

Warped Relaxion



Nayara Fonseca

DESY

Based on joint work with:

- Lima, Machado, Matheus (Phys. Rev. D 94, 015010 (2016));
- Harling, Lima, Machado, Servant (working in progress).

Blois, May 31st, 2017

Outline

1. The Relaxion Idea

2. How to generate large-scale hierarchies?

- Using N -site models : NF, Lima, Machado, Matheus; Phys.Rev. D94 (2016) 015010
- From Warped Space: NF, Harling, Lima, Machado, Servant; working in progress.

3. Concluding Remarks

The Relaxion Idea

Graham, Kaplan, Rajendran; Phys. Rev. Lett. 115, 221801 (2015)

**Before Relaxion:
Two different “solutions” to the SM hierarchy problem**

1. New dynamics at the weak scale:

- **Natural solutions;**
- **We need BSM at \sim TeV scale;**
Eg.: SUSY & Composite Higgs Models.

2. Anthropics !?

The Relaxion Idea

3. Another option: The Relaxation Mechanism

Graham, Kaplan, Rajendran; Phys. Rev. Lett. 115, 221801 (2015)

Warming up...

$$V(h, \phi) = \frac{1}{2} m_H^2(\phi) h^2 + \dots = \frac{1}{2} (-\Lambda^2 + g\Lambda\phi) h^2 + \dots$$

The Relaxion Idea

3. Another option: The Relaxation Mechanism

Graham, Kaplan, Rajendran; Phys. Rev. Lett. 115, 221801 (2015)

Warming up...

high scale

the new field

$$V(h, \phi) = \frac{1}{2} m_H^2(\phi) h^2 + \dots = \frac{1}{2} (-\Lambda^2 + g\Lambda\phi) h^2 + \dots$$

small coupling (spurion)

- ϕ scans $m_H^2(\phi)$ during the cosmological evolution;
- Arrange a mechanism so that ϕ stops where we want, precisely at the EW scale:

$$m_H^2(\phi_c) = -\Lambda^2 + g\Lambda\phi_c \ll \Lambda^2$$

The Relaxion Idea

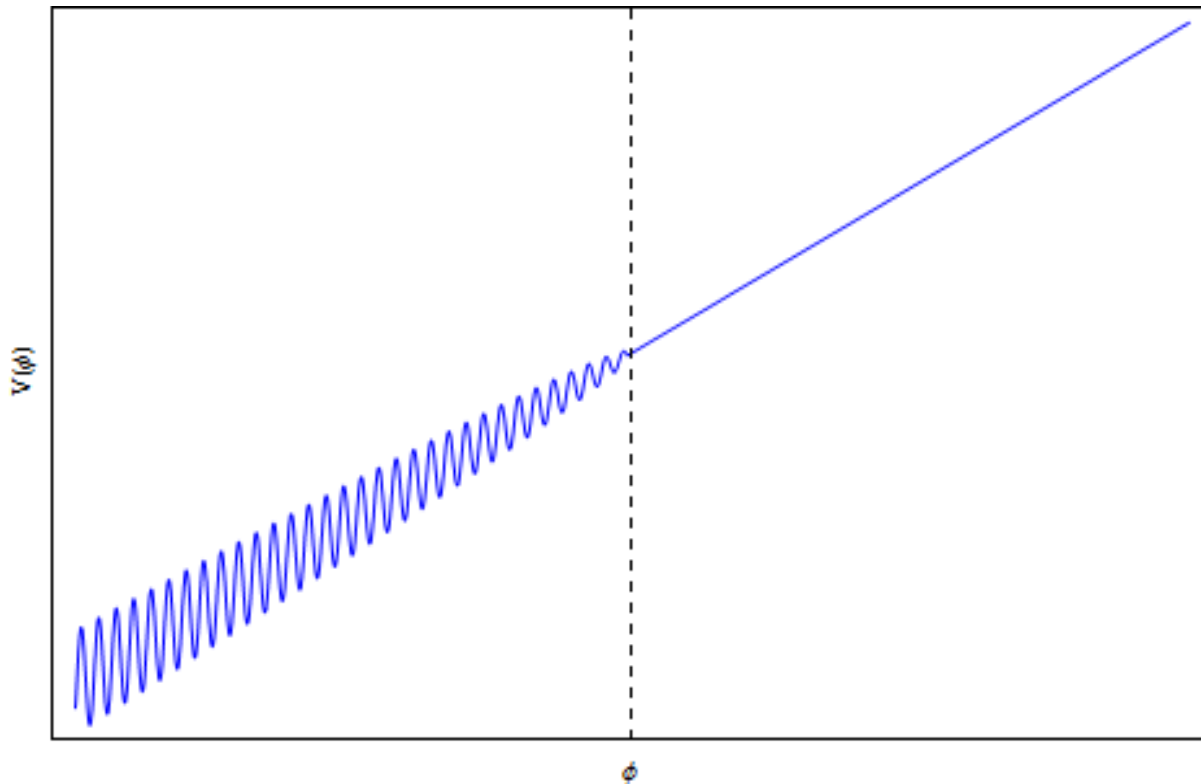
Closer Look

$$m_H^2(\phi)$$

"Slope term"

"Stopping term"

$$V(h, \phi) = \frac{1}{2} \Lambda^2 \left(\frac{g\phi}{\Lambda} - 1 \right) h^2 + g\Lambda^3 \phi + \epsilon \Lambda_c^4 \left(\frac{\langle h \rangle}{\Lambda_c} \right)^n \cos \left(\frac{\phi}{f} \right) + \dots$$



The Relaxion Idea

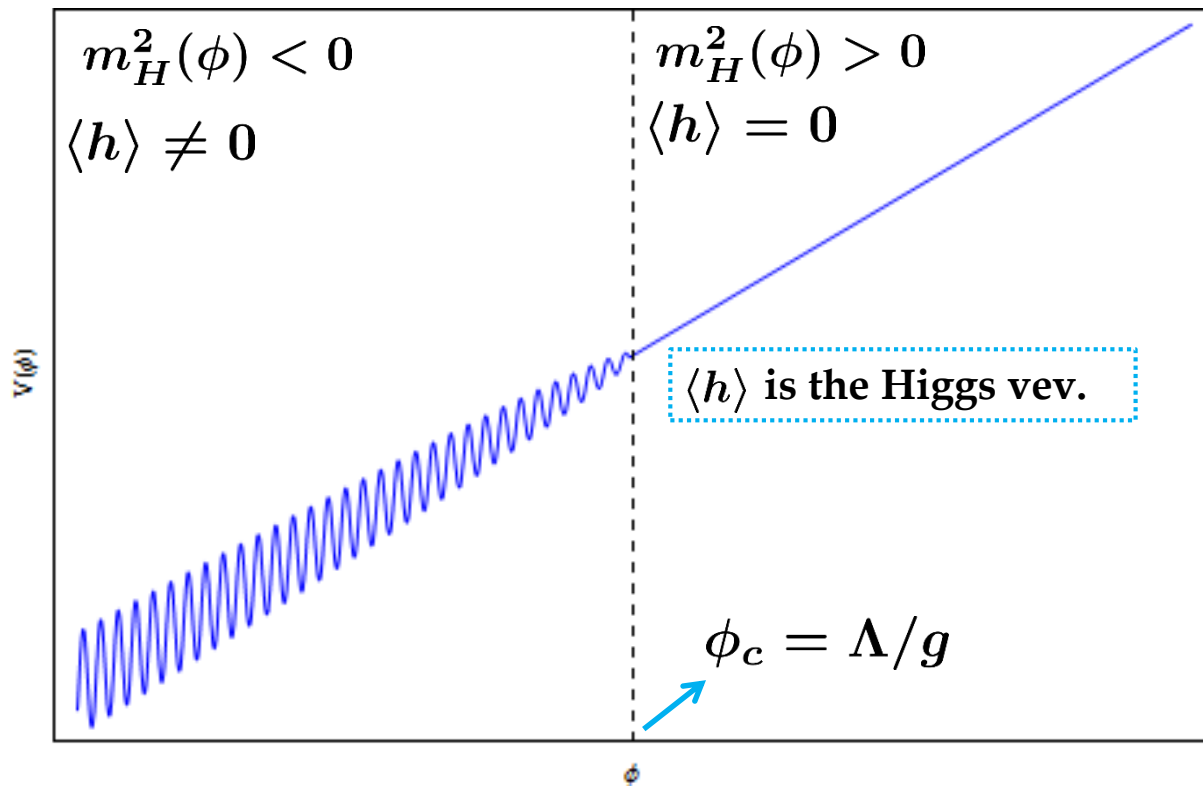
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The Relaxion Idea

