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## **$7\text{Be}(n,a)$ AND $7\text{Be}(n,p)$ CROSS-SECTION SECTION MEASUREMENT FOR THE COSMOLOGICAL LITHIUM PROBLEM AT THE $n_{\text{-TOF}}$ FACILITY AT CERN**

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One of the most important unresolved problems in Nuclear Astrophysics is the so-called “Cosmological Lithium problem”(CLiP). It refers to the large discrepancy (factor 3-5) between the abundance of primordial  $7\text{Li}$  predicted by the standard theory of Big Bang Nucleosynthesis (BBN) and the value inferred from the so-called “Spite plateau” in halo stars.

In the framework of Standard Model, a possible explanation for this longstanding puzzle is related to the incorrect estimation of the destruction rate of  $7\text{Be}$ . Indeed in the standard theory of BBN, 95% of primordial  $7\text{Li}$  is produced by the decay of  $7\text{Be}$  ( $t_{1/2}=53.2$  days), relatively late after the Big Bang, when lower temperature of Universe allows electrons and nuclei to combine into atoms. Therefore, the abundance of  $7\text{Li}$  is essentially determined by the production and destruction of  $7\text{Be}$ .

While charged-particle induced reactions responsible for the destruction of  $7\text{Be}$  have mostly been ruled out by recent measurements, data on the  $7\text{Be}(n,a)$  and  $7\text{Be}(n,p)$  reactions were so far scarce or completely missing, mainly due to experimental difficulties arising from  $7\text{Be}$  specific activity.

Recently, both reaction cross-sections have been measured at  $n_{\text{-TOF}}$  (CERN) taking advantage also of state-of-art techniques for the production of high-purity radioactive samples at ISOLDE, of high performance detection systems and, especially, of the innovative features of the new measuring station  $n_{\text{-TOF-EAR2}}$  particularly suited for challenging measurements on short-lived radioisotopes.

The two measurements provide for the first time nuclear data on  $7\text{Be}(n,a)$  and  $7\text{Be}(n,p)$  cross-section in a wide neutron energy range, namely in the energy range of interest for Nuclear Astrophysics.

**Primary author:** BARBAGALLO, Massimo (Universita e INFN, Bari (IT))

**Presenter:** BARBAGALLO, Massimo (Universita e INFN, Bari (IT))

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