



Non-linear effects

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- Books on non-linear dynamical systems
 - M. Tabor, Chaos and Integrability in Nonlinear Dynamics, An Introduction, Willey, 1989.
 - A.J Lichtenberg and M.A. Lieberman, Regular and Chaotic Dynamics, 2nd edition, Springer 1992.
- Books on beam dynamics
 - E. Forest, Beam Dynamics A New Attitude and Framework, Harwood Academic Publishers, 1998.
 - H. Wiedemann, Particle accelerator physics, 3rd edition, Springer 2007.
- Lectures on non-linear beam dynamics
 - □ A. Chao, Advanced topics in Accelerator Physics, USPAS, 2000.
 - A. Wolski, Lectures on Non-linear dynamics in accelerators, Cockroft Institute 2008.
 - W. Herr, Lectures on Mathematical and Numerical Methods for Non-linear Beam Dynamics in Rings, CAS 2013.
 - L. Nadolski, Lectures on Non-linear beam dynamics, Master NPAC, LAL, Orsay 2013.

Contents of the 1st lecture



- Accelerator performance parameters and non-linear effects
- Linear and non-linear oscillators
 - Integral and frequency of motion
 - The pendulum
- Phase space dynamics
 - Fixed point analysis
- Non-autonomous systems
 - Driven (damped) harmonic oscillator, resonance conditions
- Linear equations with periodic coefficients Hill's equations
 - Floquet solutions and normalized coordinates
- Perturbation theory
 - Non-linear oscillator
 - Perturbation by periodic function single dipole perturbation
 - □ Application to single multipole resonance conditions
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Contents of the 1st lecture Accelerator performance parameters and non-linear effects



Accelerator performance parameters



Colliders

$$L = \frac{N_b^2 k_b f_{\rm rev} \gamma}{4\pi\epsilon\beta^*} \mathcal{R}(\phi)$$

$$\bar{D} = \bar{I}E = f_N NeE$$

 $\frac{N_p}{4\pi^2\epsilon_x\epsilon_y}$

- N_b bunch population
- k_b number of bunches
- f_{rev} the revolution frequency
- γ relativistic reduced energy
- ε_n normalized emittance
- β* "betatron" amplitude function at collision point
- $R(\varphi)$ geometric reduction factor due to crossing angle

High intensity rings

□ Average beam power

- I mean current intensity
- E energy
- *f_N* repetition rate
- N number of particles/pulse

X-ray (low emittance) rings

- Brightness (photon density in phase space)
 - *N_p* number of photons
 - $\varepsilon_{x,y}$ transverse emittances

Non-linear effects limit performance of particle accelerators but impact also design cost

Non-linear effects in colliders





At injection

- Non-linear magnets (sextupoles, octupoles)
- Magnet imperfections and misalignments
- Power supply ripple
- □ Ground motion (for e+/e-)
- Electron (Ion) cloud

At collision

- Non-linear magnets (sextupoles and octupoles
- Field imperfections in the insertion quadrupoles
- Magnets in experimental areas (solenoids, dipoles)
- "Incoherent" Beam-beam effects (head on and long range)
- "Incohoront" E cloud offect

Non-linear effects in colliders





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- Limitations affecting (integrated) luminosity
 - Particle losses causing
 - Reduced lifetime
 - Radio-activation (superconducting magnet quench)
 - Reduced machine availability
 - Emittance blow-up
 - Reduced number of bunches (either due to electron cloud or long-range beam-beam)
 - Increased crossing angle
 - Reduced intensity
- **Cost** issues
 - Number of magnet correctors and families (power convertors)
 - Magnetic field and alignment tolerances
 - Design of the collimation system





$\bar{P} = \bar{I}E = f_N NeE$

- Non-linear magnets (sextupoles, octupoles)
- Magnet imperfections and misalignments
- Injection chicane
- Magnet fringe fields
- Space-charge effect



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- Non-linear magnets (sextupoles, octupoles)
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- Space-charge effect

Limitations affecting beam power

- Particle losses causing
 - Reduced intensity
 - Radio-activation (hands-on maintenance)
 - Reduced machine availability
- Emittance blow-up which can lead to particle loss
- Cost issues
 - Number of magnet correctors and families (power convertors)
 - Magnetic field and alignment tolerances
 - Design of the collimation system



$$B = \frac{N_p}{4\pi^2 \epsilon_x \epsilon_y}$$

- Chromaticity sextupoles
- Magnet imperfections and misalignments
- Insertion devices (wigglers, undulators)
- Injection elements
- Ground motion
- Magnet fringe fields
- Space-charge effect (in the vertical plane for damping rings)
 - Electron cloud (Ion) effects

Non-linear effects in low emittance rings





- Chromaticity sextupoles
- Magnet imperfections and misalignments
- Insertion devices (wigglers, undulators)
- Injection elements
- Ground motion
- Magnet fringe fields
- Space-charge effect (in the vertical plane for damping rings)
- Electron cloud (Ion) effects

Limitations affecting beam brightness

- Reduced injection efficiency
- Particle losses causing
 - Reduced lifetime
 - Reduced machine availability
- Emittance blow-up which can lead to particle loss
- Cost issues
 - Number of magnet correctors and families (power convertors)
 - Magnetic field and alignment tolerances

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Non-linear effects, JUAS, February 2017

Appendix I: Multipole expansion

Reminder: Harmonic oscillator



Described by the differential equation:

$$\frac{d^2u(t)}{dt^2} + \omega_0^2 u(t) = 0$$

The solution obtained by the substitution $u(t) = e^{\lambda t}$ and the solutions of the characteristic polynomial are $\lambda_{\pm} = \pm i\omega_0$ which yields the general solution $u(t) = ce^{i\omega_0 t} + c^* e^{-i\omega_0 t} = C_1 \cos(\omega_0 t) + C_2 \sin(\omega_0 t) = A \sin(\omega_0 t + \phi)$ Reminder: Harmonic oscillator



Described by the differential equation:

$$\frac{d^2u(t)}{dt^2} + \omega_0^2 u(t) = 0$$

The solution obtained by the substitution u(t) = e^{λt} and the solutions of the characteristic polynomial are λ_± = ±iω₀ which yields the general solution u(t) = ce^{iω₀t} + c^{*}e^{-iω₀t} = C₁ cos(ω₀t) + C₂ sin(ω₀t) = A sin(ω₀t + φ)

The amplitude and phase depend on the initial conditions $A = \frac{\left(\frac{du}{dt}(0)^2 + \omega_0^2 u(0)^2\right)^{1/2}}{\omega_0}, \quad tan(\phi) = \frac{\frac{du}{dt}(0)}{\omega_0 u(0)}$

A negative sign in the differential equation provides a solution described by an hyperbolic sine function
Note also that for no restoring force \u03c6₀ = 0 , the motion is unbounded

