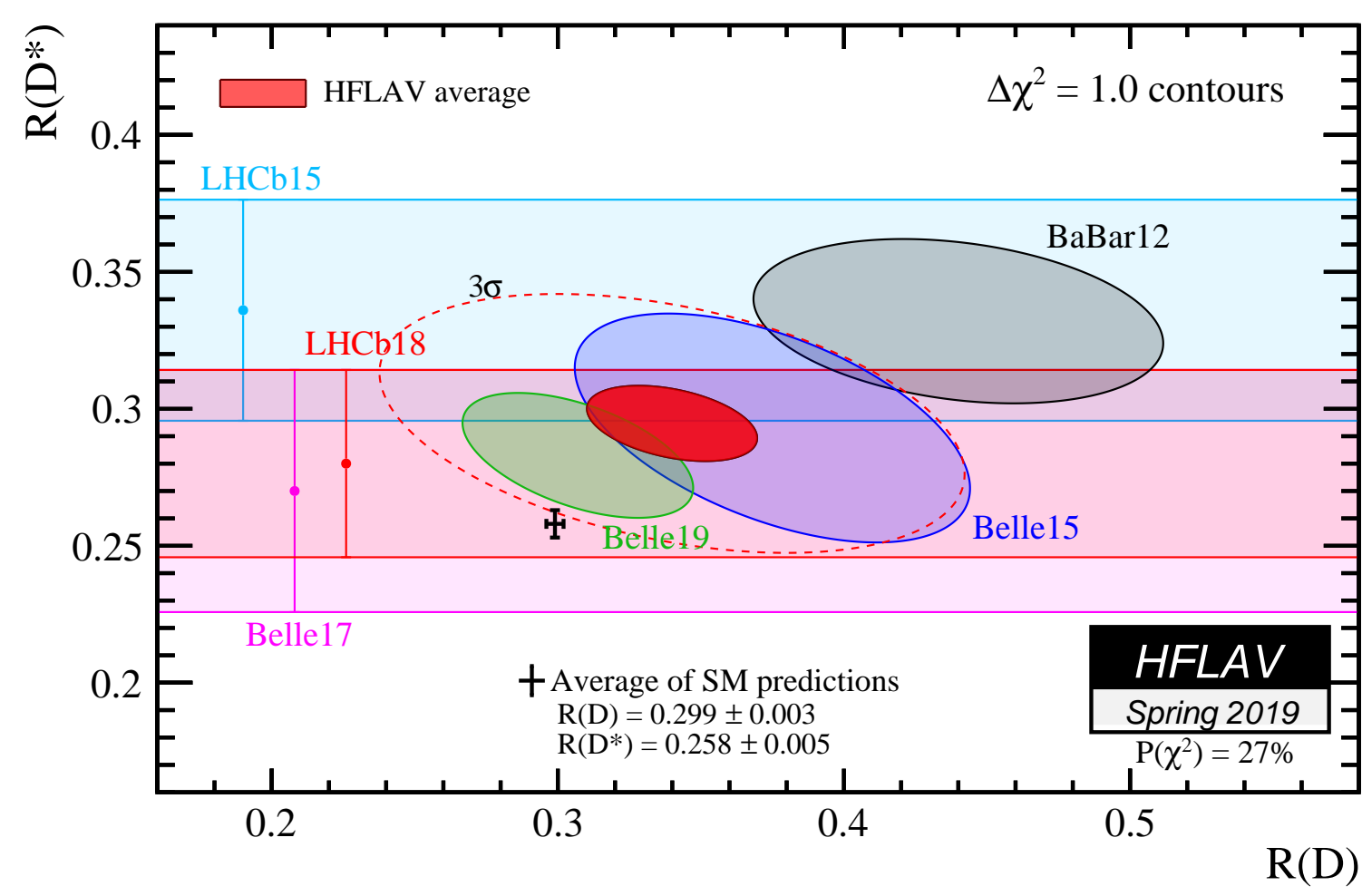


## MEASUREMENTS IN $b \rightarrow c\tau\bar{\nu}$ SECTOR

**Lepton Flavor Universality Violation:** The flavor ratios mediated by  $b \rightarrow c\tau\bar{\nu}$  transitions are

- The ratios  $R_{D^{(*)}} = \Gamma(B \rightarrow D^{(*)}\tau\bar{\nu})/\Gamma(B \rightarrow D^{(*)}l\bar{\nu})$ , ( $l = e, \mu$ ) measured by BaBar, Belle and LHCb experiments [1] indicate evidence of lepton flavor non universality. These results disagree with the Standard Model (SM) prediction at  $4.1\sigma$  level.
- In 2017, LHCb measured  $R_{J/\psi} = \Gamma(B_c \rightarrow J/\psi\tau\bar{\nu})/\Gamma(B_c \rightarrow J/\psi\mu\bar{\nu})$  and found it to differ from the SM prediction by about  $\sim 1.7\sigma$  [2].
- Belle presented updated results on  $R_D-R_{D^*}$  at Moriond 2019. Here, for the first time, they reconstructed the  $\tau$  lepton. These results are consistent with the SM [3]. Combined with earlier results, the discrepancy with the SM reduces to  $3.1\sigma$  which is still significant.



**Angular observables:** The following two angular observables are measured in  $B \rightarrow D^*\tau\bar{\nu}$  decay:

- Belle collaboration has measured the  $\tau$  polarization fraction  $P_\tau(D^*) = (\Gamma_{\lambda_\tau=+1/2} - \Gamma_{\lambda_\tau=-1/2}) / (\Gamma_{\lambda_\tau=+1/2} + \Gamma_{\lambda_\tau=-1/2})$  with a very large statistical error. This measurement is consistent with its SM prediction [4].
- Recently Belle has also measured the longitudinal  $D^*$  polarization fraction  $f_L(D^*) = \Gamma_{\lambda_{D^*}=0} / (\Gamma_{\lambda_{D^*}=0} + \Gamma_{\lambda_{D^*}=1} + \Gamma_{\lambda_{D^*}=-1})$ . The measured value is  $\sim 1.6\sigma$  higher than the SM prediction [5].

