

Illuminating Nucleon Structure Through Polarized Proton-Proton Collisions at STAR

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Lamar University

25th International Workshop on Deep Inelastic Scattering and Related Topics

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OUTLINE

- Introduction
- RHIC and STAR
- Recent Developments
- Near-term Plans
- Summary



The Fertile Field of Spin Physics

The study of spin in particle physics has unlocked doors to a deeper understanding of nucleon structure

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- Helicity

Recent results enable a better picture of gluon and sea-quark helicity

The Fertile Field of Spin Physics

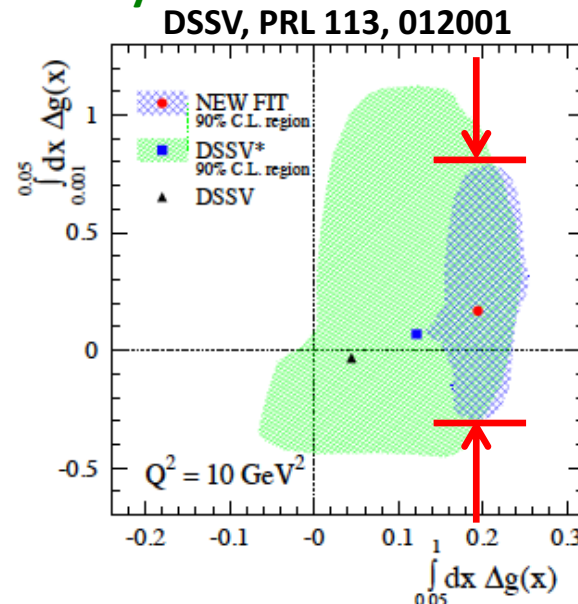
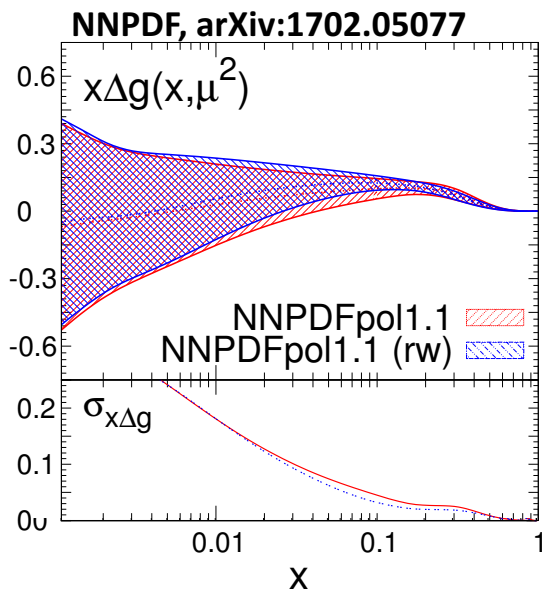
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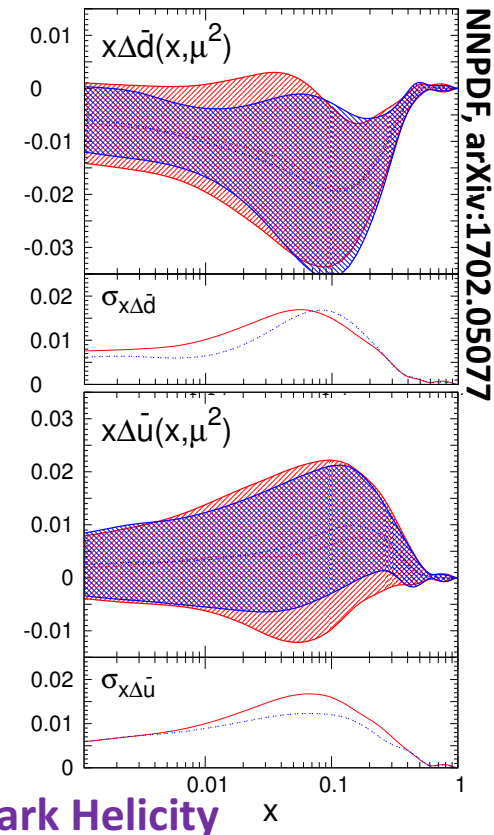
Recent results enable a better picture of gluon and sea-quark helicity

STAR data have played a key role!

Gluon Helicity



See talks by Q. Xu and B. Surrow



Sea-quark Helicity

The Fertile Field of Spin Physics

The study of spin in particle physics has unlocked doors to a deeper understanding of nucleon structure

- Helicity
- Transversity

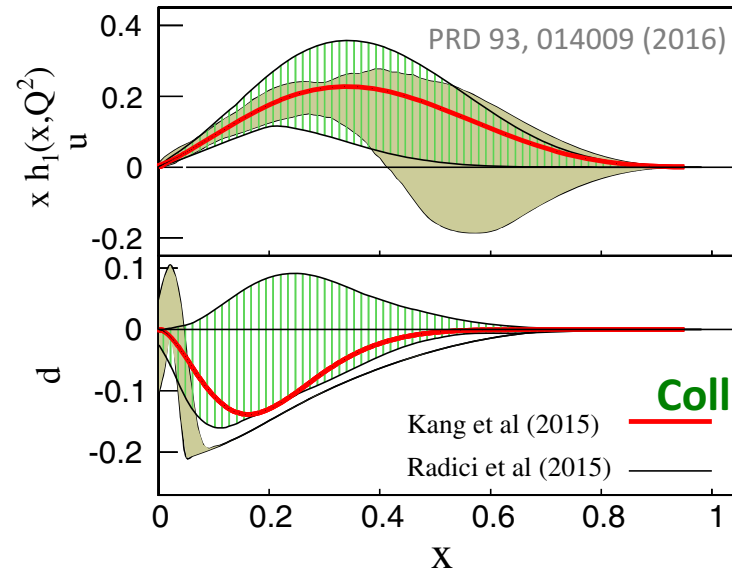
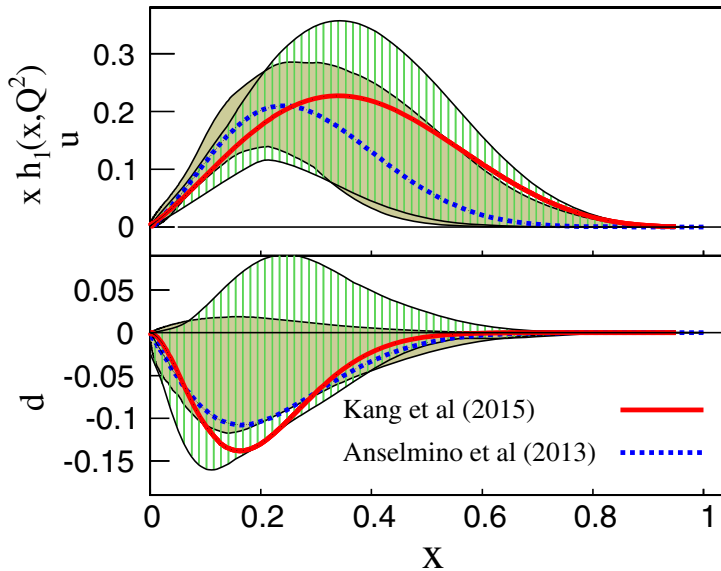
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- Helicity
- Transversity

Multiple mechanisms in play to constrain transverse spin-structure

Collins from SIDIS



Collins from SIDIS

IFF from SIDIS

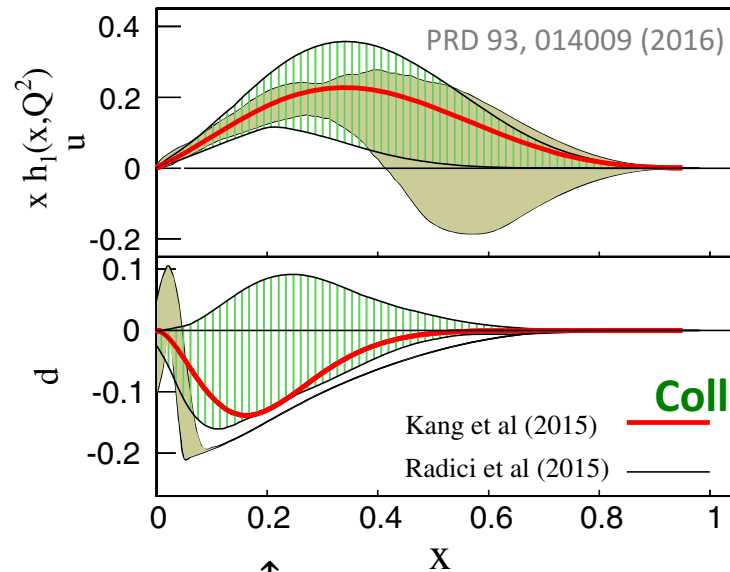
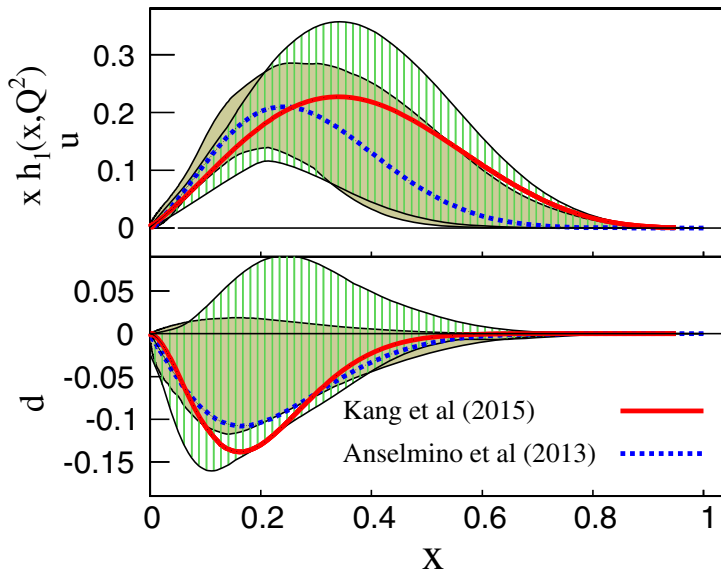
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Multiple mechanisms in play to constrain transverse spin-structure

Collins from SIDIS



Opportunities with $p^\uparrow + p$:

Expanded kinematics! Eliminate u -dominance!

Collins from SIDIS

IFF from SIDIS

(Preliminary update at this conference!)

The Fertile Field of Spin Physics

The study of spin in particle physics has unlocked doors to a deeper understanding of nucleon structure

- Helicity
- Transversity
- Higher dimensions

Non-collinear probes, e.g. TMDs, enable multidimensional imaging

The Fertile Field of Spin Physics

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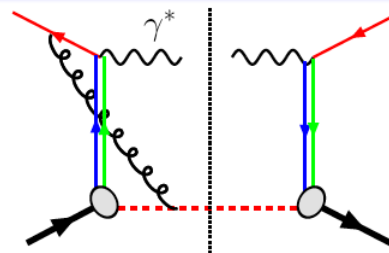
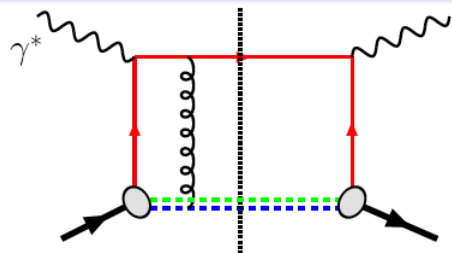
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Non-collinear probes, e.g. TMDs, enable multidimensional imaging

QCD:

DIS:
Final-state interaction

Drell-Yan or W:
Initial-state interaction



$$\text{Sivers}_{\text{DIS}} = - \text{Sivers}_{\text{Drell-Yan}} \text{ or } \text{Sivers}_{\text{W}}$$

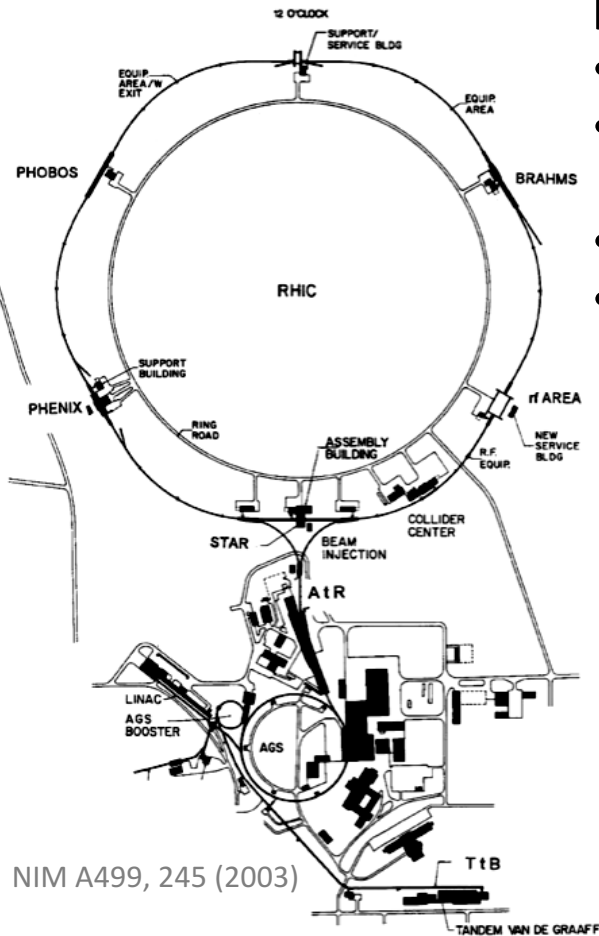
Opportunities with $p^\uparrow + p$:

Tests of Evolution? Factorization and Universality?

