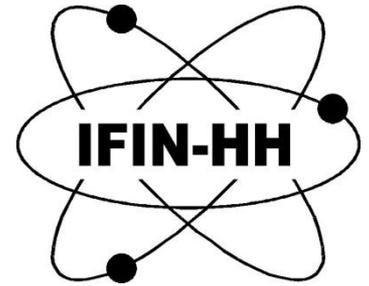




Horia Hulubei National Institute for R&D
in Physics and Nuclear Engineering
(IFIN-HH)



Alexandru Cătălin ENE

Comparison of cosmic ray models with PYTHIA at
energies near the 2nd “knee” of the cosmic ray
spectrum

Workshop on Sensors and High Energy Physics (SHEP 2016),
Suceava “Ștefan cel Mare” University



Suceava, October 2016

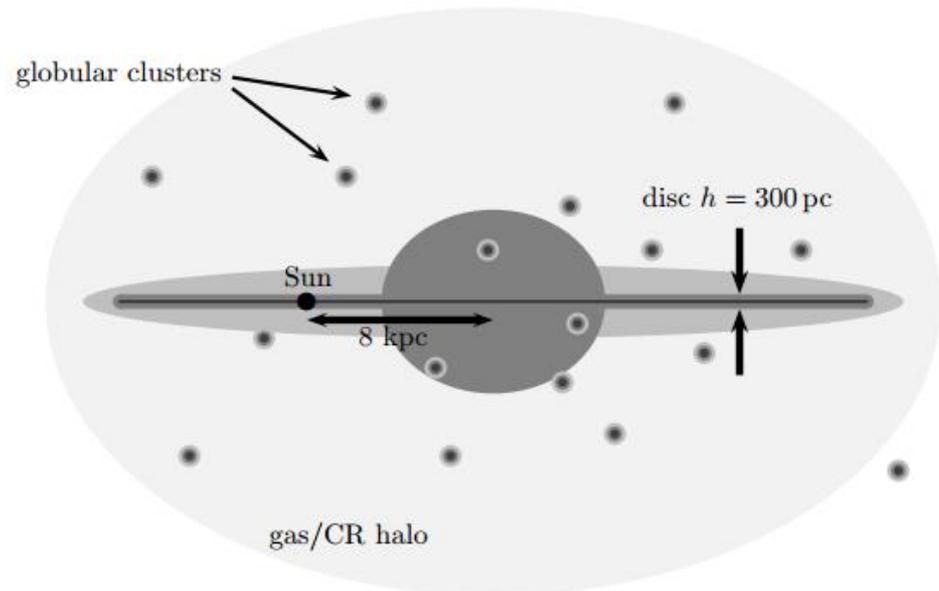


Introduction

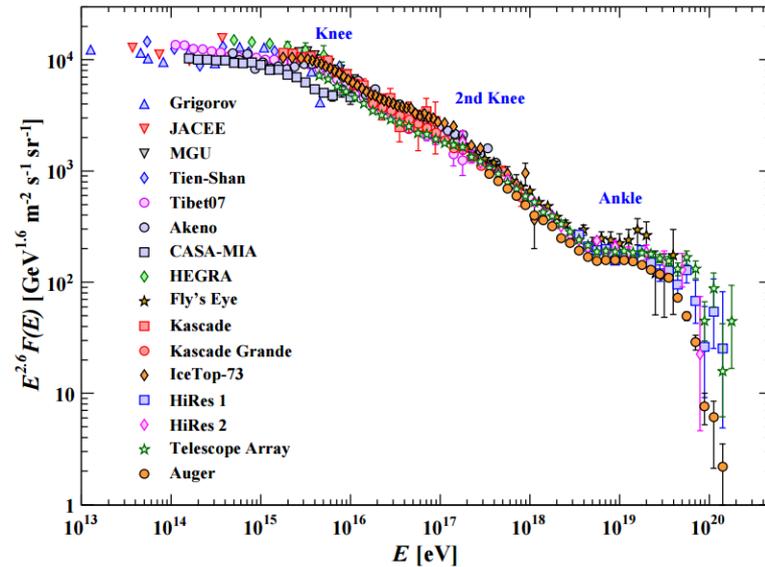
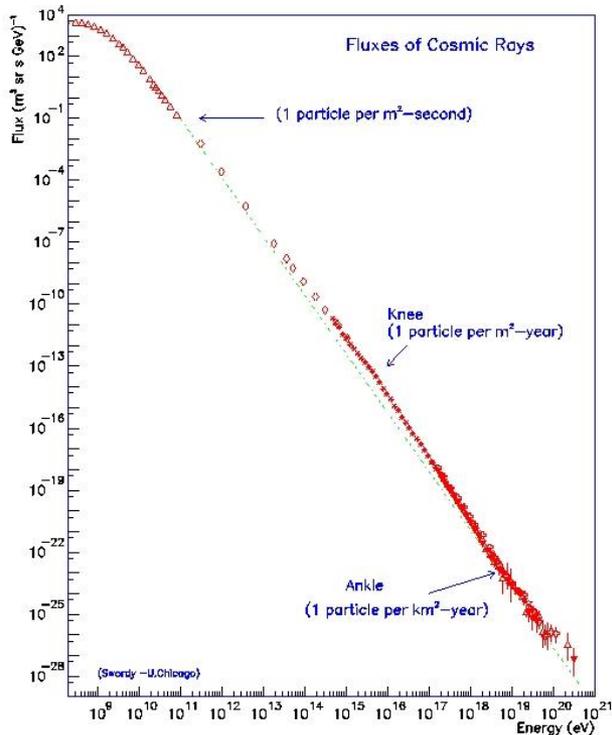
- 1912 – discovery of cosmic rays – Victor Hess;
- 1929 – Bothe and Kolhorster, tracks are curved by magnetic fields;
- 1932 – Milikan and Compton – gamma rays → cosmic rays;
- 1934 – cosmic ray primaries are positively charged;
- 1934-1938 – Rossi, Auger – coincidence measurements – extensive air showers (EAS);
- 1934 – Bethe and Heitler – electromagnetic cascade theory;
- 1949 – Fermi acceleration (galactic magnetic clouds);
- 1952-1954 – Accelerators over 1 GeV – decoupling of cosmic ray and high energy physics;
- 1954 – first measurements of high energy cosmic rays (EAS);