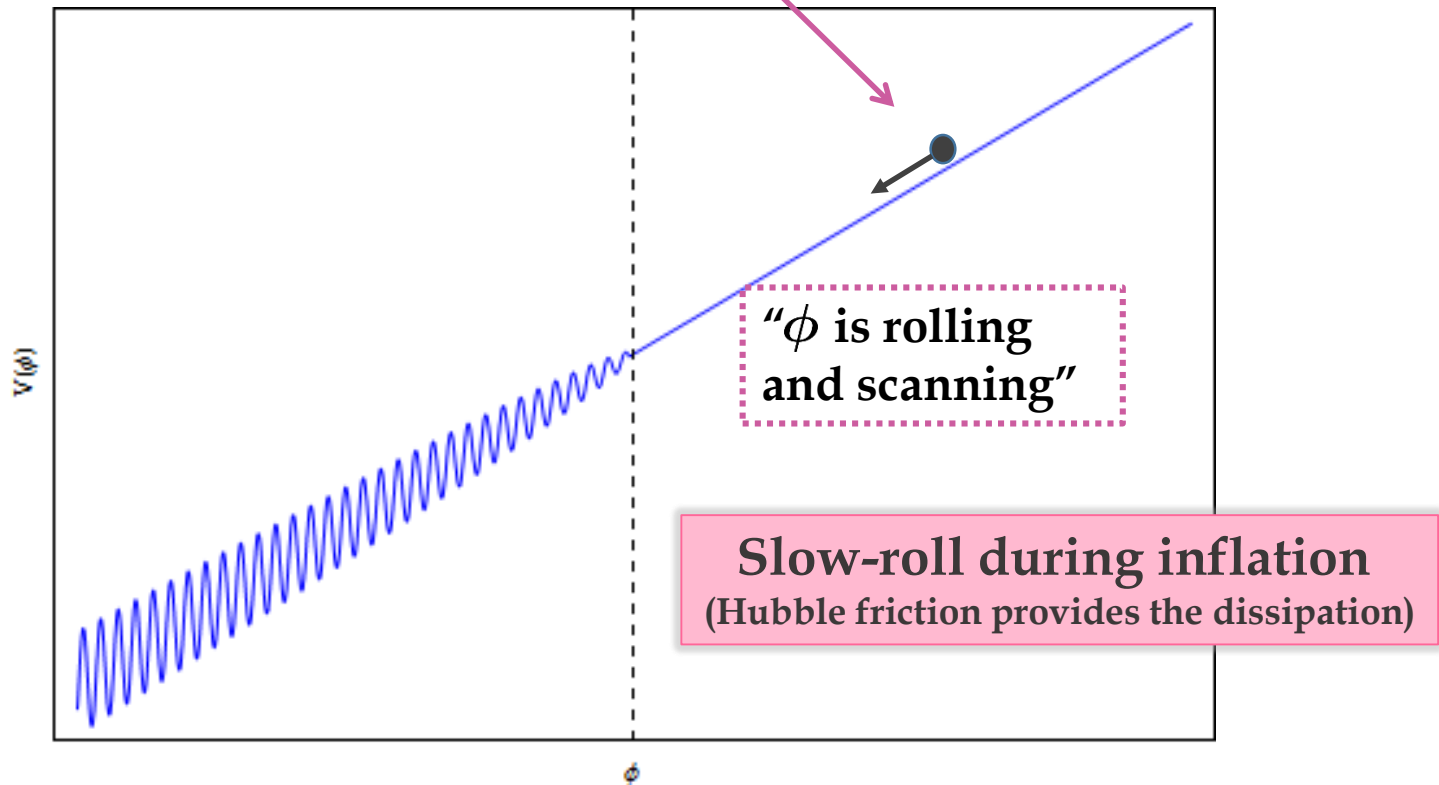
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The Relaxion Idea

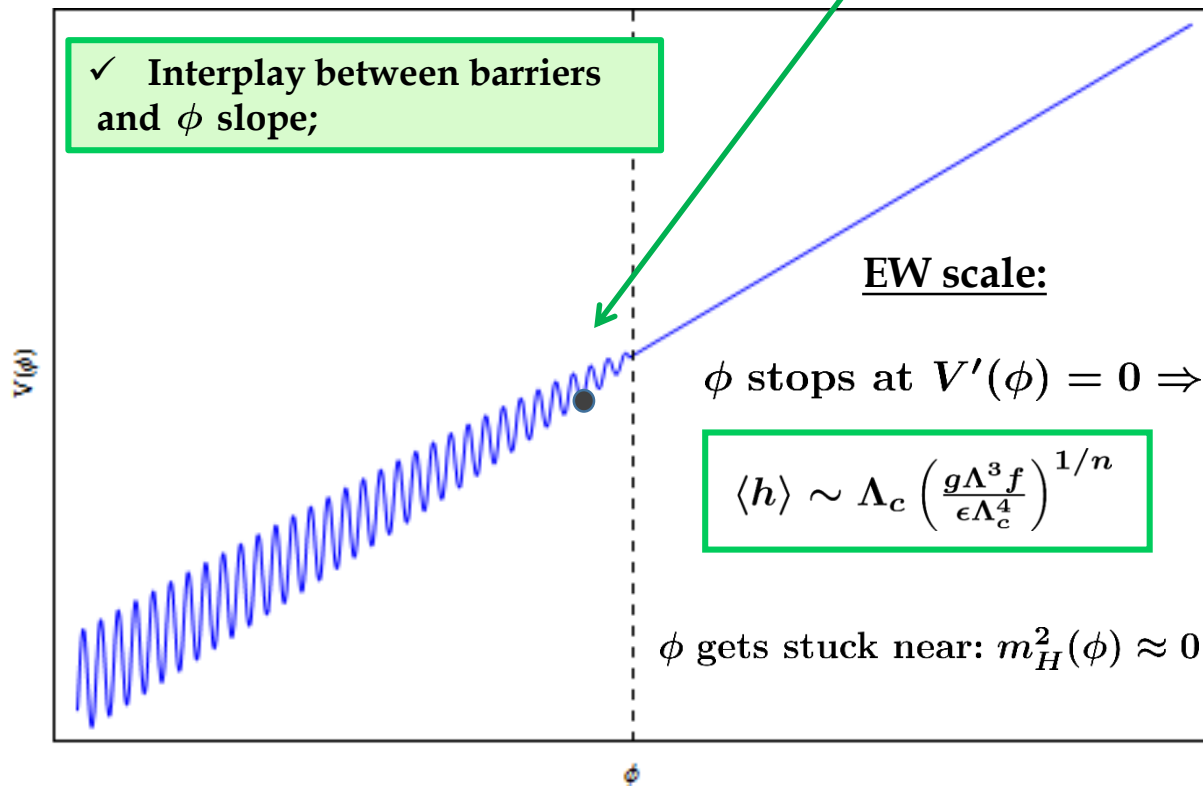
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The Relaxion Idea

