Particle Discovery Opportunities at the International Linear Collider

EPS-HEP 2019, Ghent

Moritz Habermehl¹ on behalf of the LCC Physics Working Group

12 July 2019

¹Deutsches Elektronen-Synchrotron DESY, Germany

CLUSTER OF EXCELLENCE

QUANTUM UNIVERSE







The International Linear Collider

A planned electron-positron collider

Lepton colliders are complementary to the LHC

- cleaner environment, controlled initial state
- coupling to leptons is tested

Advantages of the ILC over other planned electron-positron colliders

- mature technology
- centre-of-mass energy can be tuned and increased:
 250 GeV in initial stage, upgrades to 500 GeV and 1 TeV
- polarisation of both beams: $P(e)=\pm 80\%$, $P(e+)=\mp 30\%$
- triggerless operation
- hermeticity of detector down to lowest angles

Main Linac

Damping Rings

ILD

The BSM Physics Programme of the ILC

Large range of detailed studies



The Potential of the ILC for Discovering new Particles \rightarrow focus on 500 GeV

arxiv:1903.01629

The International Linear Collider: A Global Project → focus on 250 GeV

The BSM Physics Programme of the ILC

Large range of detailed studies



The BSM Physics Programme of the ILC

Large range of detailed studies



Searches for WIMP Dark Matter

WIMPs = Weakly Interacting Massive Particles

Interplay between search channels

- direct and indirect detection
- collider searches



Searches for WIMP Dark Matter

(GeV)

WIMPs = Weakly Interacting Massive Particles

Interplay between search channels

- direct and indirect detection
- collider searches

Example: singlet-like fermion WIMP

- likelihood analysis of Planck, PICO-2L LUX, XENON100, LEP, LHC, plus LZ, PICO250 projections
- surviving region assuming no WIMP signals are detected
 → can be tested at the ILC
- framework: effective operators
 - Λ energy scale of new physics
 - valid: testable energies $\gg \sqrt{s}$



WIMPs in the Mono-Photon Channel

e

 e^+

つ

Generic collider search

- WIMP pair production
- with a photon from initial state radiation
- single photon in an "empty" detector

WIMPs in the Mono-Photon Channel

е

 e^+

Generic collider search

- WIMP pair production
- with a photon from initial state radiation
- single photon in an "empty" detector

Example: vector-like fermion WIMP

full detector study at 500 GeV

- carefull ISR modelling
 - several photons possible
 - double-counting avoided
- ILC accelerator environment
 - luminosity spectrum
 - beam-induced background

effective operators

- vector
- axial-vector
- scalar



WIMPs in the Mono-Photon Channel

е

 e^+

Generic collider search

- WIMP pair production
- with a photon from initial state radiation
- single photon in an "empty" detector

Example: vector-like fermion WIMP

full detector study at 500 GeV

- carefull ISR modelling
 - several photons possible
 - double-counting avoided
- ILC accelerator environment
 - luminosity spectrum
 - beam-induced background

effective operators

- vector
- axial-vector
- scalar





DESY. | Particle Discovery Opportunities at the ILC | M. Habermehl, 12 July 2019

Sensitivity to WIMP Dark Matter: Operation Scenarios

Role of \sqrt{s} , luminosity and polarisation

Polarisation is crucial

- suppression of neutrino background ⁹/₉
- enhancement of signal
- provides statistically independent data sets with different polarisation configurations
 - \rightarrow control of systematic uncertainties



Sensitivity to WIMP Dark Matter: Operation Scenarios

Role of \sqrt{s} , luminosity and polarisation

Polarisation is crucial

- suppression of neutrino background
- enhancement of signal
- provides statistically independent data sets with different polarisation configuraitons
 - \rightarrow control of systematic uncertainties



Sensitivity to WIMP Dark Matter: Operation Scenarios

Role of \sqrt{s} , luminosity and polarisation

Polarisation is crucial

- suppression of neutrino background
- enhancement of signal
- provides statistically independent data sets with different polarisation configuraitons
 - \rightarrow control of systematic uncertainties

Higher centre-of-mass energies are favored

- signal cross-section rises with energy
- collecting data for decades at a moderate energy is worse than running at higher energies and collect a smaller data set



Detector Hermeticity

Background suppression crucial for BSM studies

Bhabha scattering: background for WIMPs

- huge cross-section
- particles mainly in forward region
 → challenging reconstruction
 → mono-photon signal can be mimicked

