



Partial support: CUHK VC Discretionary Fund, RGC CUHK3/CRF/10R

29<sup>th</sup> Recontres de Blois

May 28 – June 03, 2017, Blois, France



#### Recent results from Daya Bay

- The Daya Bay Reactor Neutrino Experiment
- Recent oscillation results
- Absolute reactor anti-neutrino flux and spectrum
- Search for a light sterile neutrino
- More searches

#### **Neutrino Oscillations**

- Each flavor state is a mixture of mass eigenstates
- Described by a neutrino mixing matrix

$$\begin{bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{bmatrix} = U \begin{bmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{bmatrix}$$

$$\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix} \times \begin{pmatrix} \cos\theta_{13} & 0 & e^{-i\delta_{CP}}\sin\theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{CP}}\sin\theta_{13} & 0 & \cos\theta_{13} \end{pmatrix} \times \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \times \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha/2} & 0 \\ 0 & 0 & e^{i\alpha/2+i\beta} \end{pmatrix}$$

The Maki-Nakagawa-Sakata-Pontecorvo Matrix

- A freely propagating  $v_e$  will oscillate into other types
- In general,  $|\langle v_{\mu,\tau}(t)|v_e(0)\rangle|^2 \neq 0$

$$|\langle v_e(t)|v_e(0)\rangle|^2 \approx 1 - \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta m_{31}^2 L}{4E}\right) + \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left(\frac{\Delta m_{21}^2 L}{4E}\right)$$



# The Daya Bay Reactor Neutrino Experiment

F. P. An et al., Daya Bay Collaboration, NIM A **811**, 133 (2016); PRD **95**, 072006 (2017).

# Reactor expt.: a clean way to measure $\theta_{13}$

$$P(\overline{v}_e \to \overline{v}_e) \approx 1 - \sin^2 2\theta_{13} \sin^2 \frac{\Delta m_{ee}^2 L}{4E_v} - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \frac{\Delta m_{21}^2 L}{4E_v}$$

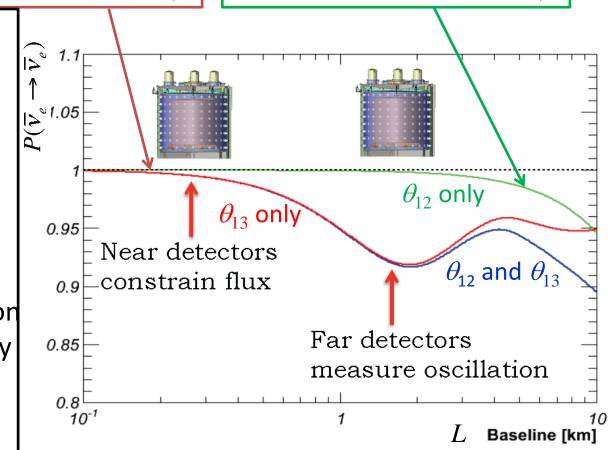
- Reactor: abundant, free, pure source of  $\overline{v}_{e}$
- disappearance of  $\overline{\nu}_{\rm e}$  at small L depends only on  $\theta_{13}$

#### Near-far configuration

Near detectors:  $\overline{v}_e$  flux and spectrum for normalization

Far detectors: near oscillation maximum for best sensitivity

Relative measurement: cancel out most systematics





# Near/far Configuration

#### Minimize systematic uncertainties:

reactor-related: cancelled by near-far ratio

detector-related: use 'identical' detectors, careful calibration

$$\frac{R_{\text{Far}}}{R_{\text{Near}}} = \left(\frac{L_{\text{Near}}}{L_{\text{Far}}}\right)^2 \frac{N_{\text{Far}}}{N_{\text{Near}}} \frac{\varepsilon_{\text{Far}}}{\varepsilon_{\text{Near}}} \left(\frac{P_{\text{surv}}(L_{\text{Far}})}{P_{\text{surv}}(L_{\text{Near}})}\right)$$

$$\overline{v}_{\text{e}} \text{ detection ratio } \frac{1/r^2}{r^2} \frac{number}{number} \frac{\text{detector}}{\text{of protons}} \frac{\text{Survival prob.}}{\text{efficiency}} \frac{1/r^2}{r^2}$$

Parameter	CHOOZ error	Near/far configuration
Reaction cross section	1.9 %	Cancelled out
Number of protons	0.8 %	Reduced to ~ 0.03%
Detection efficiency	1.5 %	Reduced to ~ 0.2%
Reactor power	0.7 %	Reduced to ~ 0.04%
Energy released per	0.6 %	Cancelled out
fission		
CHOOZ Combined	2.7 %	~ 0.21%

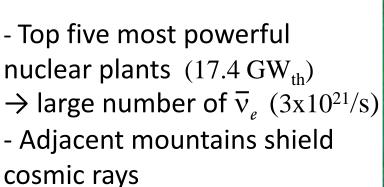
# Daya Bay (China)





#### Daya Bay Experiment

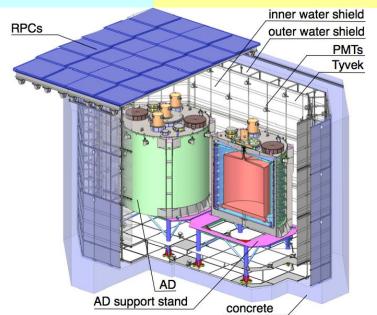






### Daya Bay detectors





RPC: muon veto

Water pool: muon veto + shielding

from environmental radiations

(2.5m water)

8 functionally identical anti-neutrino detectors (AD) to suppress systematic uncertainties

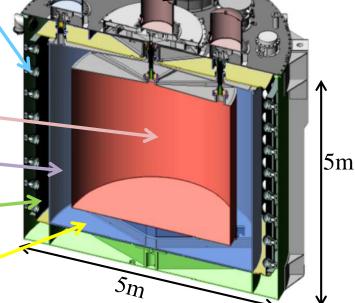
Calibration units

192 8" PMTs

#### 3 zone cylindrical vessels

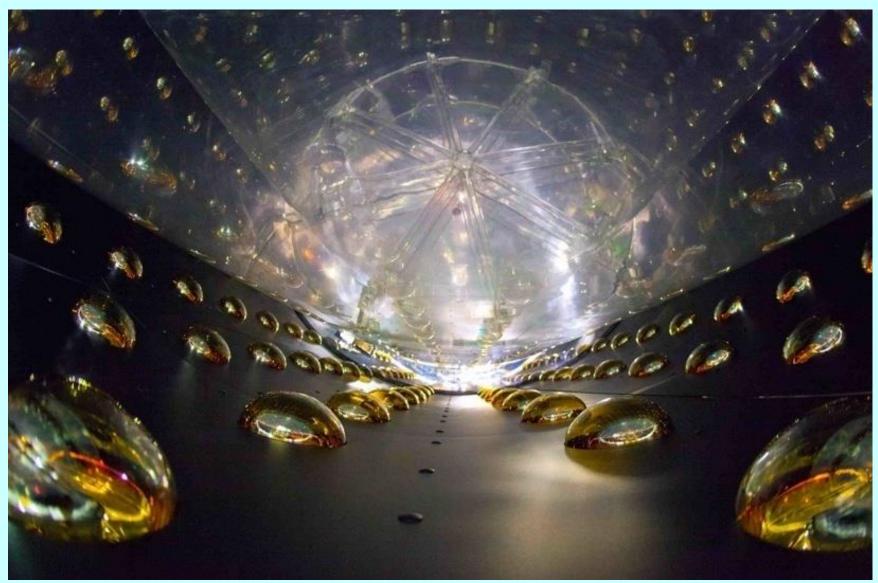
	Liquid	Mass	Function
Inner acrylic	Gd-doped liquid scint.	20 t	Antineutrino target
Outer acrylic	Liquid scintillator	20 t	Gamma catcher
Stainless steel	Mineral oil	40 t	Radiation shielding