- Stable charged particles and nuclei $\tau \sim 10^6$ yrs
- **Primaries** – accelerated at astrophysical sources – e^- (2%), p (87%), He (12%), C, O, Fe (1%) – synthesized in stars, γ, ν ;
- **Secondaries** – produced in interactions of primaries with interstellar gas – Li, Be, B (\bar{p} and e^+);
- **Solar modulation**;



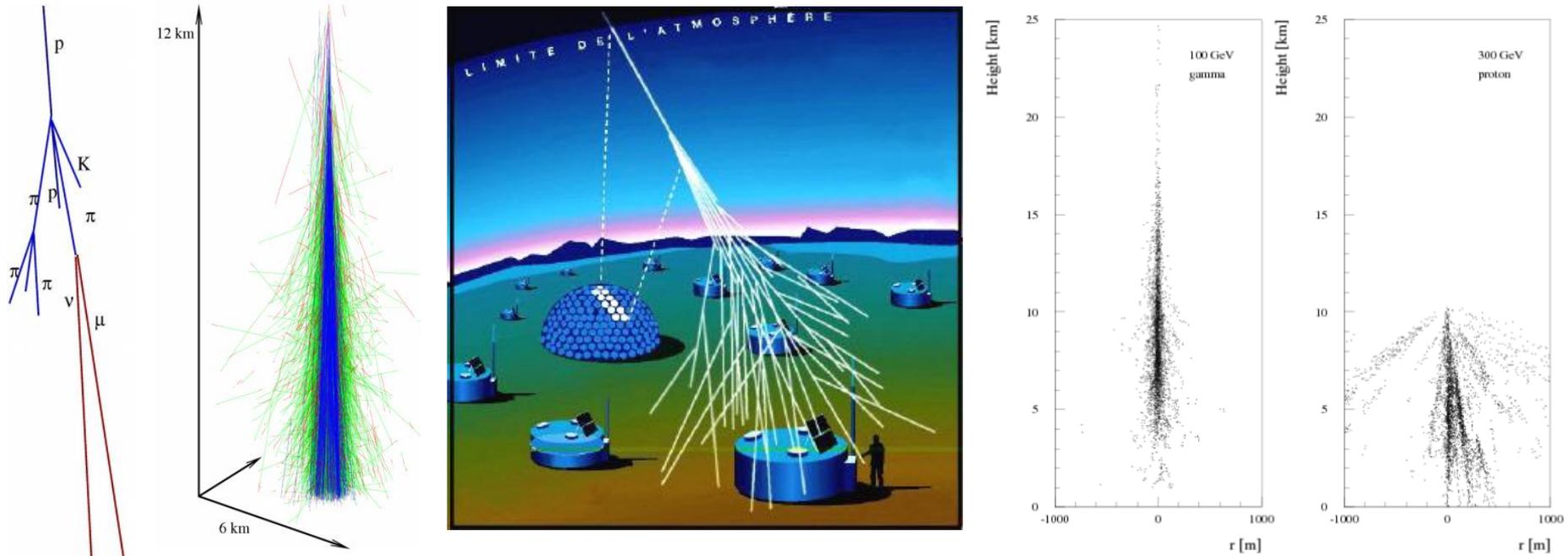
Cosmic ray spectrum



- 2 main sources – **Galactic and extragalactic**;
- Supernova remnants – SNe II, Pulsars, Active galactic nuclei (AGN), Gamma Ray Bursts;
- Shock acceleration (1st order Fermi acceleration)
- 2nd order Fermi acceleration (scattering on a moving magnetic clouds)

- **Knee** – $\sim 10^{15} - 10^{16} \text{ eV}$ → maximum energy for galactic accelerators (supernovae $\sim 10^{15} \text{ eV}$, confinement in the galaxy);
- **2nd knee** – $\sim 8 \times 10^{16} \text{ eV}$ accompanied by a transition to heavy primaries;
- **Ankle** – $\sim 10^{18.5} \text{ eV}$, **GZK effect** (high σ for protons on CMB), **photodisintegration** (nuclei)

Extensive air showers (EAS)

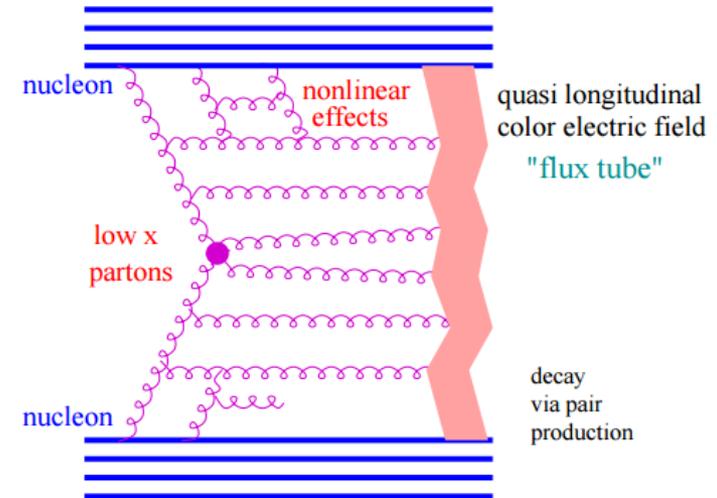


- High enough energy for its cascade to be **detected at the ground**
- **Hadronic core** – source of **em. subshowers** – $\pi^0 \rightarrow \gamma\gamma$
- Electrons, positrons – **most numerous**
- Muons – decays of charged mesons – **an order of magnitude lower**
- **Large** ground area
- **Array of detectors** (Sparse as possible – **large aperture**) on the ground for secondaries
 - **Scintillation** detectors;
 - **Cherenkov** detectors;
- Timing information – **reconstruction** of arrival direction and core of the shower;
- **Fluorescence light method** – Electrons excite nitrogen molecules – **isotropic fluorescence light** – shower evolution – the energy of the primary.

