

Towards causal relations in noncommutative spacetime

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Spacetime — the arena of classical and quantum physics

A *spacetime* \mathcal{M} is a smooth manifold, which is

- 4-dimensional,
- connected,
- Lorentzian, i.e. it is endowed with a map g

$$\mathcal{M} \ni p \mapsto g_p : T_p\mathcal{M} \times T_p\mathcal{M} \rightarrow \mathbb{R}$$

such that

- $\forall p \in \mathcal{M}$ g_p is a symmetric bilinear map of signature $(-+++)$,
- $\forall X, Y \in \mathcal{X}(\mathcal{M})$ the map $p \rightarrow g_p(X_p, Y_p)$ is smooth.
- time-oriented (will explain in a second)

Points of \mathcal{M} are suggestively called **events**.

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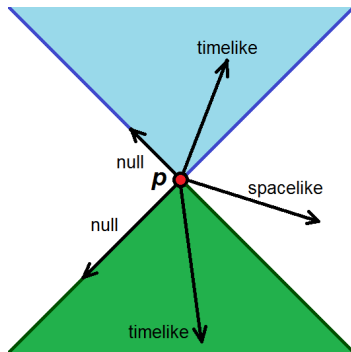
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Elements of causality theory

Classification of tangent vectors

Vector $v \in T_p\mathcal{M}$ is called

- *timelike* if $g_p(v, v) < 0$,
- *null* if $g_p(v, v) = 0$,
- *spacelike* if $g_p(v, v) > 0$,
- *causal* if $g_p(v, v) \leq 0$ and $v \neq 0$,



The set of timelike vectors tangent at p has two components. A *time-orientation* is a *continuous* choice of one of the components, containing (by definition) *future-directed timelike vectors*.

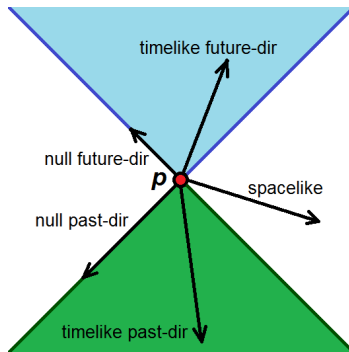
Above terms extend naturally to *vector fields* and to *piecewise C^1 curves*.

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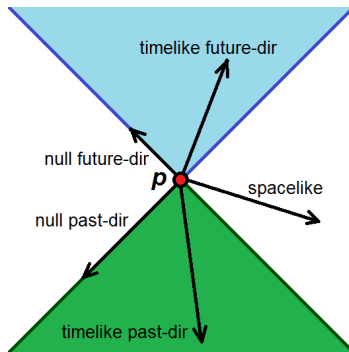
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The set of **causal** vectors tangent at p has two components. A *time-orientation* is a *continuous* choice of one of the components, containing (by definition) *future-directed causal* vectors.

Above terms extend naturally to *vector fields* and to *piecewise C^1 curves*.

Causal precedence relation \preceq

We say that $p \in \mathcal{M}$ **causally precedes** $q \in \mathcal{M}$, denoted $p \preceq q$, if there exists a (piecewise C^1) future-directed causal curve $\gamma : [0, 1] \rightarrow \mathcal{M}$ such that $\gamma(0) = p$, $\gamma(1) = q$ or if $p = q$.

Relation \preceq is a *preorder* (reflexive + transitive).

Causal future and past

The **causal future** and the **causal past** of an event $p \in \mathcal{M}$ are the sets

$$J^+(p) := \{q \in \mathcal{M} : p \preceq q\} \quad \text{and} \quad J^-(p) := \{q \in \mathcal{M} : q \preceq p\}.$$