Integral of motion

- CERN
- Rewrite the differential equation of the harmonic oscillator as a **pair of coupled 1**st order equations $\frac{du(t)}{dt} = p_u(t)$
- $\frac{dt}{dt} = -\omega_0^2 u(t)$ which can be combined to provide $\frac{dp_u(t)}{dt} = -\omega_0^2 u(t)$ or $\frac{dp_u}{dt} p_u + \omega_0^2 u \frac{du}{dt} = \frac{1}{2} \frac{d}{dt} \left(p_u^2 + \omega_0^2 u^2 \right) = 0$ or $\frac{1}{2} \left(p_u^2 + \omega_0^2 u^2 \right) = I_1$ with I_1 an integral of motion identified as the mechanical energy of the system

Integral of motion



- Rewrite the differential equation of the harmonic oscillator as a **pair of coupled 1**st order equations $\frac{du(t)}{dt} = p_u(t)$
- $\frac{dp_u(t)}{dt} = -\omega_0^2 u(t)$ which can be combined to provide $\frac{dp_u}{dt} p_u + \omega_0^2 u \frac{du}{dt} = \frac{1}{2} \frac{d}{dt} \left(p_u^2 + \omega_0^2 u^2 \right) = 0$ or $\frac{1}{2}\left(p_u^2 + \omega_0^2 u^2\right) = I_1 \text{ with } I_1 \text{ an integral of motion}$ identified as the mechanical energy of the system Solving the previous equation for p_u , the system can be reduced to a unique 1st order equation

$$\frac{du}{dt} = \sqrt{2I_1 - \omega_0^2 u^2}$$

Integration by quadrature



The last equation can be be solved as an explicit integral or "quadrature"

$$\int dt = \int \frac{du}{\sqrt{2I_1 - \omega_0^2 u^2}} \text{, yielding } t + I_2 = \frac{1}{\omega_0} \operatorname{arcsin}\left(\frac{u\omega_0}{\sqrt{2I_1}}\right)$$

or the well-known solution $u(t) = \frac{\sqrt{2I_1}}{\omega_0} \sin(\omega_0 t + \omega_0 I_2)$



- The last equation can be be solved as an explicit integral or "quadrature"
- $\int dt = \int \frac{du}{\sqrt{2I_1 \omega_0^2 u^2}} \text{ , yielding } t + I_2 = \frac{1}{\omega_0} \operatorname{arcsin} \left(\frac{u\omega_0}{\sqrt{2I_1}}\right)$ or the well-known solution $u(t) = \frac{\sqrt{2I_1}}{\omega_0} \sin(\omega_0 t + \omega_0 I_2)$ Although the previous route may seem complicated, it becomes more **natural** when **non-linear terms** appear, where a substitution of the type is not u(t) is not $u(t) = \frac{1}{\omega_0} \operatorname{arcsin} \left(\frac{u\omega_0}{\sqrt{2I_1}}\right)$
 - The ability to integrate a differential equation is not just a nice mathematical feature, but deeply characterizes the dynamical behavior of the system described by the equation



The period of the harmonic oscillator is calculated through the previous integral after integration between two extrema, i.e. when the velocity

$$rac{du}{dt} = \sqrt{2I_1 - \omega_0^2 u^2}$$
 vanishes, at $u_{\mathrm{ext}} = \pm rac{\sqrt{2I_1}}{\omega_0}$



The **period** of the harmonic oscillator is calculated through the previous integral after integration between two **extrema**, i.e. when the velocity

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m ext} = \pm rac{\sqrt{2I_1}}{\omega_0}$

The integration yields

$$T = 2 \int_{-\frac{\sqrt{2I_1}}{\omega_0}}^{\frac{\sqrt{2I_1}}{\omega_0}} \frac{du}{\sqrt{2I_1 - \omega_0^2 u^2}} = \frac{2\pi}{\omega_0}$$

The frequency (or the period) of linear systems is independent of the integral of motion (energy)

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The previous remark is not true for non-linear systems, e.g. for an oscillator with a non-linear restoring force ^{d²u}/_{dt²} + k u(t)³ = 0 The integral of motion is I₁ = ¹/₂p_u² + ¹/₄k u⁴

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The previous remark is not true for non-linear systems, e.g. for an oscillator with a non-linear restoring force $\frac{d^2u}{dt^2} + k \ u(t)^3 = 0$ The integral of motion is $I_1 = \frac{1}{2}p_u^2 + \frac{1}{4}k \ u^4$ Solving for vanishing velocity, we get $u_{\text{ext}} = \pm \left(\frac{2I_1}{k}\right)^{1/4}$ The integration yields $T = 2 \int_{-(4I_1/k)^{1/4}}^{(4I_1/k)^{1/4}} \frac{du}{\sqrt{2I_1 - \frac{1}{2}k \ u^4}} = \sqrt{\frac{1}{2\pi}} \Gamma^2(\frac{1}{4}) \left(I_1 \ k\right)^{-1/4}$

i.e. the **period** (frequency) depends on the integral of motion (energy), i.e. the maximum "amplitude"

The pendulum



An important non-linear equation which can be integrated is the one of the **pendulum**, for a string of length L and gravitational constant g

 $\frac{d^2\phi}{dt^2} + \frac{g}{L}\sin\phi = 0$ For small displacements it reduces to an harmonic

oscillator with frequency $\omega_0 = \sqrt{\frac{g}{L}}$

The integral of motion (scaled energy) is

$$\frac{1}{2} \left(\frac{d\phi}{dt}\right)^2 - \frac{g}{L}\cos\phi = I_1 = E'$$

and the quadrature is written as $t = \int \frac{d\phi}{\sqrt{2(I_1 + \frac{g}{L}\cos\phi)}}$ assuming that for t = 0, $\phi = 0$

Solution for the pendulum



The integral
$$t = \int \frac{d\phi}{\sqrt{2(I_1 + \frac{g}{L}\cos\phi)}}$$
 can be solved, using the substitution $\cos\phi = 1 - 2k^2 \sin^2 \theta$

with
$$k = \sqrt{1/2(1 + I_1 L/g)}$$
.