Cosmic ray models

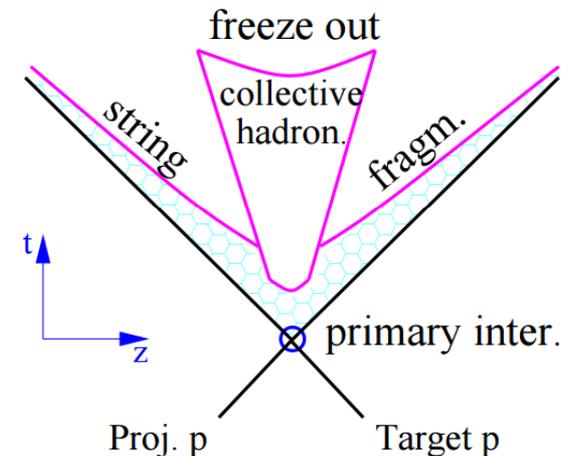
EPOS

- Minimum bias event generator;
- Heavy ion and cosmic ray interactions;
- **Parton ladder** (cut Pomeron);
- “flux tube” – relativistic string (which fragments);
- **Core-corona model (unique)**;
- Effects not present in other models;
- The core – a region with a high density of string segments
 - the hadronization is treated **collectively**;
- Corona – a region with low density of string segments.



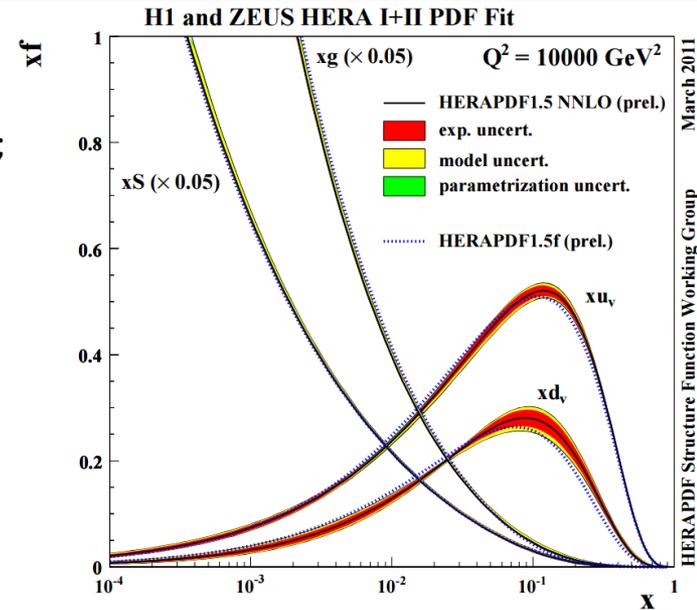
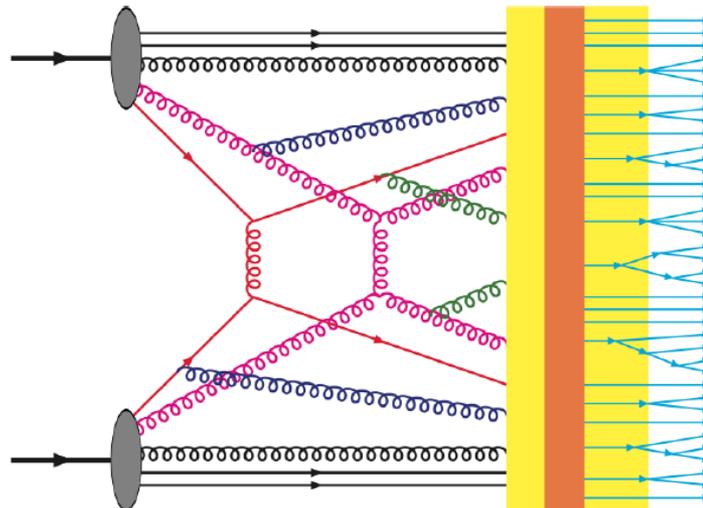
QGSJET

- Extensive air shower simulations;
- Quark-Gluon string model;
- Also nucleus-nucleus interactions and semihard processes;
- Semihard “Pomeron” approach;
- QGSJET-II **non-linear effects** (enhanced Pomeron diagrams);
- Gribov’s reggeon approach – multiple scattering processes;
- Pomeron exchanges – **microscopic parton cascades**
 - “Soft” and “semihard” Pomerons.



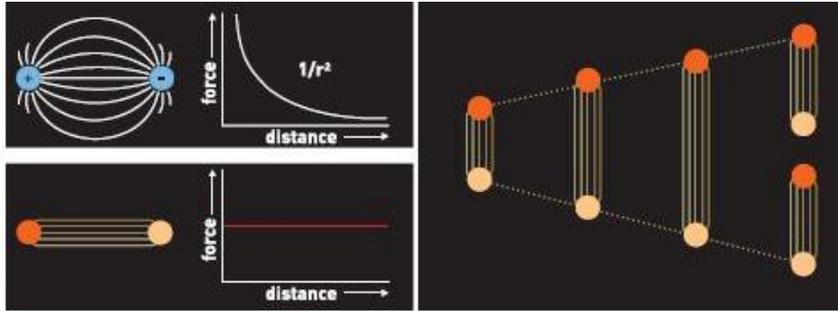
PYTHIA

1. **PDF** – Parton distribution functions;
2. **Hard process** – interaction between the most energetic partons;
3. **Initial State Radiation, ISR**
→ parton branching on hard process ingoing partons;
4. **Final State Radiation, FSR**
→ parton branching on hard process outgoing partons;
5. **Multiple Parton Interactions, MPI**
→ interactions between the other partons;
6. **Hadronization**
→ Independent fragmentation;
→ String fragmentation;
→ Cluster fragmentation.
7. **Decay.**



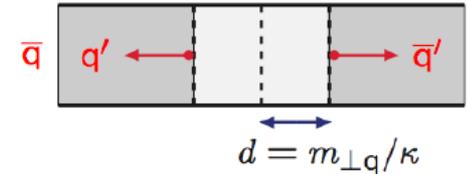
- PDF
- ME
- MPI
- ISR
- FSR
- BR
- CR
- Hadr.
- Decays
- Rescattering
- BE
- Unknown?