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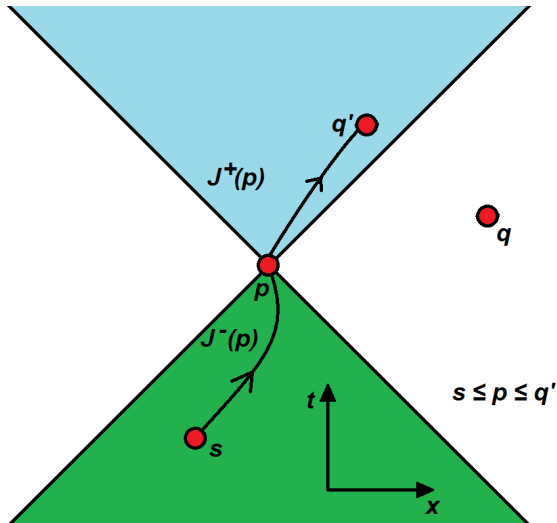
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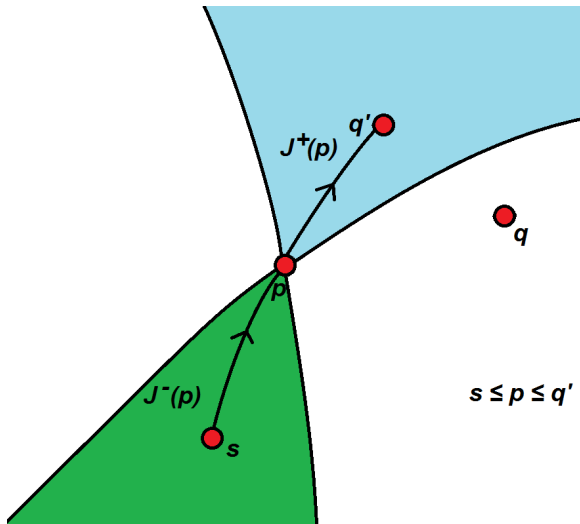
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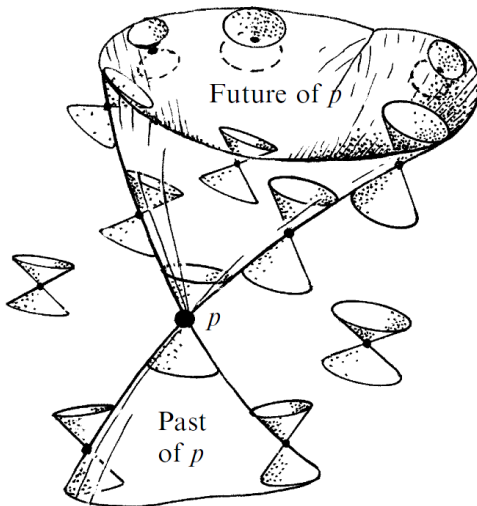
Causal structure of a spacetime in SR (Minkowski spacetime)

Elements of causality theory



Causal structure of a spacetime in GR

Elements of causality theory



Causal structure of a spacetime in GR (from Penrose's "Road to Reality")

Motivations and the goal

- Causality (in the context of Lorentzian geometry)
= certain binary relation between **events**.
- Pointlike events are **not** directly observable!
 - Measuring apparatus' imperfection
 - Wave-packets have nonzero spread
- Goal: extend the causality relation to states on a suitable (pre-) C^* -algebra \mathcal{A}
 - ↪ Besnard, Franco, Eckstein
- In this talk: $\mathcal{A} = C_0(\mathcal{M})$, and so
 - Mixed states are of the form $f \mapsto \int_{\mathcal{M}} f d\mu$ for some *Borel probability measure* μ .
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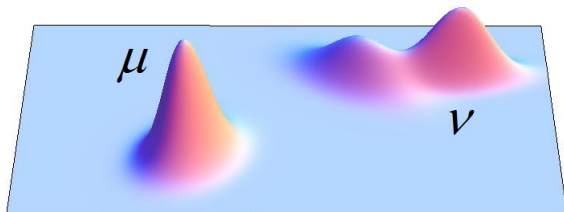
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The **causal precedence** relation \preceq on \mathcal{M}

$$p \preceq q \Leftrightarrow \exists \text{ future-directed causal curve } \gamma \text{ from } p \text{ to } q, \text{ or } p = q$$

- **Goal:** Define $\mu \preceq \nu$ for $\mu, \nu \in \mathcal{P}(\mathcal{M})$.



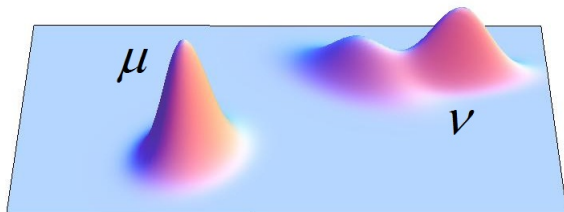
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Causality for probability measures

Causal functions

Function $f : \mathcal{M} \rightarrow \mathbb{R}$ is called **causal**, if it is nondecreasing along every future-directed causal curve.

What does it mean that $\mu \preceq \nu$ for \mathcal{M} globally hyperbolic?

For any $\mu, \nu \in \mathcal{P}(\mathcal{M})$:

$$\mu \preceq \nu \iff \int_{\mathcal{M}} f d\mu \leq \int_{\mathcal{M}} f d\nu \text{ for all causal functions } f.$$

Theorem [Minguzzi, Besnard, Franco]

For any $p, q \in \mathcal{M}$ we have that $p \preceq q$ iff $\delta_p \preceq \delta_q$ in the above sense.

