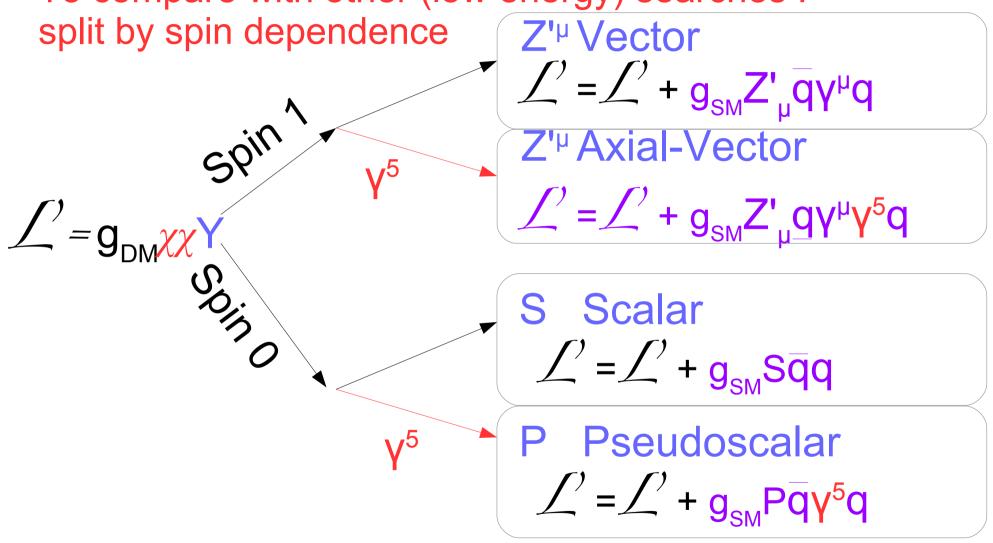
# @low mass aim is for low couplings



Reach of a few proposed beam dump experiements Can probe the interesting region

# Preserving Generality?

To compare with other (low energy) searches:



Strategy of searches in LHC does not change much

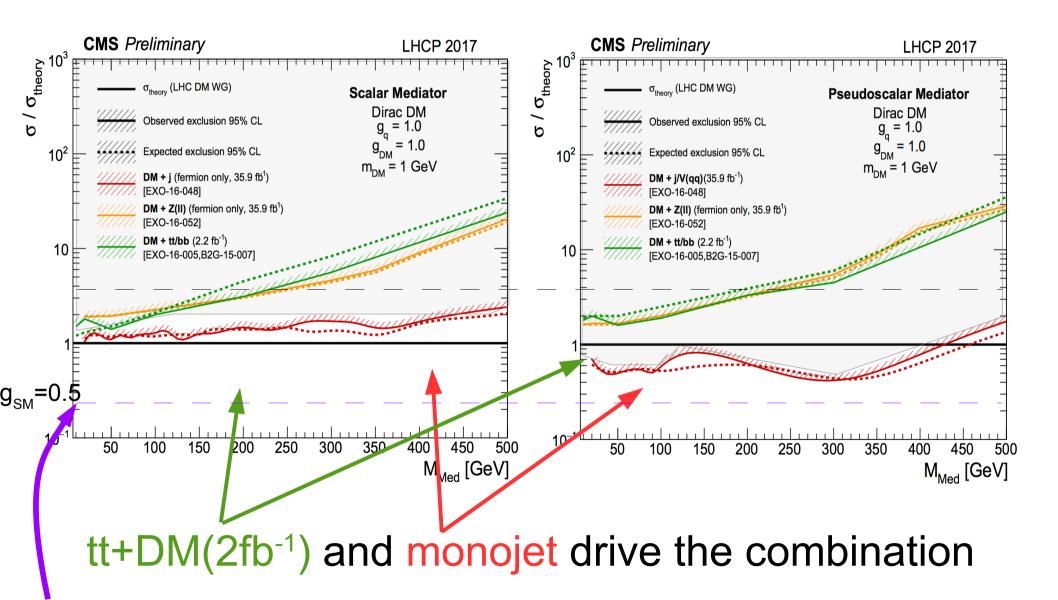
Interpretation agains Direct Detection/Indirect Changes a lot