Lund string model

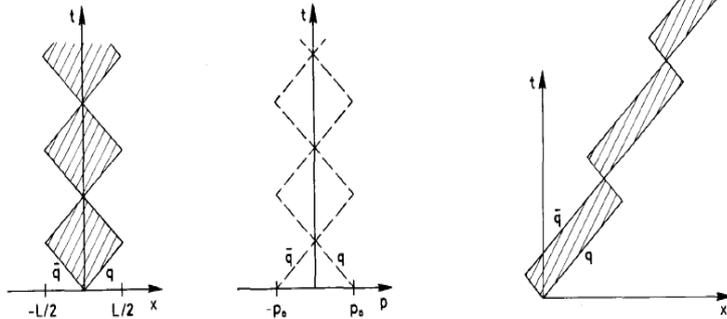


Field lines compressed to tubelike regions: **strings!**
 Linear confinement:
 $V(r) \approx \kappa r$, $\kappa \approx 1 \text{ GeV/fm}$
 Lorentz invariant.

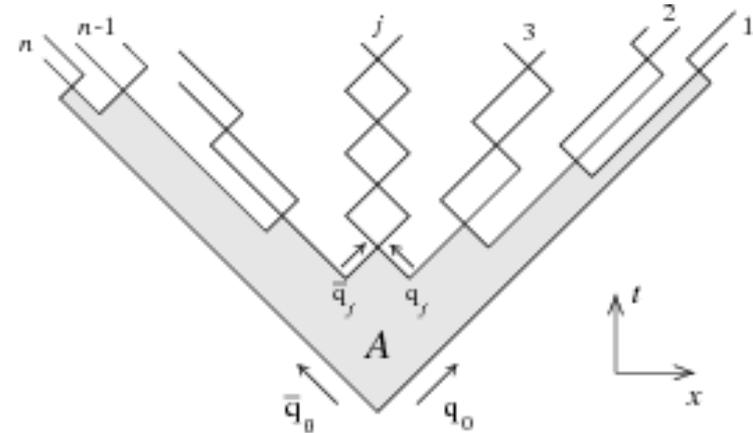
String breaking by tunneling:
 $\mathcal{P} \propto \exp(-\pi m_{\perp q}^2 / \kappa)$
 with adjacent pairs forming mesons (and baryons).



- Bound state:
 $\rightarrow \frac{dp}{dt} = \pm k;$
 \rightarrow "yo-yo" system.

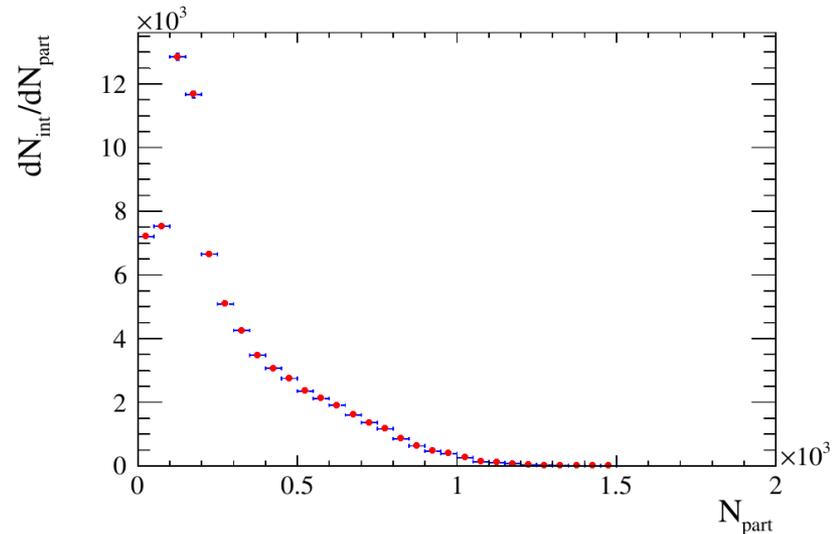
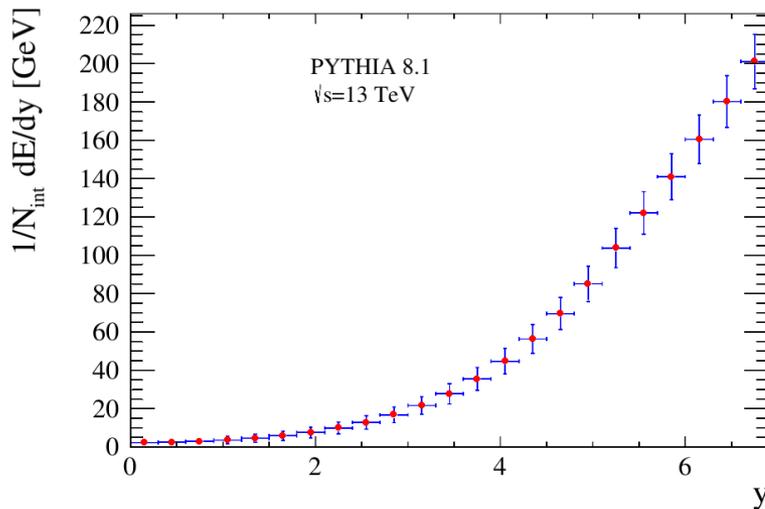


- Fragmentation:

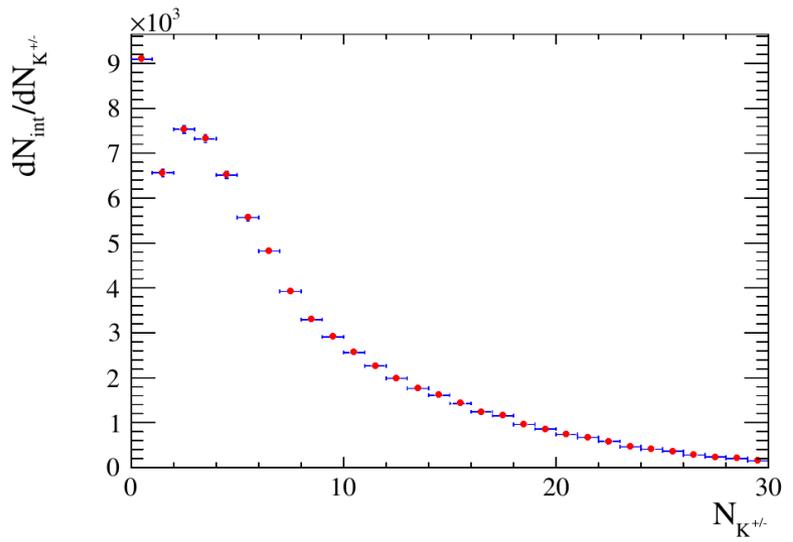
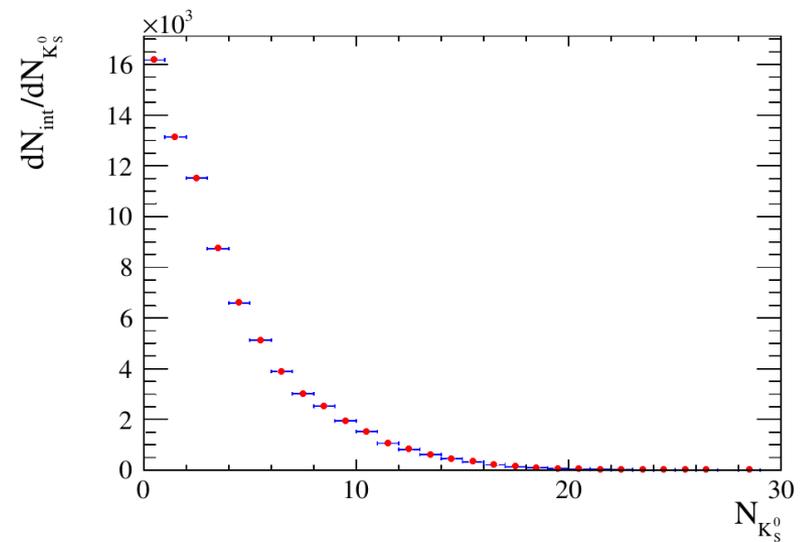
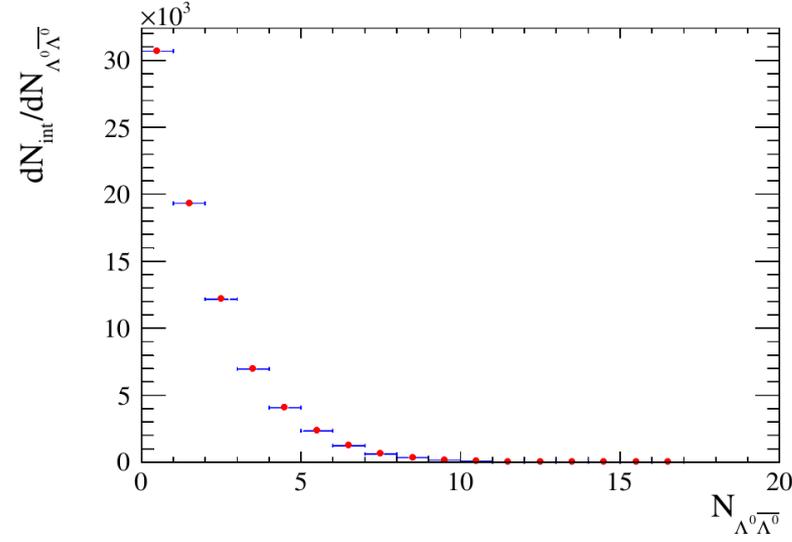