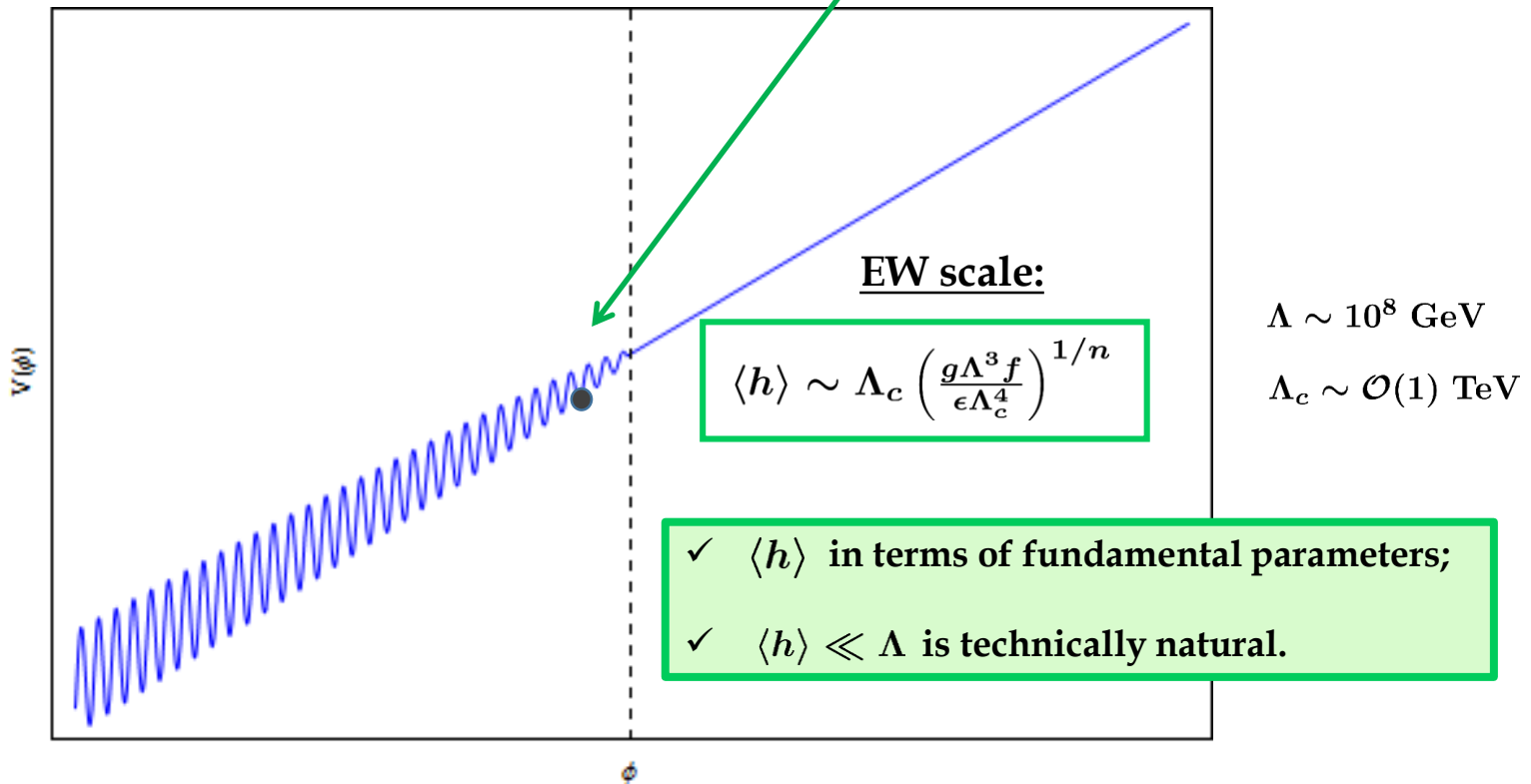
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The Relaxion Idea

“Typical” Predictions

- Natural model with $\Lambda \approx 10^8$ GeV;
- Extremely light **axion-like** states.
- Interactions with the SM through the mixing with the Higgs, suppressed by $\frac{1}{f}$ and/or g, ϵ .

Some concerns about the original idea

Let's check the symmetries

$$\mathcal{L} = \frac{1}{2}(\partial_\mu\phi)^2 - \frac{1}{2}\Lambda^2 \left(\frac{g\phi}{\Lambda} - 1\right) h^2 - g\Lambda^3\phi - \epsilon\Lambda_c^4 \left(\frac{\langle h \rangle}{\Lambda_c}\right)^n \cos\left(\frac{\phi}{f}\right)$$

UV completion? Spontaneous breaking of a Global Symmetry!

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UV completion? Spontaneous breaking of a Global Symmetry!

- $g = 0, \epsilon = 0$ \Rightarrow continuous shift symmetry: $\phi \rightarrow \phi + c$
Nambu-Goldstone boson
- $\epsilon \neq 0$ \Rightarrow breaks the continuous shift symmetry to: $\phi \rightarrow \phi + 2\pi n f$
Pseudo Nambu-Goldstone boson
- $g \neq 0$ \Rightarrow breaks the discrete shift symmetry!

Some concerns about the original idea

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UV completion? Spontaneous breaking of a Global Symmetry!

- ϵ term: $V \supset \epsilon\Lambda_c^4 \cos\left(\frac{\phi}{f}\right) \Rightarrow$ Breaks the continuous shift symmetry to:
 $\phi \rightarrow \phi + 2\pi n f$
- g terms: $V \supset g\Lambda^3\phi + \frac{1}{2}g\Lambda\phi h^2 \Rightarrow$ g is the spurion that breaks the discrete shift symmetry!

$\Rightarrow \phi$ is a pNGB (compact field range/periodicity)

$\Rightarrow g$ cannot break a gauge symmetry (the discrete shift symmetry is a redundancy)

Some concerns about the original idea

Let's check the symmetries

$$\mathcal{L} = \frac{1}{2}(\partial_\mu\phi)^2 - \frac{1}{2}\Lambda^2 \left(\frac{g\phi}{\Lambda} - 1 \right) h^2 - g\Lambda^3\phi - \epsilon\Lambda_c^4 \left(\frac{\langle h \rangle}{\Lambda_c} \right)^n \cos\left(\frac{\phi}{f}\right)$$

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The ϕ operators must be periodic!