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Better definition [M. Eckstein, TM '17]

Let \mathcal{M} be a spacetime. Then for any $\mu, \nu \in \mathcal{P}(\mathcal{M})$

$$\mu \preceq \nu \stackrel{\text{def}}{\iff} \exists \omega \in \mathcal{P}(\mathcal{M}^2) \text{ such that:}$$

- $\forall A \subseteq \mathcal{M} \quad \omega(A \times \mathcal{M}) = \mu(A), \quad \omega(\mathcal{M} \times A) = \nu(A),$
- $\omega(J^+) = 1,$

where $J^+ := \{(p, q) \in \mathcal{M}^2 \mid p \preceq q\}$.

- ω can be called a causal coupling or a causal transference plan.
- For $\mu = \delta_p, \nu = \delta_q$, the only coupling is $\omega = \delta_{(p,q)}$ and so $\delta_p \preceq \delta_q$ iff $p \preceq q$.

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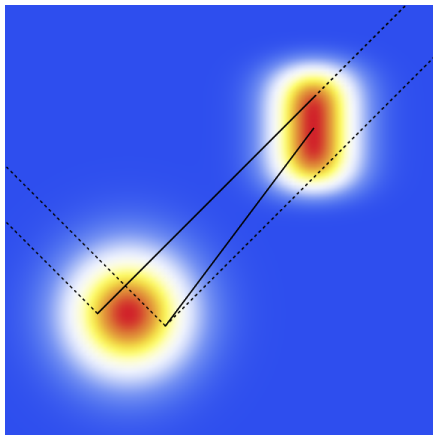
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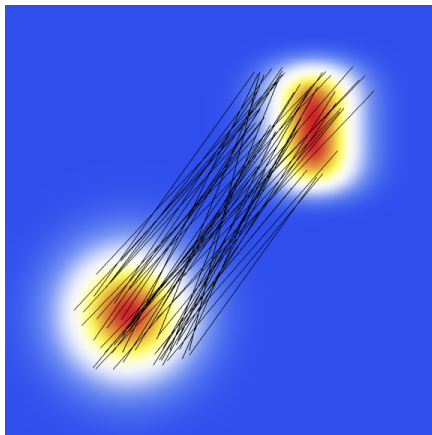
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Each infinitesimal part of the probability measure should travel along a future-directed causal curve.



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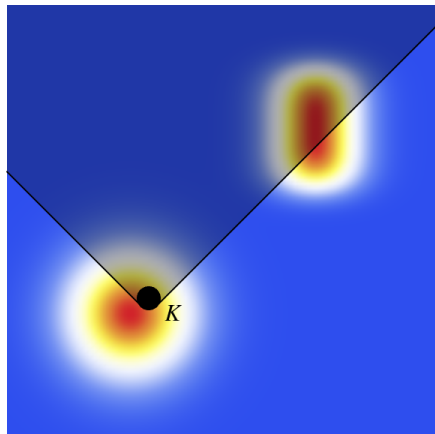
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Causality for probability measures

For \mathcal{M} causal and such that J^+ is closed topologically:

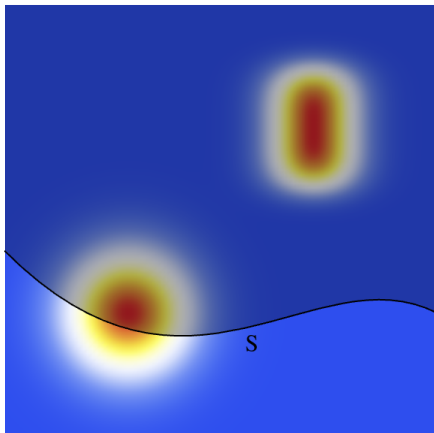
$$\mu \preceq \nu \iff \forall \text{compact } \mathcal{K} \subseteq \text{supp } \mu \quad \mu(\mathcal{K}) \leq \nu(J^+(\mathcal{K}))$$



Causality for probability measures

For \mathcal{M} globally hyperbolic (i.e. admitting Cauchy hypersurfaces):

$$\mu \preceq \nu \iff \forall \text{ Cauchy hypersurface } \mathcal{S} \quad \mu(J^+(\mathcal{S})) \leq \nu(J^+(\mathcal{S}))$$



Causal time-evolution of measures ($\mathcal{M} := \text{Minkowski}$)

Causal time-evolution of a pointlike particle

A curve $\gamma : I \rightarrow \mathcal{M}$, $\gamma(t) = (t, x(t))$ is a worldline of a physical particle if

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A map $\mu : I \rightarrow \mathcal{P}(\mathcal{M})$, $t \mapsto \mu_t$ such that $\text{supp } \mu_t \subseteq \{t\} \times \mathbb{R}^3$ for all $t \in I$ is a **causal evolution of a measure** if

