# A deeper probe of the Higgs sector

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The Standard Model Higgs and beyond
 Tests of the Higgs properties
 Interludium: D<sub>γγ</sub>
 Direct search for new states
 Conclusion

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#### The Standard Model Higgs and beyond



A very non-trivial check of the SM: test at the quantum/permille level: – constraints from data:  $M_H = 92^{+34}_{-26}$  GeV  $\leq 160$  GeV at 95% CL – experimentaly found to be:  $M_H = 125.1 \pm 0.24$  GeV (ie within  $1\sigma$ ..) In addition, it looks as it has the properties of the SM Higgs state: The triumph of the SM model of particle physics; Standarissimo?! Tokyo 9/01/2017 A deeper probe of the Higgs sector – A. Djouadi – p.2/31

#### . The Standard Model Higgs and beyond

We have a theory for the strong electroweak forces, the SM, that is:

• a relativistic quantum field theory based on a gauge symmetry, • renormalisable as proved by 't Hooft and Veltman for SEWSB, • unitary as we have now a Higgs and its mass is rather small, • perturbative up to the Planck scale as again the Higgs is light, • leads to a (meta)stable electroweak vacuum up to high scales, • compatible with (almost) all precision data available to date... s the SM the "theory of everything" and should we be satisfied with it? **No! Low** energy manifestation of a fundamental theory that solves: • "Esthetical" problems with e.g. multiple and arbitrary parameters; **gauge c**oupling unification:  $3 \neq g_i$  which do not meet a high scale. • "Experimental" problems as it does not explain all seen phenomena:  $\nu$  masses/mixing, dark matter, baryon asymmetry in the universe .... (Note: SO(10) at intermediate  $Q = 10^{11}$ GeV and axions cure these pbs) • "Theory" (or consistency) problem: the hierarchy/naturalness pbs.  $\Delta M_{\rm H}^2 \propto \Lambda^2 \approx (10^{18} {\rm ~GeV})^2$ :  $M_{\rm H}$  not stable against high scales. All these indicate that there is beyond the Standard Model (?). **Tokyo 9/01/2017** A deeper probe of the Higgs sector – A. Djouadi – p.3/31

#### . The **S**tandard Model Higgs and beyond nain avenues for solving the hierarchy or naturalness problems

I. Compositeness/substructure: All particles are composite: Technicolor  $\Rightarrow$  **H** bound state of two fermions (no more spin-0 fundamental state). **II. Extra space–time dimensions** where at least s=2 gravitons propagate.  $\Rightarrow$  effective gravity scale  $\Lambda \approx$  . **EWSB** mechanism needed: H or not H! **III.** Supersymmetry: doubling the world. - links  $s=\frac{1}{2}$  fermions to s=1 bosons, – links internal/space-time symmetries, - if made local, provides link to gravity, **– natu**ral  $\mu^2 < 0$ : radiative EWSB,  $\Rightarrow$  sparticle loops cancel  $\Lambda^2$  behavior extend EWSB sector: at least 2 doublets. **Tokyo 9/0**1/2017 A deeper probe of the Higgs sector – A. Djouadi – p.4/31





#### . The Standard Model Higgs and beyond <u>The problem is that:</u>

#### ve observe a Higgs boson with a mass of 125 GeV and no other Higgs:

 $\sigma \times BR$  rates compatible with those expected in the SM Fit of all LHC Higgs data  $\Rightarrow$ agreement at 15–30% level Results from the LHC 7–8 TeV campaign already give us:  $\mu_{tot}^{ATLAS} = 1.18 \pm 0.15$  $\mu_{tot}^{CMS} = 1.00 \pm 0.14$ 



we do not observe any new particle beyond those of SM with Higgs:
protund implications for most discussed BSM scenarios; they are in:
"Mortuary": Higgsless, 4th generation, fermio or gauge-phobic..
"Hospital": Technicolor, composite models (but see Christophe) ....
"Trouble" and strongly constrained: extra-dimensions, SUSY, ...
s an example, let us see what it implies for SUSY and the MSSM.
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. The Standard Model Higgs and beyond In the MSSM we need two doublets of complex scalar fields H<sub>1</sub> and H<sub>2</sub> to generate up/down-type fermion masses and no chiral anomalies. after EWSB, three dof for  $W_{L}^{\pm}$ ,  $Z_{L} \Rightarrow 5$  physical states:  $h, H, A, H^{\pm}$ . Only two free parameters at tree-level to describe the system  $an\beta, M_A$  $M_{h H}^{2} = \frac{1}{2} \left\{ M_{A}^{2} + M_{Z}^{2} \mp \left[ (M_{A}^{2} + M_{Z}^{2})^{2} - 4M_{A}^{2}M_{Z}^{2}\cos^{2}2\beta \right]^{1/2} \right\}$  $\mathrm{M}^2_{\mathbf{H}^\pm} = \mathrm{M}^2_{\mathbf{A}} + \mathrm{M}^2_{\mathbf{W}}$  $\tan 2\alpha = \frac{-(\mathbf{M}_{\mathbf{A}}^2 + \mathbf{M}_{\mathbf{Z}}^2)\sin 2\beta}{(\mathbf{M}_{\mathbf{A}}^2 - \mathbf{M}_{\mathbf{A}}^2)\cos 2\beta} = \tan 2\beta \frac{\mathbf{M}_{\mathbf{A}}^2 + \mathbf{M}_{\mathbf{Z}}^2}{\mathbf{M}_{\mathbf{A}}^2 - \mathbf{M}_{\mathbf{Z}}^2} \ \left(-\frac{\pi}{2} \le \alpha \le \mathbf{0}\right)$  $M_h \lesssim M_Z |\cos 2\beta| + RC \lesssim 130 \text{ GeV}, M_H \approx M_A \approx M_{H^{\pm}} \lesssim M_{EWSB}.$ • Couplings of h, H to VV are suppressed; no AVV couplings (CP). • For  $t an\beta \gg 1$ : couplings to b (t) quarks enhanced (suppressed).  $A \qquad 1/\tan\beta \qquad \tan\beta$ In decoupling limit: MSSM Higgs sector reduces to SM with a light h. Tokyo 9/01/2017 A deeper probe of the Higgs sector – A. Djouadi – p.6/31