## NEW PHYSICS OPERATORS

The effective Hamiltonian for  $b \rightarrow c\tau\bar{\nu}$  (assuming neutrinos are left-handed) at a scale  $\Lambda = 1$  TeV is given by [6]

$$H_{eff} = \frac{4G_F V_{cb}}{\sqrt{2}} \left[ \mathcal{O}_{V_L} + \frac{\sqrt{2}}{4G_F V_{cb} \Lambda^2} \sum_i \left( C_i \mathcal{O}_i + C_i^{(',')} \mathcal{O}_i^{(',')} \right) \right]$$

	Operator	Fierz Transform
$\mathcal{O}_{V_L}$	$(\bar{c}\gamma_\mu P_L b)(\bar{\tau}\gamma^\mu P_L \nu)$	
$\mathcal{O}_{V_R}$	$(\bar{c}\gamma_\mu P_R b)(\bar{\tau}\gamma^\mu P_L \nu)$	
$\mathcal{O}_{S_R}$	$(\bar{c}P_R b)(\bar{\tau}P_L \nu)$	
$\mathcal{O}_{S_L}$	$(\bar{c}P_L b)(\bar{\tau}P_L \nu)$	
$\mathcal{O}_T$	$(\bar{c}\sigma^{\mu\nu} P_L b)(\bar{\tau}\sigma_{\mu\nu} P_L \nu)$	
$\mathcal{O}'_{V_L}$	$(\bar{\tau}\gamma_\mu P_L b)(\bar{c}\gamma^\mu P_L \nu)$	$\mathcal{O}_{V_L}$
$\mathcal{O}'_{V_R}$	$(\bar{\tau}\gamma_\mu P_R b)(\bar{c}\gamma^\mu P_L \nu)$	$-2\mathcal{O}_{S_R}$
$\mathcal{O}'_{S_R}$	$(\bar{\tau}P_R b)(\bar{c}P_L \nu)$	$-\frac{1}{2}\mathcal{O}_{V_R}$
$\mathcal{O}'_{S_L}$	$(\bar{\tau}P_L b)(\bar{c}P_L \nu)$	$-\frac{1}{2}\mathcal{O}_{S_L} - \frac{1}{8}\mathcal{O}_T$
$\mathcal{O}'_T$	$(\bar{\tau}\sigma^{\mu\nu} P_L b)(\bar{c}\sigma_{\mu\nu} P_L \nu)$	$-6\mathcal{O}_{S_L} + \frac{1}{2}\mathcal{O}_T$
$\mathcal{O}''_{V_L}$	$(\bar{\tau}\gamma_\mu P_L c^c)(\bar{b}^c\gamma^\mu P_L \nu)$	$-\mathcal{O}_{V_R}$
$\mathcal{O}''_{V_R}$	$(\bar{\tau}\gamma_\mu P_R c^c)(\bar{b}^c\gamma^\mu P_L \nu)$	$-2\mathcal{O}_{S_R}$
$\mathcal{O}''_{S_R}$	$(\bar{\tau}P_R c^c)(\bar{b}^c P_L \nu)$	$\frac{1}{2}\mathcal{O}_{V_L}$
$\mathcal{O}''_{S_L}$	$(\bar{\tau}P_L c^c)(\bar{b}^c P_L \nu)$	$-\frac{1}{2}\mathcal{O}_{S_L} + \frac{1}{8}\mathcal{O}_T$
$\mathcal{O}''_T$	$(\bar{\tau}\sigma^{\mu\nu} P_L c^c)(\bar{b}^c\sigma_{\mu\nu} P_L \nu)$	$-6\mathcal{O}_{S_L} - \frac{1}{2}\mathcal{O}_T$

