Pontryagin trace anomaly

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Introduction: quantum anomalies

Property of a classical theory violated by quantization.

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- Harmful anomalies: destroying theories
	- helpful in model building
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	- violate conservation of charges (if such exist)
	- new phenomenology, implications for cosmology
- Some are 1-loop saturated \Rightarrow full perturbative result

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Being genuine quantum properties, unveiling the essence of QFT! ("What is QFT?" is still a tantalising question.)

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From Weyl symmetry to trace anomaly

Symmetry: (local) Weyl rescaling

$$
g_{\mu\nu}(x) \to e^{\omega(x)} g_{\mu\nu}(x) \qquad , \qquad \psi_j(x) \to e^{\lambda_j \omega(x)} \psi_j(x)
$$

Invariance of classical action for matter \Rightarrow traceless EM tensor:

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\delta_{\omega} S_m[\psi, g] = 0 \qquad \Longrightarrow \qquad T_{\mu}^{\mu}(x) = 0
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However, perturbative QFT calculations found anomaly [Capper, Duff 1975]:

$$
\langle \langle T_{\mu}^{\mu} \rangle \rangle = a E_2 + c (W_{\mu \nu \rho \sigma})^2 \qquad , \qquad D = 4
$$

 E_2 is 2nd Euler (Gauss-Bonnet) density, $W_{\mu\nu\rho\sigma}$ is Weyl tensor.

- The form is universal, coefficients a and c theory dependent.
- \bullet coefficients a and c not 1-loop saturated
- At the beginning, community did not believe it!

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$\langle\langle T_{\mu}^{\mu} (x) \rangle\rangle = \mathcal{A}_W (x)$

WZ consistency conditions connect trace and diff- anomalies.

General (nontrivial) solution: no (pure gravity) diff-anomalies and

 $\mathcal{A}_W = a E_2 + c (W_{\mu\nu\rho\sigma})^2 + e P_4$ [Bonora, Pasti, Tonin]

where P_4 is Pontryagin (topological) tensity

$$
P_4 = \frac{1}{32 \pi^2} \, \varepsilon^{\mu \nu \rho \sigma} \, R^{\mu \nu}{}_{\alpha \beta} \, R_{\rho \sigma \alpha \beta}
$$

- Pontryagin anomaly, being P and CP odd, is exceptional,
- \bullet but, with rare exceptions ([Nakayama 2012]), neglected \Rightarrow Why?

Status of 4D Pontryagin anomaly before 2014

All against it:

- No CP violating free relativistic QFT's
- For Weyl fermion: in 4D classical 1-1 mapping with massless Majorana fermion, and P-odd anomaly not expected for Majorana
	- No Pontryagin anomaly for Majorana fermion [Bastianelli 2016]
	- Careful: mapping is not linear unitary!
- Realisations having unitarity problems [Nakayama 2012]
- Non-compatibility with SUSY [Bonora, Giaccari 2013]

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As a consequence, Pontryagin anomaly mainly neglected by the community!

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Status of 4D Pontryagin anomaly in 2016

For free Weyl fermion in curved spacetime anomaly was found, with

$$
e_W = \pm \frac{i}{48}
$$
 [Bonora, Giaccari, Lima de Souza 2014]

Independently obtained with heat kernel method and perturbatively (1-loop) using dimensional regularisation.

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- No anomaly for Majorana and Dirac fermions.
- Note the strange-looking imaginary coefficient.
- Calculation somewhat incomplete:
	- Heat kernel calculation naive extrapolation of real fermion one
	- Simplifying field redefinition used, dim. regularisation subtleties

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Result was received (mainly) with disbelief!

Main part of the talk

Due to a controversial character of the result, we are in need of more complete, and also independent calculations.

We focused on perturbative 1-loop calculation using dimensional regularisation, and did it in two ways:

1 Repeat [Bonora at al.] calculation in more pedantic way

- without using field redefinitions
- ² Working with Dirac fermion by introducing chiral metric (MAT gravity)
	- analogue of Bardeen method used for chiral gauge anomalies

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[Bonora, Cvitan, D.P., D.Pereira, Giaccari, Štemberga, Eur. Phys. J. C 77 (2017) 511]

I. Calculation with free Weyl fermion: the theory

Action:
$$
S_L = \int d^4x \sqrt{|g|} i \overline{\psi_L} \gamma^{\mu} \left(\nabla_{\mu} + \frac{\omega_{\mu}}{2}\right) \psi_L
$$

$$
\text{EM tensor:} \qquad \qquad T^{\mu\nu} = -\frac{i}{4} \overline{\psi_L} \gamma^{\mu} \overleftrightarrow{\nabla}^{\nu} \psi_L + (\mu \leftrightarrow \nu)
$$

