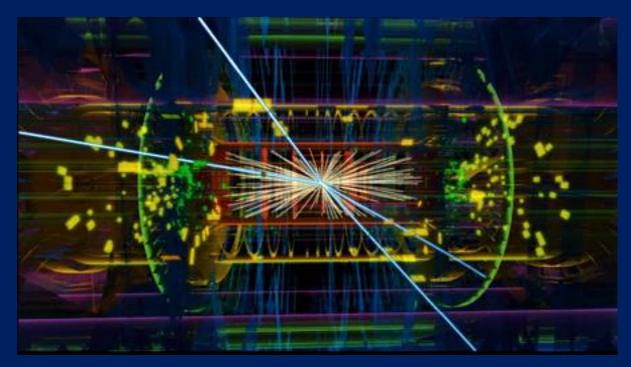
# **ASKING THE BIG QUESTIONS:** A JOURNEY INTO PARTICLE PHYSICS

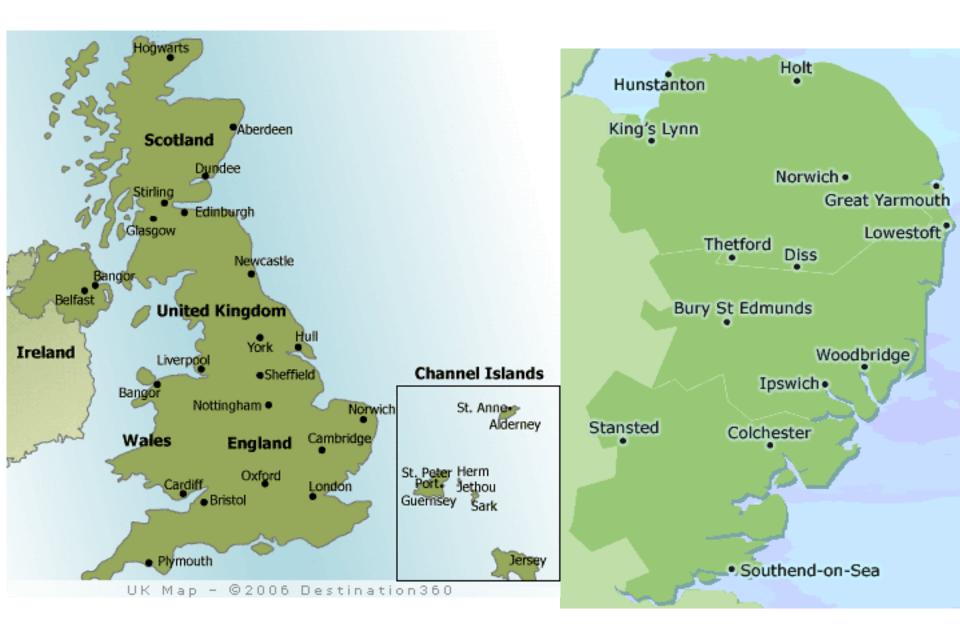
**Kate Shaw** 

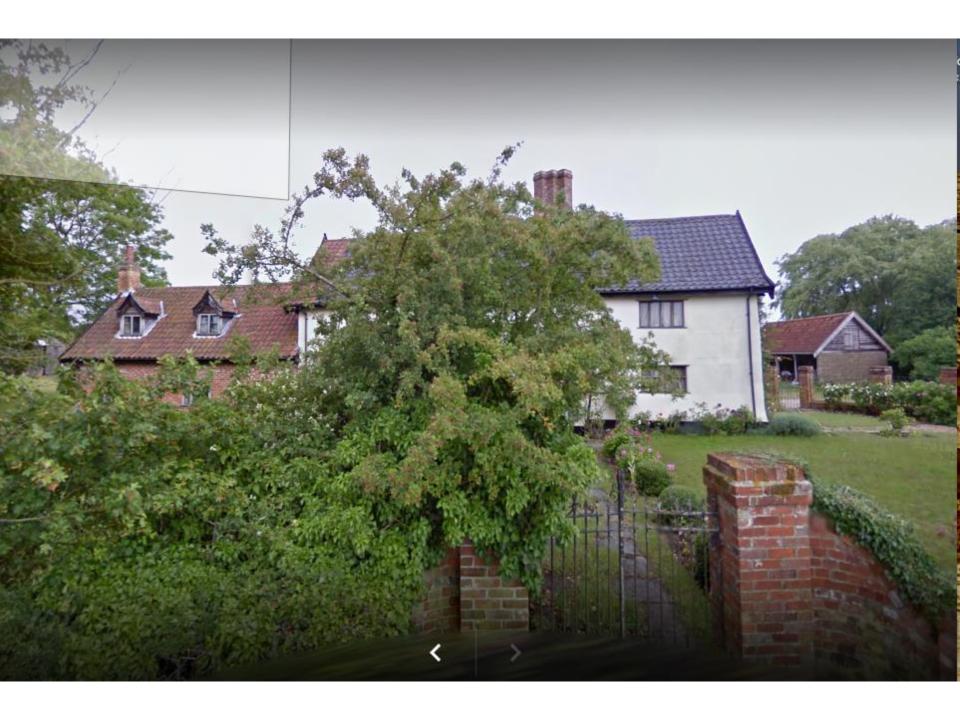
The Abdus Salam International Centre for Theoretical Physics ATLAS Collaboration



#### 7<sup>th</sup> August 2017 N International Teacher Weeks Programme 2017









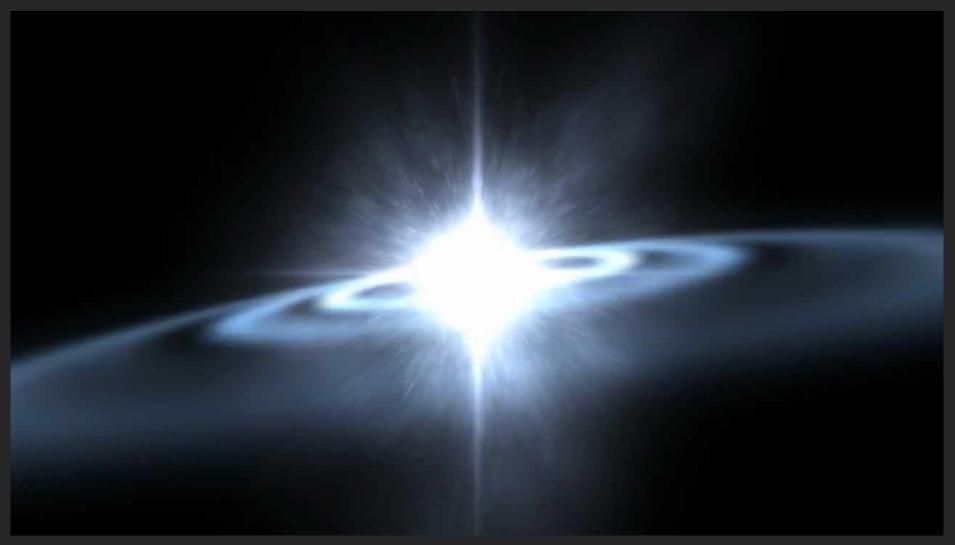


The observable Universe has around 350 billion large galaxies

It is 90 billion light years across

Particle Physicists aim to understand what the universe is made out of, how everything interacts and where everything came from!

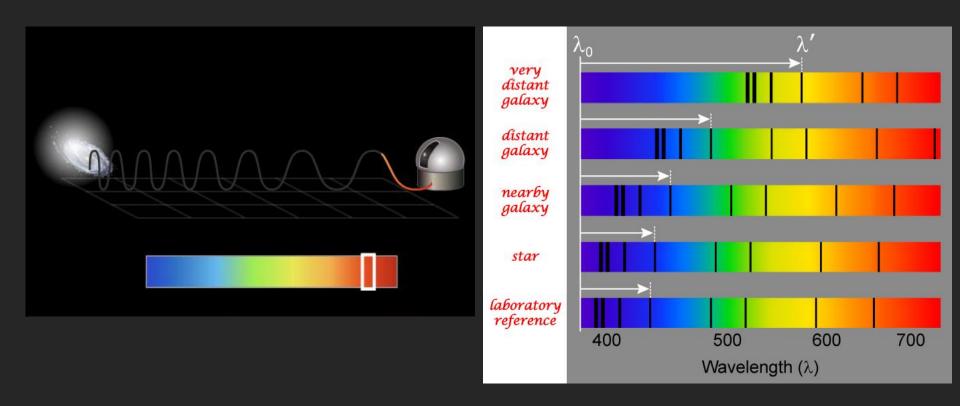
Physicists are ambitious!