## METHODOLOGY

We perform a  $\chi^2$  fit to the five observables  $R_D, R_{D^*}, R_{J/\psi}, P_\tau(D^*)$  and  $f_L(D^*)$ . The fit is done by using the CERN minimization code MINUIT. The  $\chi^2$  is defined as

$$\chi^2(C_i) = \sum (O^{th}(C_i) - O^{exp})^T \mathcal{C}^{-1} (O^{th}(C_i) - O^{exp}),$$

where  $\mathcal{C}$  is the covariance matrix.

We perform three types of fits:

- Fit-A** taking only one NP operator at a time.
- Fit-B** taking two similar NP operators at a time.
- Fit-C** taking two dissimilar NP operators at a time.

For each of the above fits, we allow only those NP solutions for which  $\chi^2_{min} \lesssim 5$  as well as  $\mathcal{B}(B_c \rightarrow \tau\bar{\nu}) < 0.1$  [7].

## FIT RESULTS BEFORE MORIOND'19

In this global fit [8], we take all the data which were available before the recent  $R_D-R_{D^*}$  results announced by Belle at Moriond'19.

NP type	Best fit value(s)	$\chi^2_{min}$
SM	$C_i = 0$	24.70
$C_{V_L}$	$0.15 \pm 0.03$	5.1
$C''_{S_L}$	$-0.52 \pm 0.10$	5.2
$(C_{V_L}, C_{V_R})$	$(0.17, 0.05)$	4.5
$(C'_{V_L}, C'_{V_R})$	$(0.12, -0.06)$	4.2
$(C''_{S_L}, C''_{S_R})$	$(-0.26, 0.16)$	5.0
$(C_{V_L}, C_{S_L})$	$(0.14, 0.09)$	4.5
$(C'_{V_L}, C'_{S_L})$	$(0.17, -0.12)$	4.5
$(C_{V_R}, C_{S_L})$	$(-0.17, 0.42)$	4.6

