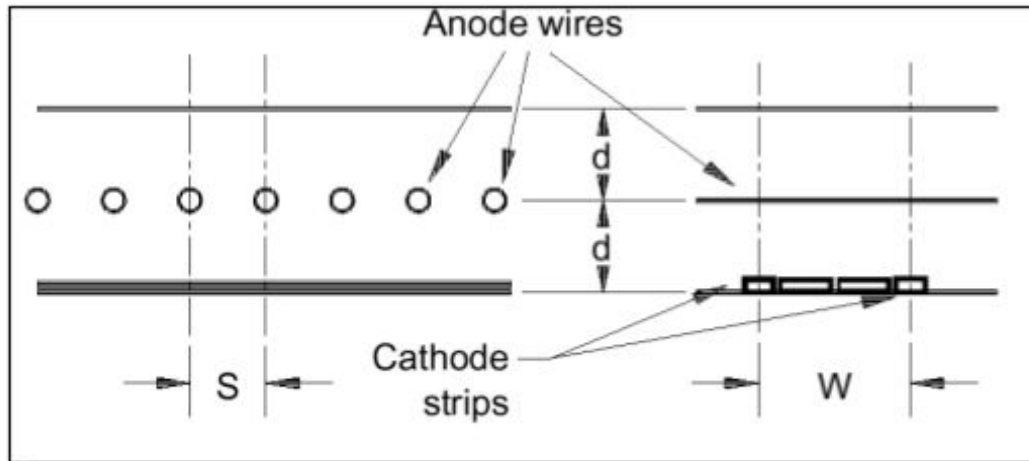


# Back to the Detectors

# Detectors(**Gaseous**)

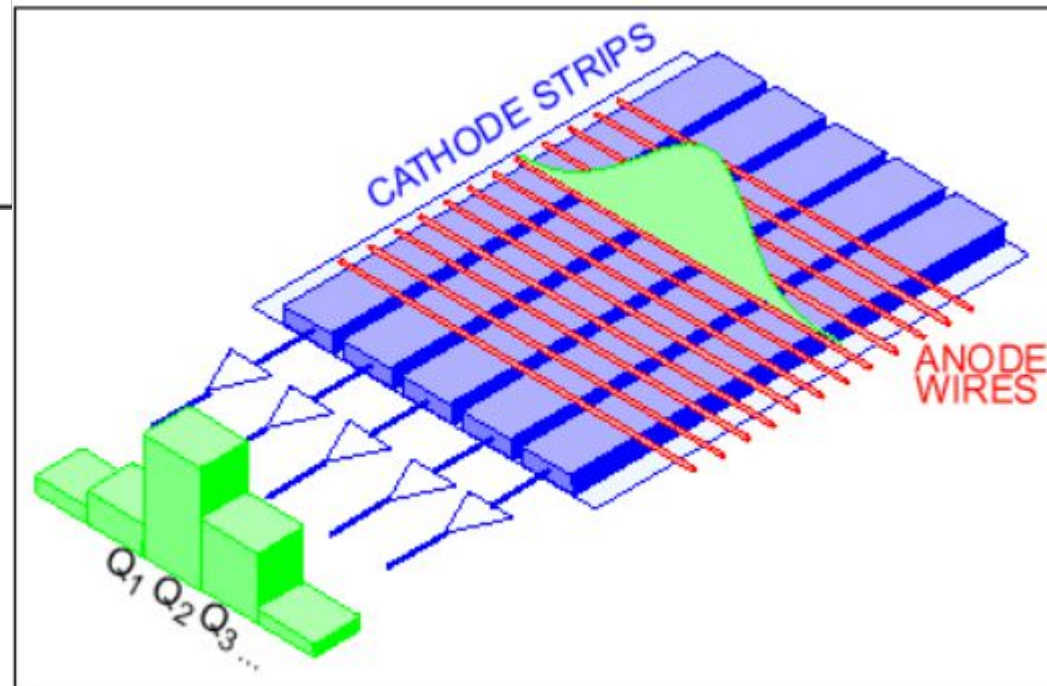
## CSC: Cathode Strip Chamber

- Determine muon position by interpolating the charge on 3 to 5 adjacent strips
- Precision (x-) strip pitch  $\sim 5.6$  mm
- Measure  $Q_1, Q_2, Q_3 \dots$  with 150:1 SNR to get  $\sigma_x \sim 60 \mu\text{m}$ .
- Second set of y-strips measure transverse coordinate to  $\sim 1$  cm.
- Position accuracy unaffected by gas gain or drift time variations



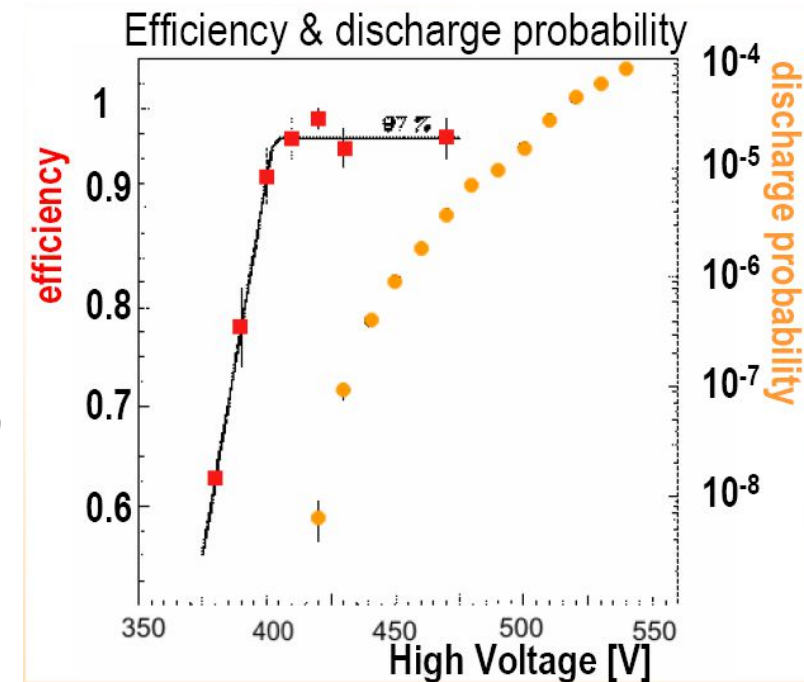
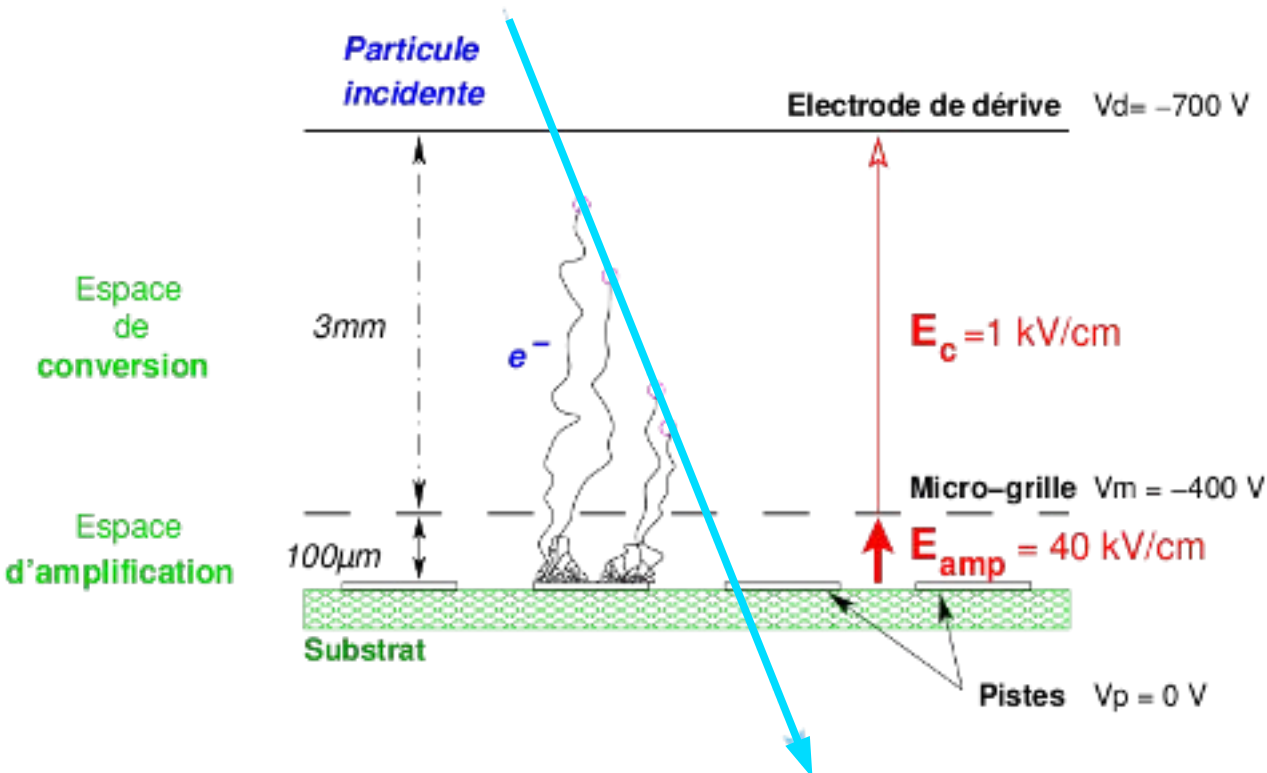
$S = 2.54$  mm

$W = 5.60$  mm



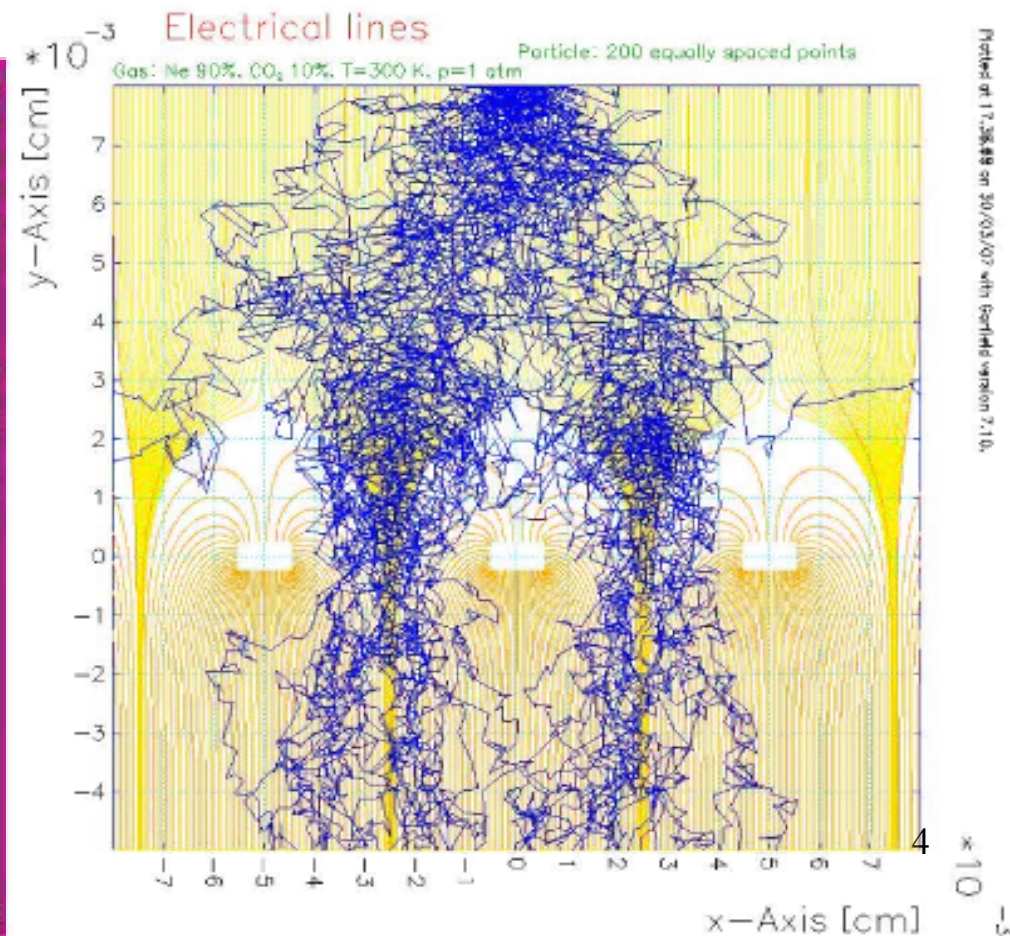
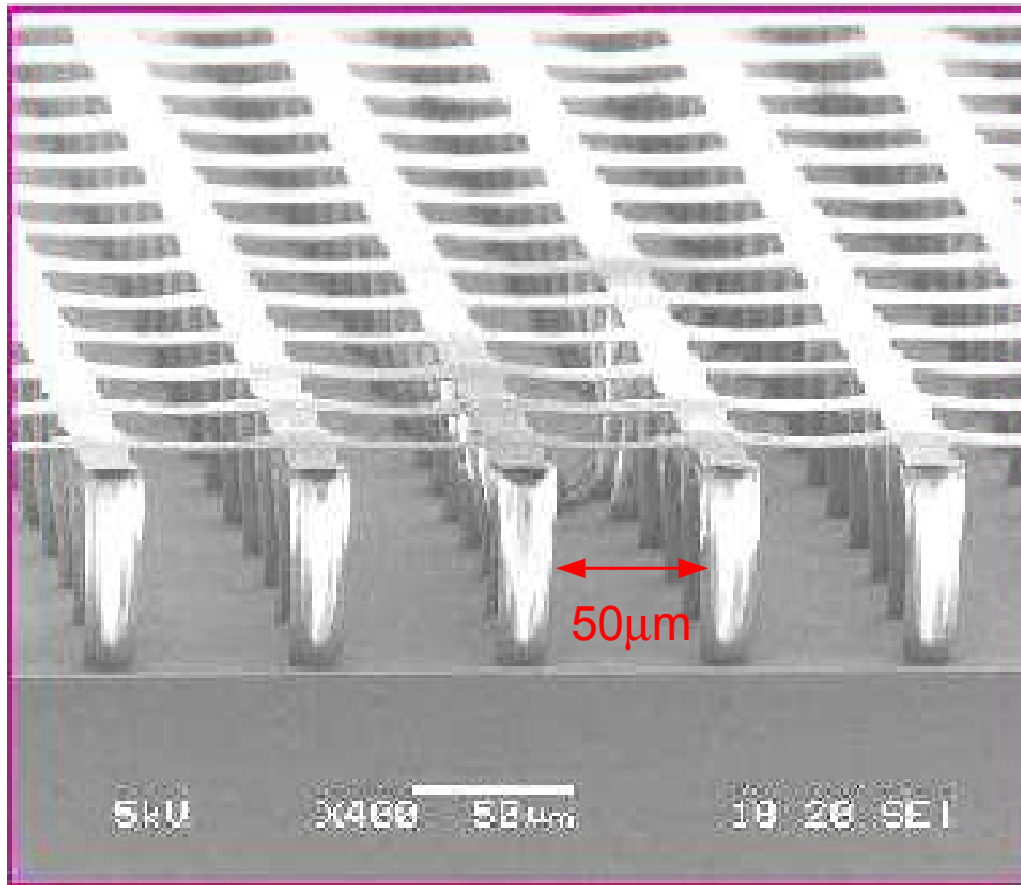
### Micro-Megas

- Giomataris I. et al., NIMA 376 (1996) 29
- Capable of operating at very high rates
- Work in magnetic fields
- Radiation hard and age well.
- Shape and readout segmentation can be adapted to the needs
- Parallel plate structures with straight-forward field shapes.
- Work at very low HV (thin amplification gap)
- Industrially produced



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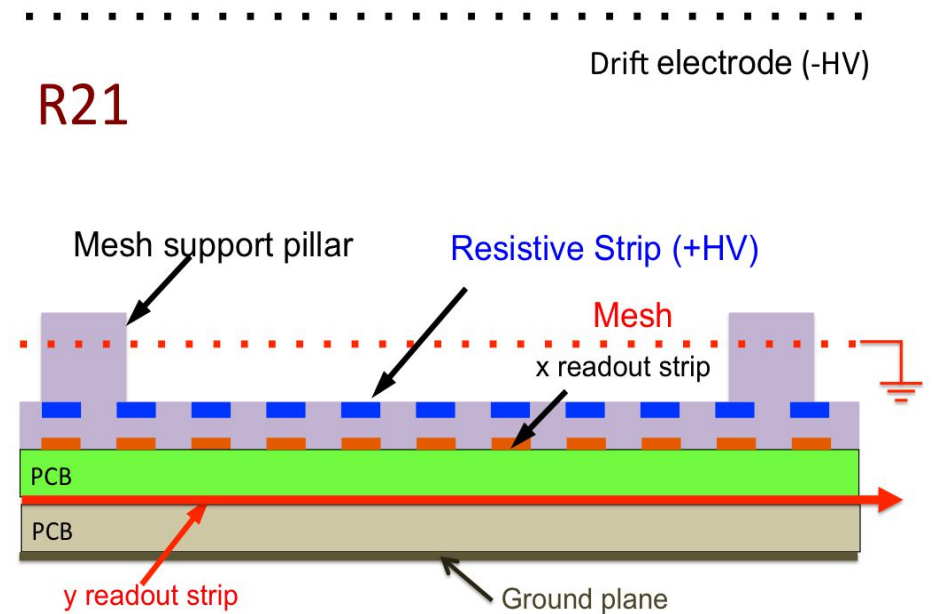
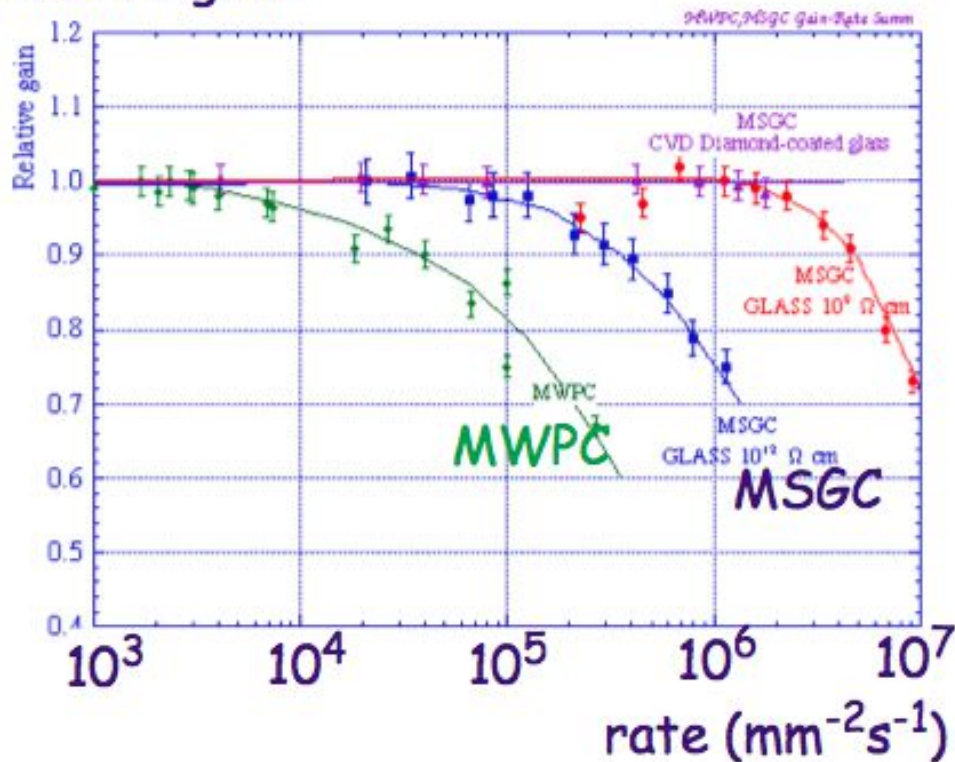




