

Monte Carlo generators for CR interactions

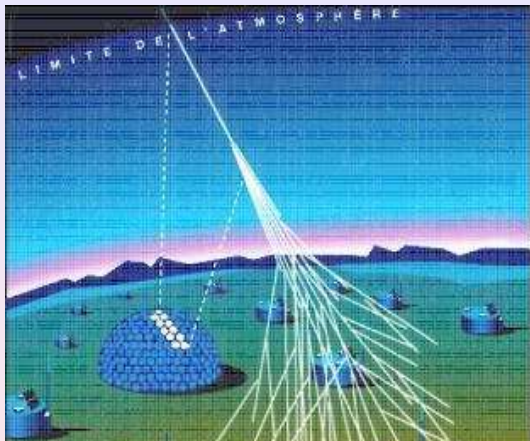
Sergey Ostapchenko

Frankfurt Institute for Advanced Studies

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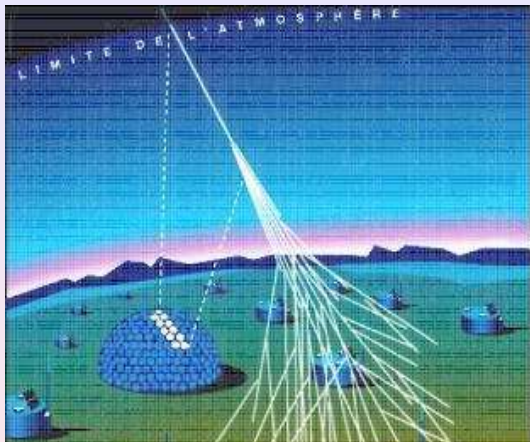
Cosmic ray studies with Extensive Air Shower technique



ground-based observations

- primary CR energy \iff charged particle density at ground
- CR composition \iff muon density ρ_μ at ground

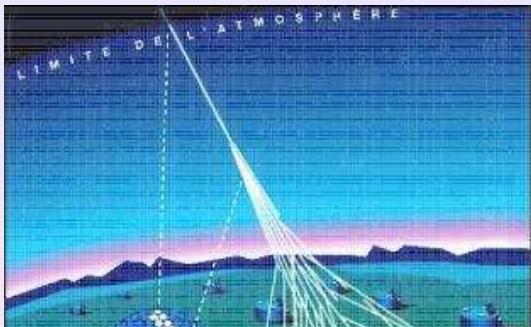
Cosmic ray studies with Extensive Air Shower technique



measurements of EAS fluorescence light

- primary CR energy \iff integrated light
- CR composition \iff shower maximum position X_{\max}

Cosmic ray studies with Extensive Air Shower technique



CR composition studies – most dependent on interaction models

- e.g. predictions for X_{\max} : **on the properties of the primary particle interaction** ($\sigma_{p\text{-air}}^{\text{inel}}$, forward particle spectra)
 - \Rightarrow most relevant to LHC studies of pp collisions
- predictions for muon density: on secondary particle interactions (cascade multiplication); mostly on $N_{\pi\text{-air}}^{\text{ch}}$
 - \Rightarrow **small potential influence of 'new physics'**

- 1 **QGSJET-II-04** [*SO, PRD83 (2011) 014018*]
 - based on the Reggeon Field Theory (RFT) approach
 - nonlinear effects: Pomeron-Pomeron interactions

Cosmic ray interaction models

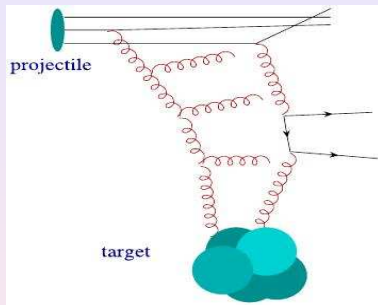
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- 2 **EPOS-LHC** [*Pierog, Karpenko, Katzy, Yatsenko & Werner, PRC92 (2015) 034906*]
 - also RFT-based but involves phenomenological solutions (e.g. parametrized saturation effects)
 - additional theoretical mechanisms (e.g. energy-momentum sharing at the amplitude level, hydrodynamics for final states)
 - generally better description of existing data (e.g. p_T -spectra)

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 - generally better description of existing data (e.g. p_T -spectra)
- 3 **SIBYLL-2.3** [*Riehn, Engel, Fedynitch, Gaisser & Stanev, arXiv:1510.00568*]
 - similar to most of the generators used at the LHC (based on the 'minijet' approach)
 - includes multiple soft interactions \Rightarrow some similarity to RFT-based models

Hadronic interactions: qualitative picture

- QCD-inspired: **interaction mediated by parton cascades**
- multiple scattering
(many cascades in parallel)
- real cascades
⇒ particle production
- virtual cascades
⇒ elastic rescattering
(just momentum transfer)



Universal interaction mechanism

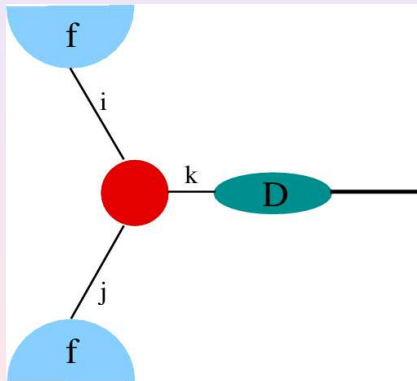
- different hadrons (nuclei) ⇒ **different initial conditions**
(parton Fock states) but same mechanism
- energy-evolution of the observables (e.g. σ_{pp}^{tot}):
due to a larger phase space for cascades to develop