Observables

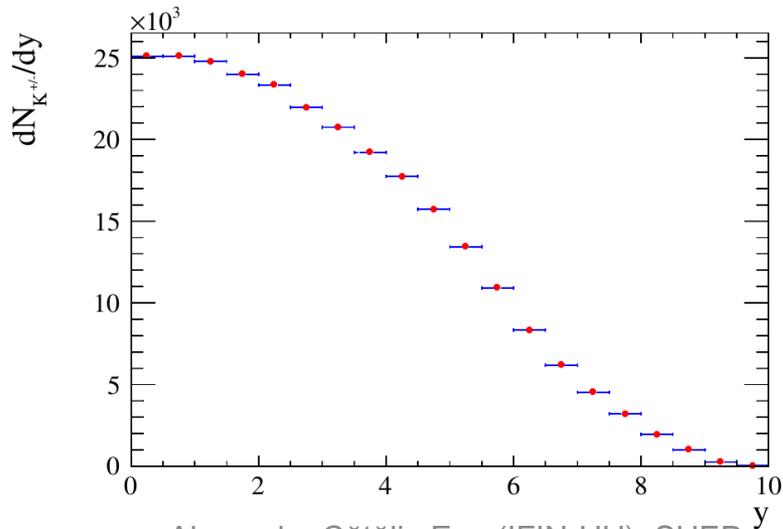
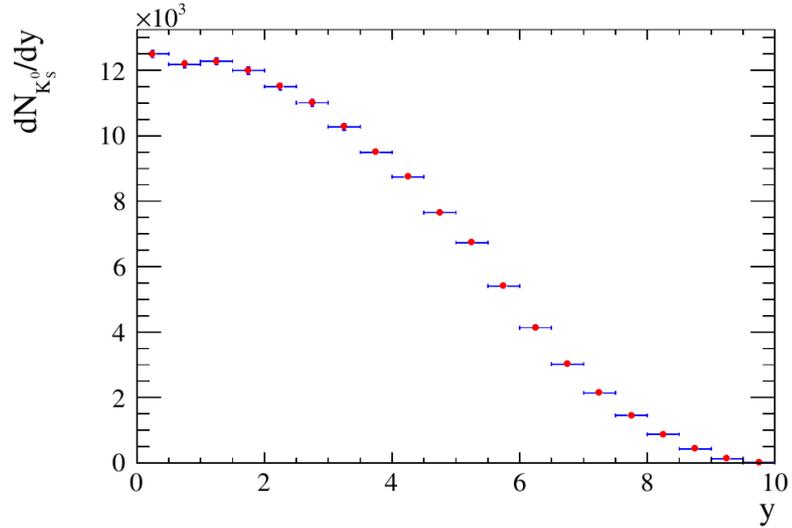
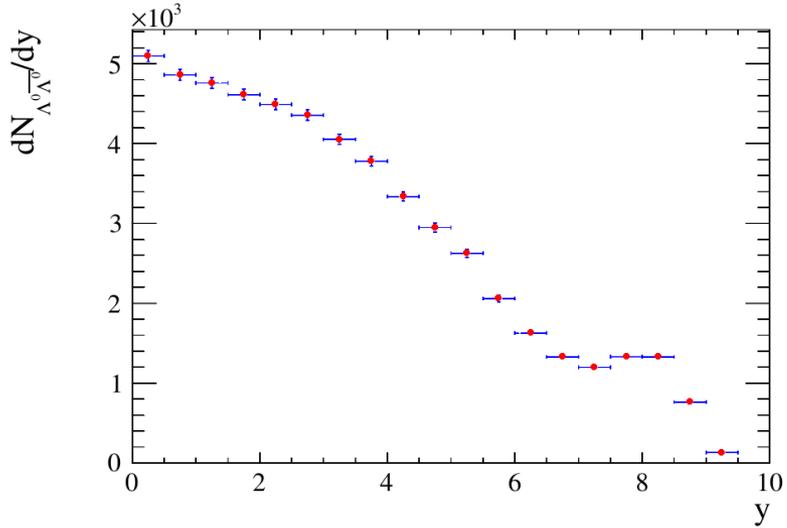
- Rapidity: $y = \frac{1}{2} \ln \left(\frac{E+p_z c}{E-p_z c} \right) \rightarrow y = \frac{1}{2} \ln \left(\frac{E+p_L}{E-p_L} \right)$
- Pseudorapidity: $\eta = -\ln \tan \frac{\theta}{2}$ or $\eta = \frac{1}{2} \ln \left(\frac{|p|+p_L}{|p|-p_L} \right)$
- Beam rapidity: $y_{beam} = \ln \left(\frac{2E_{beam}}{m_p} \right)$
- Rapidity loss: $\Delta y = y_{beam} - y \rightarrow$ Lorentz invariant
- Energy flow: $\frac{dE}{dy}, \frac{dE}{d\theta}, \frac{dE}{d\omega}$ etc.
- Transverse momentum: $p_T = \sqrt{p_x^2 + p_y^2}$



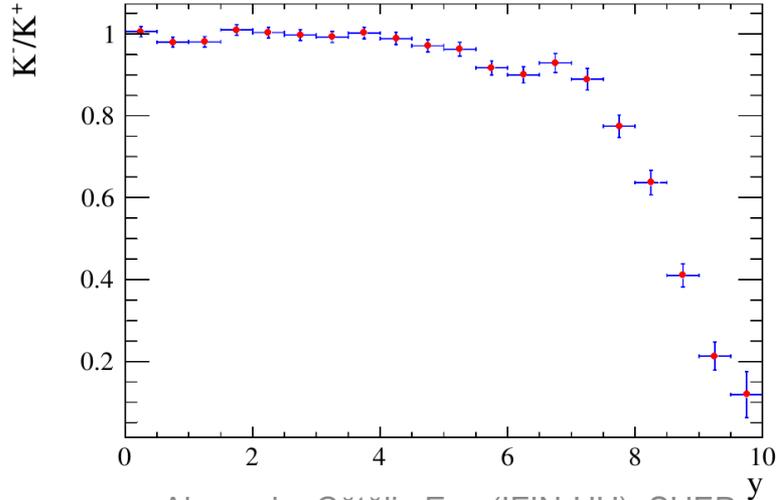
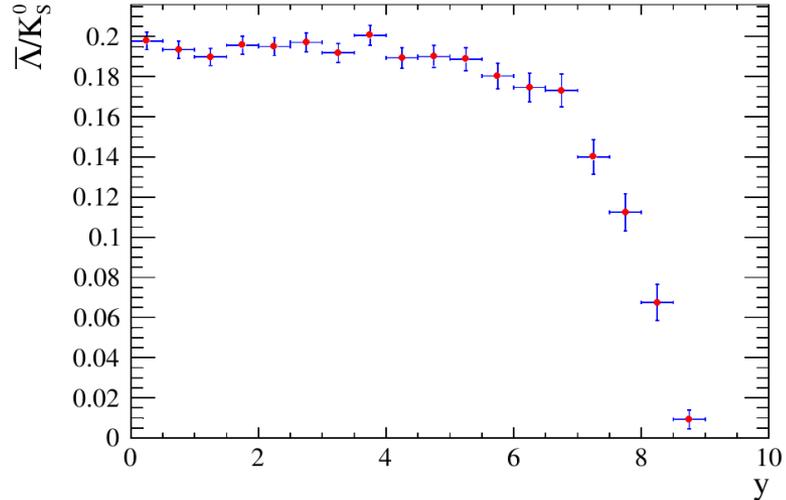
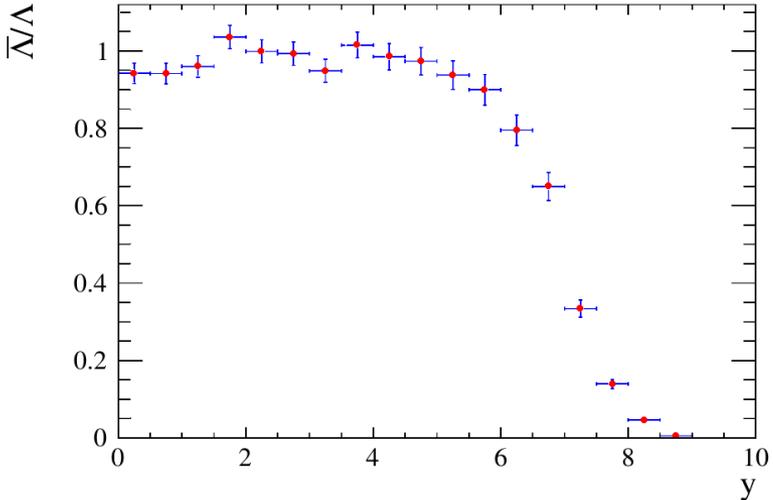
Observables