### Micro-Megas

- Capable of operating at very high rates
  - Short ion evacuation path => high rate capability
- Very precise readout structures produced using PCB technology (lithography)
- Very good spatial resolution
- Improvement (add of a layer of resistive strips above the readout structure: Spark tolerant without degrading their performance  
T. Alexopoulos et al., NIMA 640 (2011) 110-118

relative gain



# Interlude

## Muography

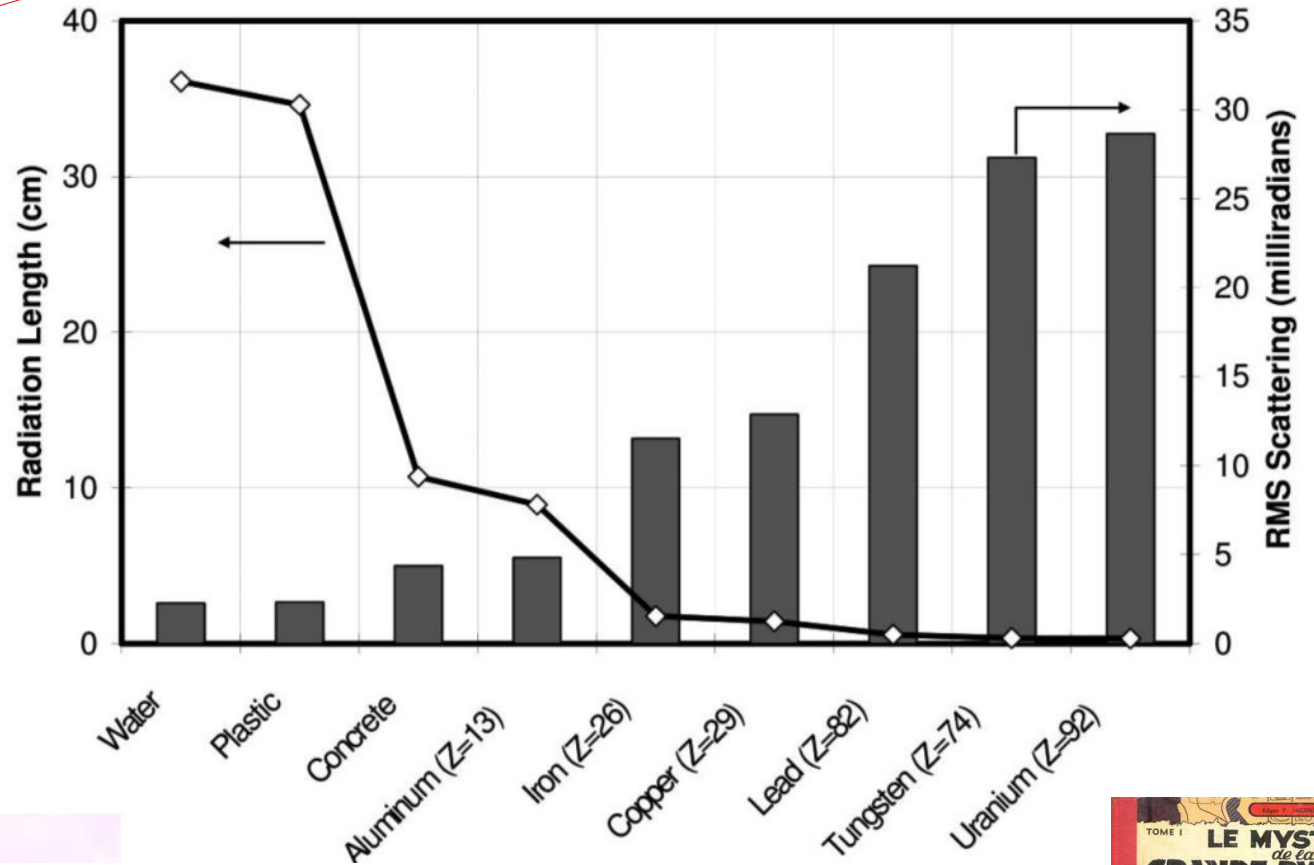
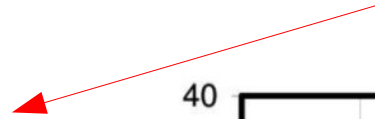
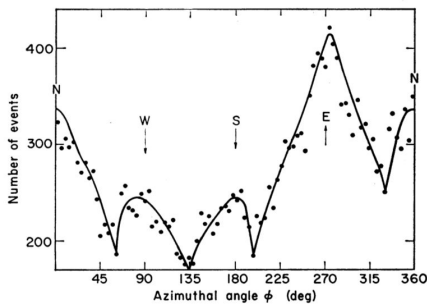
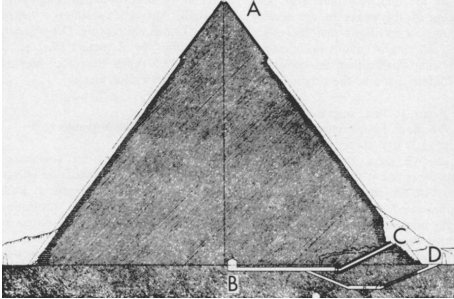
# Muography

- the probability of muon **absorption** is proportional to the density  
Muon flux → density map
- Use cosmic muons to analyse **Archaeology**, Volcanology, buildings structure,...

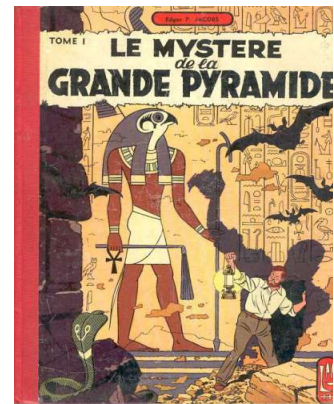
## Search for Hidden Chambers in the Pyramids

The structure of the Second Pyramid of Giza is determined by cosmic-ray absorption.

Luis W. Alvarez, Jared A. Anderson, F. El Bedwei,



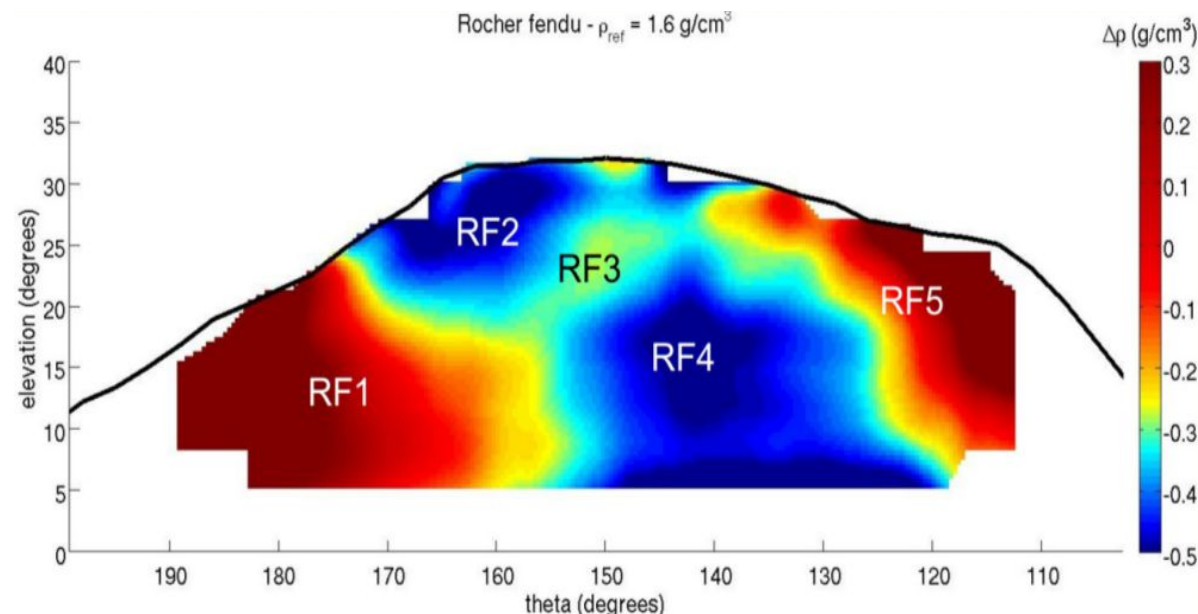
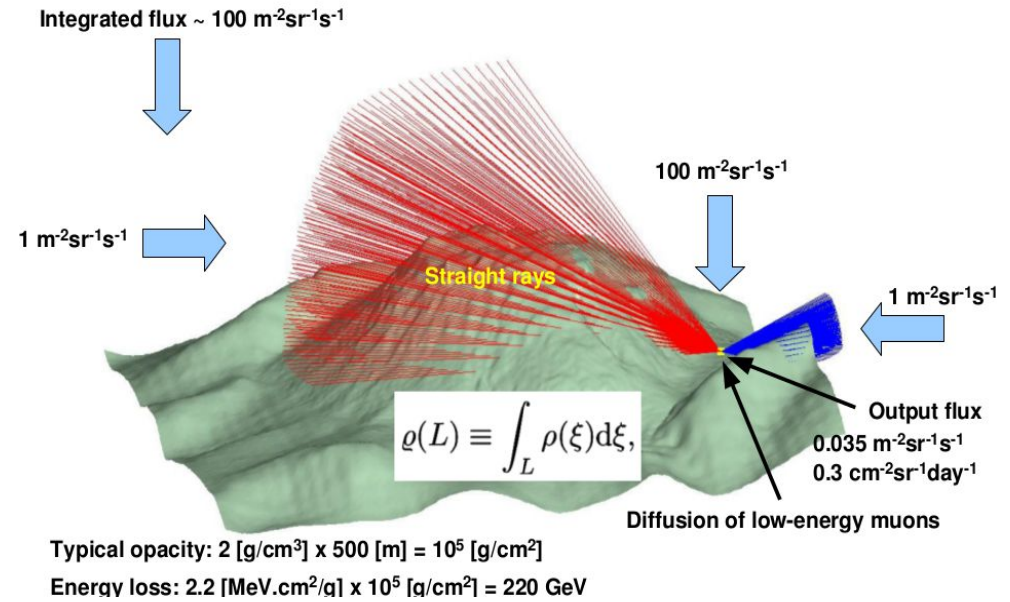
Hidden room in the pyramid? →



# Muography

## Micro Pattern Gas Detectors

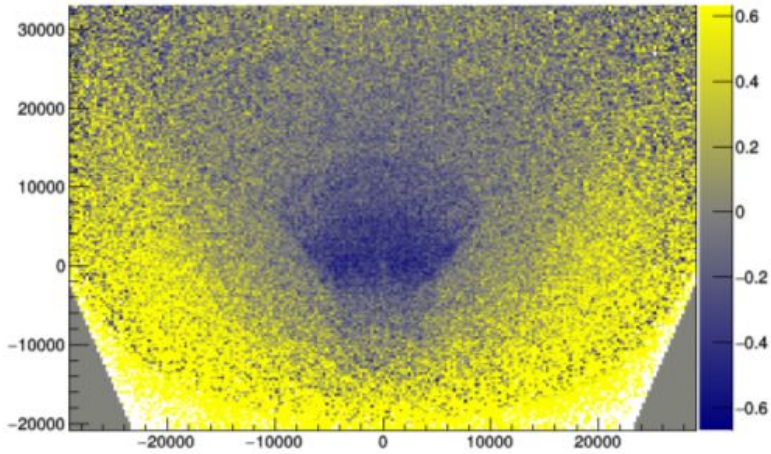
- the probability of muon **absorption** is proportional to the density  
Muon flux  $\rightarrow$  density map
- Use cosmic muons to analyse Archaeology, **Volcanology**, buildings structure,...





# Muography

- the probability of muon absorption is proportional to the density  
    Muon flux  $\rightarrow$  density map
- Use cosmic muons to analyse Archaeology, Volcanology, **buildings structure**,...

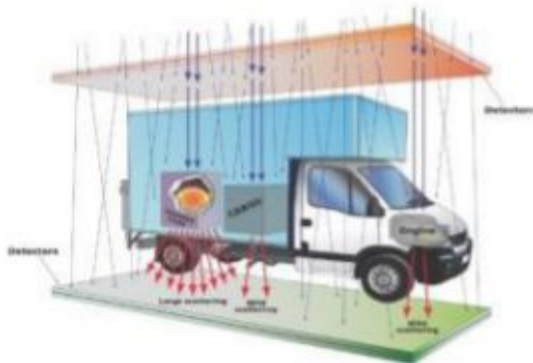
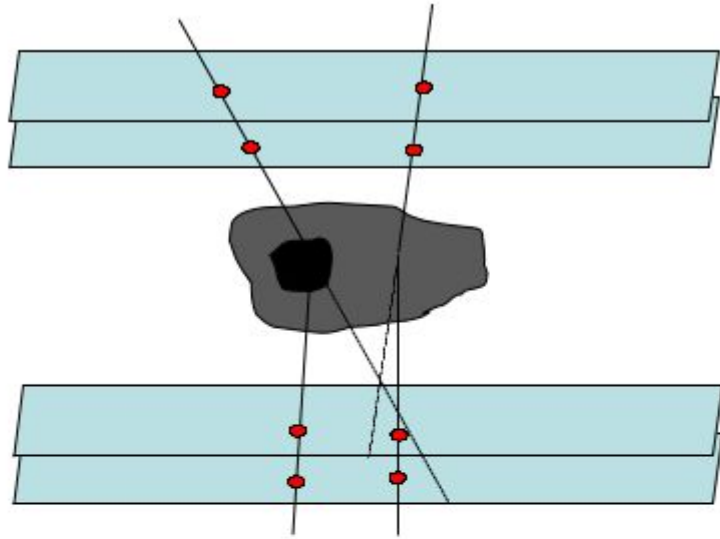


Micro-megas

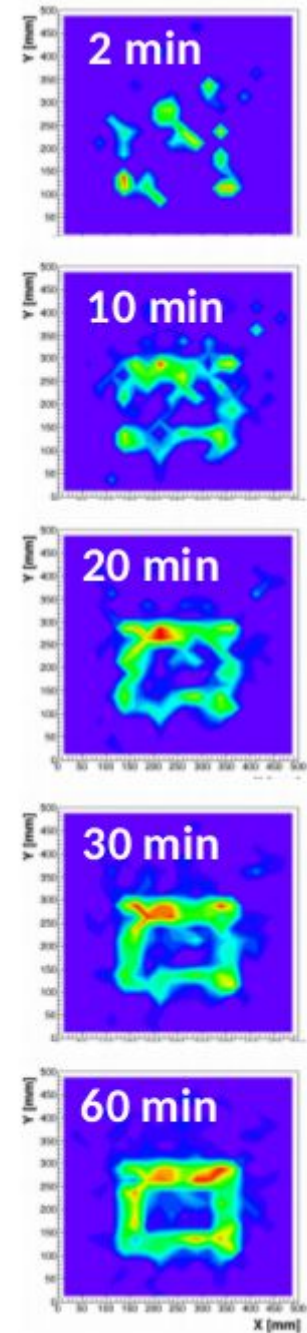


# Muography

- Use cosmic muons to analyse truck, container....
- **Multiple diffusion:**
  - 2 detectors: deviation angle
  - fast ( $\sim$ mn), 3D,

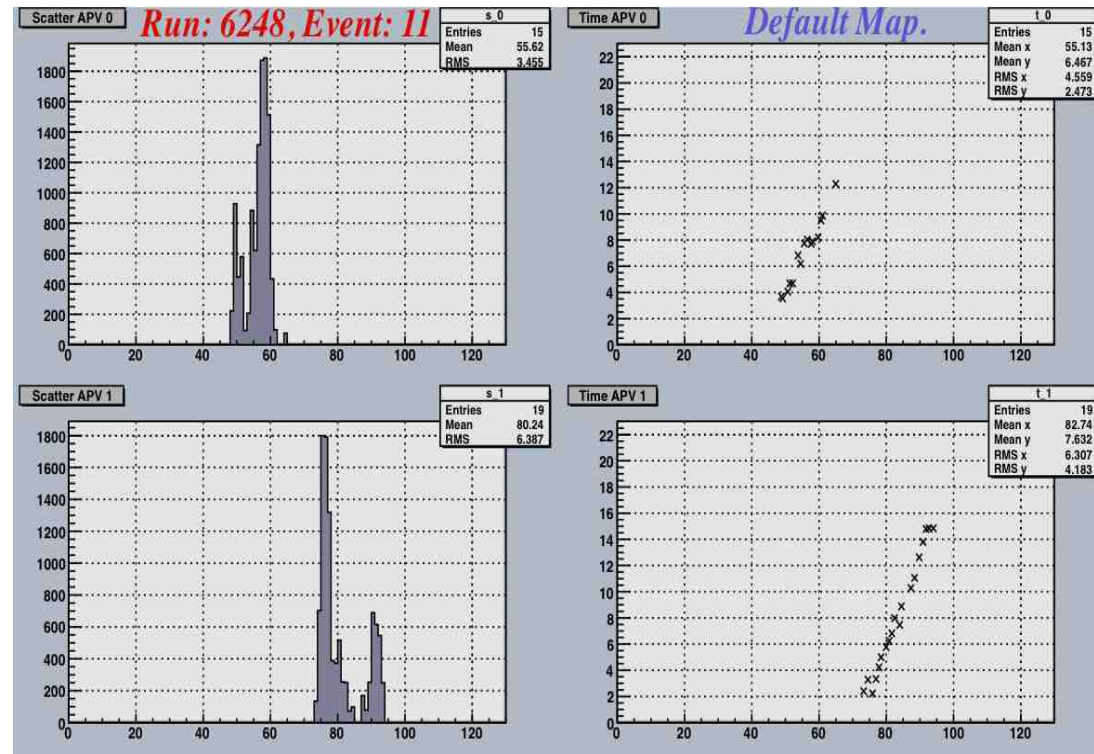
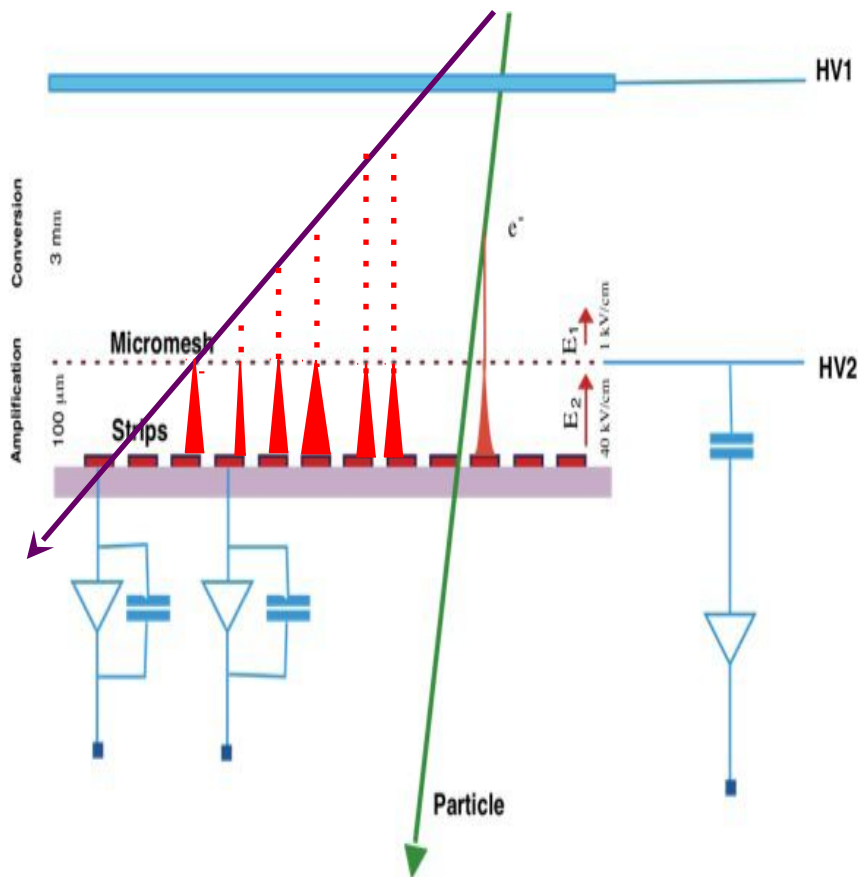


## Micro Pattern Gas Detectors



### Micro-Megas: $\mu$ TPC!!!!

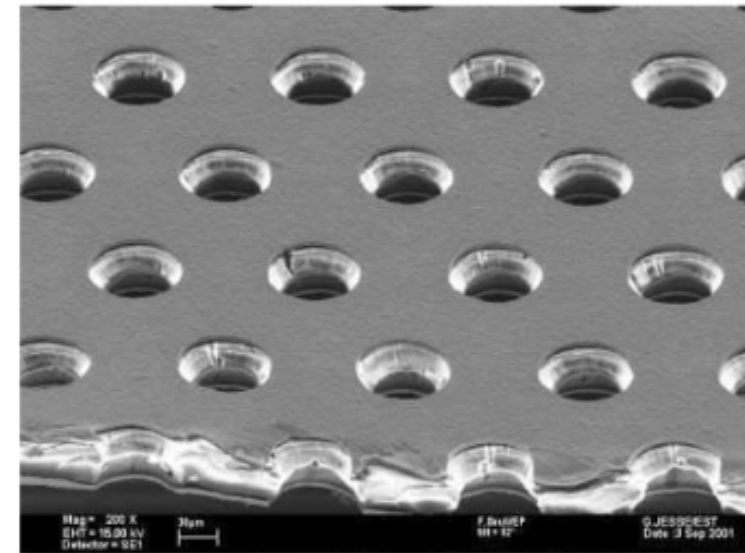
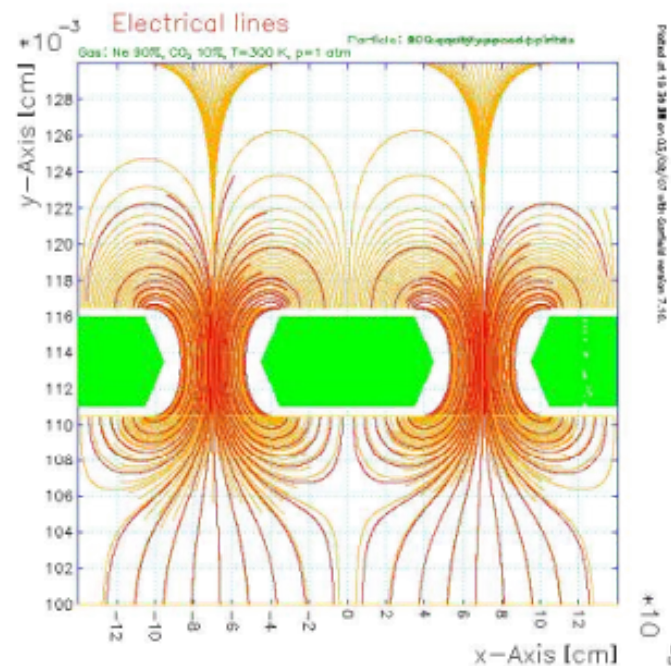
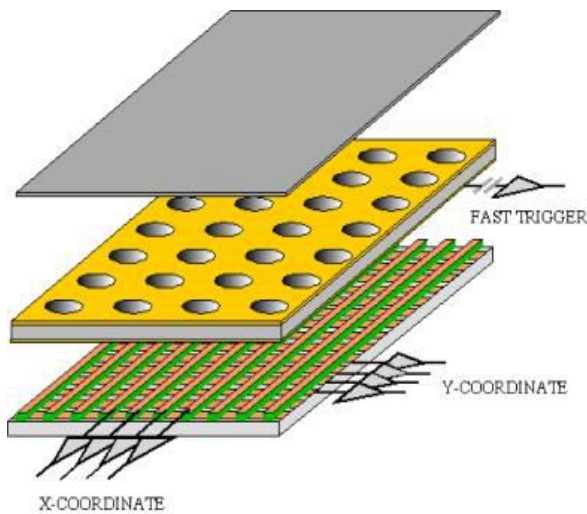
- "Wide" drift region (**typically a few mm**)
- Electric field of 100–1000 V/cm
- 100  $\mu$ m amplification gap with high electrical field (40–50 kV/cm)
  - a factor  $E_m/E_d \approx 70$ –100 is required for full mesh transparency for electrons
- Drift velocities of 5 cm/ $\mu$ s (or 20 ns/mm) electrons need 100 ns for a 5 mm gap
- Adding the time arrival of the signals => TPC-like
- Track vectors for inclined tracks





