FCC Hadron Detector Meeting



Ecal energy deposition and simple clustering

31. 8. 2016

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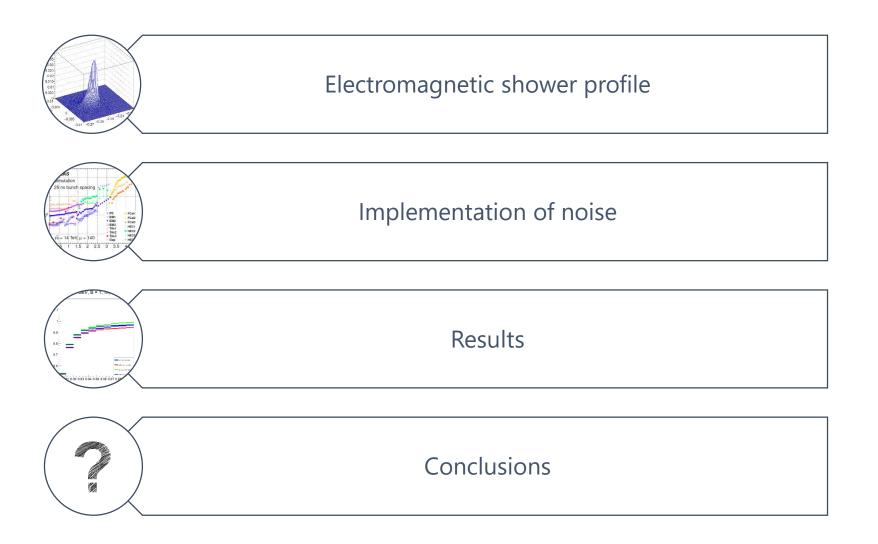
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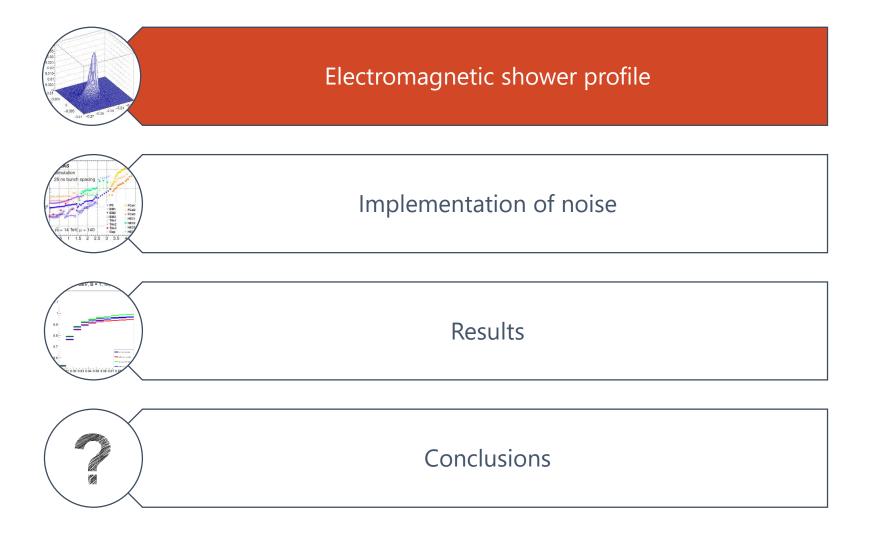
Outline





Electromagnetic shower profile





What we did



We depicted the **shower evolution** in the Electromagnetic calorimeter and determined the **influence of the magnetic field** on it.

We also **implemented** a simple **clustering algorithm** and **added noise** to the detector cells.

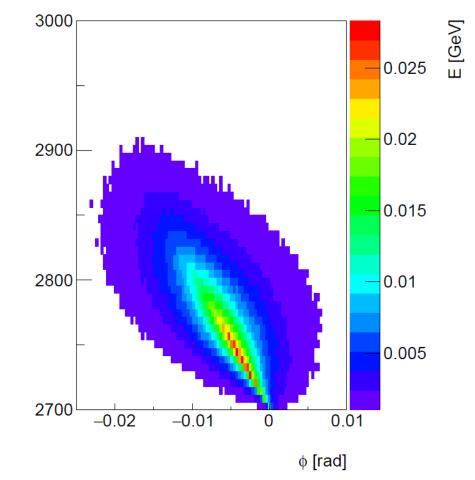
Monte Carlo simulations in the FCC software:

- Single electrons, E = 10, 100 GeV
- $\Phi \in [0, 2\pi], \eta = 0$
- B = 0, 6T
- With and without cryostat in front of the calorimeter
- Liquid Argon thickness = 4mm, Lead thickness = 2mm, R \in (2600 mm, 3500 mm)

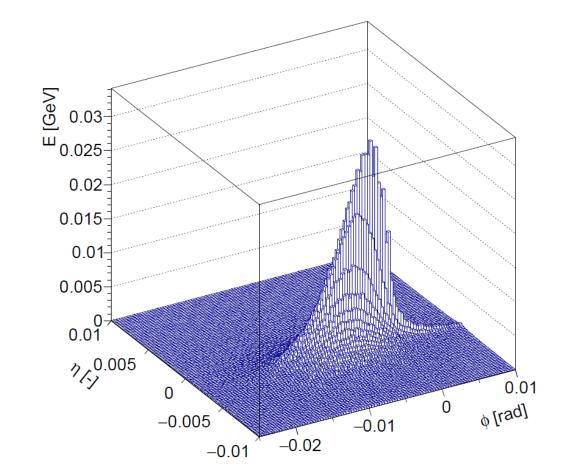


Energy deposition in the calorimeter, 10 GeV electron

E = 10 GeV, B = 1, with cryo



Edep = 10 GeV, B = 1, with cryo



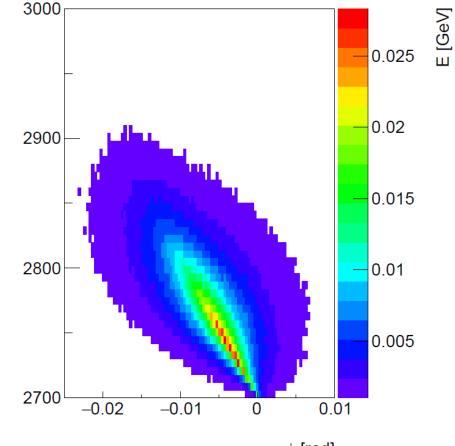
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R [mm]

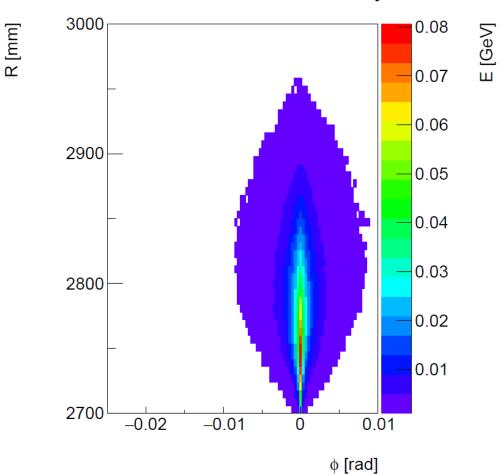


Energy deposition in the calorimeter, 10 GeV electron

E = 10 GeV, B = 1, with cryo



E = 10 GeV, B = 0, with cryo



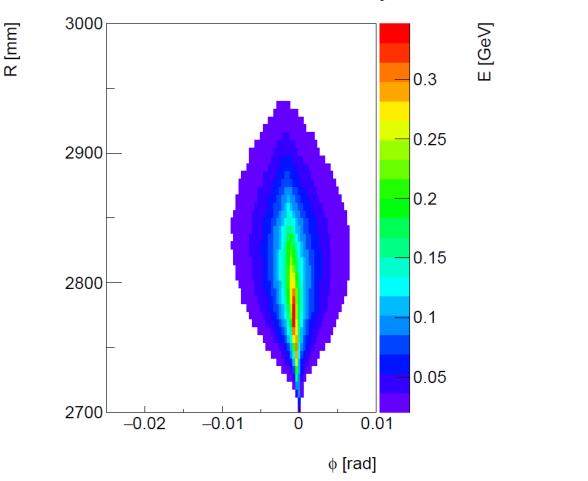
∳ [rad]

R [mm]