Top and bottom reflectors: more light, more uniform detector response



# Interior of an AD

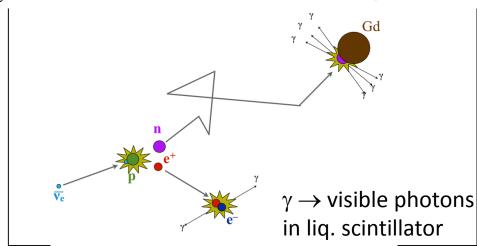


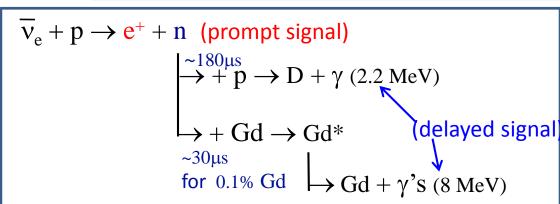


#### Anti-neutrino detection



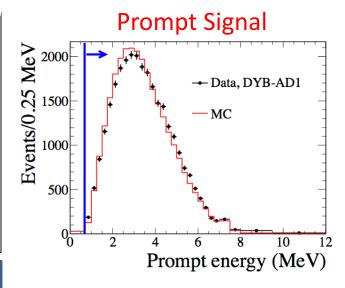
#### $\overline{v}_{e}$ detected via inverse beta-decay (IBD):



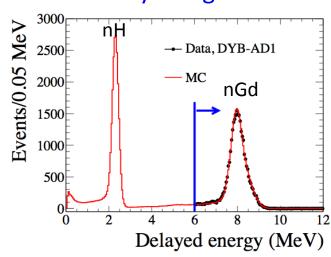


#### Powerful background rejection!

$$E_{\rm v} \approx T_{\rm e+} + T_{\rm n} + (m_{\rm n} - m_{\rm p}) + m_{\rm e+} \approx T_{\rm e+} + 1.8 \,{\rm MeV}$$