### GEM: Gas Electron Multiplier

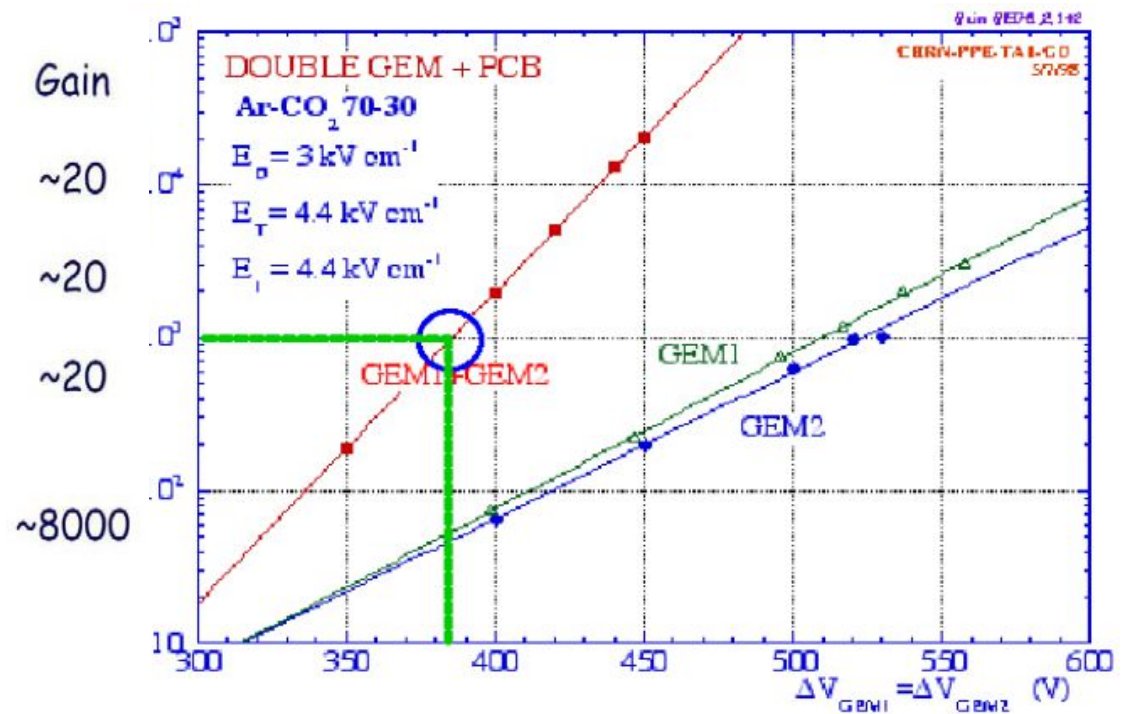
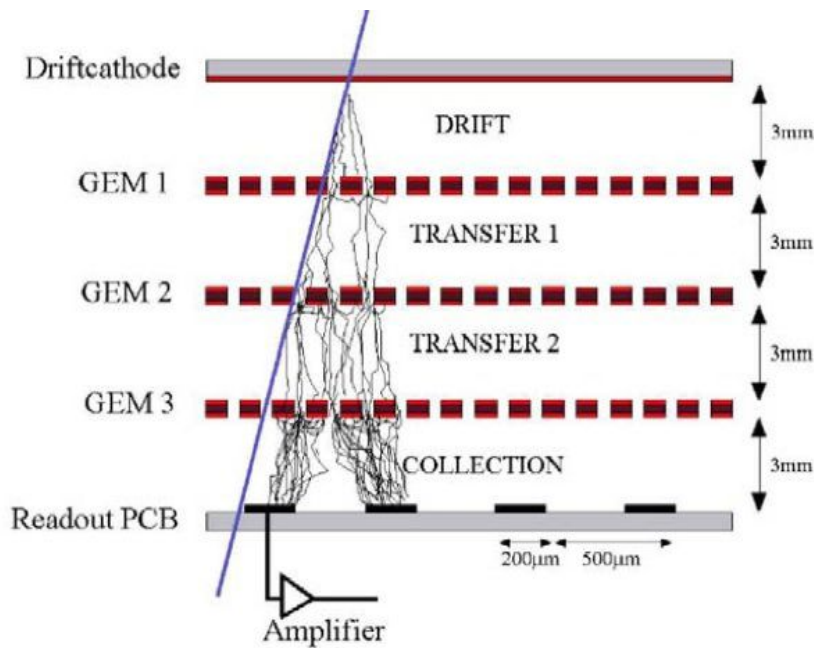
- F. Sauli at CERN, (R. Bouclier et al., NIM A 396 (1997) 50).
- Parallel plate structure with perforated Cu-clad Kapton foils.
- By applying a potential between conducting foil surfaces a strong electric field develops inside the holes
- Electron multiplication takes place in the field inside the holes
- Hole diameters are 70–120  $\mu\text{m}$
- Kapton foils are about 50  $\mu\text{m}$  thick





### GEM: Gas Electron Multiplier

- Triple GEM
- Lower voltage for the same gain
- Less spark

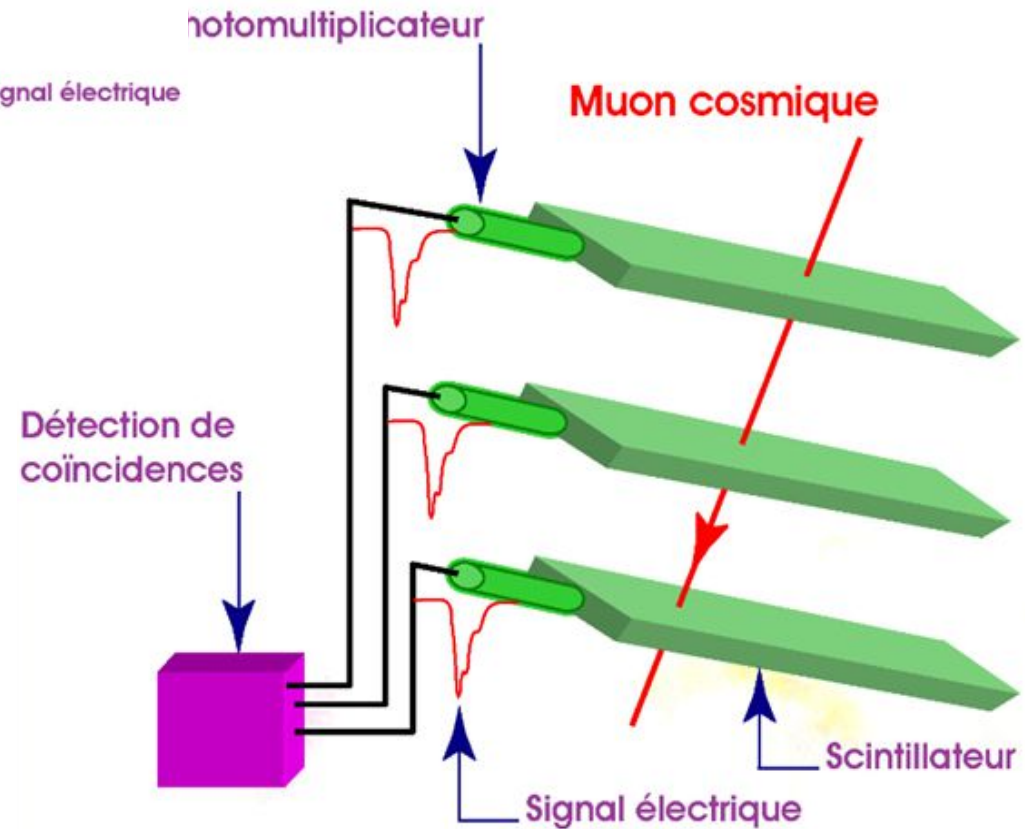
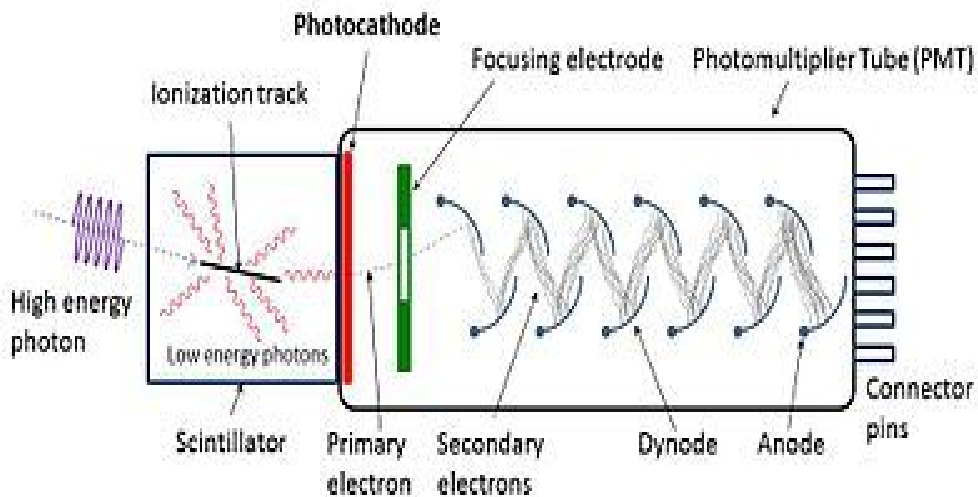
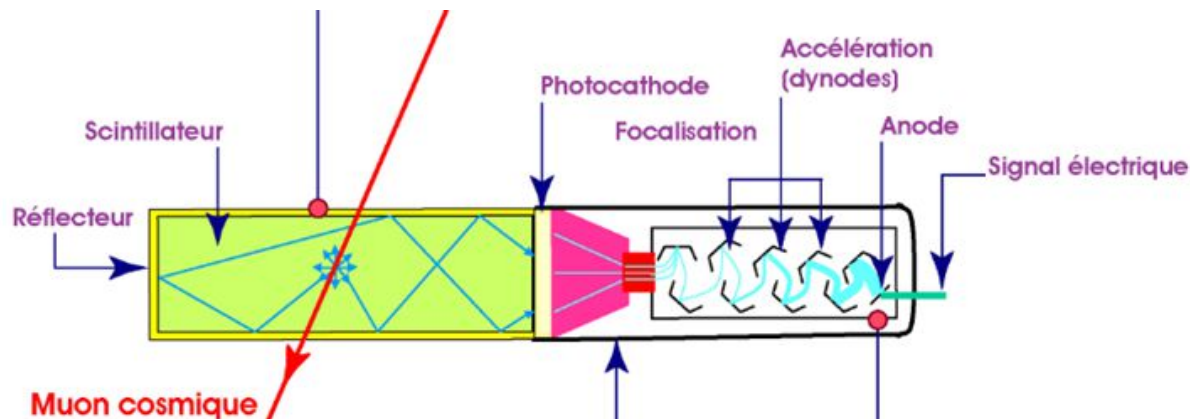


Detectors(Solid)

# Detectors(Solid)

## Scintillator

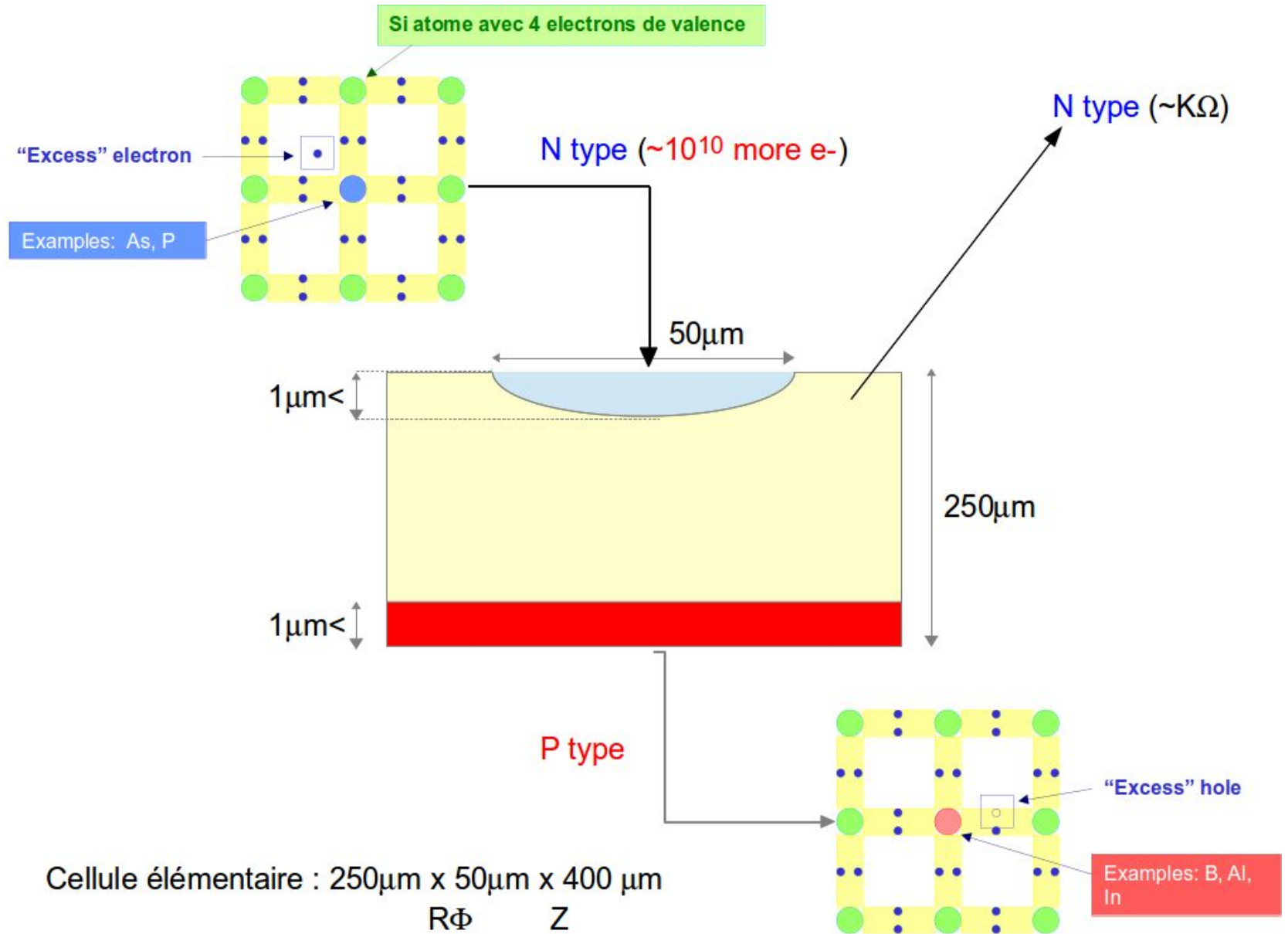
- Scintillation: atoms are excited by a muon
- Atoms are emitting photons which are detected by the photomultiplier.
- The scintillator is plastic (made from organic matter).



# Detectors(Solid)

## Silicon: Pixel

- Elementary cell



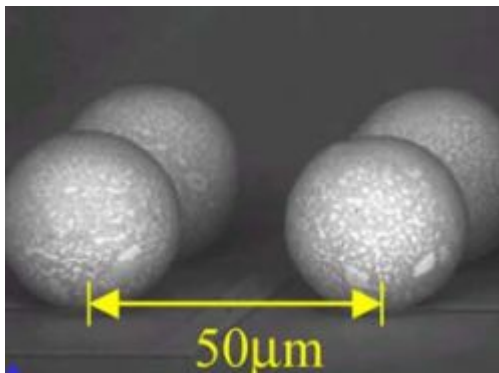
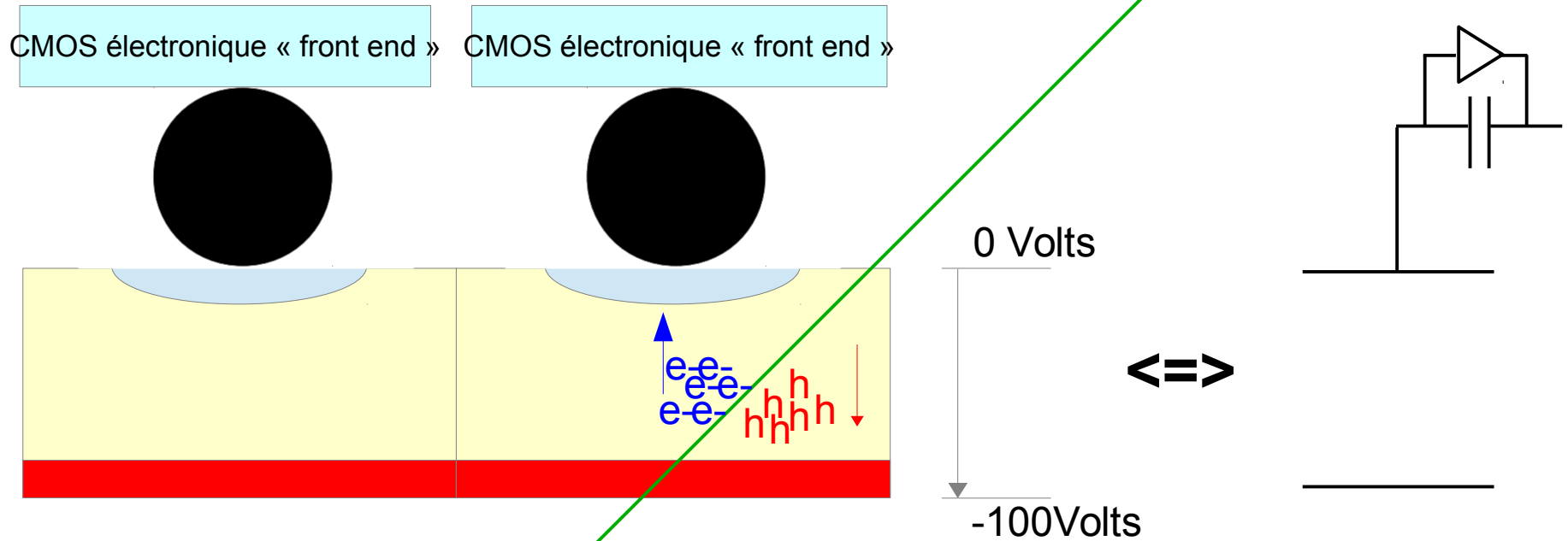
Cellule élémentaire :  $250\mu\text{m} \times 50\mu\text{m} \times 400\mu\text{m}$   
 $R\Phi$        $Z$



# Detectors(Solid)

## Silicon: Pixel

- N cells

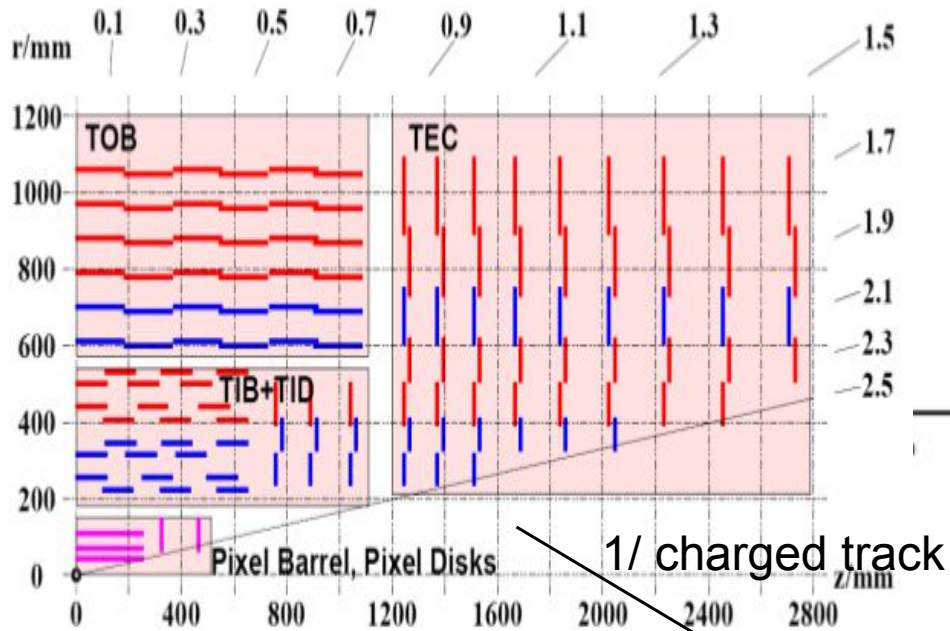
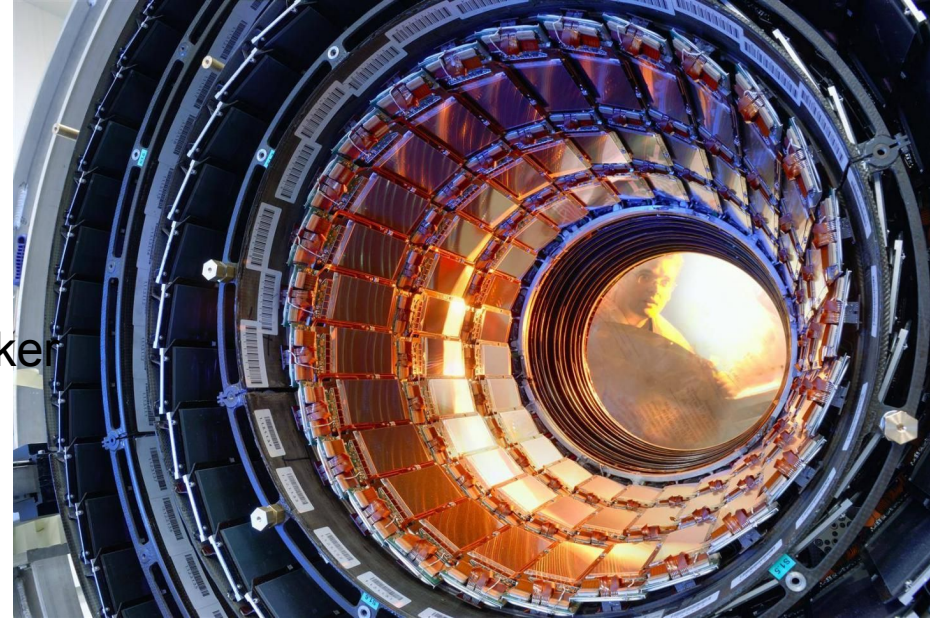


