# Next-To-Next-To-Leading Order Calculations on HPC

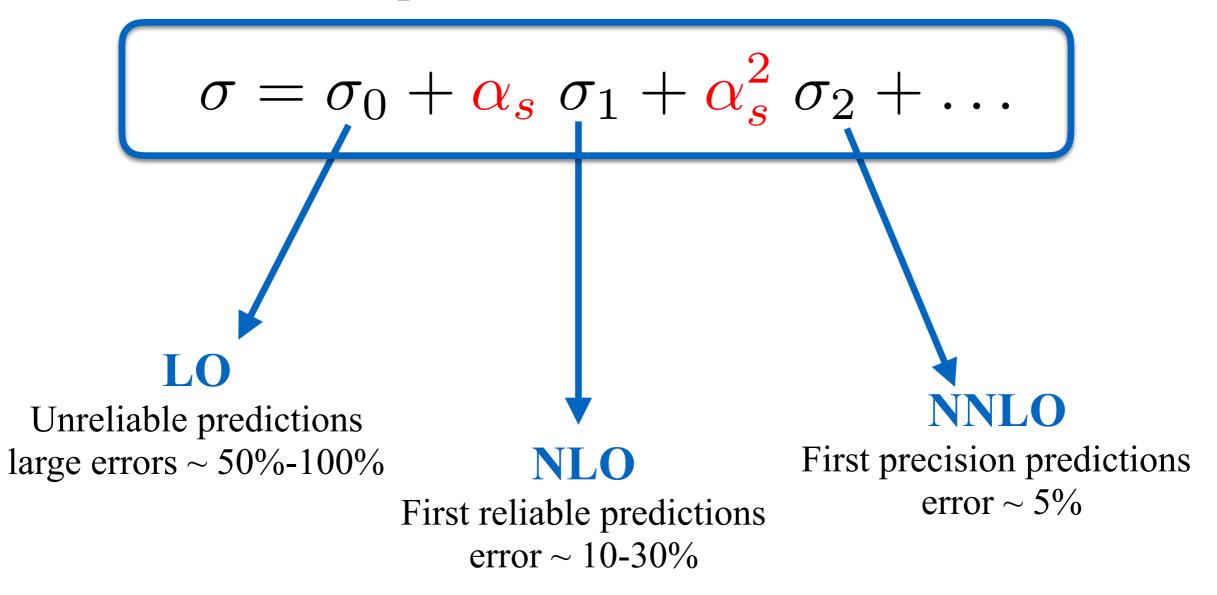
#### Radja Boughezal



Beyond Leading-order Calculations on HPC September 22, 2016

#### How we make predictions

 We expand our cross section in powers of the strong coupling constant α<sub>s</sub> (a perturbative expansion), and calculate each term in that expansion:



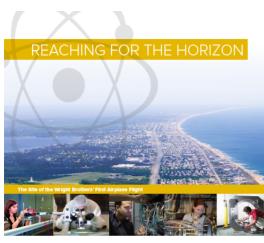
## Motivations for high-precision theory





From high energy physics:

"The full discovery potential of the Higgs will be unleashed by percent-level precision studies of the Higgs properties."



The 2015 LONG RANGE PLAN for NUCLEAR SCIENCE

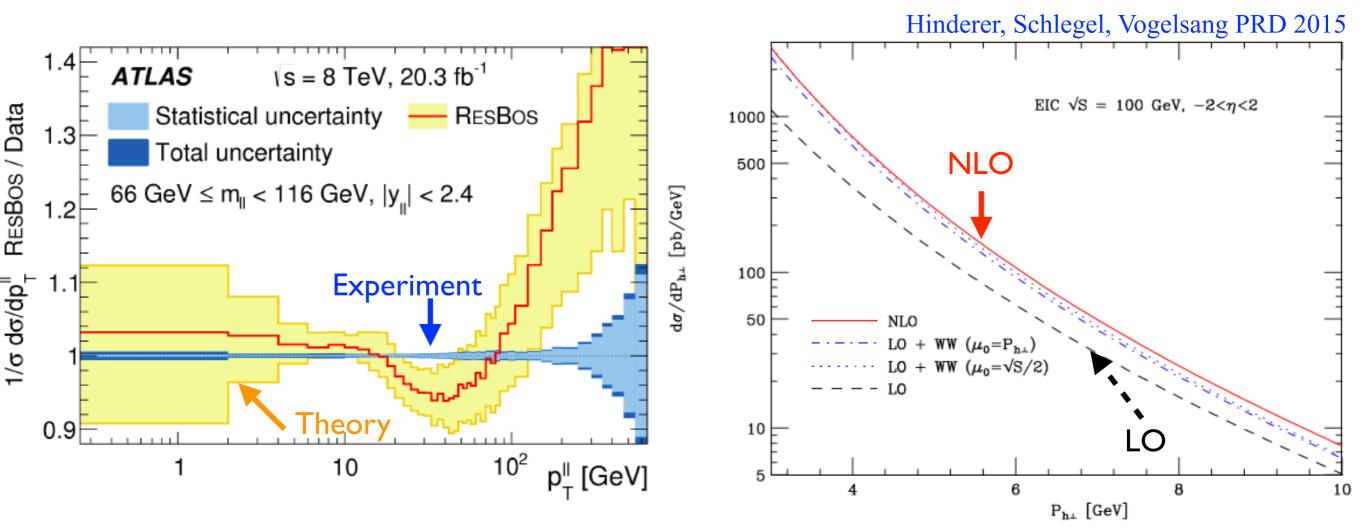
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From nuclear physics:

"To meet the challenges and realize the full scientific potential of current and future experiments, we require new investments in theoretical and computational nuclear physics."