#### **Delayed Signal**



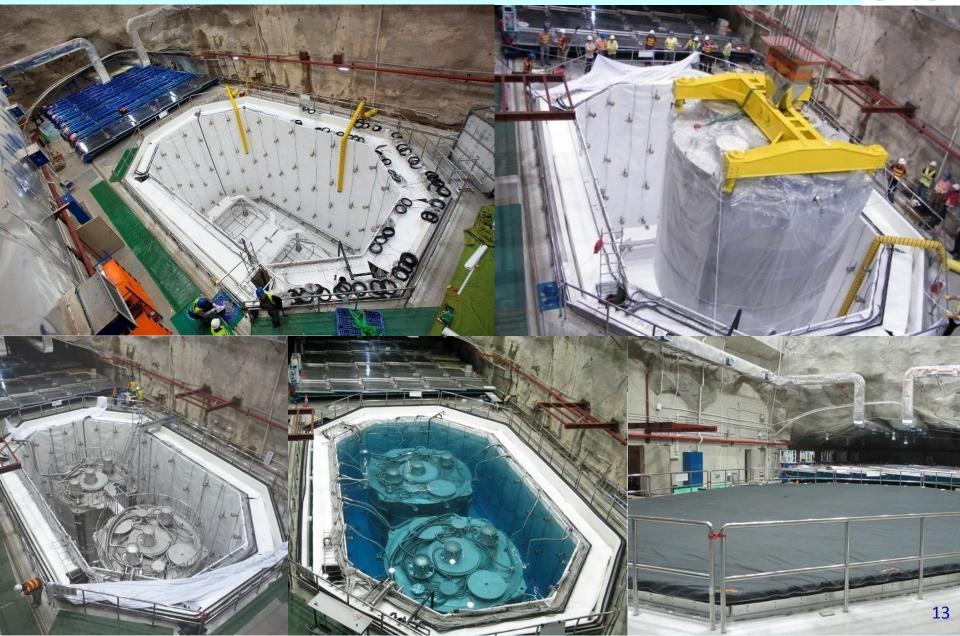
#### The Daya Bay Collaboration





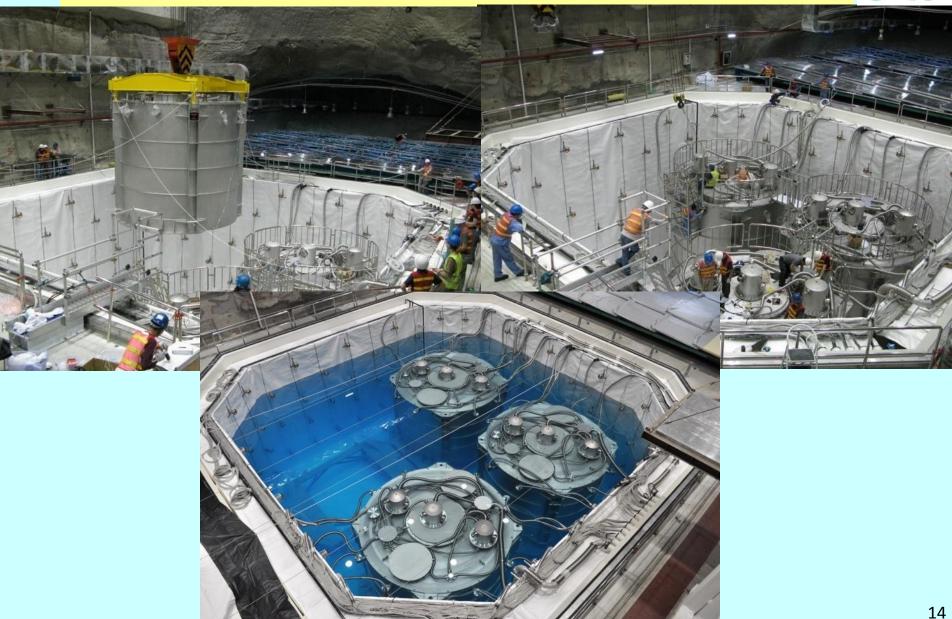
# AD Installation - Near Hall





# AD Installation - Far Hall

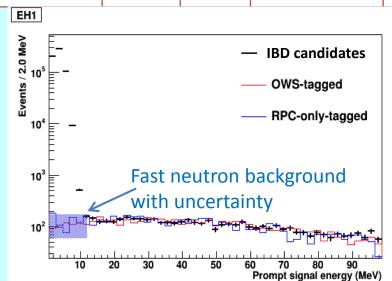


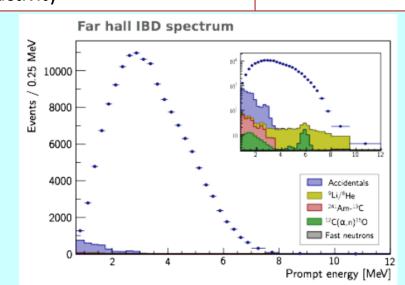


# Background



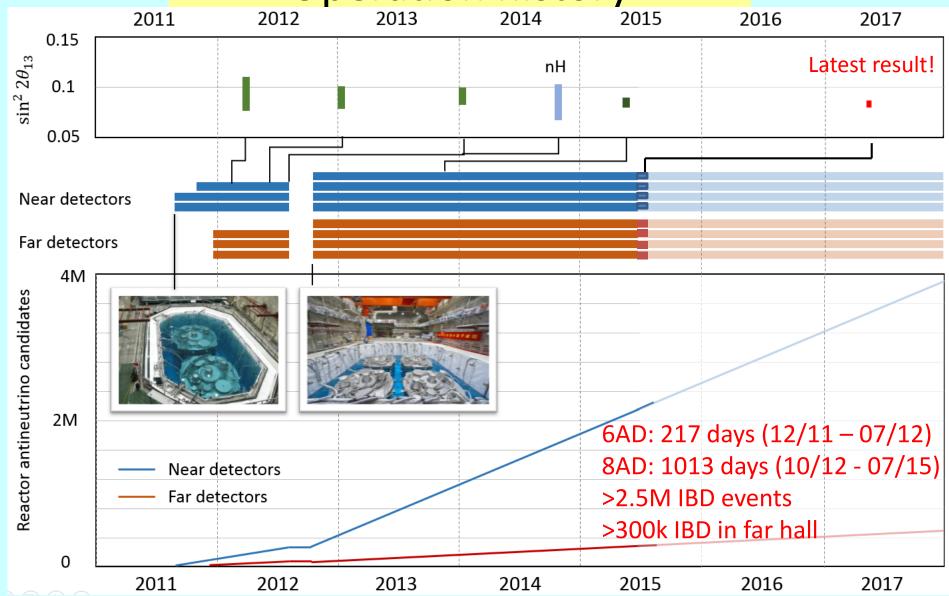
Background	Near	Far	Uncertainty	Method	Improvement	
Accidentals	1.4%	2.3%	~1%	Statistically calculated from uncorrelated singles	Extend to larger data set	
<sup>9</sup> Li/ <sup>8</sup> He	0.4%	0.4%	~44%	Measured with after-muon events	Extend to larger data set	
Fast neutrons	0.1%	0.1%	~13%	Measured from RPC+OWS tagged muon events	Model independent measurement	
AmC source	0.03%	0.2%	~45%	MC benchmarked with single gamma and strong AmC source	Two sources are taken out in Far site ADs	
α-n	0.01%	0.1%	~50%	Calculated from measured radioactivity	Reassess systematics	







#### **Operation history**





# Signal and background summary

	E	H1	EH2		EH3			
	AD1	AD2	AD3	AD8	AD4	AD5	AD6	AD7
$\Delta N_{\rm p}$ [%]	$0.00 \pm 0.03$	$0.13 \pm 0.03$	$-0.25 \pm 0.03$	$0.02 \pm 0.03$	$-0.12 \pm 0.03$	$0.24 \pm 0.03$	$-0.25 \pm 0.03$	$-0.05 \pm 0.03$
				Selection	on A			
$\bar{\nu}_e$ candidates	597616	606349	567196	466013	80479	80742	80067	66862
DAQ live time [days]	1117.178	1117.178	1114.337	924.933	1106.915	1106.915	1106.915	917.417
$\epsilon_{\mu}$	0.8255	0.8221	0.8573	0.8571	0.9824	0.9823	0.9821	0.9826
$ar{\epsilon}_{ m m}$	0.9744	0.9747	0.9757	0.9757	0.9759	0.9758	0.9756	0.9758
Accidentals [day <sup>-1</sup> ]	$8.46\pm0.09$	$8.46\pm0.09$	$6.29 \pm 0.06$	$6.18 \pm 0.06$	$1.27\pm0.01$	$1.19\pm0.01$	$1.20\pm0.01$	$0.98 \pm 0.01$
Fast neutron [AD <sup>-1</sup> day <sup>-1</sup> ]	0.79	± 0.10	0.57 =	± 0.07		0.05	$\pm 0.01$	
<sup>9</sup> Li, <sup>8</sup> He [AD <sup>-1</sup> day <sup>-1</sup> ]	2.46	± 1.06	1.72 =	± 0.77		0.15	$\pm 0.06$	
<sup>241</sup> Am- <sup>13</sup> C, 6-AD [day <sup>-1</sup> ]	$0.27\pm0.12$	$0.25\pm0.11$	$0.28 \pm 0.13$		$0.22 \pm 0.10$	$0.21\pm0.10$	$0.21 \pm 0.10$	
<sup>241</sup> Am- <sup>13</sup> C, 8-AD [day <sup>-1</sup> ]	$0.15\pm0.07$	$0.16\pm0.07$	$0.13 \pm 0.06$	$0.15\pm0.07$	$0.04\pm0.02$	$0.03\pm0.02$	$0.03\pm0.02$	$0.05\pm0.02$
$^{13}C(\alpha, n)^{16}O \text{ [day}^{-1}]$	$0.08 \pm 0.04$	$0.07 \pm 0.04$	$0.05 \pm 0.03$	$0.07 \pm 0.04$	$0.05 \pm 0.03$	$0.05 \pm 0.03$	$0.05 \pm 0.03$	$0.05 \pm 0.03$
$ar{ u}_e$ rate, $R_{ar{ u}}$ [day <sup>-1</sup> ]	$653.03 \pm 1.37$	$665.42 \pm 1.38$	599.71 ± 1.12	$593.82 \pm 1.18$	$74.25 \pm 0.28$	$74.60 \pm 0.28$	$73.98 \pm 0.28$	$74.73 \pm 0.30$

F. P. An et al., Daya Bay Collaboration, PRD **95**, 072006 (2017).

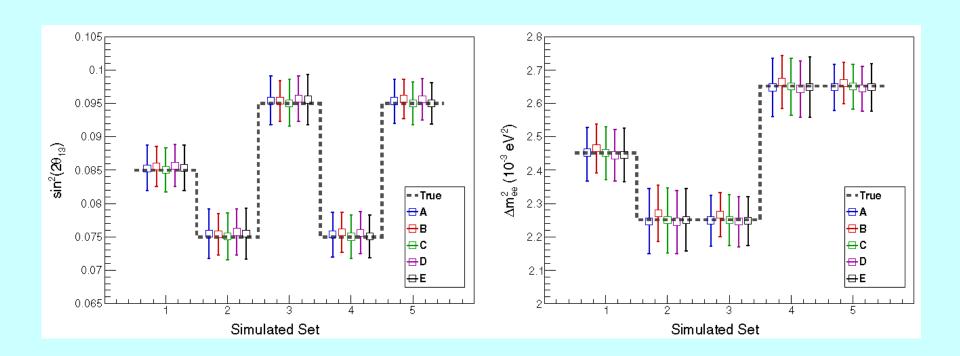


# Recent Oscillation Results

F. P. An et al., Daya Bay Collaboration, PRD **95**, 072006 (2017).

#### Oscillation results

5 independent analysis methods, all consistent with each other and validated by simulated data generated with various  $\sin^2\!2\theta_{13}$  and  $\Delta m^2_{\rm ee}$ 