- ~20000 electrons per track
- Times electron collection ~ 5ns

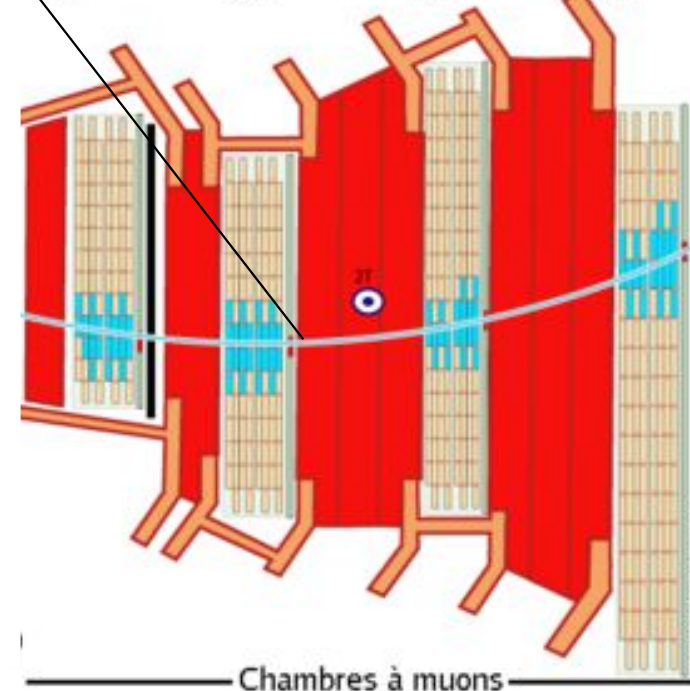
# Detectors(Solid)

## Silicon Tracker: CMS

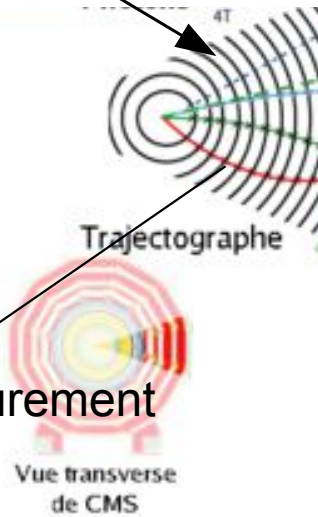
- 11 layers
- 200 m<sup>2</sup> of active silicon for CMS tracker



2/ Muon confirmation



3/ Muon momentum measurement

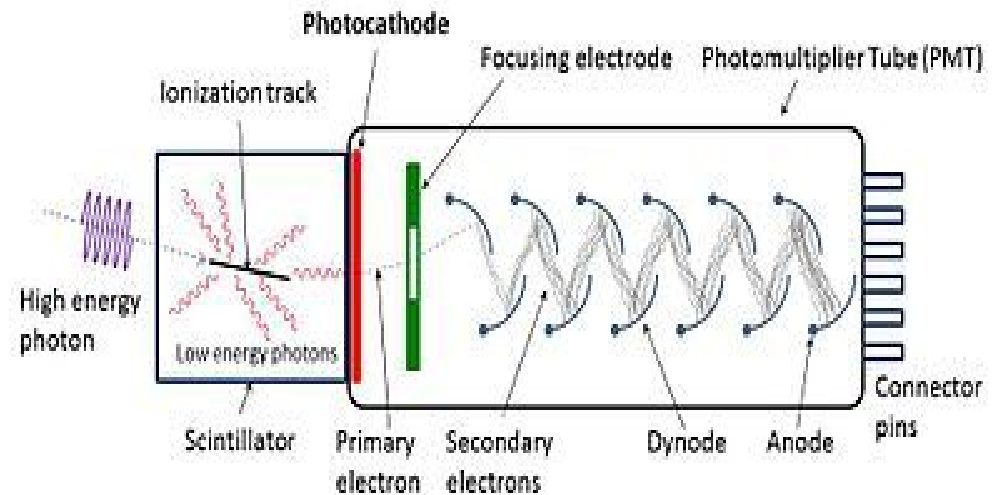
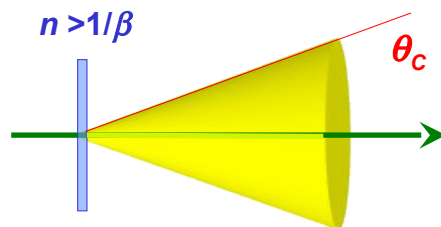
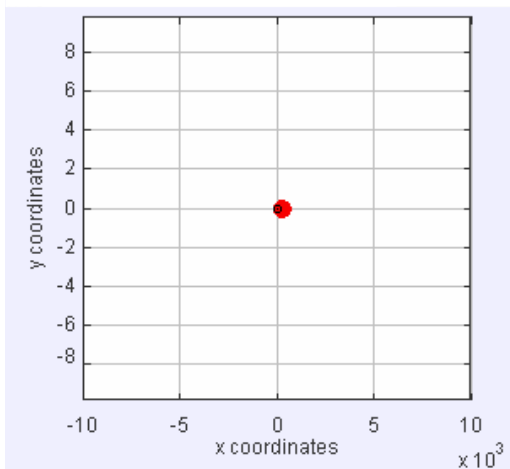
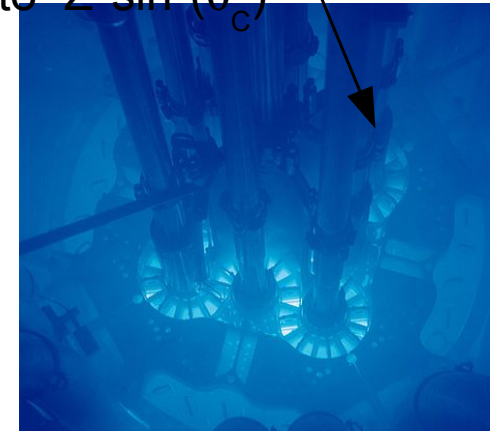
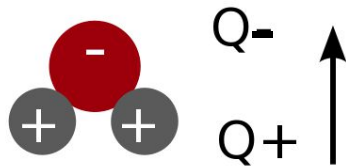


Detectors(Liquid)

# Detectors(Liquid)

## Cherenkov radiation: Photo Multiplier (PM)

- Relativistic charged particles through a medium of refractive index  $n > 1 / \beta$
- Relativistic means that the particle moves faster than the light in the medium
- Cherenkov radiation is tangent to a cone  $\theta_c$  around the trace:  $\cos(\theta_c) = 1 / n\beta$ 
  - Radiation is due to the polarization of the medium and a dynamic variation of the dipole moment of the molecules of the medium (ie water)
  - Number of photons (Frank-Tamm) is proportional to  $Z^2 \sin^2(\theta_c)$

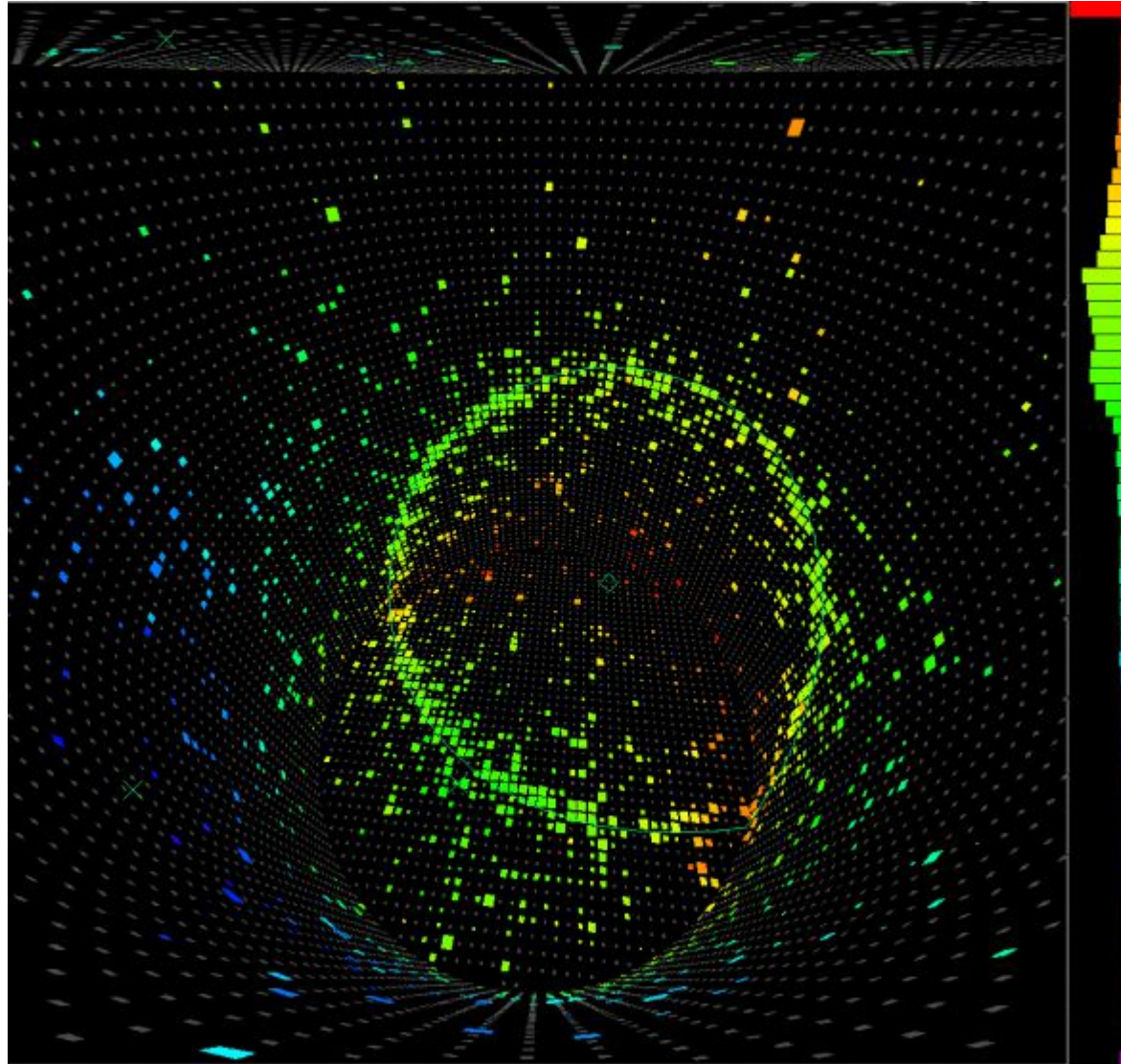




# Detectors(Liquid)

## Photo Multiplier: Cherenkov radiation

- Mini-Boon

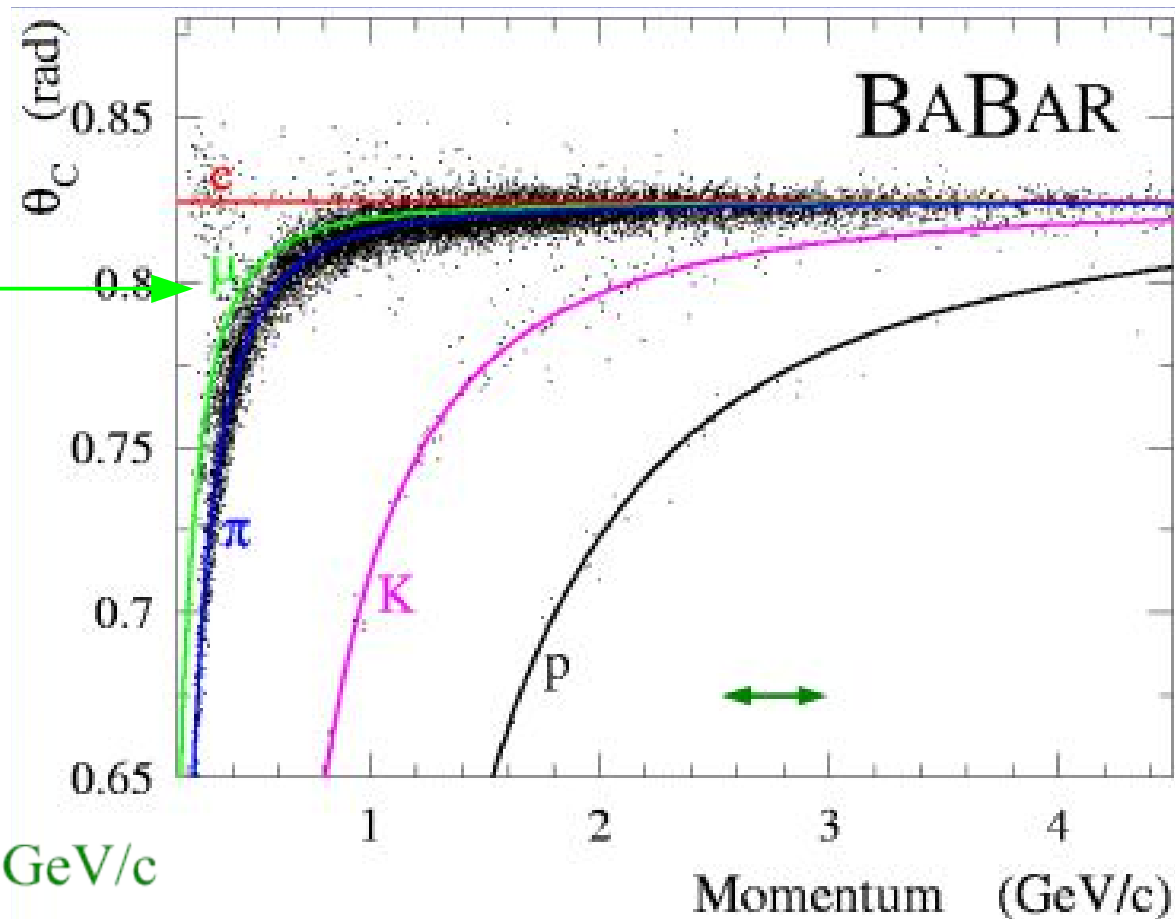


# Detectors(Liquid)

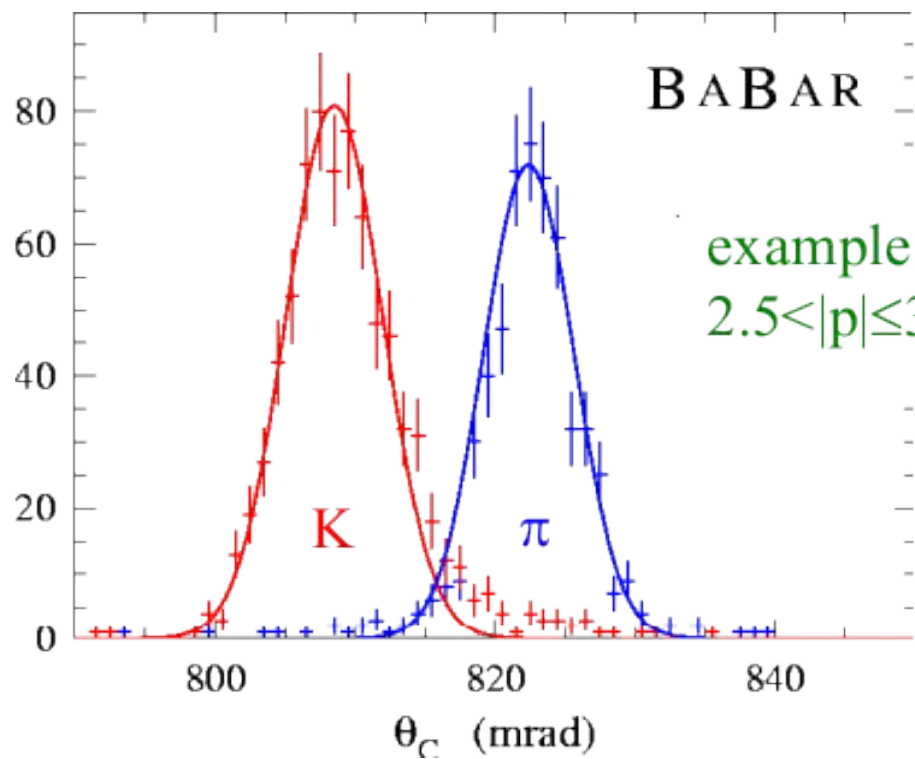
## Photo Multiplier: Cherenkov radiation

- Babar identification

Muon →



example:  
 $2.5 < |p| \leq 3 \text{ GeV/c}$

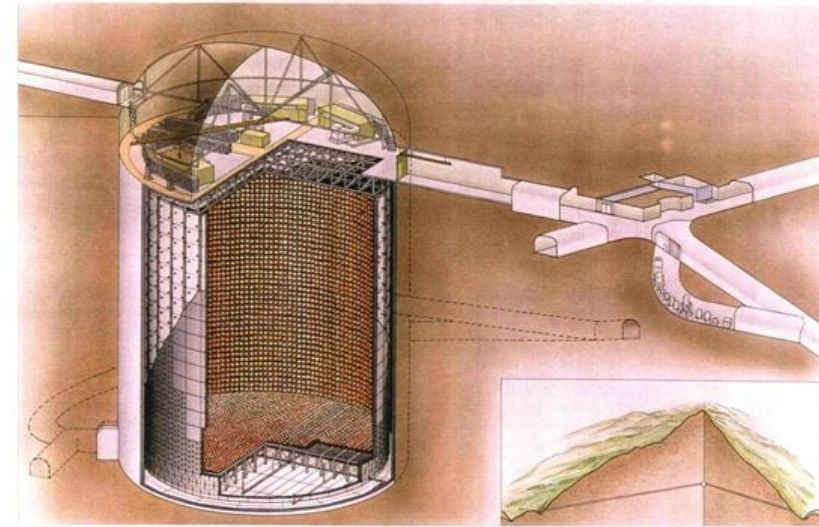
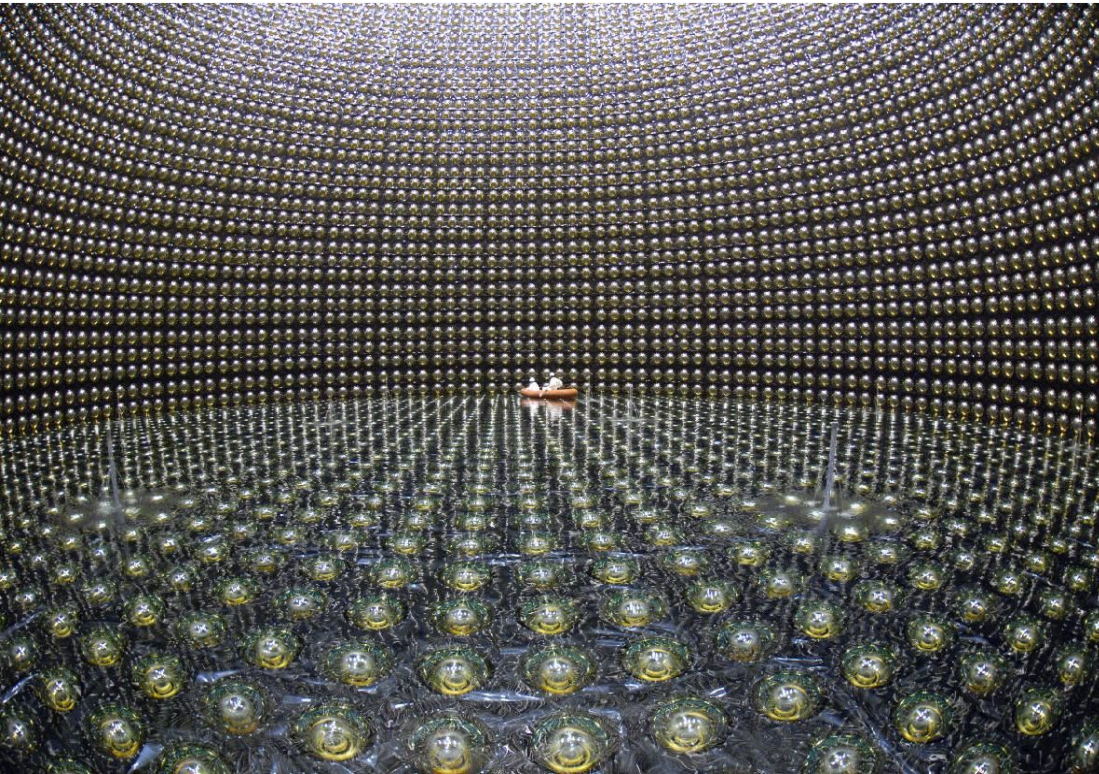




