# Clustering algorithm

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> FCC hadron detector meeting 14 December 2016

### Status of calorimeter reconstruction

In official repository of FCC Software. <u>Reconstruction documentation</u>

**Digitisation:** 

- Merging energy deposits in cells \_
- Calibration to the EM scale \_
- Noise -

**Clustering algorithm:** 

- Sliding window algorithm -
  - Electron / photon reconstruction \_
  - Used in ATLAS [ATLAS note] -

### Noise in the ECAL

#### **Electronic noise**

- Scales with detector capacitance [FCC meeting]

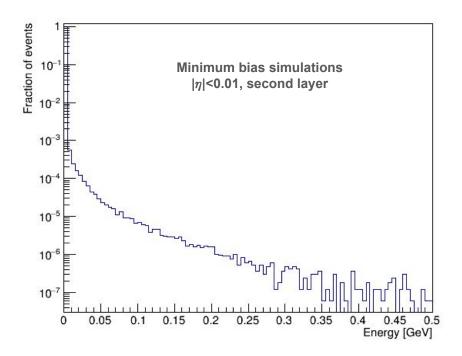
#### Pile-up

- In-time pile-up: a number of minimum bias events in the same bunch crossing
- **Out-of-time pile-up**: contribution from minimum bias events from the previous bunch crossing (drift time)
- Depends on  $\Delta \eta \ge \Delta \phi$ , radial layer, material in front of the calorimeter, energy of the minimum bias collisions
  - Difficult to be extrapolated from ATLAS

### Pile-up noise

## Estimation based on FCC full simulations

- Simulation of minimum bias (MB) collisions in ECAL (one per bunch crossing)
- Corrections
  - Scaling with avg. number of MB events per bunch crossing  $(\sqrt{\mu})$
  - Out-of-time pile-up

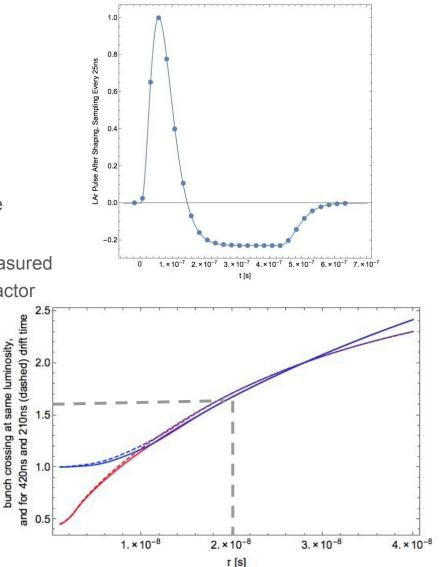


### Out-of-time pile-up

pile-up noise factor for 5ns and 25ns

#### Pulse-shape in ATLAS

- Drift time 420 ns
- Pulse is digitized every 25 ns
- Bipolar shaping
  - The sum of in-time pile-up and out-of time pile-up yields zero
  - Fluctuations will be higher than those measured for the in-time pile-up only → correction factor



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#### **Correction factor: 1.6**

(shaping time ~20 ns, for 5 ns and 25 ns bunch crossings)

### Pile-up noise

Simulations of min. bias event: *RMS*<sub>1MB</sub> per cell

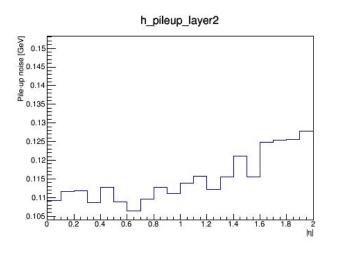
Final pile-up noise:  $RMS_{1MB}^*$  1.6 \*  $\sqrt{\mu}$  ( $\mu$  = 1000)