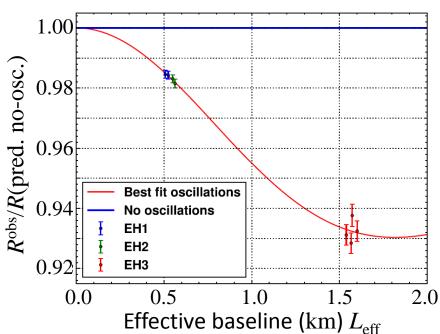


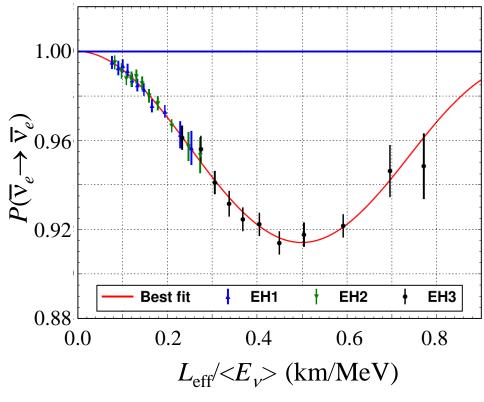
#### Oscillation results



$$P_{ee} = 1 - \sin^2 2\theta_{13} \sin^2 (\Delta m_{ee}^2 L/4E_{\nu}) - \sin^2 2\theta_{12} \cos^4 2\theta_{13} \sin^2 (\Delta m_{21}^2 L/4E_{\nu})$$

- Far/near relative measurement
- Oscillation parameters measured with rate + spectral distortion
- Both consistent with neutrino oscillation interpretation

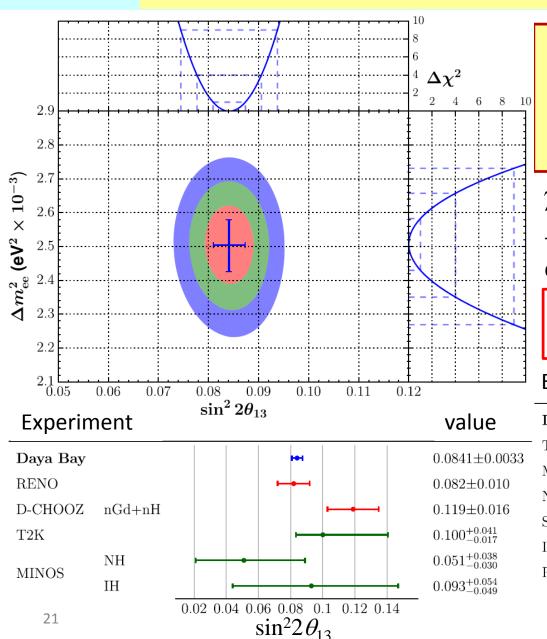


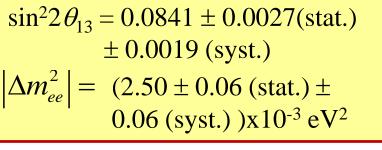


F. P. An et al., Daya Bay Collaboration, PRD **95**, 072006 (2017).

#### Oscillation results



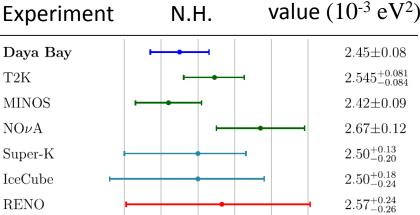




$$\chi^2/NDF = 232.6/263$$

- Most precise measurement (< 4%) of  $\sin^2 2\theta_{13}$  and  $|\Delta m^2_{ee}| \rightarrow$ 

 $\Delta m_{32}^2 = (2.45\pm0.08) \times 10^{-3} \text{ eV}^2 \text{ (N.H.)}$  $(-2.56\pm0.08) \times 10^{-3} \text{ eV}^2 \text{ (I.H.)}$ 



2.6

 $\Delta m^2_{32}(10^{-3} \text{ eV}^2)$ 

2.7

2.8

#### Independent $\theta_{13}$ measurement with nH

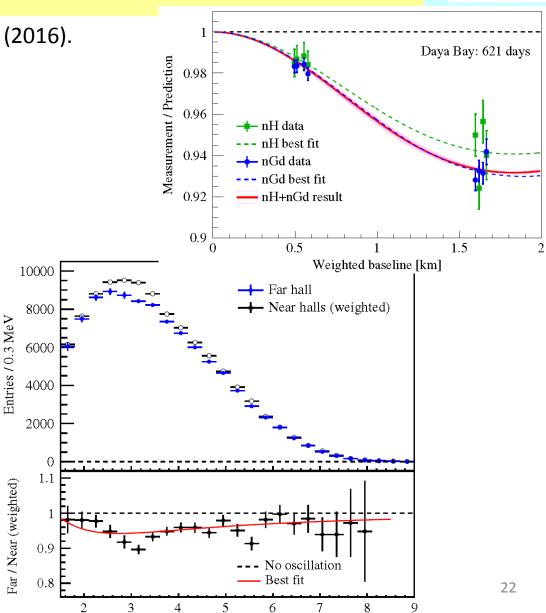


Daya Bay Collaboration, PRD93, 072011 (2016).

- Independent measurement, different systematics
- Longer capture time, lower delayed energy (2.2 MeV)
   →high accidental background
- → higher prompt energy cut
   (> 1.5 MeV) + prompt-to-delay
   distance cut (< 0.5 m)</li>
- nH:  $\sin^2 2\theta_{13} = 0.071 \pm 0.011$
- Combined nH + nGd:

$$\sin^2 2\theta_{13} = 0.082 \pm 0.004$$

•  $3^{\rm rd}$  world's most precise measurement of  $\theta_{13}$  after Daya Bay nGd and RENO



Prompt Energy [MeV]

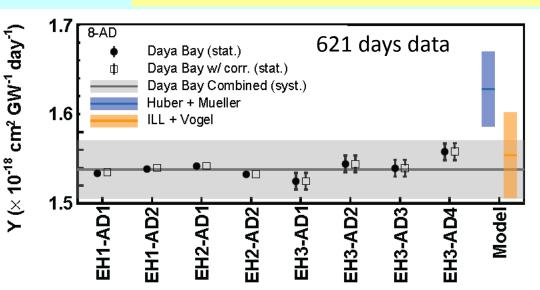


# Absolute reactor anti-neutrino flux and spectrum

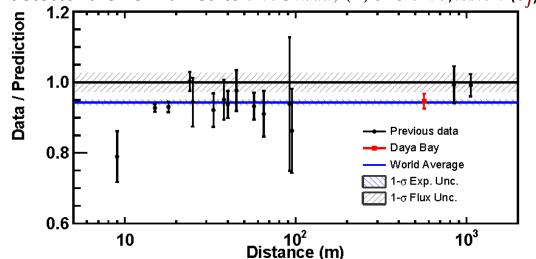
F. P. An et al., Daya Bay Collaboration, PRL **116**, 061801 (2016); Chinese Physics C **41**(1), 13002 (2017); arXiv:1704.01082v1, PRL 2017.