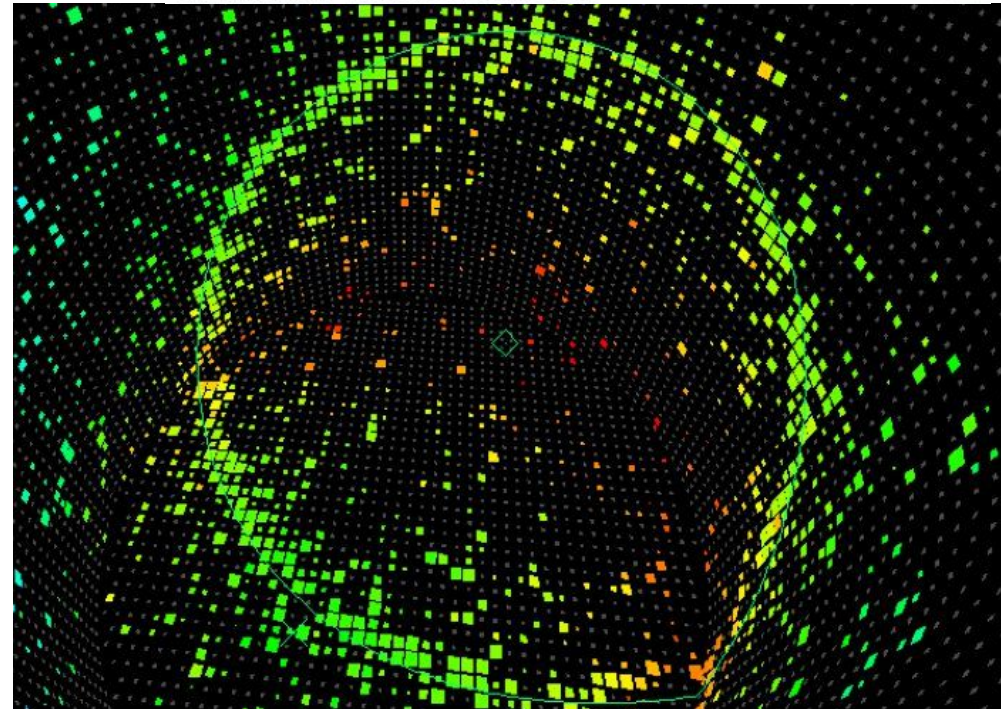
# Detectors(Liquid)

## Photo Multiplier: Cherenkov radiation

- T2K



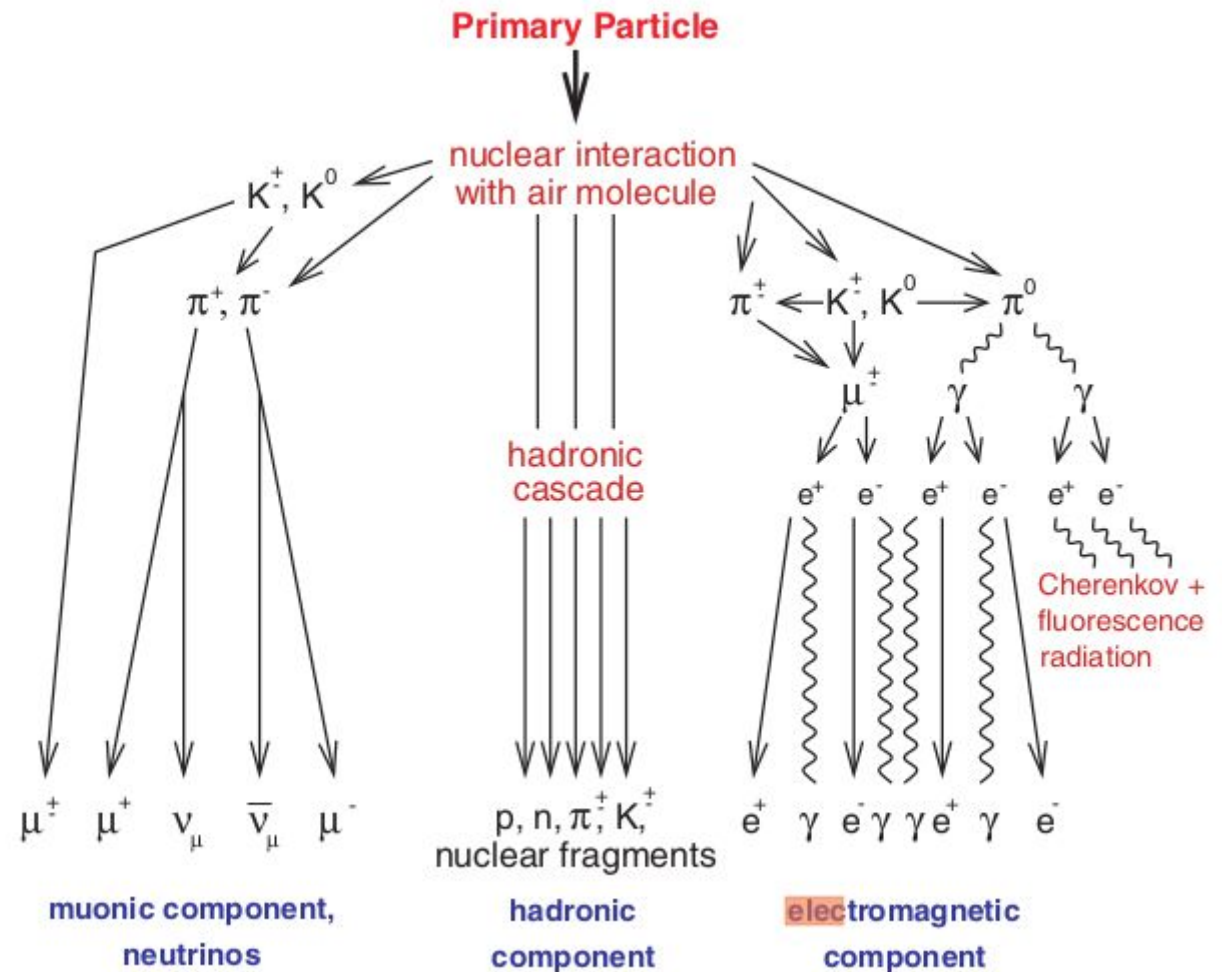
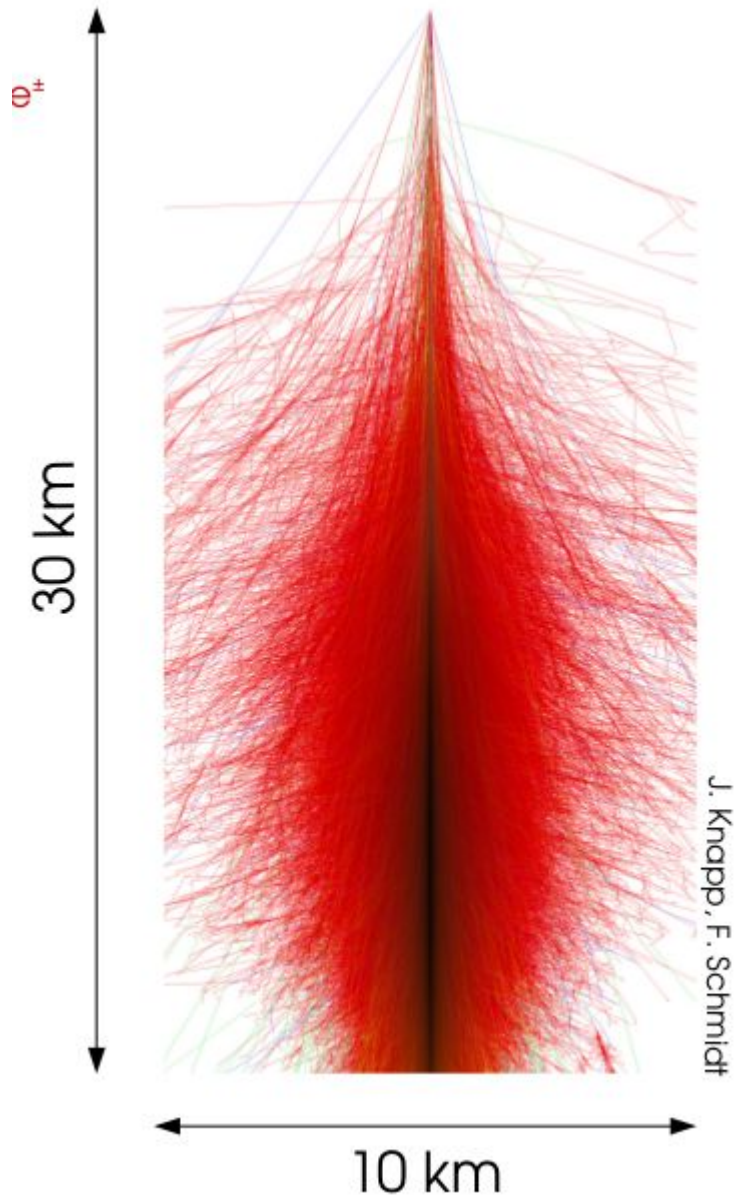
SUPERKAMIOKANDE INSTITUTE FOR COSMIC RAY RESEARCH UNIVERSITY OF TOKYO  
(c) Kamioka Observatory, ICRR(Institute for Cosmic Ray Research), The University of Tokyo





## Photo Multiplier: Cherenkov radiation

- Back to Cosmics



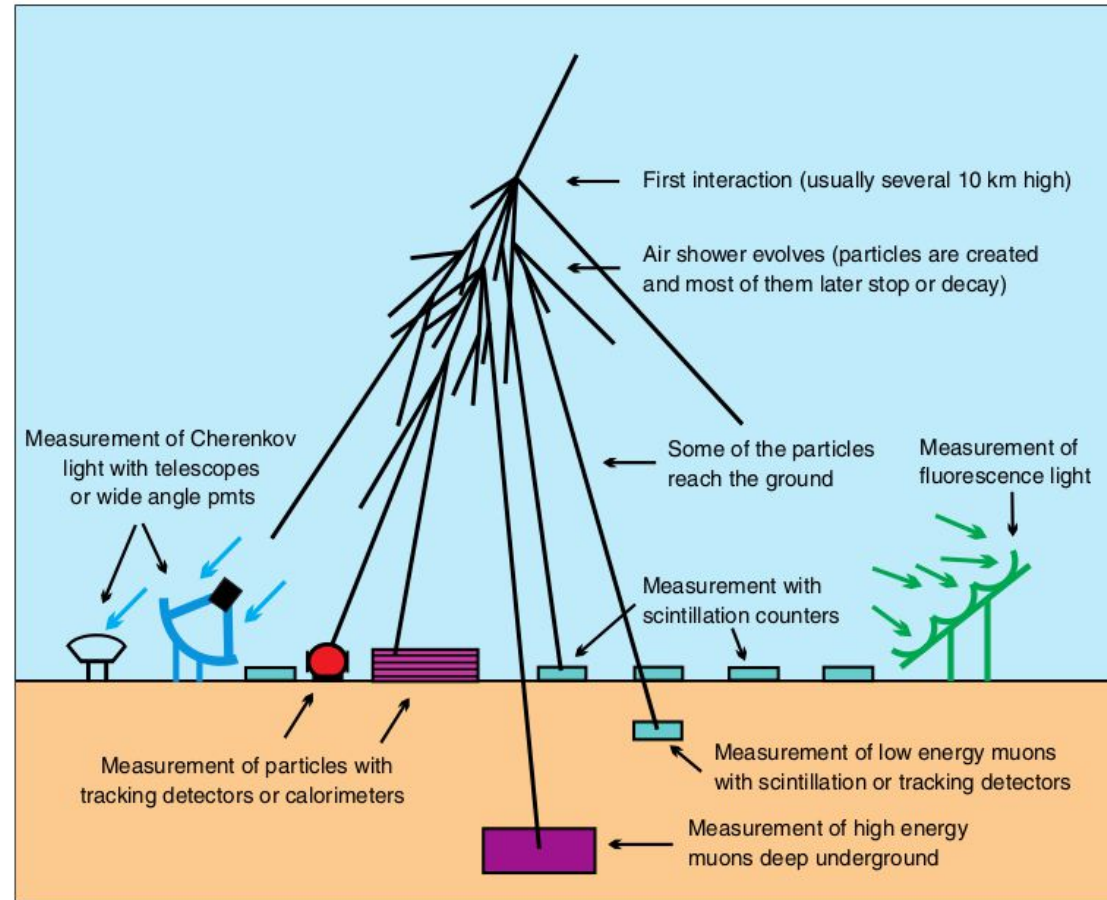
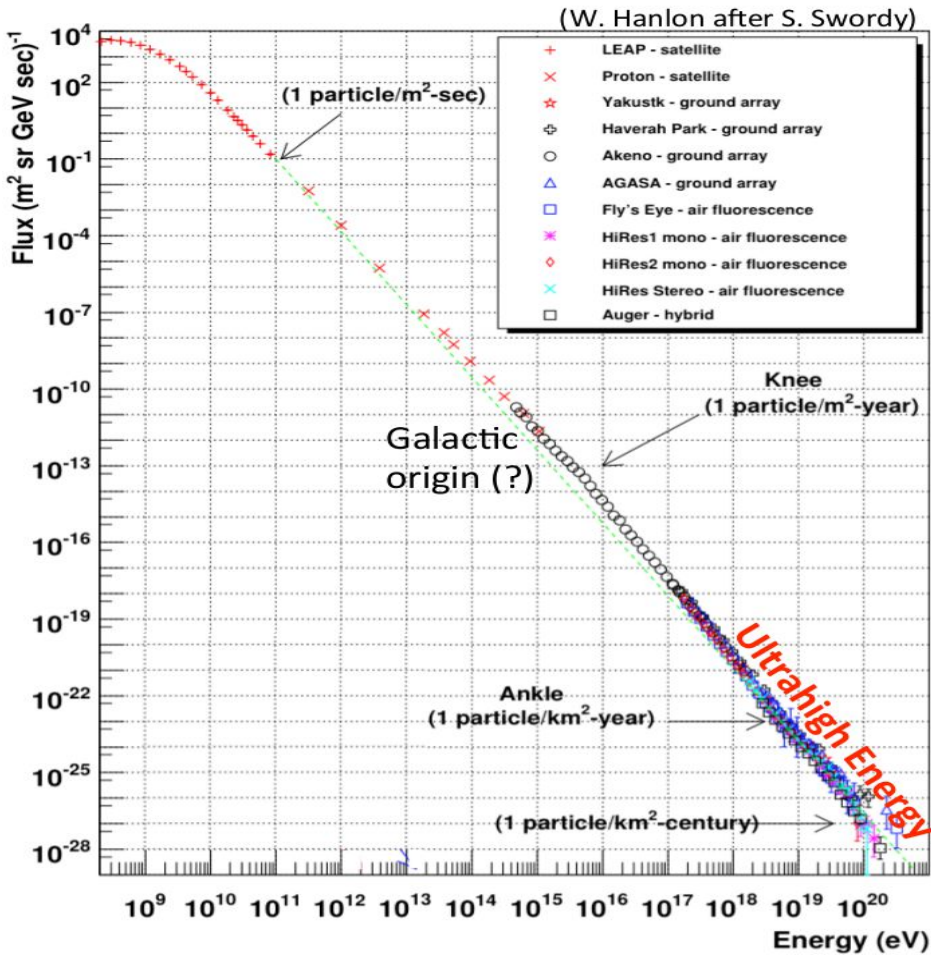


# Detectors(Liquid)

Francois Montanet  
Experimental Astroparticle Physics

## Photo Multiplier: Cherenkov radiation

- Back to Cosmics



100K light years  
Milky Way Galaxy

5M light years  
Local Galactic Group

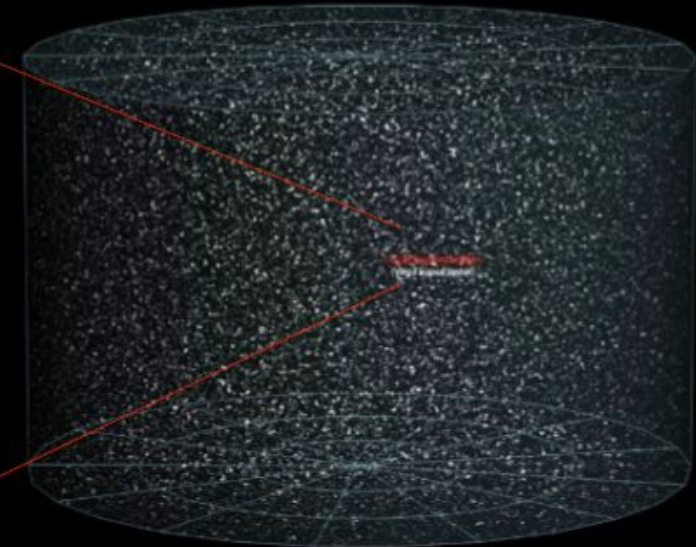
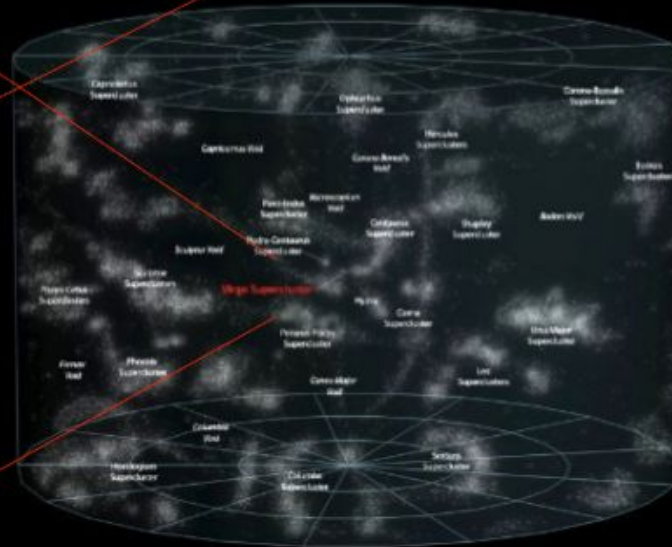
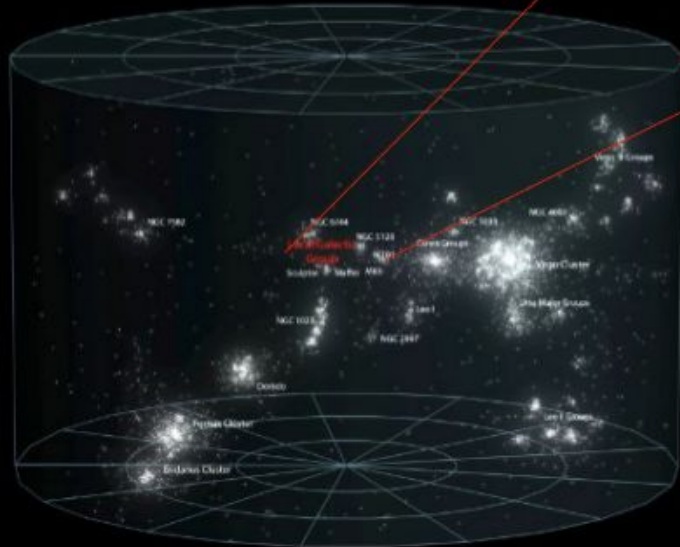
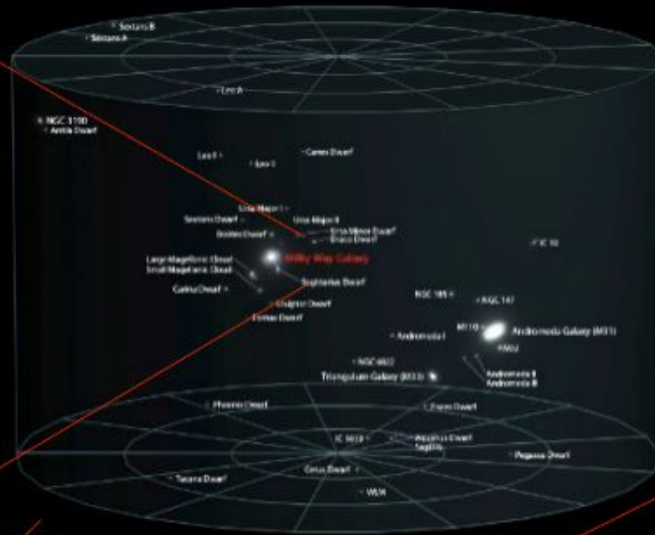
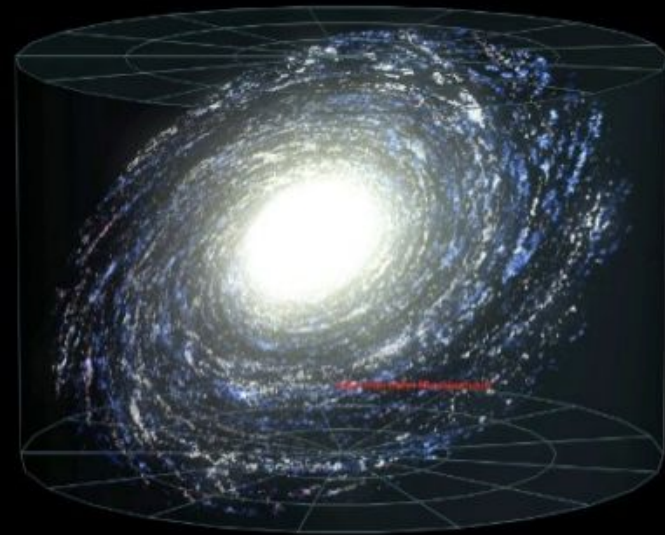
Universe becomes  
opaque for high energy  
Photons:

$$\gamma + \gamma_{\text{background}} \rightarrow e^+ + e^-$$

100M light years  
Virgo Supercluster

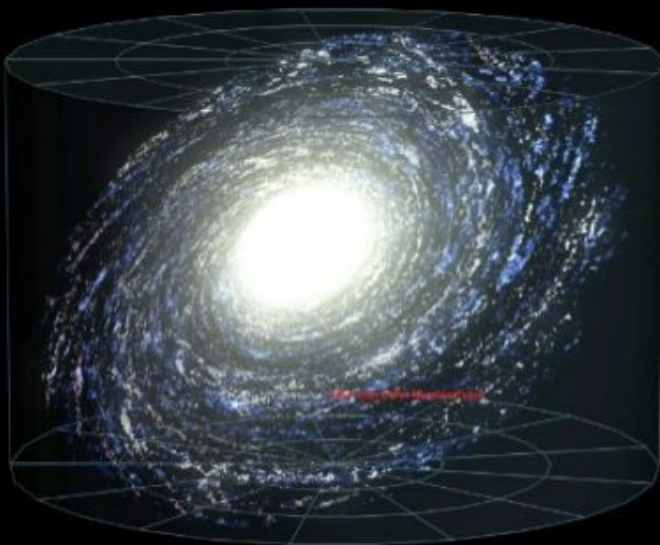
1G light years  
Local Superclusters

Observable Universe

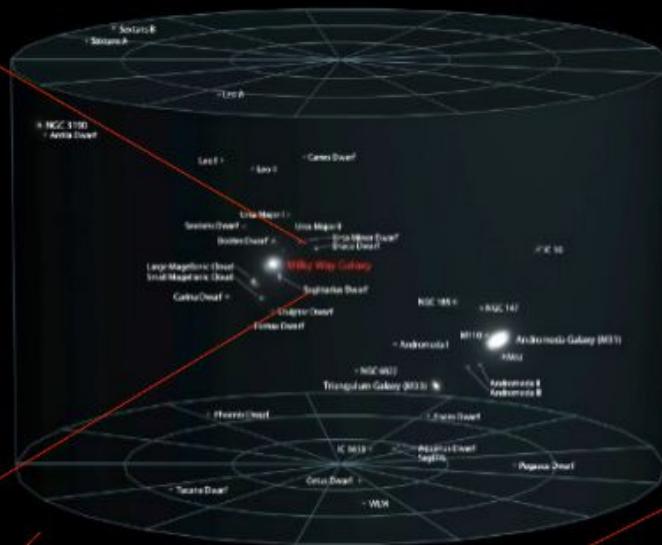




100K light years  
Milky Way Galaxy



5M light years  
Local Galactic Group

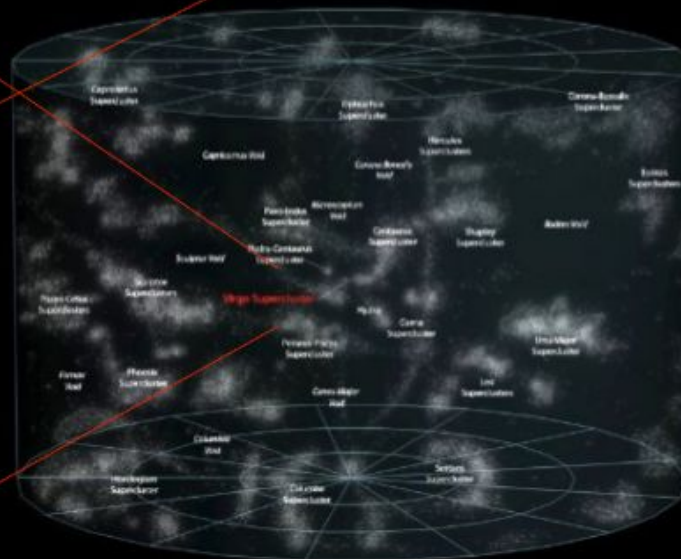


Universe is transparent to neutrinos at all energies

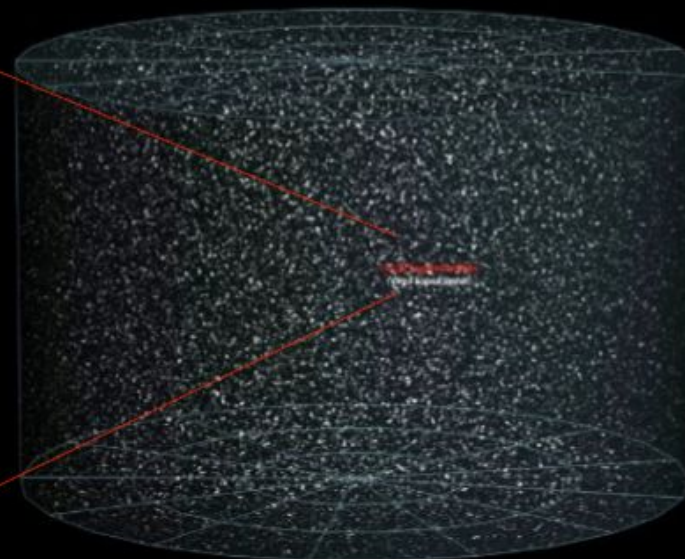
100M light years  
Virgo Supercluster



1G light years  
Local Superclusters



Observable Universe



# Detectors(Liquid)

## Photo Multiplier: Cherenkov radiation

- The muon is detected via Cherenkov radiation in the water or ice

### IceCube Neutrino Observatory

86 strings

60 Optical Modules per string

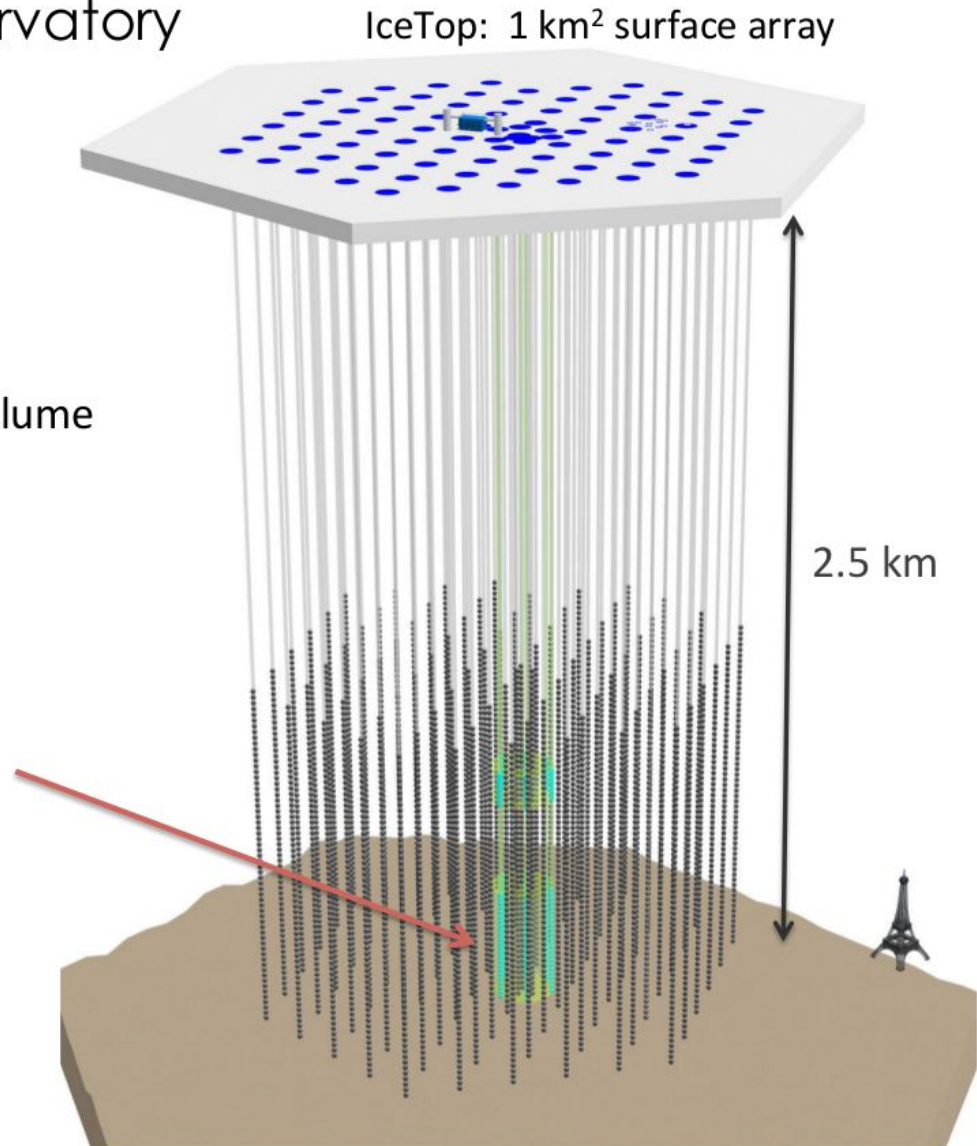
5 160 total modules in Ice

1 km<sup>3</sup> = Gigaton instrumented volume

Began full operations May 2011

**DeepCore**  
Low-energy Extension

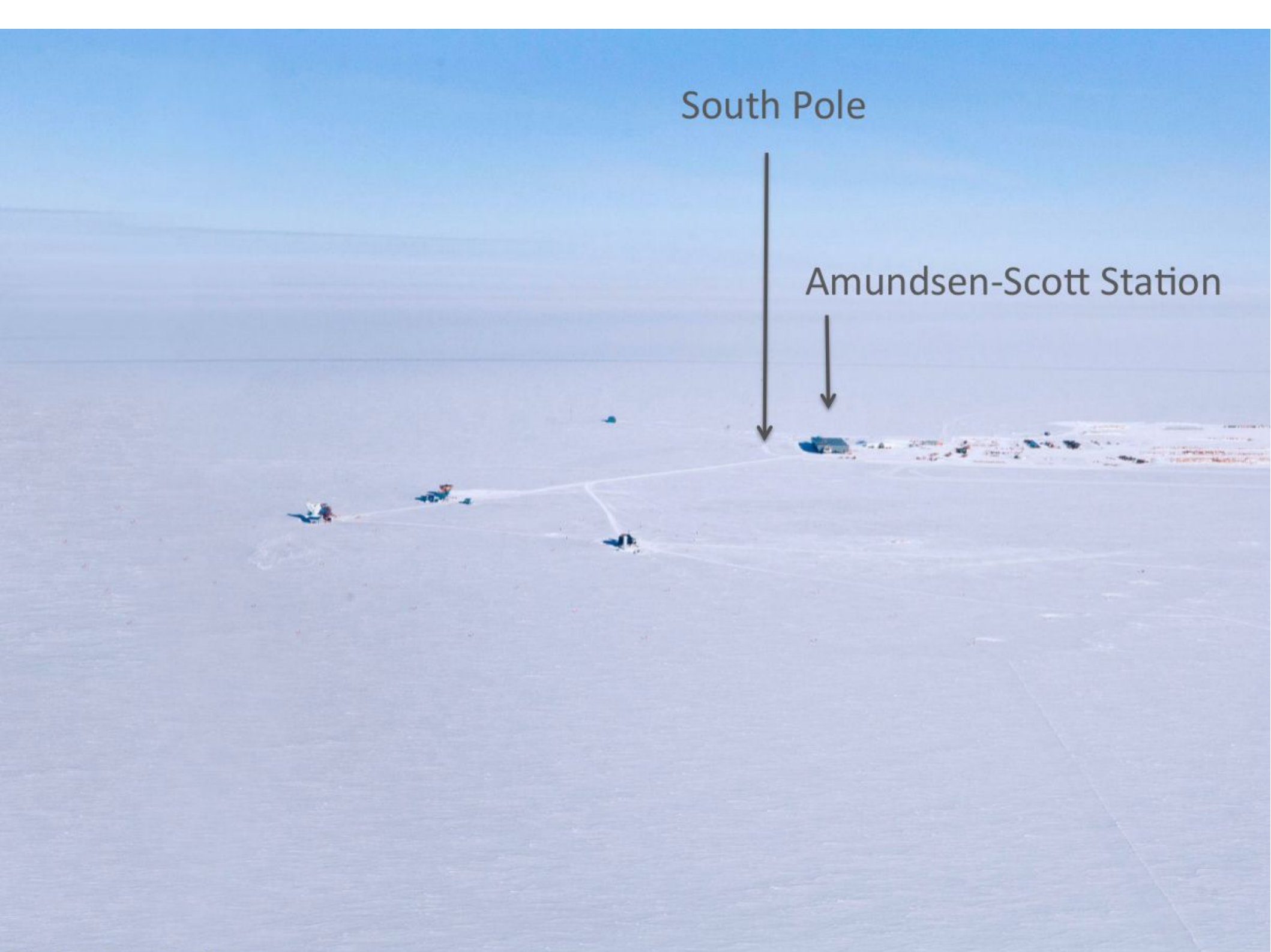
*Dark Matter,  
Neutrino Oscillations*





South Pole

Amundsen-Scott Station



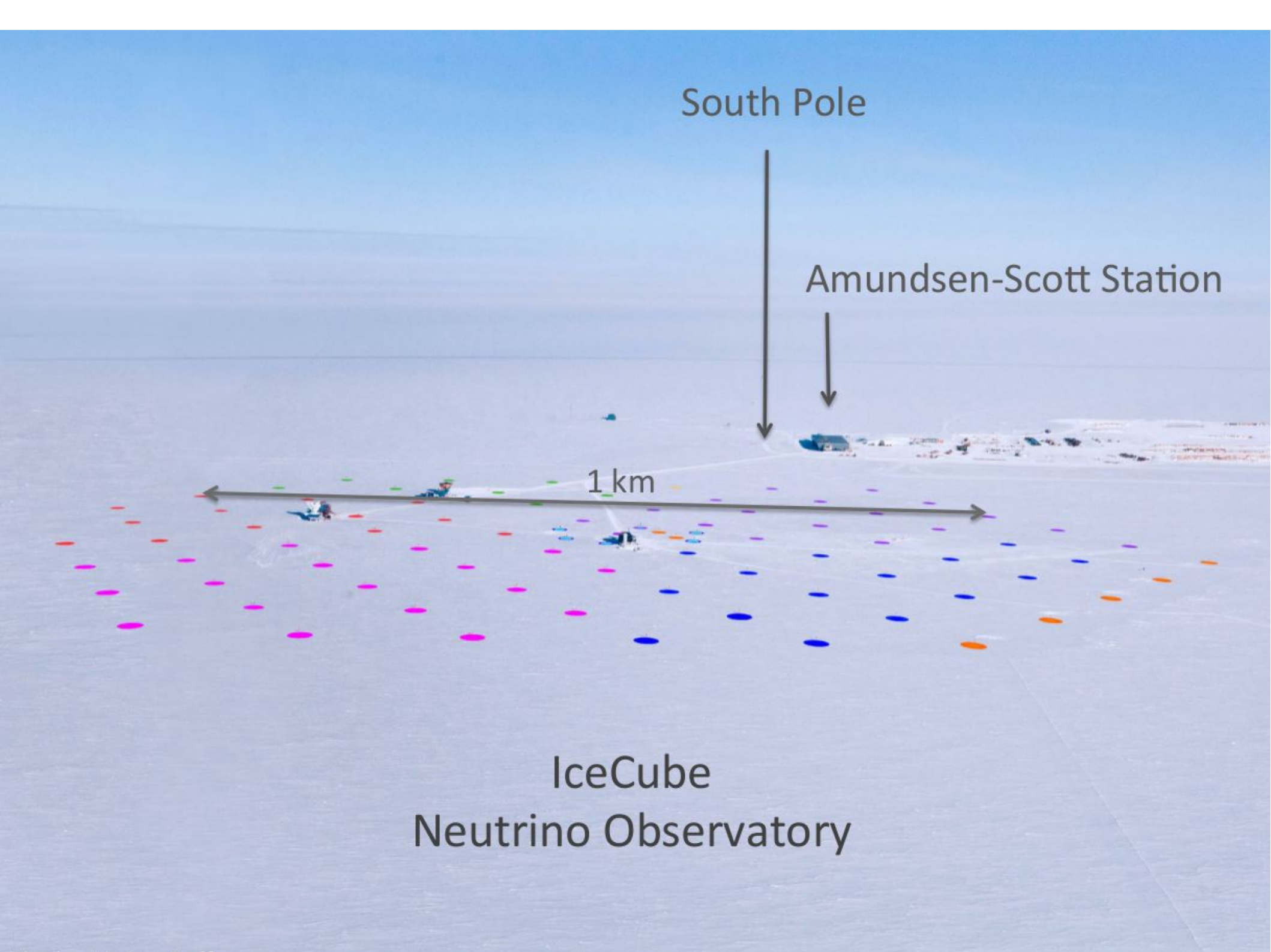


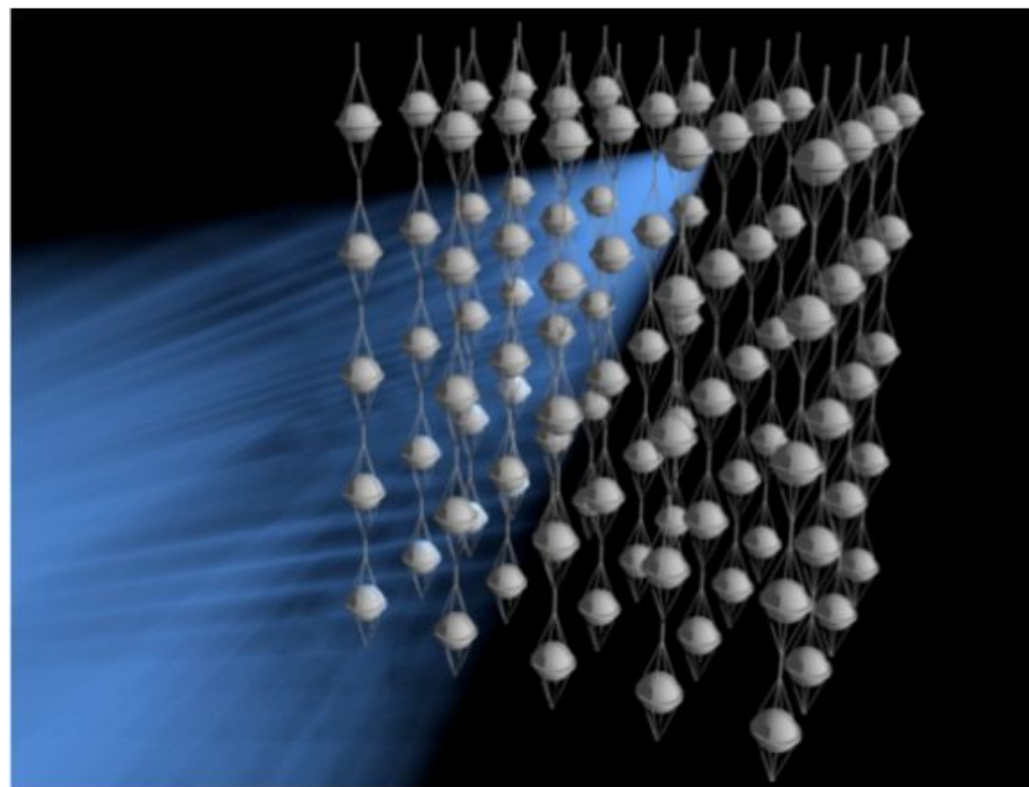
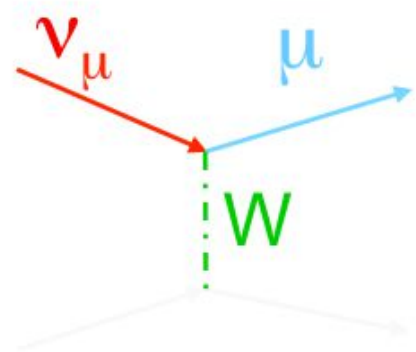
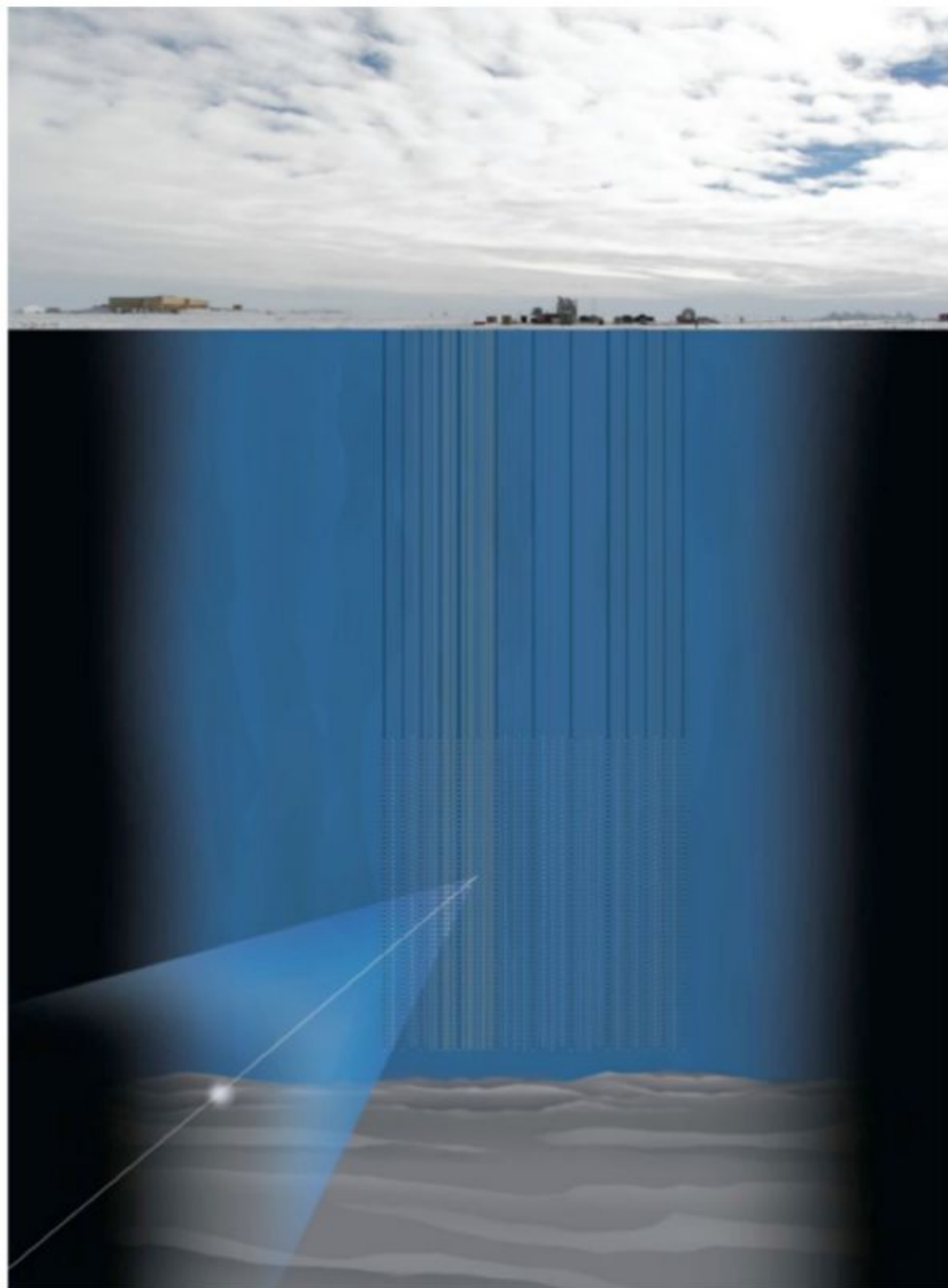
South Pole