# Monojet search Straddling SM and BSM

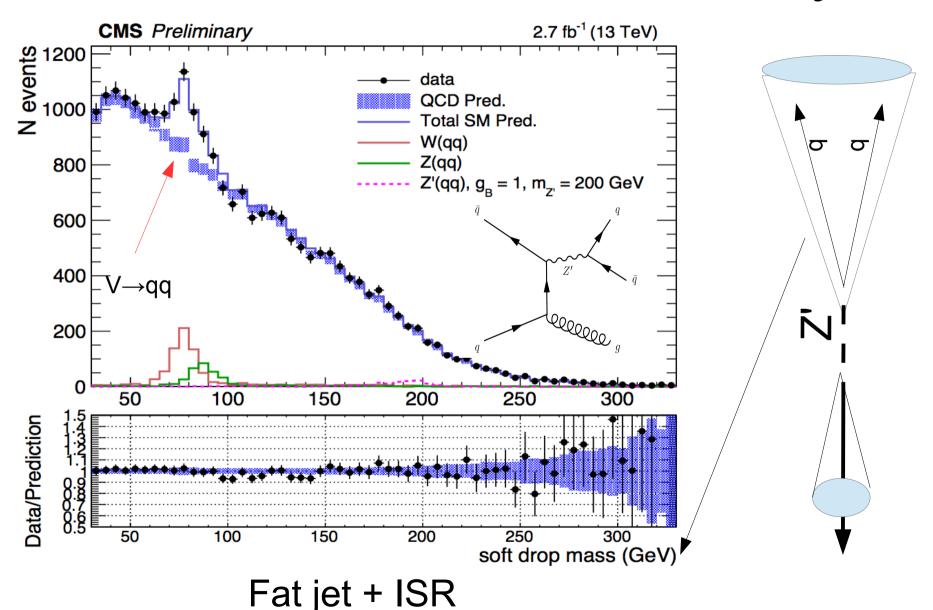
#### No EWSB

# Comparing all channels

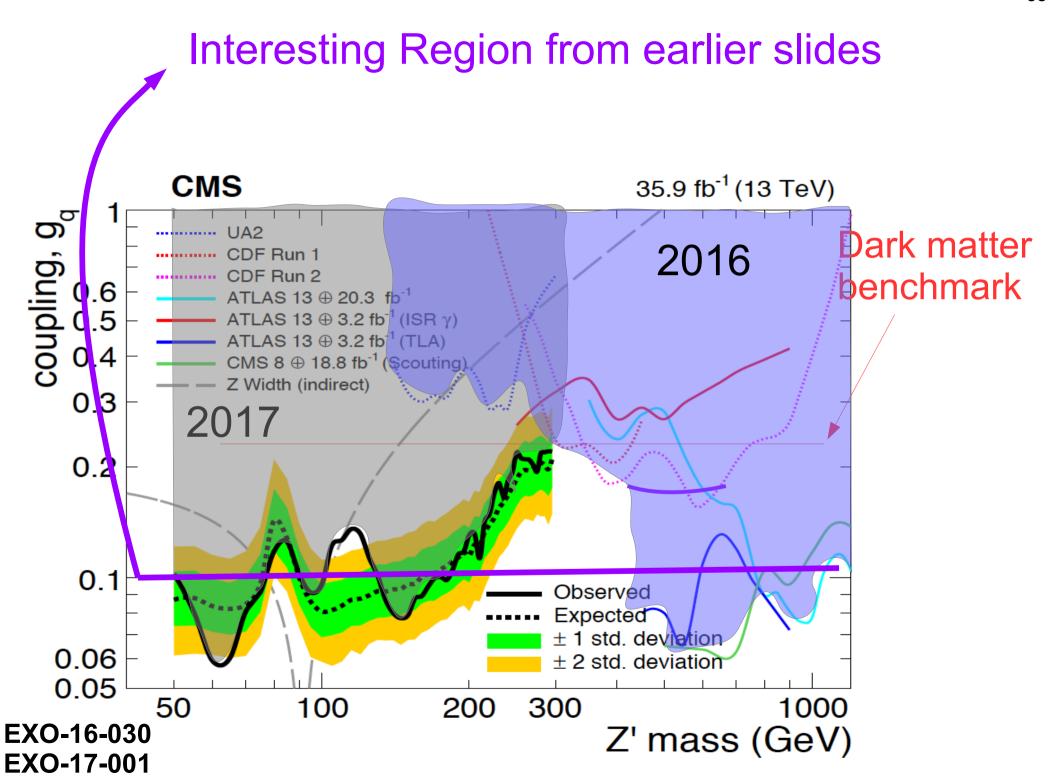


Not far from an intermediate benchmark of  $g_{SM} = 0.5$ 

#### Going all the way down



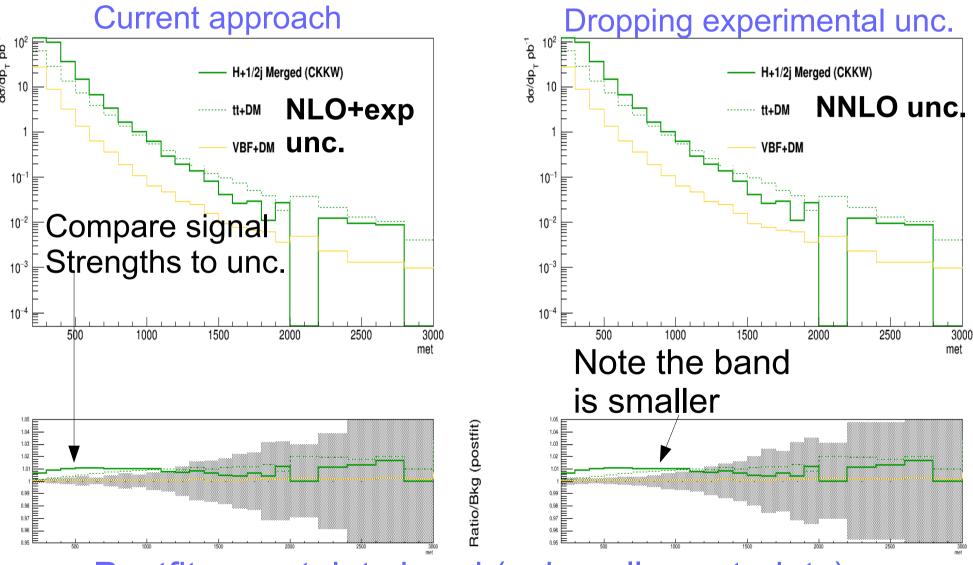
For this plot we invented a new substructure var



#### Where are we most sensitive?

10 fb<sup>-1</sup>: Signal sensitivity to unceratinty

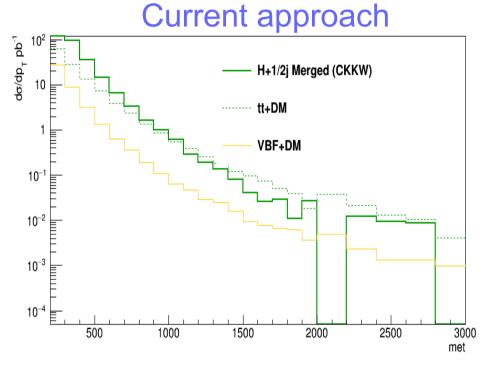
Ratio/Bkg (postfit)

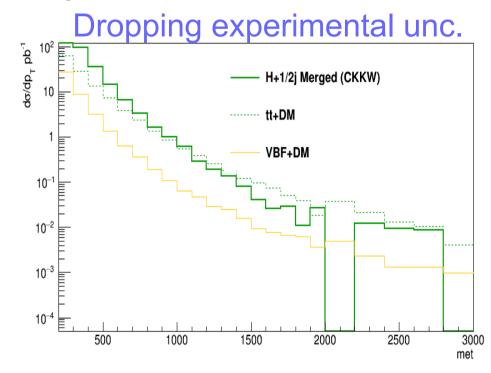


Postfit uncertainty band (using all constraints)

#### Where are we most sensitive?

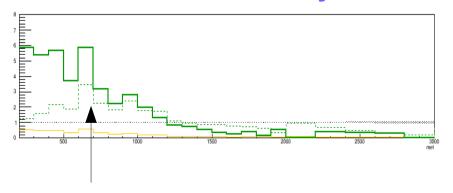
10 fb<sup>-1</sup>: Changing ratio to Bin/postfit unc. σ