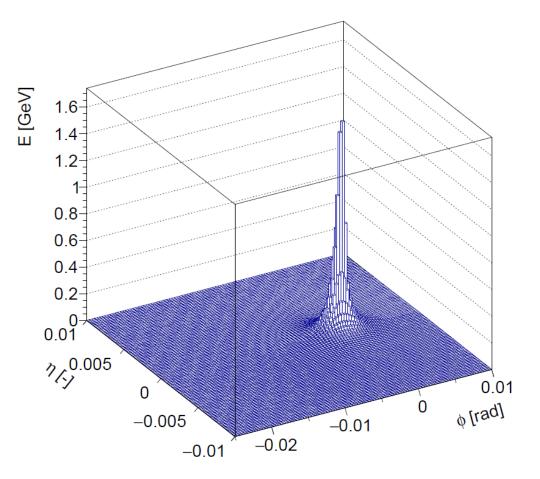


Energy deposition in the calorimeter, 100 GeV electron

E = 100 GeV, B = 1, with cryo



Edep = 100 GeV, B = 1, with cryo

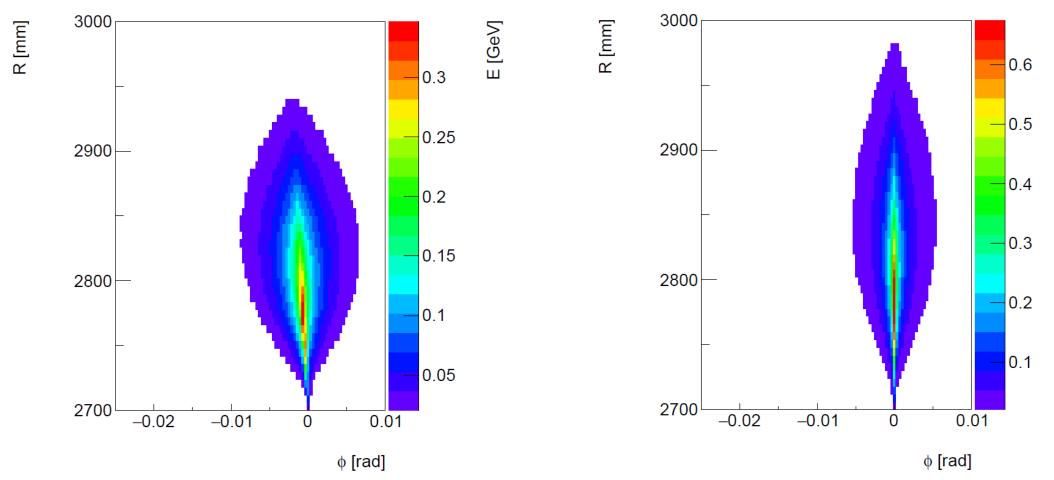




E [GeV]

Energy deposition in the calorimeter, 100 GeV electron

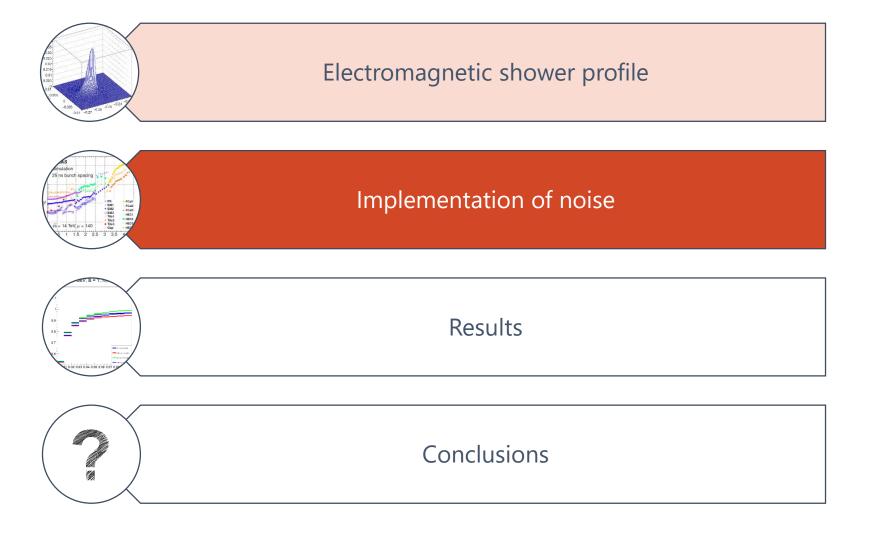
E = 100 GeV, B = 1, with cryo



E = 100 GeV, B = 0, with cryo

Implementation of clustering and noise

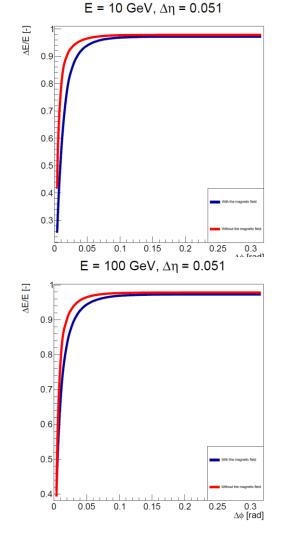




Reminder: Implementation of clustering

Then, we implemented a **simple clustering** algorithm in Φ and η that would help us find a good initial cell size.