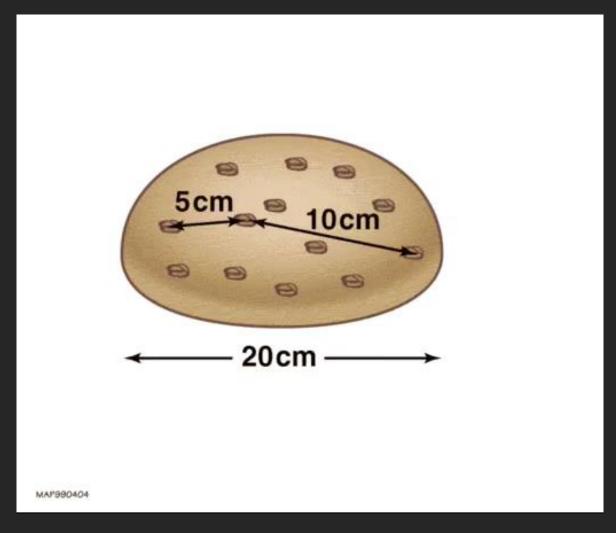
BIG BANG 13.8 Billion years ago



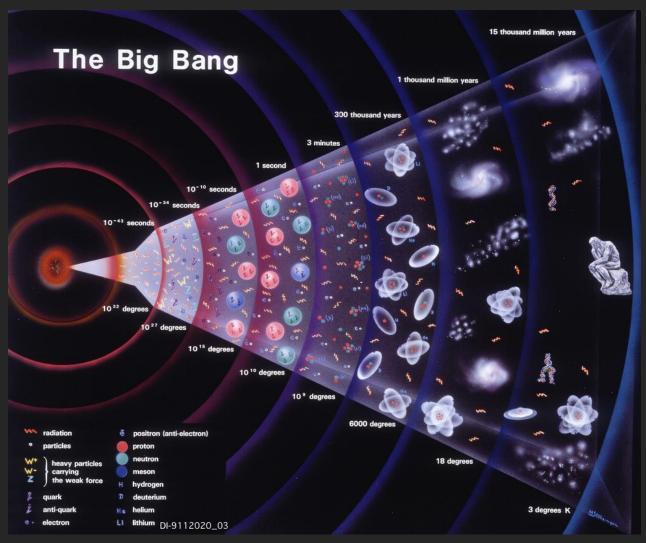
Edwin Hubble, American Astronomer, first to prove there are galaxies outside of the Milky Way (1924)



Hubble showed that the further away a galaxy is from us (or any point in space), the **faster it appears to move** (1929)



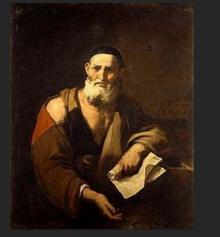
Space is stretching at a constant rate! Hubble constant: 42 miles/second/3 million light year steps



BIG BANG 13.8 Billion years ago



Democritus , Ancient Greek philosopher, ~5 BC



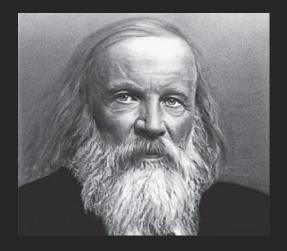
Leucippus , Ancient Greek philosopher, ~5 BC

"The universe is composed of two elements: the atoms and the void in which they exist and move."

'All matter consists of invisible particles called 'atoms'

The word *atom* comes from *atomos*, an ancient Greek word meaning *indivisible* 

'Atoms are indestructible, solid but invisible, homogenous, and differ in size, shape, mass, position and arrangement'



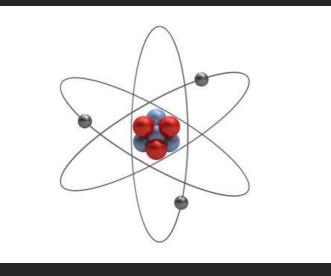
In 1869 Russian chemist Dimitri Mendeleev published the periodic table, arranging chemical elements by atomic mass.

~ 80 elements – WHY SO MANY? Is there a simpler structure?

