The growth with energy of vector meson photo-production cross-sections and low x evolution

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based on:

I. Bautista, A. Ferndandez Tellez, MH.; [arXiv:1607.05203] (PRD **94** 054002)



Outline

Introduction

Ingredients of our study NLO BFKL gluon density Impact factor $\gamma \rightarrow V$ Amplitude: Real part from imaginary part

Results & Conclusions

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photo-production of J/Ψ and Υ : explore proton at ultra-small x



- measured at HERA (ep) and LHC (pp, ultra-peripheral pPb)
- exclusive process, but allows to relate to inclusive gluon

reach values down to $x=4\times 10^{-6} \rightarrow$ (unique ?) opportunity to explore the low x gluon



- describes data or not \rightarrow re-fit
- if yes: do we really see saturation effects?

i.e. BK type evolution
$$\frac{d}{d\ln 1/x}G(x) = K \otimes G(x) - \underbrace{G \otimes G}_{\text{present, relevant?}}$$

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DGLAP vs. saturation (II)



- $J/\Psi \rightarrow \Upsilon \simeq$ evolution 2.4 GeV² \rightarrow 22.4 GeV²
- high density effects die away in collinear limit
- DGLAP unstable at ultra-small x and small scales ...
- \blacktriangleright convinced: pdf studies highly valuable \rightarrow constrain pdf's at ultra-small x
- useful benchmark for saturation searches (?)

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Will argue in the following ...

- ► a far better dilute (!) benchmark might (?) be given by BFKL evolution (→ applies for UPCs@LHC, might be different for a future (US-)EIC ... → phase space!)
- ▶ why? BFKL \equiv low x evolution without high density/saturation effects
- available up to NLO [Fadin, Lipatov; PLB 429 (1998) 127]; [Ciafaloni, Camici; PLB 430 (1998) 349], resummation schemes for coll. logs exist & to some degree well explored [Salam; hep-ph/9806482], ...
- ▶ not only explored in n = 0 sector → additional constrains from e.g. angular decorrelation studies of jets → see talks on Wednesday

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The framework of this BFKL study

procedure:

a) calculate diff. Xsec. at t = 0

 \rightarrow exclusive scattering amplitude can be expressed through *inclusive* gluon distribution

b) parametrize t dependence
$$\frac{d\sigma(t)}{dt} = \frac{d\sigma(t=0)}{dt} \cdot e^{-|t|B_D(W)}$$
,
slope $B_D(W) = b_0 + 4\alpha' \ln \frac{W}{W_0}$ + fix parameters by (HERA) data (here: values proposed by [Jones, Martin, Ryskin, Teubner; 1307.7099, 1312.6795])



The setup: diff. Xsec. at t = 0

a) imaginary part of scattering amplitude:

- unintegrated gluon density from NLO BFKL fit to combined HERA data [MH, Salas, Sabio Vera; 1209.1353; 1301.5283]
- ▶ impact factor $\gamma \rightarrow J/\Psi, \Upsilon$ from light-front wave function used in dipole model studies

[Kowalski, Motyka, Watt; hep-ph/0606272]

- b) real part:
 - SmA(W², t) dominant, real part can be numerically large
 → recover real part using dispersion relation



The underlying NLO BFKL fit to DIS data



	virt. photon impact factor	$Q_0/{\sf GeV}$	δ	\mathcal{C}	$\Lambda_{\rm QCD}/~{ m GeV}$
fit 1	leading order (LO)	0.28	8.4	1.50	0.21
fit 2	LO with kinematic improvements	0.28	6.5	2.35	0.21

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data: [H1 & ZEUS collab. 0911.0884]

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Solve BFKL equation in conjugate (γ) Mellin space

$$G\left(x,\boldsymbol{k}^{2},M\right) = \frac{1}{\boldsymbol{k}^{2}} \int_{\frac{1}{2}-i\infty}^{\frac{1}{2}+i\infty} \frac{d\gamma}{2\pi i} \hat{g}\left(x,\frac{M^{2}}{Q_{0}^{2}},\frac{\overline{M}^{2}}{M^{2}},\gamma\right) \left(\frac{\boldsymbol{k}^{2}}{Q_{0}^{2}}\right)^{\gamma}$$

re-introduce two scales: hard scale of process (M) and scale of running coupling (\overline{M})

 \hat{g} : operator in γ space!