Observables

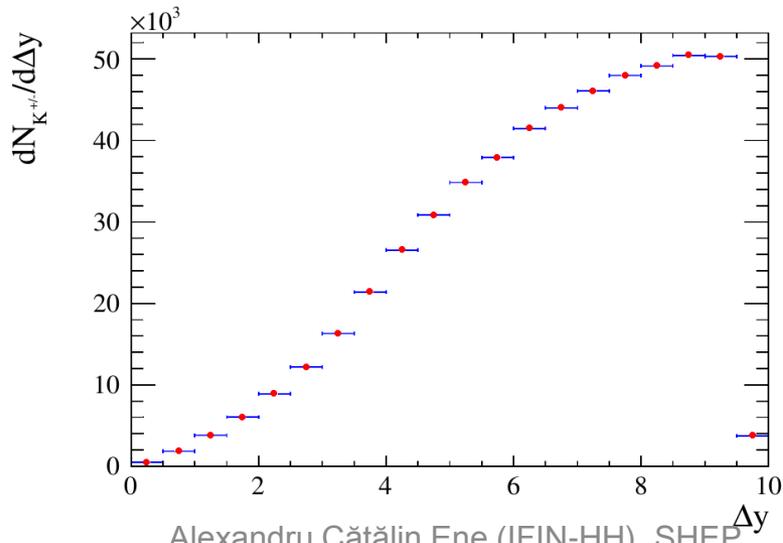
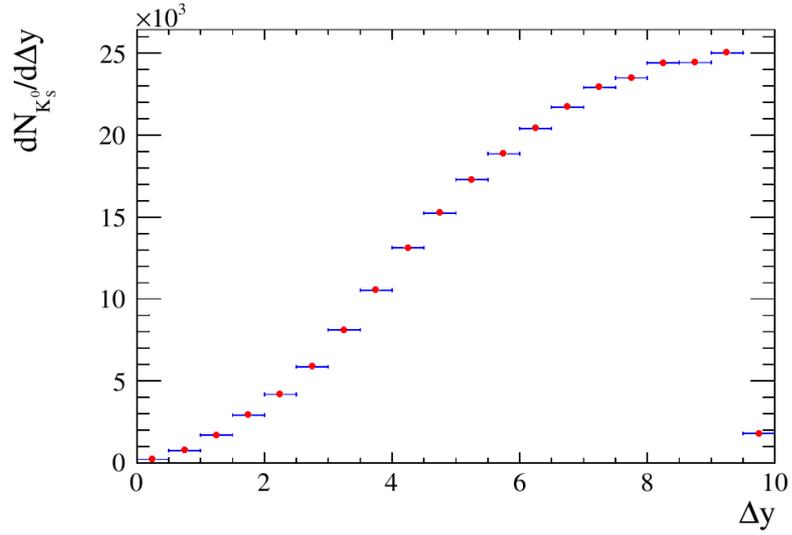
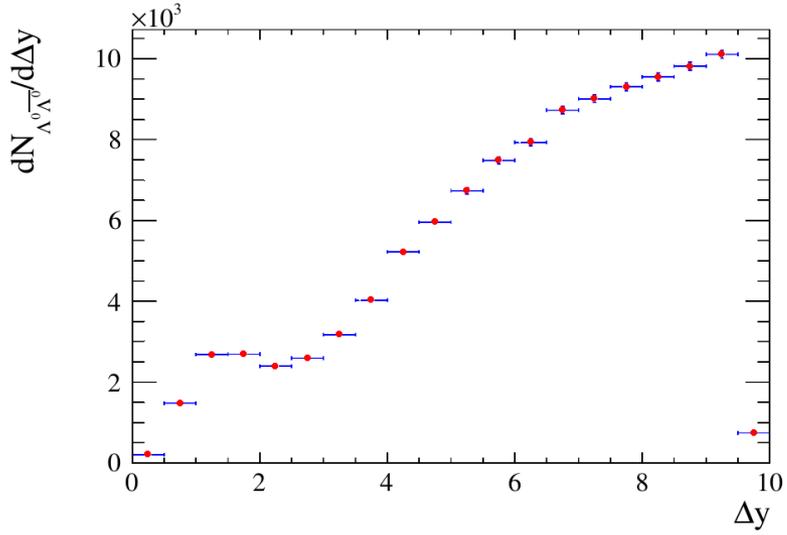