The integral then becomes

$$t = \sqrt{\frac{L}{g}} \int_0^\theta \frac{d\theta}{\sqrt{1 - k^2 \sin^2 \theta}}$$

It is solved using **Jacobi elliptic functions**, with the final result:

$$\phi(t) = 2 \arcsin\left[k \sin\left(t\sqrt{\frac{g}{L}},k\right)\right]$$

Period of the pendulum



For recovering the **period**, the integration is performed between the two extrema, i.e. $\phi = 0$ $\phi = \frac{1}{2} \frac{1}{2} \frac{1}{2} \frac{1}{2} \int_{0}^{\pi/2} \frac{d\theta}{\sqrt{1 - k^2 \sin^2 \theta}} = 4\sqrt{\frac{L}{g}} \mathcal{K}(k)$

The **period** is

i.e. the **complete elliptic integral** multiplied by four times the period of the second of the seco

with

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, the "amplitude"

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Phase space dynamics



- Valuable description when examining trajectories in **phase space** (u, p_u)
- Existence of integral of motion imposes geometrical constraints on phase flow
- For the simple harmonic oscillator

$$H = \frac{1}{2} \left(p_u^2 + \omega_0^2 u^2 \right)$$

phase space curves are **ellipses** around the equilibrium point parameterized by the integral of motion Hamiltonian (energy)





Phase space dynamics



- Valuable description when examining trajectories in **phase space** (u, p_u)
- Existence of integral of motion imposes geometrical constraints on phase flow
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phase space curves are **ellipses** around the equilibrium point parameterized by the integral of motion Hamiltonian (energy)

By simply **changing** the **sign** of the potential in the harmonic oscillator, the phase trajectories become **hyperbolas**, symmetric around the equilibrium point where two straight lines cross, moving towards and away from it







Conservative non-linear oscillators have Hamiltonian

$$H = E = \frac{1}{2}p_u^2 + V(u)$$

with the potential being a general (polynomial) function of positions **Equilibrium points** are associated with extrema of the potential



Non-linear oscillators





Conservative non-linear oscillators have Hamiltonian

$$H = E = \frac{1}{2}p_u^2 + V(u)$$

with the potential being a general (polynomial) function of positions

- Equilibrium points are associated with extrema of the potential
- Considering three non-linear oscillators
 - Quartic potential (left): two minima and one maximum
 - **Cubic** potential (center): one minimum and one maximum
 - Pendulum (right): periodic minima and maxima



Fixed point analysis



- Consider a general second order system $\frac{\frac{du}{dt}}{\frac{dp_u}{dt}} = f_1(u, p_u)$ $\frac{\frac{dp_u}{dt}}{\frac{dt}{dt}} = f_2(u, p_u)$
- Equilibrium or "**fixed**" points $f_1(u_0, p_{u0}) = f_2(u_0, p_{u0}) = 0$ are determinant for topology of trajectories at their vicinity





- Consider a general second order system $\frac{\frac{du}{dt}}{\frac{dp_u}{dt}} = f_1(u, p_u)$ $\frac{\frac{dp_u}{dt}}{\frac{dt}{dt}} = f_2(u, p_u)$
- Equilibrium or "fixed" points f₁(u₀, p_{u0}) = f₂(u₀, p_{u0}) = 0 are determinant for topology of trajectories at their vicinity
 The linearized equations of motion at their vicinity are

$$\frac{d}{dt} \begin{bmatrix} \delta u \\ \delta p_u \end{bmatrix} = \mathcal{M}_J \begin{bmatrix} \delta u \\ \delta p_u \end{bmatrix} = \begin{bmatrix} \frac{\partial f_1(u_0, p_{u0})}{\partial u} & \frac{\partial f_1(u_0, p_{u0})}{\partial p_u} \\ \frac{\partial f_2(u_0, p_{u0})}{\partial u} & \frac{\partial f_2(u_0, p_{u0})}{\partial p_u} \end{bmatrix} \begin{bmatrix} \delta u \\ \delta p_u \end{bmatrix}$$

Jacobian matrix

Fixed point nature is revealed by **eigenvalues** of \mathcal{M}_J , i.e. solutions of the characteristic polynomial $\det |\mathcal{M}_J - \lambda \mathbf{I}| = 0$

Fixed point for conservative systems



- For conservative systems of 1 degree of freedom, the second order characteristic polynomial for any fixed point has two possible solutions:
 - Two complex eigenvalues with opposite sign, corresponding to elliptic fixed points. Phase space flow is described by ellipses, with particles evolving clockwise or anti-clockwise



Fixed point for conservative systems



- For conservative systems of 1 degree of freedom, the second order characteristic polynomial for any fixed point has two possible solutions:
 - Two complex eigenvalues with opposite sign, corresponding to elliptic fixed points. Phase space flow is described by ellipses, with particles evolving clockwise or anti-clockwise
 - Two real eigenvalues with opposite sign, corresponding to hyperbolic (or saddle) fixed points. Flow described by two lines (or manifolds), incoming (stable) and outgoing (unstable)



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separating bounded librations and unbounded rotations

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Non-autonomous systems



Consider a linear system with explicit dependence in time $\frac{d^2 u}{dt^2} + \omega_0^2 u = F(t)$

Time now is **not** an **independent** variable but can be considered as an **extra dimension** leading to a completely new type of behavior Non-autonomous systems



Consider a linear system with explicit dependence in time $\frac{d^2 u}{dt^2} + \omega_0^2 u = F(t)$

- Time now is not an independent variable but can be considered as an extra dimension leading to a completely new type of behavior
- Consider two independent solutions of the homogeneous equation $u_1(t)$ and $u_2(t)$
 - The general solution is a sum of the **homogeneous** solutions $u_h(t) = c_1 u_1(t) + c_2 u_2(t)$ and a **particular** solution, $u_p(t) = c_3 u_1(t) + c_4 u_2(t)$, where the coefficients are computed as

$$c_{3} = \int \frac{u_{2}(t)F(t)}{W(t)}dt \quad , c_{4} = \int \frac{u_{1}(t)F(t)}{W(t)}dt$$

with the Wronskian of the system

$$W(t) = u_1(t)\frac{du_2(t)}{dt} - u_2(t)\frac{du_1(t)}{dt}$$

Driven harmonic oscillator



Consider periodic force pumping energy into the system

$$\frac{d^2u(t)}{dt^2} + \omega_0^2 u(t) = \frac{F}{m}\cos(\omega t)$$

General solution is a combination of the homogeneous and a particular solution found as $_{F}$

$$u(t) = u_0 \sin(\omega_0 t + \phi_0) + \frac{F}{m(\omega_0^2 - \omega^2)} \cos(\omega t)$$

Driven harmonic oscillator



Consider periodic force pumping energy into the system

$$\frac{d^2u(t)}{dt^2} + \omega_0^2 u(t) = \frac{F}{m}\cos(\omega t)$$

General solution is a combination of the homogeneous and a particular solution found as

$$u(t) = u_0 \sin(\omega_0 t + \phi_0) + \frac{T}{m(\omega_0^2 - \omega^2)} \cos(\omega t)$$

Obviously a **resonance** condition appears when driving frequency hits the oscillator eigen-frequency.

Driven harmonic oscillator



Consider periodic force pumping energy into the system

$$\frac{d^2u(t)}{dt^2} + \omega_0^2 u(t) = \frac{F}{m}\cos(\omega t)$$

General solution is a combination of the **homogeneous** and a **particular** solution found as F

$$u(t) = u_0 \sin(\omega_0 t + \phi_0) + \frac{1}{m(\omega_0^2 - \omega^2)} \cos(\omega t)$$

Obviously a **resonance** condition appears when driving frequency hits the oscillator eigen-frequency.