Perturbative expansion in $h_{\mu\nu}=g_{\mu\nu}-\eta_{\mu\nu}$ up to h^3 :

$$
S_{int} = \int d^{4}x \left[-\frac{i}{4} \overline{\psi_{L}} h_{a}^{m} \gamma^{a} \overleftrightarrow{\partial}_{m} \psi_{L} + \frac{3i}{16} \overline{\psi_{L}} (h^{2})_{a}^{m} \gamma^{a} \overleftrightarrow{\partial}_{m} \psi_{L} - \frac{5i}{32} \overline{\psi_{L}} (h^{3})_{a}^{m} \gamma^{a} \overleftrightarrow{\partial}_{m} \psi_{L} \right. \\ - \frac{1}{16} \epsilon^{mabc} \overline{\psi_{L}} \gamma_{c} \gamma_{5} \psi_{L} \left(h_{m}^{\sigma} \partial_{a} h_{b\sigma} + (h^{2})_{m}^{\sigma} \partial_{b} h_{a\sigma} - h_{m}^{\rho} h_{a}^{\sigma} \partial_{\sigma} h_{\rho b} - \frac{1}{2} h_{m}^{\rho} \partial_{a} h_{\rho \sigma} h_{c}^{\sigma} \right) \\ + \frac{1}{2} (\text{tr} h) \left(\frac{i}{2} \overline{\psi_{L}} \gamma^{m} \overleftrightarrow{\partial}_{m} \psi_{L} - \frac{i}{4} \overline{\psi_{L}} h_{a}^{m} \gamma^{a} \overleftrightarrow{\partial}_{m} \psi_{L} + \frac{3i}{16} \overline{\psi_{L}} (h^{2})_{a}^{m} \gamma^{a} \overleftrightarrow{\partial}_{m} \psi_{L} - \frac{1}{16} \epsilon^{mabc} \overline{\psi_{L}} \gamma_{c} \gamma_{5} \psi_{L} h_{m}^{\sigma} \partial_{a} h_{b\sigma} \right) \\ + \left(\frac{1}{8} (\text{tr} h)^{2} - \frac{1}{4} (\text{tr} h^{2}) \right) \left(\frac{i}{2} \overline{\psi_{L}} \gamma^{m} \overleftrightarrow{\partial}_{m} \psi_{L} - \frac{i}{4} \overline{\psi_{L}} h_{a}^{m} \gamma^{a} \overleftrightarrow{\partial}_{m} \psi_{L} \right) \\ + \left(-\frac{1}{8} (\text{tr} h) (\text{tr} h^{2}) + \frac{1}{48} (\text{tr} h)^{3} + \frac{1}{6} (\text{tr} h^{3}) \right) \frac{i}{2} \overline{\psi_{L}} \gamma^{m} \overleftrightarrow{\partial}_{m} \psi_{L} + \dots \right] \\ + \frac{1}{2} \gamma_{c} \gamma_{c} \gamma_{c} \gamma_{c} \
$$

I. Calculation with free Weyl fermion: Weyl anomaly

Full perturbative calculation up to 3-point functions:

• no assumptions, no field redefinition

$$
\langle \langle T_{\mu}^{\mu} \rangle \rangle_{\text{P-odd}} = \frac{\mp i}{768 \pi^2} \varepsilon^{\mu\nu\lambda\rho} \left(\partial_{\mu} \partial_{\sigma} h_{\nu}^{\alpha} \partial_{\lambda} \partial_{\alpha} h_{\rho}^{\sigma} - \partial_{\mu} \partial_{\sigma} h_{\nu}^{\alpha} \partial_{\lambda} \partial^{\sigma} h_{\alpha_{\rho}} \right) + O(h^3)
$$

Covariantisation gives:

$$
\langle\langle T_{\mu}^{\mu}\rangle\rangle_{\text{P-odd}} = \mp \frac{1}{16} P_4
$$

Opposite sign and three times larger then in [Bonora et al. 2014] Inconsistency? Not yet. First one has to check diff-covariance.

I. Calculation with Weyl fermion: Diff-anomaly and result

We have to check the conservation of EM tensor

 $\nabla_{\mu}\left\langle \left\langle T^{\mu\nu}\right\rangle \right\rangle _{\textrm{P-odd}}=\mathcal{A}_{\textrm{diff}}$

We obtain $A_{\text{diff}} \neq 0$, which can be canceled by the local counterterm:

$$
{\cal C}=\frac{1}{6}\int\mathrm{d}^4x\,h^{\mu}_{\mu}\,{\cal A}^{(0)}_W
$$

But, this counterterm adds to trace anomaly $-\frac{4}{3}\mathcal{A}_{W}^{(0)}$, so the final result is:

$$
\langle\langle T_{\mu}^{\mu}\rangle\rangle_{\rm P-odd}=\pm\frac{1}{48}\,P_4\qquad ,\qquad \nabla_{\mu}\,\langle\langle\,T^{\mu\nu}\rangle\rangle_{\rm P-odd}=0
$$

In full agreement with [Bonora et al. 2014, 2015]

II. Calculation with Dirac fermion. MAT gravity

Avoid subtleties with Weyl fermions by using Dirac field.