Observables

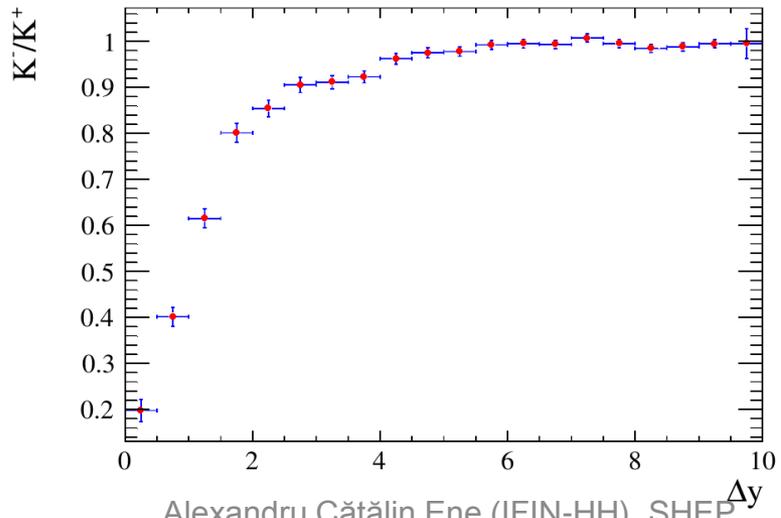
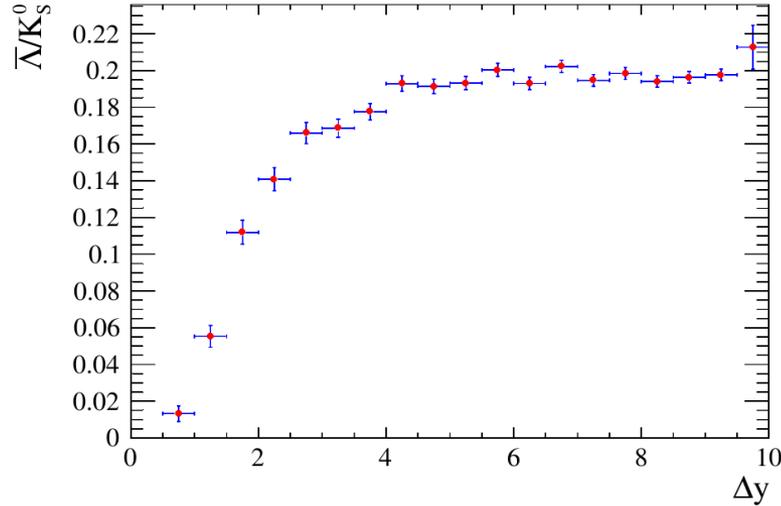
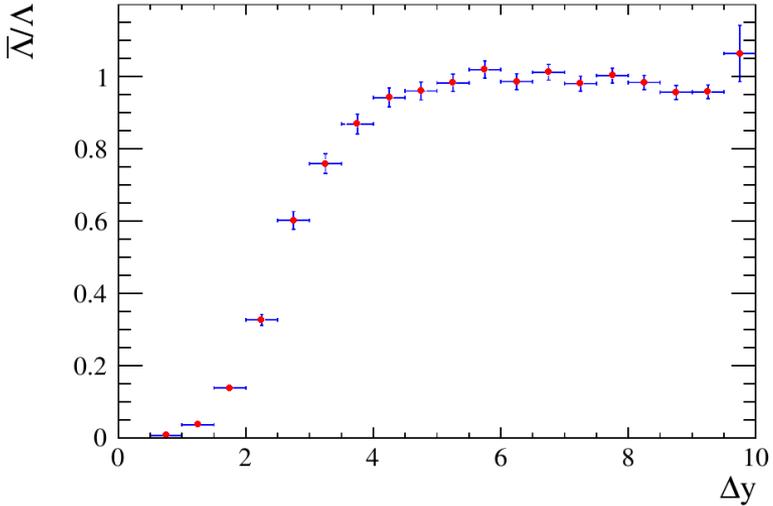


Alexandru Cătălin Ene (IFIN-HH), SHEP
2016

Observables



Observables



Future plans

We will compare the output from PYTHIA with the one from the cosmic ray models and try to explain the differences.

References

<http://physik.uibk.ac.at/hephy/Hess/more/VFHess.gif>
<http://astroparticle.uchicago.edu/siteold/sciam1.jpg>
<http://moriond.in2p3.fr/J05/trans/sunday/engel1.pdf>
<http://images.iop.org/objects/phw/news/7/10/13/cascade.jpg>
<http://s2.quickmeme.com/img/fc/fc485d49a8ceacf0a8d546b6a56d5c0b9640a2f68eccbe81755a4eebae0701a1.jpg>
<http://slice.seriousseats.com/images/20081203-original-cosmic-rays.jpg>
<http://www.frontline.in/static/html/fl2610/images/20090522261008505.jpg>
https://inspirehep.net/record/475420/files/figures_lunddecay.png
<http://home.thep.lu.se/~torbjorn/pythia81html/Welcome.html>
http://www.scholarpedia.org/article/Parton_shower_Monte_Carlo_event_generators
B. Andersson, G. Gustafson, G. Ingelman și T. Sjöstrand. Parton fragmentation and string dynamics. Physics Reports 97, 1983.
T. Sjöstrand, Soft QCD (and PYTHIA), ATLAS SM Workshop, Annecy, 4th February 2015.
<http://home.thep.lu.se/~torbjorn/ppp2011/lec7.pdf>
<http://pdg.lbl.gov/2015/reviews/rpp2015-rev-cosmic-rays.pdf>
<http://moriond.in2p3.fr/J05/trans/sunday/engel1.pdf>
<https://arxiv.org/pdf/0801.4376v1.pdf>
<https://arxiv.org/pdf/hep-ph/0412332v1.pdf>
<https://arxiv.org/pdf/1306.0121v2.pdf>
<https://arxiv.org/pdf/hep-ph/0603175v2.pdf>
<http://icecube.wisc.edu/~tmontaruli/801/lect14.pdf>
http://www.hep.shef.ac.uk/edaw/PHY206/Site/2012_course_files/phy206rlec7.pdf