In the limit of $\omega
ightarrow \omega_0$ the solution becomes

$$u(t) = \hat{u}_0 \sin(\omega_0 t + \hat{\phi}_0) + \frac{F}{2m\omega_0} t \sin(\omega_0 t)$$

The 2nd secular term implies unbounded growth of amplitude at resonance





Consider now a simple harmonic oscillator where the **frequency** is **time-dependent**

$$H = \frac{1}{2} \left(p_u^2 + \omega_0^2(t) u^2 \right)$$

- Plotting the evolution in phase space, provides trajectories that intersect each other
- The phase space has time as extra dimension



Phase space for time-dependent systems



Consider now a simple harmonic oscillator where the **frequency** is **time-dependent**

$$H = \frac{1}{2} \left(p_u^2 + \omega_0^2(t) u^2 \right)$$

- Plotting the evolution in phase space, provides trajectories that intersect each other
- The phase space has time as extra dimension
- By rescaling the time to become $= \omega_0 t$ and considering every integer interval of the new p_{u_s} "time" variable, the phase space looks like the one of the harmonic oscillator
 - This is the simplest version of a **Poincaré surface of section**, which is useful for studying geometrically phase space of multi-dimensional systems



Phase space for time-dependent systems



Consider now a simple harmonic oscillator where the **frequency** is **time-dependent**

$$H = \frac{1}{2} \left(p_u^2 + \omega_0^2(t) u^2 \right)$$

- Plotting the evolution in phase space, provides trajectories that intersect each other
- The phase space has time as extra dimension
- By **rescaling** the **time** to become $= \omega_0 t$ and considering every integer interval of the **new** p_{u} "**time**" variable, the **phase space** looks like the one of the **harmonic oscillator**
 - This is the simplest version of a **Poincaré surface of section**, which is useful for studying geometrically phase space of multi-dimensional systems
 - The **fixed point** in the surface of section is now a periodic orbit



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 A very important class of equations especially for beam dynamics (but also solid state physics) are linear equations with periodic coefficients

Linear equation with periodic coefficients

$$\frac{d^2u}{dt^2} + K(t)u = 0$$



George Hill

with $K(t) = K(t \neq p\mathbf{F})$ indic function of time

 A very important class of equations especially for beam dynamics (but also solid state physics) are linear equations with periodic coefficients

$$\frac{d^2u}{dt^2} + K(t)u = 0$$



George Hill

with $K(t) = K(t + p\mathbf{F})$ indic function of time

These are called **Hill's equations** and can be thought as equations of harmonic oscillator with time dependent (periodic) frequency

There are two solutions that can be written as $u(t) = \Re \left\{ w(t)e^{i\psi(t)} \right\}$ with $w(t) = w(t - periodic but also <math>e^{i\psi(t+T)} With = e^{i\sigma}$ constant which implies that $\frac{d\psi}{dt}(t+T) = \frac{d\psi}{dt}e$ for $\frac{d\psi}{dt}$

Linear equation with periodic coefficients

The solutions are derived based on **Floquet** theory

Amplitude, phase and invariant



Differentiating the solutions twice and substituting to Hill's equation, the following two equations are obtained

$$\frac{d^2w}{dt^2} - w(\frac{d\psi}{dt})^2 + K(t)w = 0$$
$$2\frac{dw}{dt}\frac{d\psi}{dt} + w\frac{d^2\psi}{dt^2} = 0$$

The 2nd one can be integrated to give $\frac{d\psi}{dt} = \frac{1}{w^2}$, i.e. the relation between the "**phase**" and the **amplitude**

Amplitude, phase and invariant



Differentiating the solutions twice and substituting to Hill's equation, the following two equations are obtained

$$\frac{d^2w}{dt^2} - w(\frac{d\psi}{dt})^2 + K(t)w = 0$$
$$2\frac{dw}{dt}\frac{d\psi}{dt} + w\frac{d^2\psi}{dt^2} = 0$$

- The 2nd one can be integrated to give $\frac{d\psi}{dt} = \frac{1}{w^2}$, i.e. the relation between the "**phase**" and the **amplitude**
- Substituting this to the 1st equation, the amplitude equation is derived (or the **beta function** in accelerator jargon) $\frac{d^2w}{dt^2} + K(t)w - \frac{1}{w^3} = 0$
 - By evaluating the quadratic sum of the solution and its derivative an invariant can be constructed, with the form $I(u, \frac{du}{dt}, t) = \left[\frac{u^2}{w^2} + \left(w\frac{du}{dt} \frac{dw}{dt}u\right)^2\right]$

Normalized coordinates



- Recall the Floquet solutions $u(s) = \sqrt{\epsilon\beta(s)}\cos(\psi(s) + \psi_0)$ for betatron motion $u'(s) = -\sqrt{\frac{\epsilon}{\beta(s)}}(\sin(\psi(s) + \psi_0) + \alpha(s)\cos(\psi(s) + \psi_0))$
- Introduce new variables

$$\mathcal{U} = \frac{u}{\sqrt{\beta}} , \quad \mathcal{U}' = \frac{d\mathcal{U}}{d\phi} = \frac{\alpha}{\sqrt{\beta}}u + \sqrt{\beta}u' , \quad \phi = \frac{\psi}{\nu} = \frac{1}{\nu}\int \frac{ds}{\beta(s)}$$

In matrix form $\begin{pmatrix} \mathcal{U}\\ \mathcal{U}' \end{pmatrix} = \begin{pmatrix} \frac{1}{\sqrt{\beta}} & 0\\ \frac{\alpha}{\sqrt{\beta}} & \sqrt{\beta} \end{pmatrix} \begin{pmatrix} u\\ u' \end{pmatrix}$

Normalized coordinates



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- Introduce new variables
- $\mathcal{U} = \frac{u}{\sqrt{\beta}} , \quad \mathcal{U}' = \frac{d\mathcal{U}}{d\phi} = \frac{\alpha}{\sqrt{\beta}}u + \sqrt{\beta}u' , \quad \phi = \frac{\psi}{\nu} = \frac{1}{\nu} \int \frac{ds}{\beta(s)}$ In matrix form $\begin{pmatrix} \mathcal{U} \\ \mathcal{U}' \end{pmatrix} = \begin{pmatrix} \frac{1}{\sqrt{\beta}} & 0 \\ \frac{\alpha}{\sqrt{\beta}} & \sqrt{\beta} \end{pmatrix} \begin{pmatrix} u \\ u' \end{pmatrix}$ Hill's equation becomes $\frac{1}{\nu^2 \beta^{3/2}} \left(\frac{d^2 \mathcal{U}}{d\phi^2} + \nu^2 \mathcal{U} \right) = 0$ System becomes harmonic oscillator with frequency $\mathcal{U}^2 + \mathcal{U'}^2 = \epsilon$ $\begin{pmatrix} \mathcal{U} \\ \mathcal{U}' \end{pmatrix} = \sqrt{\epsilon} \begin{pmatrix} \cos(\nu\phi) \\ -\sin(\nu\phi) \end{pmatrix}$