$$\begin{split} \hat{g}\left(x,\frac{M^2}{Q_0^2},\overline{\frac{M}{M^2}},\gamma\right) &= \frac{\mathcal{C}\cdot\Gamma(\delta-\gamma)}{\pi\Gamma(\delta)} \cdot \left(\frac{1}{x}\right)^{x\left(\gamma,\frac{M^2}{M^2}\right)} \cdot \\ &\left\{1+\frac{\bar{\alpha}_s^2\beta_0\chi_0\left(\gamma\right)}{8N_c}\log\left(\frac{1}{x}\right)\left[-\psi\left(\delta-\gamma\right)+\log\frac{M^2}{Q_0^2}-\partial_\gamma\right]\right\}, \end{split}$$

resummed NLO BFKL eigenvalue with optimal scale setting (\rightarrow modifies $\chi_1(\gamma)$):

$$\chi\left(\gamma, \frac{\overline{M}^2}{M^2}\right) = \bar{\alpha}_s \chi_0\left(\gamma\right) + \bar{\alpha}_s^2 \tilde{\chi}_1\left(\gamma\right) - \frac{1}{2} \bar{\alpha}_s^2 \chi_0'\left(\gamma\right) \chi_0\left(\gamma\right) + \chi_{RG}(\bar{\alpha}_s, \gamma, \tilde{a}, \tilde{b}) - \frac{\bar{\alpha}_s^2 \beta_0}{8N_c} \chi_0(\gamma) \log \frac{\overline{M}^2}{M^2}.$$

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Impact factor from boosted Gaussian wave function

use relation dipole amplitude \leftrightarrow unintegrated gluon *e.g.* [Kutak, Stasto; hep-ph/040811]

$$2\int d^2 \boldsymbol{b} \mathcal{N}(x,r,b) = \frac{4\pi}{N_c} \int \frac{d^2 \boldsymbol{k}}{\boldsymbol{k}^2} \left(1 - e^{i\boldsymbol{k}\cdot\boldsymbol{r}}\right) \alpha_s G(x,\boldsymbol{k}^2)$$

$$\Im \mathfrak{M} \mathcal{A}_{T}^{\gamma^{*}p \to Vp}(W, t=0) = 2 \int d^{2}\boldsymbol{b} \int d^{2}\boldsymbol{r} \int_{0}^{1} \frac{dz}{4\pi} (\Psi_{V}^{*}\Psi)_{T}(r) \cdot \mathcal{N}(x, r, b)$$
$$= \alpha_{s}(\overline{M} \cdot Q_{0}) \int_{\frac{1}{2}-i\infty}^{\frac{1}{2}+i\infty} \frac{d\gamma}{2\pi i} \int_{0}^{1} \frac{dz}{4\pi} \hat{g}\left(x, \frac{M^{2}}{Q_{0}^{2}}, \overline{M^{2}}, Q_{0}, \gamma\right) \cdot \Phi_{V,T}(\gamma, z, M) \cdot \left(\frac{M^{2}}{Q_{0}^{2}}\right)^{\gamma}$$

yields (using boosted Gaussian wave-functions iwth Brodsky-Huang-Lepage prescription)

$$\Phi_{V,T}(\gamma, z, M) = e\hat{e}_f 8\pi^2 \mathcal{N}_T \frac{\Gamma(\gamma)\Gamma(1-\gamma)}{m_f^2} \left(\frac{m_f^2 \mathcal{R}^2}{8z(1-z)}\right)^2 e^{-\frac{m_f^2 \mathcal{R}^2}{8z(1-z)}} e^{\frac{m_f \mathcal{R}^2}{2}} \left(\frac{8z(1-z)}{M^2 \mathcal{R}^2}\right)^{\gamma} \\ \left[U\left(2-\gamma, 1, \frac{m_f^2 \mathcal{R}^2}{8z(1-z)}\right) + [z^2 + (1-z)^2]\frac{(2-\gamma)}{2}U\left(3-\gamma, 2, \frac{m_f^2 \mathcal{R}^2}{8z(1-z)}\right)\right],$$

[U(a, b, z) hypergeometric function of the second kind or Kummer's function]

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Real part from imaginary part using dispersion relation ...

$$\mathcal{A}(W^2, t) = \xi^{(\tau=+)}(\lambda) \cdot \Im \mathsf{m} \mathcal{A}(W^2, t), \qquad \xi^{\tau=+}(\lambda) = \left(i + \tan \frac{\lambda \pi}{2}\right)$$

commonely used: $\lambda = \text{const} \rightarrow \text{constant ratio of real & imaginary part}$