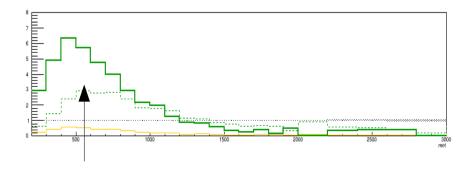




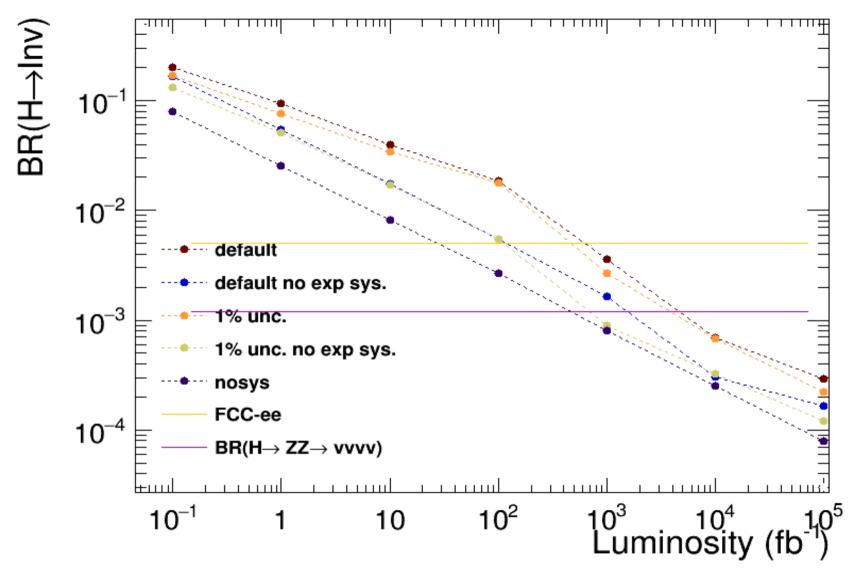
#### In both cases monojet dominates tt+H signal for sensitivity

Ratio/ggH





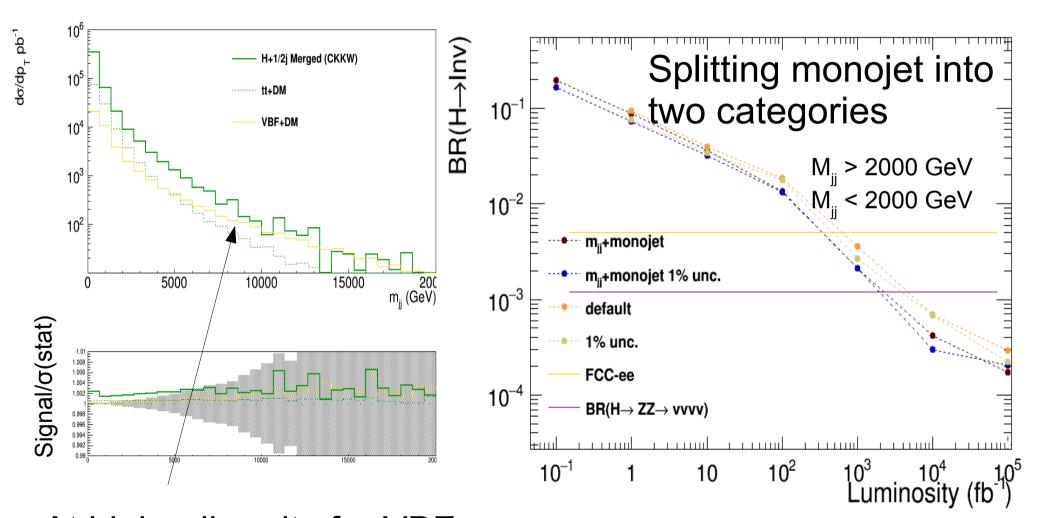
#### How do things scale?



Cross the SM neutrino wall at FCC with < 1 ab<sup>-1</sup>

# Can we extend things?

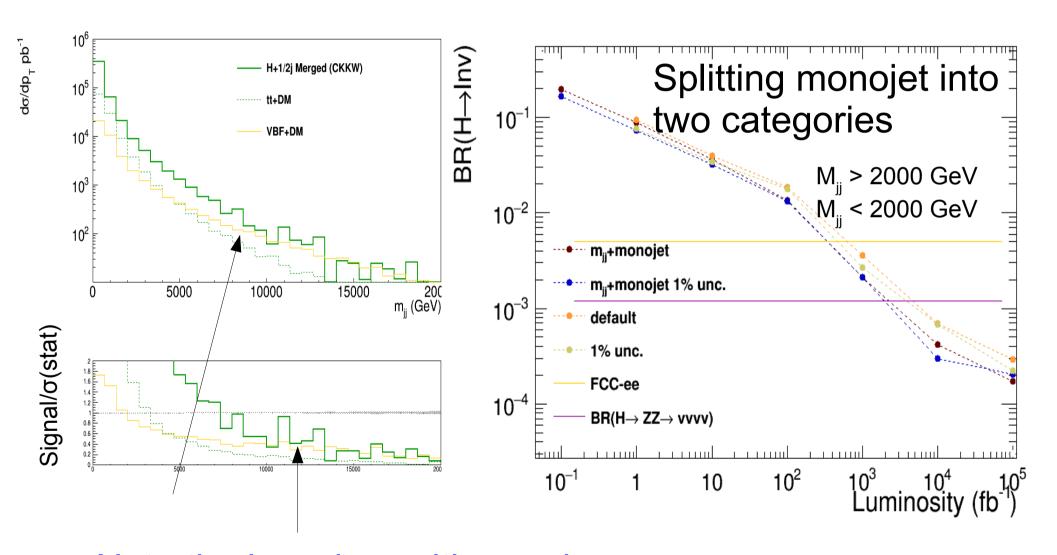
Can consider targetting the VBF final state?