Relativistic Heavy Ion Collider

RHIC as Polarized-proton Collider

- “Siberian Snakes” → mitigate depolarization resonances
- Spin rotators provide choice of spin orientation
independent of experiment
- Spin direction varies bucket-to-bucket (9.4 MHz)
- Spin pattern varies fill-to-fill

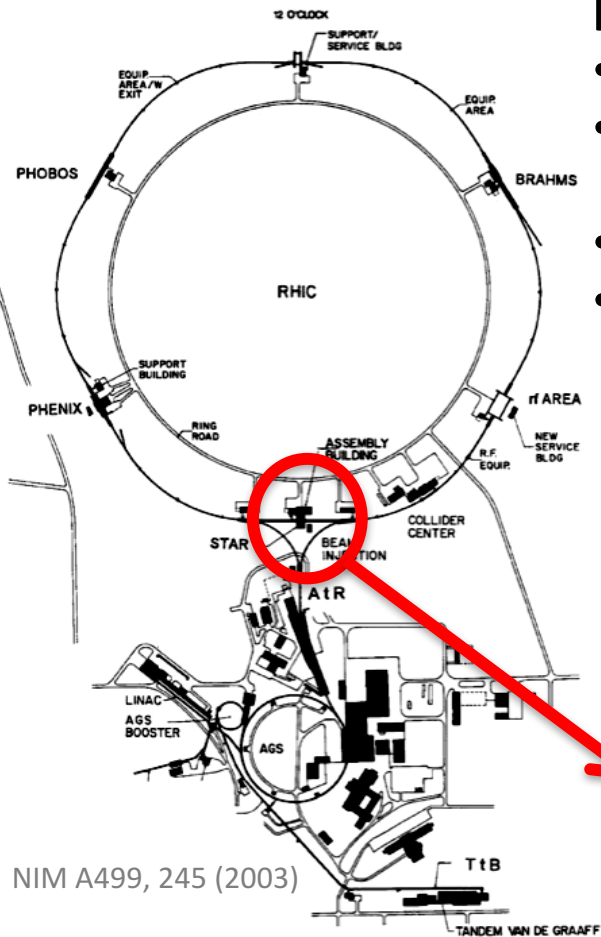


NIM A499, 245 (2003)

Solenoidal Tracker at RHIC

RHIC as Polarized-proton Collider

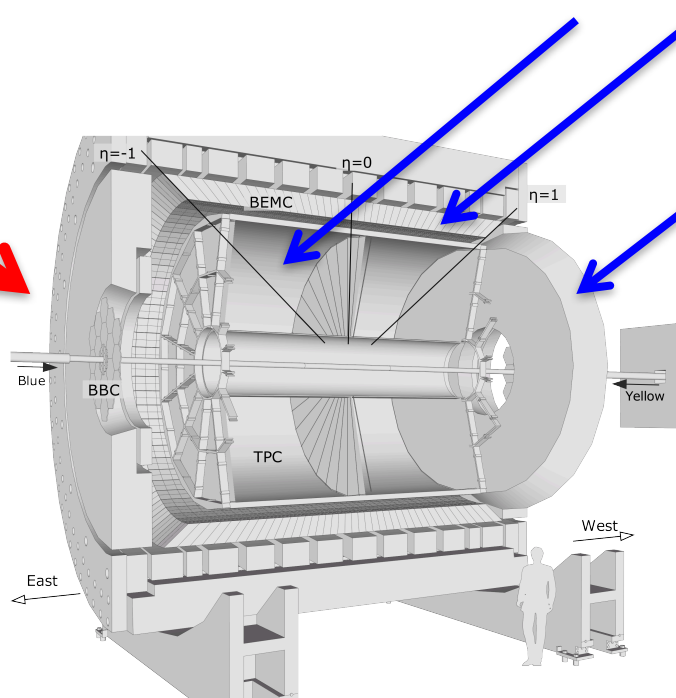
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NIM A499, 245 (2003)

Jets, di-hadrons, weak bosons

TPC + Barrel + Endcap EMC



π^0 , photons, jet-like measurements

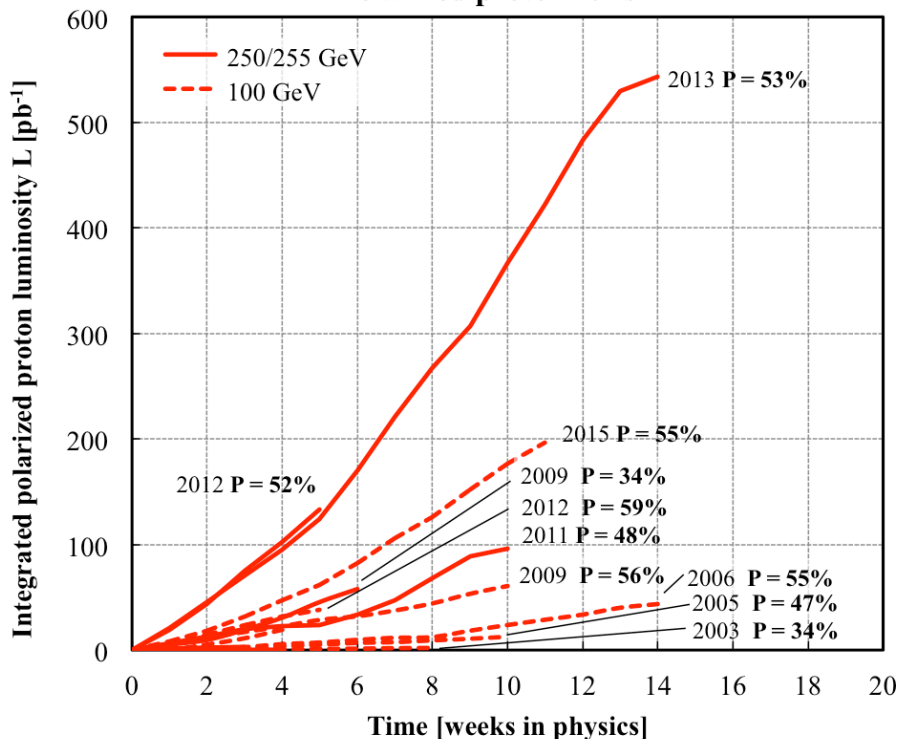
FMS

Trigger on calorimeter energy

Polarized-proton Datasets at RHIC

Unique opportunities to probe nucleon spin structure!

Polarized proton runs



Transverse Luminosity Recorded

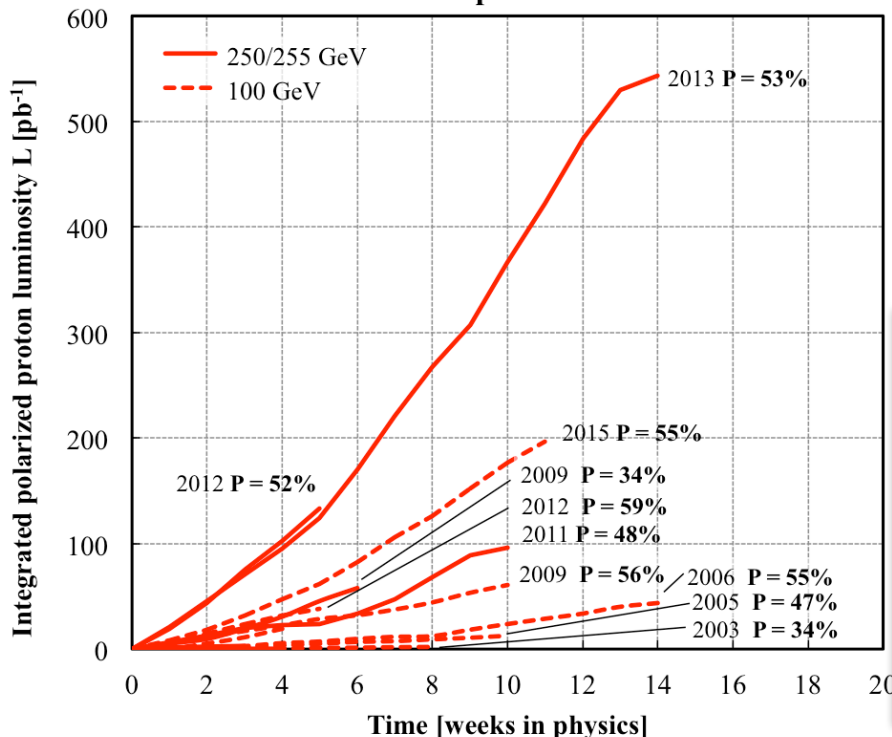
Year	\sqrt{s} [GeV]	STAR	PHENIX	$\langle P \rangle$ [%]
2006	200	8.5 pb ⁻¹	2.7 pb ⁻¹	57
2006	62.4	0.2 pb ⁻¹	0.02 pb ⁻¹	48
2008	200	7.8 pb ⁻¹	5.2 pb ⁻¹	45
2011	500	25 pb ⁻¹	--	53/54
2012	200	22 pb ⁻¹	9.7 pb ⁻¹	61/58
2015	200	53 pb ⁻¹	52 pb ⁻¹	53/57
2015	200 pAu	0.42 pb ⁻¹	0.20 pb ⁻¹	60
2015	200 pAl	1.0 pb ⁻¹	--	54

PHENIX numbers for $|z_{\text{vtx}}| < 40$ cm

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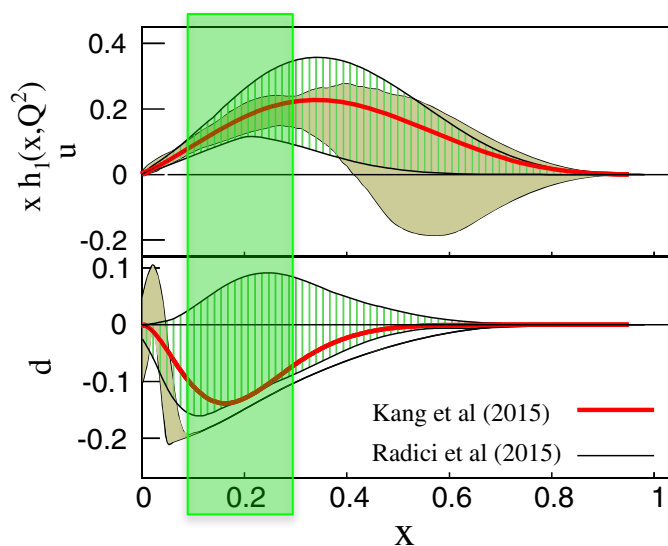
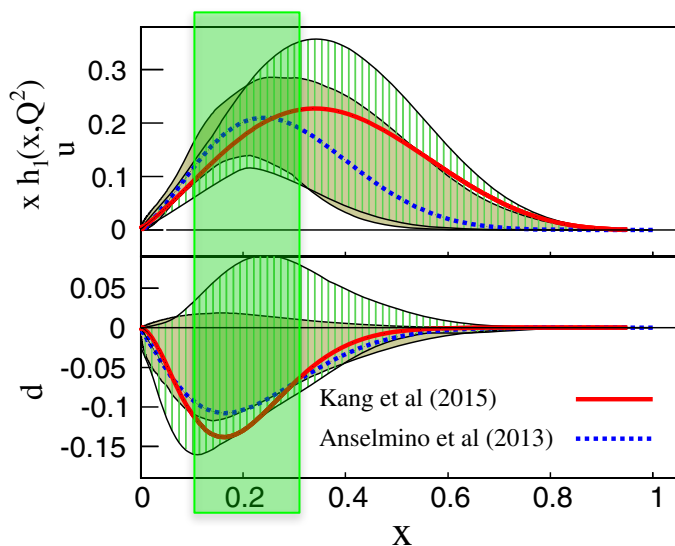
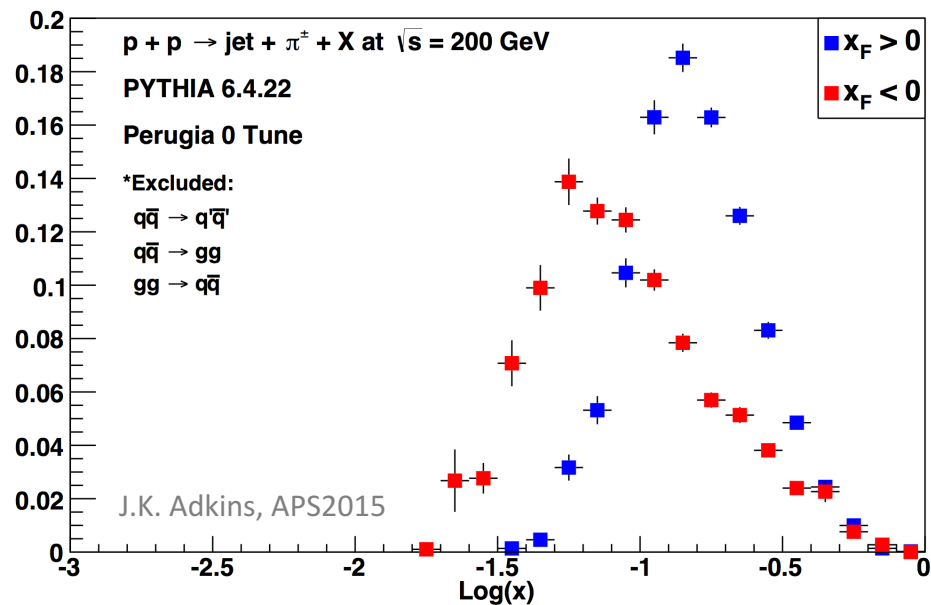
Dramatically increased figure of merit in recent years

Integrating more 510 GeV data, as we speak!