Arbey, Battaglia, AD, Mahmoudi, Quevillon (2012) Arbey, Battaglia, AD, Mahmoudi, Quevillon (2012) A general MSSM and higher in constrained models. Tokyo 9/01/2017 A deeper probe of the Higgs sector – A. Djouadi – p.7/31

#### . The Standard Model Higgs and beyond

This is backed up by direct searches of SUSY particles at the LHC:

#### the SUSY scale $M_{SUSY} \gtrsim O(1 \text{ TeV})$ in most experimental searches..

	Model	$e, \mu, \tau, \gamma$	Jets	E <sup>miss</sup> <sub>T</sub>	$\int \mathcal{L} dt [\mathbf{fb}]$	Mass limit	Reference
S	MSUGRA/CMSSM $\tilde{q}\tilde{q}, \tilde{q} \rightarrow q \tilde{\chi}_{1}^{0}$ (compressed)	0 0 1 x	2-6 jets 2-6 jets 0-1 jet	Yes Yes	20.3 20.3 20.3	<i>q̃. ĝ</i> <i>q̃. ĝ</i> <i>q̃. ĝ</i> <i>q̃. ĝ</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i> <i>s</i>	1405.7875 1405.7875 1411 1559
arche	$\widetilde{g}\widetilde{g}, \widetilde{g} \rightarrow q\widetilde{q}\widetilde{\chi}_{1}^{0}$ $\widetilde{g}\widetilde{g}, \widetilde{g} \rightarrow q\widetilde{q}\widetilde{\chi}_{1}^{0}$ $\widetilde{g}\widetilde{g}, \widetilde{g} \rightarrow qq\widetilde{\chi}_{1}^{\pm} \rightarrow qqW^{\pm}\widetilde{\chi}_{1}^{0}$	0 1 e,µ	2-6 jets 3-6 jets	Yes	20.3 20	k         1.33 TeV         m(k <sup>*</sup> ) <sub>1</sub> = 0.6(x <sup>*</sup> )           k         1.2 TeV         m(k <sup>*</sup> ) <sub>1</sub> = 0.5(m(k <sup>*</sup> ) <sub>1</sub> +m(z))	1405.7875 1501.03555
ve Se	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq(\ell\ell/\ell\nu/\nu\nu)\tilde{\chi}_1^0$ GMSB ( $\tilde{\ell}$ NLSP)	2 e,μ 1-2 τ + 0-1 ℓ	0-3 jets 0-2 jets	Yes	20 20.3	ĝ         1.32 TeV         m(ξ <sup>2</sup> )=0 GeV           ĝ         1.6 TeV         tanβ >20	1501.03555 1407.0603
Inclus	GGM (bino NLSP) GGM (wino NLSP) GGM (biggsing-bing NLSP)	$2\gamma$ 1 $e, \mu + \gamma$	- - 1 h	Yes Yes Yes	20.3 4.8	<u>8 1.28 TeV</u> m(X <sup>*</sup> )⊳50 GeV <u>8 619 GeV</u> m(X <sup>*</sup> )⊳50 GeV <u>900 GeV</u> m(X <sup>*</sup> )⊳20 GeV	ATLAS-CONF-2012-14- ATLAS-CONF-2012-14- 1211 1167
	GGM (higgsino NLSP) Gravitino LSP	2 e, μ (Z) 0	0-3 jets mono-jet	Yes Yes	5.8 20.3	%         690 GeV         mk(1)=210 GeV           %         690 GeV         m(NLSP)=200 GeV           F <sup>1/2</sup> scale         865 GeV         m(G)>1.8 × 10 <sup>-4</sup> eV, m(g)=1.5 TeV	ATLAS-CONF-2012-152 1502.01518
3 <sup>rd</sup> gen. Ĩ med.	$\tilde{g} \rightarrow b \bar{b} \tilde{\chi}_{1}^{0}$ $\tilde{g} \rightarrow t \bar{t} \tilde{\chi}_{1}^{0}$	0	3 b 7-10 jets	Yes Yes	20.1 20.3	β         1.25 TeV         m(k <sup>2</sup> <sub>1</sub> )×400 GeV           ĝ         1.1 TeV         m(k <sup>2</sup> <sub>1</sub> )<50 GeV	1407.0600 1308.1841
	$\tilde{g} \rightarrow t \bar{t} \tilde{\chi}_{1}^{0}$ $\tilde{g} \rightarrow b t \tilde{\chi}_{1}^{+}$	0-1 <i>e</i> ,μ 0-1 <i>e</i> ,μ	3 b 3 b	Yes Yes	20.1 20.1	ĝ         1.34 TeV         m(ξ <sup>0</sup> )<400 GeV           ĝ         1.3 TeV         m(ξ <sup>0</sup> )<300 GeV	1407.0600 1407.0600
arks tion	$ \begin{array}{l} \tilde{b}_1 \tilde{b}_1,  \tilde{b}_1 \rightarrow b \tilde{\chi}_1^0 \\ \tilde{b}_1 \tilde{b}_1,  \tilde{b}_1 \rightarrow t \tilde{\chi}_1^\pm \end{array} $	0 2 <i>e</i> , μ (SS)	2 b 0-3 b	Yes Yes	20.1 20.3	μ         100-620 GeV         m( $\tilde{k}_1^0$ )<90 GeV           μ         275-440 GeV         m( $\tilde{k}_1^2$ )=2 m( $\tilde{k}_1^0$ )	1308.2631 1404.2500
. squa	$\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow b \tilde{\chi}_1^{\pm}$ $\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow W b \tilde{\chi}_1^0 \text{ or } t \tilde{\chi}_1^0$	1-2 e,μ 2 e,μ	1-2 b 0-2 jets	Yes Yes	4.7 20.3	Ji         110-167 GeV         230-460 GeV         m(k <sup>2</sup> <sub>1</sub> ) = 2m(k <sup>0</sup> <sub>1</sub> ), m(k <sup>2</sup> <sub>1</sub> ) = 55 GeV           i         90-191 GeV         215-530 GeV         m(k <sup>2</sup> <sub>1</sub> ) = 16 eV           i         90-191 GeV         m(k <sup>2</sup> <sub>1</sub> ) = 16 eV         i	1209.2102, 1407.0583 1403.4853, 1412.4742
<sup>rd</sup> gen irect p	$t_1 t_1, t_1 \rightarrow \mathcal{X}_1$ $\tilde{t}_1 \tilde{t}_1, \tilde{t}_1 \rightarrow c \tilde{\chi}_1^0$ $\tilde{t}_1 \tilde{t}_1 (natural GMSB)$	0 m 2 e, μ (Z)	ono-jet/c- 1 b	tag Yes Yes	20.3 20.3	r1         210-500 GeV         m(x) = 166V <i>i</i> _1         90-240 GeV         m(i) > 85 GeV <i>i</i> _1         150-580 GeV         m(i) > 85 GeV	1407.0608 1403.5222
σ	$\tilde{t}_2 \tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$ $\tilde{t}_2 = \tilde{t}_2 - \tilde{t}_2 \rightarrow \tilde{t}_1^0$	3 e, μ (Z)	1 <i>b</i>	Yes	20.3	Ž20-600 GeV         m(ξ <sup>0</sup> )/200 GeV           200-600 GeV         m(ξ <sup>0</sup> )/200 GeV	1403.5222