At high mjj purity for VBF can become quite high

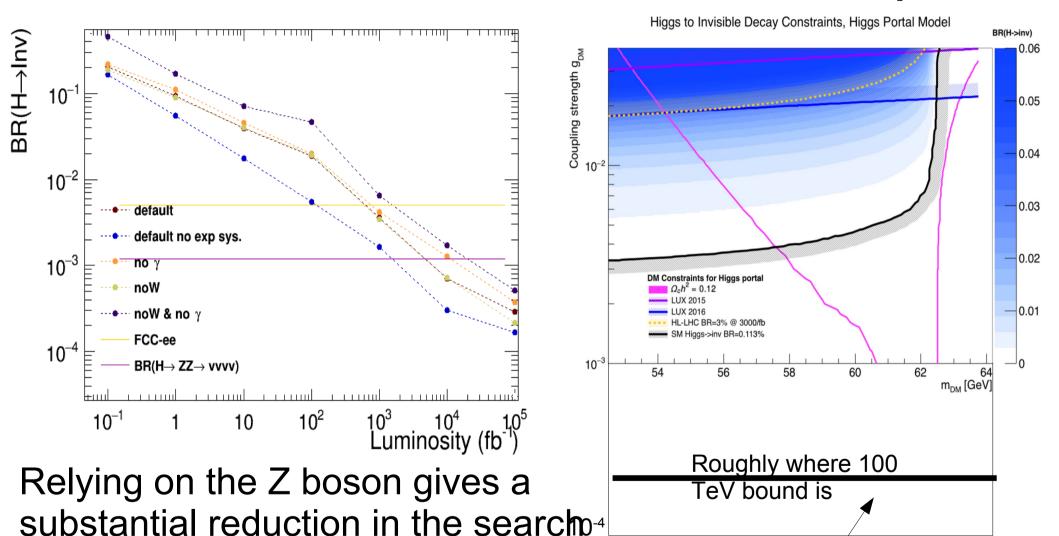
# Can we extend things?

Can consider targetting the VBF final state?



Note the broad sensitive region

#### What is the impact?



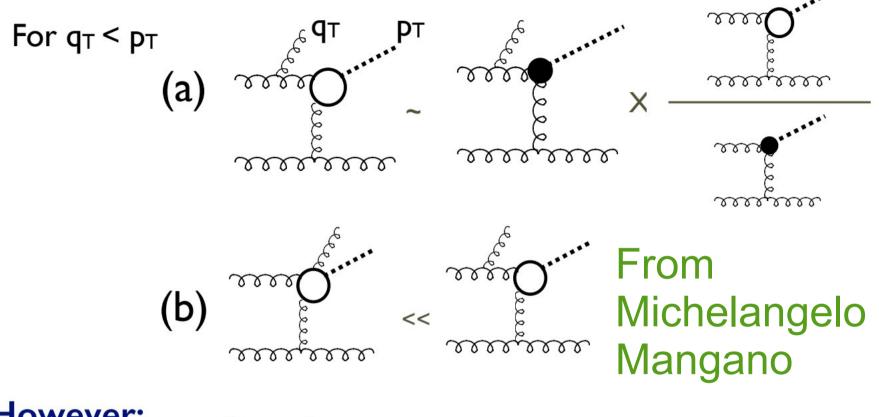
Equivalent mass splitting to be < 1 GeV (given relic)

#### Conclusion

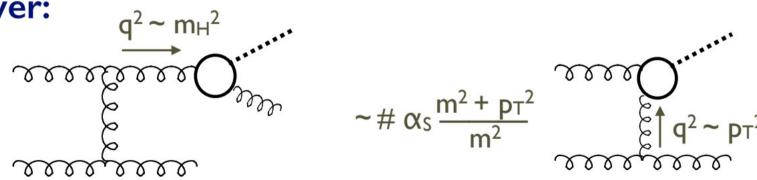
- Currently investigating H→Invisible
  - Monojet and tt+H are the dominant productions
  - Modern approach allows for scaling of limits
    - Result scales with luminosity
    - Systematic choice is critical for search
- Improving the search:
  - Better understanding of the Higgs p<sub>⊤</sub> needed

- For Higgs Invisible we find that :
  - We can reach the neutrino wall SM H→Invisible
    - Best BR(H→Invisible) < 1-2x10<sup>-4</sup>

#### Justification for this approximation:



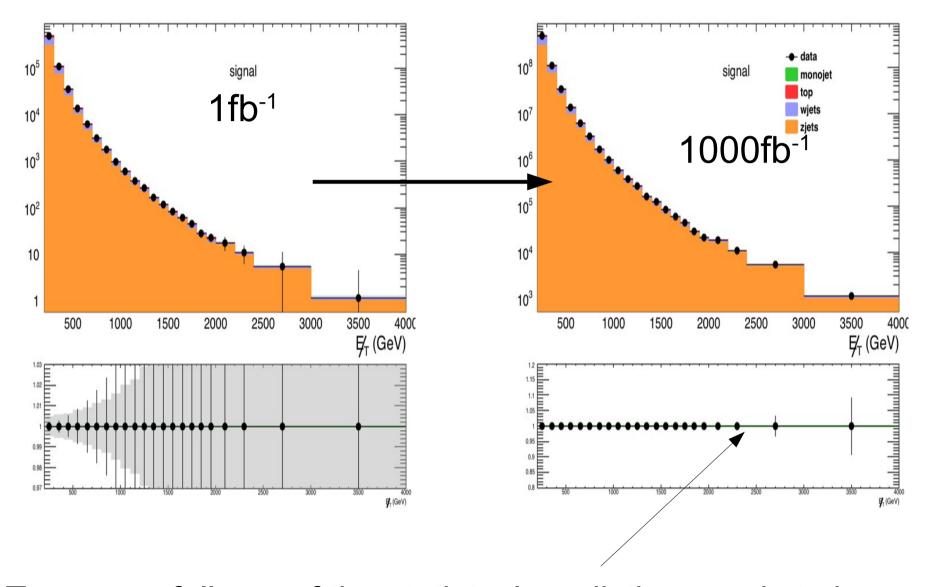
**However:** 



These diagrams will eventually take over at very large  $p_T$ .

They are not covered by the "merging" approach .... this should be looked at in some more detail ....