### Motivations for high-precision theory

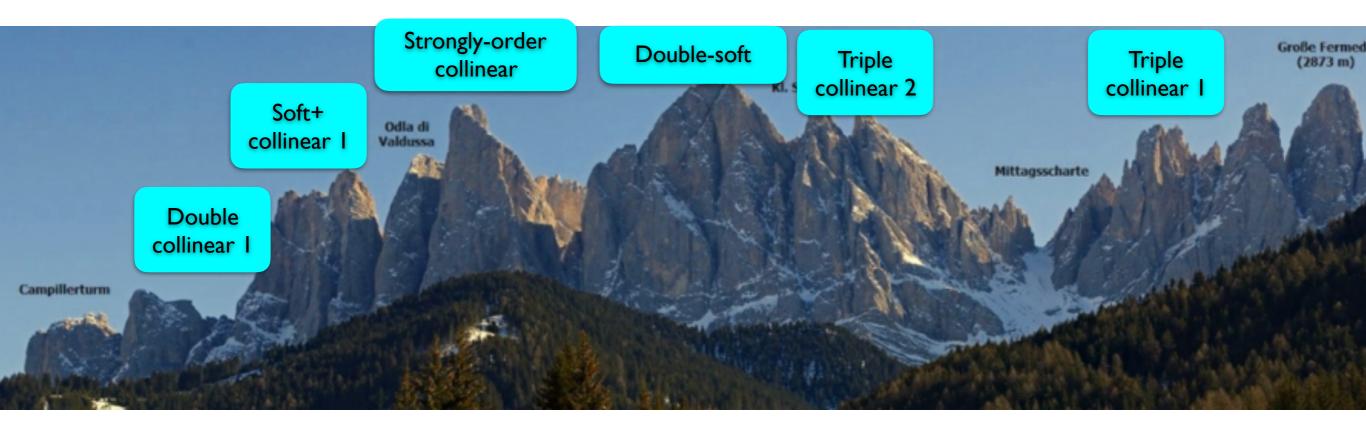
• Current data, and anticipated future data, clearly demonstrate the need for predictions beyond the current next-to-leading order (NLO) standard



Theory errors more than an order of magnitude larger than experimental ones NLO corrections for EIC hadron production increase LO result by more than a factor of 2

### Anatomy of a NNLO calculation

• Next-to-next-to-leading order calculations (NNLO) in a nutshell: a collection of integrals over integrands with the following structure:

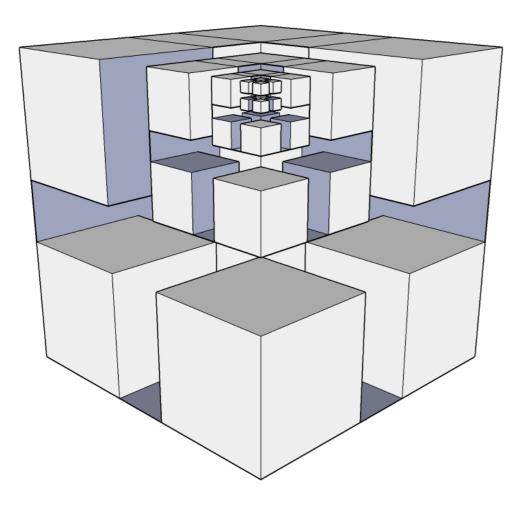


- Proliferation of possible singular configurations leads to a high-dimensional integrand with many peaks scattered throughout integration space
- Techniques require significant "by-hand" manipulation; can they be automated?
- Are the theoretical algorithms scalable to effectively use HPCs?

## An example of a traditional approach

• Sector-improved residue subtraction: an example of a more traditional approach. The integration region is divided with carefully chosen variable changes into O(200) independent sectors; each sector handles a particular configuration

Binoth, Heinrich 2000; Anastasiou, Melnikov, Petriello 2004; Czakon, 2010; RB, Melnikov, Petriello 2011



- Each sector has a different complexity and runtime; some run quickly on a single core, some require many cores to reach the requisite precision
- Significant by-hand manipulation that differs for each process

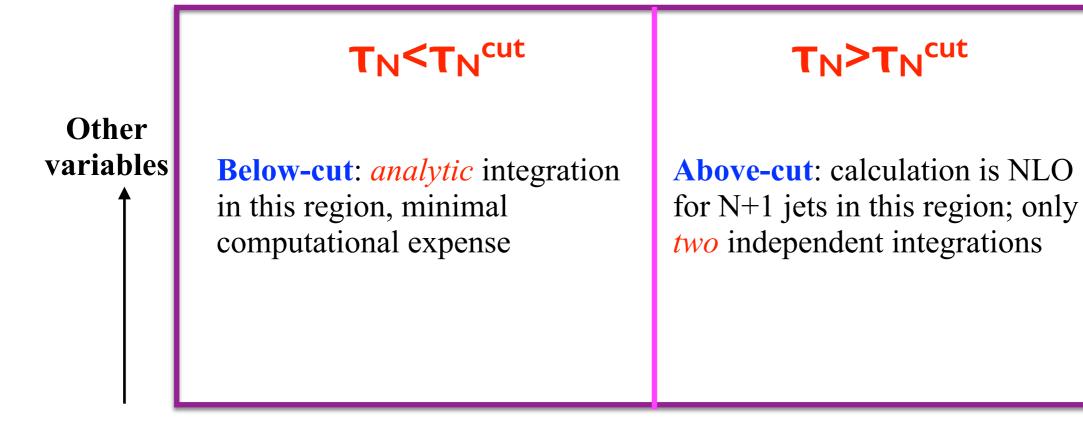
Difficult to automate and parallelize

# A new idea

#### N-jettiness subtraction

RB, Focke, Liu, Petriello PRL 2015; Gaunt, Stahlhofen, Tackmann, Walsh JHEP 2015

Divide integration region into two pieces:



ΤN

## The N-jettiness scorecard

#### • Advantages:

✓ Reduction of integrations: 200→3 (only 2 that have computational costs)
 ✓ Can reuse existing NLO frameworks for above-cut region (below-cut region is trivial to code)

Much less process-dependent work, easier to release and maintain public codes incorporating many processes

#### • Challenges:

Stringent precision requirement on the 2 non-trivial integrations (~0.01%)

(technical reason: cancellation of  $log(\tau_N^{cut})$  between above/below cut regions introduces numerical noise)

• NNLO results:

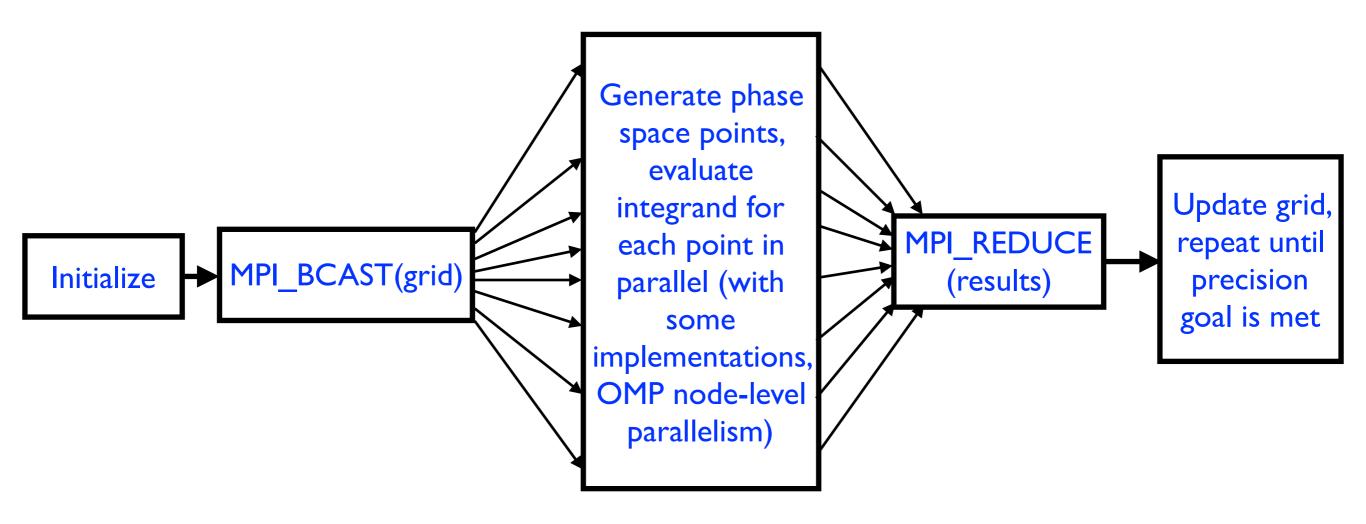
(all in past 1.5 years!)

- W+jet: first result (RB, Focke, Liu, Petriello PRL 2015)
- Z+jet: first result (RB, Campbell, Ellis, Focke, Giele, Liu, Petriello PRL 2016)
- EIC jets: first result (Abelof, RB, Liu, Petriello, 2016)
- Higgs+jet (RB, Focke, Giele, Liu, Petriello PLB 2015)
- Higgs,W,Z,ZH,WH (RB, Campbell, Ellis, Focke, Giele, Liu, Petriello, Williams 2016, MCFM 8.0)
- γγ: corrected previous result in literature (Campbell, Ellis, Li, Williams, 2016)

Use of HPC instrumental in addressing this challenge!

### Overview of HPC implementation

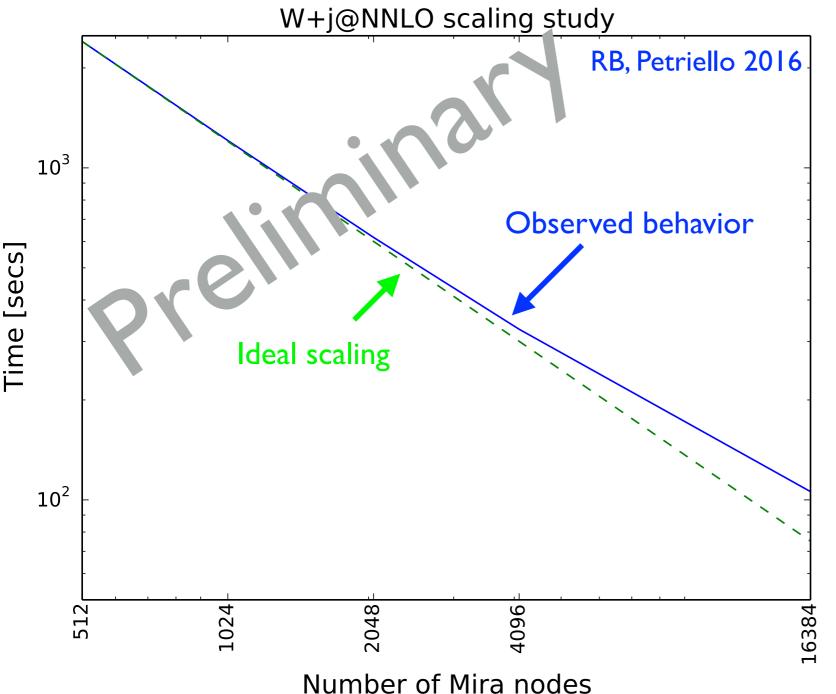
- Use adaptive Monte Carlo integration (VEGAS) for integrations
- Fundamental assumption: evaluation of integrand dominates runtime