Real part from imaginary part using dispersion relation ...

$$\mathcal{A}(W^2, t) = \xi^{(\tau=+)}(\lambda) \cdot \Im \mathsf{m} \mathcal{A}(W^2, t), \qquad \xi^{\tau=+}(\lambda) = \left(i + \tan\frac{\lambda\pi}{2}\right)$$

commonely used: $\lambda = \text{const} \rightarrow \text{constant ratio of real & imaginary part}$ more precise: reconstruct real part using ω -Mellin transform, $\omega \leftrightarrow W^2$:

$$\mathcal{A}(W^2, t) = \int_{\delta - i\infty}^{\delta + i\infty} \frac{d\omega}{2\pi i} \left(\frac{1}{x}\right)^{\omega} \left(i + \tan\frac{\omega\pi}{2}\right) a(\omega, t), \quad x = \frac{M_V^2}{W^2 - m_p^2}$$

partial wave $a(\omega,t)$ can be fixed from imaginart part

$$\begin{aligned} a(\omega,0) &= \alpha_s \int\limits_{\frac{1}{2}-i\infty}^{\frac{1}{2}+i\infty} \frac{d\gamma}{2\pi i} \left(\frac{M^2}{Q_0^2}\right)^{\gamma} \int\limits_{0}^{1} \frac{dz}{4\pi} \,\Phi_{V,T}(\gamma,z) \frac{\mathcal{C}\cdot\Gamma(\delta-\gamma)}{\pi\Gamma(\delta)} \cdot \left\{\frac{1}{\omega-\chi\left(\gamma,\frac{\overline{M}^2}{M^2}\right)} \right. \\ &+ \frac{\bar{\alpha}_s^2 \beta_0 \chi_0\left(\gamma\right) / (8N_c)}{\left[\omega-\chi\left(\gamma,\frac{\overline{M}^2}{M^2}\right)\right]^2} \left[-\psi\left(\delta-\gamma\right) - \frac{d\ln\left[\Phi_{V,T}(\gamma,z)\right]}{d\gamma}\right] \right\} \end{aligned}$$

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Amplitude: Real part from imaginary part



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Results & Conclusions

comparison to data: Υ production



- ► provide study for two hard scales: photoproduction scale: $M_{pp} = M_V/2$ impact factor motivated: $M_{if}^2 = 8\mathcal{R}_V^{-2}$
- \blacktriangleright fix normalization by low energy H1 data point \rightarrow K-factor; no further adjustments

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Results & Conclusions

comparison to data: J/Ψ production



► NEW (wrt. [Bautista, Fernando Tellez, MH; 1607.05203]): 13 TeV LHCb data

- ▶ fix normalization by low energy ALICE data point → K-factor believe: related to HERA fit (massless, n_f = 4, (C₁/C₂)² = 2.45)
- ▶ often included (not here): GPD motivated factor (" $x' \neq x$ "); known for collinear [Shuvaev, Golec-Biernat, Martin, Ryskin, hep-ph/9902410]
 - \rightarrow to be calculated for k_T factorized BFKL impact factor
 - \sim kinematic improvements for $\gamma \rightarrow V$

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Concluding remarks

BFKL fits [MH, Salas, Sabio Vera; 1301.5283] somehow approach its limits, but so far works very well \rightarrow keep in mind: most simple combination of existing NLO BFKL fit & existing VM impact factor

first suggestion: no need for saturation effects, linear NLO evolution sufficient

why so hard to manifest saturation? two possible reasons:

a) BFKL simply appropiate framework,... saturation effects not (yet) present

b) observable $\sim N^{\text{dipole}} \Leftrightarrow G_{\text{ugd}}^{\text{BFKL}} \Rightarrow \text{high density effects (if at all present)}$ only through evolution

→ but evolve not even an order of magnitude w.r.t. HERA data;

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Results & Conclusions

A possible way out ...

observables with higher order correlators of Wilson lines \rightarrow inclusive observables (no gap) + resolved final sates (*e.g.* inclusive di- & tri-hadrons/jets)

 $\mathcal{N} \sim 1 - \frac{1}{N_c} \operatorname{tr} \left[V(\boldsymbol{x}) V^{\dagger}(\boldsymbol{y}) \right] \qquad \leftrightarrow \qquad G^{\mathsf{BFKL}}(x, \boldsymbol{k})$