ot _	$\begin{split} &\tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow \tilde{\ell}\nu(\ell\tilde{\nu}) \\ &\tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow \tilde{\ell}\nu(\ell\tilde{\nu}) \\ &\tilde{\chi}_{1}^{+}\tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow \tilde{\tau}\nu(\tau\tilde{\nu}) \end{split}$	2 e,μ 2 τ	0	Yes	20.3 20.3	λ         140-465 GeV         m(k1)=0 GeV, m(k, k)=0.5(m(k1)+m(k1))           k1         100-350 GeV         m(k1)=0 GeV, m(k, k)=0.5(m(k1)+m(k1))           k1         100-350 GeV         m(k1)=0 GeV, m(k, k)=0.5(m(k1)+m(k1))	1403.5294 1407.0350
E M dire	$\tilde{\chi}_{1}^{\pm}\tilde{\chi}_{2}^{0} \rightarrow \tilde{\ell}_{L} \nu \tilde{\ell}_{L} \ell(\tilde{\nu}\nu), \ell \tilde{\nu} \tilde{\ell}_{L} \ell(\tilde{\nu}\nu)$ $\tilde{\chi}_{1}^{\pm}\tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} Z \tilde{\chi}_{1}^{0}$ $\tilde{\chi}_{1}^{\pm} \tilde{\chi}_{2}^{0} \rightarrow W \tilde{\chi}_{1}^{0} Z \tilde{\chi}_{1}^{0}$	3 e,μ 2-3 e,μ	0 0-2 jets	Yes Yes	20.3 20.3	ξ <sup>2</sup> <sub>1</sub> , k <sup>2</sup> <sub>2</sub> 700 GeV         m(k <sup>2</sup> <sub>1</sub> )=m(k <sup>2</sup> <sub>2</sub> ), m(k <sup>2</sup> <sub>1</sub> )=0.5(m(k <sup>2</sup> <sub>1</sub> )+m(k <sup>2</sup> <sub>1</sub> ))           ξ <sup>2</sup> <sub>1</sub> , k <sup>2</sup> <sub>2</sub> 420 GeV         m(k <sup>2</sup> <sub>1</sub> )=m(k <sup>2</sup> <sub>1</sub> ), m(k <sup>2</sup> <sub>1</sub> )=0.5(m(k <sup>2</sup> <sub>1</sub> )+m(k <sup>2</sup> <sub>1</sub> ))           ξ <sup>2</sup> <sub>1</sub> , k <sup>2</sup> <sub>2</sub> 420 GeV         m(k <sup>2</sup> <sub>1</sub> )=m(k <sup>2</sup> <sub>1</sub> ), m(k <sup>2</sup> <sub>1</sub> )=0.5(m(k <sup>2</sup> <sub>1</sub> )+m(k <sup>2</sup> <sub>1</sub> ))	1402.7029 1403.5294, 1402.7029
_	$\begin{array}{c} \chi_1 \chi_2 \rightarrow W \chi_1 h \chi_1, h \rightarrow b b / W W / \tau \tau / \\ \tilde{\chi}_2^0 \tilde{\chi}_3^0, \tilde{\chi}_{2,3}^0 \rightarrow \tilde{\ell}_{\mathrm{R}} \ell \end{array}$	γγ <i>e</i> ,μ,γ 4 <i>e</i> ,μ	0-2 0	Yes	20.3	At μX z         250 GeV         m(X) = m(X2), m(X) = 0, sleptons decoupled           K <sup>1</sup> <sub>2</sub> z         620 GeV         m(K <sup>2</sup> <sub>2</sub> )=m(K <sup>2</sup> <sub>3</sub> ), m(K <sup>2</sup> <sub>1</sub> )=0, sleptons decoupled	1501.07110 1405.5086
ong-lived	Direct $\chi_1^+ \chi_1^-$ prod., long-lived $\chi_1^+$ Stable, stopped $\tilde{g}$ R-hadron Stable $\tilde{g}$ R-hadron	Disapp. trk 0 trk	1 jet 1-5 jets	Yes	20.3 27.9	χ₁         270 GeV         m(ξ1)/m(ζ1)/=160 MeV, τ(ζ1)/=0.2 ns           χ̄         832 GeV         m(ζ1)/=100 GeV, 10 μs <r(ζ)< th="">         m(ζ1)/=100 GeV, 10 μs<r(ζ)< th="">           φ         1 27 TaV         1         100 GeV, 10 μs<r(ζ)< th="">         100 GeV, 10 μs<r(ζ)< th="">         100 GeV, 10 μs         100 GeV, 10 μs</r(ζ)<></r(ζ)<></r(ζ)<></r(ζ)<>	1310.3675 1310.6584 1411.6795
	GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu}) + \tau(e, \tilde$	,μ) 1-2 μ 2 γ	-	- Yes	19.1 20.3	δ         537 GeV         10 <tar th="" β<50<=""> <math>k_{\mu}^{\mu}</math>         435 GeV         2<rt <math="">k_{\mu}^{\mu})         3 ns, SPS8 model</rt></tar>	1411.6795 1409.5542
_	$\tilde{q}\tilde{q}, \tilde{\chi}_{1}^{0} \rightarrow qq\mu \text{ (RPV)}$ LFV $pp \rightarrow \tilde{\gamma}_{\tau} + X, \tilde{\gamma}_{\tau} \rightarrow e + \mu$	1 μ, displ. vtx	-	-	20.3	ĝ         1.0 TeV         1.5 <cr<156 br(μ)="1," m(ξ<sup="" mm,="">0)=108 GeV           p.         1.61 TeV         λ',, =0.10, λ<sub>1</sub>, =0.05</cr<156>	ATLAS-CONF-2013-09
>	LFV $pp \rightarrow \tilde{v}_{\tau} + X, \tilde{v}_{\tau} \rightarrow e(\mu) + \tau$ Bilinear RPV CMSSM	1 e, μ + τ 2 e, μ (SS)	- 0-3 b	- Yes	4.6 20.3	γ.         1.1 TeV         λ <sub>11</sub> 0.10. λ <sub>1(D)10</sub> =0.05           φ. φ         1.35 TeV         M(g)=m(g), cT_SP<1 mm	1212.1272 1404.2500
RP	$ \begin{split} \tilde{\chi}_{1}^{+} \tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow W \tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow e e \tilde{\nu}_{\mu}, e \mu \tilde{\nu}_{e} \\ \tilde{\chi}_{1}^{+} \tilde{\chi}_{1}^{-}, \tilde{\chi}_{1}^{+} \rightarrow W \tilde{\chi}_{1}^{0}, \tilde{\chi}_{1}^{0} \rightarrow \tau \tau \tilde{\nu}_{e}, e \tau \tilde{\nu}_{\tau} \end{split} $	4 e,μ 3 e,μ + τ	-	Yes Yes	20.3 20.3	k <sup>1</sup> 750 GeV m(k <sup>2</sup> )>0.2×m(k <sup>2</sup> ),λ <sub>121</sub> ≠0 k <sup>2</sup> 450 GeV m(k <sup>2</sup> )>0.2×m(k <sup>2</sup> ),λ <sub>121</sub> ≠0 m(k <sup>2</sup> )>0.2×m(k <sup>2</sup> ),λ <sub>121</sub> ≠0	1405.5086 1405.5086
	$g \rightarrow qqq$ $\tilde{g} \rightarrow \tilde{t}_1 t, \tilde{t}_1 \rightarrow bs$	0 2 <i>e</i> , μ (SS)	6-7 jets 0-3 b	Yes	20.3 20.3	g         916 GeV         BH(p)=BR(c)=0%           ž         850 GeV	AILAS-CONF-2013-09 1404.250
Othor	Scalar charm, $\tilde{c} \rightarrow c \tilde{\chi}_{1}^{0}$	0	2 c	Yes	20.3		1501.01325