Sensitivity to Transversity at STAR

Access to transversity in interesting region!

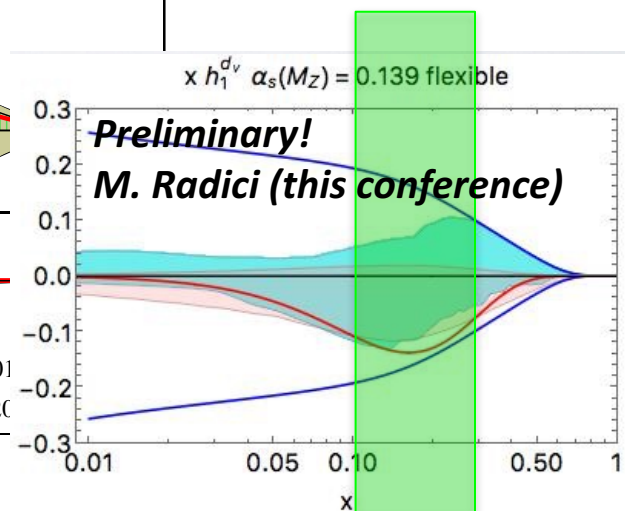
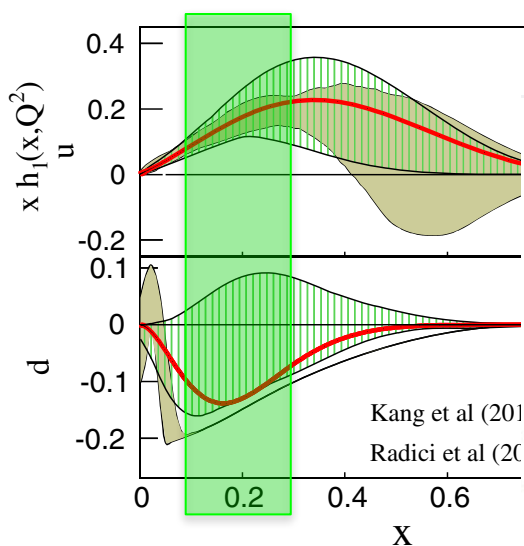
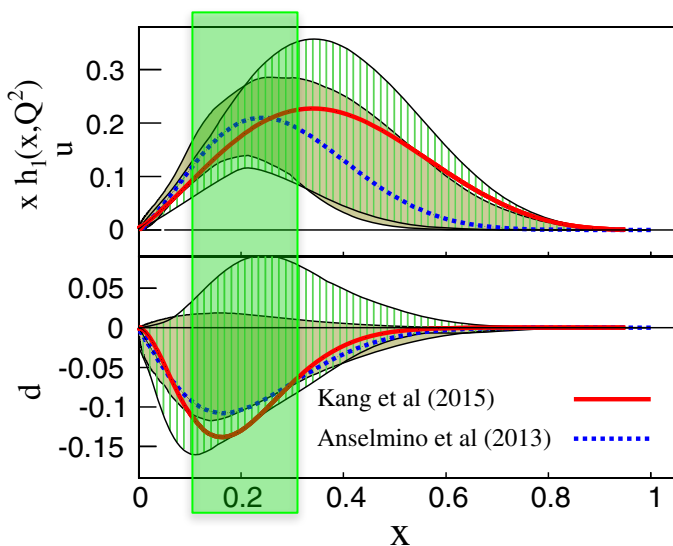
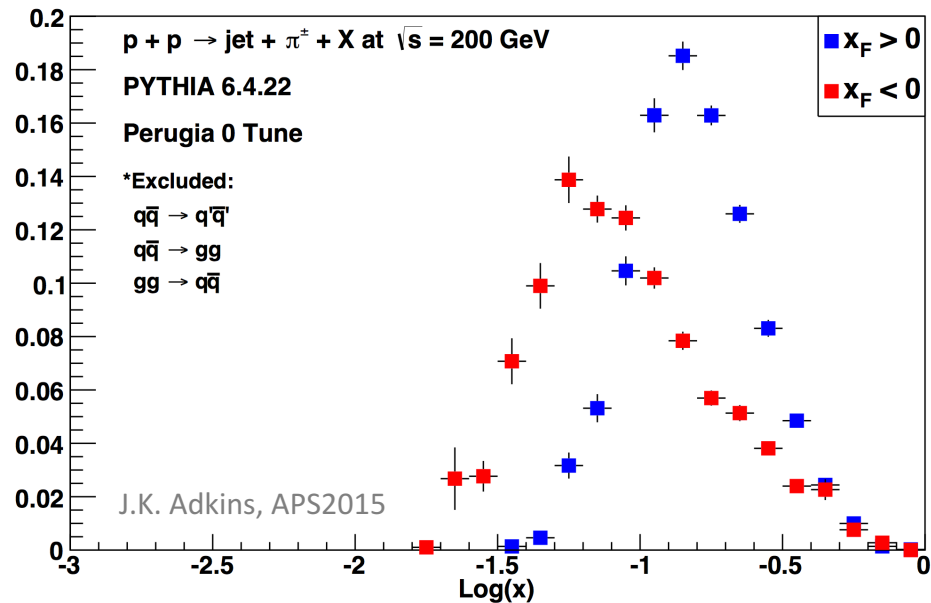
- Limited constraints
- Potentially large effects
- Sensitivity to evolution
- *Insight into nature of Collins mechanism!*



Sensitivity to Transversity at STAR

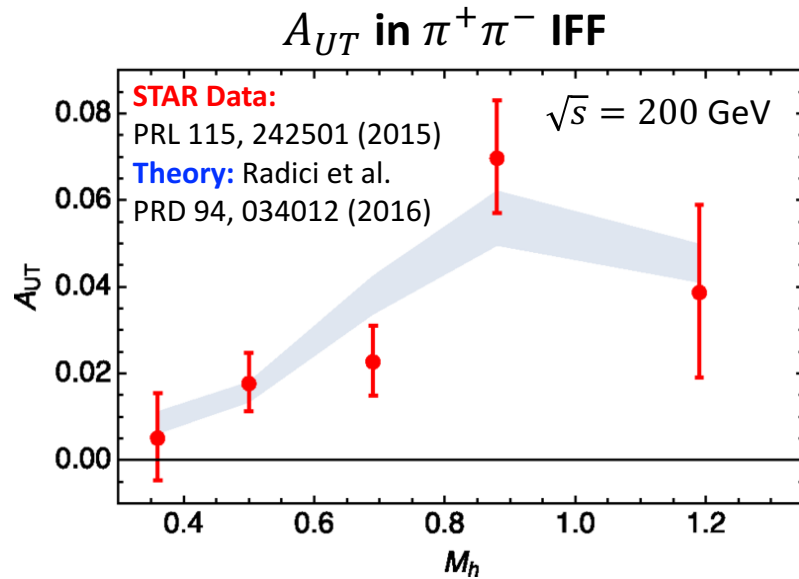
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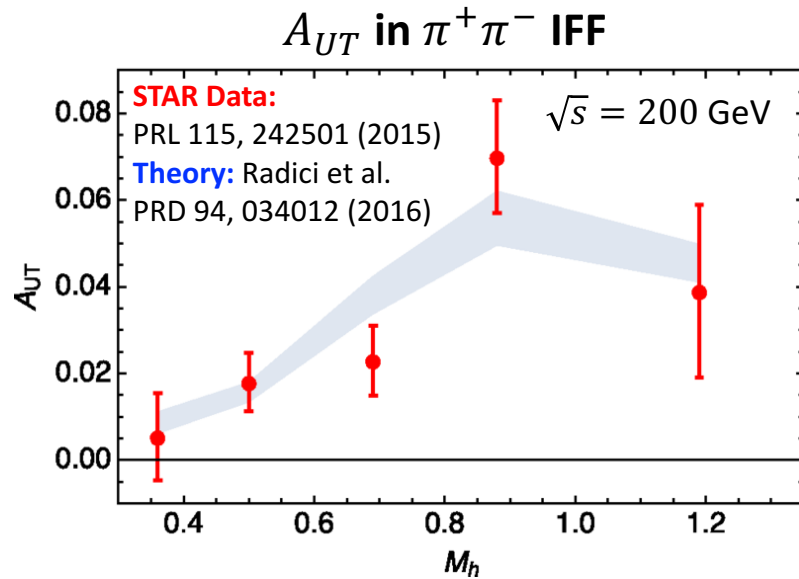
Comparison of Early STAR Data to Theory

A clear message from early results:
Access to transversity effects at RHIC



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Compare with models
based on *SIDIS*/ e^+e^-
Band represents 68% of replicas deduced
by fitting *SIDIS* and e^+e^- data

Overall **agreement** in terms of
invariant mass

→ **Same mechanism as in *SIDIS*!**

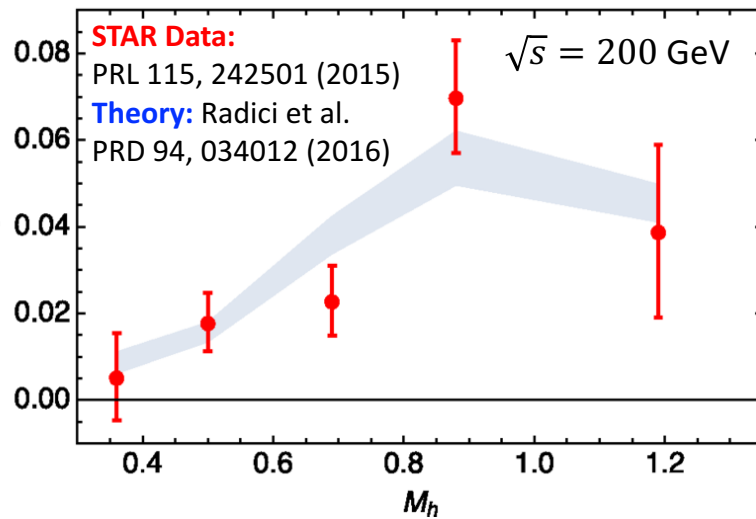
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A_{UT} in $\pi^+\pi^-$ IFF