 $\mathcal{Q}^{(4)} \sim 1 - \tfrac{1}{N_c} \mathrm{tr} \left[V(\boldsymbol{x}) V^{\dagger}(\boldsymbol{y}) V(\boldsymbol{y}') V^{\dagger}(\boldsymbol{x}') \right] \quad \leftrightarrow \quad G + \# G^2 + \# G^4 + \dots$

Results & Conclusions

A possible way out ...

observables with higher order correlators of Wilson lines \rightarrow inclusive observables (no gap) + resolved final sates (e.g. inclusive di- & tri-hadrons/jets)



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Supplementary material

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Pomeron: effective degree of freedom which describes the rise of cross-sections with energy

BFKL Pomeron: microscopic description in terms of quarks & gluons \rightarrow requires process with hard scale $Q^2 \gg Q_0^2 \Rightarrow \alpha_s(Q^2) \ll 1$

requires:

expansion of perturbative amplitudes in $1/s\,$

+ resummation of enhanced terms $\left(\alpha_s(Q^2)\ln s\right)^n\sim 1$ to all orders in α_s

→ BFKL equation

- LL: [Fadin, Kuraev, Lipatov, PLB 60 (1975) 50] [Balitsky, Lipatov, SJNP (1978 822)]
- NLL: [Fadin, Lipatov; PLB 429 (1998) 127]; [Ciafaloni, Camici; PLB 430 (1998) 349]



Phenomenology of BFKL evolution

 at LHC: first success in the description of angular decorrelation of multi-jet observables

[Ducloué, Szymanowski, Wallon, 1312.2624], [Celiberto, Ivanov, Murdaca, Papa; 1504.08233]

[Caporale, Chachamis, Murdaca, Sabio Vera; 1508.07711]

→ see talks by Francesco G. Celiberto

▶ test conformal spin $n \neq 0$ components of the BFKL kernel interesting in its own right, perturbatively stable & test of the underlying framework

 ▶ n = 0 component: rise of perturbative cross-sections
 → essentially only studied in fits to inclusive HERA DIS data [Kowalski, Lipatov, Ross; 1005.0355; 1205.6713]
 [Hentschinski, Salas, Sabio Vera; 1209.1353, 1301.5283]

[Levin, Potashnikova; 1307.7823]

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Description within BFKL framework



- exclusive process: vaccuum
 qauntum # exchange between VM
 & proton
- ► Appropiate theoretical framework: non-forward BFKL t ≠ 0
- kernels known up to NLO, but not explored in phenomenological studies & not sufficiently well understood how to use them

use procedure of *e.g.* DGLAP study [Jones, Martin, Ryskin, Teubner; 1307.7099, 1312.6795] → relate *exclusive* photo-production to *inclusive* gluon distribution

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Solve BFKL equation in conjugate (γ) Mellin space

$$\chi_{RG}(\bar{\alpha}_{s},\gamma,a,b) = \bar{\alpha}_{s}(1+a\bar{\alpha}_{s})\left(\psi(\gamma)-\psi(\gamma-b\bar{\alpha}_{s})\right) - \frac{\bar{\alpha}_{s}^{2}}{2}\psi''(1-\gamma) - \frac{b\bar{\alpha}_{s}^{2}\cdot\pi^{2}}{\sin^{2}(\pi\gamma)} + \frac{1}{2}\sum_{m=0}^{\infty}\left(\gamma-1-m+b\bar{\alpha}_{s}-\frac{2\bar{\alpha}_{s}(1+a\bar{\alpha}_{s})}{1-\gamma+m} + \sqrt{(\gamma-1-m+b\bar{\alpha}_{s})^{2}+4\bar{\alpha}_{s}(1+a\bar{\alpha}_{s})}\right)$$

resums (anti-) collinear 'logs' (= γ -poles) of $\bar{\alpha}_s \chi_0(\gamma) + \bar{\alpha}_s \chi_1(\gamma) - \frac{1}{2} \bar{\alpha}_s^2 \chi'_0(\gamma) \chi_0(\gamma)$ [Salam; hep-ph/9806482], [Sabio Vera; hep-ph/0505128]