Perturbation of Hill's equations



Hill's equations in normalized coordinates with harmonic perturbation, using $\mathcal{U} = \mathcal{U}_x$ or \mathcal{U}_y and $\phi = \phi_x$ or ϕ_y $\frac{d^2\mathcal{U}}{d\phi^2} + \nu^2\mathcal{U} = \nu^2\beta^{3/2}F(\mathcal{U}_x(\phi_x), \mathcal{U}_y(\phi_y))$

where the *F* is the Lorentz force from perturbing fields

- Linear magnet imperfections: deviation from the design dipole and quadrupole fields due to powering and alignment errors
- Time varying fields: feedback systems (damper) and wake fields due to collective effects (wall currents)
- Non-linear magnets: sextupole magnets for chromaticity correction and octupole magnets for Landau damping
- Beam-beam interactions: strongly non-linear field
- □ **Space charge effects**: very important for high intensity beams
- non-linear magnetic field imperfections: particularly difficult to control for super conducting magnets where the field quality is entirely determined by the coil winding accuracy

■ From Gauss law of magnetostatics, a vector potential exist ∇ ⋅ B = 0 → ∃A : B = ∇ × A
■ Assuming transverse 2D field, vector potential has only one component A_s. The Ampere's law in vacuum (inside the beam pipe) ∇ × B = 0 → ∃V : B = −∇V
■ Using the previous equations, the relations between field components and potentials are

$$B_x = -\frac{\partial V}{\partial x} = \frac{\partial A_s}{\partial y} , \quad B_y = -\frac{\partial V}{\partial y} = -\frac{\partial A_s}{\partial x}$$



Multipole expansion II



From the complex potential we can derive the fields

$$B_y + iB_x = -\frac{\partial}{\partial x}(A_s(x,y) + iV(x,y)) = -\sum_{n=1}^{\infty} n(\lambda_n + i\mu_n)(x + iy)^{n-1}$$
Setting $b_n = -n\lambda_n$, $a_n = n\mu_n$
 $B_y + iB_x = \sum_{n=1}^{\infty} (b_n - ia_n)(x + iy)^{n-1}$

Multipole expansion II



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$$B_{y} + iB_{x} = -\frac{\partial}{\partial x}(A_{s}(x, y) + iV(x, y)) = -\sum_{n=1}^{\infty} n(\lambda_{n} + i\mu_{n})(x + iy)^{n-1}$$
Setting $b_{n} = -n\lambda_{n}$, $a_{n} = n\mu_{n}$

$$B_{y} + iB_{x} = \sum_{n=1}^{\infty} (b_{n} - ia_{n})(x + iy)^{n-1}$$
Define normalized **multipole** coefficients
 $b'_{n} = \frac{b_{n}}{10^{-4}B_{0}}r_{0}^{n-1}$, $a'_{n} = \frac{a_{n}}{10^{-4}B_{0}}r_{0}^{n-1}$
on a reference radius r_{0} , 10⁻⁴ of the main field to get
 $B_{y} + iB_{x} = 10^{-4}B_{0}\sum_{n=1}^{\infty} (b'_{n} - ia'_{n})(\frac{x + iy}{r_{0}})^{n-1}$
Note: $n' = n - 1$ is the US convention

Contents of the 1st lecture



- Accelerator performance parameters and non-linear effects
- Linear and non-linear oscillators
 - □ Integral and frequency of motion
 - □ The pendulum
 - Damped harmonic oscillator
- Phase space dynamics
 - Fixed point analysis
- Non-autonomous systems
 - Driven harmonic oscillator, resonance conditions
- Linear equations with periodic coefficients Hill's equations
- Perturbation theory
 - Non-linear oscillator
 - Perturbation by periodic function single dipole perturbation
 - □ Application to single multipole resonance conditions
 - □ Examples: single quadrupole, sextupole, octupole perturbation
 - General multi-pole perturbation example: linear coupling
 - Resonance conditions and working point choice

Summary

Appendix: Damped harmonic oscillator





Completely integrable systems are exceptional

For understanding dynamics of general non-linear systems composed of a part whose solution $u_0(t)$ is known and a part parameterized by a small constant ϵ , **perturbation theory** is employed

Perturbation theory



Completely integrable systems are exceptional

- For understanding dynamics of general non-linear systems composed of a part whose solution $u_0(t)$ is known and a part parameterized by a small constant ϵ , **perturbation theory** is employed
- The general idea is to **expand** the solution in a power series $u(t) = u_0(t) + \epsilon u_1(t) + \epsilon^2 u_2(t) + \ldots$ and **compute recursively** the corrections $u_1(t), u_2(t), \ldots$ hoping that a few terms will be sufficient to find an accurate representation of the general solution



Completely integrable systems are exceptional

- For understanding dynamics of general non-linear systems composed of a part whose solution $u_0(t)$ is known and a part parameterized by a small constant ϵ , **perturbation theory** is employed
- The general idea is to expand the solution in a power series u(t) = u₀(t) + \epsilon u₁(t) + \epsilon² u₂(t) + ... and compute recursively the corrections₁(t), u₂(t), ... hoping that a few terms will be sufficient to find an accurate representation of the general solution
 This may not be true for all times, i.e. facing series convergence problems
 In addition, any series expansion breaks in the

effects



Consider a non-linear harmonic oscillator, $\frac{d^2u}{dt^2} + \omega_0^2 u - \frac{1}{6}\epsilon\omega_0^2 u^3 = 0$

This is just the pendulum expanded to 3rd order in Note that ϵ is a dimensionless measure of smallness, which may represent a scaling factor of u (e.g. $\epsilon = 1$ without loss of generality)



- Consider a non-linear harmonic oscillator, $\frac{d^2u}{dt^2} + \omega_0^2 u \frac{1}{\epsilon} \epsilon \omega_0^2 u^3 = 0$
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- Expanding $u(t) = u_0(t) + \epsilon u_1(t) + \epsilon^2 u_2(t) + \dots$ and separating the equations with equal power in ϵ :
 - $\Box \text{ Order 0: } \frac{d^2 u_0}{dt^2} + \omega_0^2 u_0 = 0 \Rightarrow u_0(t) = A\cos(\omega_0 t)$ **Order 1**: $\frac{d^2 u_1}{dt^2} + \omega_0^2 u_1 = \frac{\omega_0^2 u_0^3}{6} = \frac{\omega_0^2 A^3}{6} \cos^3(\omega_0 t) = \frac{\omega_0^2 A^3}{24} \left(\cos(3\omega_0 t) + 3\cos(\omega_0 t)\right)$



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- Expanding $u(t) = u_0(t) + \epsilon u_1(t) + \epsilon^2 u_2(t) + \dots$ and separating the equations with equal power in ϵ :

□ Order 0:
$$\frac{d^2u_0}{dt^2} + \omega_0^2u_0 = 0 \Rightarrow u_0(t) = A\cos(\omega_0 t)$$

□ Order 1: $\frac{d^2u_1}{dt^2} + \omega_0^2u_1 = \frac{\omega_0^2u_0^3}{6} = \frac{\omega_0^2A^3}{6}\cos^3(\omega_0 t) = \frac{\omega_0^2A^3}{24}(\cos(3\omega_0 t) + 3\cos(\omega_0 t))$
■ The 2nd equation has a particular solution with two terms. A well behaved one $u_{1a}(t) = -\frac{A^3}{192}\cos(3\omega t)$ and $u_{1b}(t) = \frac{A^3}{64}(\omega_0 t\cos(\omega_0 t) + 2\cos(\omega_0 t))$
the first part of which grows linearly with time (secular term)
■ But this cannot be true, the pendulum does not present such behavior. What did it go wrong?