• Performance measure: strong scaling of the implementation

### Mira scaling study

• Here is a representative example: scaling study of the W+jet@NNLO process on Mira (similar results for other processes, and on Edison@NERSC)



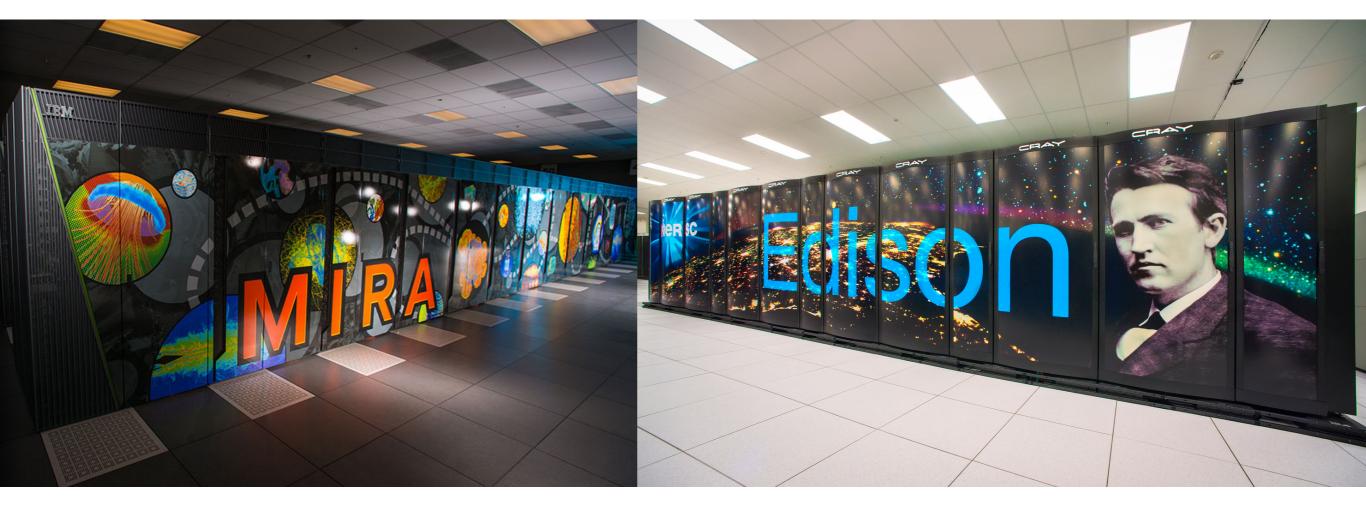
- 20% deviation from ideal scaling for a 16K Mira partition
- Known cause: VEGAS grid adaptation on a single core
- Can be improved, but we usually run single integrations on 4K or less partitions
- Future: scaling improves as we tackle more complicated processes!

### N-jettiness implementations

• MCFM 8.0 (RB, Campbell, Ellis, Focke, Giele, Liu, Petriello, Williams 2016)

- Implementation of color-singlet processes: H, W, Z, WH, ZH, γγ
- Fortran77+Fortran90
- MPI+OpenMP VEGAS algorithm
- Ongoing development for color singlet+jet processes
- DISTRESS (Abelof, RB, Liu, Petriello 2016)
  - Implementation of jet production in deeply-inelastic scattering (DIS)
  - Fortran 90
  - MPI VEGAS algorithm (OMP under development)
  - Ongoing development for other DIS processes

#### HPC resources



- Director's Discretionary QCDHPC (PI: RB)
- ALCC EnergyFEC (PIs: Childers, LeCompte)

- m2405: QCDNNLO (PI: RB)
- m1758: THEP-HPC (PI: Hoeche)

#### HPC resources

Results of 8 papers, two community reports and contributions to 4 experimental analyses were produced on these facilities between 2015-2016. They are vital for us!

- Director's Discretionary QCDHPC (PI: RB)
- ALCC EnergyFEC (PIs: Childers, LeCompte)

- m2405: Ultra-precise Predictions for the Terascale (PI: RB)
- m1758: HPC for QCD Theory (PI: Hoeche)

### Details of running on Mira

- Typically run for several  $\tau_N^{cut}$  values to monitor systematic effects on the integration; also for several different seeds for the random number generator
- All jobs should finish at the same time to allow for systematic error study, statistical averaging, and analysis (all done off-site with in-house python scripts)
- Sweet spot: single-grid integrations on 1024/2048-node partitions, bundled into 16K ensemble jobs (may change for future applications!)