optimal scale setting $\rightarrow \gamma$ -dependent running coupling

$$\bar{\alpha}_s\left(\overline{M}\cdot Q_0,\gamma\right) = \frac{4N_c}{\beta_0\left[\log\left(\frac{\overline{M}\cdot Q_0}{\Lambda^2}\right) + \frac{1}{2}\chi_0(\gamma) - \frac{5}{3} + 2\left(1 + \frac{2}{3}Y\right)\right]},$$

also use parametrization of running coupling in the infra-red [Webber; hep-ph/9805484]

$$\alpha_s\left(\mu^2\right) = \frac{4\pi}{\beta_0 \ln \frac{\mu^2}{\Lambda^2}} + f\left(\frac{\mu^2}{\Lambda^2}\right), \quad f\left(\frac{\mu^2}{\Lambda^2}\right) = \frac{4\pi}{\beta_0} \frac{125\left(1 + 4\frac{\mu^2}{\Lambda^2}\right)}{\left(1 - \frac{\mu^2}{\Lambda^2}\right)\left(4 + \frac{\mu^2}{\Lambda^2}\right)^4},$$

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relate 2 pictures of the BFKL Pomeron

a) exclusive photo-production of vector mesons:



'uncut' Pomeron: diffractive/elastic scattering (amplitude level)

 $\mathcal{A}(s,t)$



b) proton structure functions:

'cut' Pomeron: high multiplicity events (total X-sec.)

$$\sigma_{\rm tot} = \tfrac{1}{s} \Im \mathsf{m} \mathcal{A}(s,t=0)$$

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Wave function and impact factor

Vector mesons & dipole models ...

factorization into light-front wave function & dipole amplitude

e.g. [Kowalski, Motyka, Watt; hep-ph/0606272]

$$\Im \mathsf{m}\mathcal{A}_{T,L}^{\gamma^* p \to V p}(W,t=0) = 2 \int d^2 \boldsymbol{r} \int d^2 \boldsymbol{b} \int_0^1 \frac{dz}{4\pi} \, (\Psi_V^* \Psi)_{T,L} \, \mathcal{N}\left(x,r,b\right),$$

light-front wave function overlap

$$(\Psi_V^*\Psi)_T = \frac{\hat{e}_f e N_c}{\pi z (1-z)} \left\{ m_f^2 K_0(m_f r) \phi_T(r,z) - \left[z^2 + (1-z)^2 \right] m_f K_1(m_f r) \partial_r \phi_T(r,z) \right\}$$

scalar parts of VM wave function: boosted Gaussian wave-functions with Brodsky-Huang-Lepage prescription

$$\phi_T^{1s}(r,z) = \mathcal{N}_T z (1-z) \exp\left(-\frac{m_f^2 \mathcal{R}_{1s}^2}{8z(1-z)} - \frac{2z(1-z)r^2}{\mathcal{R}_{1s}^2} + \frac{m_f^2 \mathcal{R}_{1s}^2}{2}\right)$$

free parameters fixed through normalization condition & leptonic decay width $\Gamma_{e^-e^+}$:

Meson	$m_f/{ m GeV}$	\mathcal{N}_T	\mathcal{R}^2/GeV^{-2}	$M_V/{ m GeV}$	$8\mathcal{R}^{-2}/GeV^2$	$rac{1}{4}M_V^2/{ m GeV^2}$
J/ψ	$m_c = 1.27$	0.596	2.45	3.097	3.27	2.40
Υ	$m_b = 4.2$	0.481	0.57	9.460	15.38	22.42

use parameters obtained by [Armesto, Rezaeian; 1402.4831], [Goncalves, Moreira, Navarra; 1408.1344] Martin Hentschinski (UDLAP) BFKL & the growth of the VM Xsec. 04/04/1017 25 / 30

Data

Existing description of data

... work pretty well

- $\blacktriangleright~J/\Psi$: power-law fit to HERA data $\sigma\sim W^{0.67}$ [LHCb Collaboration; 1401.3288]
- collinear factorization: NLO fits [Jones, Martin, Ryskin, Teubner; 1307.7099]
- saturation models: IPsat, bCGC, rcBK

[Armesto, Rezaeian; 1402.4831], [Goncalves, Moreira, Navarra; 1405.6977, 1408.1344]

See also [Fiore, Jenkovszky, Libov, Machado; 1408.0530], [Cisek, Schäfer, Szczurek; 1405.2253]

BFKL special:

don't fit W -dependence, but calculate from perturbative low x evolution don't evoke saturation (= effects beyond BFKL)