#### Reactor antineutrino flux





Measured IBD events (background subtracted) in each detector are normalized to  $cm^2/GW/day$  (Y) and  $cm^2/fission$  ( $\sigma_f$ ).



Daya Bay's reactor antineutrino flux measurement is consistent with previous short baseline expts.

4-AD (near halls) measurement  $Y = (1.53 \pm 0.03) \times 10^{-18} \, \mathrm{cm^2 GW^{-1} day^{-1}}$   $\sigma_f = (5.91 \pm 0.12) \times 10^{-43} \, \mathrm{cm^2 fission^{-1}}$ 

Compare to flux model
Data/Prediction (Huber+Mueller)

 $0.946 \pm 0.020$ Data/Prediction (ILL+Vogel)  $0.992 \pm 0.021$ 

Effective baseline (near sites)  $L_{\text{eff}} = 573 \text{m}$ 

Effective fission fractions  $F_i$ 

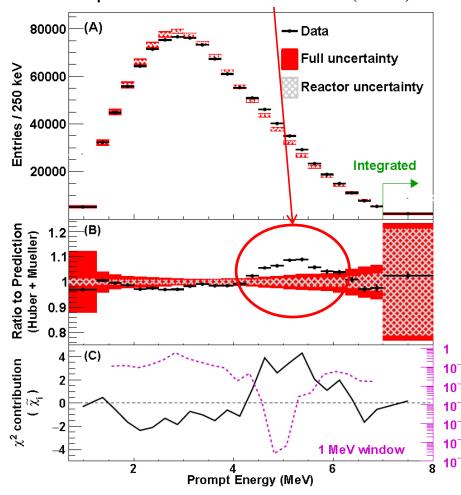
<sup>235</sup> U	238U	<sup>239</sup> Pu	<sup>241</sup> Pu
0.561	0.076	0.307	0.056

Global comparison of measurement and prediction (Huber+Mueller)

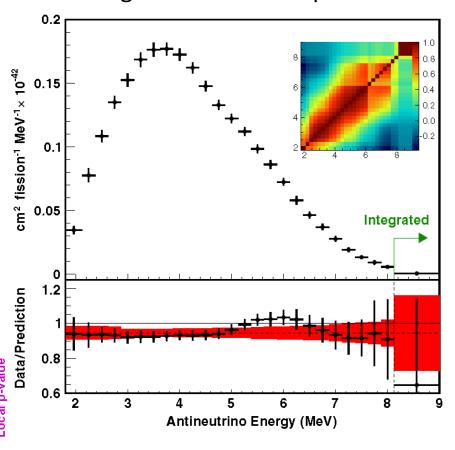


#### Reactor antineutrino spectrum

- Absolute positron spectral shape is NOT consistent with the prediction. A bump is observed in 4-6 MeV (4.4  $\sigma$ ).



 Extract a generic observable reactor antineutrino spectrum by removing the detector response



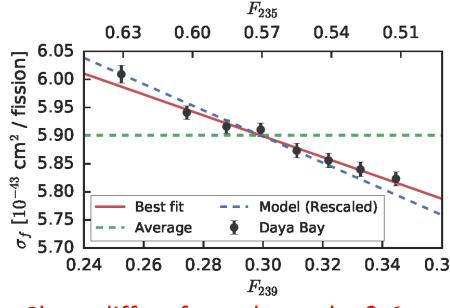
#### Reactor antineutrino flux evolution



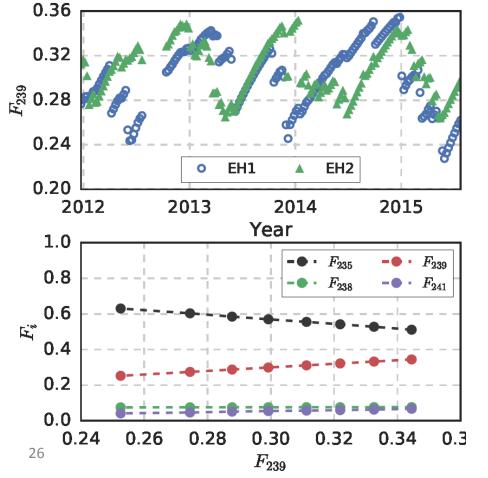
arXiv:1704.01082v1, PRL 2017.

Effective fission fraction for 
$$i^{\text{th}}$$
 isotope  $F_i(t) = \sum_{r=1}^6 \frac{W_{\text{th},r}(t) \bar{p}_r f_{i,r}(t)}{L_r^2 \overline{E}_r(t)} \bigg/ \sum_{r=1}^6 \frac{W_{\text{th},r}(t) \overline{p}_r}{L_r^2 \overline{E}_r(t)}$  changes in time as fuel evolves:

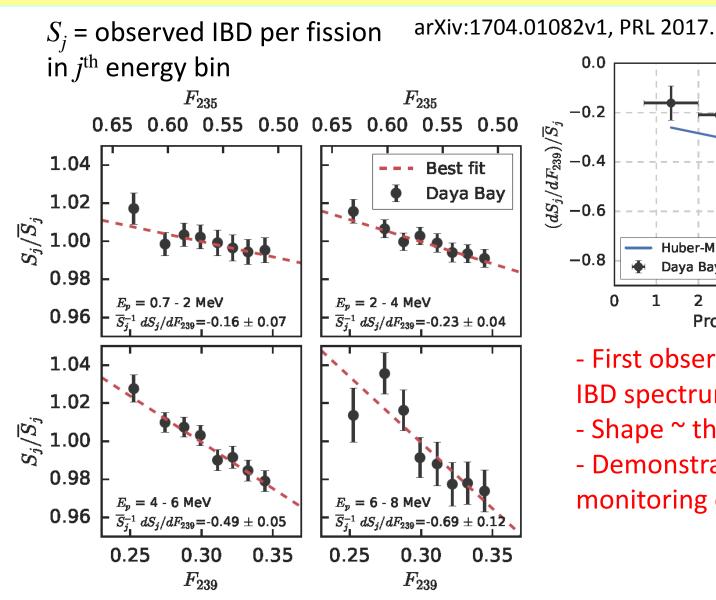
 $\sigma_t(t) = \sum_i \sigma_i F_i(t)$  also evolves IBD yield ith isotope

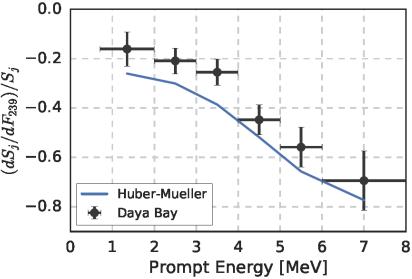


Slope differs from theory by  $2.6\sigma$ Sterile  $\nu$  only incompatible at  $2.6\sigma$ favor: overestimation of <sup>235</sup>U yield