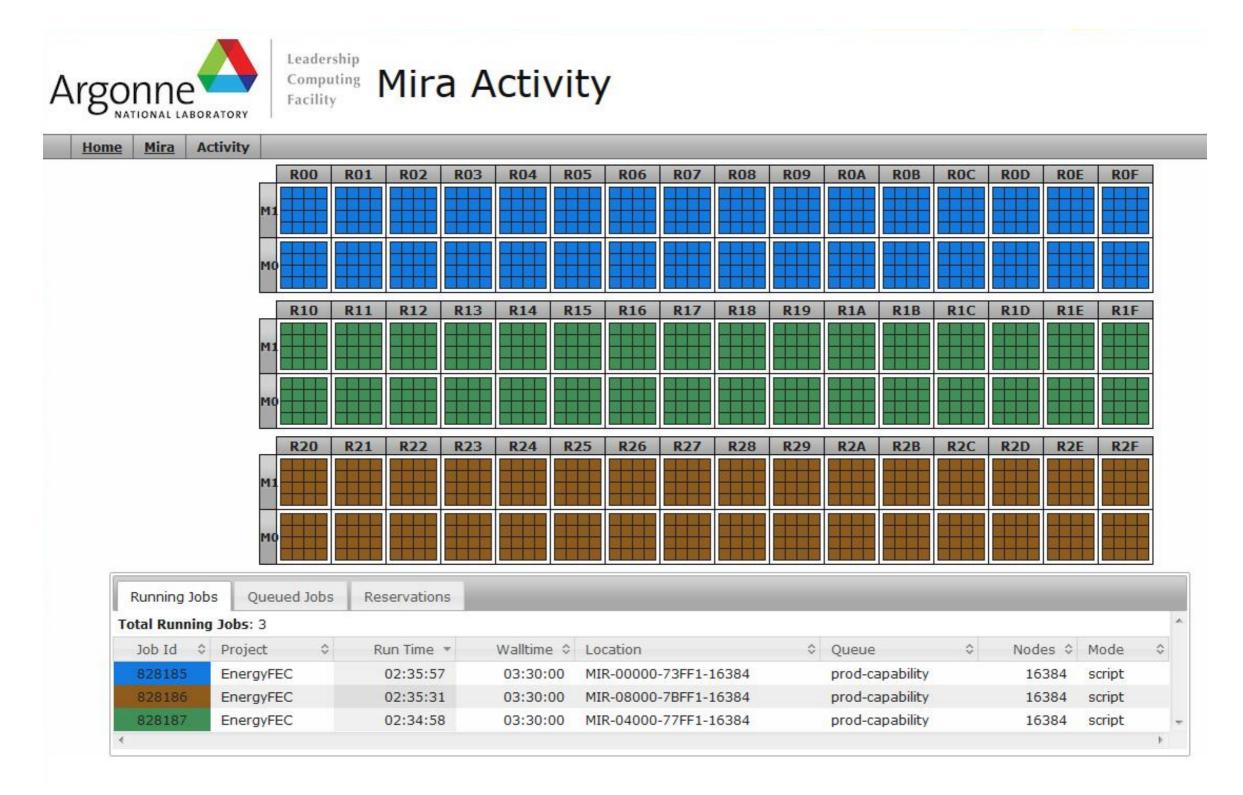
#### Example single job:

```
qsub -A EnergyFEC -t 45 -n 1024 --env I_MPI_PLATFORM=auto
--mode c64 ./distress y0=1d-14 tau1cut=0.00005 seed=800
outfile=tau00005_s800.dat
```

Example ensemble job:

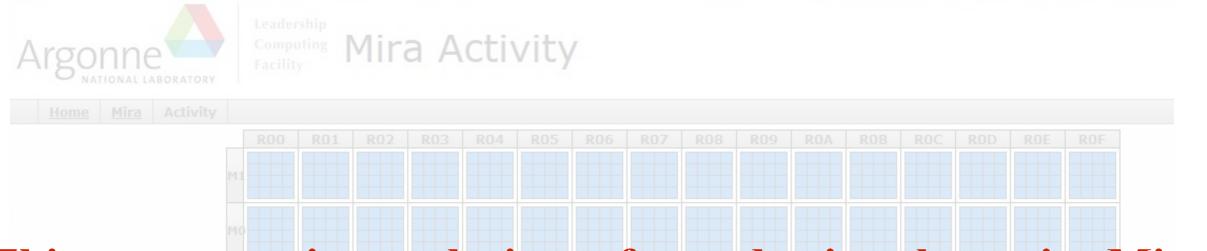
```
qsub -A EnergyFEC -t 60 -n 16384 --disable_preboot
--mode script enscript_etajet.sh
```

#### Details of running on Mira



#### We will happily use the entire machine!

### Details of running on Mira



#### This was an entire analysis performed using the entire Mira!

#### A comparison of NNLO QCD predictions with 7 TeV ATLAS and CMS data for V+jet processes

Radja Boughezal (Argonne), Xiaohui Liu (Maryland U., College Park), Frank Petriello (Argonne & Northwestern U.). Feb 17, 2016. 8 pp. Published in Phys.Lett. B760 (2016) 6-13



otal Running	Jobs: 3						
			MIR-00000-73FF1-16384		16384	script	
			MIR-08000-7BFF1-16384		16384	script	
		02:34:58	MIR-04000-77FF1-16384		16384	script	