\*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 $\sigma$  theoretical signal cross section uncertainty.

#### ⇒ ATLAS/CMS depressing tables (update: B. Petersen).... Tokyo 9/01/2017 A deeper probe of the Higgs sector – A. Djouadi – p.8/31

#### The Standard Model Higgs and beyond Also backed up indirectly by the measurement of the Higgs properties: fits of the h couplings $\Rightarrow$ constraints on the MSSM $[M_A, \tan\beta]$ plane.



with dominant RC only (see e.g. above for  $M_h$ ) the h couplings read:  $\mathbf{g_{h\bar{t}t}} = \cos \alpha / \sin \beta$  $\mathbf{g}_{\mathbf{h}\bar{\mathbf{b}}\mathbf{b}} = \cos\alpha / \sin\beta$  $\mathbf{g}_{\mathbf{h}\mathbf{V}\mathbf{V}} = \sin(\beta - \alpha)$  $\alpha \approx \mathbf{f}(\tan\beta, \mathbf{M}_{\mathbf{A}}, \mathbf{M}_{\mathbf{h}})$ like  $M_H$  and  $M_H^{\pm}$ as in so-called hMSSM **AD**, Quevillon, Maiani ... (2013) – A. Djouadi – p.9/31

he next question is then. "is Particle Physics closed"? Answer is not ) Need to check that H is indeed responsible of EWSB (and SM-like?) — measure its fundamental properties in the most precise way: • its mass and total decay width (invisible width due to dark matter?), • its spin-parity quantum numbers (CP violation for baryogenesis?), • its couplings to fermions and gauge bosons and check if they are only proportional to particle masses (no new physics contributions?), • its self-couplings to reconstruct the potential  $V_S$  that makes EWSB. Possible for  $M_H \approx 125$  GeV as all production/decay channels useful.





- A. Djouadi - p.10/31

#### . Tests of the Higgs properties A check of spin parity quantum numbers

**Spin:** clear situation (no suspense) as the new state decays into  $\gamma\gamma \Rightarrow$  not s=1 from Landau–Yang and s=2 (KK graviton?) unlikely..