#### Reactor antineutrino spectrum evolution





- First observation of change in IBD spectrum with  $F_{239}$  at 5.1 $\sigma$
- Shape ~ theory
- Demonstration of neutrino monitoring of reactors



# Search for a light sterile neutrino

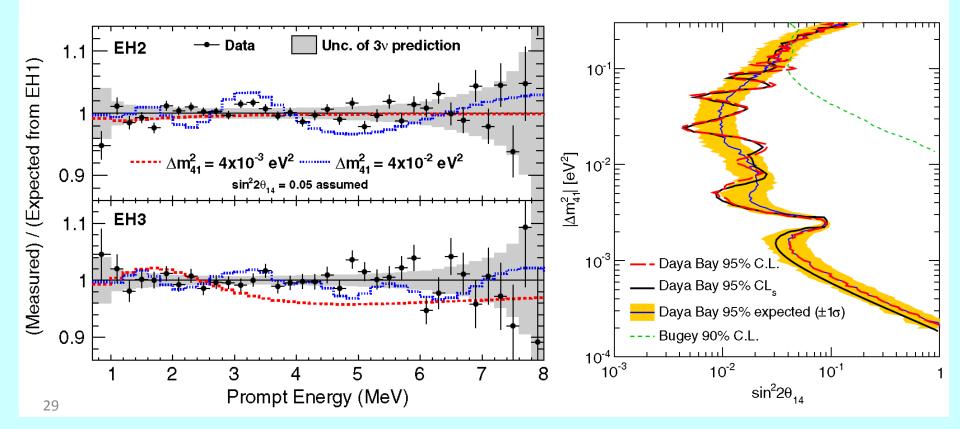
F. P. An et al., Daya Bay Collaboration, PRL **117**, 151802 (2016); PRL **113**, 141802 (2014).

Daya Bay and MINOS Collaborations, PRL 117, 151801 (2016).

# Search for a light sterile neutrino

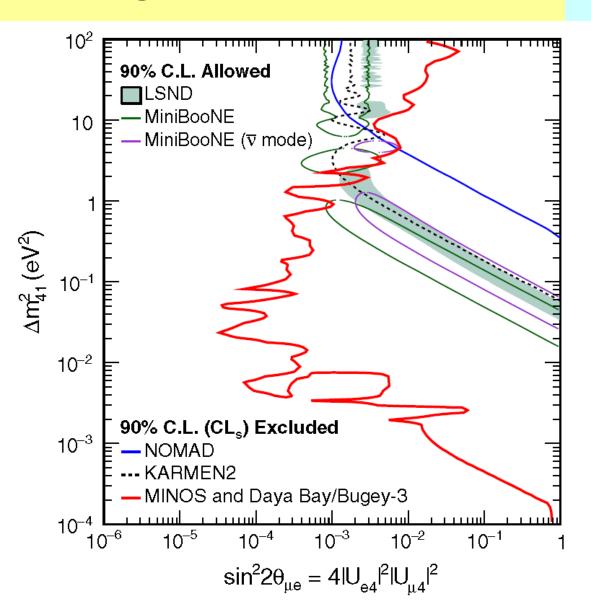


- Sterile neutrino: additional oscillation mode  $\theta_{14}$
- 3 expt. halls → multiple baselines
  - Relative measurement at EH1 (~350m), EH2 (~500m), EH3 (~1600m)
  - Unique sensitivity at  $10^{-3} \text{ eV}^2 < \Delta m_{41}^2 < 0.1 \text{ eV}^2$
- most stringent limit on  $\sin^2 2\theta_{14}$  for  $2 \times 10^{-4}$  eV<sup>2</sup>  $< \Delta m_{41}^2 < 0.2$  eV<sup>2</sup>



# Search for a light sterile neutrino

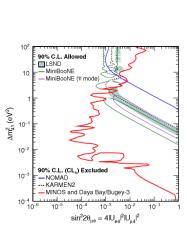
- Combined constraints on  $\sin^2 2\theta_{14}$  from  $\overline{\nu}_e$  disappearance in Daya Bay and Bugey with constraints on  $\sin^2 2\theta_{24}$  from  $\overline{\nu}_\mu$  disappearance in MINOS
- Set constraints over 6 orders of magnitude in  $\Delta m^2_{41}$  . Strongest constraint to date
- Exclude parameter space allowed by MiniBooNE and LSND for  $\Delta m_{41}^2 < 0.8 \ {\rm eV^2}$

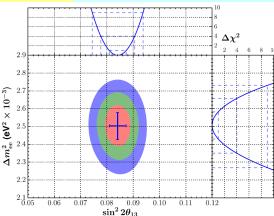


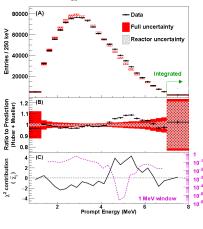


#### Summary

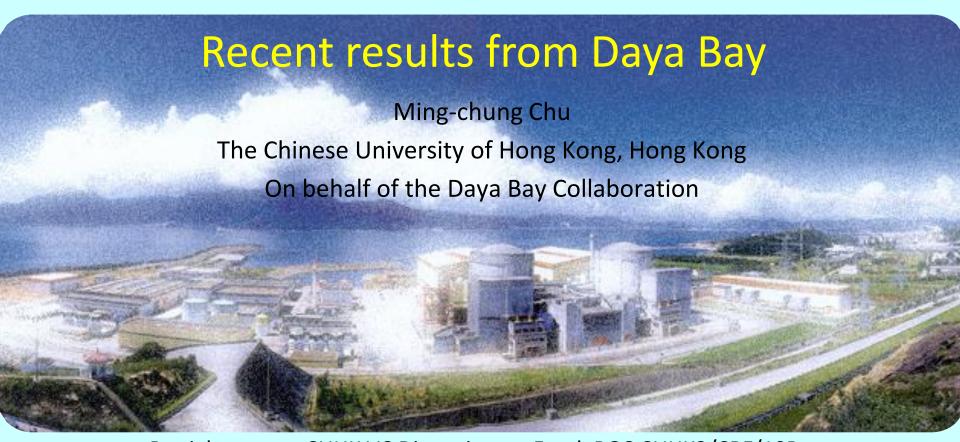
- Daya Bay 1230 days of data, > 2.5M IBD events
  - Most precision measurement of  $\sin^2 2\theta_{13}$ : 3.9%
  - Most precision measurement of  $\left|\Delta m_{ee}^2\right|$ : 3.4%
  - Oscillation results confirmed with independent nH rate measurement (621 days)
- reactor antineutrino flux and spectrum
  - Flux: consistent with previous short baseline experiments, but not with theoretical prediction
  - Spectrum: 4.4 $\sigma$  deviation from prediction in [4, 6] MeV e<sup>+</sup> energy
  - **Evolution observed**. Favors  $\sigma_{235}$  wrong.
- Set new limit to light sterile neutrinos
- Will continue till 2020