Hadronic interactions: input from pQCD & problems

- pQCD: **collinear factorization applies for inclusive spectra**

$$\frac{d^3\sigma_{pp\rightarrow h}}{dp^3} = \sum_{i,j,k} f_{i/p} \otimes \sigma_{ij\rightarrow k} \otimes f_{j/p} \otimes D_{h/k}$$

- separates short- & long-distance dynamics
- pQCD predicts evolution of PDFs ($f_{i/p}$) & FFs ($D_{h/k}$)
- \Rightarrow allows to simulate perturbative (high p_t) part of parton cascades (initial & final state emission)

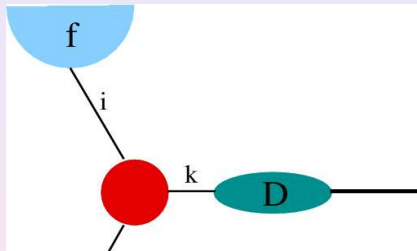


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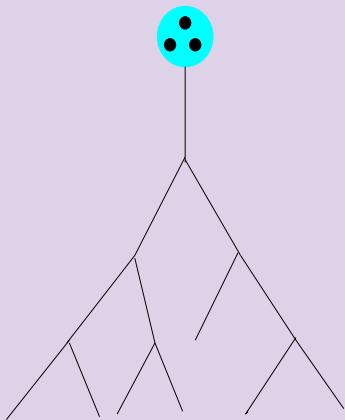
What is beyond and why the models are so different?

- nonperturbative (low p_t) parton evolution ('soft' rescatterings; very initial stage of 'semihard' cascades)
- multiple scattering aspect
- nonlinear effects (interactions between parton cascades)
- constituent parton Fock states & hadron 'remnants'

Hadronic interactions: nonperturbative Fock states

1. (Implicitly) always same nonperturbative Fock state (typical for models used at colliders, also SIBYLL model)

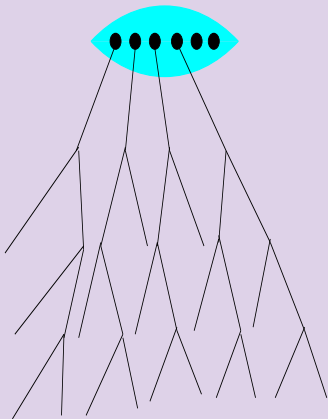
- multiple parton cascades originate from the same initial parton state
- **multiple scattering has small impact on forward spectra**
 - new branches emerge at small x
($G(x, q^2) \propto 1/x$)
- \Rightarrow **Feynman scaling & limiting fragm. for forward production**
- higher $\sqrt{s} \Rightarrow$ more abundant central particle production
- **forward & central production – decoupled from each other**
 - (decreasing number of cascade branches for increasing x)



Hadronic interactions: nonperturbative Fock states

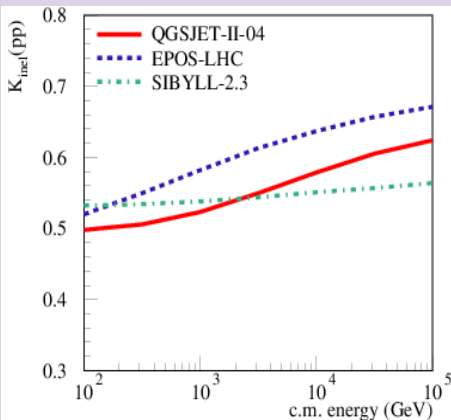
2. $p = \sum$ of multi-parton Fock states [EPOS & QGSJET(-II)]

- many cascades develop in parallel (already at nonperturbative stage)
- higher $\sqrt{s} \Rightarrow$ larger Fock states come into play: $|qqq\rangle \rightarrow |qqq\bar{q}q\rangle \rightarrow \dots |qqq\bar{q}q\dots\bar{q}q\rangle$
 - \Rightarrow softer forward spectra (energy sharing between constituent partons)
- forward & central particle production - strongly correlated
 - e.g. more activity in central detectors \Rightarrow larger Fock states \Rightarrow softer forward spectra



Why of importance for air shower predictions?

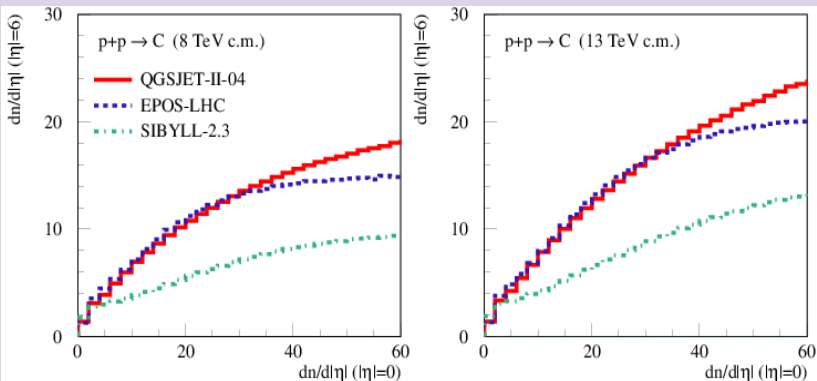
Main cause: energy-dependence of the nucleon 'inelasticity'



- SIBYLL: K_{pp}^{inel} - weak energy dependence
 - for increasing \sqrt{s} , mostly central production enhanced
- smaller K^{inel} \Rightarrow stronger 'leading particle' effect
- \Rightarrow slower shower development (larger X_{max})