**Prombers: CP-even, CP-odd, or mixture? more important issue: CPV in Higgs sector.) ATLAS and CMS MELA analyses for pure CP → pure CP-even favored at** ≈ 3*σ* level. **But problems with this (too simple) picture: pure CP-odd does not couple to VV@tree-level.** 

Indirect probe: through  $\hat{\mu}_{ZZ} = 1.1 \pm 0.4$   $g_{HVV} = c_V g_{\mu\nu}$  gives upper bound on CP  $\eta_{CP} \equiv 1 - c_V^2 \gtrsim 0.5@68\% CL$ Direct probe:  $g_{Hff}$  more democratic.

 $\begin{array}{ll} \mbox{spin-correlations in $q\bar{q} \rightarrow HZ \rightarrow b\bar{b}ll$\\ \mbox{or later in $q\bar{q}/gg \rightarrow Ht\bar{t} \rightarrow b\bar{b}t\bar{t}.$\\ \hline \mbox{Lxtremely challenging even at $HL-LHC...$\\ \hline \mbox{Fokyo $9/01/2017} & A deeper probe of the Higgs sector } \end{array}$ 





• A much more precise measurement of the **n various H production+decay channels But rather large errors mainly due to:** - experimental: stats, system., lumi... - theory: PDFs, HO/scale, jetology... total error about 15–20% in  $\mathrm{gg} 
ightarrow \mathrm{H}$ Hjj contaminates VBF (now 30%)..  $\Rightarrow$  ratios of  $\sigma$ xBR: many errors out! Deal with width ratios  $\Gamma_{\mathbf{X}}/\Gamma_{\mathbf{Y}}$ - **TH** on  $\sigma$  and some EX errors - parametric errors in BRs - TH ambiguities from  $\Gamma_{\rm H}^{\rm tot}$  $c_f$ • Achievable accuracy: -0.3- now: 20–30% on some ratios -1.0

- future: few % at HL–LHC?



But will this be sufficient to probe BSM physics (at.high scales..)? Fokyo 9/01/2017 A deeper probe of the Higgs sector – A. Djouadi – p.12/31

- Total width:  $\Gamma_{\rm H} = 4$  MeV, too small to be resolved experimentally.
- **very l**oose bound from interference  $gg \rightarrow ZZ$  (a factor 2–5 at most).
- **no way to access it indirectly (via production rates) in a precise way.**
- Invisible decay width: more easily accessible at the LHC

#### **Direct m**easurement:

 $q\bar{q} \rightarrow HZ$  and  $qq \rightarrow Hqq$ ;  $H \rightarrow inv$ Combined HZ+VBF search from CMS  $BR_{inv} \lesssim 50\% @95\%$  CL for SM Higgs Also promising in the future: monojets

# $\mathbf{gg} \rightarrow \mathbf{H} + \mathbf{j} \rightarrow \mathbf{j} + \mathbf{E}_{\mathbf{T}}$

**Indirect** measurement:

again assume SM-like Higgs couplings constrain width from signal strengths  $BR_{inv} \lesssim 50\%@95\%CL$  for  $c_f = c_V = 1$ Improvement in future: 10%@HL-LHC?

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⇒ difficult to be beaten by anything else for ≈ 125 GeV Higgs ⇒ welcome to the e<sup>+</sup>e<sup>-</sup> precision machine! But let's get back to the near future: what can we do at HL-LHC? Tokyo 9/01/2017 A deeper probe of the Higgs sector – A. Djouadi – p.15/31

#### . Interludium: $D_{\gamma\gamma}$

• Precise measurement of Higgs couplings in various H channels:

example of the cleanest detection channels:  $\mathbf{H} \to \overline{\gamma\gamma}, \ \mathbf{H} \to \mathbf{ZZ}^* \to 4\ell^{\pm}$ 

<b>chan</b> nel	atlas	cms
$\mu_{\gamma\gamma}$	$1.17 \ {}^{+0.23}_{-0.23} \ {}^{+0.16}_{-0.11} \ \ ({}^{+0.12}_{-0.08})$	$1.14 egin{array}{cccc} +0.21 & +0.16 & (+0.09) \ -0.21 & -0.10 & (-0.05) \end{array}$
$\mu_{\mathbf{Z}\mathbf{Z}}$	$1.46 \ {}^{+0.35}_{-0.31} \ {}^{+0.19}_{-0.13} \ ({}^{+0.18}_{-0.11})$	$0.93 \ {}^{+0.26}_{-0.23} \ {}^{+0.13}_{-0.09}$

**Is** this enough to probe effects of new physics or BSM? ot in the case of weakly interacting theories like 2HDM, SUSY, etc... expect effects at  $\approx \frac{C_{new} \alpha_W}{\pi} \approx \frac{M_h^2}{M^2} \approx 1\%$ ; ATLAS Simulation  $\sqrt{s} = 14 \text{ TeV}: \left[ \text{Ldt} = 300 \text{ fb}^{-1}; \right] \text{Ldt} = 3000 \text{ fb}^{-1}$ is 1% accuracy achievable at HL-LHC (3ab<sup>-1</sup>)? fb<sup>-1</sup> extrapolated from 7+8 TeV H→µµ • Statistical error:  $20\%/\sqrt{3 \times 100} \lesssim 1-2\%$ ttH,H→μμ VBF.H→ττ (projection OK with ATLAS+CMS combo)  $H \rightarrow ZZ$ • Systematical error: can be made  $\leq 1\%$ ? VBF,H→ WW  $H \rightarrow WW$ some errors are common (luminosity, etc....).  $VH, H \rightarrow \gamma \gamma$ ttH,H→γγ • Theoretical uncertainty (if it is  $\gg 1\%$ ):  $VBF,H\rightarrow\gamma\gamma$ will be then by far the crucial/limiting issue! Н→үү (+j) Η→γγ How big is it? Can it be reduced? Removed? 0.2 0.4 0.6 0.8 0 A deeper probe of the Higgs sector – A. Djouadi – p.16/31 **Fokyo 9/01/2017** 