Overall *agreement* in terms of
invariant mass

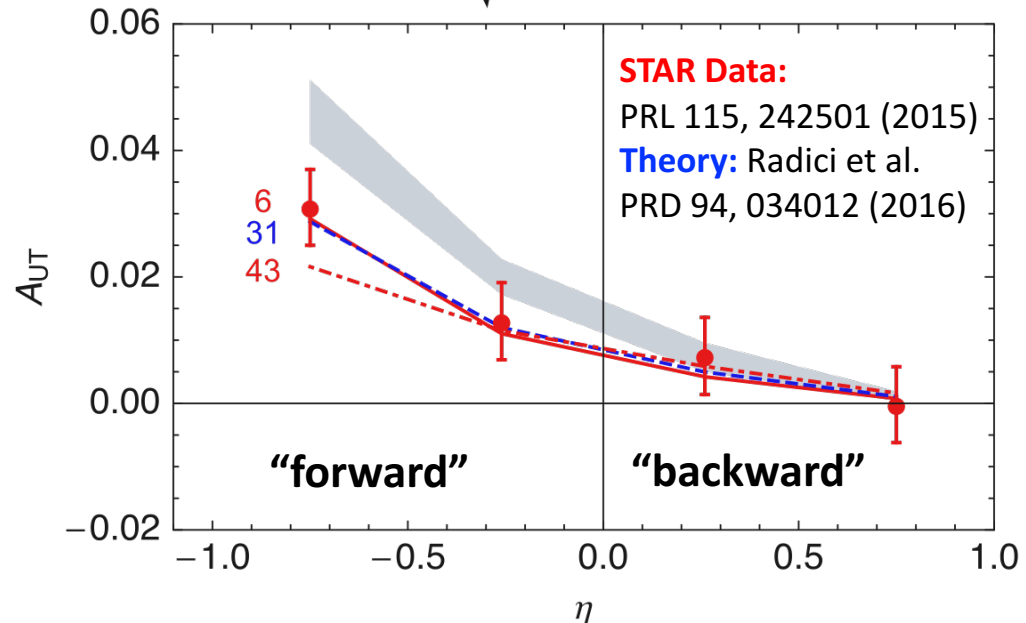
→ *Same mechanism as in SIDIS!*

Deviation at more *forward* scattering

→ *Tension with SIDIS?*

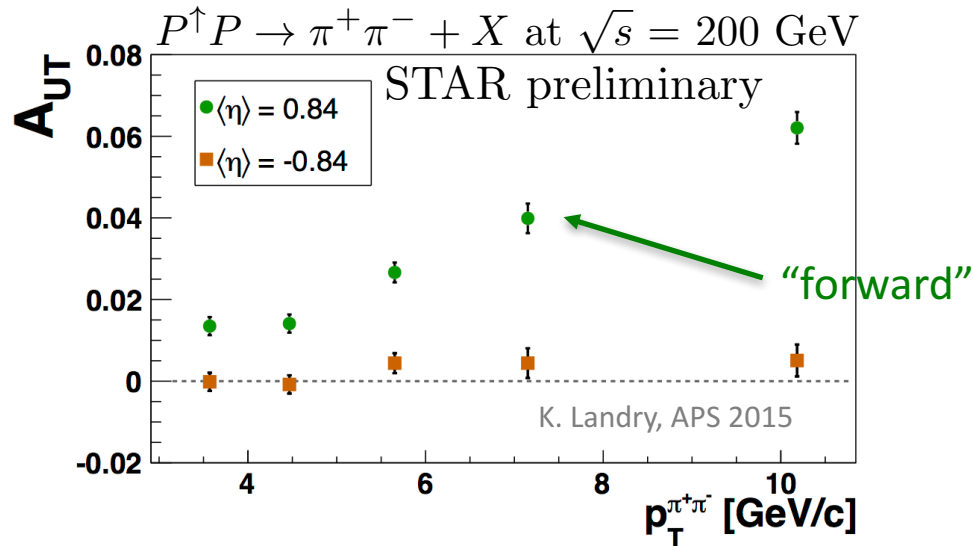
→ *More information needed on D_g^1 ?*

$\sqrt{s} = 200$ GeV



Recent IFF Results at STAR

Much larger datasets collected at 500 and 200 GeV in 2011 and 2012

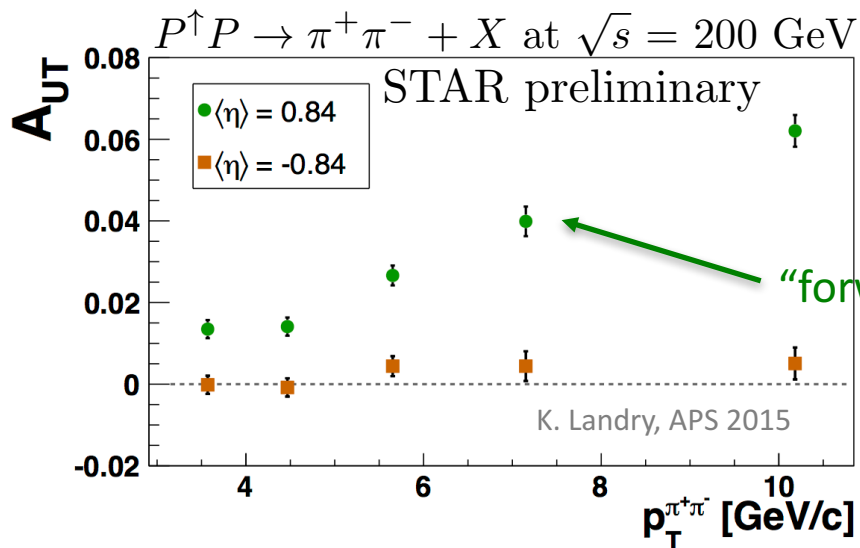


**Significant non-zero
di-hadron asymmetries at
 $\sqrt{s} = 200$ and 500 GeV!**

- Increasing with pion p_T

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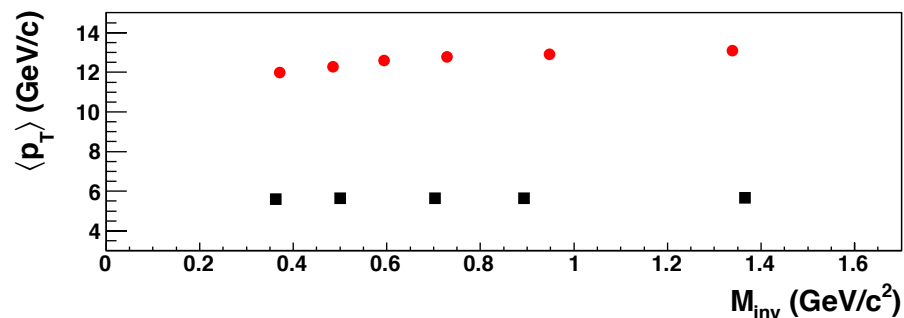
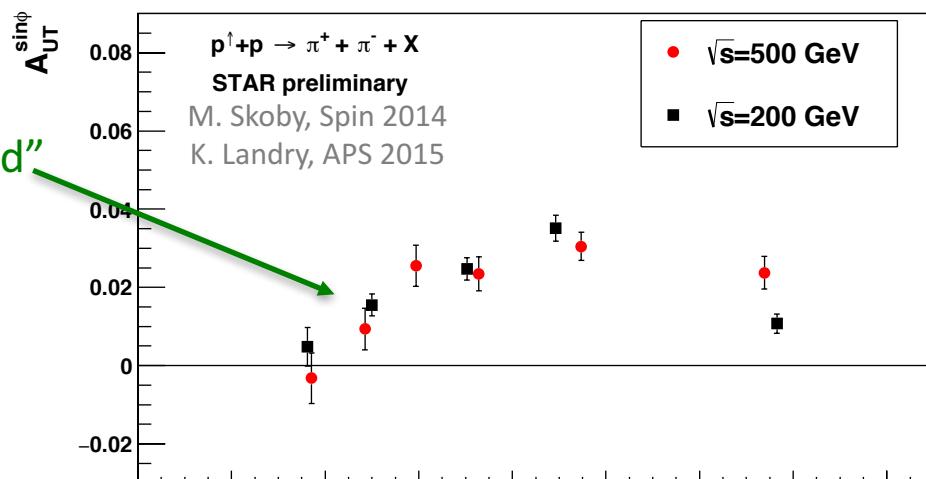
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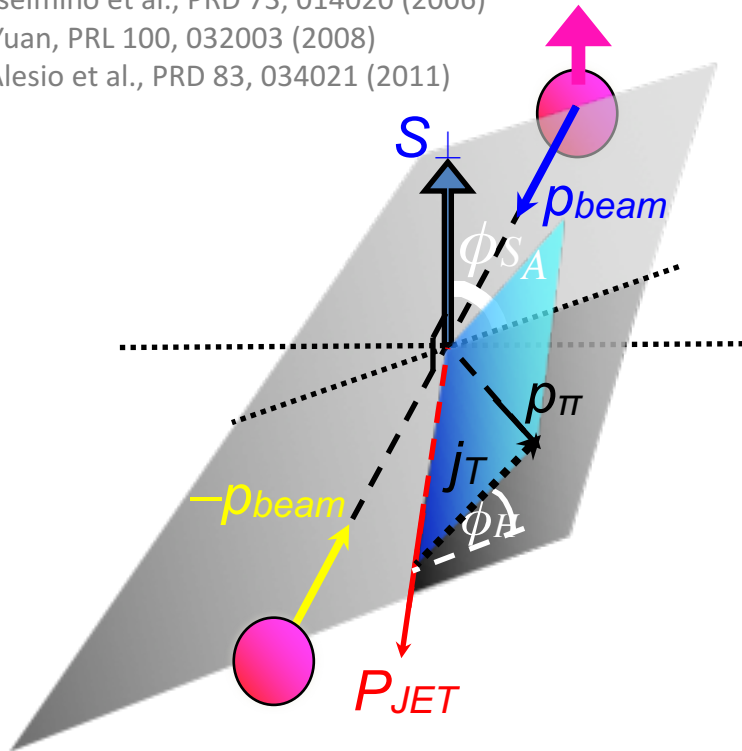
- Increasing with pion p_T

Consistent behavior when scaled for $2\langle p_T \rangle / \sqrt{s}$



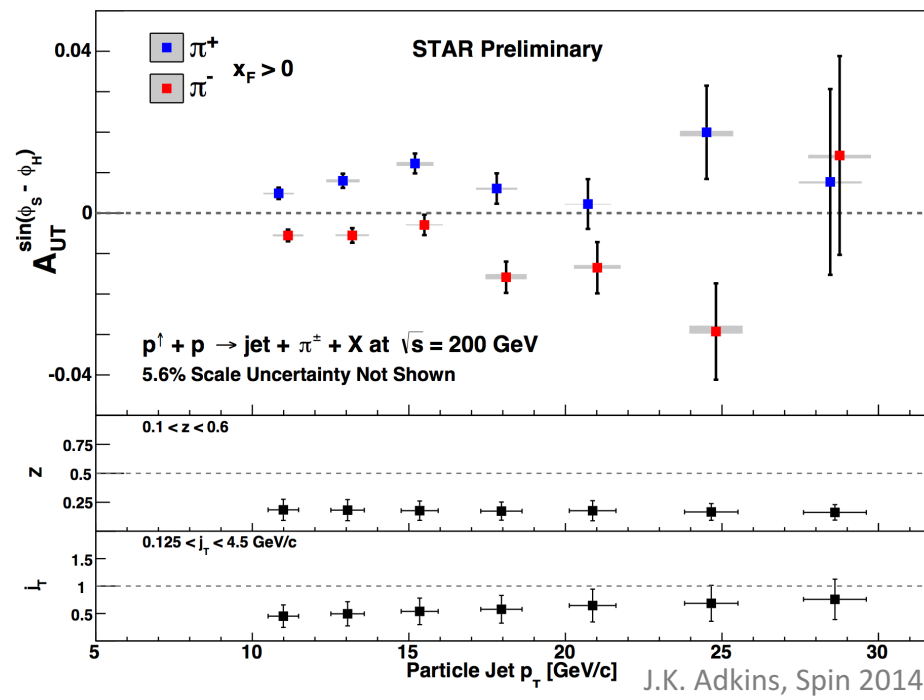
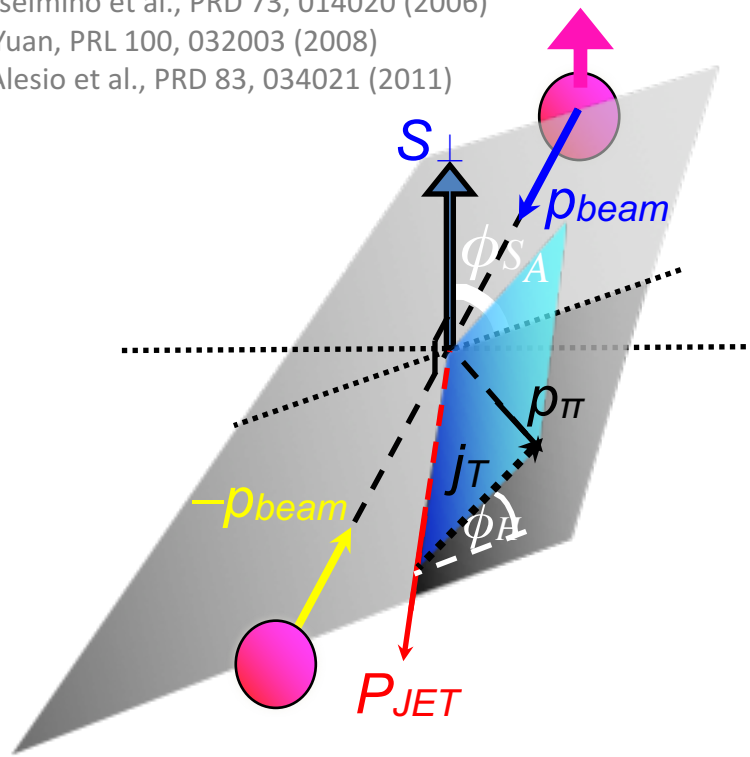
Newest Collins Results at $\sqrt{s} = 200$ GeV