'Smoking gun' test: signal correlations in CMS & TOTEM

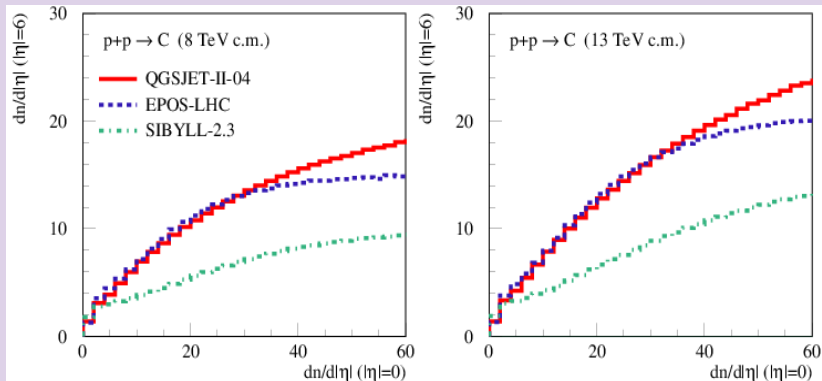
Cross-correlation of $dN_{pp}^{\text{ch}}/d|\eta|$ at $\eta = 0$ ($p_t > 0.1$ GeV) and $\eta = 6$



- strong correlation for QGSJET-II-04 & EPOS-LHC (apart from the tails of the multiplicity distributions)
- twice weaker correlation for SIBYLL-2.3

'Smoking gun' test: signal correlations in CMS & TOTEM

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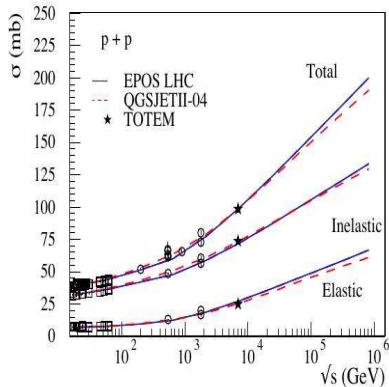
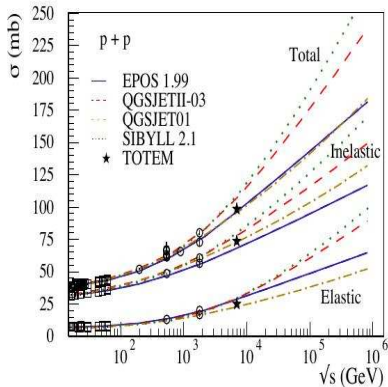


- strong correlation for QGSJET-II-04 & EPOS-LHC (apart from the tails of the multiplicity distributions)

Alternatively: discrimination by LHCf & ATLAS (see extra slides)

All the models: updated with Run 1 data of LHC

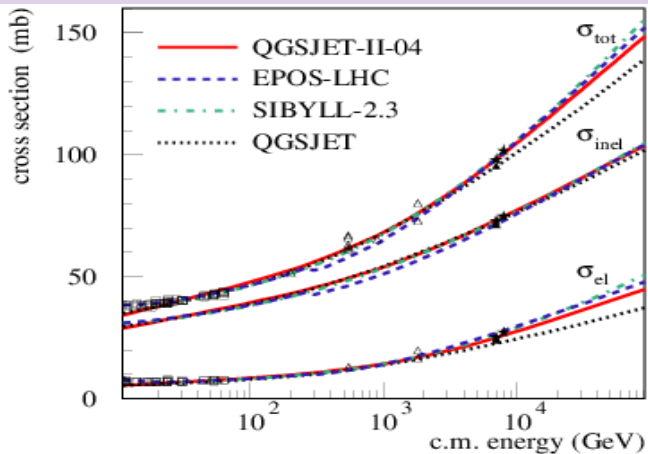
Most important: data of TOTEM & ATLAS ALFA for $\sigma_{pp}^{\text{tot/el}}$



[R. Engel, talk at "Composition-2015"]

All the models: updated with Run 1 data of LHC

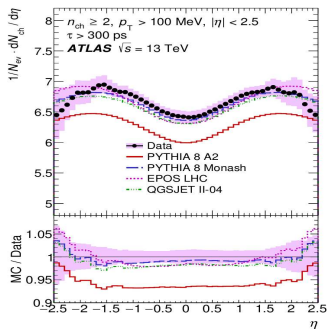
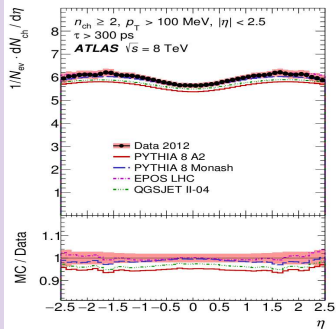
Now: very similar high energy extrapolations for all the models



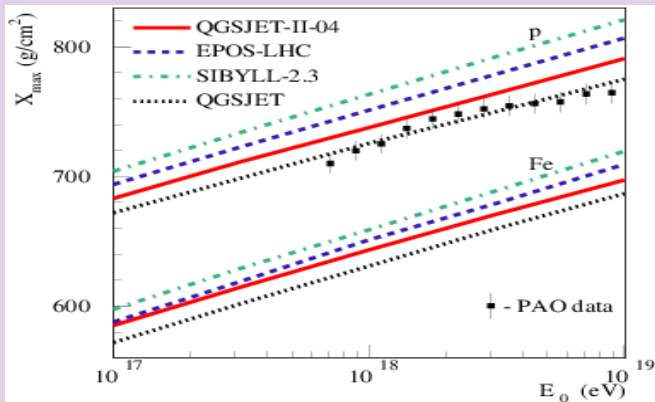
- NB: $\sigma_{p\text{-air}}^{\text{inel}}$ defines where the cascade starts