#### We will happily use the entire machine!

### Details of running on Edison

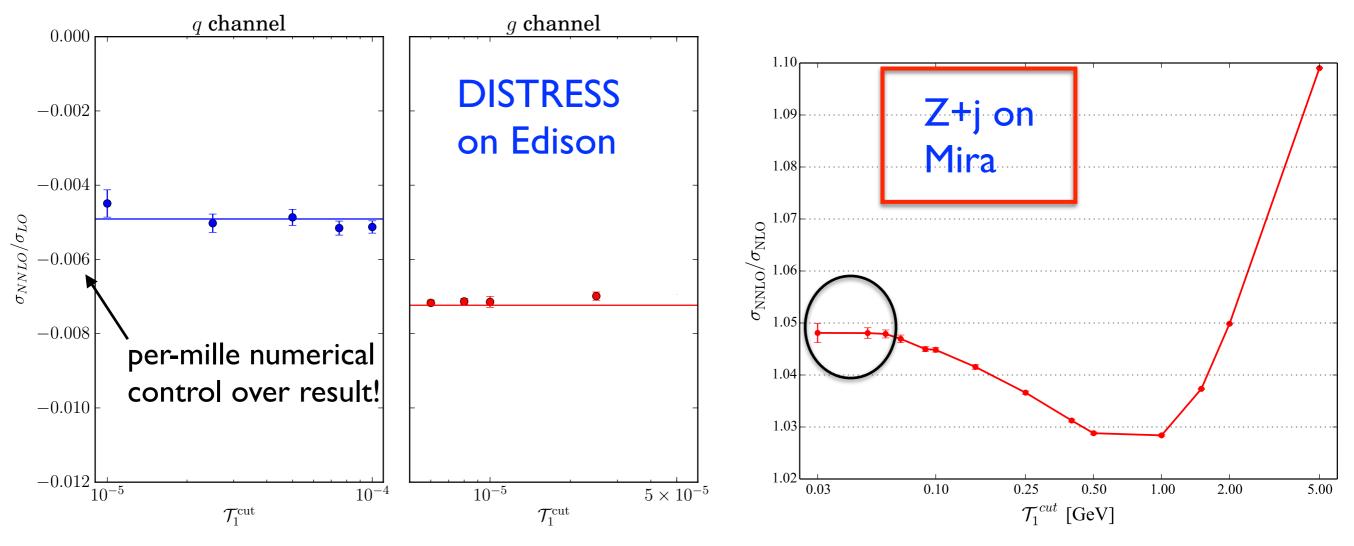
- Same run parameter considerations as on Mira (multiple  $\tau_N^{cut}$ , VEGAS seeds)
- Observe good scaling with DISTRESS: run 684-node jobs to benefit from reduced charge factor (1.2 versus 2.0)

```
Example single job submission script:
  #!/bin/bash -l
  #SBATCH -p regular
  #SBATCH -N 684
  #SBATCH -A m2405
  #SBATCH -t 01:45:00
  #SBATCH -j tau00001_s7800_R
  cd $SLURM_SUBMIT_DIR
```

```
srun -n 16416 ./distress y0=1d-15 tau1cut=0.00001 seed=7800
outfile='tau00001_s7800_R.dat'
```

#### Validation of results

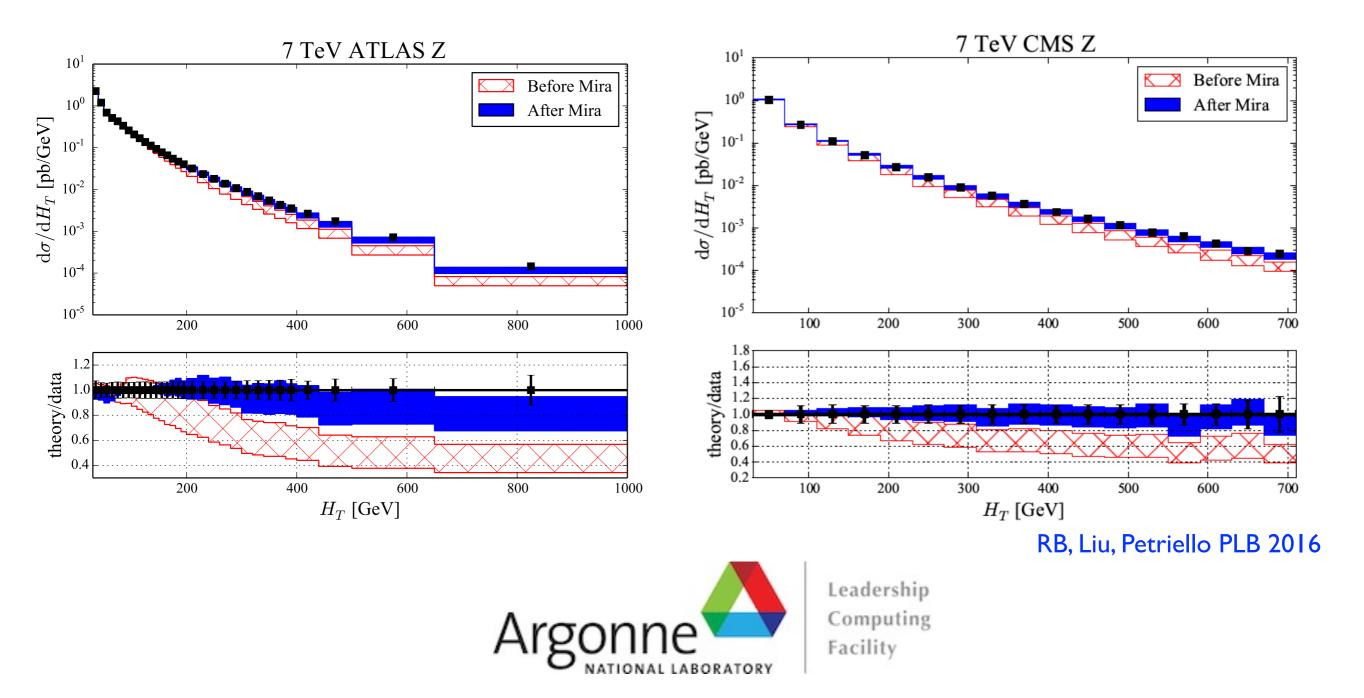
• A powerful check of the formalism is the independence of the final result from  $\tau_N^{cut}$ , which checks the implementation of almost all parts of the calculation