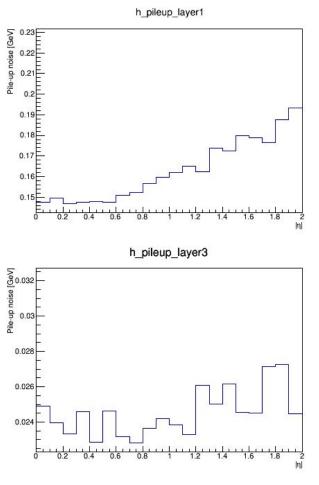
Three radial layers, cell size 0.01 x 0.01 in  $\Delta \eta \times \Delta \phi$ 

Notes:

- Largest pile-up in the first radial layer
- Pile-up increases with  $\boldsymbol{\eta}$







### Sliding window algorithm

0. Calorimeter towers with fixed  $\eta \ge \varphi$  size

- 1. Seeding
  - Scanning the  $\eta \ge \phi$  tower map with a fixed size window for local maxima
  - If energy inside window is above threshold  $\rightarrow$  mark as pre-cluster

#### 2. Barycentre position calculation

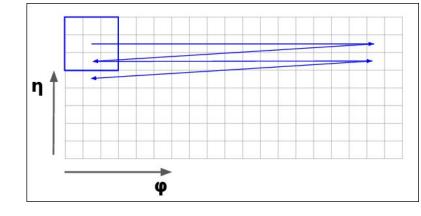
- Energy-weighted position for each pre-cluster

#### 3. Duplicates removal

- If two pre-clusters are next to each other, the pre-cluster with lower energy is removed

#### 4. Cluster building

Each step (1-4) can use window of different size (centred around the tower seed)



### Proof of principle

-0.26

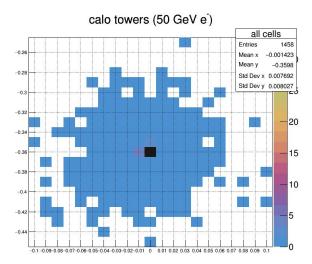
-0.1

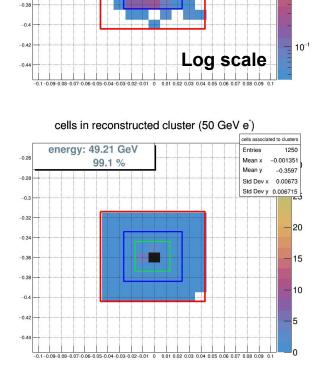
-0.3

-0.3

-0.36

Sizes of the windows (seeding, position calculation, duplicate removal) and the energy threshold are to be optimised.





calo towers (50 GeV e)

energy: 49.65 GeV

all cells

Std Dev x 0.007692

Std Dev y 0.008027

1458

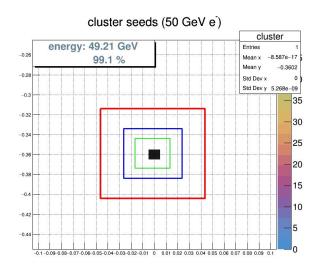
-0.001423

-0.3598

Entries

Mean x

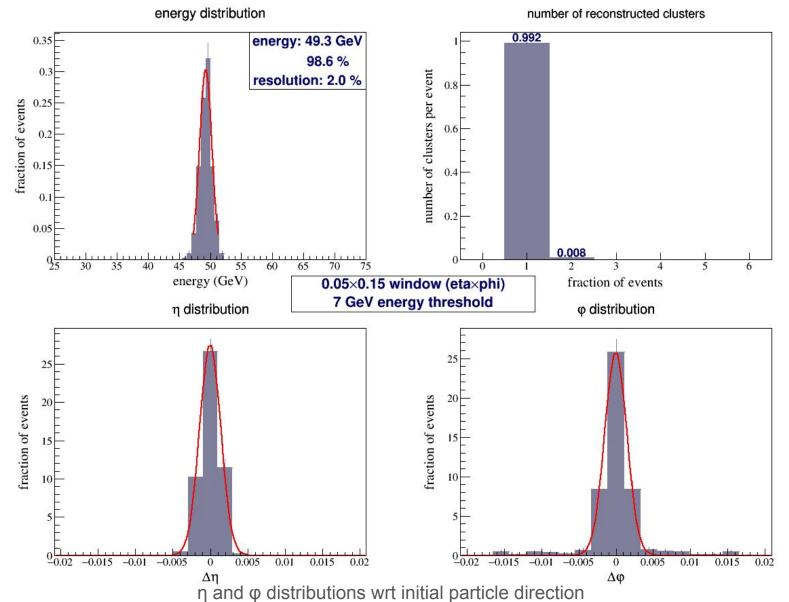
Mean y



### Simple noise

no tracker, no cryostat

Gaussian noise (all cells:  $\sigma$  = 10 MeV)

