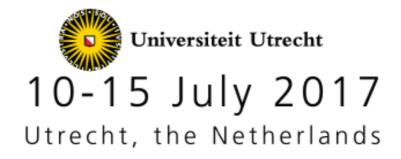
## **Gravitational Waves & Neutron Stars** - Discussion Session -

David.Blaschke@gmail.com JINR Dubna & MEPhl Moscow, Russia & University of Wroclaw, Poland

- 1. Jo van den Brand: "Gravitational waves measurements"
- 2. Luciano Rezzolla "From gravitational-wave spectroscopy to nuclear EOS"
- 3. Thomas Klaehn: "Strange-quark matter in neutron stars?"
- 4. Stefan Schramm: "Dense and warm matter in compact stars ..."
- 5. David Blaschke: "Discussion"

17<sup>th</sup> International Conference on Strangeness in **Quark Matter** 













Russian Science Foundation

H. Stroebele: How likely is it that s-quarks (and no s-bar) exist and survive in neutron stars in a QGP or in hyperons. How large is then the ratio s/(u+d) in neutron stars and in the universe?

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It is not impossible that there is no strangeness in the Universe except for particle collision Experiments! See talk by T. Klaehn

#### Arguments:

Onset of hypernuclear matter could be preceded by that of (two-flavor) quark matter. Strange quarks are likely to appear sequentially (at T=0) due to their larger mass. Once strangeness degree of freedom appears in quark matter, the EoS is softened and compact stars become unstable against gravitational collapse.

Increasing density

H. Stroebele: How likely is it that s-quarks (and no s-bar) exist and survive in neutron stars in a QGP or in hyperons. How large is then the ratio s/(u+d) in neutron stars and in the Universe?

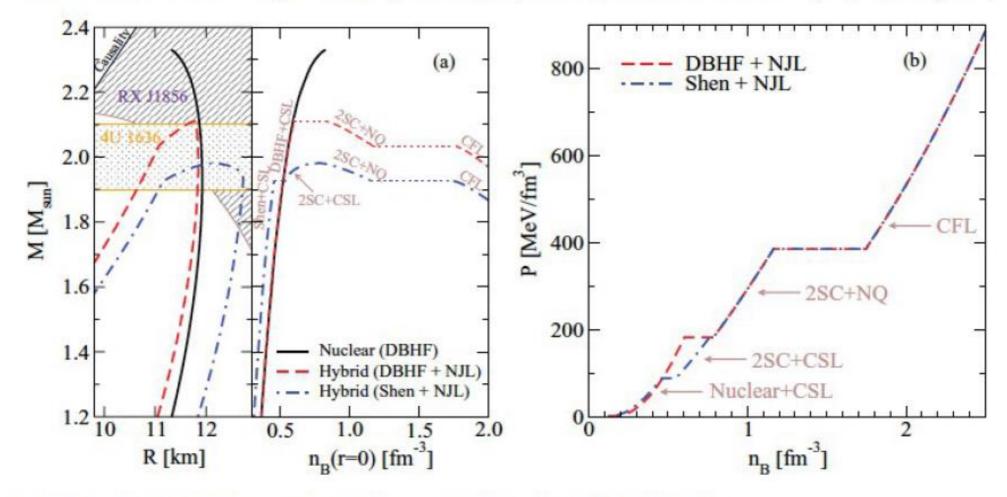
There could also be single flavor quark matter, mixed with nuclear matter (d-quark dripline)

pure nuclear matter d-quark drip 2-flavor quark matter 3-flavor quark matter (NM) (NM + d-CSL phase)(2SC phase) (CFL phase)

D.B., F. Sandin, T. Klaehn, J. Berdermann, PRC 80 (2009) 065807

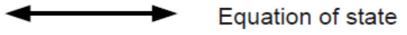
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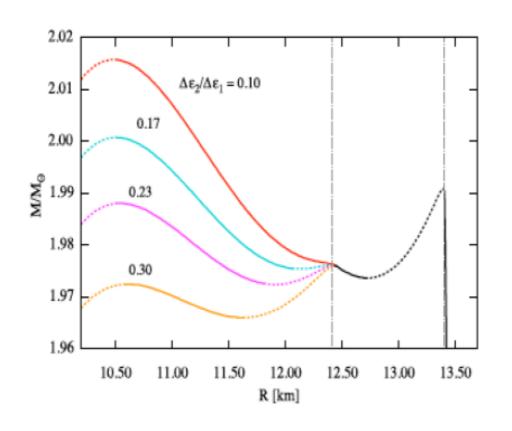
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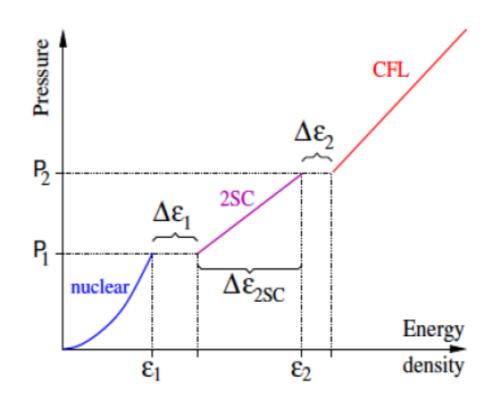


D.B., F. Sandin, T. Klaehn, J. Berdermann, PRC 80 (2009) 065807

Measuring Mass vs. Radius







## High-mass twins:

D. Blaschke et al., PoS CPOD 2013

S. Benic et al., A&A 577 (2015) A50

#### **High-mass triples:**

M. Alford and A. Sedrakian, arxiv:1706.01592

## Advertisement:

### THOR EU COST Action CA15213

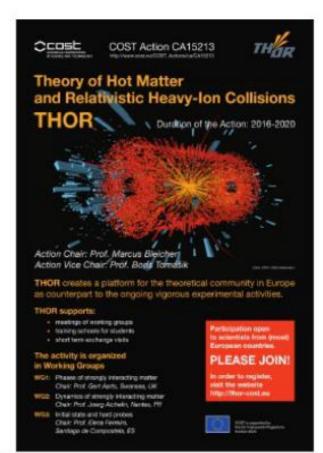
 Theory of Hot Matter and Relativistic Heavy Ion Collisions http://thor-cost.eu

## NewCompStar EU COST Action MP1304

 Theory of Compact Stars (ending 2017) http://compstar.uni-frankfurt.de

## PHAROS EU COST Action CA16214

 The multi-messenger physics and astropysics of neutron stars



## 6th International Workshop on

# Compact Stars in the QCD Phase Diagram VI

(Cosmic matter in heavy-ion collision laboratories?)

**Dates:** 26.-29. September 2017

Venue: Dubna, Russian Federation

Organizers: D. Blaschke, H. Grigorian

Website: http://www.quarknova.ca/CSQCD.html

http://theor.jinr.ru/meetings/2017 (t.b.u.)

## Previous meetings:

Copenhagen (2001), Beijing (2009), Guaruya (2012), Prerow (2014), Gran Sasso (2016)

## Topics:

- QCD phase diagram for HIC vs. Astrophysics
- Quark deconfinement in HIC vs. supernovae, neutron stars and their mergers
- Strangeness in HIC and in compact stars