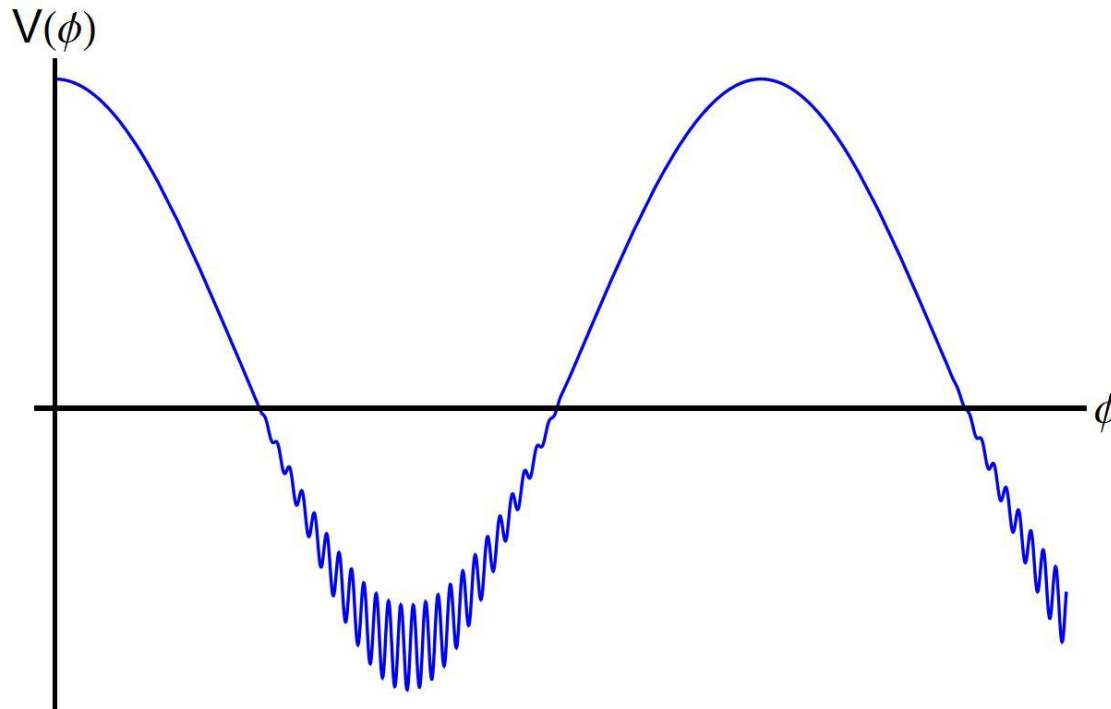
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The ϕ operators must be periodic!

Effectively, it is enough to have a hierarchy of decay constants: $F = n f \gg f$

$$V \sim A \cos\left(\frac{\phi}{F}\right) + B(h) \cos\left(\frac{\phi}{f}\right)$$



Eg.: D. E. Kaplan, R. Rattazzi; Phys.Rev. D93 (2016) no.8, 085007
K. Choi, S. H. Im; JHEP 1601 (2016) 149

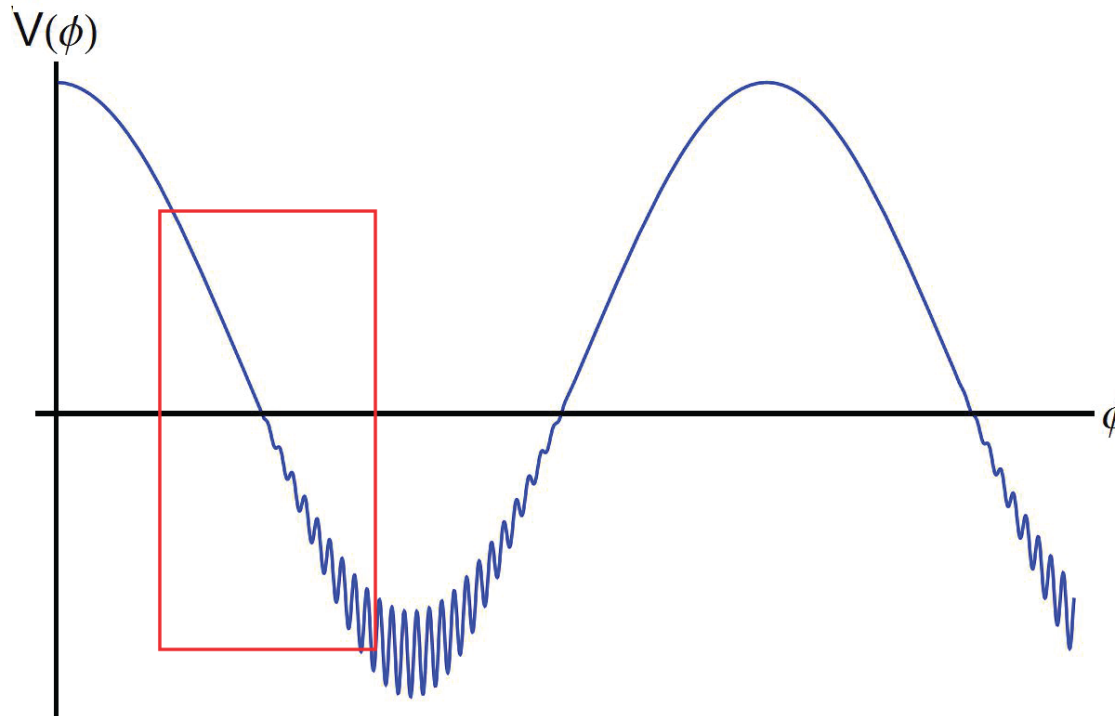
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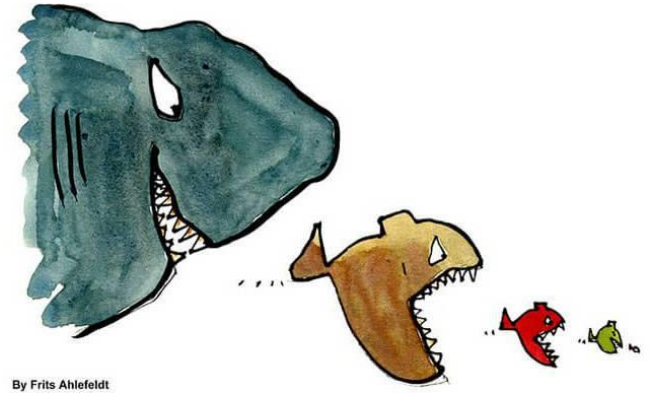
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- **Using N -site models**: NF, Lima, Machado, Matheus; Phys.Rev. D94 (2016) 015010.
- **From Warped Space**: NF, Harling, Lima, Machado, Servant; working in progress.