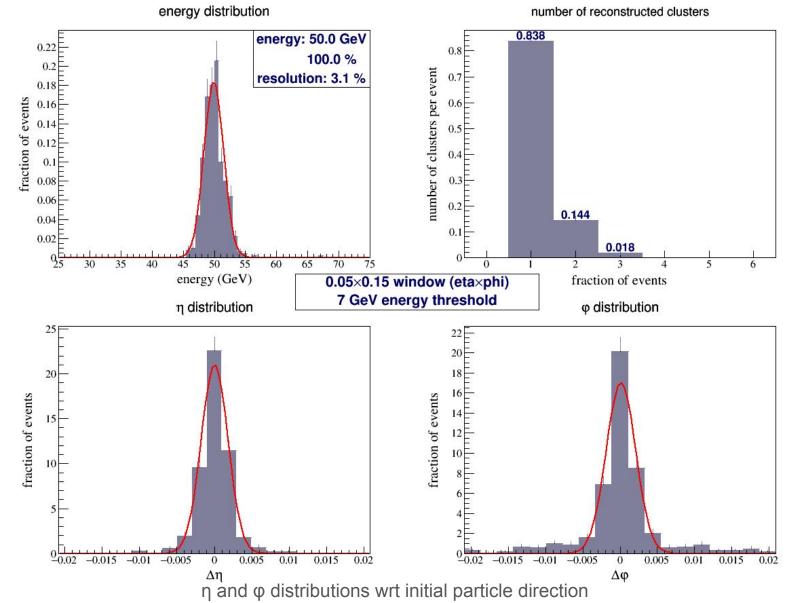


### More realistic noise

#### 500 single electron events (50 GeV) B = 0

no tracker, no cryostat

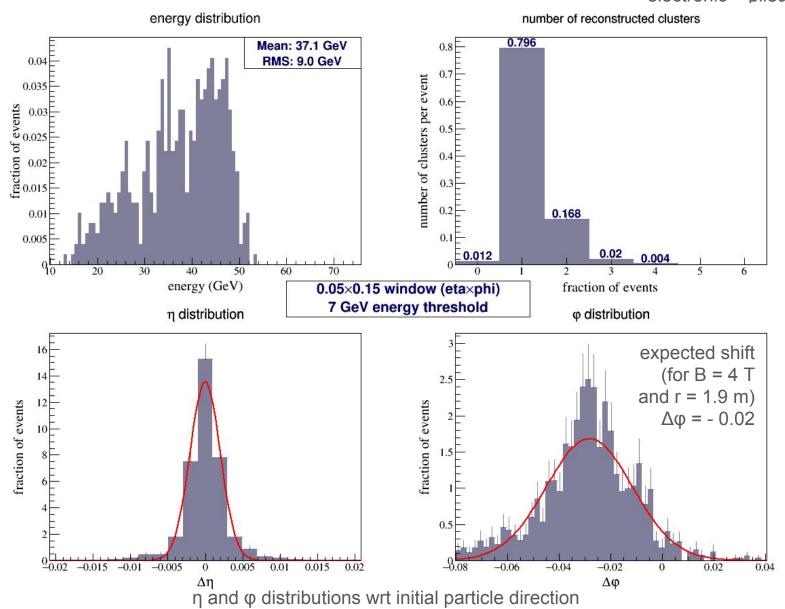
electronic + pileup noise



### With tracker and B field

#### 500 single electron events (50 GeV) B = 4 T with tracker + cryostat

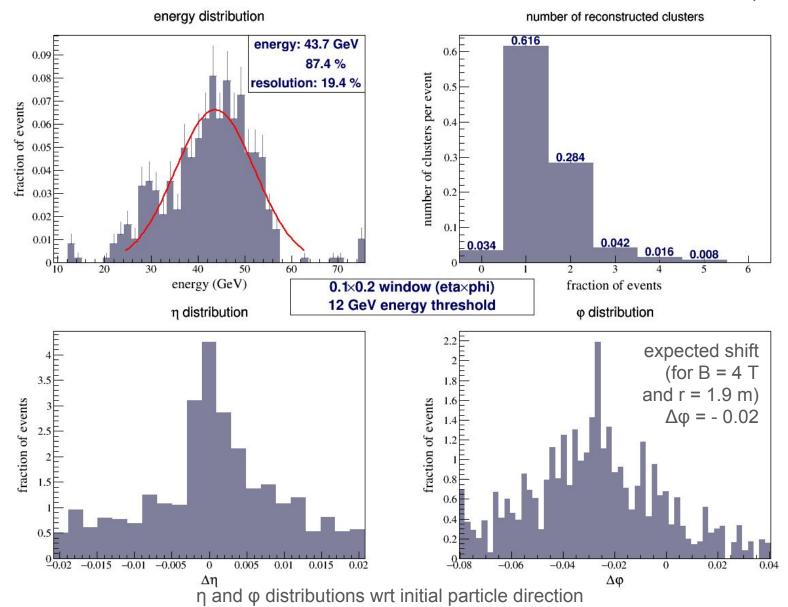
electronic + pileup noise



Work in progress

### Larger window

electronic + pileup noise



Work in progress

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#### (just started)

#### Parameter optimisation

window 0.1 x 0.2

window 0.05 x 0.15

energy of cluster duplicates energy of cluster duplicates 3500 0.05×0.15 window (eta×phi) 0.1×0.2 window (eta×phi) 3000 6 GeV energy threshold 6 GeV energy threshold 3000 69 duplicate(s)/event 191 duplicate(s)/event 2500 number of clusters 12000 12000 12000 of clusters 2000 clusters energy threshold under 1000 4 GeV 500 500 Same trend for  $0^{L}_{0}$  $0^{\perp}_{0}$ 12 14 2 4 6 8 10 12 14 2 4 6 8 10 all energies E (GeV) E (GeV) -> contribution energy of cluster duplicates energy of cluster duplicates from noise 1400-0.1×0.2 window (eta×phi) 0.05×0.15 window (eta×phi) 120-6 GeV energy threshold 6 GeV energy threshold 1200 1 duplicate(s)/event 57 duplicate(s)/event 100 number of clusters number of clusters energy 80 threshold 60 6 GeV 40400 20 200 0 0 12 14 2 4 6 8 10 12 14 2 6 8 10 4 E (GeV) E (GeV)

> Reconstruction parameters may depend on each other Starting with seeding window size and energy threshold

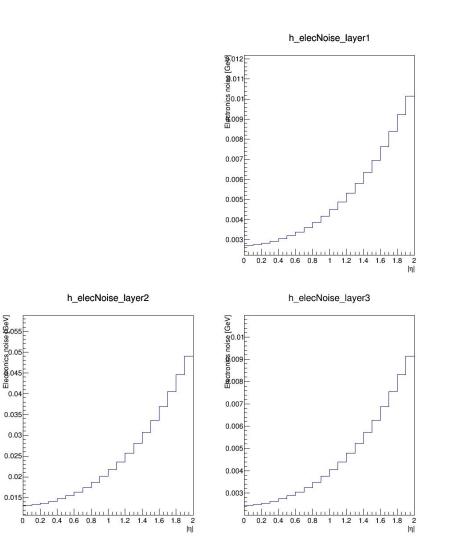
### Summary

- Pile-up noise estimation in ECAL updated
- Sliding window algorithm implemented
  - Validation done without any material in front of the detector
- Ongoing work: Optimisation of sliding window parameters
  - Reconstruction efficiency and fake clusters rejection
  - Aim to reconstruct as low energetic particles as possible