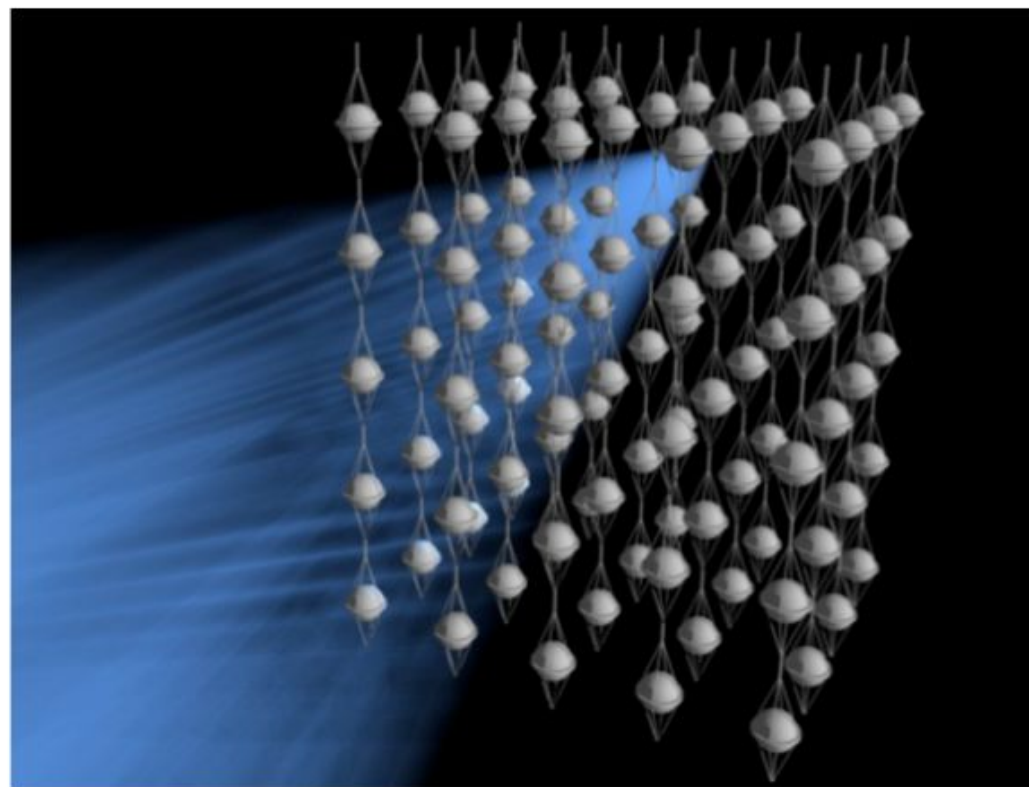
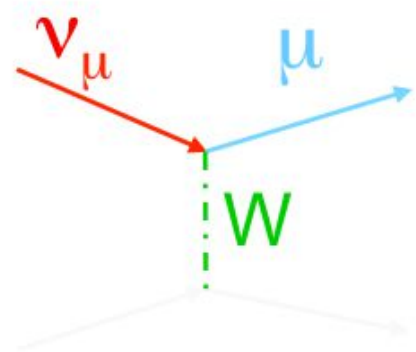
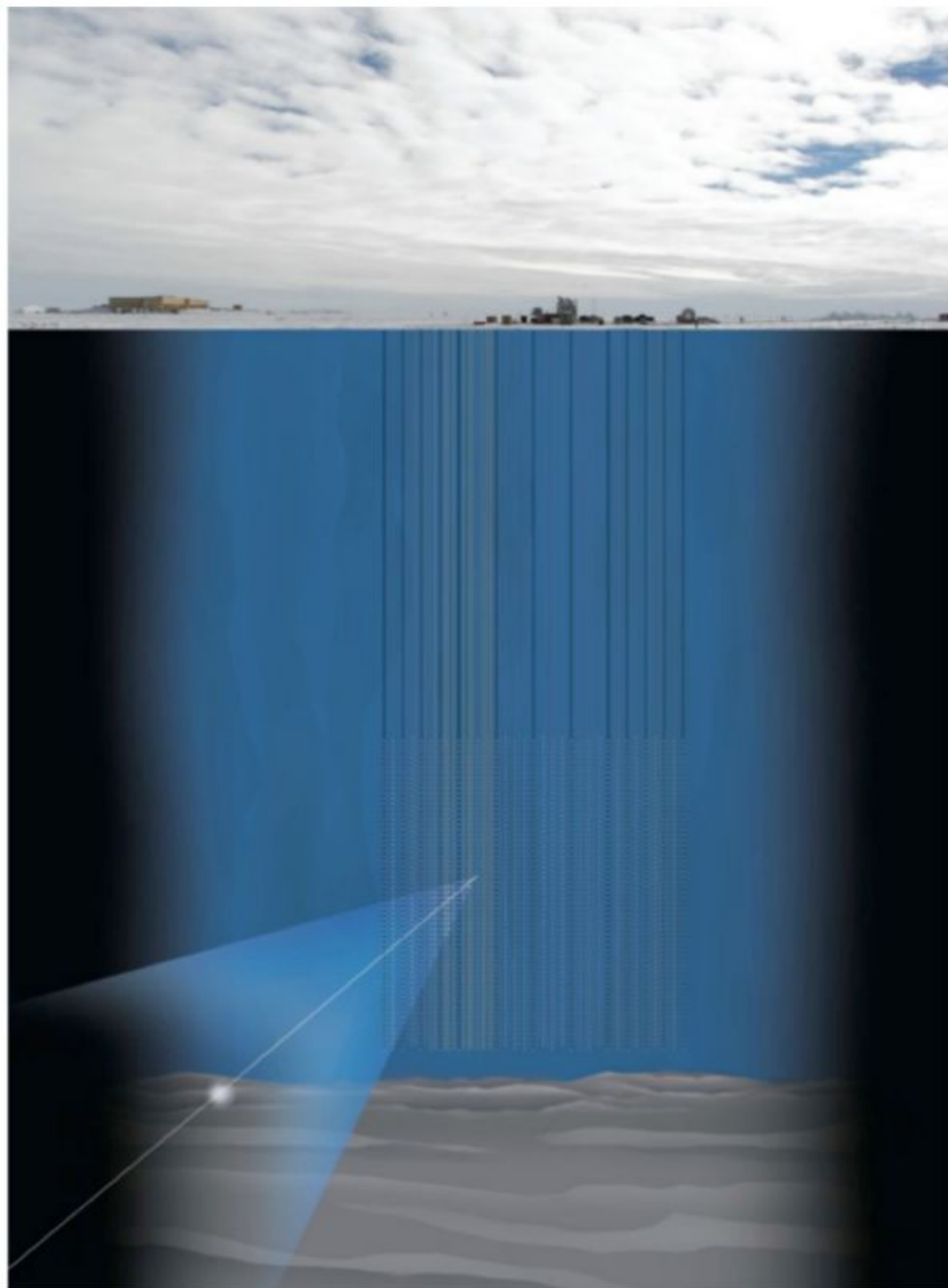
Amundsen-Scott Station