#### Whats the precision?



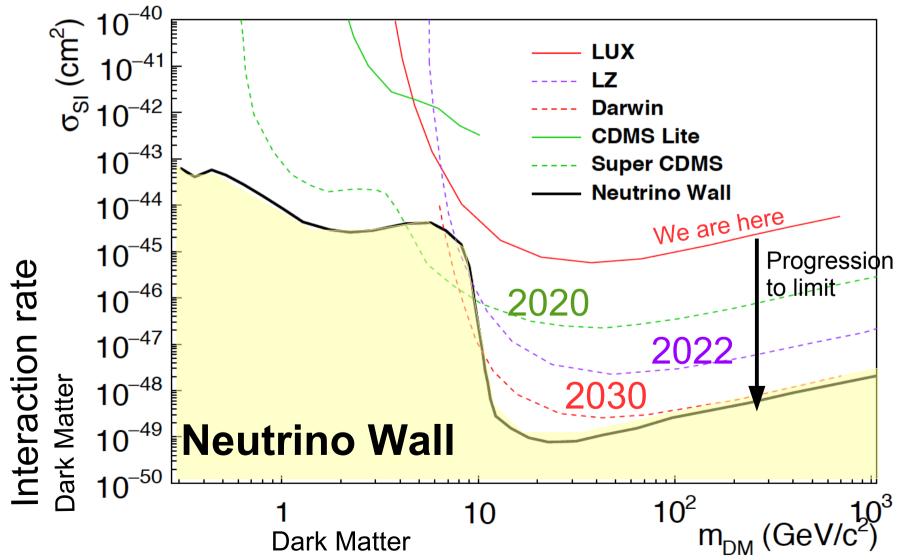
To ensure full use of the statistical prediction needs to be a at few % level → This works now with fully correlated shape

# Targeting Dark matter

- Currently there are 3 industries looking for DM
  - Direct detection
  - Indirect detection
  - Collider searches
- For each of these approaches:
  - Benchmarks have been established to drive search
    - For collider this is not as well formed
- For collider searches:
  - New benchmark to be established based precision SM
    - Turns out DM search is best way to measure high p<sub>T</sub> V prod
  - This talk looks at this benchmark for the 100 TeV

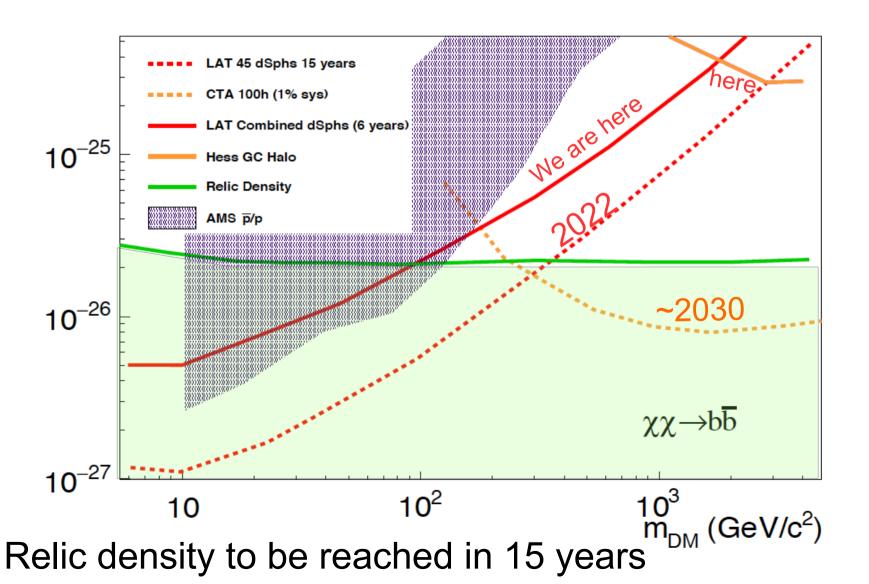
#### What are the ulitmate bounds?

- Ultimate bounds exist for each experiment
  - Direct detection this ultimate bound is the neurino wall



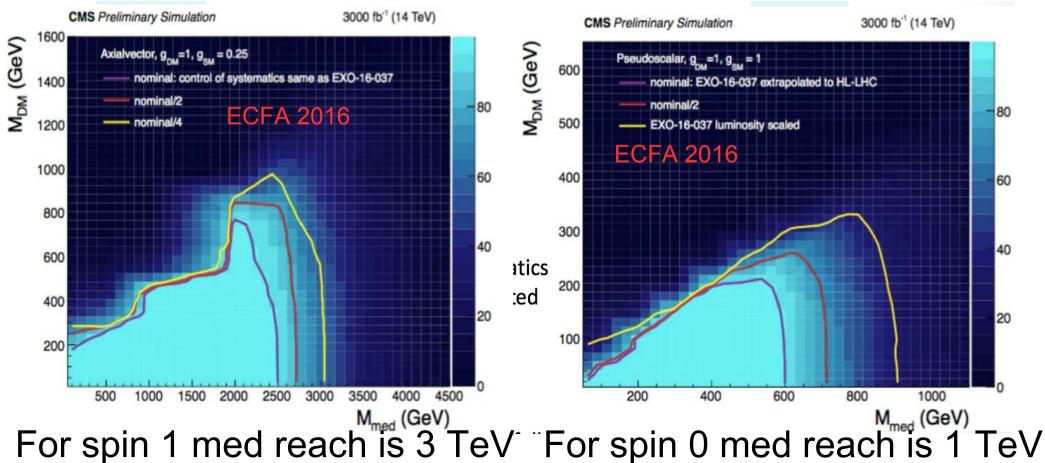
#### Whatr are ultimate bounds?