сшій окисель		Группа II. <b>R<sup>2</sup>O<sup>2</sup> яля RO</b>	Группа III. R <sup>2</sup> O <sup>3</sup>	Группа IV. <b>R<sup>2</sup>O<sup>4</sup> вля RO<sup>2</sup></b>	Группа V. <b>R</b> <sup>2</sup> <b>O</b> <sup>5</sup>	Группа VI. <b>R<sup>2</sup>O<sup>6</sup> или RO</b> <sup>3</sup>	Группа VII. <b>R</b> <sup>2</sup> <b>O</b> <sup>7</sup>		па VIII. (перехол или RO <sup>4</sup>	ъкъ I)			11	Rb Cs	Li
азующій соли	$\begin{array}{c} \mathbf{H} = \mathbf{I} \\ \mathbf{H}^{2}0, \mathbf{H}\mathbf{H}, \mathbf{H}\mathbf{C}\mathbf{I}, \end{array}$	100		RH <sup>4</sup>	RH <sup>3</sup>	$RH^2$	RH	02			HX SO			Ca	Ве
Типичес. (	HªNAHªCARÓH.		D 11	C 10					Тало твердое, Тало газообра	SHOC BLR JCTVS	tee.		- Er?	Dip	H
	$\frac{\text{Li}=7}{\text{ICI,LiOH,Li}^{2}\text{O}}$	Be=9,4 BeCl <sup>‡</sup> BeO <sub>*</sub>		CH+C"H=n-+-?	NHANH4CLN20	0=16 онно 20,0707	FH,BF#SiFt		I=K, Ag Mª K=Cl,ONO;OH	=Ca. Pb				PZr	
Раль І.	Na=23	Mg=24	A1=27.3	Si=28	P=31	ом;о <sup>*</sup> R, нок. S=32	CaF <sup>*</sup> <sub>*</sub> KF,KHF <sup>*</sup> . Cl=35,	,5 Ge	2		50				0
. Jana	NaCl,NaHO,Na <sup>2</sup> C Na <sup>2</sup> SO <sup>4</sup> Na <sup>2</sup> CO <sup>3</sup>	MgSO <sup>4</sup> MgNH <sup>4</sup> PO <sup>4</sup>	A12Cl5A12O5 KA1S2O*I2H2C	KAISi <sup>3</sup> O <sup>*</sup> <sub>2</sub> SiO <sup>*</sup> <sub>2</sub>	* PH <sup>‡</sup> PCl <sup>‡</sup> PCl <sup>*</sup> P <sup>2</sup> O <sup>3</sup> <sub>2</sub> P <sup>2</sup> O <sup>4</sup> , Ca <sup>3</sup> P	OI SOISU'X Ba'S	Of CIOH, CIO+H, A	gCl.					Ta –	Np 1	N
(Ряль 2.	KCLKOH.K <sup>2</sup> 0	Ca=40 CaSO1CaOnSiO1	?44=Eb?	Ti=48(\$0?)	V=51 VOCIAV#01VO1	Cr=52 CrClaCrCl5Cr <sup>3</sup> O <sup>3</sup>	MnK*O*MnKO*	Fe=56 FeKiOfFeSt	Co=59 CoXICoXI	Ni=59	Cu=63		R D	Mo	0
	NO <sup>3</sup> K <sup>2</sup> PtCl <sup>4</sup> K <sup>2</sup> SiF <sup>4</sup> Cu=63	$CaCl;CaO_*CaCO_*$ Zn=65		FeTi03Ti0S04	$Pb^{a}V^{a}O_{a}^{a}VO_{a}$ Ci As=75	o';K*Cr0;Cr0*Cl1 N Se=78	InCl <sup>3</sup> MnO <sub>*</sub> MnO <sup>2</sup> Br=80	FeO Fe <sup>±</sup> O <sup>±</sup> FeK <sup>+</sup> Cy <sup>6</sup>	CoX <sup>3</sup> 5NH <sup>3</sup>	NiSO46H*O NiK*Uy*	Cu <sup>*</sup> O, CuO, CuKCy <sup>2</sup>		11	Mb	H
Разь 3,	CuX,CuX 2	ZnCl <sup>‡</sup> ZnO <sub>s</sub> ZnCO ZnSO <sup>‡</sup> ZnEt <sup>‡</sup>	?68=El?	?72=Es? ?II,EsO <sup>1</sup> ?	AsH <sup>2</sup> AsCl <sup>2</sup> As <sup>4</sup> As <sup>4</sup> O <sup>5</sup> As <sup>4</sup> S <sup>3</sup>	0 <sup>4</sup> SeH1SeO1SeO SeM1SeM1O <sup>4</sup>	BrH.BrM.						1 Os	Fe Ru	
р (Рядъ 4,	Rb=85	Sr=87 SrCl1Sr0,SrH <sup>3</sup> 0 <sup>3</sup>	?88==Yt?(92)	Zr=90 ZrCl1Zr01ZrX <sup>4</sup> .	Nb=94	Mo=96 MoCliMoSiMoOd	100		Rh=104 RhCl{RhCl	Pd=106	Ag=108				
	Rb <sup>1</sup> PtCl <sup>*</sup> Ag==108	SrS04SrC04 Cd=112	?Yt*0 <sup>2</sup> YtX <sup>3</sup> ? In=113			M <sup>3</sup> MoO <sup>4</sup> nMoO <sup>3</sup>	(128?) I=127	RuO 2RnCla	Rh*O*RhX* RhK*Cy*	PdH #HO PdI #PClz PdK #V <sup>4</sup>	AgNO <sup>*</sup> AgX AgCl <sub>*</sub> Ag <sup>*</sup> O <sub>*</sub> AgKCy <sup>2</sup>		Ir	Co Rh	
Н. {Радъ 5.	AgX,AgCla		InCl <sup>3</sup> In <sup>±</sup> O <sup>3</sup> <sub>0</sub>	SnCliSnCliSnO. SnX iSnNa <sup>+</sup> O <sup>2</sup>	<ul> <li>SbH:SbCl?Sb?f</li> </ul>	DE TeHTTeCl4T	eO? IH.IAg.IH	103	nun oj	Ture is	AgROJ		Pt -	Ni Pd	
<ul> <li>(Рядъ 6.</li> </ul>	Cs=133	Ba=137 BaCltBaH=03Ba0 ?1		Ce=140(138?) CeCl <sup>2</sup> Ce <sup>3</sup> 0 <sup>2</sup> CeO <sub>2</sub>	142								Au -	Ag	Na
	CsCl,CsOH. $Cs^*PtCl_{\#}^{\circ}$	BaSO BaSO BaSO BaSO BaSO BaSO BaSO BaSO	?La <sup>3</sup> O <sup>3</sup> <sub>*</sub> LaX <sup>3</sup> ?	CeX <sup>3</sup> CeX <sup>3</sup> CeK <sup>*</sup> X <sup>9</sup>	142	146	148	150	151	152	153				
PARE 7.	153	158	160	162	164	166	168						Hg		Mg .
- (Радъ 8.					Ta=182	W=184		Os=193	Ir=198?	Pt=197	Au=197		- 11	In	Al
₩	175	177	?ErOPErX??	180—Di?—La(187) ?DiO‡DiX <sup>4</sup> ?		$\frac{\mathrm{WCl}_{\mathbb{A}}^{*}\mathrm{WCl}_{\mathbb{T}}^{*}\mathrm{WO}_{\mathbb{A}}^{*}}{\mathrm{K}^{*}\mathrm{WO}^{+}\mathrm{n}\mathrm{WO}_{\mathbb{A}}^{*}}$	190	OsCI4OsCI3	IrClair 101	PtCl3PtC X4	AuCl <sup>3</sup> AuCl Au <sup>4</sup> O <sup>3</sup> <sub>#</sub> Au <sup>4</sup> O <sub>8</sub>		Pb	- Sn	Si
Паріотя Рядъ 9.	Au=197 AuX,AuX	$\operatorname{Hg=200}_{\operatorname{HgCl_{\#}HgCl_{\#}Hg}}$	O. TICL TI*O.TI*	Ož PbClzPbO_PbO	1 BiCl3Bi2OEBi2O7	H± 210	212	OsK <sup>4</sup> Cy <sup>6</sup>	IrK <sup>a</sup> Cy <sup>6</sup>	PtK=Cy4	AuKCy <sup>2</sup>		Bi	As	P
а (Ряль 10.		HgO, HgX inHg	O TI*SO;TICI	PbEttPbSO+PbK	"O" BiX BiOX, BiNO"	U = 240	<u></u>			Malen			111	Te	S
a '	220	225		ThCl <sup>‡</sup> ThO <sup>±</sup> ThX <sup>‡</sup> Th(SO <sup>+</sup> ) <sup>3</sup>	235	UCIAUOIUO*X* UOIM*U*OI	245	246	248	249	250	4	111	IB	0

Particle Physics seeks to understand fundamental particle and forces

#### The atom



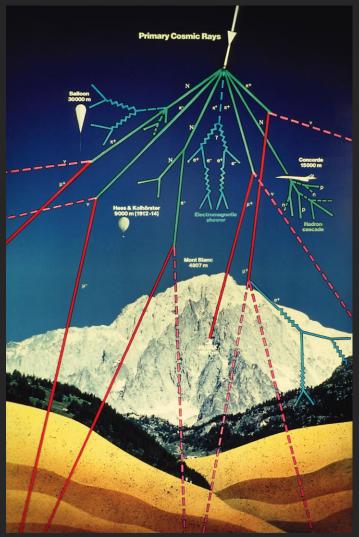
- 1897 electron discovered by J.J. Thompson
- 1911 nucleus of the atom discovered by Ernest Rutherford, and nucleus of hydrogen a proton
- 1932 neutron discovered by James Chadwick

 Photon – particle of the electromagnetic force suggested by Einstein in 1905

However other particles not part of the atom seemed to appear!

- 1932 positron detected (predicted by Dirac in 1928)
- 1934 neutrinos established in theory detected in 1956

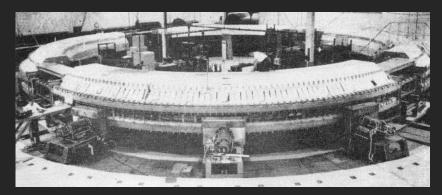
#### Cosmic rays



- 1937 muons were discovered in cosmic rays
   Who ordered that?
- 1947 pions ( a type of meson) were also discovered!!

Things were getting out of hand! What were all these particles?

#### Atom smashers (1950s)

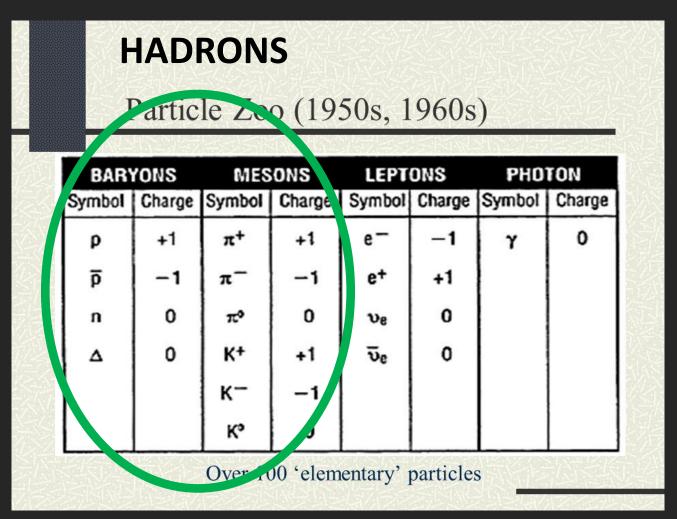


The Brookhaven Cosmotron, at 1.3 – 3.3 GeV accelerator

Plethora of NEW particles started coming out of these atom smashes! Sigma particles, rho particles, Delta particles, kaons, Lambda...