Anselmino et al., PRD 73, 014020 (2006)
F. Yuan, PRL 100, 032003 (2008)
D'Alesio et al., PRD 83, 034021 (2011)



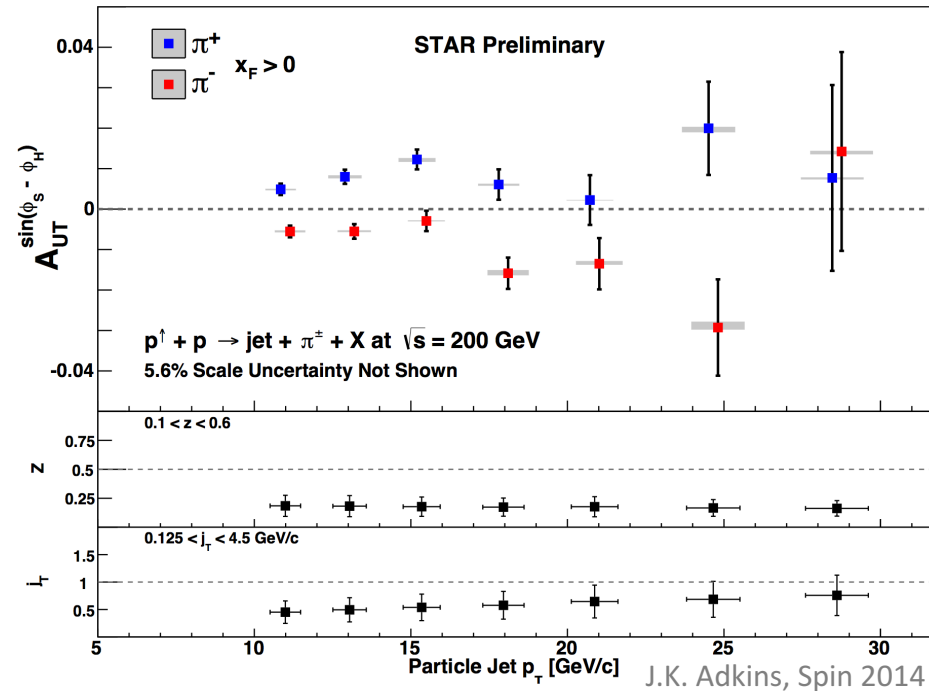
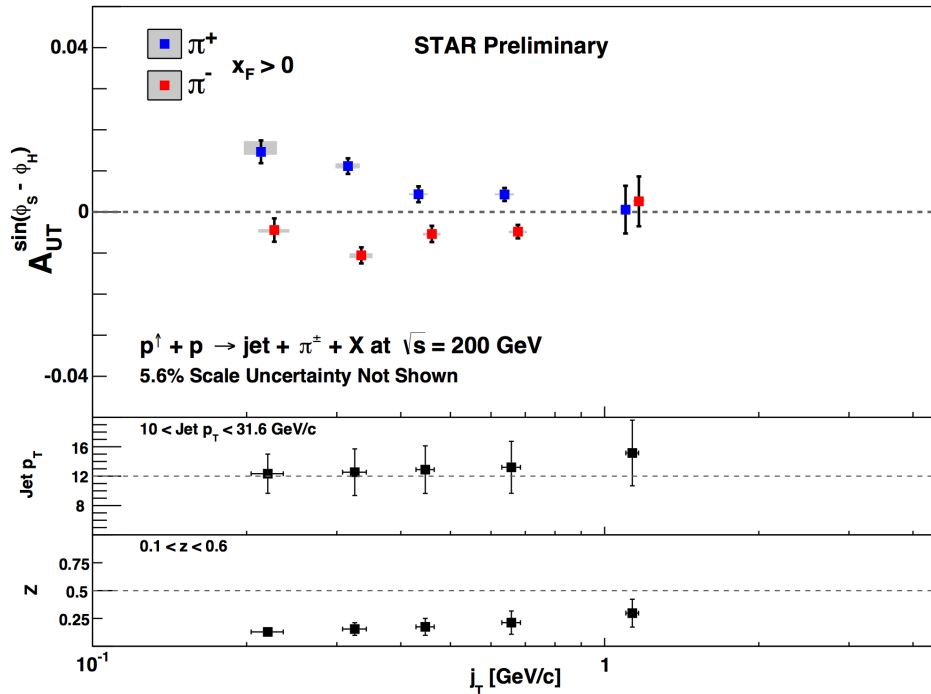
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**Clear first observation of
 Collins asymmetry in $p + p$!**

Newest Collins Results at $\sqrt{s} = 200$ GeV



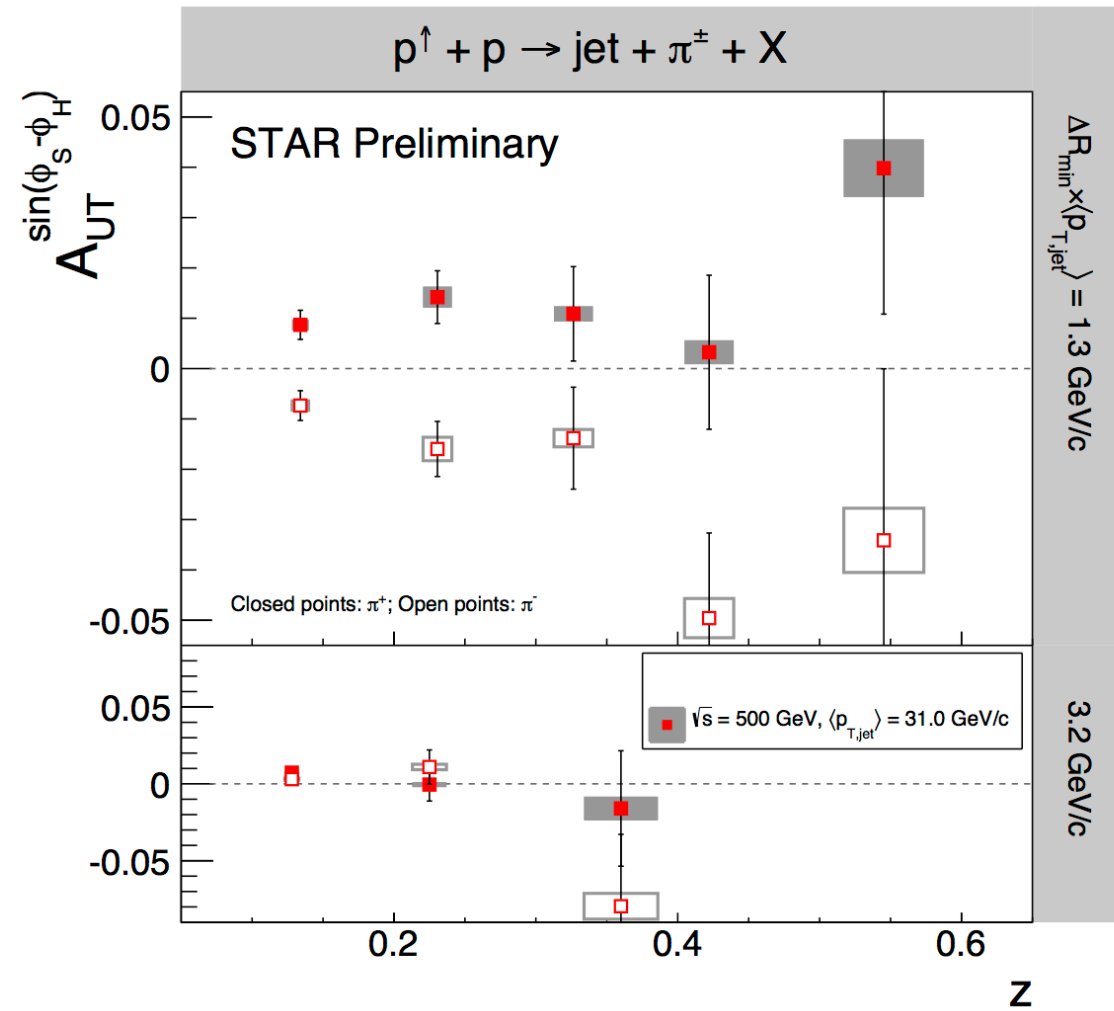
STRONG dependence upon j_T

$$j_{T,\min} \approx z \times \Delta R_{\min} \times \langle p_T \rangle,$$

$$\Delta R = \sqrt{(\eta_{\text{jet}} - \eta_{\pi})^2 + (\phi_{\text{jet}} - \phi_{\pi})^2}$$

Clear first observation of Collins asymmetry in $p + p$!

STAR Collins Results at $\sqrt{s} = 500$ GeV

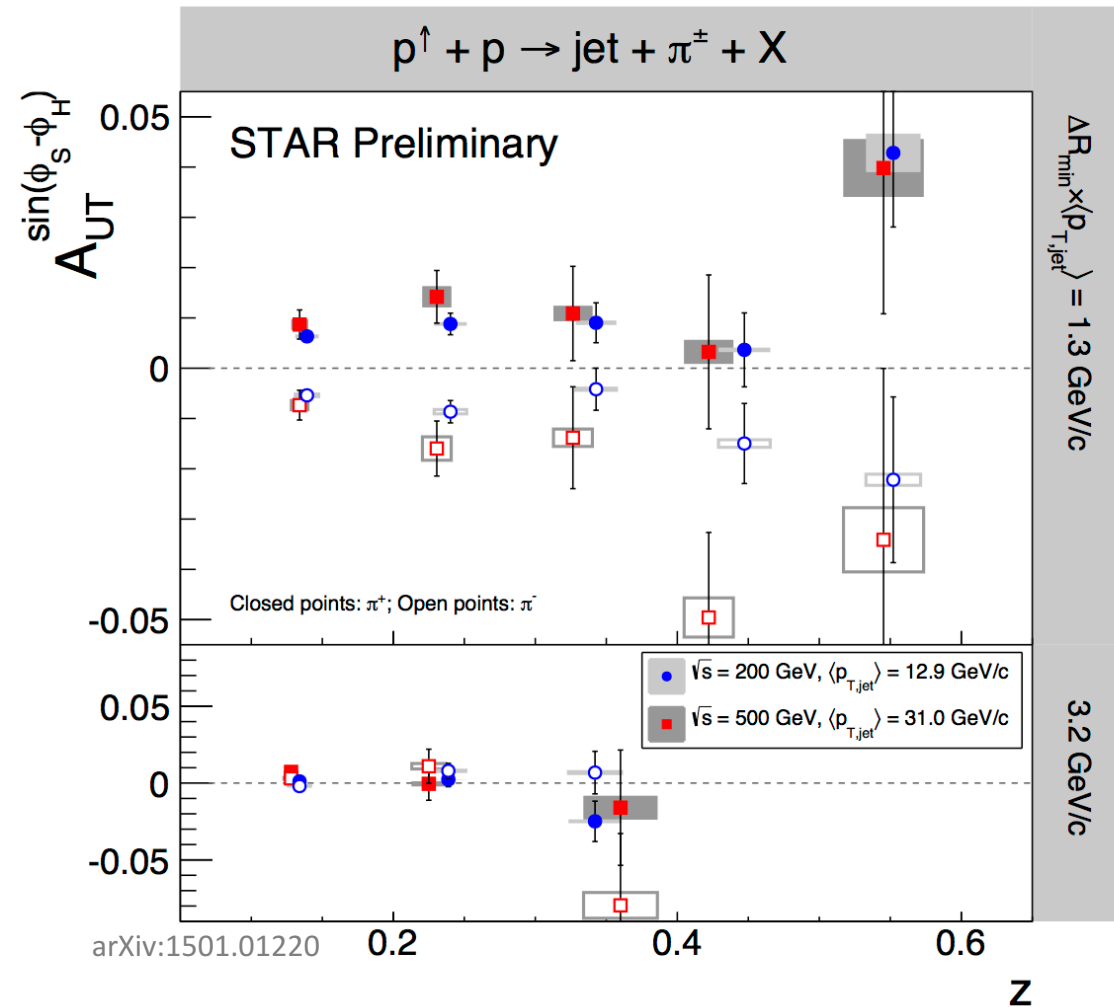


Non-zero Collins asymmetries observed

at $\sqrt{s} = 500$ GeV!