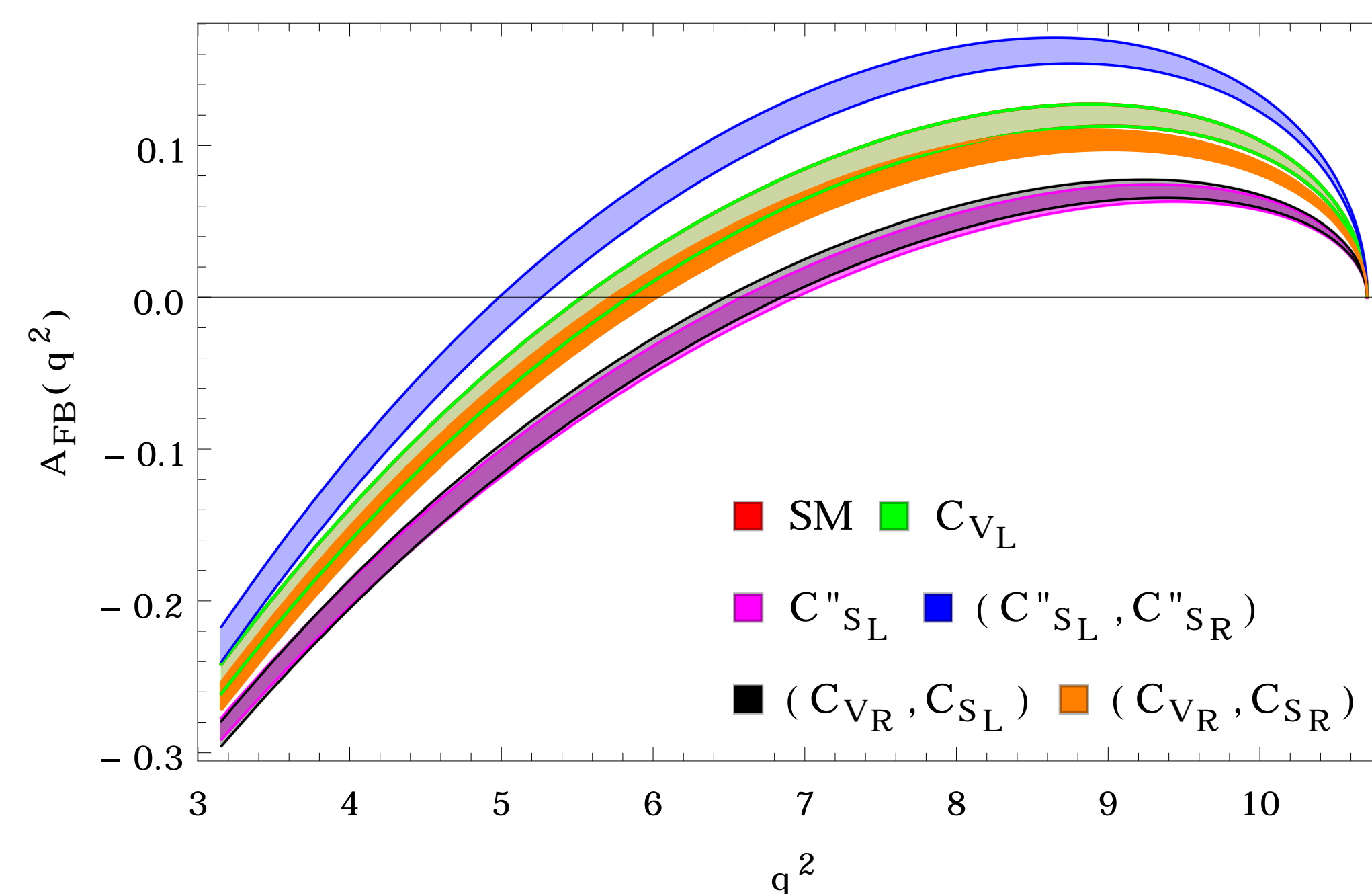
NOTE: The tensor solution [9, 10] which was allowed before the measurement of  $D^*$  polarization fraction, is now ruled out at the level of  $5\sigma$  [8].

## METHODS TO DISCRIMINATE THE NP OPERATORS

Angular observables are the standard and powerful tools to discriminate between the NP operators [11].

NP type	$P_\tau(D^*)$	$f_L(D^*)$	$A_{FB}(D^*)$	$q^2[A_{FB}(q^2) = 0] \text{ GeV}^2$	$\mathcal{B}(B_c \rightarrow \tau\bar{\nu})$
SM	$-0.499 \pm 0.004$	$0.45 \pm 0.04$	$-0.011 \pm 0.007$	5.7	$2.15 \times 10^{-2}$
$C_{V_L}$	$-0.499 \pm 0.004$	$0.46 \pm 0.04$	$-0.011 \pm 0.007$	5.7	$2.50 \times 10^{-2}$
$C''_{S_L}$	$-0.493 \pm 0.003$	$0.44 \pm 0.05$	$-0.062 \pm 0.010$	6.8	$1.14 \times 10^{-6}$
$(C''_{S_L}, C''_{S_R})$	$-0.494 \pm 0.005$	$0.47 \pm 0.04$	$0.027 \pm 0.008$	5.0	$7.93 \times 10^{-2}$
$(C_{V_R}, C_{S_L})$	$-0.526 \pm 0.004$	$0.45 \pm 0.04$	$-0.061 \pm 0.006$	6.7	$2.23 \times 10^{-3}$
$(C_{V_R}, C_{S_R})$	$-0.468 \pm 0.005$	$0.47 \pm 0.04$	$-0.023 \pm 0.006$	5.8	0.12

NOTE: Neither the  $\tau$  nor the  $D^*$  polarization fraction has any capability to distinguish between the allowed solutions listed above. The zero crossing point of the forward-backward asymmetry  $A_{FB}(q^2)$  and the branching ratio  $\mathcal{B}(B_c \rightarrow \tau\bar{\nu})$  together can uniquely identify each NP operator.