All the models: updated with Run 1 data of LHC

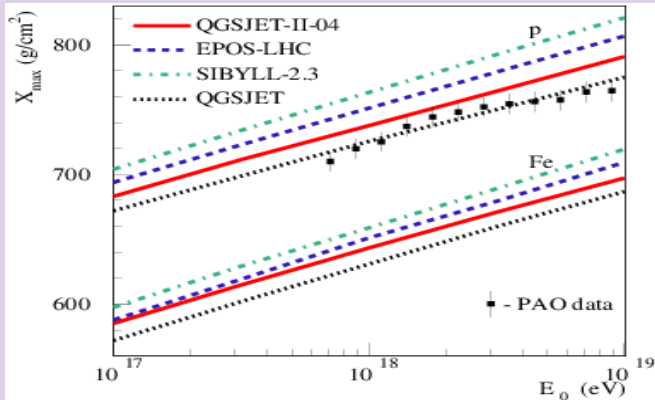
Also for central production: $dn_{pp}^{\text{ch}}/d\eta$ vs. ATLAS ($\sqrt{s} = 8, 13$ TeV)



Model predictions for EAS, e.g. X_{\max} : yet large differences



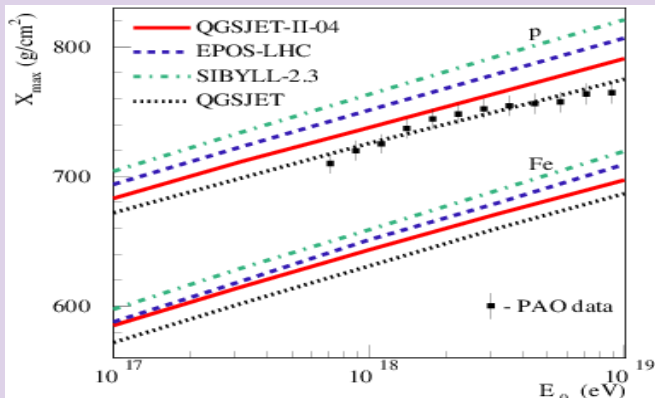
Model predictions for EAS, e.g. X_{\max} : yet large differences



Deepest X_{\max} of SIBYLL-2.3 – mainly due to the smallest $K_{p\text{-air}}^{\text{inel}}$

- direct consequence of the assumptions on parton Fock states
 - can be discriminated at LHC (central-forward correlations)

Model predictions for EAS, e.g. X_{\max} : yet large differences



For other models: treatment of proton diffraction?

- $\sigma_{pp}^{\text{diff}}$ impacts recalculation from pp to pA (AA)
 - $\sigma_{p\text{-air}}^{\text{inel}}$ – due to inelastic screening
 - directly related to $\sigma_{p\text{-air}}^{\text{diff}}$, hence, also to $K_{p\text{-air}}^{\text{inel}}$ – due to small 'inelasticity' of diffractive collisions (especially for target SD)

Impact of diffraction uncertainties on X_{\max} predictions

[SO, PRD 89, 074009 (2014)]

Presently: tension between CMS & TOTEM concerning σ_{pp}^{SD}

	TOTEM	CMS
M_X range, GeV	7 – 350	12 – 394
$\sigma_{pp}^{\text{SD}}(\Delta M_X)$, mb	$\simeq 3.3$	4.3 ± 0.6
$\frac{d\sigma_{pp}^{\text{SD}}}{dy_{\text{gap}}}$, mb	0.42	0.62

- \Rightarrow may be regarded as the characteristic uncertainty for σ_{pp}^{SD}

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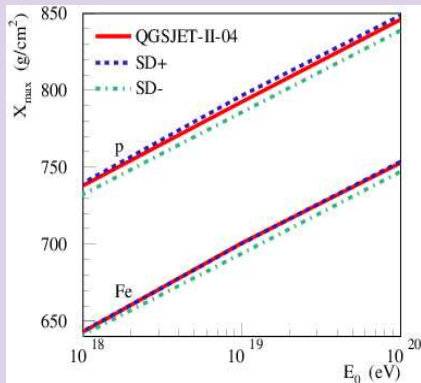
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Two alternative model versions (tunes): SD+ & SD-

- SD+: increased high mass diffraction (HMD)
 - to approach CMS results
 - slightly smaller LMD – to soften disagreement with TOTEM
- SD-: **smaller LMD (by 30%)**, same HMD
- similar $\sigma_{pp}^{\text{tot/el}}$ & central particle production in both cases

Impact of diffraction uncertainties on X_{\max} predictions

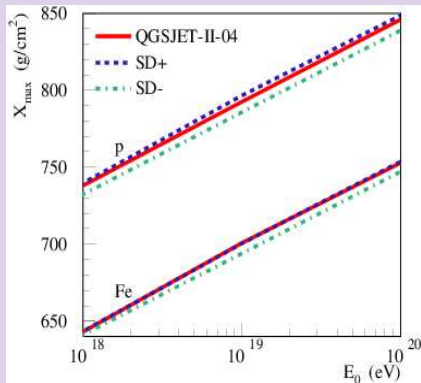
Characteristic differences: $\Delta X_{\max} \simeq 10 \text{ g/cm}^2$



- option SD-:
 - smaller inelastic screening
 \Rightarrow larger $\sigma_{p\text{-air}}^{\text{inel}}$
 - smaller diffraction for $p\text{-air}$
 \Rightarrow larger $K_{p\text{-air}}^{\text{inel}}$
 - \Rightarrow smaller X_{\max} (all effects in the same direction)
- option SD+: opposite effects