Indirect detection ultimate bound is relic density

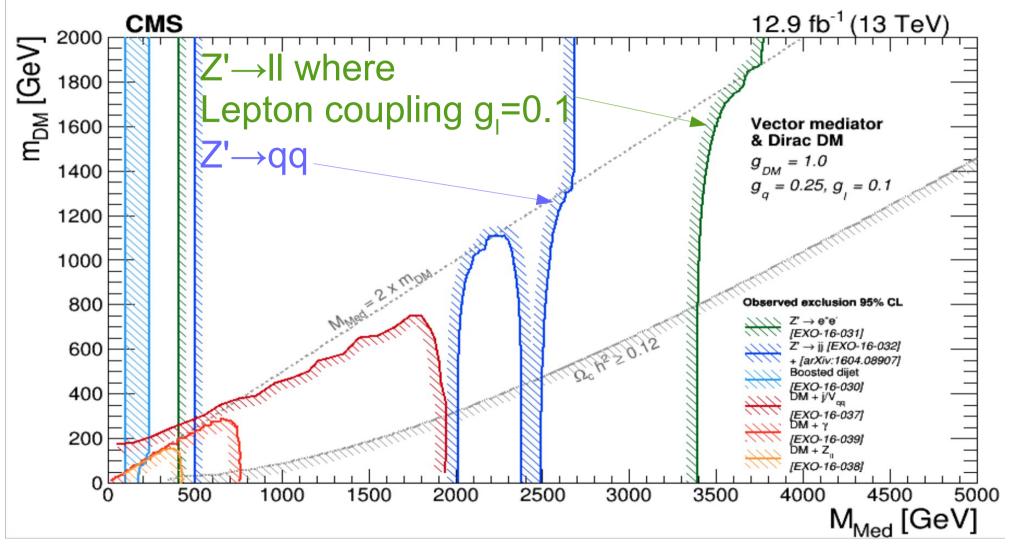


#### And for collider?

- For LHC
  - Our bounds are a bit more model dependent
  - We can start from the 14 TeV projections



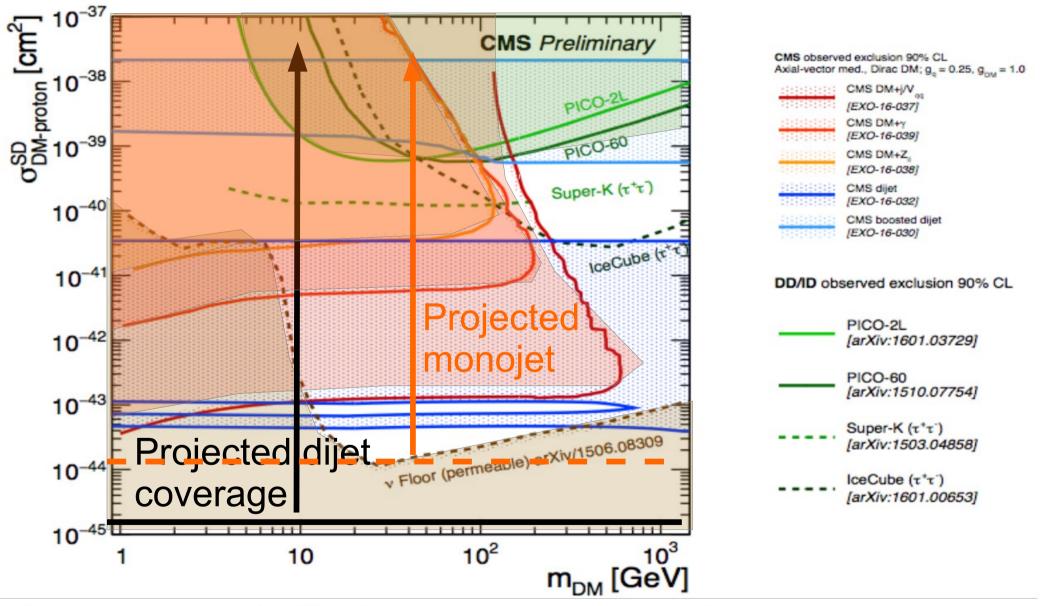
#### What about visible bounds?



Dijet reach : 2.5 TeV  $\rightarrow$  Becomes 5 TeV with 3ab<sup>-1</sup>  $M_{now} (L_{future}/L_{now})^{1/8}$ 