- solid lines: inclusive structure functions
- Obtained from DISTRESS after integration over phase space, also known analytically
- Result taken from circled region, where result becomes independent of the cut

### Scientific highlights

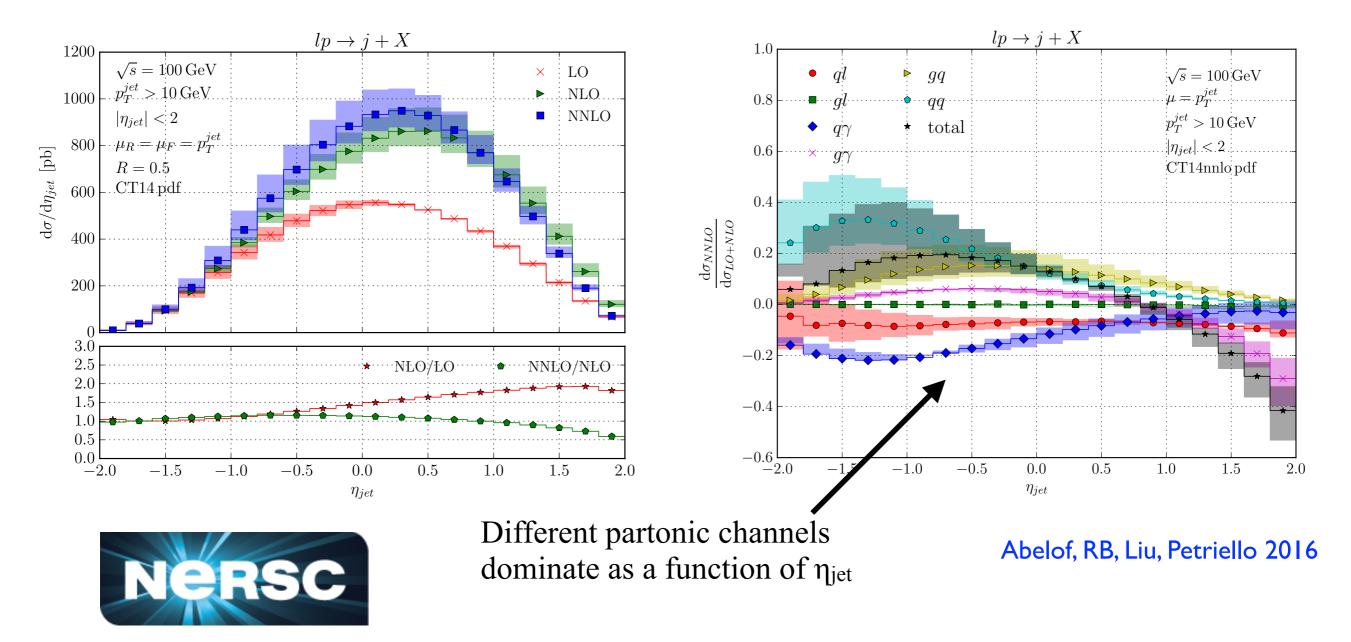
• Runs on Mira for the first time satisfactorily describe data for Z+jet coming from ATLAS and CMS