Impact of diffraction uncertainties on X_{\max} predictions

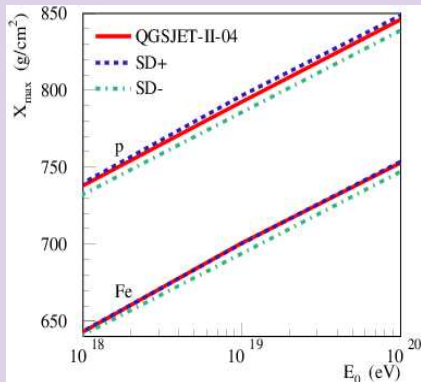
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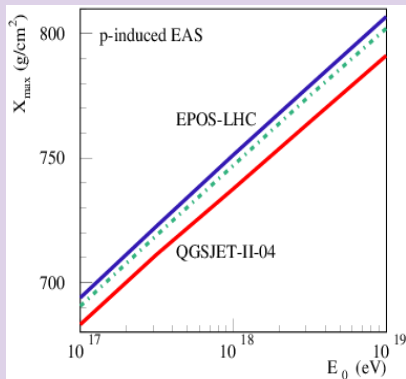
Model differences for X_{\max} twice bigger (reach 20 g/cm^2)

- other interaction properties relevant?
- may be checked using the “cocktail” approach:
using different models for certain interactions in air showers

Other sources of model uncertainties for X_{\max}

Let us compare X_{\max} of EPOS-LHC & QGSJET-II-04

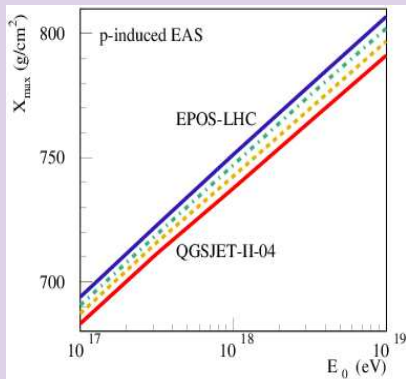
- and construct 'mixture models'
- use QGSJET-II for $\sigma_{p\text{-air}}^{\text{inel}}$ & leading nucleon spectrum (EPOS-LHC for the rest)
- $\Delta X_{\max} \leq 5 \text{ g/cm}^2$ - in agreement with above



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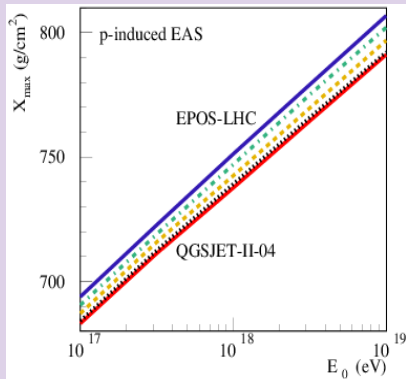
- QGSJET-II for $\sigma_{p\text{-air}}^{\text{inel}}$ & leading nucleon spectrum
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- now QGSJET-II for the complete 1st interaction (EPOS-LHC for the rest)
- $\Delta X_{\max} \leq 5 \text{ g/cm}^2$
- reason: harder pion spectra in $p\text{-air}$ in EPOS-LHC



Other sources of model uncertainties for X_{\max}

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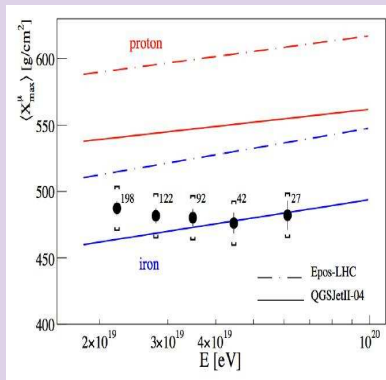
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- now QGSJET-II for the complete 1st interaction
- $\Delta X_{\max} \leq 5 \text{ g/cm}^2$
- remaining difference: copious $\bar{p}p$ - & $\bar{n}n$ -pair production and higher diffraction for π -air collisions in EPOS-LHC



Constraining pion interactions by cosmic ray data

PAO measurement of maximal muon production depth X_{\max}^{μ}

- models predict deeper X_{\max}^{μ} than observed
 - e.g. one needs primary iron for QGSJET-II-04
 - or primary gold for EPOS-LHC...

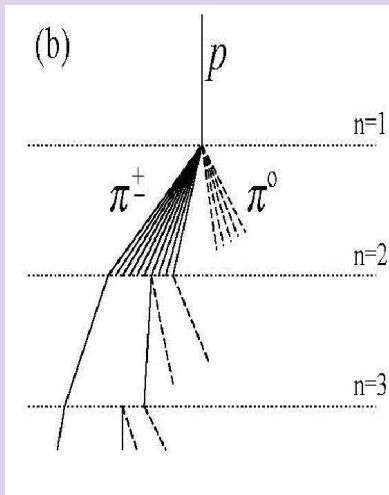


[from M. Roth, "Composition-2015" talk]

Constraining pion interactions by cosmic ray data

X_{\max}^{μ} : effects of inelastic & diffractive π – air cross sections

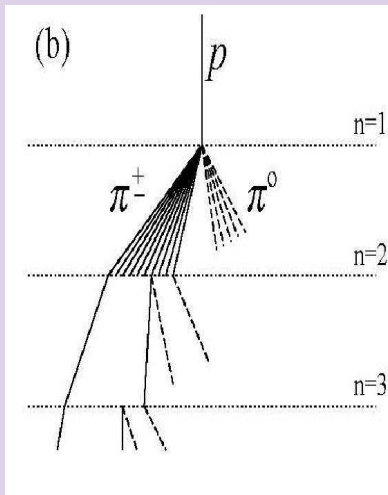
- NB: muons originate from a **multi-step hadron cascade**
- smaller $\sigma_{\pi\text{-air}}^{\text{inel}}$ \Rightarrow larger **distances between the cascade steps**
 - \Rightarrow deeper X_{\max}^{μ}
- larger diffraction in π – air \Rightarrow similar effect