#### . Interludium: $D_{\gamma\gamma}$

LO<sup>*a*</sup>: already at one loop CD: exact NLO<sup>*b*</sup>: K  $\approx$  1.7 EFT NLO<sup>*c*</sup>: good approx. EFT NNLO<sup>*d*</sup>: K  $\approx$  2 EFT NNLL<sup>*e*</sup>:  $\approx$  + (5%) EFT N3LO<sup>*f*</sup>:  $\approx$  3%. W: EFT NLO: <sup>*g*</sup>:  $\approx$  ± very small exact NLO<sup>*h*</sup>:  $\approx$  ± a few % QCD+EW<sup>*i*</sup>: a few % istributions: a few programs<sup>*j*</sup>

<sup>a</sup>Georgi+Glashow+Machacek+Nanopoulos <sup>b</sup>Spira+Graudenz+Zerwas+AD (exact)<sub>2</sub> <sup>c</sup>Spira+Zerwas+AD; Dawson (EFT) <sup>d</sup>Harlander+Kilgore, Anastasiou+Melnik <sup>c</sup>Catani+de Florian+Grazzini+Nason <sup>f</sup>Anustasiou et al. (2015)! <sup>g</sup>Gantoino+AD; Degrassi et al. <sup>h</sup>Actis+Passarino+Sturm+Uccirati <sup>i</sup>Anustasiou et al.; Grazzini, Nason... **Fokyo 9/01/2017** A deeper probe of the Higgs sector



#### **Interludium:** $D_{\gamma\gamma}$ **Despite of that.** the gg $\rightarrow$ H cross section still affected by uncertainties

Higher-order or scale uncertainties:
 K-factors large ⇒ HO could be important
 HO estimated by varying scales of process

 $\mu_0/\kappa \leq \mu_{\mathbf{R}}, \mu_{\mathbf{F}} \leq \kappa \mu_0$ at IHC:  $\mu_0 = \frac{1}{2}M_H, \kappa = 2 \Rightarrow \Delta_{\text{scale}}^{\text{NNLO}} \approx 10\%$ • gluon PDF+associated  $\alpha_s$  uncertainties: **PDF(g)** at high-x less constrained by data  $\alpha_s$  uncertainty (WA, DIS?) affects  $\sigma \propto \alpha_s^2$  $\Rightarrow$  still discrepancies between NNLO PDFs **PDF4LHC recommend:**  $\Delta_{pdf} \approx 10\%$  @lHC Uncertainty from EFT approach at NNLO  $m_{loop} \gg M_H$  good for top if  $M_H \lesssim 2m_t$ but not for b ( $\approx 10\%$ ) and W/Z loops **Estimate from (exact) NLO:**  $\Delta_{\rm EFT} \approx 5\%$ total  $\Delta \sigma^{
m NNLO}_{
m gg 
ightarrow 
m H 
ightarrow 
m X} \approx 10-20\%$  @lHC **LHC-HxsWG: Baglio+AD**  $\Rightarrow$ A deeper probe of the Higgs sector <mark>Fokyo 9/0</mark>1/2017



#### **Interludium:** $D_{\gamma\gamma}$ **Production cross sections**

**Fokyo 9/01/2017** 

 $gg \rightarrow H$  by far dominant process  $(\approx 85\%$  of the events before cuts) 100  $\Rightarrow O(10\%)$  total TH uncertainty ..... 10 **followed by cleaner VBF+VH modes:** only  $\leq 15\%$  of rate before cuts... smaller TH error only for inclusive...<sup>0.1</sup>  $\Rightarrow O(10\%)$  for total uncertainty? **LHCXSWG** (2011), Baglio et al (2015) **Decay** branching ratios **Dominant decay**  $H \rightarrow b\bar{b} \approx 60\%$ **Affected by QCD+parametric errors:** from  $m_b$  and  $\alpha_s$  only, a few  $\% \Rightarrow$ migrate to O(5%) error in other modes such as  $\mathbf{H} \to \gamma \gamma, \mathbf{ZZ}, \mathbf{WW}, \tau \tau$ (partial widths very precise  $\leq 1\%$ ). **too** large theory uncertainties (even if reduced by a factor of 2)...

A deeper probe of the Higgs sector – A.



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#### . Interludium: $\mathbf{D}_{\gamma\gamma}$

Best way to eliminate theory uncertainty: use ratios of signal rates.

 $\begin{array}{l} \mathbf{H} \rightarrow \mathbf{VV} \text{ with } \mathbf{V} \rightarrow \ell \text{ as reference and } \mathbf{H} \rightarrow \mathbf{XX} \text{ with } \mathbf{H} \text{ produced in p:} \\ \mathbf{D}_{\mathbf{XX}} = \sigma^{\mathbf{p}}(\mathbf{pp} \rightarrow \mathbf{H} \rightarrow \mathbf{XX}) / \sigma^{\mathbf{p}}(\mathbf{pp} \rightarrow \mathbf{H} \rightarrow \mathbf{VV}) \end{array}$ 

 $= \sigma^{\mathbf{p}}(\mathbf{p}\mathbf{p} \rightarrow \mathbf{H}) \times \mathbf{B}\mathbf{R}(\mathbf{H} \rightarrow \mathbf{X}\mathbf{X}) / \sigma^{\mathbf{p}}(\mathbf{p}\mathbf{p} \rightarrow \mathbf{H}) \times \mathbf{B}\mathbf{R}(\mathbf{H} \rightarrow \mathbf{V}\mathbf{V})$ 