#### BACKUP

### Minimum bias energy distribution

#### **Electronic noise**



#### 500 single electron events (50 GeV) B = 0no tracker, no cryostat

#### electronic + pileup noise

energy\_diff

Mean 0.7703

Std Dev 0.2911

di i

0.9 1 ΔE/E

eta\_diff

Mean -0.0135

phi\_diff

R\_diff

Std Dev 0.9954

A B

Mean 0.09792

92

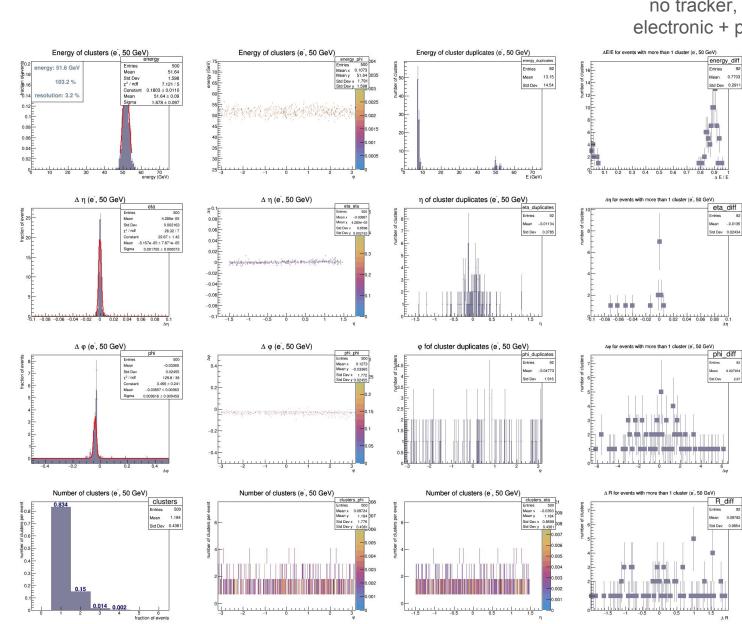
0.007934

intries. 02

Std Dev 2.67

92 Entries

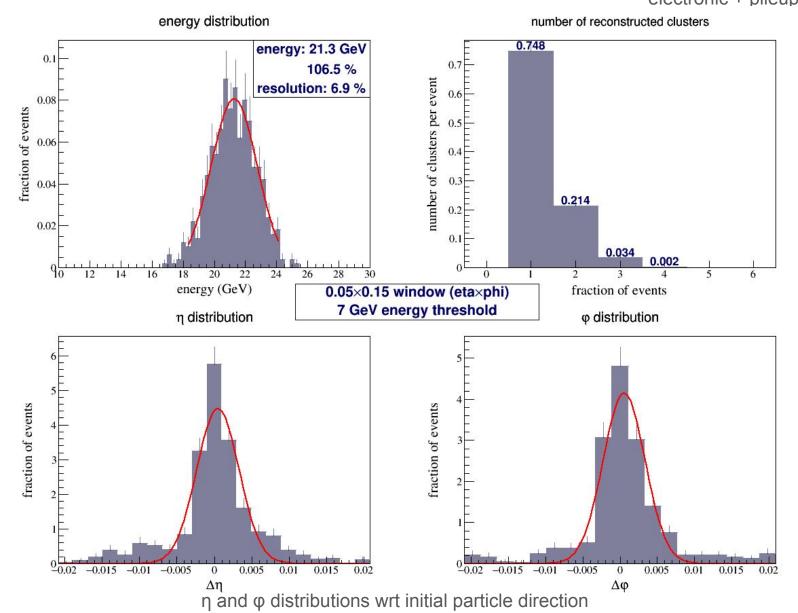
Entries 92



50 GeV

20 GeV

no tracker, no cryostat electronic + pileup noise



10 GeV

#### 500 single electron events (10 GeV) B = 0

no tracker, no cryostat electronic + pileup noise