- Strong dependence on $\Delta R_{\min}(j_{T,\min})$

STAR Collins Results at $\sqrt{s} = 200$ and 500 GeV

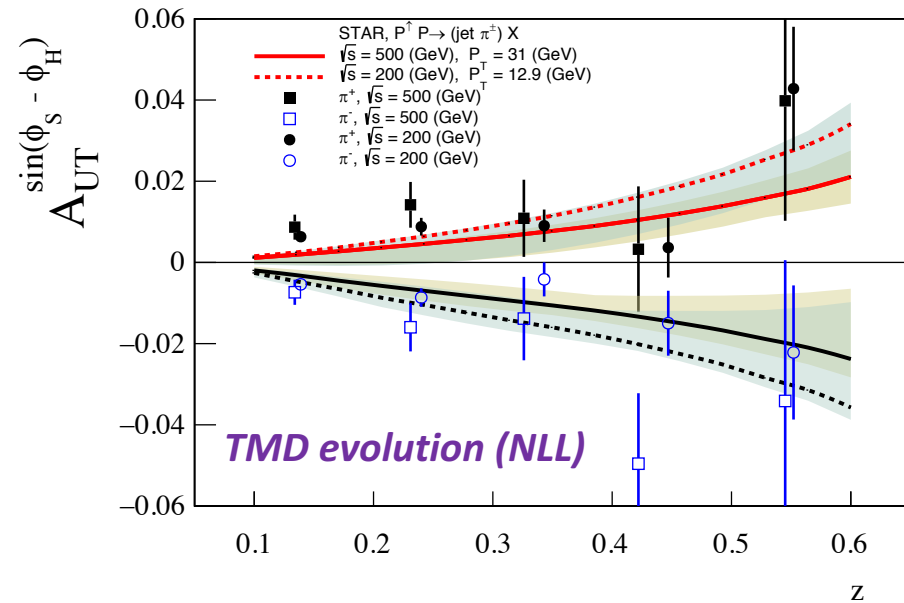


Non-zero Collins asymmetries observed at $\sqrt{s} = 500 \text{ GeV}$!

- Strong dependence on $\Delta R_{\min}(j_{T,\min})$
- Consistent with $\sqrt{s} = 200 \text{ GeV}$ results for consistent cuts and x_T

At the current precision, Collins results from $p + p$ appear consistent with x_T scaling

STAR Collins Results at $\sqrt{s} = 200$ and 500 GeV



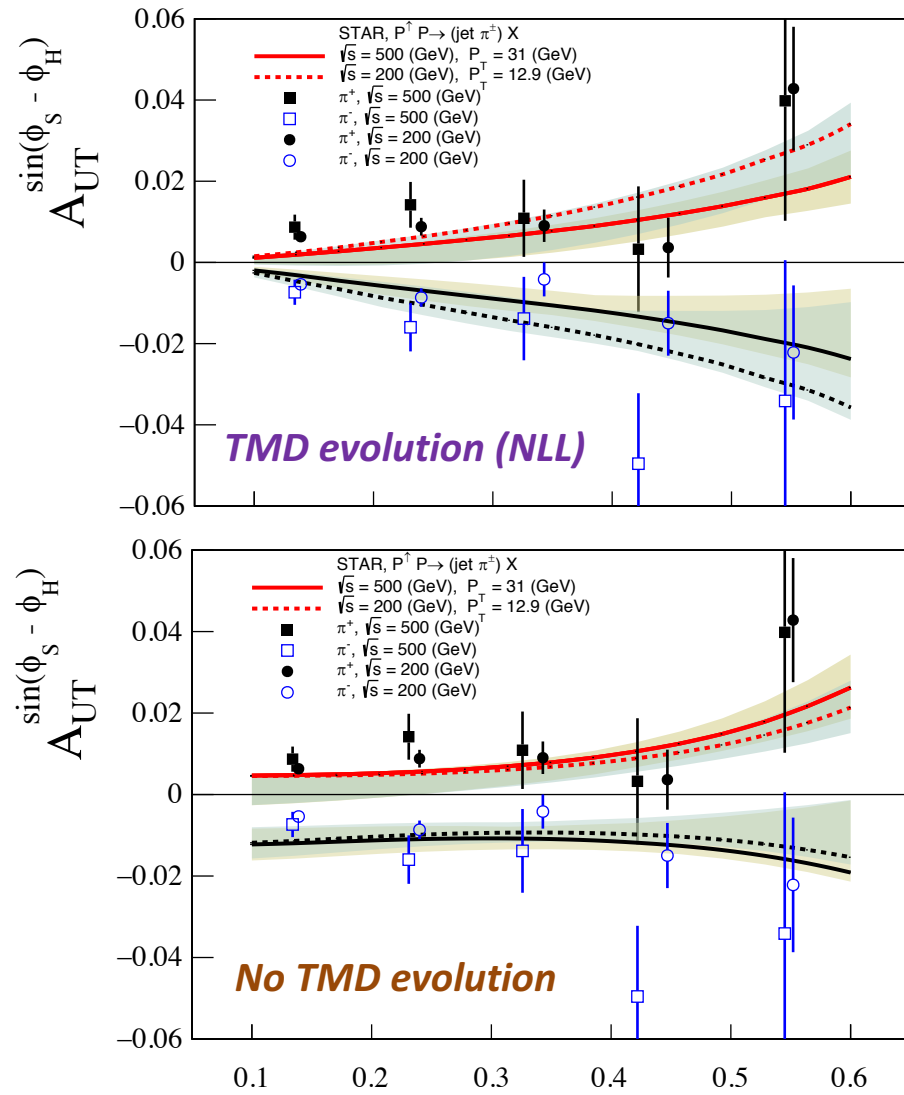
Compare with models based on
SIDIS/ e^+e^-

- Assume **universality** and **robust factorization**
- **One model with TMD evolution up to NLL**

Theory: Kang, Prokudin, Ringer, Yuan 2017 in preparation

Data: **STAR Preliminary**

STAR Collins Results at $\sqrt{s} = 200$ and 500 GeV



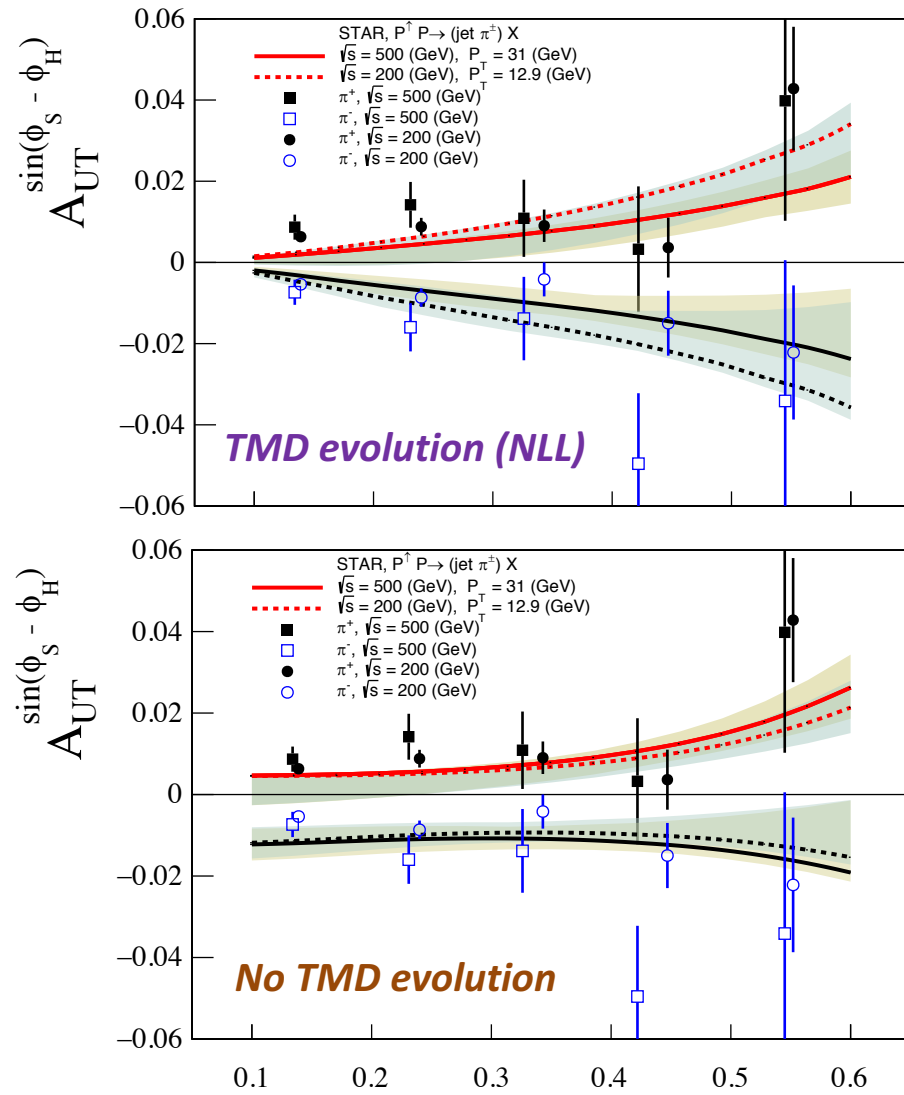
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Compare with models based on *SIDIS*/ e^+e^-

- Assume **universality** and **robust factorization**
- One model with TMD evolution up to NLL**
- One model without TMD evolution**

Generally decent agreement between models and STAR data!

Slight preference for no evolution?

“Beauty is in the eye of the beholder!”

Theory: Kang, Prokudin, Ringer, Yuan 2017 in preparation

Data: **STAR Preliminary**

Weak Boson Asymmetries at $\sqrt{s} = 500$ GeV

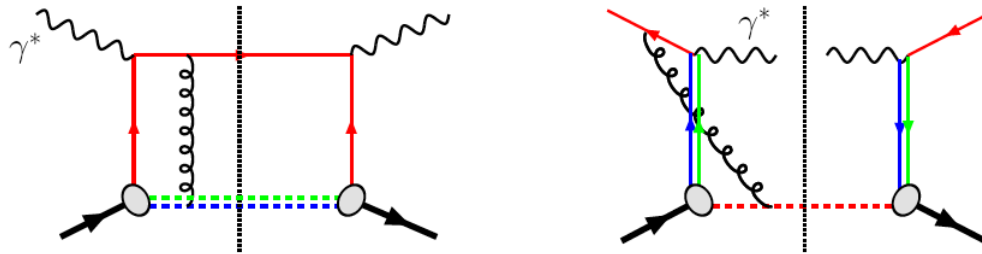
Color interactions in QCD

Non-universality of the “Sivers” function

QCD:

DIS:
Final-state interaction

Drell-Yan or W:
Initial-state interaction



$$\text{Sivers}_{\text{DIS}} = - \text{Sivers}_{\text{Drell-Yan}} \text{ or } \text{Sivers}_{\text{W}}$$

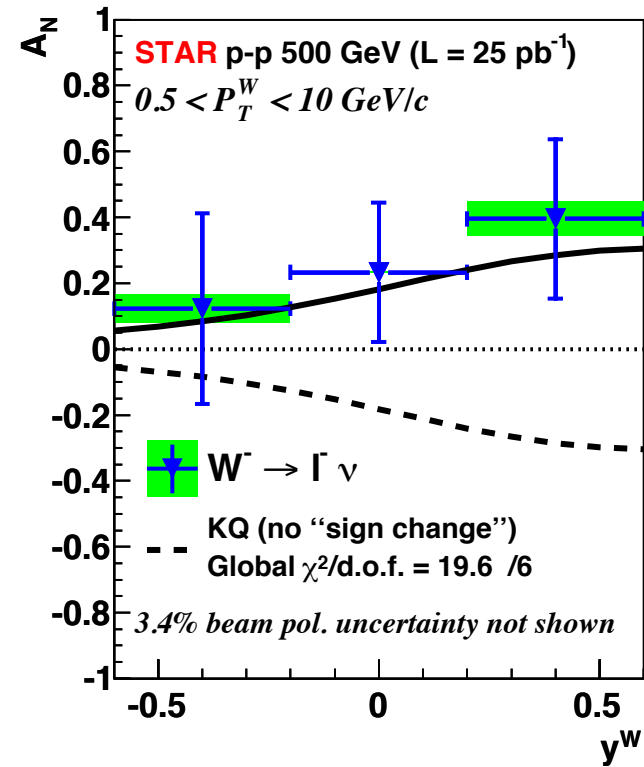
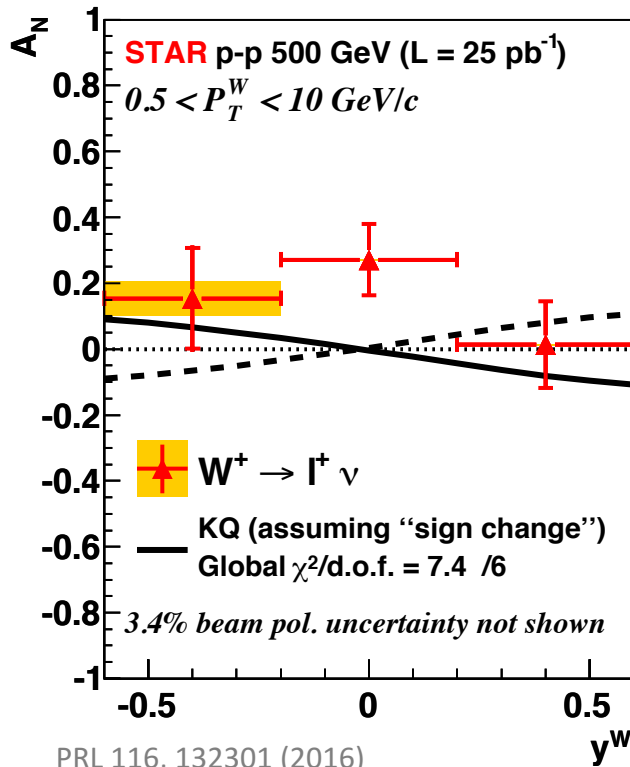
A_N for direct photon also has a closely related sign change

*Opportunity to see the repulsive interaction between
like color charges for the first time!*

Can explore all of these observables in 500 GeV $p + p$ collisions at RHIC!