1 km

IceCube  
Neutrino Observatory

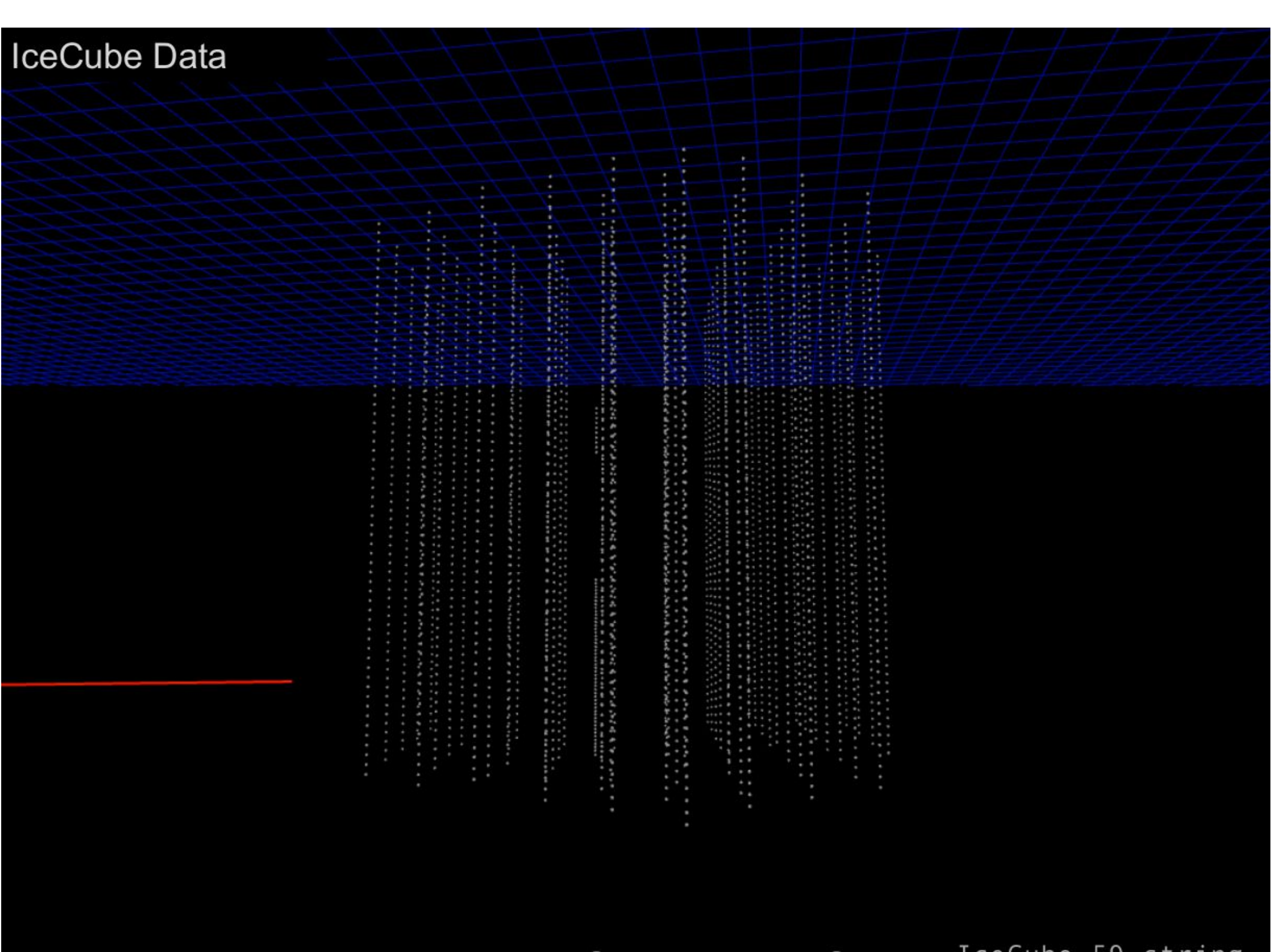




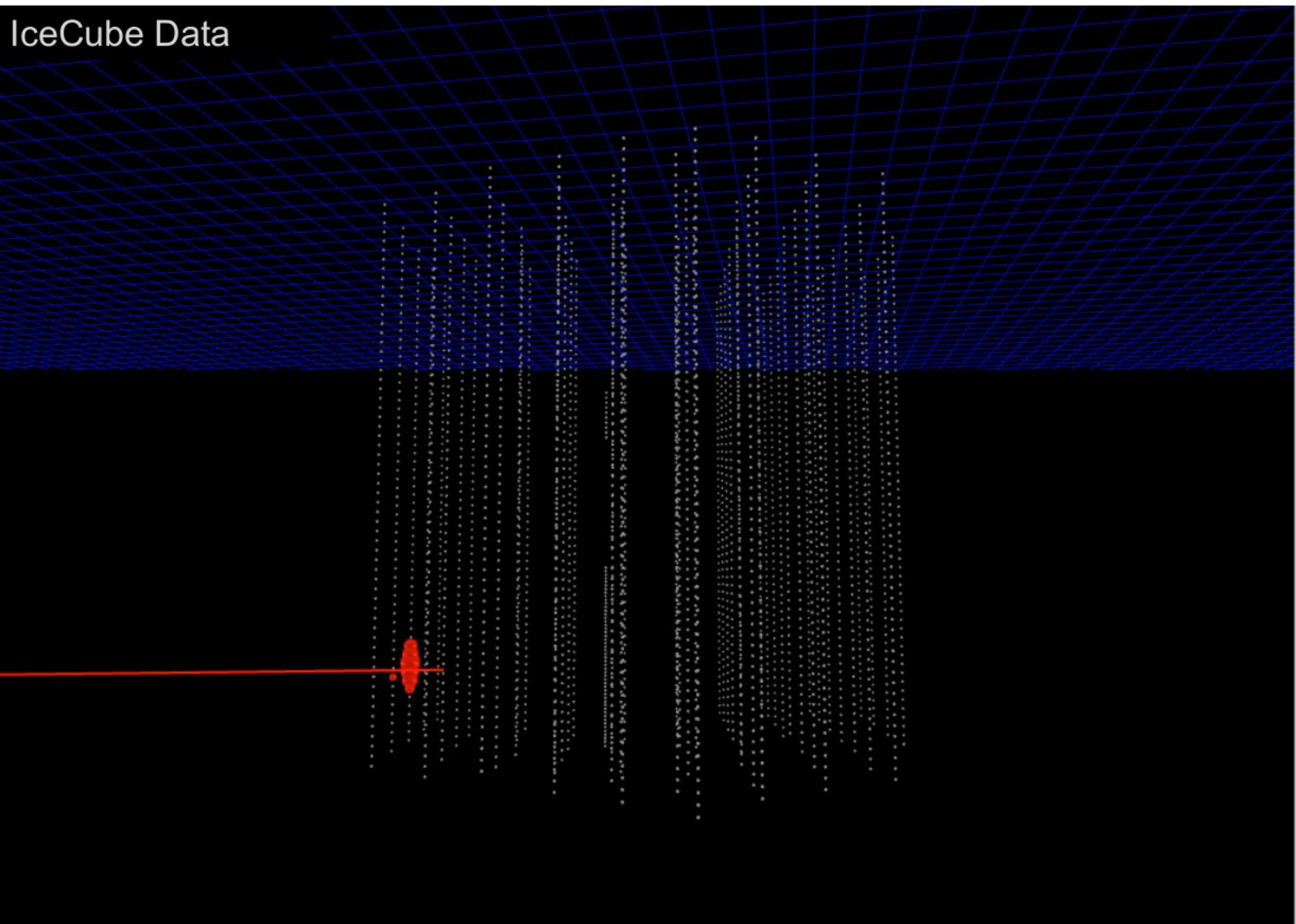




# IceCube Data

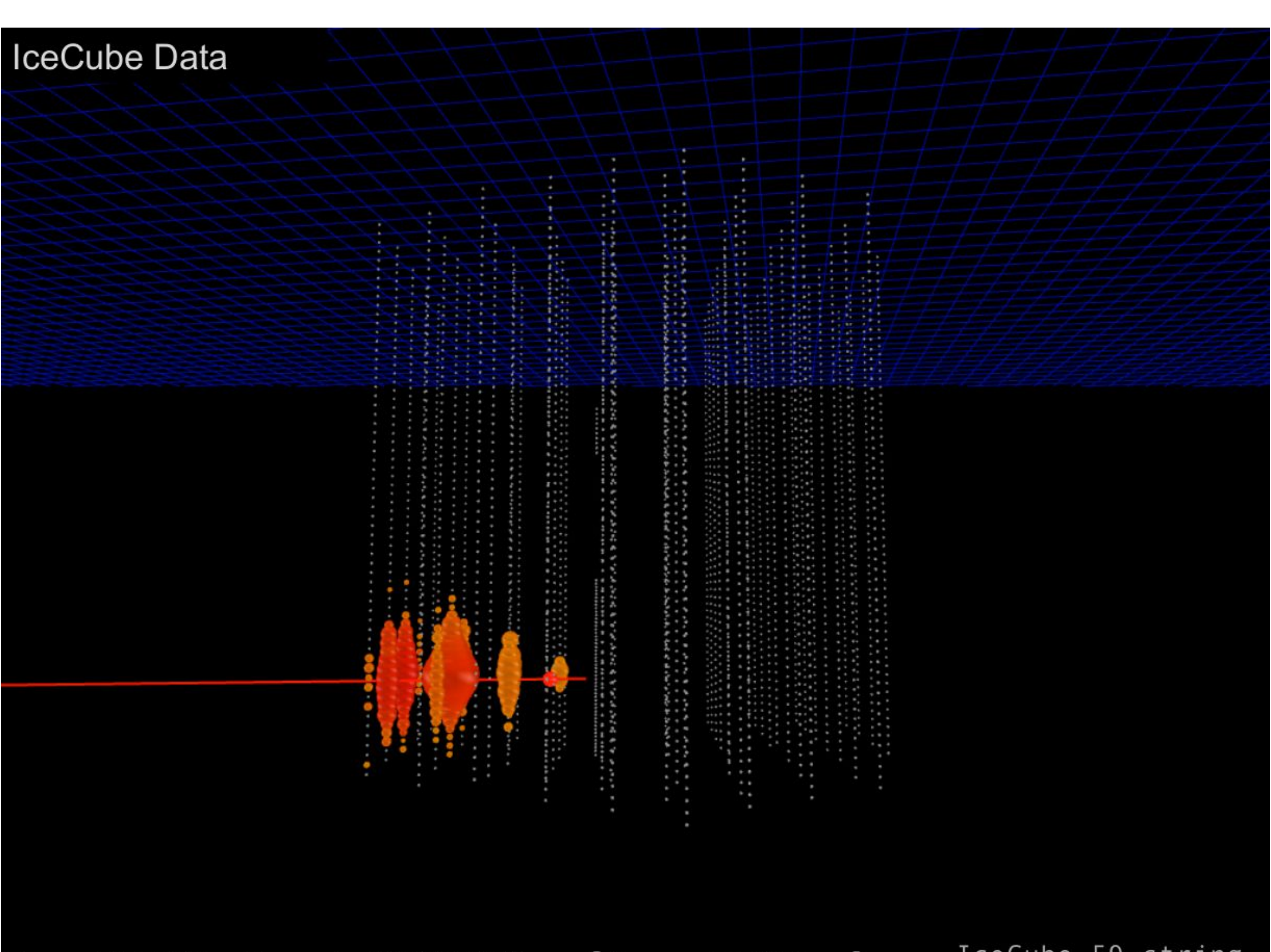


# IceCube Data

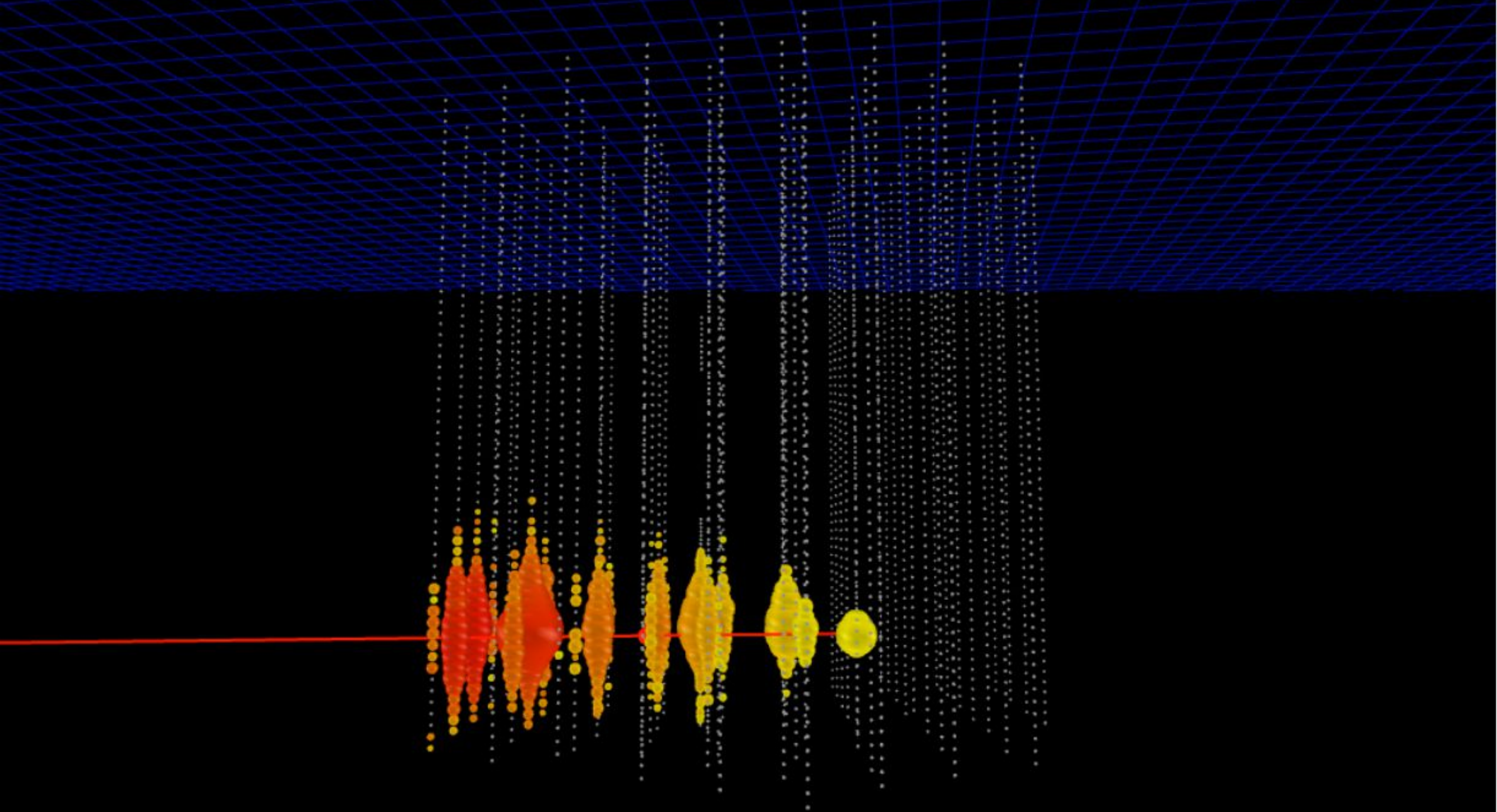




# IceCube Data

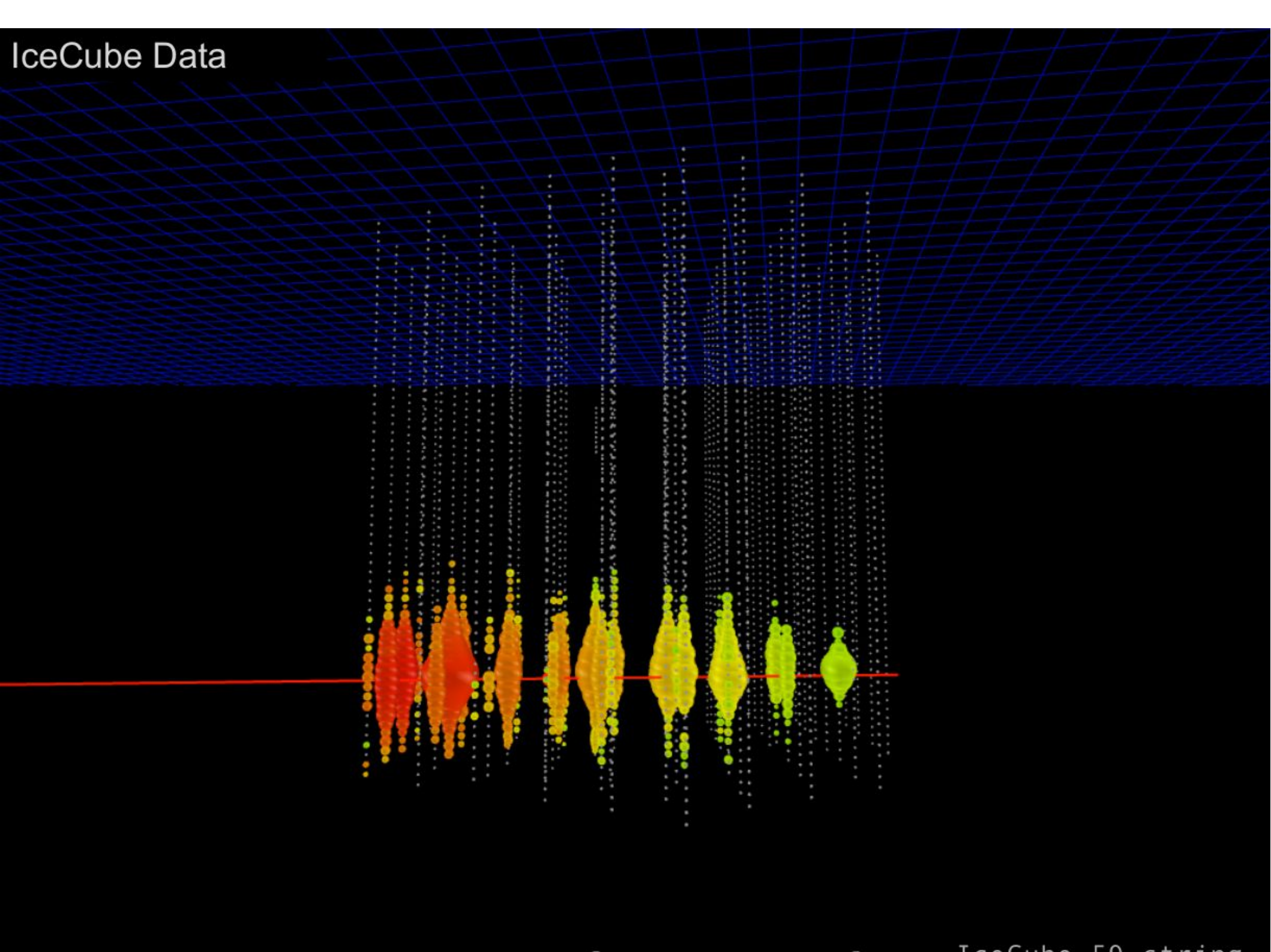


# IceCube Data



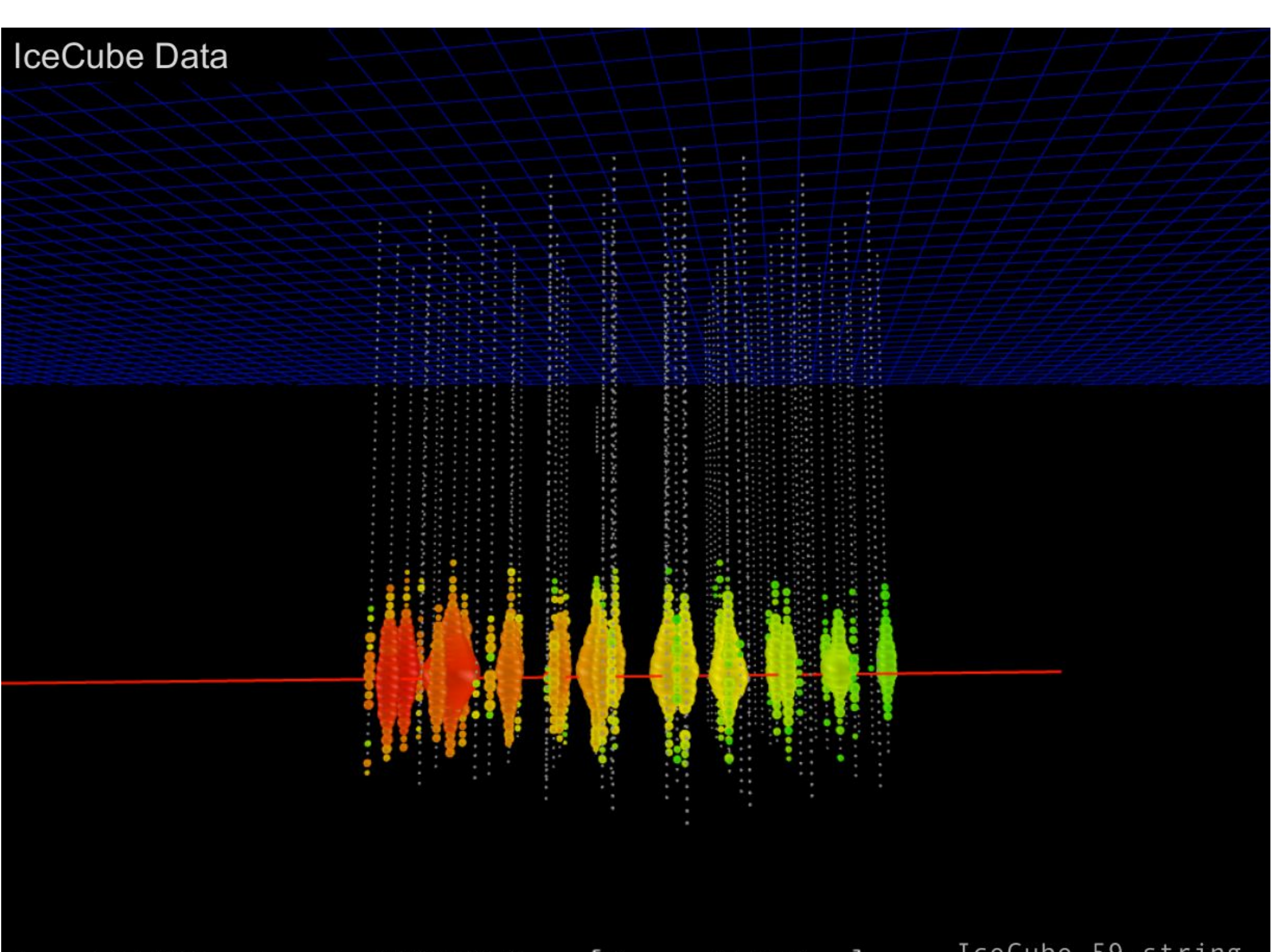


IceCube Data



IceCube\_50\_string

IceCube Data



IceCube\_50\_string



# Muon

## Beam

- In accelerators muons are abundantly produced in hadronic interactions
  - $pp \rightarrow \pi + \dots$  and  $\pi \rightarrow \mu \nu_\mu$
- Today muon beams are available at many places in Europe, Asia, and America.
- High energy muon beams, e.g., at CERN SPS, FNAL
- Low/medium energy: PSI, TRIUMF, Los Alamos, BNL, DUBNA, RAL, ...



# Muon

## Muon cooling for a Higgs Factory at CERN ?

### New boson sparks call for 'Higgs factory'

Jul 5, 2012 [15 comments](#)



Former CERN boss Carlo Rubbia wants a muon collider

CERN's discovery of a new fundamental particle – most likely a Higgs boson – was barely hours old when physicists speaking at this year's Lindau Nobel Laureate Meeting in Germany argued the case for a new facility to measure its properties in detail. Speaking out in favour of a new machine was former CERN boss Carlo Rubbia, who shared the 1984 Nobel Prize for Physics for the discovery of the W and Z bosons. "The technology is there to construct a Higgs factory," he claimed. "You don't need €10bn; it could be done relatively cheaply." *Saclay, feb. 2015*

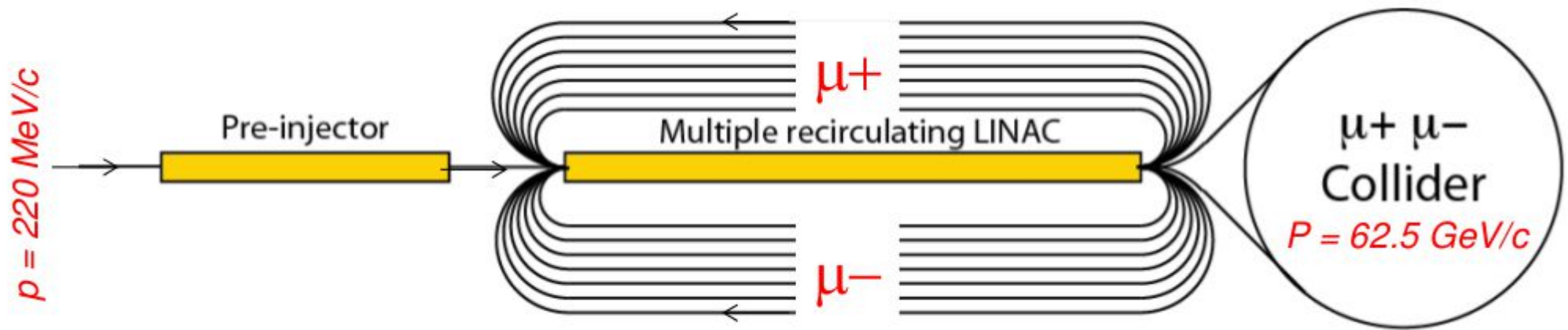
physicsworld.com  
BEST SPECIALIST SITE FOR JOURNALISM 2011

"With a Higgs of 125 GeV we need only a modest machine, perhaps not a large linear collider." Rubbia points out that muons colliding at a combined energy of roughly 125 GeV would suffice – just over half the energy of LEP and requiring a machine with a much smaller radius.



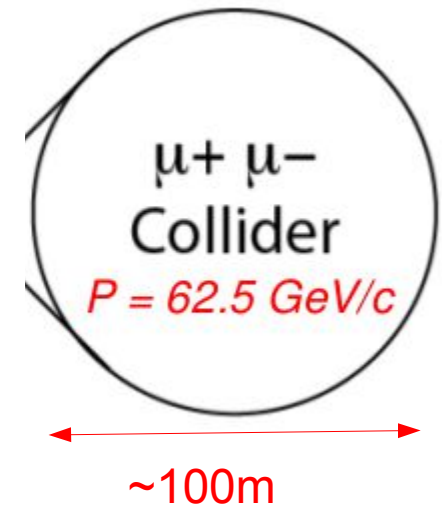
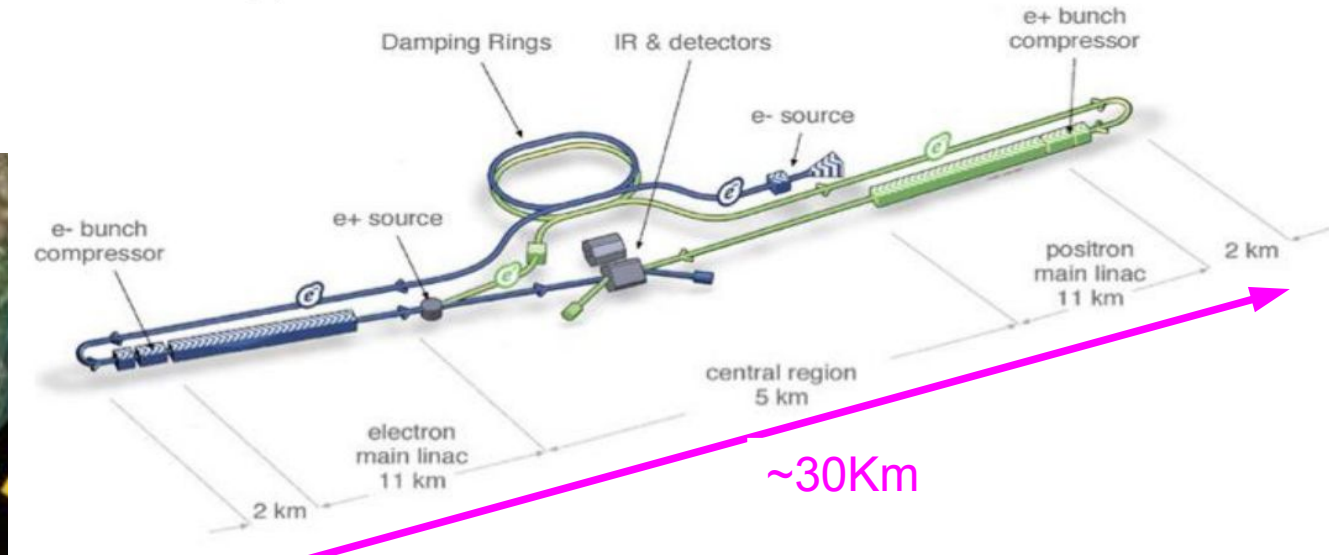
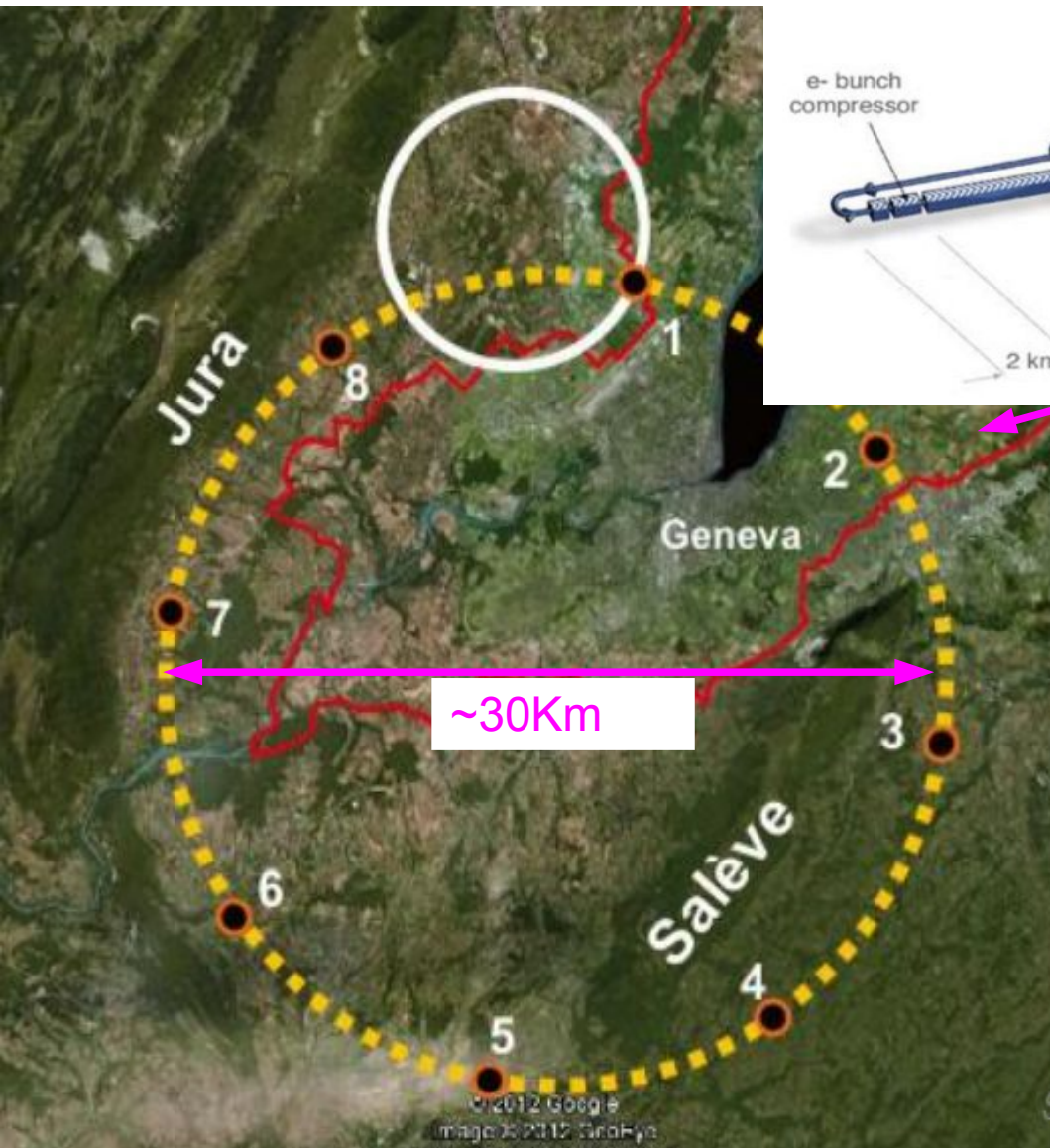
# Muon

## Muon cooling for a Higgs Factory at CERN ?



# Muon

## Muon cooling for a Higgs Factory at CERN ?





# Summary

- Muon is an elementary particle describe by the SM
  - All its parameters are well known
  - Some tension on  $g_{\mu}-2$  (3.6 sigma)
  - Muons is used in many domain: Astrophysics, particle physics
    - Atmospheric showers, Trigger, Veto,...
- Muon detection
  - started with cloud chambers and Geiger-Müller
  - Detection mechanism always the same: **Ionisation**
  - The main break-through in tracking detectors: MWPC
  - Spark chambers parallel-plate chambers has lead to RPCs and now MPGDs\*
  - MPDGs are probably the new generation of muon detectors being robust
  - GEM, MicroMegas, THGEM...
    - Radiation hard and showing no signs of ageing
    - High rates
    - Excellent spatial resolution
    - Fast (trigger)
- Muo-graphy
- Future: Muon collider?