3. Concluding Remarks



Realizing the Relaxion with N -site Models (' N -Relaxion')

NF, Lima, Machado, Matheus; Phys.Rev. D94 (2016) 015010

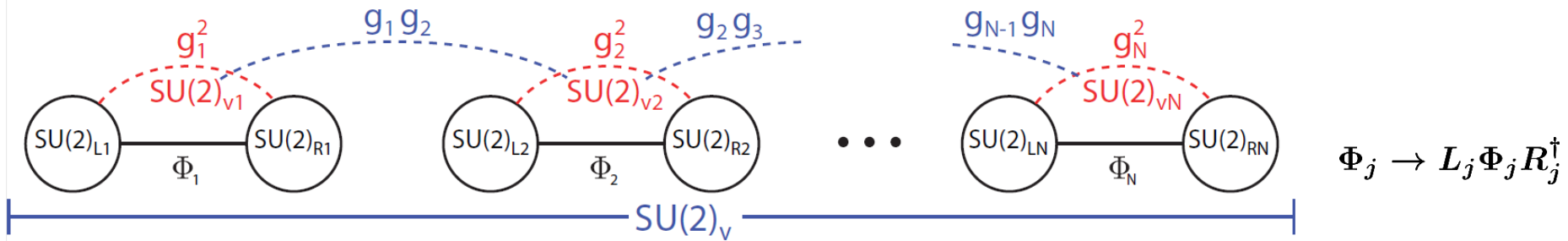
Main goals

- Obtain a realization from an N -site model (find a model close to a deconstructed extra-dimension);
- pNGB as the relaxion;
- Generate an effective scale F much larger than f ;
- Generalize to non-abelian symmetries.

N-Relaxion

NF, Lima, Machado, Matheus; Phys.Rev. D94 (2016) 015010

- pNGB as the relaxion;
- Oscillations with hierarchical periods.



- In the low energy limit, these fields are non-linearly realized:

$$\Phi_j = \frac{f}{2} e^{i\vec{\pi}_j \cdot \vec{\sigma}} / f, \quad \vec{\pi}_j \text{ are the Nambu-Goldstone bosons.}$$

$$\vec{\pi}^T \cdot M_\pi^2 \cdot \vec{\pi} \equiv \sum_{j=1}^{N-1} f^2 (g_j \vec{\pi}_j - g_{j+1} \vec{\pi}_{j+1})^2$$

$$\vec{\pi}^T \equiv \{\vec{\pi}_1, \dots, \vec{\pi}_N\}$$

N-Relaxion: Message

NF, Lima, Machado, Matheus; Phys.Rev. D94 (2016) 015010

The zero mode η_0 (the would-be relaxion), and its profile is

$$\vec{\eta}_0 = \sum_{j=1}^N \frac{q^{N-j}}{\sqrt{\sum_{k=1}^N q^{2(k-1)}}} \vec{\pi}_j$$

Exponentially
localized in the
last site

$$g_j \rightarrow q^j, \quad 0 < q < 1$$

$$f_j \equiv f q^{j-N} \mathcal{C}_N$$

Identical to the one obtained for a pNGB in the deconstruction of AdS₅.

$$\mathcal{L}_\eta = \sum_{j=1}^N \left[\frac{1}{2} \partial_\mu \vec{\eta}_0 \cdot \partial^\mu \vec{\eta}_0 + f^4 (2 - \delta_{j,1} - \delta_{j,N}) q^{2j} \cos \frac{\eta_0}{f_j} \right] + \sum_{j=1}^{N-1} f^4 q^{2j+1} \sin \frac{\eta_0}{f_j} \sin \frac{\eta_0}{f_{j+1}}$$

$\eta_0 \equiv \sqrt{\vec{\eta}_0 \cdot \vec{\eta}_0}$

Oscillating with different scales

$$f_j \equiv f q^{j-N} \mathcal{C}_N \begin{cases} \rightarrow f_{\max} = f_1 \approx f/q^{N-1} \\ \rightarrow f_{\min} = f_N \approx f \end{cases}$$

Warped Relaxion

NF, Harling, Lima, Machado, Servant; working in progress.

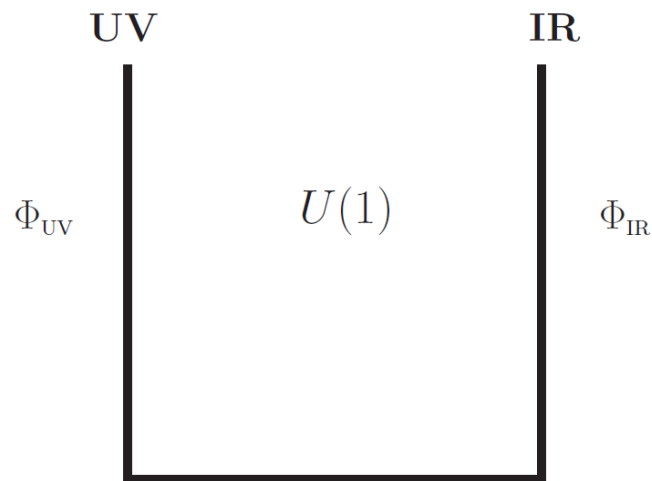
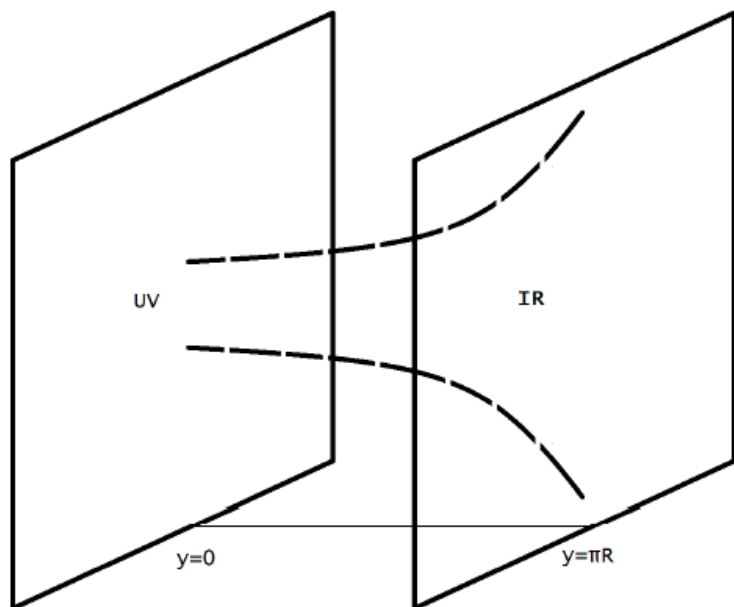
Main goals

- Obtain a relaxion realization from a warped extra-dimension (AdS) model;
- Hierarchy from the geometry;
- pNGB as the relaxion;

Warped Relaxion

NF, Harling, Lima, Machado, Servant; working in progress.

$U(1)$ gauge symmetry in a slice of the AdS_5 space that is spontaneously broken at the UV and IR branes by boundary scalar fields



$$ds^2 = e^{-2ky} \eta_{\mu\nu} dx^\mu dx^\nu - dy^2$$

$$ds^2 = \frac{1}{(kz)^2} (\eta_{\mu\nu} dx^\mu dx^\nu - dz^2) \text{ conformal coordinates}$$

$$z = e^{ky}/k$$

Warped Relaxion

NF, Harling, Lima, Machado, Servant; working in progress.

