Observing superradiance in a vortex flow

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<u>Rotational superradiant scattering in a vortex flow</u>, T.T. *et al* Nature Physics (June 2017), doi : 10.1038/nphys4151

Use a hydrodynamic system to mimic gravity effect

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<u>Good review</u> Barcelo *et al,* arXiv : gr-qc 0505065

- BEC
- Fluids of Light
- Water

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Analogue gravity

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Black holes



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Black holes



How do black holes loose their energy ?



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- Fluids of Light
- Water



- Zel'Dovich (1971)
- Rotating Black Holes
- Many other systems

<u>Review :</u> Brito *et al,* Superradiance. Lect. Notes Phys., 906:pp.1-237,2015.

Rotating Black Holes

Black Hole Regions Ergosphere Horizon Gravitational Spacetime Quiet region: Distortion negligible gravitational influence Singularity



During 70's : Bekenstein and Hawking Black Hole Thermodynamics



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$$S_{BH} = \frac{k_B A}{4 {l_P}^2}$$
 First law : $\delta M = \frac{\kappa}{8\pi} \delta A + \Omega_H \delta J$
Second law : $\delta A \ge 0$



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Wave scattering : m, ω



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$$\frac{\delta J}{\delta M} = \frac{m}{\omega}$$

$$\delta M\left(\frac{\omega - m\Omega_H}{\omega}\right) = \frac{\kappa}{8\pi}\delta A$$



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Credits : W. Unruh



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When $c_{Flow} < c_{wave}$: can propagate in all direction



<u>Credits</u> : V	V. Unruh				
When	C _{Flow}	<	C _{wave}	:	can propagate in all direction
When	C _{Flow}	≥	C _{wave}	: A	can't propagate in all direction nalogue horizon



Mathematically equivalent !

R.Schützhold and W.G. Unruh. Gravity wave analogues of black holes. PRD, 66:044019, Aug. 2012



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Navier-Stokes + Continuity + Boundary conditions



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Flow : Inviscid fluid, Incompressible, Irrotational Waves : Small amplitude, Shallow water waves



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Scalar field propagating on an effective metric $g^{\mu\nu} = \begin{pmatrix} 1 & \overrightarrow{v_B} \\ \overrightarrow{v_B}^T & \overrightarrow{v_B} \times \overrightarrow{v_B} - gh_B.Id \end{pmatrix}$

Example in Samaria Gorges



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Try to propagate against the flow (blue shift)

Analogue horizon












Irrotational

Axisymmetric Stationary



Irrotational

Axisymmetric



$\begin{array}{c} -5 \\ (fl) \\ (fl$

Irrotational

Axisymmetric



Irrotational

Axisymmetric







Irrotational

Axisymmetric







Irrotational

Axisymmetric







Condition of superradiance :
$$\omega - m\Omega < 0$$

Azimuthal number

Condition of superradiance : $\omega - m\Omega < 0$

We want to excite azimuthal waves



Condition of superradiance : $\omega - m\Omega < 0$



Condition of superradiance : $\omega - m\Omega < 0$

$$e^{i\vec{k}.\vec{r}} = \sum_{m=-\infty}^{\infty} i^m J_m(kr) e^{im\theta}$$

Condition of superradiance : $\omega - m\Omega < 0$

 $m = -\infty$

Compare in/out part of the different m's

Shallow water regime : $\lambda \gg h_B$

<u>Shallow water regime</u> : $\lambda \gg h_B$ <u>Small amplitude</u> : $A \ll h_B$

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Pattern projector

Two cameras

Stereo-photogrammetry

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Analogue Black Hole :

- Irrotational
- Stationary
- Axisymmetric

- Shallow water
- Small Amplitude

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<u>Analogue Black Hole :</u>

- Irrotational
- Stationary
- Axisymmetric

<u>Waves</u>:

- Shallow water
- Small Amplitude

Numerical simulation by Dolan, S. & Oliveira, E. PRD87, 124038 (2013).

















Results

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Free surface at various frequency ω (2.87, 3.04, 3.27, 3.45, 3.70, 4.11 Hz)

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Dispersion (T.T. in preparation)

Rotationnal flow (S. Patrick in preparation)

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New ideas on gravity...?

Rotationnal flow (S. Patrick in preparation)

Thank you !