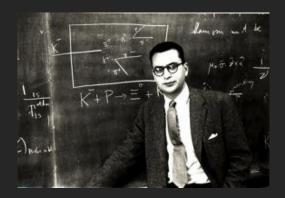


#### Particle Zoo

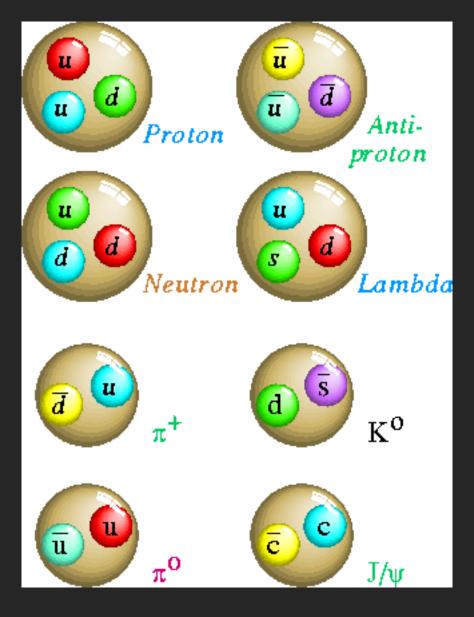


#### **BREAK THROUGH**

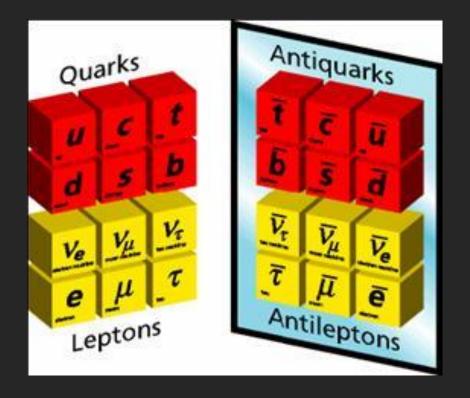
1960s Murray Gell-Mann of Caltech said hadrons are composed of more fundamental particles which he called quarks.



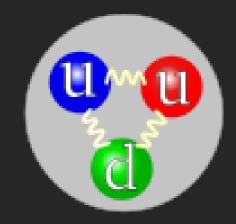
The word comes from a line in Finnegans Wake, a book written by James Joyce.



#### Fundamental matter particles



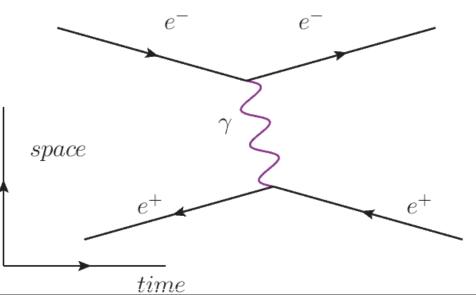
Quarks lived together in twos or threes, making up hadrons like the proton



Fundamental forces described using relativistic quantum field theories

Electromagnetism: Quantum electrodynamics describes how light and matter interact

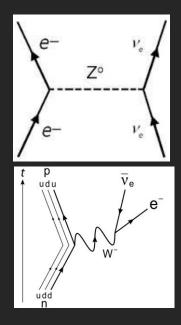
Extremely successful theory! e.g. measurements of the fine structure constant show agreement to within ten parts in a billion (10<sup>-8</sup>)



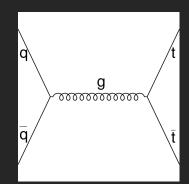
When electrically charged particles interact they exchange a photon

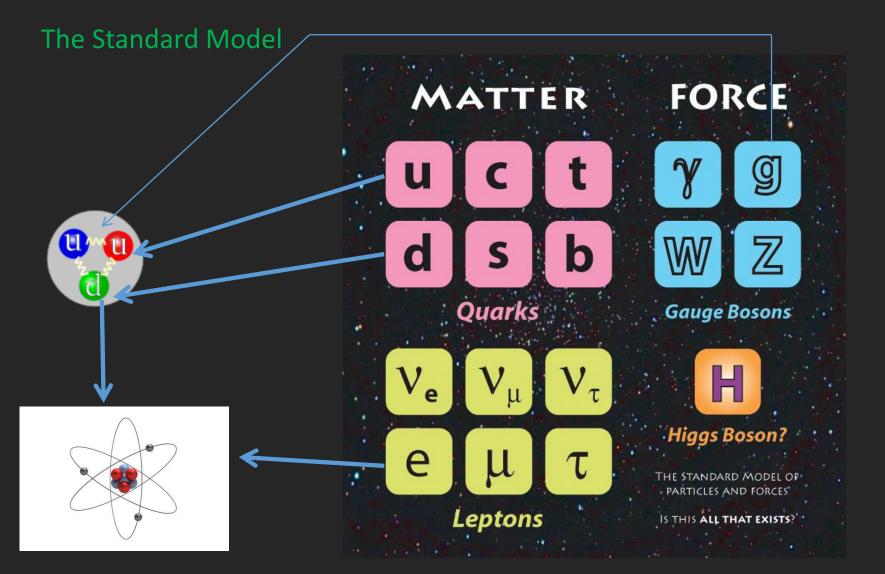
Fundamental forces described using relativistic quantum field theories

A unified theory with electromagnetism was obtained in 1968 – the electroweak theory W and Z predicted and discovered at CERN 1983

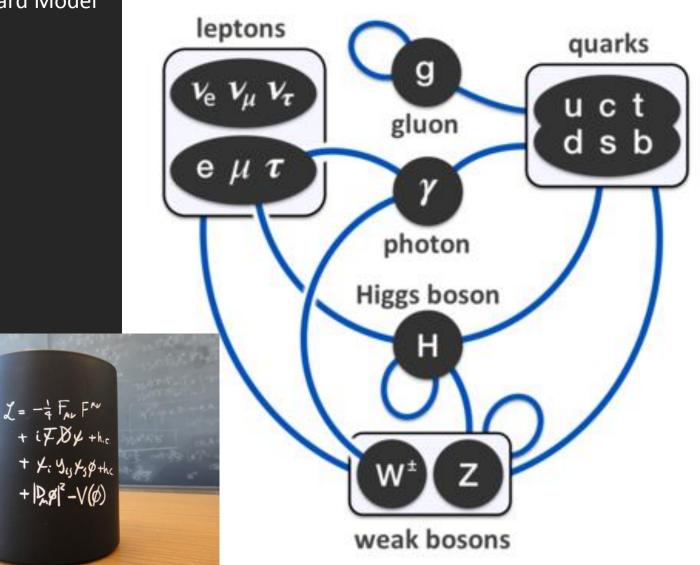


Finally the Strong Force – a very much more complicated force – was described in Quantum Chromodynamics