Constraining pion interactions by cosmic ray data

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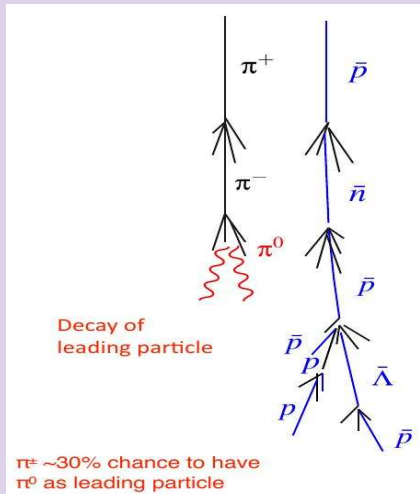
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Constraining pion interactions by cosmic ray data

X_{\max}^{μ} : relation to (anti-)baryon production

- no decay for p & \bar{p} (n & \bar{n})
⇒ few more cascade steps
- but: impact on X_{\max}^{μ} IFF
 $N_{p,\bar{p},n,\bar{n}}$ comparable to N_{π} !
(the case of EPOS)

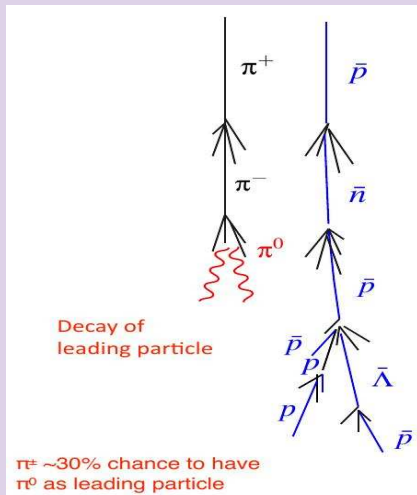


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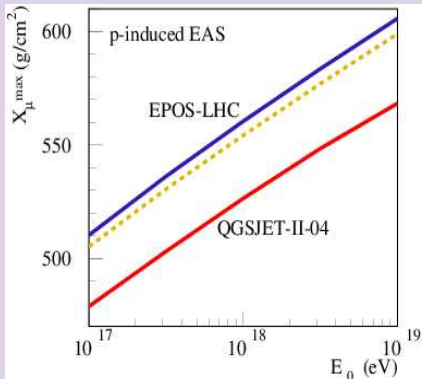


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Constraining pion interactions by cosmic ray data

Difference of X_{max}^{μ} : EPOS-LHC / QGSJET-II-04, using “cocktail”

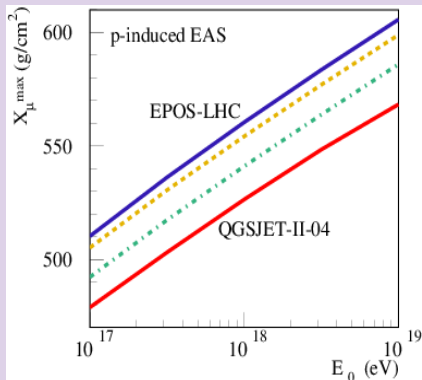
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Constraining pion interactions by cosmic ray data

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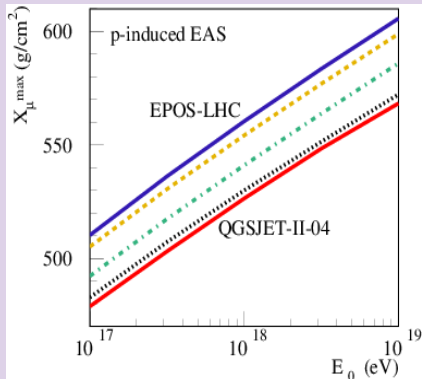
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- now QGSJET-II also for $\bar{p}p$ & $\bar{n}n$ production in π -air – largest effect



Constraining pion interactions by cosmic ray data

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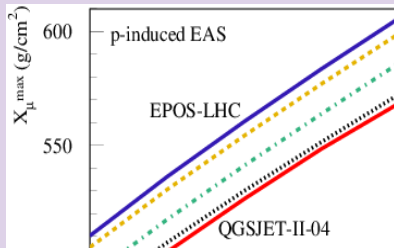
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- small effect: X_{\max}^{μ} difference – due to pion-air collisions
- largest effect: copious $\bar{p}p$ & $\bar{n}n$ production in EPOS
- remaining difference: π^{\pm} & K^{\pm} spectral shapes & diffraction in π - & K -air



Constraining pion interactions by cosmic ray data

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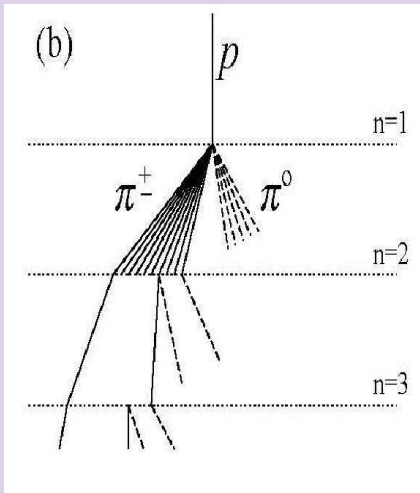


Model-dependence of X_{\max}^{μ} : same features of π -air as for X_{\max}

- X_{\max}^{μ} – even more sensitive!
- \Rightarrow can be used to constrain model approaches
- e.g. copious $\bar{p}p$ & $\bar{n}n$ production and large pion diffraction – disfavored by Auger data