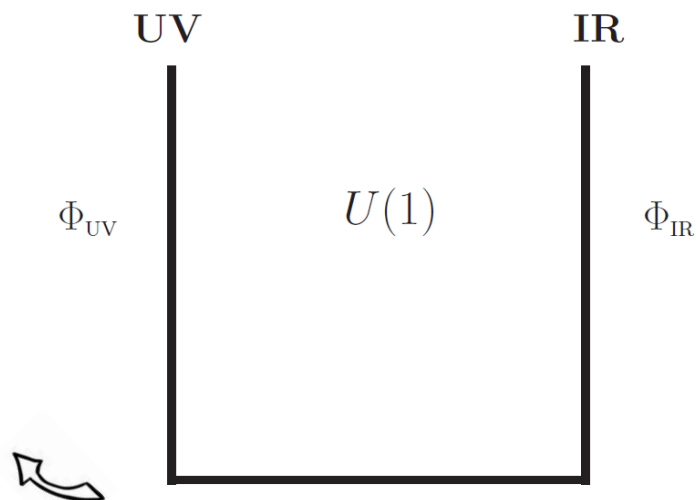
$U(1)$ gauge symmetry in a slice of the AdS_5 space that is spontaneously broken at the UV and IR branes by boundary scalar fields

$$\Phi_i(x) = \frac{1}{\sqrt{2}}[v_i + h_i(x)]e^{i\pi_i(x)/v_i}$$

$$v_i \rightarrow \infty$$

$$A_\mu|_{UV,IR} = 0; \partial_5(A_5/z)|_{UV,IR} = 0$$

Only a global subgroup of the original $U(1)$ gauge symmetry survives.



See eg.:hep-ph/0510275

Warped Relaxion

NF, Harling, Lima, Machado, Servant; working in progress.

The boundary conditions are chosen such that there is a zero mode of A_5 in the spectrum

Scalar component of a 5D gauge boson

$$A_5(x, z) = \mathcal{N} z A_5^0(x)$$

$$\mathcal{N} \approx k e^{-kL} \sqrt{2kL}, \quad kL \gg 1$$

- L is the size of the extra dimension;
- k is the curvature;
- g_5 is the 5D gauge coupling.

Warped Relaxion

NF, Harling, Lima, Machado, Servant; working in progress.

Anomalous couplings can be induced by brane-localized fermions appropriately charged under the residual global symmetry

$$\sim \int dz \frac{A_5}{k} \epsilon^{\mu\nu\rho\sigma} G_{\mu\nu}(x) G_{\rho\sigma}(x) \delta(z - z_i)$$



New strongly-interacting
gauge sectors

Warped Relaxion

NF, Harling, Lima, Machado, Servant; working in progress.

After canonical normalization:

$$\frac{A_5^0(x)}{f_{\text{UV}}} G \tilde{G}; \quad f_{\text{UV}} \approx k e^{kL}$$

$$\frac{A_5^0(x)}{f_{\text{IR}}} G' \tilde{G}'; \quad f_{\text{IR}} \approx k e^{-kL}$$

$$\tilde{G}_{\mu\nu}^a = \frac{1}{2} \epsilon_{\mu\nu\rho\sigma} G^{\rho\sigma,a}$$



The warp factor can explain the large hierarchy between the decay constants in the relaxion potential

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3. Concluding Remarks

Concluding Remarks & Outlook

- UV **sensibility** to the Higgs mass: one of the leading motivation for new physics at the LHC;
- No compelling evidence of BSM at the LHC current data! 🙄



Why is
this stable?

Concluding Remarks & Outlook

- UV **sensibility** to the Higgs mass: one of the leading motivation for new physics at the LHC;
- No compelling evidence of BSM at the LHC current data! 🙄
- Relaxation models: **proof of concept**. If self-consistent, then the hierarchy problem cannot be an argument for new physics at the TeV scale.



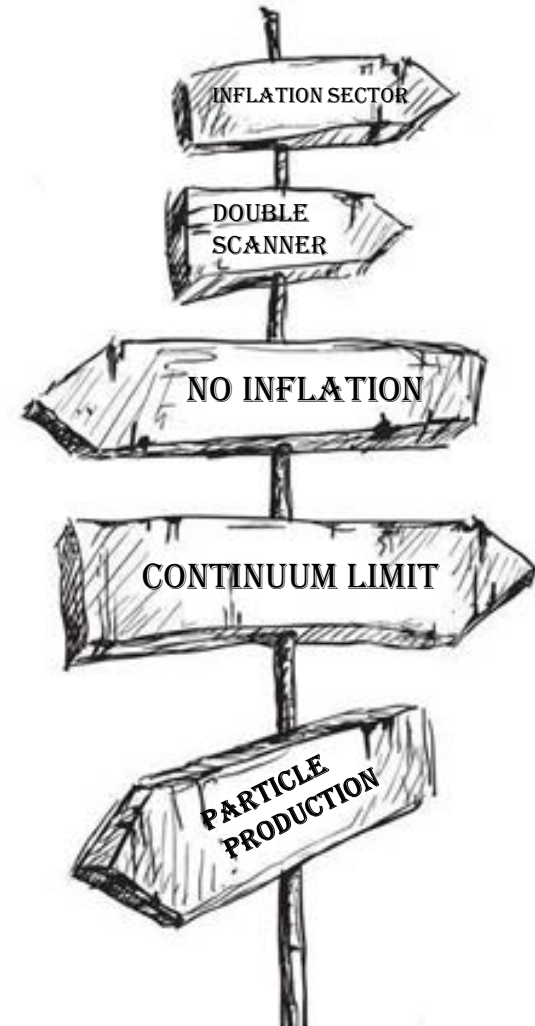
Chiricahua National Monument, Arizona

Concluding Remarks & Outlook

- **N-Relaxion:** *N*-site model generating a large-scale hierarchy; ✓
NF, Lima, Machado, Matheus, Phys. Rev. D 94, 015010 (2016).

- pNGB as the relaxion;
- Oscillations with hierarchical periods.

- **Warped Relaxion: a continuum limit;**
NF, Harling, Lima, Machado, Servant; working in progress.



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- **N-Relaxion:** *N*-site model generating a large-scale hierarchy; ✓

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- **pNGB as the relaxion;**
- **Oscillations with hierarchical periods.**

- **Warped Relaxion: a continuum limit;**

NF, Harling, Lima, Machado, Servant; working in progress.

- **Many possible directions:**

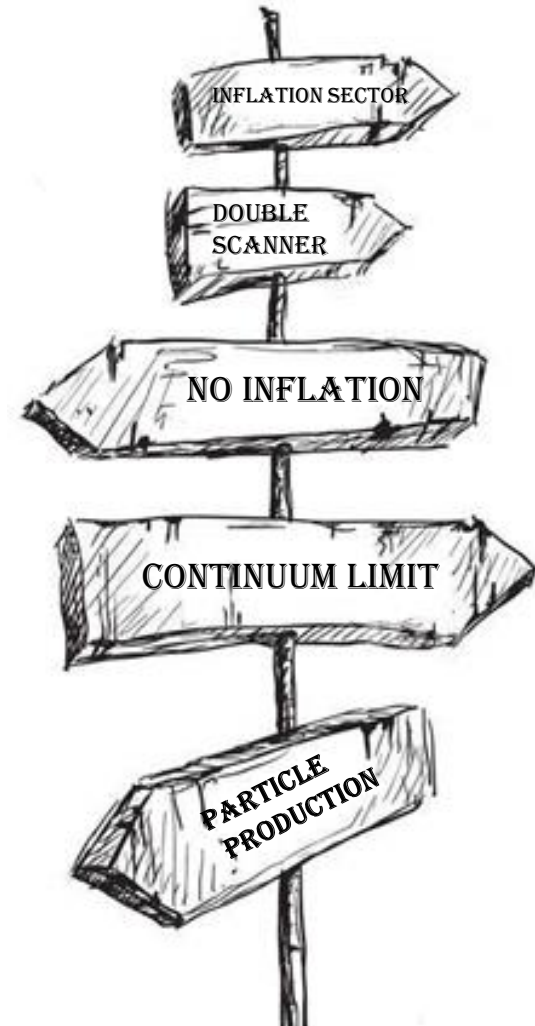
- **Alternatives to inflation?**

Eg.: Hook, Marques-Tavares JHEP 1612 (2016) 101

- **Can we apply this mechanism to generate other large-scale hierarchies?**

Abbott's attempt to solve the CC problem; Phys. Lett. B 150 , 427 (1985)

- ...