Partial support: CUHK VC Discretionary Fund, RGC CUHK3/CRF/10R

29<sup>th</sup> Recontres de Blois

May 28 – June 03, 2017, Blois, France

# backup



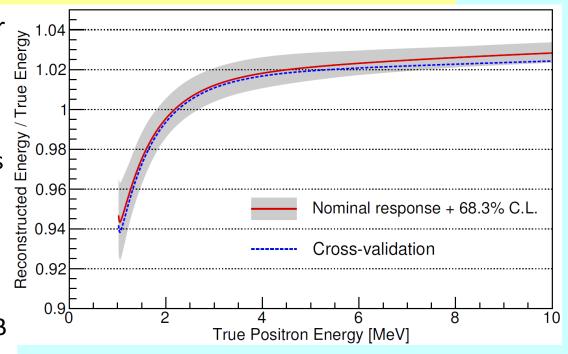
#### More searches

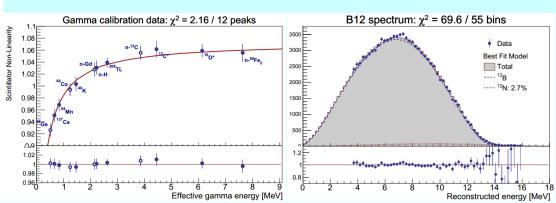
- Precision measurement of spectral distortion:
  - neutrino decoherence
  - sterile neutrino mixing
  - CPT violation/NSI
  - mass-varying neutrinos
- Precision measurement of neutrino rate:
  - sidereal modulation (CPT violation, ...)
  - supernova neutrinos
- High energy events:
  - neutron-anti-neutron oscillation





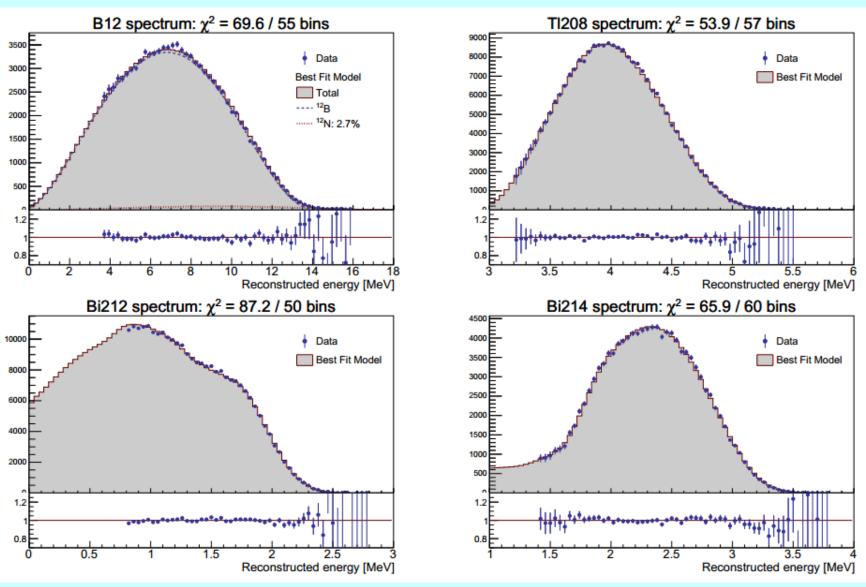
- Particle-dependent scintillator nonlinearity: modeled with Birks' law and Cherenkov fraction
- Charge-dependent electronics nonlinearity: modeled with MC and single channel FADC measurement
- Nominal model: fit to monoenergetic gamma lines and <sup>12</sup>B beta-decay spectrum
- Cross-validation model: fit to <sup>208</sup>Th, <sup>212</sup>Bi, <sup>214</sup>Bi beta-decay spectrum, Michel electron
- Uncertainty < 1% above 2 MeV</li>





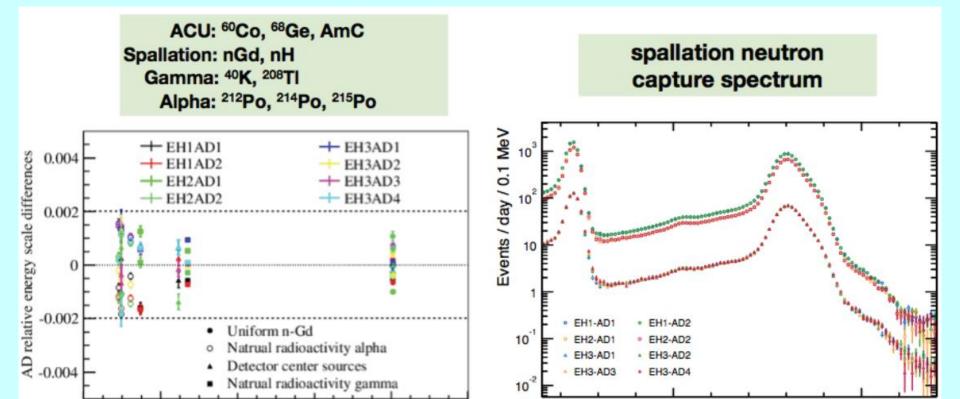
### Detector energy response model





# AD Calibration





Less than 0.2% variation in reconstructed energy between detectors.

Reconstructed energy (MeV)

Reconstructed Energy (MeV)



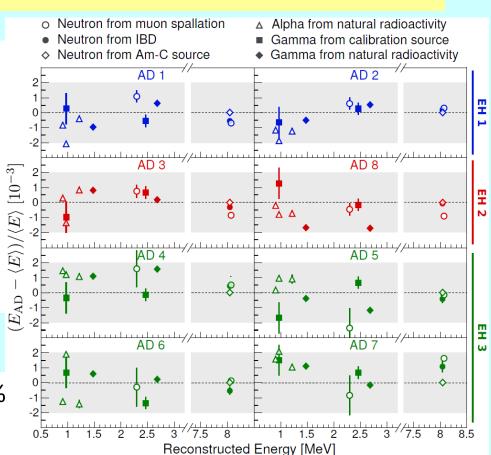
### **Energy calibration**

- PMT gain: Single electrons from photocathode
- Absolute energy scale: AmC at AD center
- Time variation: <sup>60</sup>Co at AD center
- Non-uniformity: <sup>60</sup>Co at different positions
- Alternative calibration: spallation neutrons

- Relative energy scale uncertainty: 0.2%
  - <sup>68</sup>Ge, <sup>60</sup>Co, AmC: detector center

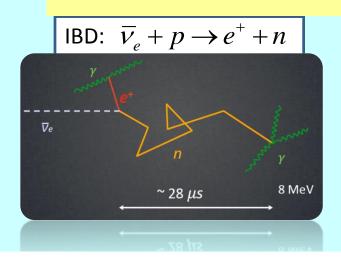
40K 208TI nH. 1m vartay cut

- nGd from IBD and muon spallation: Gd-LS region
- $\alpha$  from polonium decay: Gd-LS vertex cut