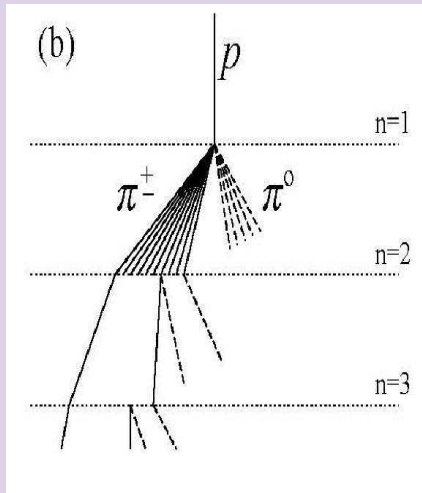
Muon component of air showers & PAO muon excess

- NB: N_μ results from a **multi-step hadron cascade**
 - ~ 1 cascade step per energy decade
- $N_\mu \propto E_0^{\alpha_\mu} = \prod_{i=1}^{\text{int}(\lg E_0)} 10^{\alpha_\mu}$
- each order of magnitude: factor $10^{\alpha_\mu} \simeq 8$ ($\alpha_\mu \simeq 0.9$)



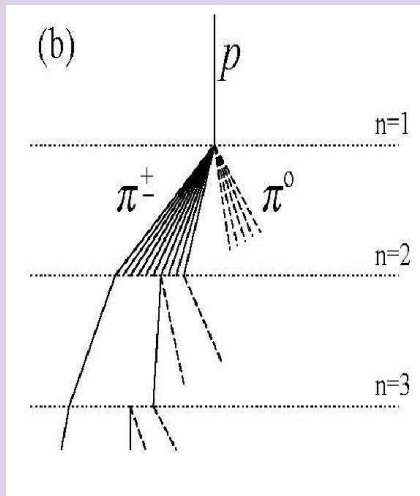
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- each order of magnitude: factor $10^{\alpha_\mu} \simeq 8$ ($\alpha_\mu \simeq 0.9$)
- \Rightarrow higher N_μ requires to change π – air interactions over a wide energy range (see the talk of Jan Ebr)



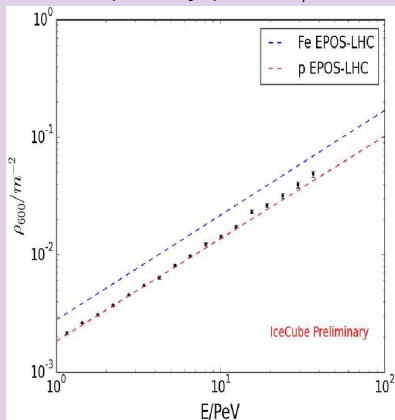
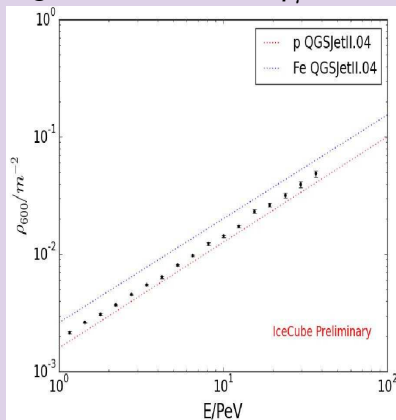
\Rightarrow muon excess will emerge also at lower energies

Muon component of air showers & PAO muon excess

⇒ muon excess will emerge also at lower energies

Problem: muon excess not seen up to 10^{17} eV

E.g., IceCube data on ρ_μ – consistent with primary protons/helium



Muon component of air showers & PAO muon excess

Muon excess produced by 1-2 cascade steps between 10^{17} & 10^{19} ?

- e.g. if we double N^{ch} for the 1st interaction?
 - **< 10% increase for N_μ !** [SO, talk at C2CR, Prague 2005]

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- proton-air cross section at UH energies: $\sigma_{p\text{-air}}^{\text{inel}} \sim 1/2 \text{ b}$
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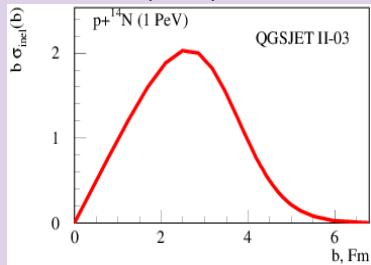
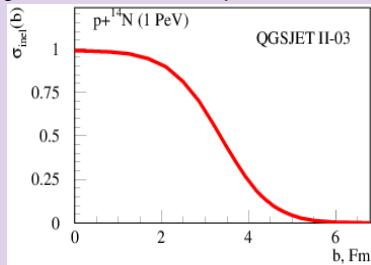
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- proton-air cross section at UH energies: $\sigma_{p\text{-air}}^{\text{inel}} \sim 1/2 \text{ b}$
- to be detected by air shower techniques:
new physics should impact the bulk of interactions
- \Rightarrow to emerge with barn-level cross section
 - presently at LHC: nothing at fb level

Muon component of air showers & PAO muon excess

NB: signals of new physics may be discriminated by PAO

p -air: interaction profile & distribution of the impact parameter b :

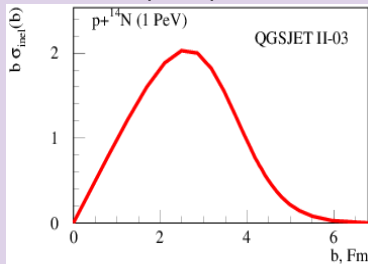
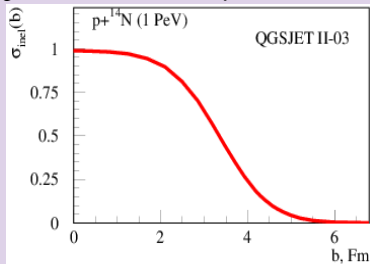


- \Rightarrow interactions dominated by peripheral (large b) collisions
- at large b : low parton density
 - \Rightarrow not suitable for new physics to emerge