 $= \mathbf{BR}(\mathbf{H} \rightarrow \mathbf{XX}) / \mathbf{BR}(\mathbf{H} \rightarrow \mathbf{VV}) = \Gamma(\mathbf{H} \rightarrow \mathbf{XX}) / \Gamma(\mathbf{H} \rightarrow \mathbf{VV})$ 

To first approximation:  $D_{\mathbf{X}\mathbf{X}} = c_{\mathbf{X}}^2/c_{\mathbf{V}}^2$ 

Works only if one selects exactly same kinematical configuration
(i.e. same "fiducial cross sections") for the two channels X and V!
the theoretical uncertainties from the cross sections drop out;
the parametric uncertainties from the branching ratios drop out;
the theoretical ambiguities in the Higgs total width also drop out;

⇒ D<sub>XX</sub> measures only the ratio of partial decay widths. • Extremely clean theoretically, although some information will be lost • And maybe it has also some advantages from the experimental side? Hest probe by far is D<sub>γγ</sub> which measures deviations of the γγ loop  $D_{\gamma\gamma} = \frac{\sigma(pp \rightarrow H \rightarrow \gamma\gamma)}{\sigma(pp \rightarrow H \rightarrow VV)} = \frac{\Gamma(H \rightarrow \gamma\gamma)}{\Gamma(H \rightarrow VV)} = \frac{d_{\gamma\gamma}c_{\gamma}^2/c_{V}^2}{AD(2012)}$ Fokyo 9/01/2017 A deeper probe of the Higgs sector - A. Djouadi - p.20/31

**Interludium:**  $D_{\gamma\gamma}$  $\sum_{\gamma} \frac{G_{\mu} \alpha^2 M_{H}^3}{128 \sqrt{2} \pi^3} \left| \sum_{f} N_{c} e_{f}^2 A_{\frac{1}{2}}^H(\tau_{f}) + A_{1}^H(\tau_{W}) \right|^2$  $\begin{array}{ll} \gamma(Z) & \mathbf{A_{1/2}^{H}}(\tau) = \mathbf{2}[\tau + (\tau - \mathbf{1})\mathbf{f}(\tau)] \, \tau^{-2} \\ \mathbf{\mathcal{M}} & \mathbf{A_{1}^{H}}(\tau) = -[\mathbf{2}\tau^{2} + \mathbf{3}\tau + \mathbf{3}(\mathbf{2}\tau - \mathbf{1})\mathbf{f}(\tau)] \, \tau^{-2} \end{array}$ Loop decay. In SM: only W- and top-loops are relevant (others small) • For  $m_i \to \infty \Rightarrow A_{1/2} = \frac{4}{3}$  and  $A_1 = -7$ : W loop dominating! (approximation  $\tau_W \to 0$  valid only for  $M_H \lesssim 2M_W$ : relevant here). yy width counts the number of charged particles coupling to Higgs! **Contribution**  $A_s^p$  of particle p of spin s with Higgs coupling  $g_{Hpp}$ :  $A_0^p = -rac{1}{3}g_{Hpp}^2/m_P^2, A_{1/2}^p = +rac{4}{3}g_{Hpp}^2/m_P^2, A_1^p = -7g_{Hpp}^2/m_P^2,$ If  $g_{Hpp} \propto m_p \Rightarrow A_0^p \rightarrow +rac{1}{3}, A_{1/2}^p \rightarrow -rac{4}{3}, A_1^p \rightarrow +7.$ **Small/c**alculated QCD and EW corrections: only of order of percent. +Spira+Zerwas, Vicini et al., Passarino et al., AD+Gambino, Denner et al.,.. In SM with W,t loops:  ${f c}_{\gamma}pprox {f 1.26 imes} |{f c}_{f W}-{f 0.21\, c}_{f t}|$ Assuming custodial symmetry  $g_{HZZ} = g_{HWW} = c_V, D_{\gamma\gamma} = c_{\gamma}^2/c_V^2$  is  $\mathrm{c}^2_{_{\gamma}}/\mathrm{c}^2_{_{\mathbf{V}}}pprox 6.5 imes |1-rac{1}{5}\mathrm{c_t}/\mathrm{c_V}|^2$ with  $c_V = c_t = 1$  in SM. Any new physics effects will alter this value.

 Formula Comparison of the Higgs sector
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#### **Interludium:** $\mathbf{D}_{\gamma\gamma}$

#### Vill D<sub>1</sub> be the g-2 of the LHC?

Examples of BSM searches with the observable if measured at 1% level AD, J. Quevillon and R. Vega-Morales, arXiv:1509.03913 Model independent search through an effective Lagrangian approach.

$$\mathcal{L} = \frac{\mathbf{H}}{\mathbf{v}} \Big( \mathbf{c}_{\mathbf{V}} (\mathbf{2} \mathbf{M}_{\mathbf{W}}^{2} \mathbf{W}_{\mu}^{+} \mathbf{W}^{-\mu} + \mathbf{M}_{\mathbf{Z}}^{2} \mathbf{Z}_{\mu} \mathbf{Z}^{\mu}) - \mathbf{m}_{\mathbf{t}} \mathbf{\overline{t}} (\mathbf{c}_{\mathbf{t}} + \mathbf{i} \mathbf{\widetilde{c}}_{\mathbf{t}} \gamma^{\mathbf{5}}) \mathbf{t} \Big)$$

$$+rac{\mathbf{c}_{\gamma\gamma}}{4}\mathbf{F}^{\mu
u}\mathbf{F}_{\mu
u}+rac{\mathbf{ ilde{c}}_{\gamma\gamma}}{4}\mathbf{ ilde{F}}^{\mu
u}\mathbf{F}_{\mu
u}\Big)$$