For this we have to invent chiral coupling to gravity \Rightarrow MAT gravity Metric-axial-tensor (MAT) gravity based on chiral metric:

 $G_{\mu\nu} = g_{\mu\nu} + \gamma_5 f_{\mu\nu}$, $f_{\mu\nu}$ is additional spin-2 field

One defines chiral vierbein, connections, Riemann tensor, . . . in the usual way from $G_{\mu\nu}$.

Two diff-symmetries and two Weyl-symmetries \Rightarrow two conserved traceless spin-2 tensors.

(MAT gravity can be viewed as a bimetric gravity with $g^{(\pm)}_{\mu\nu}=g_{\mu\nu}\pm f_{\mu\nu}$.)

II. Calculation with Dirac fermion. The theory

We put Dirac fermion in MAT gravity background

$$
S_D=\int\mathrm{d}^4x\,i\overline{\psi}\sqrt{|\bar{\mathsf{G}}|}\gamma^a\hat{E}^{\mu}_a\left(\partial_{\mu}+\frac{1}{2}\Omega_{\mu}\right)\psi
$$

We obtained (dim. regularisation) for the trace anomalies

$$
\langle\langle T_{\mu}^{\mu}\rangle\rangle = \frac{i}{768 \pi^2} \epsilon^{\mu\nu\lambda\rho} R_{\mu\nu}^{(1)\sigma\tau} R_{\lambda\rho\sigma\tau}^{(2)}
$$

$$
\langle\langle T_{5\mu}^{\mu}\rangle\rangle = \frac{i}{768 \pi^2} \frac{1}{2} \epsilon^{\mu\nu\lambda\rho} \left(R_{\mu\nu}^{(1)\sigma\tau} R_{\lambda\rho\sigma\tau}^{(1)} + R_{\mu\nu}^{(2)\sigma\tau} R_{\lambda\rho\sigma\tau}^{(2)} \right)
$$

where $R^{(1,2)}_{\mu\nu\lambda\rho} = \frac{1}{2} \left(R^{(+)}_{\mu\nu\lambda\rho} \pm R^{(-)}_{\mu\nu\lambda\rho} \right)$

We are interested in the blue one, which is the standard trace anomaly.

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(a) If we take
$$
g_{\mu\nu}^{(-)} = \eta_{\mu\nu} \Rightarrow \omega^{(-)a}{}_{b\mu} = 0 = \Gamma^{(-)\mu}{}_{\nu\rho}
$$

- Dirac MAT action \Rightarrow action in flat space for Ψ_R plus action in curved metric $g^{(+)}$ for ψ_L
- trace anomaly belongs to the left Weyl fermion, and is

$$
\langle \langle T_{\mu}^{\mu} \rangle \rangle = \frac{i}{768 \pi^2} \frac{1}{2} \epsilon^{\mu \nu \lambda \rho} R_{\mu \nu}^{(+)} \sigma^{\tau} R_{\lambda \rho \sigma \tau}^{(+)}
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(b) If we take $g_{\mu\nu}^{(+)} = \eta_{\mu\nu}$ we obtain anomaly for the right Weyl fermion $\langle \langle T_{\mu}^{\mu} \rangle \rangle = -\frac{i}{768}$ $768\,\pi^2$ 1 $\frac{1}{2} \epsilon^{\mu\nu\lambda\rho} R^{(-)\sigma\tau}_{\mu\nu} R^{(-)}_{\lambda\rho\sigma}$ λρστ

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(c) If we take $g^{(+)}_{\mu\nu} = g^{(-)}_{\mu\nu} = g_{\mu\nu}$ we obtain a standard Dirac field in curved space, and

 $\langle\langle T_{\mu}^{\mu}\rangle\rangle=0$

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Confirms previously obtained results

We have extended evidence that Weyl fermion in 4D has Pontryagin trace anomaly.

If true, this has important and strange consequences:

- Gravitational perturbative CP violation
- Weyl and massless Majorana not equivalent in curved spacetime.
- Strange imaginary coefficient, is unitarity violated?
- **•** If yes, anomaly is harmful \Rightarrow only Dirac and Majorana fermions \Rightarrow minimal SM with Weyl fermions not consistent
- **•** If harmless, leads to CP-violating imaginary term in effective action
	- CP-violating gravitational Schwinger mechanism?
	- No effect on 0th-order cosmology, could affect perturbations (GWs) [Mauro, Shapiro 2015]

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• could be important for black holes

Outlook

Assuring the anomaly is not a quirk of dimensional regularisation:

- **•** Perturbative calculation using other schemes
	- Pauli-Villars in MAT-gravity
- Point-splitting and heat kernel
- If sticks, understand its nature and further consequences.
- If bogus, understand what happened with dimensional regularisation.

Result looks really strange, but we've seen stranger stuff in QM.