LHC: reach ultra-small x values $\simeq 4 \cdot 10^{-6}$ not constrained by HERA

Data

Comparison to data

- provide results for both HERA fits (standard (fit 1) & kinematic improved (fit 2) LO impact factor)
- ▶ hard scale M²:
 - photoproduction scale $M_{
 m pp}=M_V/2$

$$\begin{split} & \left(M_{\rm pp}^2 \right)_{J/\Psi} = 2.40 \ {\rm GeV^2} \\ & \left(M_{\rm pp}^2 \right)_{\Upsilon} = 22.42 \ {\rm GeV^2} \end{split}$$

- impact factor motivated: $M_{if}^2 = 8 \mathcal{R}_V^{-2}$ – eliminates $(...)^{\gamma}$ factor & minimizes NLO running coupling correction related to impact factor

$$\begin{split} & \left(M_{\mathrm{if}}^2\right)_{J/\Psi} = 3.27 \ \mathrm{GeV^2} \\ & \left(M_{\mathrm{if}}^2\right)_{J/\Psi} = 15.38 \ \mathrm{GeV^2} \end{split}$$

- ▶ (hard) running coupling scale $\overline{M} = M$, but vary in range $[M^2/2, M^2 \cdot 2]$ to check stability of result
- ▶ fix normalization by low energy ALICE (J/ Ψ) and H1 (Υ) data point → K-factor

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Observations:

- ► K-factor: small for fit 2, sizeable for fit 1 – likely related to the impact factors in used in the HERA fit (massless, n_f = 4, (C₁/C₂)² = 2.45)
- ▶ common correction not included: GPD motivated factor to take into account $x' \neq x$; currently calculated for collinear pdf [Shuvaev, Golec-Biernat, Martin, Ryskin, hep-ph/9902410] → to be calculated for k_T factorized BFKL impact factor



very good description of W-dependence

 $W_{J/\Psi} > 471 \text{ GeV } \& W_{\Upsilon} > 669 \text{ GeV} \equiv$ beyond region of incl. HERA fit (from $x = 4.3 \cdot 10^{-5}$ to $x = 3.5 \cdot 10^{-6}$) \rightarrow direct test of BFKL evolution

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Caveats

- both BFKL HERA fit & VM photoproduction use LO impact factor
 large corrections at NLO possible
- ▶ BFKL HERA fit for $n_f = 4$ mass-less quarks

both effects should affect the normalization, not so much $W\mbox{-dependence}$

unintegrated gluon density can develop instability at ultra-small x:

$$G\left(x, \boldsymbol{k}^{2}, M\right) = \frac{1}{\boldsymbol{k}^{2}} \int_{\frac{1}{2} - i\infty}^{\frac{1}{2} + i\infty} \frac{d\gamma}{2\pi i} \quad \hat{g}\left(x, \gamma\right) \quad \left(\frac{\boldsymbol{k}^{2}}{Q_{0}^{2}}\right)^{\gamma}$$
$$f(x, \gamma) \sim \left(\frac{1}{x}\right)^{\chi\left(\gamma, \frac{M^{2}}{M^{2}}\right)} \cdot \left\{1 + \frac{\bar{\alpha}_{s}^{2} \beta_{0} \chi_{0}\left(\gamma\right)}{8N_{c}} \log\left(\frac{1}{x}\right) \left[-\psi\left(\delta - \gamma\right) + \log\frac{M^{2}}{Q_{0}^{2}} - \partial_{\gamma}\right]\right\},$$

▶ will enter at some point region $\alpha_s^2 \ln(1/x) \sim 1 \implies$ control of such terms will become necessary

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DGLAP

(physical) DGLAP evolution [MH, Stratmann; arXiv:1311.2825, to be completed] direct evolution of *e.g.* doublet (F_2, F_L) ; renormalization scale dependence $\mu_R = k \cdot Q^2 \dots$ stable for $Q^2 = 30 \text{ GeV}^2 \rightarrow 100 \text{GeV}^2 \dots$



DGLAP

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for $Q^2 = 2 \text{ GeV}^2 \rightarrow 10 \text{GeV}^2$ at small x highly instable reason: un-resummed $\ln 1/x$

convinced:

- \blacktriangleright pdf studies highly valuable ... \rightarrow constrain pdf's at ultra-small x
- benchmark for saturation searches (?)

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