Perturbation of non-linear oscillator
 It was already shown that the pendulum has an amplitude dependent frequency, so the frequency has to be developed as well (Poincaré-Linstead method):

 $\omega = \omega_0 + \epsilon \omega_1 + \epsilon^2 \omega_2 + \dots$



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 $\omega = \omega_0 + \epsilon \omega_1 + \epsilon^2 \omega_2 + \dots$

Assume that the solution is a periodic function of $\tau = \omega t$ which becomes the new independent variable. The equation at zero order gives the solution $u_0(\tau) = A\cos(\tau)$ and at leading perturbation order becomes $\omega_0 \frac{d^2 u_1}{d\tau^2} + \omega_0 u_1 = -2\omega_1 \frac{d^2 u_0}{d\tau^2} + \frac{\omega_0}{6} u_0^3 = \frac{\omega_0 A^3}{24} \cos(3\tau) + \left(\frac{\omega_0 A^3}{8} + 2A\omega_1\right) \cos(\tau)$



It was already shown that the pendulum has an amplitude dependent frequency, so the frequency has to be developed as well (Poincaré-Linstead method):

 $\omega = \omega_0 + \epsilon \omega_1 + \epsilon^2 \omega_2 + \dots$

 Assume that the solution is a periodic function of τ = ωt which becomes the new independent variable. The equation at zero order gives the solution u₀(τ) = A cos(τ) and at leading perturbation order becomes
 ω₀ d²u₁/dτ² + ω₀u₁ = -2ω₁ d²u₀/dτ² + ω₀/6 u₀³ = ω₀A³/24 cos(3τ) + (ω₀A³/8 + 2Aω₁) cos(τ)
 The last term has to be zero, if not it gives secular terms_{ω1}thus A²ω₀/16 which reveals the reduction of

the frequency with the oscillation amplitude

Finally, the solution $u_1(t) = \frac{A^3}{192} \left(\cos(\omega_0 t) - \cos(3\omega_0 t) \right)$ is the leading order correction due to the non-linear term 68

Perturbation by periodic function



- In beam dynamics, perturbing fields are **periodic functions**
- The problem to solve is a generalization of the driven harmonic oscillator, $\frac{d^2u}{dt^2} + \omega_0^2 u(t) = g(t)$ with a general periodic function g(t), with frequency ω

 $m = +\infty$ The right side can be **Fourier analyzed**: $g(t) = \sum a_m e^{im\omega t}$ $m = -\infty$

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 $m = +\infty$

The right side can be **Fourier analyzed**: $g(t) = \sum a_m e^{im\omega t}$

- The homogeneous solution is $u_h(t) = u_0(t) \sin(\omega_0 t + \phi_0)$
- The **particular** solution can be found by considering that u(t)has the same form asg(t) : $u_p(t) = \sum u_{pm}e^{im\omega t}$
- By substituting the following relation is derived for the Fourier coefficients of the particular solution $u_{pm} = \frac{a_m}{\omega_0^2 - m^2 \omega^2}$
 - There is a **resonance condition** for infinite number of frequencies satisfying $\omega_0^2 = m^2 \omega^2$

Perturbation by single dipole



Hill's equations in normalized coordinates with single dipole perturbation:

$$\frac{d^2\mathcal{U}}{d\phi^2} + \nu_0^2\mathcal{U} = \nu_0^2\beta^{3/2}b_1(\phi) = \overline{b_1}(\phi)$$

The dipole perturbation is periodic, so it can be expanded in a Fourier series

$$\overline{b_1}(\phi) = \sum_{m=-\infty}^{\infty} \overline{b_{1m}} e^{im\phi}$$

Perturbation by single dipole



Hill's equations in normalized coordinates with single dipole perturbation:

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The dipole perturbation is periodic, so it can be expanded in a Fourier series

$$\overline{b_1}(\phi) = \sum_{m=1}^{\infty} \overline{b_{1m}} e^{im\phi}$$

Note, as before that a periodic kick introduces infinite number of integer driving frequencies

The resonance condition occurs when $\nu_0 = m$ i.e. **integer tunes** should be avoided (remember orbit distortion due to single dipole kick)


For a generalized multi-pole perturbation, Hill's equation is: $\frac{d^2 \mathcal{U}}{d\phi^2} + \nu_0^2 \mathcal{U} = \nu_0^2 \beta^{\frac{n}{2}+1} b_n(\phi) \mathcal{U}^{n-1} = \overline{b_n}(\phi) \mathcal{U}^{n-1}$ $\blacksquare \text{ As before, the multipole coefficient}_{\overline{b_n}(\phi)} = \sum_{n=1}^{\infty} \overline{b_{nm}} e^{im\phi}$

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As before, the multipole coefficient can be expanded in Fourier series

$$\overline{b_n}(\phi) = \sum_{m=-\infty}^{\infty} \overline{b_{nm}} e^{im\phi}$$

Following the perturbation steps, the zero-order solution is given by the homogeneous equation U₀ = W₁e^{iν₀φ} + W₋₁e^{-iν₀φ}
 Then the position can be expressed as

$$\mathcal{U}_0^{n-1} = \sum_{k=0}^{n-1} \binom{n-1}{k} W_1^{n-1-k} W_{-1}^k e^{i(n-1-2k)\nu_0 \phi}$$



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Following the perturbation steps, the zero-order solution is given by the homogeneous equation $U_0 = W_1 e^{i\nu_0\phi} + W_{-1} e^{-i\nu_0\phi}$ Then the position can be expressed as

 $\overline{W}_{q} \qquad q = -n+1, -n+3, \dots, n-1$ $\mathcal{U}_{0}^{n-1} = \sum_{k=0}^{n-1} \binom{n-1}{k} W_{1}^{n-1-k} W_{-1}^{k} e^{i(n-1-2k)\nu_{0}\phi} = \sum_{q=-n+1}^{n-1} \overline{W}_{q} e^{iq\nu_{0}\phi}$ a=-n+1

with $\overline{W}_{n-2} = \overline{W}_{n-4} = \overline{W}_{n-6} = \cdots = \overline{W}_{-n+2} = 0$