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#### . Interludium: $D_{\gamma\gamma}$ $\propto rac{4}{3} imes {f g}_{{f h}\chi_{f i}^+\chi_{f i}^-}/m_{\chi_{f i}^\pm} \propto 1/m_{\chi_{f i}^\pm}^2$ $1 \propto rac{1}{3} imes {f g_{f h ilde au_{f i} ilde au_{f i}}}/m_{ ilde au_{f i}}^2 \!\propto\! m_ au {f X_ au}/m_{ ilde au_{f i}}^2$ $\Delta D_{\gamma\gamma}$ $\Delta D_{\gamma\gamma}$ 20 0.02 0.05 0.01 0.500 1000-0.300 15 0.005 0.100 800 $m_{\tilde{t}_2}$ [TeV] 0.050 tan*β*=60 600 0.040 10 *X*<sub>τ</sub><0 0.030 400 0.020 5 0.010 200 $\tan\beta = 1$ -0.05 -0.01 0.005 200 400 600 800 1000 $m_{\chi_1^{\pm}} \, \, [\text{GeV}]$ 60 120 140 160 80 100 180 200 $m_{\tilde{\tau}_1}$ [GeV]

o limit on charginos and stau's from LHC direct searches in some cases. Fokyo 9/01/2017 A deeper probe of the Higgs sector – A. Djouadi – p.24/31

#### . Interludium: $\mathbf{D}_{\gamma\gamma}$



#### **Direct** search for new states w that the Higgs boson is found, is Particle Physics "closed"? Not 2) Fully probe the TeV scale that is relevant for the hierarchy problem **cont**inue to search for heavier H bosons and new (super)particles. or instance in the MSSM: search for the heavier Higgs bosons: Fig. **Improve** "standard" searches for the heavier MSSM Higgs bosons: • Searches for the $pp \rightarrow A/H/(h) \rightarrow \tau \tau$ resonant process: $\Rightarrow$ constrains high tan $\beta$ for low M<sub>A</sub> values. • Searches for charged Higgs in $t \rightarrow bH^+ \rightarrow b\tau\nu$ decays: • Search for the heavier Higgsses in H $\rightarrow$ ZZ,WW (and $\gamma\gamma$ !) modes:



# **Direct search for new states**



• More Higgs particles:  $\Phi = h, H, A, H^{\pm}$ :

licated in the MSSM

- some couple almost like the SM Higgs,
- but some are more weakly coupled.
- In general same production as in SM but also new/more complicated processes (rates can be smaller or larger than in SM).
- Possibly many different decay modes, (and clean decays eg into  $\gamma\gamma$  suppressed).
- Impact of light SUSY particles?

⇒ very complicated situation in general.
But simpler in the decoupling regime:

– h as in SM with  $\mathbf{M_{h}}\!=\!\mathbf{115}\!-\!\mathbf{130}~GeV$ 

- dominant mode:  $gg, b\bar{b} \rightarrow H/A \rightarrow \tau\tau$ .

It is even more tricky in beyond MSSM, and also in many non-SUSY extensions...

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#### **Direct search for new states**

Now that the Higgs boson is found, is Particle Physics "closed"? No! 2) Fully probe the TeV scale that is relevant for the hierarchy problem ⇒ continue to search for heavier H bosons and new (super)particles. r instance in gtghe MSSM: search for the heavier Higgs bosons: Fig.

B) Look at other channels for H/A not present in the SM case:
Searches for the interesting channels A → hZ and H → hh for low tanβ and not too heavy A,H states (below tt threshold)
Searches for heavy H/A → tt resonances (beware of interference) for low tanβ and heavy A,H states (far above the tt threshold)



#### **Direct search for new states**

The "money Higgs plot" at the end of the LHC could look like:



## . Direct search for new states

Search for supersymmetric particles not only strong but also electroweak): - squarks and gluinos up to a few TeV, - chargino/neutralino/sleptons to 1 TeV, - LSP/DM neutralino up to few 100 GeV, including in non minimal scenarios). See David Shih; example of CMS for  $\tilde{t}/\chi_1^0 \Rightarrow$ 

• Search for any new heavy particle (predicted in all BSM extensions...):

<mark>- new m</mark>ulti–TeV Z′ bosons

- Kaluza–Klein excitations
- Techni-fermions and bosons
- top (composite) partners
- unexpected ones (LQ, new f, ..)

See Brian; ex: ATLAS for  $\mathbf{Z}' \rightarrow \ell \ell \Rightarrow$ 

t-t production LSP mass [GeV] CMS Preliminary 450 √s = 8 TeV SUSY 2013 Expected 400 SUS-13-004 0-lep+1-lep (Razor) 19.3 fb<sup>-1</sup> ( $\tilde{t} \rightarrow t \tilde{\chi}^0$ ) 350 SUS-13-011 1-lep (leptonic stop)19.5 fb<sup>-1</sup> ( $\widetilde{t} \rightarrow t \, \widetilde{\chi}_{-}^{0}$ ) S-13-011 1-lep (leptonic stop)19.5 fb<sup>-1</sup> ( $\tilde{t} \rightarrow b \ \tilde{\chi}_{*}^{*}$ , x=0.25) 300 250 200 150 100 50 200 300 400 500 600 700 100 800 σ B [pb] ATLAS --- Expected limit √s = 8 TeV Expected ± 1σ 10  $Z' \rightarrow H$ Expected  $\pm 2\sigma$ Observed limit  $10^{-2}$ \_\_\_\_Ζ'<sub>Ψ</sub> 10<sup>-3</sup> 10  $L dt = 20.3 \text{ fb}^{-1}$ dt = 20.5 fb 2.5 3.5 M<sub>7</sub> [TeV]

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#### **Conclusion**

Hence, we need to continue search for New Physics and falsify the SM:

indirectly via high precision Higgs measurements (HL-LHC, ILC, ...),
directly via heavy particle searches at high-energy (HE-LHC, CLIC),
and we should plan/prepare/construct the new facilities already now.



So let's move forward: it is still action time! (or as the experimentalists usually say: stay tuned...) Fokyo 9/01/2017 A deeper probe of the Higgs sector – A. Djouadi – p.31/31