Dilepton reach: 3.5 TeV → Becomes 7 TeV with 3ab<sup>-1</sup> and coupling 0.1

#### How does this compare?

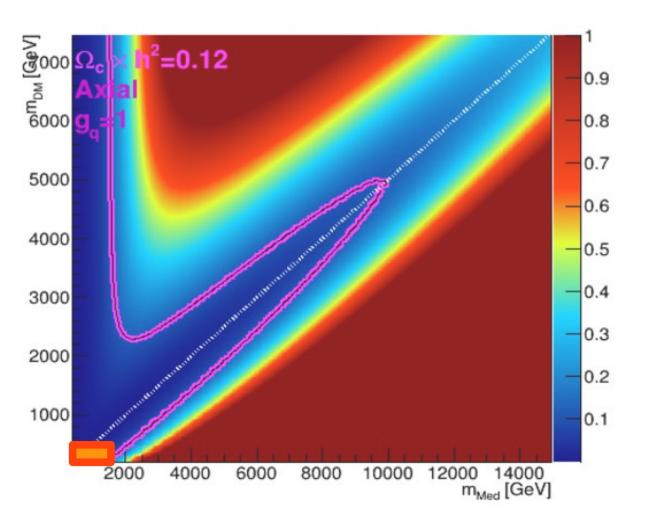


Dijet alone at 14 TeV exceed spin-dependent direct detection

# How about relic density?

For a benchmark model we start to cover with LHC: At FCC this is definitive

Note: Model is oversimplified bounds can loosen w/particles



Mediator mass Maxes out around 8-9 TeV

Smaller for coupling g<sub>a</sub><1

Approximate Dilepton reach

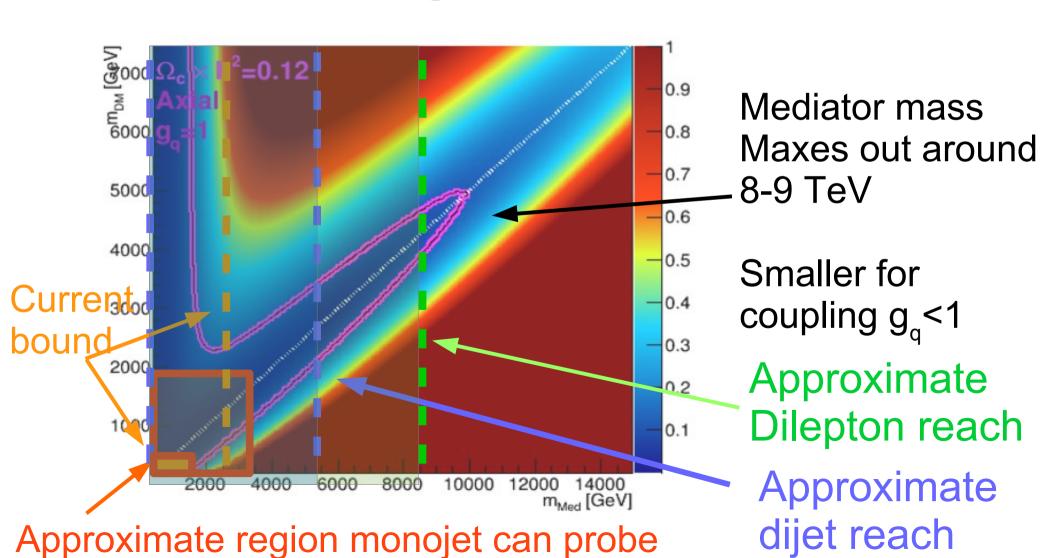
Approximate dijet reach

Approximate region monojet can probe

# How about relic density?

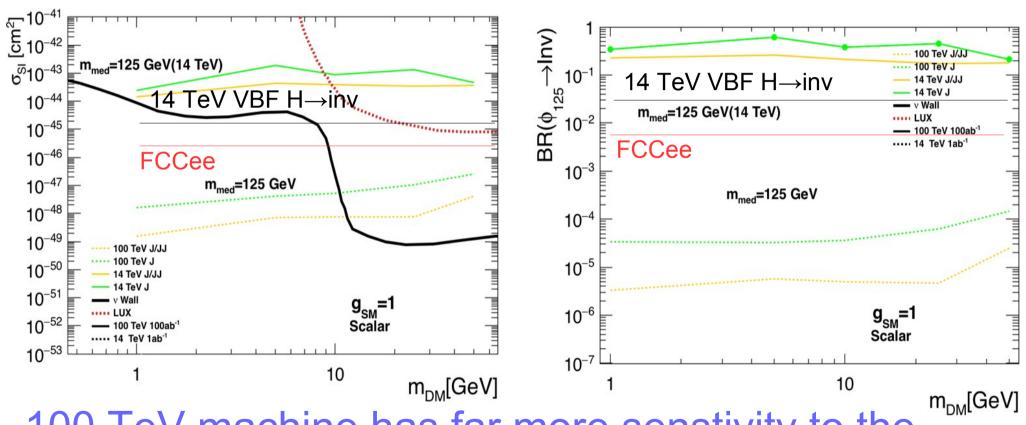
For a benchmark model @ LHC : Spin 1 Axial-vector med At FCC this is definitive

Note: Model is oversimplified bounds can loosen w/particles



# Higgs to invisible

A nice benchmark is the Higgs invisible:



100 TeV machine has far more senstivity to the invisible decays of a Higgs

https://arxiv.org/abs/1603.07739

https://indico.cern.ch/event/438866/contributions/1085169/attachments/1258088/1858101/FCCwck Hinv MDG 14042016.pdf

# Looking beyond monojet final state

- Deep understanding of monojet extends to many models
  - Disappear track
    - Monojet+track
  - Displaced jets

Final State	Analysis	section
jet+MET	Wino, Higgsino DM	4.3.1 - 4.3.4
jet+MET	Higgs Portal	4.3.5
jet+MET	Simplified Vector/Axial	4.4.1 - 4.4.3
jet+MET	Simplified Scalar/Pseudo	4.4.1 - 4.4.3
jet+MET	Gluion/stop coannihalation	4.5.1
VBF jets +MET	Wino, Higgsino DM	4.3.1 - 4.3.2
VBF jets +MET	Higgs Portal	4.3.5
photon+MET	Wino	4.3.2
Disappearing tracks	Wino,Higgsino	4.3.1 - 4.3.2
Disappearing tracks	Fiveplet DM	4.3.3
Disappearing tracks	Relic-Neutralino	4.3.4
lepton+ $\gamma$ +MET	Relic-Neutralino	4.3.4
$Z_D  o ll + (Z_D  o ll)$	Dark Photons	4.4.4, 4.6.3
displaced jets	Dark QCD/Hidden Valley	4.6.2
long lived charged particle	Super-WIMPS/Gravitino	4.6.4
dijet	Simplified Vector/Axial	4.4.1 - 4.4.3

**Table 5:** Overview of the final states and the associated model, with a link to the respective section.

#### What about the cross sections?

The relative rate to all processes is similar

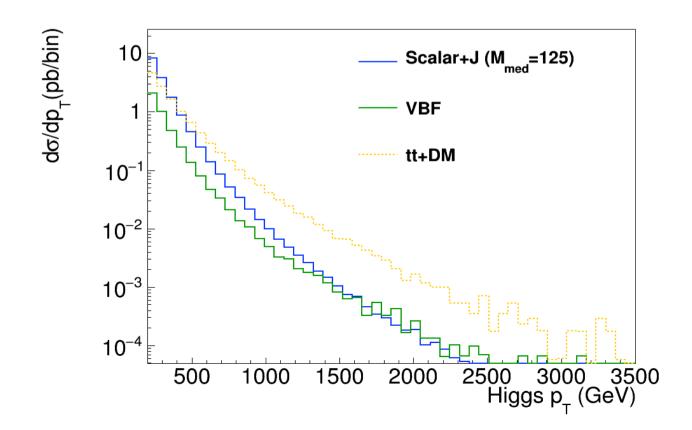
```
-\sigma(100 \text{ TeV}/14 \text{ TeV}) : ggH : 14.7
-\sigma(100 \text{ TeV}/14 \text{ TeV}) : VBF : 18.6
- \sigma(100 \text{ TeV}/14 \text{ TeV}) : WH : 9.8
-\sigma(100 \text{ TeV}/14 \text{ TeV}): ZH : 12.5
-\sigma(100 \text{ TeV}/14 \text{ TeV}): \text{ttH} : 60.8
- \sigma(100 \text{ TeV}/14 \text{ TeV}) : \text{bbH} : 14.8
- \sigma(100 \text{ TeV}/14 \text{ TeV}) : HH : 42.0

    Except for ttH
```

- Means we expect VBF to give similar improvement
- Benchmarking agains ggH means ttH/VBF have a lot of room to gain