- The  $A_{FB}$  is defined as  $A_{FB}(q^2) = \frac{1}{d\Gamma/dq^2} \left[ \int_0^1 \frac{d^2\Gamma}{dq^2 d\cos\theta_\tau} d\cos\theta_\tau - \int_{-1}^0 \frac{d^2\Gamma}{dq^2 d\cos\theta_\tau} d\cos\theta_\tau \right]$ .
- The bands of  $A_{FB}(q^2)$  for the SM and  $\mathcal{O}_{V_L}$  solution are exactly overlapped with each other.
- The zero crossing points of  $A_{FB}(q^2)$  are:
  - $\sim 5 \text{ GeV}^2$  for  $(C''_{S_L}, C''_{S_R})$  solution,
  - $\sim 5.8 \text{ GeV}^2$  for  $C_{V_L}$  and  $(C_{V_R}, C_{S_R})$  solutions,
  - $\sim 6.8 \text{ GeV}^2$  for  $C''_{S_L}$  and  $(C_{V_R}, C_{S_L})$  solutions.
- $\mathcal{B}(B_c \rightarrow \tau\bar{\nu}) \sim 2\%$  for  $C_{V_L}$  solution whereas it is  $> 10\%$  for  $(C_{V_R}, C_{S_R})$  solution.
- It is more difficult to distinguish between  $C''_{S_L}$  and  $(C_{V_R}, C_{S_L})$  solutions. It requires the measurement of  $\mathcal{B}(B_c \rightarrow \tau\bar{\nu})$  at the level of  $10^{-3}$ .

## SUMMARY

- Present data in  $b \rightarrow c\tau\bar{\nu}$  sector implies that five distinct NP solutions with different Lorentz structures are allowed.
- As the tension in  $R_D-R_{D^*}$  has been come down to  $3.1\sigma$ , the coefficient of  $\mathcal{O}_{V_L}$  solution is now reduced by  $\sim 33\%$ .
- The operator  $\mathcal{O}_{V_L}$  has the same Lorentz structure as the SM. Therefore, its predictions for all the angular observables are the same as those of the SM.
- The zero crossing point of  $A_{FB}(q^2)$  and  $\mathcal{B}(B_c \rightarrow \tau\bar{\nu})$  are powerful tools to distinguish between the five solutions.
- To measure  $A_{FB}$ , the reconstruction of  $\tau$  momentum is necessary. This is difficult but this challenge should be taken up.
- Also  $B_c \rightarrow \tau\bar{\nu}$  is quite sensitive to NP. In present and future experiments, it should be tested with a good accuracy.

## REFERENCES

- [1] Heavy Flavor Averaging Group, <https://hflav.web.cern.ch>.
- [2] R. Aaij *et al.* [LHCb Collaboration], Phys. Rev. Lett. **120** (2018) no.12, 121801
- [3] A. Abdesslem *et al.* [Belle Collaboration], arXiv:1904.08794
- [4] S. Hirose *et al.* [Belle Collaboration], Phys. Rev. Lett. **118** (2017) no.21, 211801
- [5] K. Adamczyk [Belle and Belle-II Collaborations], arXiv:1901.06380 [hep-ex].
- [6] M. Freytsis, Z. Ligeti and J. T. Ruderman, Phys. Rev. D **92** (2015) no.5, 054018
- [7] A. G. Akeroyd and C. H. Chen, Phys. Rev. D **96** (2017) no.7, 075011
- [8] A. K. Alok, D. Kumar, S. Kumbhakar and S. Uma Sankar, arXiv:1903.10486 [hep-ph].
- [9] A. K. Alok, D. Kumar, S. Kumbhakar and S. Uma Sankar, Phys. Rev. D **95** (2017) no.11, 115038
- [10] A. K. Alok, D. Kumar, J. Kumar, S. Kumbhakar and S. Uma Sankar, JHEP **1809** (2018) 152
- [11] A. K. Alok, D. Kumar, S. Kumbhakar and S. Uma Sankar, Phys. Lett. B **784** (2018) 16