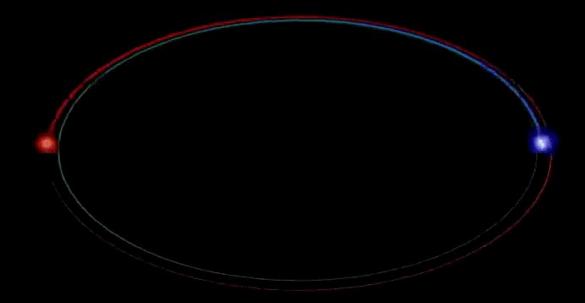




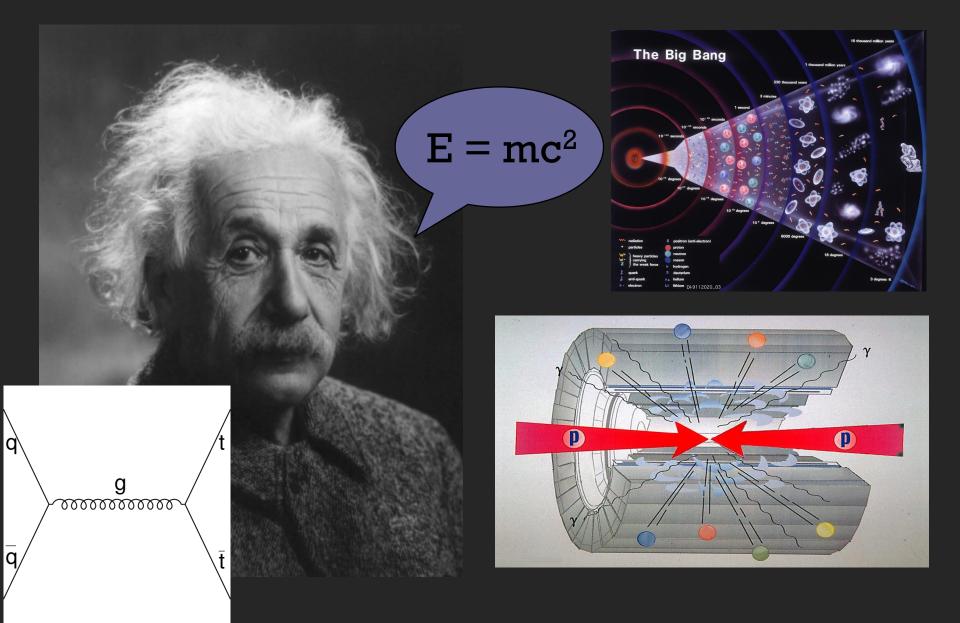
The Standard Model



## **HOW DO WE STUDY PARTICLE PHYSICS?**

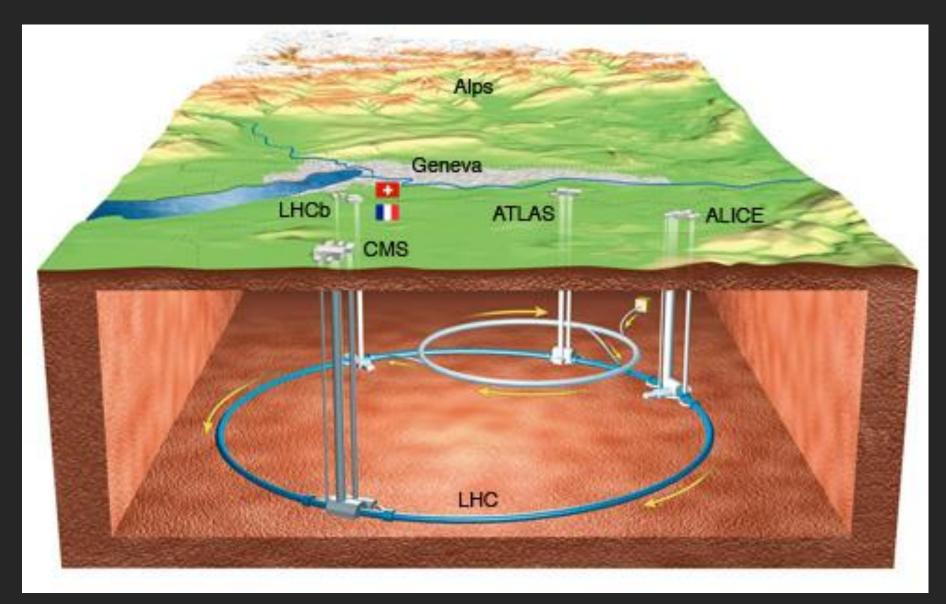


## **HOW DO WE STUDY PARTICLE PHYSICS?**





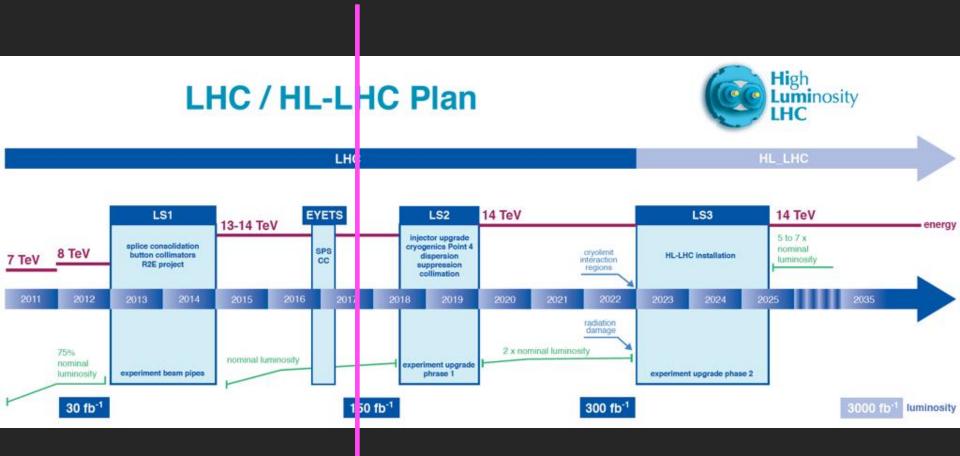


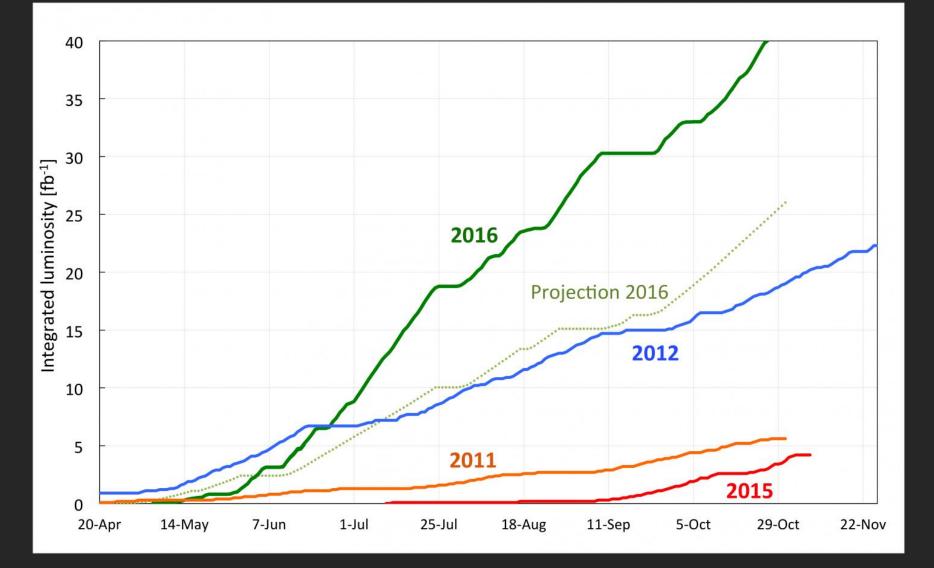


#### THE ATLAS EXPERIMENT



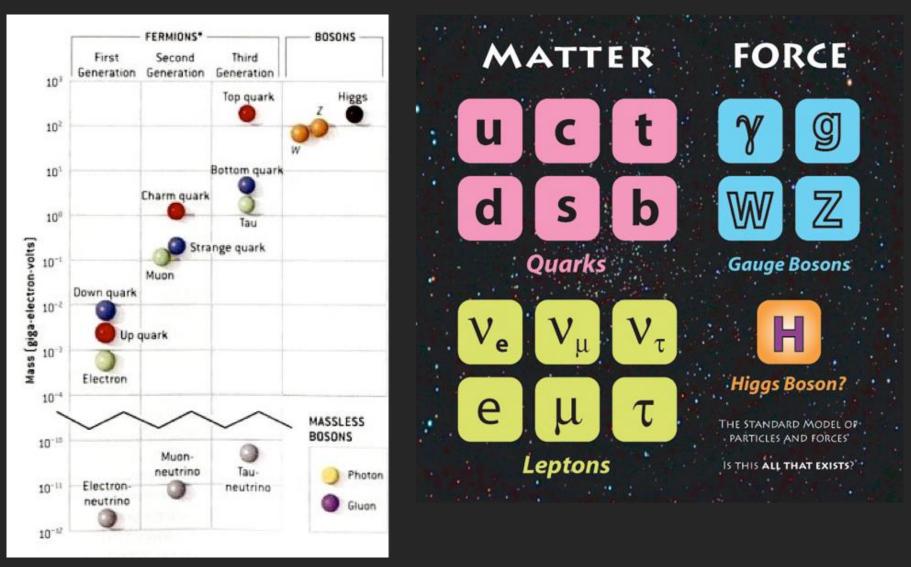
ATLAS comprises 5000 scientists from about 180 institutions around the world, representing 38 countries