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The first order solution is written as n-1 $m = \infty$ $\frac{d^2\mathcal{U}_1}{d\phi^2} + \nu_0^2\mathcal{U}_1 = \overline{b_n}(\phi)\mathcal{U}_0^{n-1} = \sum_{i=1}^{n-1} \sum_{m=\infty}^{m=\infty} \overline{b_{nm}}\overline{W}_q e^{i(m+q\nu_0)\phi}$ 76 $q = -n + 1 m = -\infty$



Following the discussion on the periodic perturbation, the solution can be found by setting the leading order solution to be periodic with the same frequency as the right hand

side

n-1 $m = \infty$ $\mathcal{U}_1 = \sum \qquad \sum \quad \mathcal{U}_{1mq} e^{i(m+q\nu_0)\phi}$ $q = -n + 1 m = -\infty$



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$$\mathcal{U}_1 = \sum_{q=-n+1}^{n-1} \sum_{m=-\infty}^{m=\infty} \mathcal{U}_{1mq} e^{i(m+q\nu_0)\phi}$$

Equating terms of equal exponential powers, the Fourier amplitudes are found to satisfy the relationship

$$\mathcal{U}_{1mq} = \frac{\nu_{nm} \nu_{q}}{\nu_{0}^{2} - (m + q\nu_{0})^{2}}$$



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side

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$$\mathcal{U}_1 = \sum_{q=-n+1}^{n-1} \sum_{m=-\infty}^{m=\infty} \mathcal{U}_{1mq} e^{i(m+q\nu_0)q}$$

Equating terms of equal exponential powers, the Fourier amplitudes are found to satisfy the relationship

$$\mathcal{U}_{1mq} = \frac{\nu_{nm} \nu_{q}}{\nu_{0}^{2} - (m + q\nu_{0})^{2}}$$

 $m\pm |q|\nu_0=\nu_0$ This provides the resonance condition

or $\nu_0 = \overline{1 \pm |q|}$ which means that there are resonant

frequencies for and "infinite" number of rationals

Tune-shift for single multi-pole



Note that for even multi-poles and q=1 or m=0, there is a Fourier coefficient \overline{b}_{n0} , which is independent of ϕ and represents the average value of the periodic perturbation

The perturbing term in the r.h.s. is

$$\overline{b}_{n0}\overline{W}_{1}e^{i\nu_{0}\phi} = \nu_{0}^{2}\beta^{\frac{n}{2}+1}b_{n0}\binom{n-1}{\frac{n}{2}-1}W_{1}^{n-1}W_{-1}^{\frac{n}{2}-1}e^{i\nu_{0}\phi}$$

which can be obtained for $k = \frac{n}{2} - 1$ (it is indeed an integer only for even multi-poles)

Tune-shift for single multi-pole



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which can be obtained for $k = \frac{n}{2} - 1$ (it is indeed an integer only for even multi-poles)

Following the approach of the perturbed non-linear harmonic oscillator, this term will be secular unless a perturbation in the frequency is considered, thereby resulting to a **tune-shift** equal to

$$\nu = -\frac{\nu_0 \beta^{\frac{n}{2}+1} b_{n0}}{2} \binom{n-1}{\frac{n}{2}-1} \widetilde{W}^{n-2} \quad \text{with} \quad \widetilde{W}^2 = W_1 W_{-1}$$

This tune-shift is amplitude dependent for n > 2

Example: single quadrupole perturbation



Consider single quadrupole kick in one normalized plane: $\frac{d^2 \mathcal{U}}{d\phi^2} + \nu_0^2 \mathcal{U} = \nu_0^2 \beta^2 b_2(\phi) \mathcal{U} = \overline{b_2}(\phi) \mathcal{U}$ The quadrupole perturbation can be expanded in a Fourier series

$$\overline{b_2}(\phi) = \sum_{m = -\infty} \overline{b_{2m}} e^{im\phi}$$

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i.e. integer and half-integer tunes should be avoided



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Single Sextupole Perturbation



Consider a localized sextupole perturbation in the horizontal plane

$$\frac{d^2\mathcal{U}}{d\phi^2} + \nu_0^2\mathcal{U} = \nu_0^2\beta^{\frac{5}{2}}b_3(\phi)\mathcal{U}^2 = \overline{b_3}(\phi)\mathcal{U}^2$$

After replacing the perturbation by its Fourier transform and inserting the unperturbed solution to the right hand side $\frac{d^2 \mathcal{U}_1}{d\phi^2} + \nu_0^2 \mathcal{U}_1 = \sum_{q=-2}^2 \sum_{m=-\infty}^\infty \overline{W}_q \overline{b_{3m}} e^{i(m+q\nu_0)\phi} \text{ with } \overline{W}_{-1} = \overline{W}_1 = 0$ Single Sextupole Perturbation



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After replacing the perturbation by its Fourier transform and inserting the unperturbed solution to the right hand side $\frac{1}{2}$

$$\frac{d^2 \mathcal{U}_1}{d\phi^2} + \nu_0^2 \mathcal{U}_1 = \sum_{q=-2}^{2} \sum_{m=-\infty}^{\infty} \overline{W}_q \overline{b_{3m}} e^{i(m+q\nu_0)\phi} \text{ with } \overline{W}_{-1} = \overline{W}_1 = 0$$

$$\mathbf{3^{rd} integer} \to 3\nu_0 = m \text{ for } q = -2$$

Resonance conditions: integer $\rightarrow \nu_0 = m$ for q = 0, 2

Single Sextupole Perturbation



Consider a localized sextupole perturbation in the horizontal plane

$$\frac{d^2\mathcal{U}}{d\phi^2} + \nu_0^2\mathcal{U} = \nu_0^2\beta^{\frac{5}{2}}b_3(\phi)\mathcal{U}^2 = \overline{b_3}(\phi)\mathcal{U}^2$$

After replacing the perturbation by its Fourier transform and inserting the unperturbed solution to the right hand side $\frac{d^2 \mathcal{U}_1}{d\phi^2} + \nu_0^2 \mathcal{U}_1 = \sum_{q=-2}^2 \sum_{m=-\infty}^\infty \overline{W}_q \overline{b_{3m}} e^{i(m+q\nu_0)\phi} \text{ with } \overline{W}_{-1} = \overline{W}_1 = 0$ $\mathbf{3^{rd} integer} \to 3\nu_0 = m \text{ for } q = -2$

Resonance conditions: integer $\rightarrow \nu_0 = m$ for q = 0, 2Note that there is **not a tune-spread associated**. This is only true for small perturbations (**first order** perturbation treatment)

Although perturbation treatment can provide approximations for evolution of motion, there is **no exact solution** General multi-pole perturbation