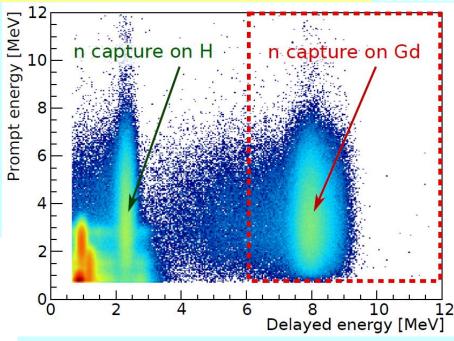


#### Antineutrino candidates selection





- Reject PMT flashers
- **Coincidence** in energy and time with multiplicity = 2
  - **Energy**:  $0.7 \text{ MeV} < E_p < 12.0 \text{ MeV}$ ,  $6.0 \text{ MeV} < E_d < 12.0 \text{ MeV}$
  - Time:  $1 \mu s < \Delta t_{p-d} < 200 \mu s$
- Muon anticoincidence
  - Water pool muon: reject 0.6 ms
  - AD muon (>20 MeV): reject 1 ms
  - AD shower muon (>2.5 GeV): reject 1 s



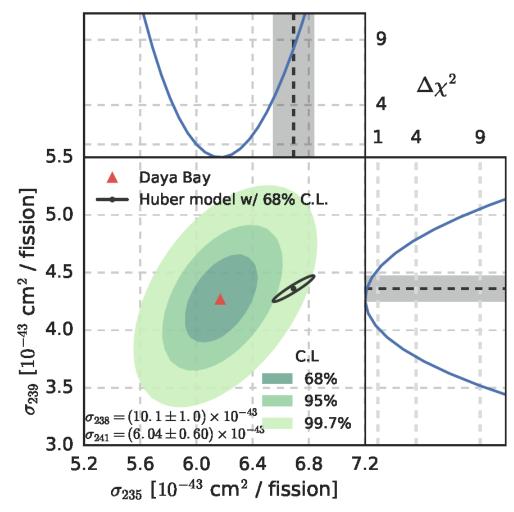
	efficiency	correlated	uncorrelated
target protons		0.47%	0.03%
flasher cut	99.98%	0.01%	0.01%
delayed energy cut	90.9%	0.6%	0.12%
prompt energy cut	99.88%	0.10%	0.01%
multiplicity cut		0.02%	< 0.01%
capture time cut	98.6%	0.12%	0.01%
Gd capture fraction	83.8%	0.8%	< 0.1%
spill-in	105.0%	1.5%	0.02%
livetime	100.0%	0.002%	< 0.01%
combined	78.8%	1.9%	0.2%

# Reactor antineutrino flux and spectrum evolution 13

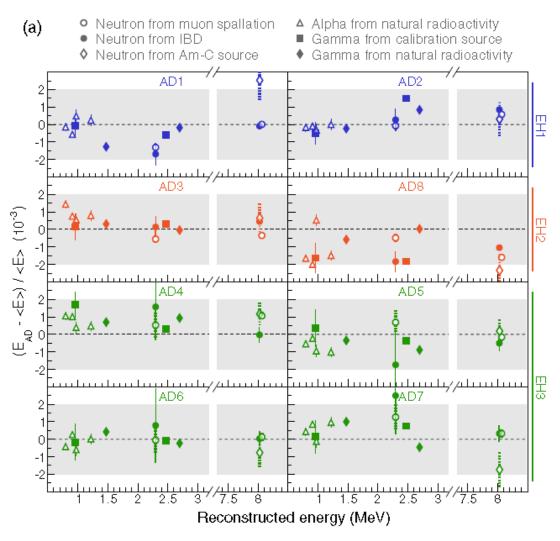


arXiv:1704.01082v1, PRL 2017.

Likely due to overestimation of <sup>235</sup>U yield

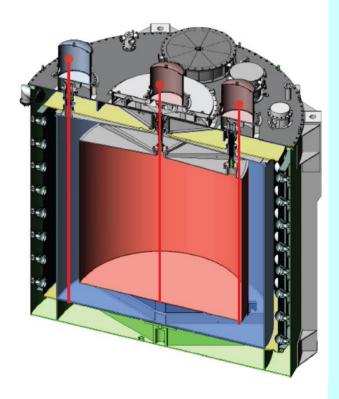


#### **Detector calibration**



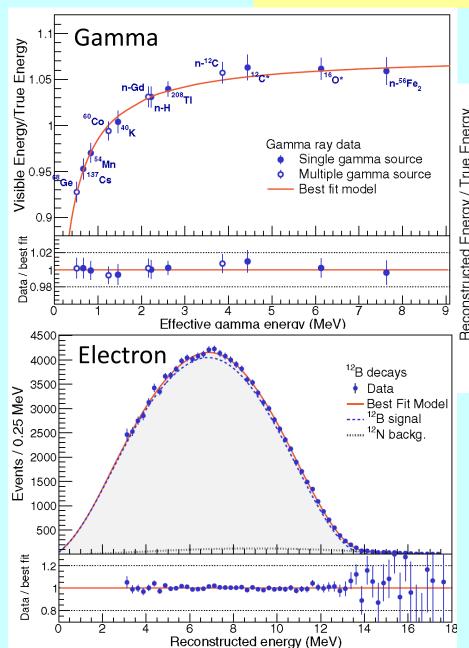
Calibration using <sup>68</sup>Ge (1.02MeV), <sup>60</sup>Co (2.5MeV), <sup>241</sup>Am-<sup>13</sup>C (8MeV), LED, spallation neutrons

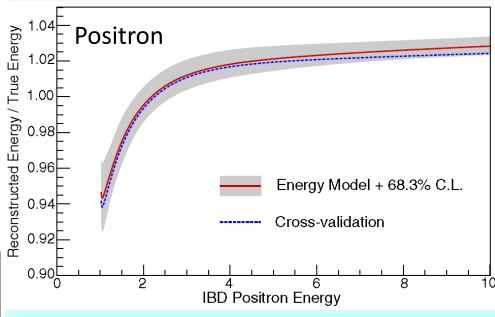
ACU-C ACU-A ACU-B



Relative energy scale uncertainty < 0.2%

# **Energy non-linearity**





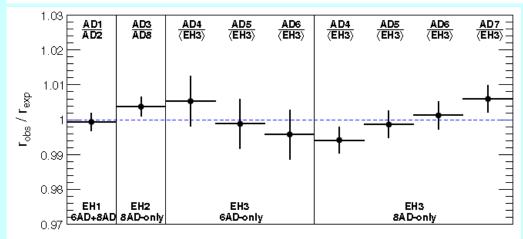
- Measured  $\gamma$  and e responses
- Derive  $e^+$  energy model from  $\gamma$  and e responses using simulation Uncertainty ~ 1% (correlated among detectors)

#### **Systematics**

#### **Detector efficiency**

	Efficiency	Correlated	Uncorrelated
Target protons	-	0.92%	0.03%
Flasher cut	99.98%	0.01%	0.01%
Delayed energy cut	92.7%	0.97%	0.08%
Prompt energy cut	99.8%	0.10%	0.01%
Multiplicity cut		0.02%	0.01%
Capture time cut	98.7%	0.12%	0.01%
Gd capture fraction	84.2%	0.95%	0.10%
Spill in	104.9%	1.00%	0.02%
Live time	-	0.002%	0.01%
Combined	80.6%	1.93%	0.13%

Correlated uncertainties cancelled out in relative measurement



Uncorrelated uncertainties cross-checked by multiple detectors in the same hall