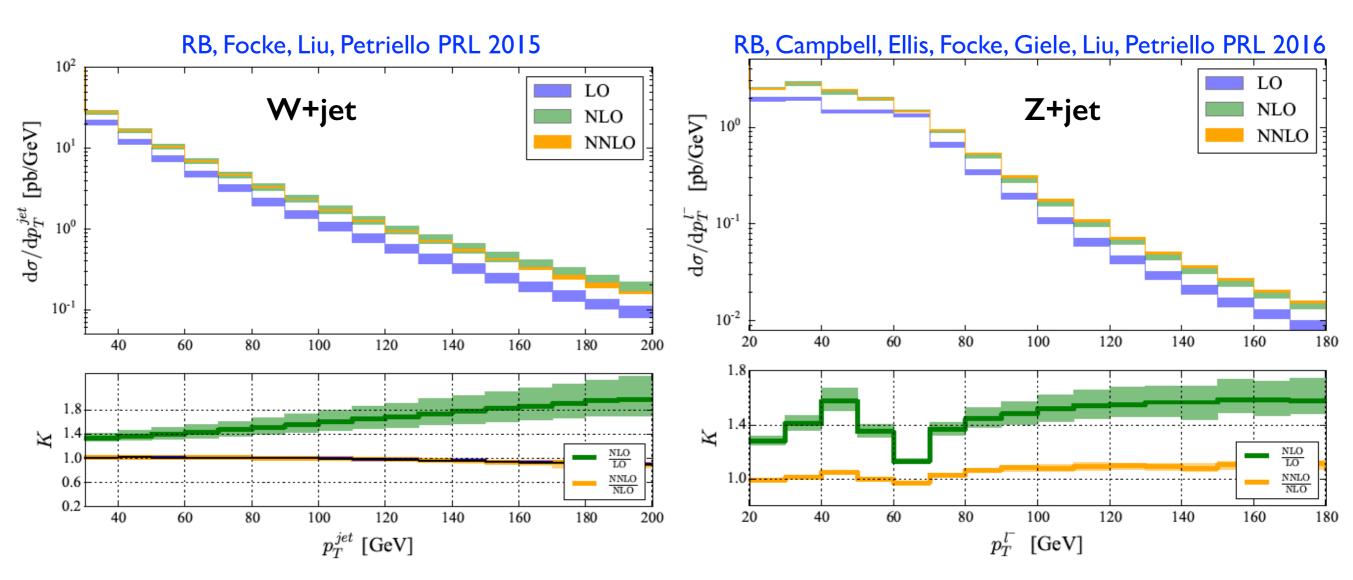
### Scientific highlights

• Large corrections observed for jet production at an Electron Ion collider tamed at NNLO

• **DISTRESS** allows for a detailed investigation of how the different parts of the corrections combine to give the full result



## Scientific highlights



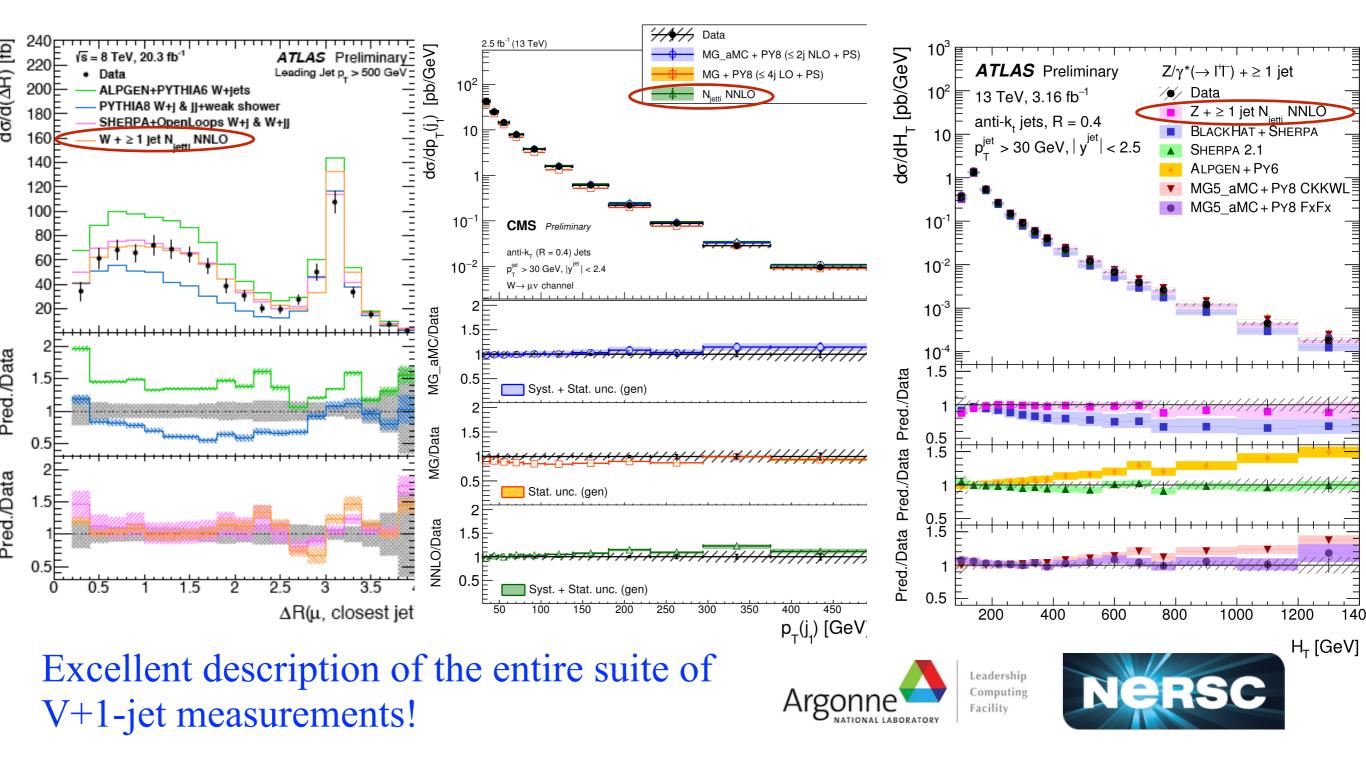
• First full calculations for W+jet@NNLO and Z+jet@NNLO; dramatic reduction of theory uncertainty upon inclusion of NNLO corrections





#### Impact on the experimental program

• Both ATLAS and CMS have incorporated these calculations into their Run I and Run II analyses; a sampling of results is shown below



#### To-do-list

- Theoretical:
  - Theoretical refinements that will reduce  $\tau_N^{cut}$  dependence
- Computational:
  - Everything compiled using GNU; can we use a different compiler?
  - Public release of V+jet, Higgs+jet codes
  - Public release of DISTRESS