Weak Boson Asymmetries at $\sqrt{s} = 500$ GeV

Exploratory Measurement from a small dataset

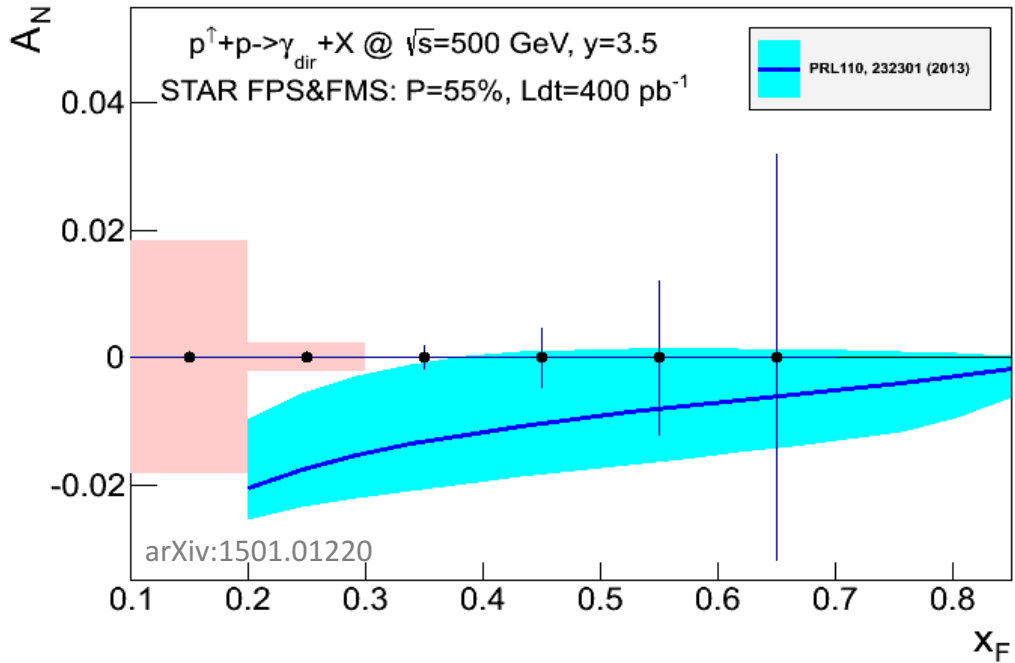
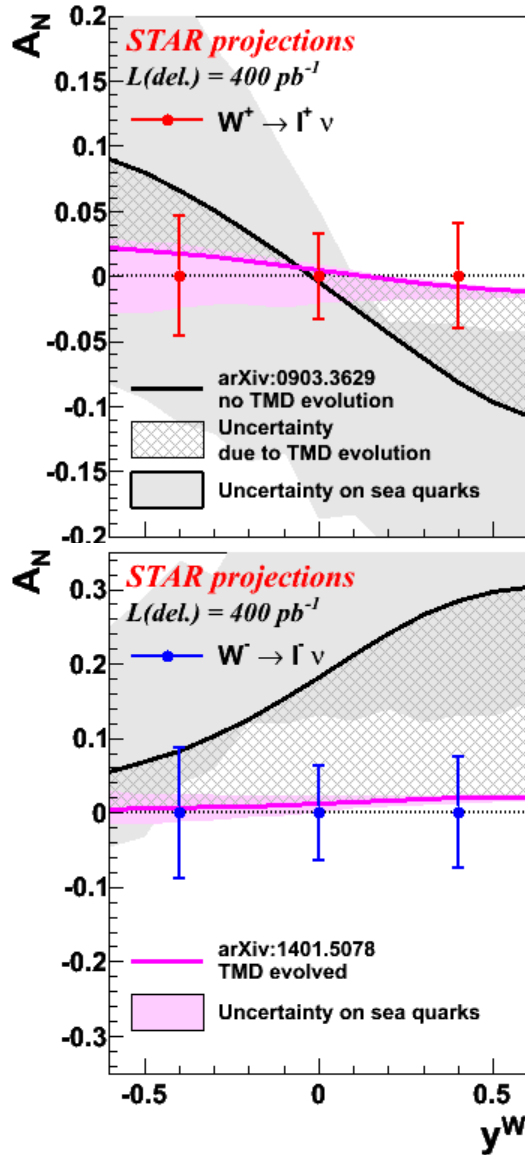


First Measurement of A_N for Weak Bosons!

Global fit to the (unevolved) KQ prediction:

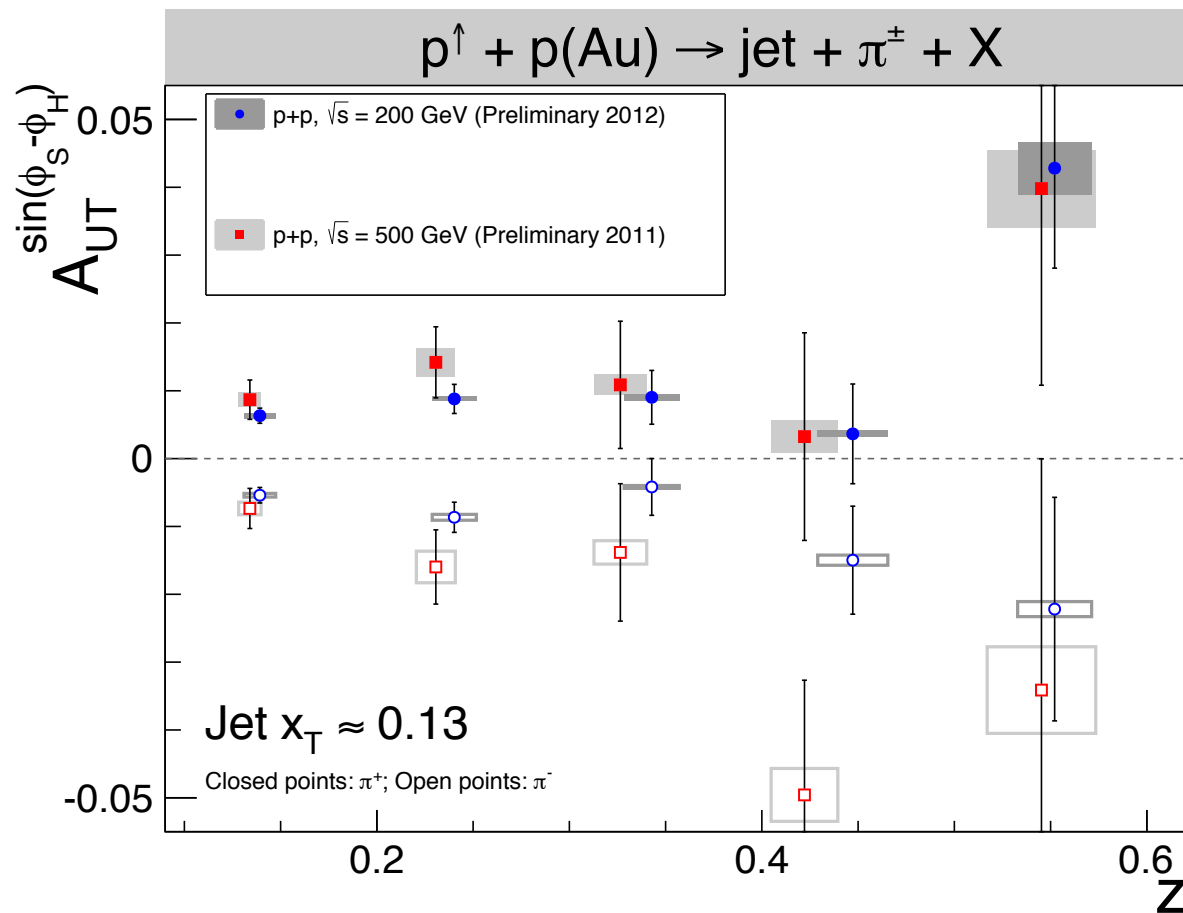
- **solid line:** assume Sivers sign change: $\chi^2/\nu = 7.4/6$
- **dashed line:** assume no sign change: $\chi^2/\nu = 19.6/6$

The Near-term Future: Sivers Sign-change+Evolution



STAR currently taking data at $\sqrt{s} = 510 \text{ GeV}$!
 $W^{\pm} A_N$ can be sensitive to Sivers sign-change if TMD-evolution suppression factor ~ 5 or less
Evaluate sign-change+evolution through W^{\pm}/Z , forward direct- γ (twist-3), and forward Drell-Yan
Forward direct- γ at 200 GeV already in the bag!

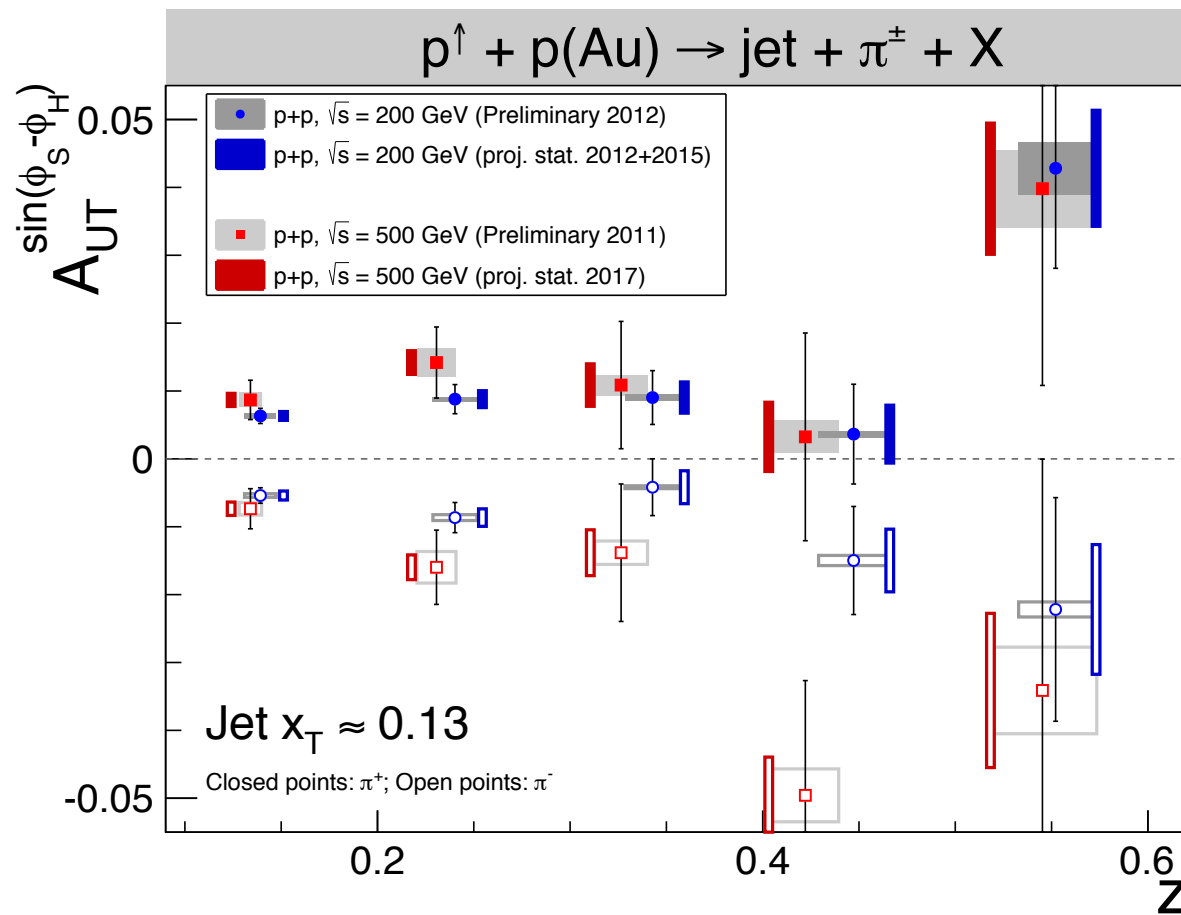
The Near-term Future: Collins Evolution



Preliminary 2011 and 2012 Collins asymmetries suggest x_T scaling

Implications for TMD evolution?