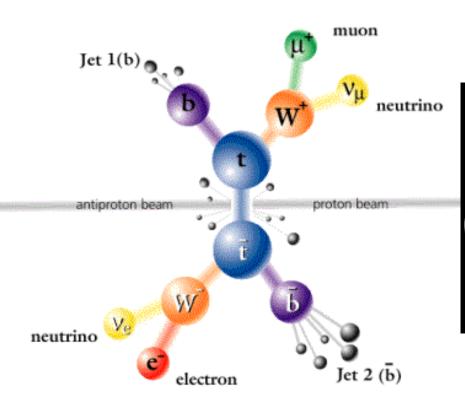


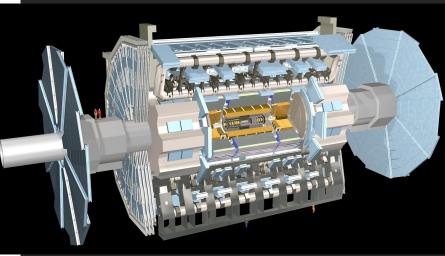
### THE ATLAS EXPERIMENT

#### How do we find new particles?



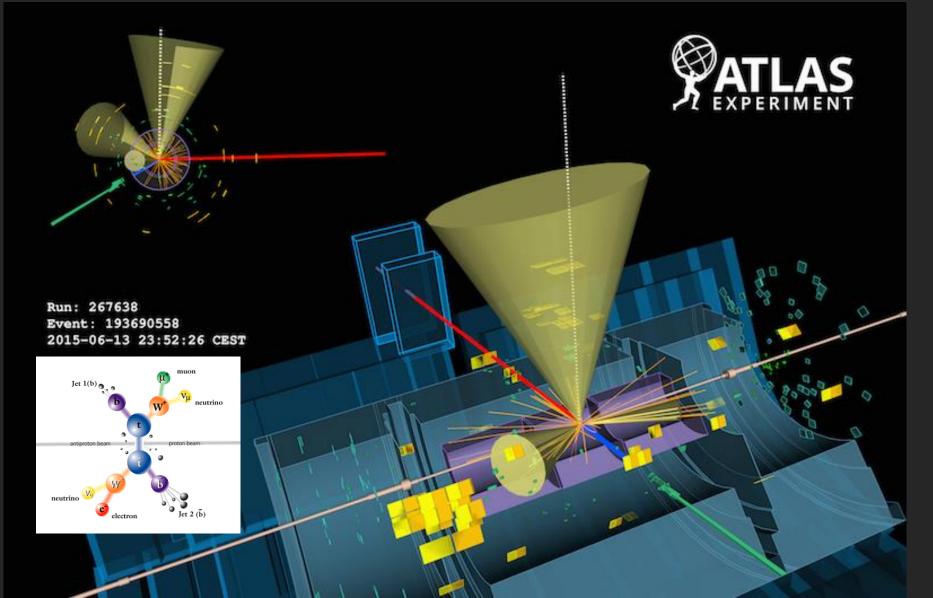
# **THE ATLAS EXPERIMENT** How do we find particles?



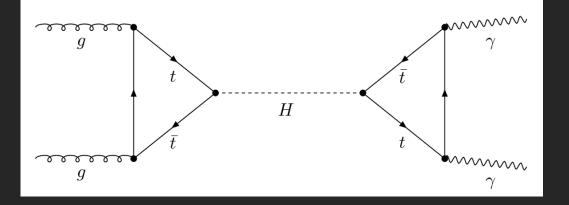


### THE ATLAS EXPERIMENT

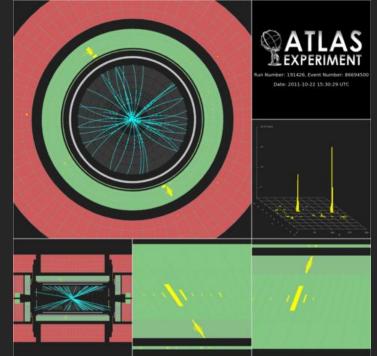
#### How do we find particles?