Muon component of air showers & PAO muon excess

NB: signals of new physics may be discriminated by PAO

p -air: interaction profile & distribution of the impact parameter b :



Assume new physics to emerge in 10% of most central collisions

- and result in EAS with a factor of 10 higher muon density...
 - \Rightarrow 90% muon excess ($\langle \rho_{\mu} \rangle = 0.1 * 10\rho_{\mu}^{(0)} + 0.9 * \rho_{\mu}^{(0)} = 1.9\rho_{\mu}^{(0)}$)
- \Rightarrow large fluctuations of muon density: $\sigma_{\rho_{\mu}}/\rho_{\mu} \simeq 100\%$
- \Rightarrow can be easily discriminated in PAO data (for usual EAS: $\sigma_{\rho_{\mu}}/\rho_{\mu} \simeq 10 \div 15\%$)

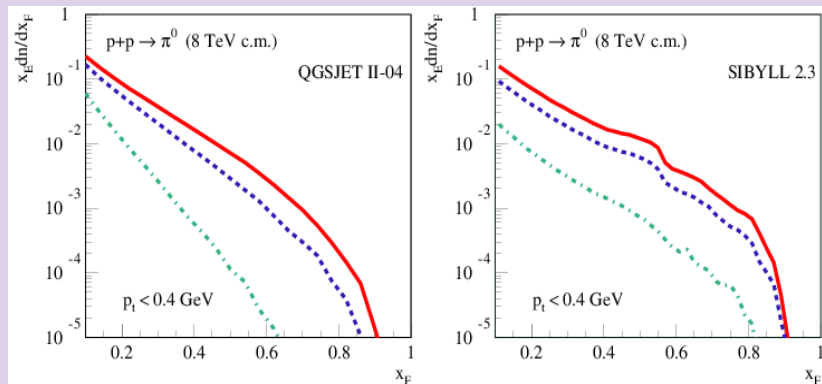
Summary

- 1 LHC studies of pp collisions constrained interaction models
 - most important for CR physics: $\sigma_{pp}^{\text{tot/el}}$ by TOTEM & ATLAS
 - of importance: to resolve the diffraction issue
- 2 Differences for predicted $K_{p\text{-air}}^{\text{inel}}$ ($\Rightarrow X_{\text{max}}$):
model assumptions for constituent parton Fock states
 - can be discriminated by combined measurements with central & forward-looking detectors at the LHC
- 3 Present uncertainties for EAS predictions:
largely due to the treatment of pion-air interactions
 - can be constrained by X_{max}^{μ} measurements in CR experiments
- 4 Present PAO data on X_{max}^{μ} :
disfavor model features which lead to deep X_{max}
- 5 PAO muon excess implies a higher N_{μ} at lower energies
 - more exotic options may be discriminated by studying fluctuations of muon density at ground

Extra slides

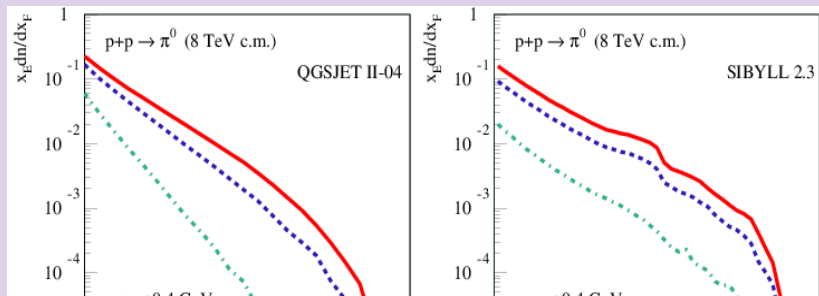
Tests at LHC: correlations of central & forward production

Alternatively, forward π^0 spectra in LHCf for different ATLAS triggers ($\geq 1, 6, 20$ charged hadrons of $p_t > 0.5$ GeV & $|\eta| < 2.5$)



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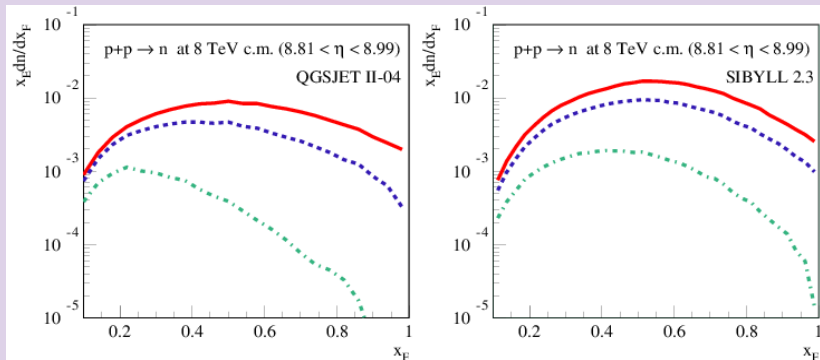


Compare QGSJET-II-04 (left) to SIBYLL 2.3 (right)

- enhanced multiple scattering
⇒ softer pion spectra
- ⇒ violation of limiting fragmentation (energy sharing between constituent partons)
- nearly same spectral shape for all the triggers
- ⇒ perfect limiting fragmentation (central production decoupled)

Tests at LHC: correlations of central & forward production

Neutron spectra in LHCf ($8.99 < \eta < 9.22$) for same triggers



- remarkably universal spectral shape in SIBYLL-2.3 (decoupling of central production)
 - closely related to the small 'inelasticity' of the model
- strong suppression of forward neutrons in QGSJET-II-04
 - higher central activity \Rightarrow more constituent partons involved \Rightarrow less energy left for the proton 'remnant'

σ_{inel} & forward hadron spectra for pion-nitrogen collisions