- Impose a large neighborhood over the bin with maximum energy deposition.
- Integrate the deposited energy over the clustering window for every possible position in this neighborhood.
- Store the maximum energy deposit.
- Vary the dimensions of the clustering window in Φ.





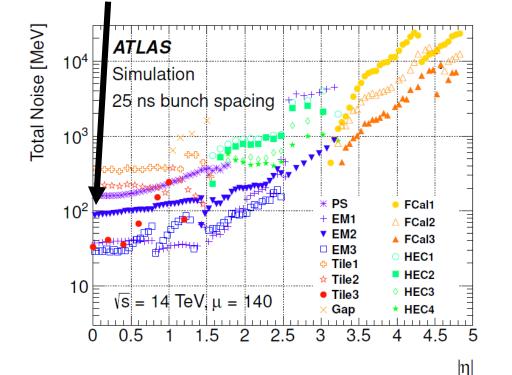
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Implementation of noise

Then, we implemented a **uniform Gaussian noise** to the detector cells in Φ and η .

First approximation was obtained from the ATLAS upgrade simulations where $\sigma = 100 \text{ MeV}$ (0.025 $\Phi \times 0.025 \eta$).

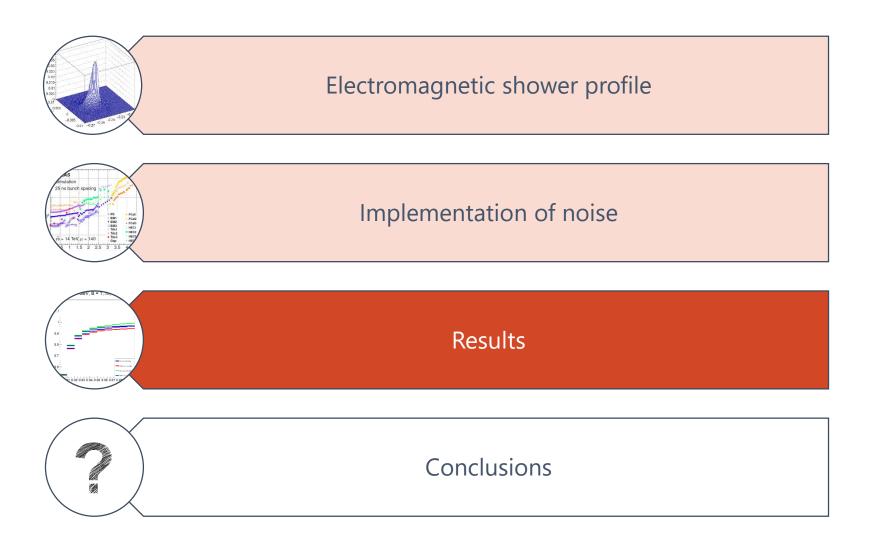
- For our binning of 0.005 Φ x 0.005 η, σ had to be adjusted to 20 MeV.
- There is no segmentation in R at the moment.
- For the largest cluster size considered (0.1 $\Phi \times 0.05 \eta$), $\sigma = 300 \text{ MeV}$