# **THE ATLAS EXPERIMENT** How did we find the Higgs?

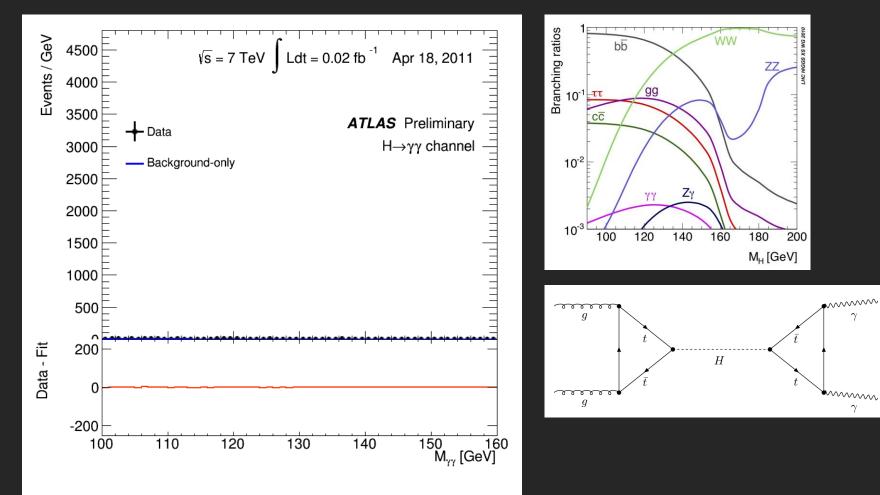


#### Example: Higgs decays into two photos



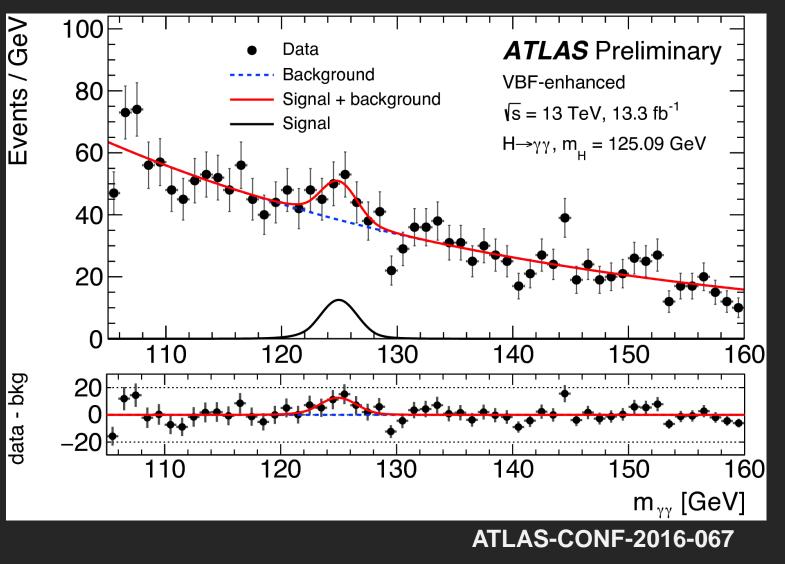
#### THE ATLAS EXPERIMENT

#### How did we find the Higgs?

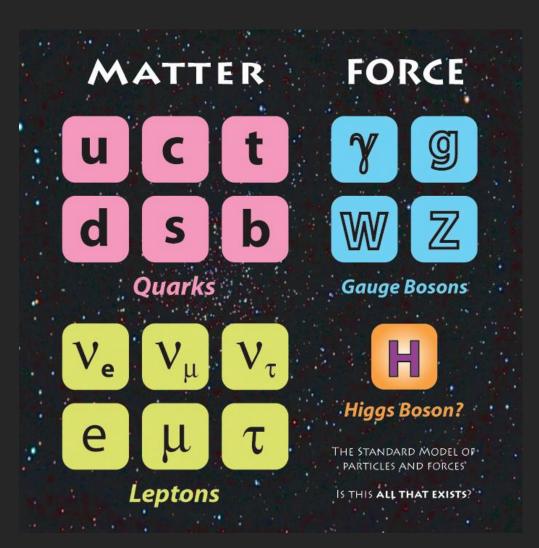


#### THE ATLAS EXPERIMENT

#### How did we find the Higgs?

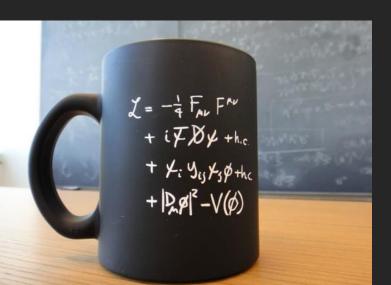


#### **PARTICLE PHSYICS**

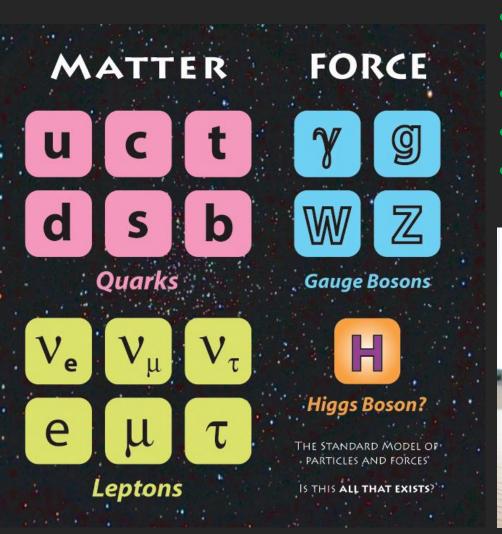


#### GREAT JOB PARTICLE PHYSICIS!!!!





#### **PARTICLE PHSYICS**



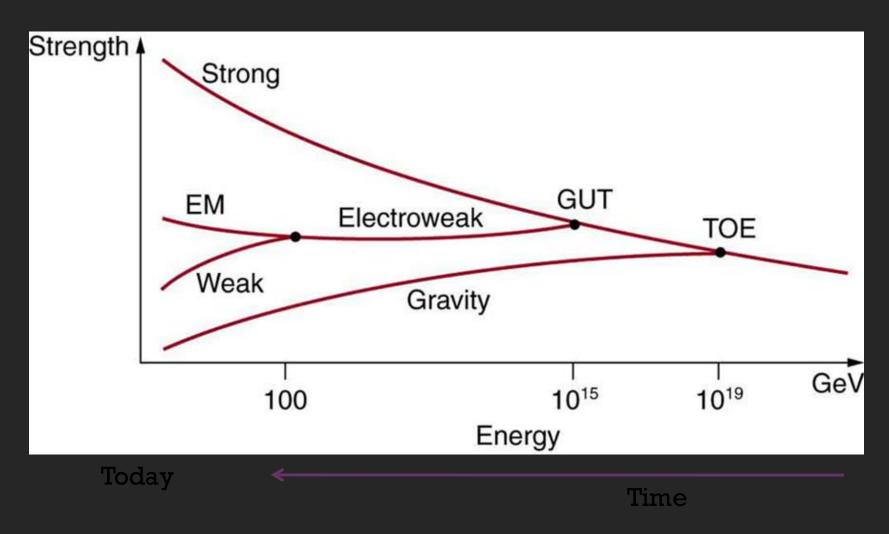
Issues:

- Gravity
- Dark Energy, Dark Matter
- Neutrino masses
- Matter-antimatter asymmetry
- Hierarchy problem

Z= - + FAN FAN + iFDy +h.c. + Ki Yij Kig the  $+ |\underline{D}, \varphi|^2 - \vee (\phi)$ 

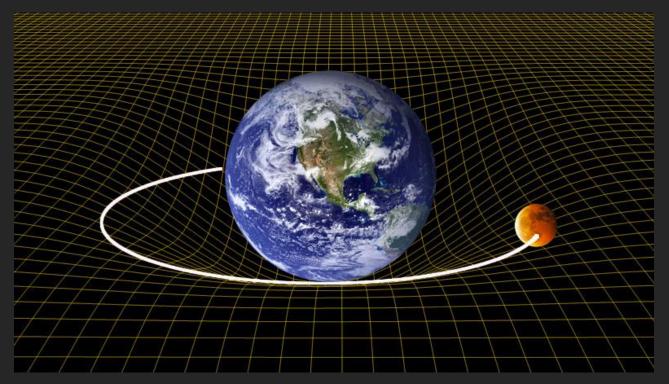
#### **PARTICLE PHYSICS**

Fundamental forces described using relativistic quantum field theories



#### THE PROBLEM WITH GRAVITY

#### Gravity is described by General Relativity! We currently don't have a quantum theory of gravity!



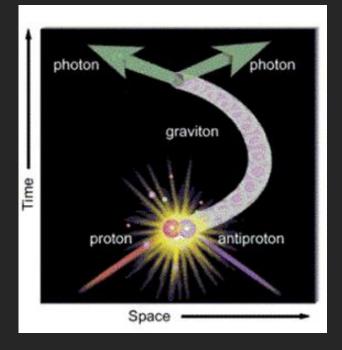
Problem in making a quantum theory of gravity due to something called renormalisation – infinities occur!

Other approaches include string theory and loop quantum gravity

#### THE PROBLEM WITH GRAVITY

Hypothetical force carrier for Gravity: The Graviton

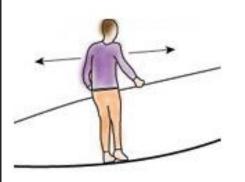
- Graviton: spin 2 massless (tensor) boson
- Cross-section would be very low!
- ATLAS and CMS are still searching for it!



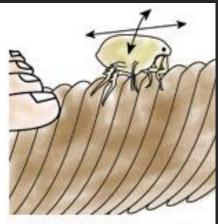
#### **EXTRA DIMENSIONS?**

Einstein's general theory of relativity tells us that space can expand, contract, and bend.

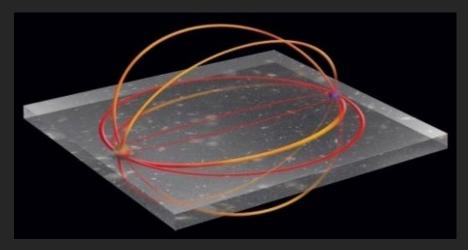
If one dimension were to contract to a size smaller than an atom, it would be hidden from our view.



An acrobat can only move in one dimension along a rope..



...but a flea can move in two dimensions.

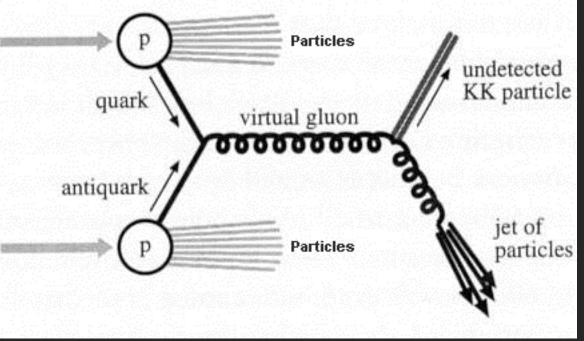


One possibility for gravity being so weak is because part of it spreads to extra dimensions

#### THE GRAVITON & EXTRA DIMENSIONS?

If gravitons exist, it should be possible to create them at the LHC, but they would rapidly disappear into extra dimensions.

#### Example – Graviton produced in association with a jet



The Graviton propagates in the extra dimensions

Only one jet of hadrons would be observed in our four-dimensional world.

We would see of monojet events at colliders.

## **NEUTRINO MASSES**

Neutrinos are the least understood particles in the standard model. They come in 3 flavours, were found to oscillate between these flavours in the 90's and 00's, indicating they have a **non-zero mass!** 

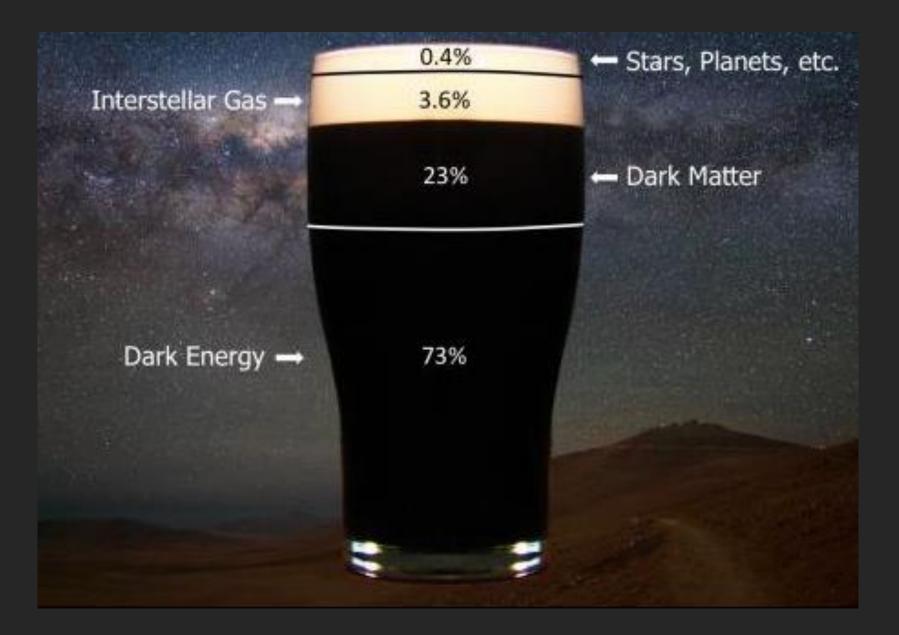
#### This is not predicted by the Standard Model.