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Causal time-evolution of measures ($\mathcal{M} := \text{Minkowski}$)

The Schrödinger equation

$$i\hbar\partial_t\psi = \hat{H}\psi$$

yields the $\mathcal{P}(\mathcal{M})$ -valued map

$$d\mu_t = \delta_t \times |\psi(t, x)|^2 d^3x.$$

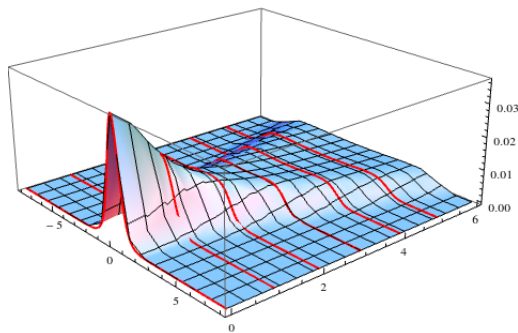
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Causal time-evolution of measures ($\mathcal{M} := \text{Minkowski}$)

Theorem [M. Eckstein, TM '17]

Suppose $\rho(t, x)$ satisfies the *continuity equation*

$$\partial_t \rho + \nabla \cdot \rho \mathbf{v} = 0$$

with a velocity field such that $\|\mathbf{v}(t, x)\| \leq c$. Then μ_t defined via

$$d\mu_t = \delta_t \times \rho(t, x) d^3x$$

evolves causally.

- Dirac's Hamiltonian respects causality.
- Same for Białyński-Birula's "Dirac-like" Hamiltonian of the wave function of a photon.
- However, the $\sqrt{\hat{p}^2 + m^2}$ Hamiltonian violates causality!

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with a velocity field such that $\|\mathbf{v}(t, x)\| \leq c$. Then μ_t defined via

$$d\mu_t = \delta_t \times \rho(t, x) d^3x$$

evolves causally.

- Dirac's Hamiltonian respects causality.
- Same for Białyński-Birula's "Dirac-like" Hamiltonian of the wave function of a photon.
- However, the $\sqrt{\hat{p}^2 + m^2}$ Hamiltonian violates causality!

Causal time-evolution of measures ($\mathcal{M} := \text{Minkowski}$)

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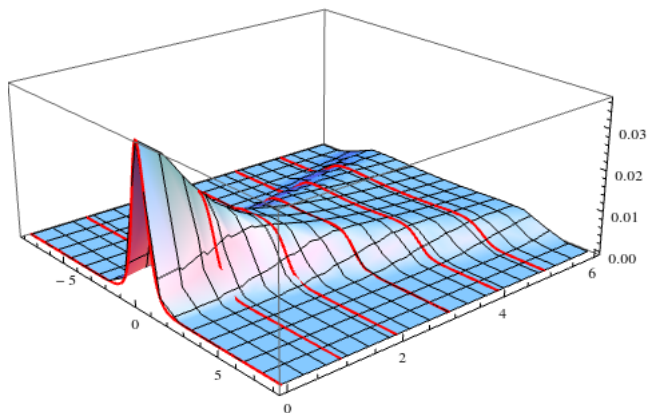
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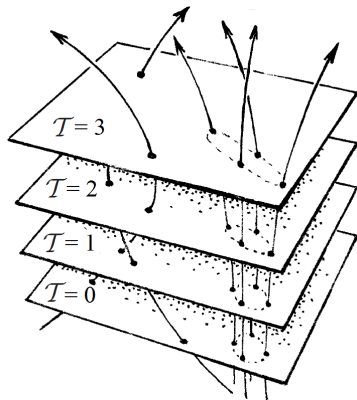
What about other observers?

Causal time-evolution of measures ($\mathcal{M} := \text{Minkowski}$)

Theorem [TM '17]

Consider a map $t \mapsto \mu_t \in \mathcal{P}(\mathcal{M})$ satisfying $\text{supp } \mu_t \subseteq \{t\} \times \mathbb{R}^3$ for every $t \in I$. TFAE:

- The map $t \mapsto \mu_t$ is *causal*, i.e.
 $\forall s, t \in I \quad s \leq t \Rightarrow \mu_s \preceq \mu_t$.
- There exists a **probability measure on the space of worldlines**, from which one can recover μ_t for all $t \in I$.



Adapted from Penrose's "Road to Reality"

- Any spacetime \mathcal{M} possesses an inherent **causal structure**, encoded by the binary relation \preceq .
- This relation can be **naturally** extended onto $\mathcal{P}(\mathcal{M})$, encapsulating the requirement that probability can only flow subluminally.
- **Claim:** Probability distributions obtained from QM wave packets should respect this causal relation.
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Conclusions






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Thank you for your attention!

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