

# Pedal coordinates, Dark Kepler and other force problems

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- Bertrand theorem (1873) : *No other central force problem has the property that all bounded trajectories are also closed.*

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**(Mahomed and Vawda in 2000)**

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$$\varphi \rightarrow \varphi + \omega \int r d\varphi.$$



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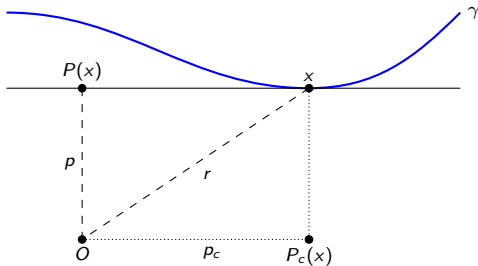
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If not the transform  $E_\alpha^* B_\alpha$  can be used to convert the problem into the previous case.

# Pedal coordinates



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$$\kappa := \frac{r^2 + 2r'_\varphi{}^2 - rr''_\varphi}{(r^2 + r'_\varphi{}^2)^{\frac{3}{2}}}, \quad \Rightarrow \quad \kappa = \frac{1}{rr'}.$$

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$$C := ES \xrightarrow{\frac{1}{2}} P$$

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## Example

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## Kepler problem in General relativity

$$r'_{\varphi}{}^2 = \frac{r^4}{b^2} - \left(1 - \frac{r_s}{r}\right) \left(\frac{r^4}{a^2} + a^2\right), \quad (4)$$

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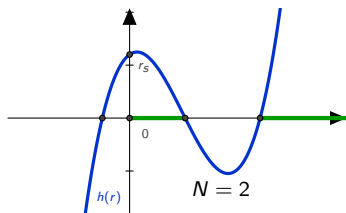
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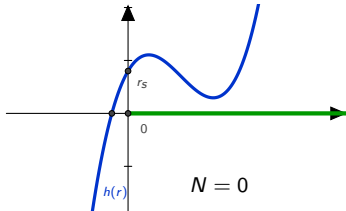
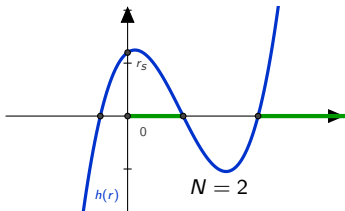
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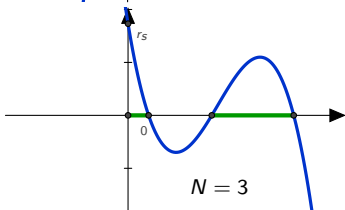
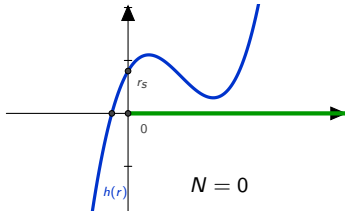
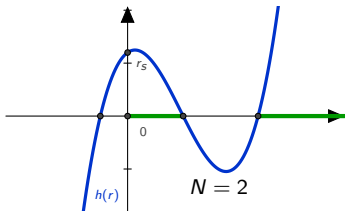
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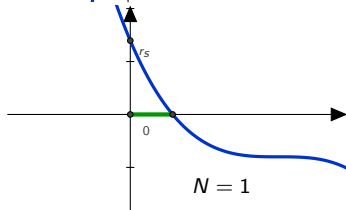
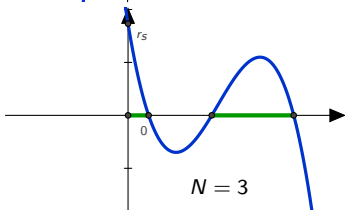
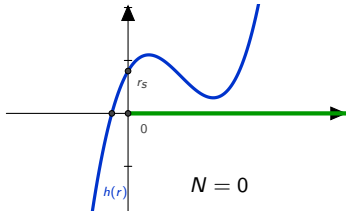
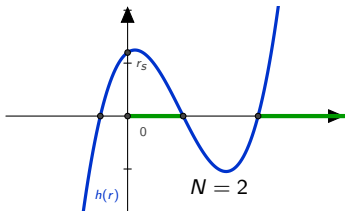
or

$$0 \leq dr^3 + \frac{r_s}{a^2} r^2 - r + r_s =: h(r).$$









P. Blaschke: *Pedal coordinates, Dark Kepler and other force problems*, Journal of Mathematical Physics, DOI: 10.1063/1.4984905

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