Scientists are studying these elusive particles in numerous experiments.



T2K, Japan

#### DARK ENERGY & DARK MATTER



Galaxies moving too fast for the cluster to be bound by only the visible matter of its galaxies

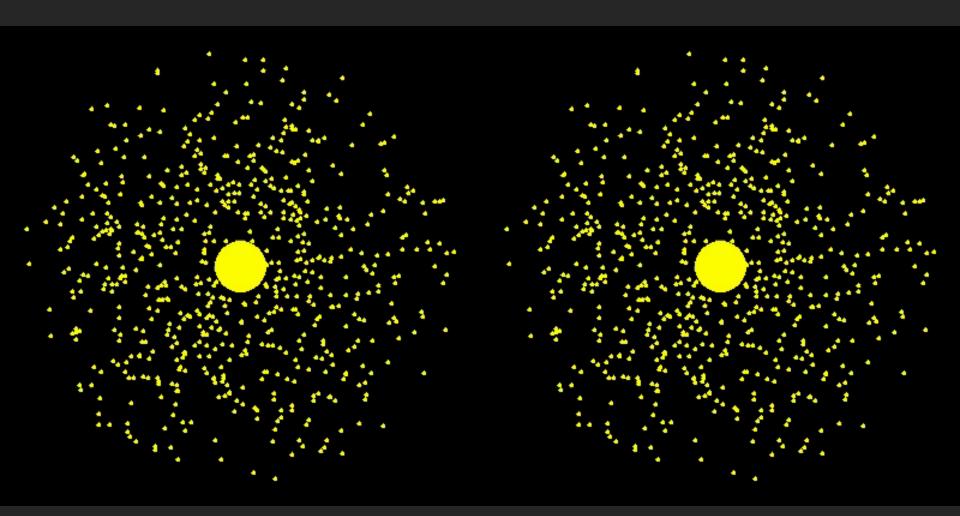
Coma Cluster, contains over 1000 identified galaxies, 321 million light years away from Earth, ~ 90 % dark matter Fritz Zwicky, Swiss Astronomer In 1933 inferred existence of *dunkle Materie* 

Galaxies must be held together by some dunkle Materie



dark matter halo bulge Sun bulge disk Milky Way

1970s Vera Rubin: Investigated rotational curves of spiral galaxies



Galaxies must be held together by some dunkle Materie

By 1980 Dark Matter was widely recognised as an unsolved problem in astronomy

What do we know about dark matter?

- Its dark, it does not absorb, reflect or emit light! This means it does not interact with the electromagnetic force
- It interacts with Gravity



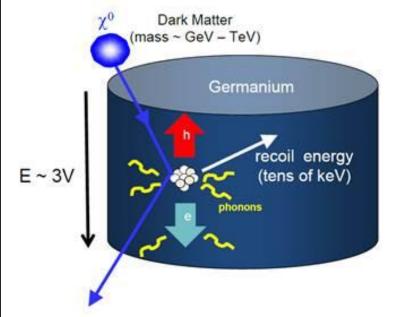
We know what it is *not*:

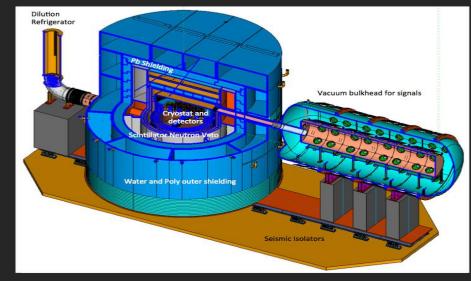
- clouds of normal matter without stars
- antimatter which produces unique gamma rays when it interacts with normal matter
- made of black holes as these are confined objects, dark matter is scattered

A heavy candidate is the WIMP: Weakly Interacting Massive Particles

- If not too massive these should be detectable at the LHC
- A theory named SUPERSYMMETRY which solves the hierarchy problem provides a neutral, heavy, weakly interacting particle, which is an ideal candidate for the WIMP!

#### Light candidates: Axion, Gravitino, light moduli, dilatons





**Cryogenic Dark Matter Search (CDMS)** 

#### BUT – nothing has been seen in direct searches!



Asimina Arvanitaki, Stavros Niarchos Foundation Aristarchus Chair in Theoretical Physics at Perimeter Institute, Canada

- Suggests novel method to search for ultra light Dark Matter
- Using a type of gravitational wave detector (that use differential measurement of two atom interferometer)

Gravitational wave detection



1984

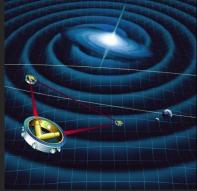
Livingston

#### arXiv:1606.04541



Asimina Arvanitaki, Stavros Niarchos Foundation Aristarchus Chair in Theoretical Physics at Perimeter Institute, Canada

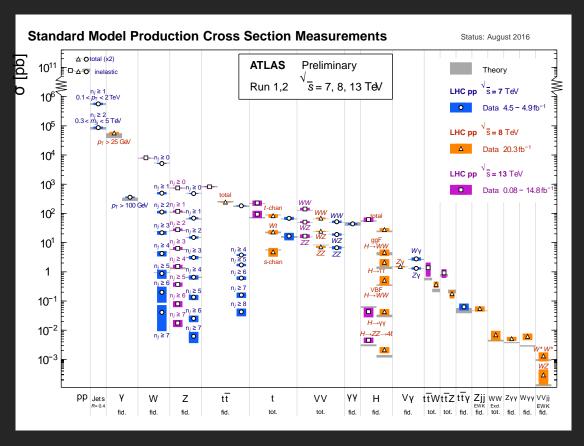
- Packed particles, specifically bosons, in a given space will overlap and begin to behave like a wave
- Ultra light dark matter can cause temporal oscillations in fundamental constants
- Frequency set by DM mass, amplitude set by local DM density



## Looking for cracks in the SM

Testing Standard Model predictions at the LHC

Using the Higgs, the Top and others and a probe for New Physics!



## Looking for cracks in the SM



Dirac famously predicted that, for electrons and muons, the factor relating the magnetic moment to spin would have a value of 2.

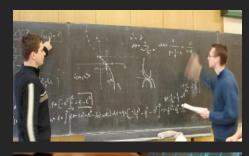
A recent measurement for the muon differs with a significance of  $3.5\sigma$ 

The muon g-2 experiment at Fermilab, US, aims to measure the anomalous magnetic moment of the muon to unprecedented accuracy

UK institutes include Liverpool, UCL, Lancaster, Oxford

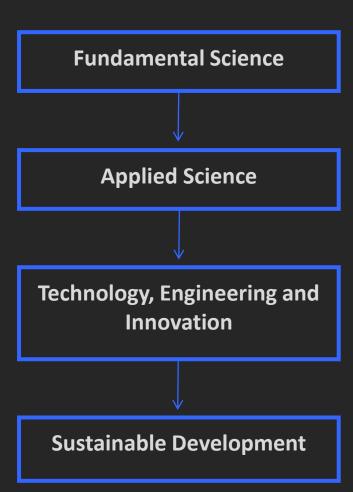
## **Particle Physics for Society**

Because particle physics asked **BIG questions** we need new unique and innovative ideas and equipment which has had a huge impact in society!

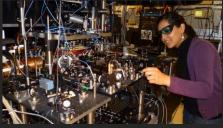












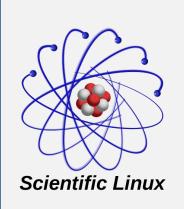


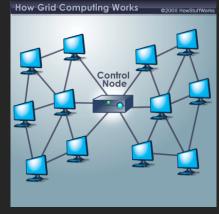
#### Technology



# World Wide Web developed at CERN







**GRID** Computing

**MRI and PET** 

Education – physics/science undergraduates / PhDs





The need to solve environmental and developmental problems requires scientists

Education and investment into educational, technological and cultural institutions play a key role in growing a knowledge-based economy

Higher education, research and innovation fuel national development

**Education** – Particle Physics can inspire young people!



Science teaches young people to think for themselves, ask questions, query assumptions!

Particle physics is an awe-inspiring combination of ambitious research, technical skill and global cooperation, allowing us to continue to ask the BIG **QUESTIONS!** 

Kate.shaw@cern.ch Facebook: https://www.facebook.com/kate.shaw.73744 Twitter: @KateShawOnline

# THANK YOU