Thanks!

QCD Relaxion

P. W. Graham, D. E. Kaplan, S. Rajendran; Phys. Rev. Lett. 115, 221801 (2015)

ϕ is the QCD axion, $\mathcal{L} \supset \frac{g_s^2}{32\pi^2} \frac{\phi}{f} G_{\mu\nu} \tilde{G}^{\mu\nu}$

Instanton effects generate: $V(\phi, H) \sim m_u(H) \langle q\bar{q} \rangle \cos(\phi/f)$

$$\Lambda_c = \Lambda_{QCD} \quad \epsilon = Y_u \quad V_{\text{stop}} = \epsilon \Lambda_c^4 \left(\frac{\langle h \rangle}{\Lambda_c} \right)^n \cos \frac{\phi}{f} \quad n = 1$$

$$\Lambda < 10^7 \text{ GeV} \left(\frac{10^9 \text{ GeV}}{f} \right)^{1/6}$$

$$10^9 \text{ GeV} < f < 10^{12} \text{ GeV}$$

Star cooling DM abundance

QCD Relaxion

P. W. Graham, D. E. Kaplan, S. Rajendran; Phys. Rev. Lett. 115, 221801 (2015)

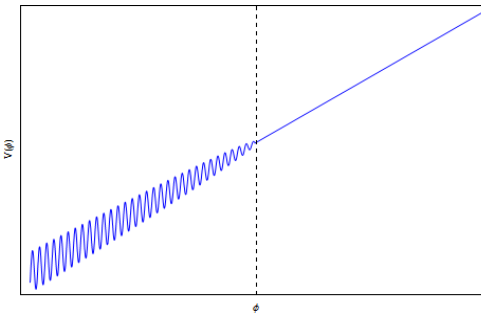
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Instanton effects generate: $V(\phi, H) \sim m_u(H) \langle q\bar{q} \rangle \cos(\phi/f)$

- **But this model spoils the axion solution to the strong CP problem** ($\theta_{\text{QCD}} < 10^{-10}$);
- **If the relaxion is the QCD axion, its vev determines the QCD theta parameter.**

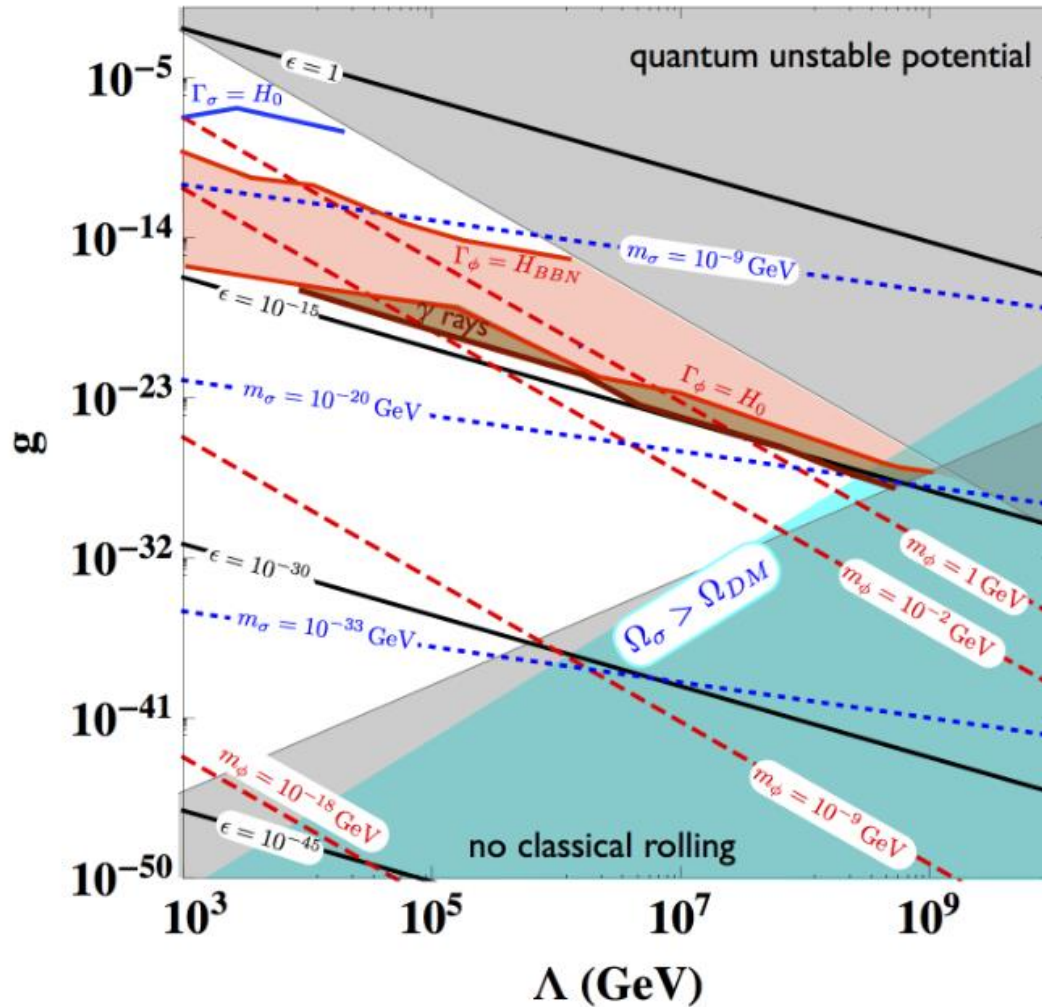
$$\Rightarrow \theta_{\text{QCD}} = \langle \frac{\Delta\phi}{f} \rangle \sim \mathcal{O}(1) \quad !?$$

(Due to the tilt of the potential)



Ways to solve this: adding dynamics at the end of inflation (removes the slope of the potential). Cutoff scale at most ~ 30 TeV.

Typical predictions

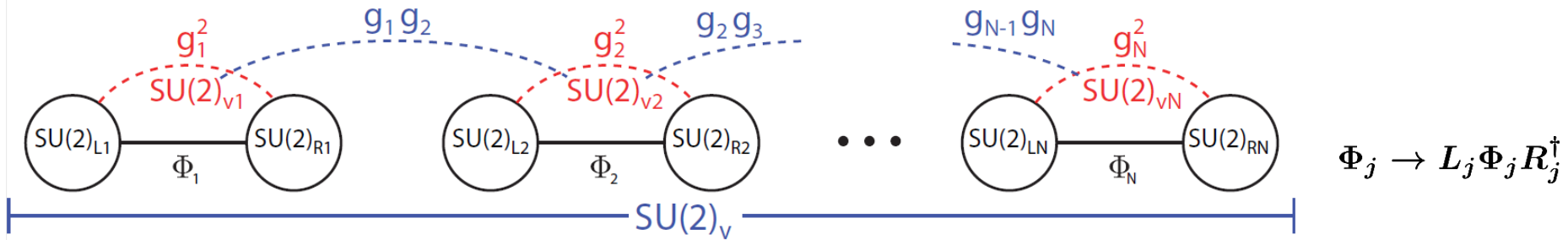


J.R. Espinosa, C. Grojean, G. Panico, A. Pomarol O. Pujolàs, G. Servant;
Phys.Rev.Lett. 115 (2015) no.25, 251803

N-Relaxion

NF, L. de Lima, C. S. Machado, R. D. Matheus; Phys.Rev. D94 (2016) no.1, 015010

- pNGB as the relaxion;
- Oscillations with hierarchical periods.