Equations of motion including any multi-pole error term, in both planes

$$\frac{d^2 \mathcal{U}_x}{d\phi_x^2} + \nu_{0x}^2 \mathcal{U}_x = \overline{b_{n,r}}(\phi_x) \mathcal{U}_x^{n-1} \mathcal{U}_y^{r-1}$$

Expanding perturbation coefficient in Fourier series and inserting the solution of the unperturbed system on the rhs gives the following series: $\mathcal{U}_{x}^{n-1} \approx \mathcal{U}_{0x}^{n-1} = \sum_{n=1}^{n-1} \overline{W}_{q_{x}} e^{iq_{x}\nu_{0}\phi_{x}}$

General multi-pole perturbation



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$$\frac{d^2\mathcal{U}_x}{d\phi_x^2} + \nu_{0x}^2\mathcal{U}_x = \sum_{m,q_x,q_y} \overline{b_{nrm}} W_{q_x}^x W_{q_y}^y e^{i(m+q_x\nu_{0x}+q_y\nu_{0y})\phi_x}$$

In principle, same perturbation steps can be followed for getting an approximate solution in both planes Example: Linear Coupling



For a localized skew quadrupole we have

$$\frac{d^2 \mathcal{U}_x}{d\phi_x^2} + \nu_{0x}^2 \mathcal{U}_x = \overline{b_{1,2}}(\phi_x)\mathcal{U}_y$$

Expanding perturbation coefficient in Fourier series and inserting the solution of the unperturbed system gives the following equation:

$$\frac{d^2 \mathcal{U}_x}{d\phi_x^2} + \nu_{0x}^2 \mathcal{U}_x = \sum_{m=-\infty}^{\infty} \sum_{q_y=-1}^{q_y=1} \overline{b_{12m}} W_{q_y}^y e^{i(m+q_y\nu_{0y})\phi_x} \text{ with } W_0^y = 0$$

Example: Linear Coupling



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The coupling resonance are found for $q_y = \pm 1$ Linear sum resonance Linear difference resonance

$$m = \nu_{0x} + \nu_{0y}$$

$$m = \nu_{0x} - \nu_{0y}$$

General resonance conditions



The general resonance conditions is $m + q_x \nu_{0x} + q_y \nu_{0y} = \nu_{0x}$ or $m + q'_x \nu_{0x} + q_y \nu_{0y} = 0$, with order $|q_x| + |q_y| + 1$

The same condition can be obtained in the vertical plane

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the order grows (the rational numbers form a dense set inside the real numbers), but Fourier amplitudes should ^o



Systematic and random resonance



If lattice is made out of N identical cells, and the perturbation follows the same periodicity, resulting in a reduction of the resonance conditions to the ones satisfying $q_x \nu_{0x} + q_y \nu_{0y} = jN$ These are called systematic resonances

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Practically, any (linear) lattice perturbation breaks super-periodicity and any random resonance can be excited

Careful choice of the working point is necessary



Contents of the 1st lecture



- Accelerator performance parameters and non-linear effects
- Linear and non-linear oscillators
 - □ Integral and frequency of motion
 - □ The pendulum
 - Damped harmonic oscillator
- Phase space dynamics
 - Fixed point analysis
- Non-autonomous systems
 - Driven harmonic oscillator, resonance conditions
- Linear equations with periodic coefficients Hill's equations
 - Floquet solutions and normalized coordinates
- Perturbation theory
 - Non-linear oscillator
 - Perturbation by periodic function single dipole perturbation
 - □ Application to single multipole resonance conditions
 - □ Examples: single quadrupole, sextupole, octupole perturbation
 - General multi-pole perturbation example: linear coupling
 - Resonance conditions and working point choice
 - Summary

Non-linear effects, JUAS, February 2017

Appendix: Damped harmonic oscillator





- Accelerator performance depends heavily on the understanding and control of non-linear effects
- The ability to integrate differential equations has a deep impact to the dynamics of the system
- Phase space is the natural space to study this dynamics
- Perturbation theory helps integrate iteratively differential equations and reveals appearance of resonances
- Periodic perturbations drive infinite number of resonances
- There is an amplitude dependent tune-shift at 1st order for even multi-poles
- Periodicity of the lattice very important for reducing number of lines excited at first order

Damped harmonic oscillator I



Damped harmonic oscillator:

$$\frac{d^2u(t)}{dt^2} + \frac{\omega_0}{Q}\frac{du(t)}{dt} + \omega_0^2u(t) = 0$$

 $\Box Q = \frac{1}{2\zeta}$ is the ratio between the stored and lost energy per cycle with the damping ratio

 $\Box \, \omega_0$ is the eigen-frequency of the harmonic oscillator

General solution can be found by the same ansatz $u(t) = e^{\lambda t}$

leading to an auxiliary 2nd order equation

$$\lambda^2 + rac{\omega_0}{Q}\lambda + \omega_0^2 = 0$$
 with solutions
 $\lambda_{\pm} = -rac{\omega_0}{2Q}(-1 \pm \sqrt{1 - 4Q^2}) = -\omega_0\zeta(-1 \pm \sqrt{1 - rac{1}{\zeta^2}})_{\text{sec}}$

Damped harmonic oscillator II



- Three cases can be distinguished
 - □ Overdamping (real ζ i.e. 1 Q or 1/2): The system exponentially decays to equilibrium (slower for larger damping ratio values)
 - □ Critical damping (ζ = 1): The system returns to equilibrium as quickly as possible without oscillating.
 - □ Underdampiàg (complex ζ i.e. 1 $Q > \Phi / 2$): The system oscillates with the amplitude gradually decreasing to zero, with a slightly different frequency / than 2^2



Note that there is no integral of motion, in that case, as the energy is not conserved (dissipative system) Damped oscillator with periodic driving



Consider periodic force pumping energy into the system

$$\frac{d^2u(t)}{dt^2} + \frac{\omega_0}{Q}\frac{du(t)}{dt} + \omega_0^2u(t) = \frac{F}{m}\cos(\omega t)$$

The solution of the homogeneous system is

 $u_h(t) = u_0(t)e^{-\omega_0\zeta t}\sin(\omega_0\sqrt{1-\zeta^2}t+\phi_0)$

Damped oscillator with periodic driving



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 $u_h(t) = u_0(t)e^{-\omega_0\zeta t}\sin(\omega_0\sqrt{1-\zeta^2}\ t+\phi_0)$ The particular solution is

$$u_p(t) = \frac{F\cos(\omega t + \phi'_0)}{m \,\omega_0^2 \sqrt{(1 - \frac{\omega^2}{\omega_0^2})^2 + 4\zeta^2 \frac{\omega^2}{\omega_0^2}}}$$

- The homogeneous solution vanishes $f dr \to \infty$, leaving only the particular one, for which there is an **amplitude** maximum for $\omega_0 = \omega$ but no divergence
- In that case, the energy pumped into the system compensates the friction, and a **steady state** is reached representing a **limit cycle**