The Near-term Future: Collins Evolution

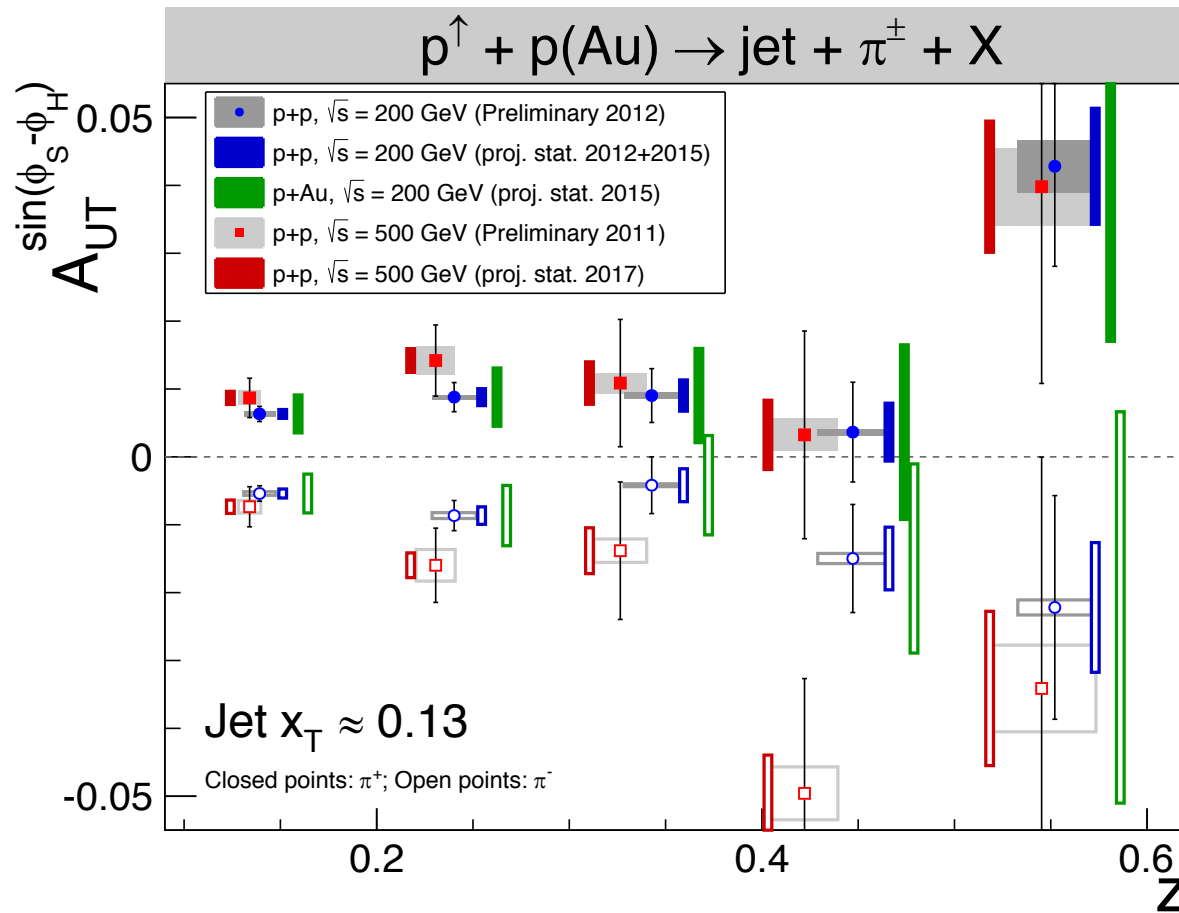


Higher precision in 2015 and 2017 will allow more precise comparison!

Preliminary 2011 and 2012 Collins asymmetries suggest x_T scaling

Implications for TMD evolution?

The Near-term Future: $p + A$ Collins



Higher precision in 2015 and 2017 will allow more precise comparison!

First $p^\uparrow + Au$ run!
Should allow for first glimpse of Collins in $p + A$
 \rightarrow Explore hadronization

Preliminary 2011 and 2012 Collins asymmetries suggest x_T scaling
Implications for TMD evolution?

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 - Possible x_T scaling for Collins and IFF asymmetries
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- **Recent and near-future runs offer even more potential**
 - Investigation of Sivers/twist-3 in W , γ , and Drell-Yan
 - Substantially increased precision for Collins and IFF at 200 and 510 GeV
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Stay tuned for more new results from RHIC transverse spin!

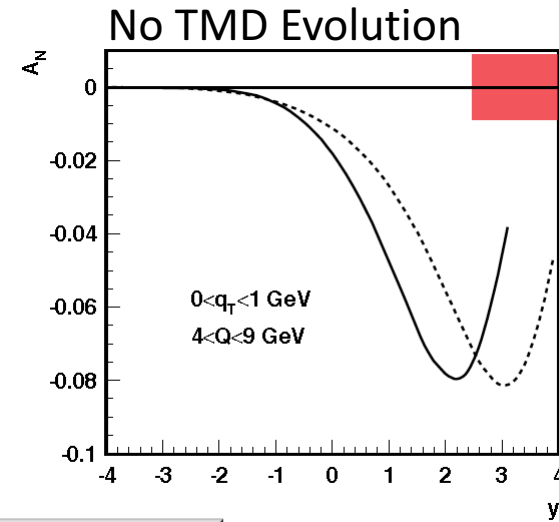
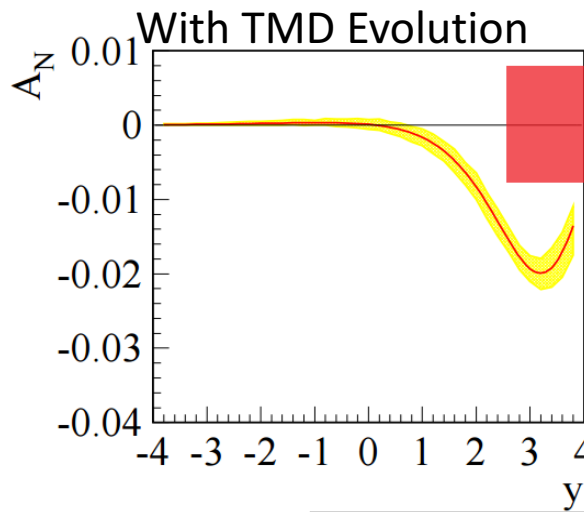
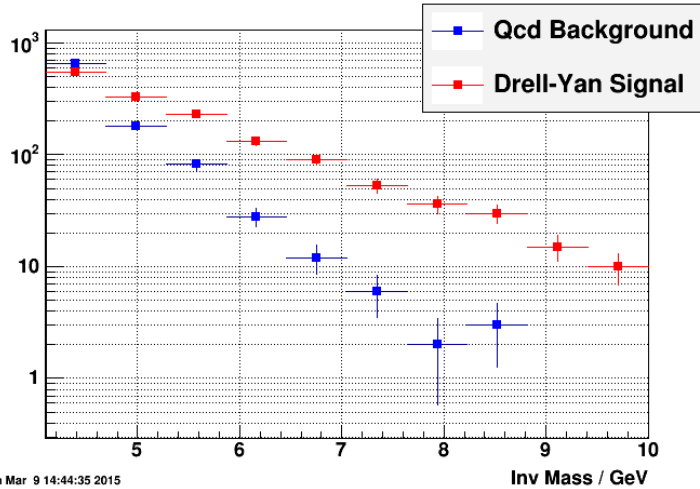
Back-up Slides

Drell-Yan at STAR For 500 GeV

Kinematics:

DY e^+e^- in $2.5 < \eta < 4.0$ $4.0 < M_{e^+e^-} < 9.0 \text{ GeV}/c^2$

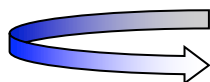
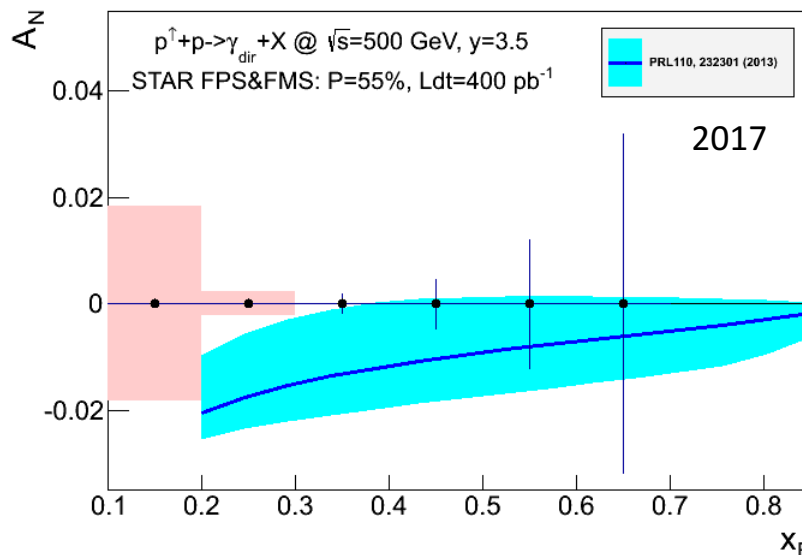
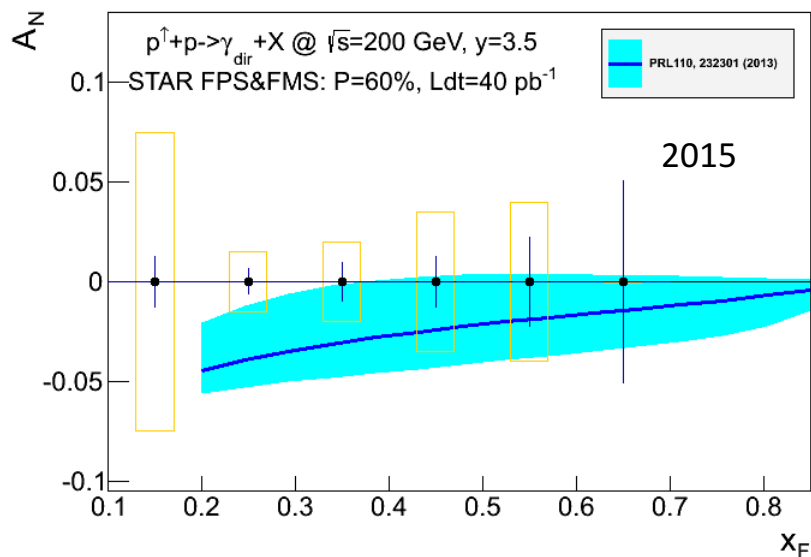
After analysis $2.5 < \eta < 4.0$:



Run-17 $\int L_{\text{del}} = 400 \text{ pb}^{-1}$
 $\rightarrow A_N$ for DY to ± 0.008

Assembled by
E.C. Aschenauer

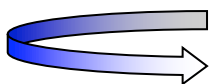
Complementary Channel: A_N direct photon



A_N for direct photon production:

- sensitive to sign change, but in TWIST-3 formalism
- not sensitive to TMD evolution
- no sensitivity to sea-quarks; mainly u_v and d_v at high x
- collinear objects but more complicated evolutions than DGLAP
- indirect constraint on Sivers fct.

$$-\int d^2 k_{\perp} \frac{|k_{\perp}^2|}{M} f_{1T}^{\perp q}(x, k_{\perp}^2)|_{SIDIS} = T_{q,F}(x, x)$$



Not a replacement for $A_N(W^{+/-}, Z^0, DY)$ measurement
but an important complementary piece in the puzzle

Assembled by
E.C. Aschenauer