- Φ_j gets a vev $\langle \Phi_j \rangle = \frac{f}{2}$, spontaneously breaking $SU(2)_{L_j} \times SU(2)_{R_j} \rightarrow SU(2)_{V_j}$;
- $g_j g_{j+1}$ terms break $SU(2)_{V_j} \times SU(2)_{V_{j+1}} \rightarrow SU(2)_{V_{j,j+1}}$;

Small symmetry breaking terms

$$\mathcal{L}_\Phi = \sum_{j=1}^N \text{Tr} \left[\partial_\mu \Phi_j^\dagger \partial^\mu \Phi_j + \frac{f^3}{2} (2 - \delta_{j,1} - \delta_{j,N}) g_j^2 (\Phi_j + \Phi_j^\dagger) \right] - \frac{f^2}{2} \sum_{j=1}^{N-1} g_j g_{j+1} \text{Tr} \left[(\Phi_j - \Phi_j^\dagger)(\Phi_{j+1} - \Phi_{j+1}^\dagger) \right]$$

N-Relaxion

NF, L. de Lima, C. S. Machado, R. D. Matheus; Phys.Rev. D94 (2016) no.1, 015010

Mass Matrix


$$\mathcal{L}_\Phi = \sum_{j=1}^N \text{Tr} \left[\partial_\mu \Phi_j^\dagger \partial^\mu \Phi_j + \frac{f^3}{2} (2 - \delta_{j,1} - \delta_{j,N}) g_j^2 (\Phi_j + \Phi_j^\dagger) \right] - \frac{f^2}{2} \sum_{j=1}^{N-1} g_j g_{j+1} \text{Tr} \left[(\Phi_j - \Phi_j^\dagger) (\Phi_{j+1} - \Phi_{j+1}^\dagger) \right]$$

Writing in terms of the $\vec{\pi}_j$'s, we obtain the mass matrix for the pNGBs:

$$\vec{\pi}^T \cdot M_\pi^2 \cdot \vec{\pi} \equiv \sum_{j=1}^{N-1} f^2 (g_j \vec{\pi}_j - g_{j+1} \vec{\pi}_{j+1})^2$$

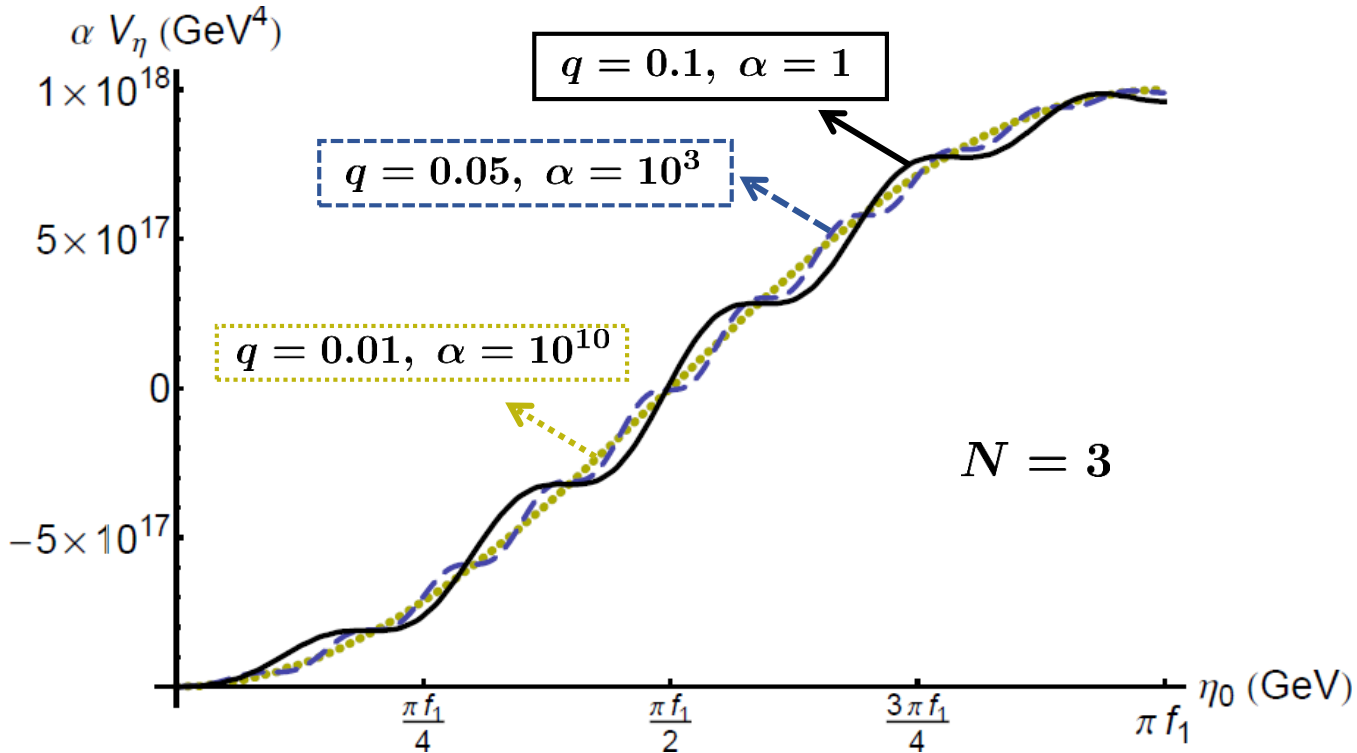
$$\vec{\pi}^T \equiv \{\vec{\pi}_1, \dots, \vec{\pi}_N\}$$

✓ The parametrization $g_j \rightarrow q^j$, $0 < q < 1$ results in a mass matrix that is identical to the one obtained for a pNGB Wilson line in the deconstruction of AdS₅.


$$M_\pi^2 = f^2 \begin{pmatrix} q^2 & -q^3 & 0 & \dots & 0 & 0 \\ -q^3 & 2q^4 & -q^5 & \dots & 0 & 0 \\ 0 & -q^5 & 2q^6 & \dots & 0 & 0 \\ \vdots & \vdots & \vdots & \dots & \vdots & \vdots \\ 0 & 0 & 0 & \dots & 2q^{2(N-1)} & -q^{2N-1} \\ 0 & 0 & 0 & \dots & -q^{2N-1} & q^{2N} \end{pmatrix}$$

N-Relaxion

NF, L. de Lima, C. S. Machado, R. D. Matheus; Phys.Rev. D94 (2016) no.1, 015010



- Quite large values of q to exaggerate the features of the potential;
- the slope quickly gets smooth;
- before the inclusion of the Higgs the relaxion is able to roll down.