Results

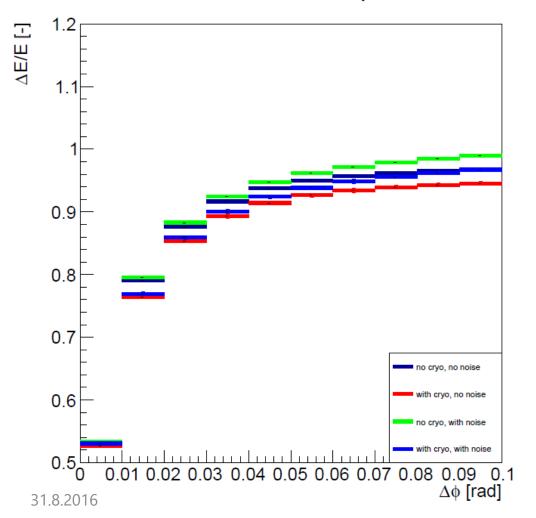




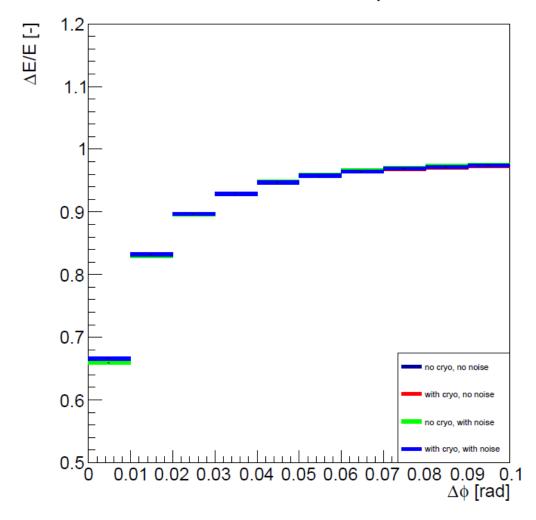
Results



E = 10 GeV, B = 1, Δη = 0.05

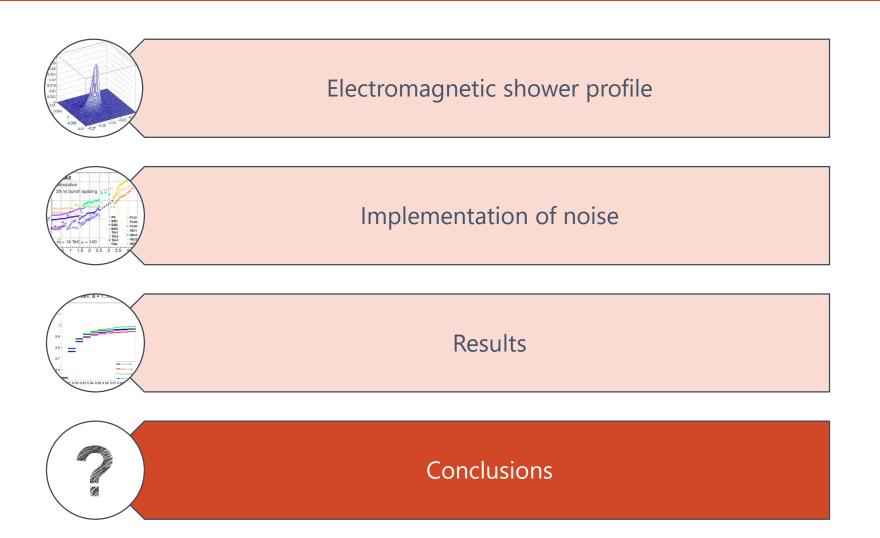


E = 100 GeV, B = 1, Δη = 0.05



Conclusions





Conclusions



Geometry of the shower, simple clustering and simulated noise was studied.

- Imposing a magnetic field in the detector changes the resulting showers both in their slope and shape.
- This effect is smaller for **100 GeV electrons** but still **non-negligible**.
- Addition of the noise increases the energy collected by the simple clustering by ~300 MeV for the largest clustering window (0.1 Φ x 0.05 η). More sophisticated clustering algorithm is needed.
- Addition of dead material in front of the detector (5 cm cryostat) decreases the deposited energy by ~300
 MeV in the presence of the magnetic field.
- Next: Repeat the simulation with **new dimensions** and **include tracker**.

Thank you for your attention

Ecal thickness vs energy resolution

• Energy resolution: $\frac{\sigma}{E} = \frac{a}{\sqrt{E}} \oplus \frac{b}{E} \oplus c$

a: stochastic term (statistical fluctuations) b: noise term (electronic noise) – we set it to zero c: constant term (leakage, uniformity)

- Simplified Ecal simulations with single electrons at different energies (from 20 ٠ to 1000 GeV)
 - No B field, no cryostat, no noise
- Difference in the resolution between 30 X_0 and 25 X_0 is very small
 - Current default is 30 X_0 (in ATLAS at $\eta = 0$: 22 X_0)
 - Ecal depth of 25 X₀ could be considered

ECAL depth (X_{θ})	а	С
25	10.1%	0.53%
27	10.4%	0.36%
30	10.6%	0.20%

Example: 6 mm LAr + 4 mm Pb