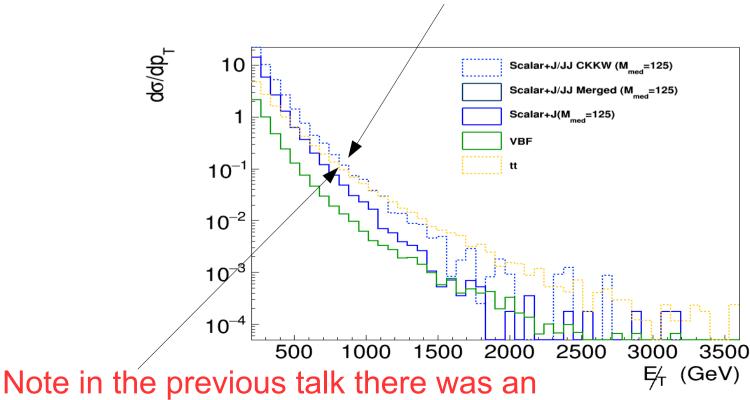
# What are the production modes?

- At 100 TeV:
  - ttH is hugely enhanced
  - When compared with H+1j form gluon fusion it wins



# What are the production modes?

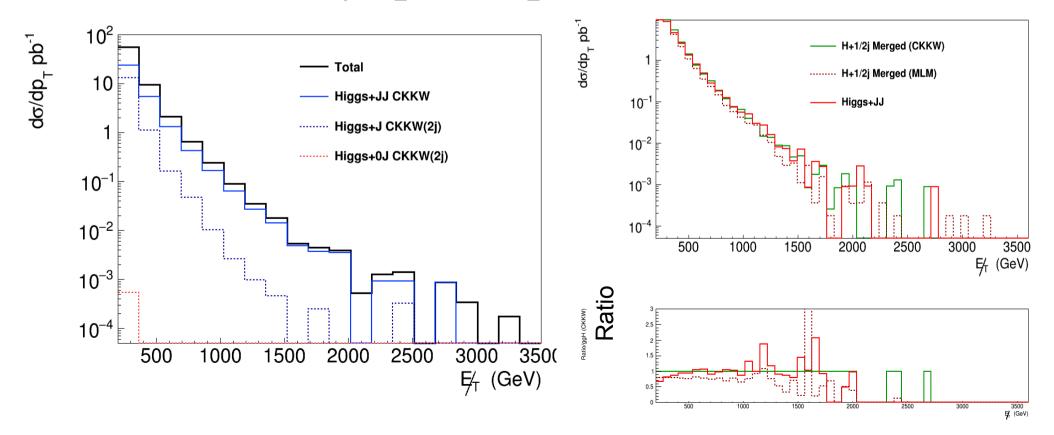
- At 100 TeV:
  - ttH is hugely enhanced
  - When compared with H+1j form gluon fusion it wins
  - However H+2j is also large



issue in the 2jet generation (was a bug)

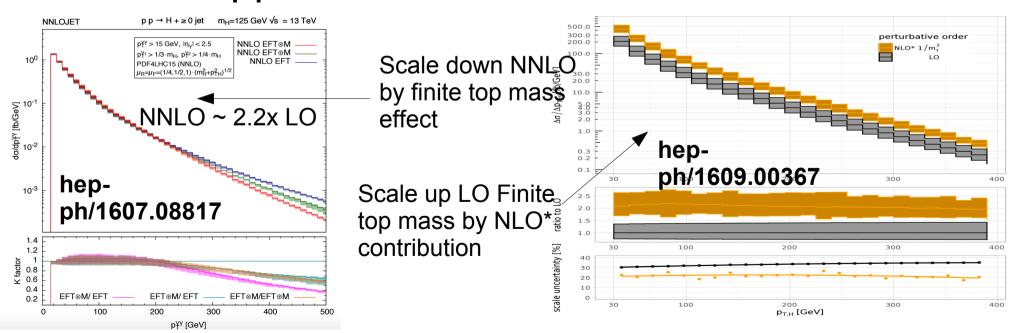
# Cross checking the 2jet model

- When this was previously present
  - There was a bug (turns out the impact is small!)
- At 100 TeV :
  - Different setups give roughly the same yield



# A note about Higgs p<sub>T</sub>

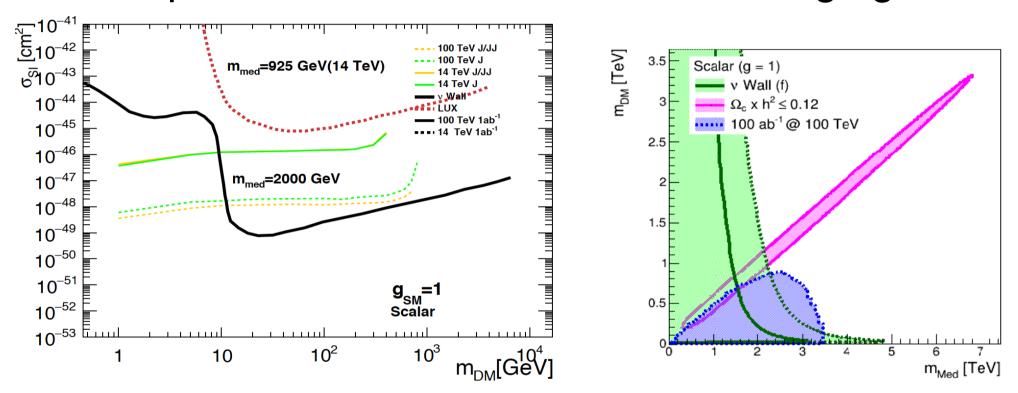
- We are using MG@NLO 0/1/2 jet LO finite m<sub>t</sub>
  - This generation accounts for the finite top mass
  - Finite top mass is the dominiant effect at high p<sub>T</sub>
  - However generation is alson only LO
- Several approaches to take into account NLO



We will scale the result by x2 after the fact

# Spin 0

For spin-0 the bounds are more challenging



Projections with a scalar simplfiled model indicate:

Close to crossing the neutrino wall with the FCC detector Close to reaching the bounds of direct detection

Rest of this talk: How do we impove these bounds? We will do this in the context of Higgs invisible