Overlay

- pairs produced from beamstrahlung photons
- forward region is polluted

Hermeticity study: WIMPs as example

- sensitivity depends on number of reconstructed Bhabha scattering events
- with a larger "blind region" around beam pipes: testable energy reach is significantly reduced

forward detector BeamCal at 3.6 m to IP



Dark photon search

 $\textbf{e+e-} \rightarrow \gamma A \text{`} \rightarrow \gamma \mu \text{+} \mu \text{-}$

95% exclusion limit (no deviation from SM observed)

- amplitute ε: mixing of dark photon with hypercharge gauge boson
- assumes full ILC programme
 → shown mass range dominated by run at 250 GeV



Dark photon search

 $\textbf{e+e-} \rightarrow \gamma A \text{`} \rightarrow \gamma \mu \text{+} \mu \text{-}$

95% exclusion limit (no deviation from SM observed)

- amplitute ε: mixing of dark photon with hypercharge gauge boson
- assumes full ILC programme
 → shown mass range dominated by run at 250 GeV

Role of polarisation

- polarisation not considered in study
 → sensitivity around Z mass could improve
- for invisible decay products: polarisation crucial to suppress SM neutrino background



Additional Higgs Bosons

Search for extra light scalars

Recoil against Z

- lepton collider: initial state 4-mom. is known
- Z reconstructed from its decay products • \rightarrow scalar can be reconstructed as the missing 4-mom.
- reconstruction of the scalar without "looking at the scalar" • \rightarrow highly model-independent

Exclusion limit

- $sin^{2}(\theta)$: cross-section relative to a SM Higgs with correspoding mass
- full detector simulation •
 - \rightarrow for low masses: detector effects play a role



Gev

 $\mathsf{M}_{\mathsf{c}^0}$

200

Supersymmetry

UV-complete BSM models

Assumption: discovery of several new particles

- nature of new physics can be tested
- conclusions about underlying parameters

Detector simulations of the complete model

- complete particle spectrum
- model branching ratios (instead of just assuming 100% for dominant channel)



Supersymmetry

UV-complete BSM models

Assumption: discovery of several new particles

- nature of new physics can be tested
- conclusions about underlying parameters

Detector simulations of the complete model

- complete particle spectrum
- model branching ratios (instead of just assuming 100% for dominant channel)

Loop-hole free searches

arxiv:1308.1461

GeV

250

150

100

50

0

^{ds}] 200

- next-to-lightest SUSY particle searches
- in kinematic reach: discovery guaranteed or immutable limits
- in case of a discovery: accurate measurements



SUSY: Precision Measurements

Fitting the dark matter relic density $\boldsymbol{\varOmega}$

Stau coannihilation scenario

- lightest supersymmetric particle (LSP) is the dark matter
- study at 500 GeV

arxiv:1508.04383

• $\widetilde{\tau_1} \rightarrow \tau \, \widetilde{\chi_1^0}$ endpoint \rightarrow stau mass precision 0.15%



SUSY: Precision Measurements

Fitting the dark matter relic density Ω

Stau coannihilation scenario

- lightest supersymmetric particle (LSP) is *the* dark matter •
- study at 500 GeV

arxiv:1508.04383

• $\widetilde{\tau_1} \rightarrow \tau \widetilde{\chi_1^0}$ endpoint \rightarrow stau mass precision 0.15%

With SUSY observables as input dark matter relic density Ω can be fitted

- using micrOMEGAs (arxiv:1305.0237) •
- input: discovery of all sleptons, sneutrinos, $\tilde{\chi}_1^0$, $\tilde{\chi}_2^0$, $\tilde{\chi}_1^{\pm}$. with permille mass precision
- fit precision $2\% \rightarrow$ same precision as Planck •
- clear identification of LSP as dark matter constituent



SUSY: Precision Measurements II

Testing gaugino mass unification

talk by S. Sasikumar

Natural SUSY with light Higgsinos

- not excluded natural SUSY model
- mirage mediated breaking model
 → unification below GUT scale

Running gaugino masses

- gaugino masses are fitted (at 1 TeV)
 - using Fittino (arxiv:hep-ph/0412012)
 - input: Higgsino masses and polarised cross-sections
- evolving gaugino masses to higher scales
 - using renormalization group equations
 - bands show 1σ confidence levels
- unification at GUT scale can be excluded from fit



Conclusion

new particle that couple prefentially to leptons

→ chirality of new process
 can be tested

The ILC has the potential to find ...







polarisation: control of systematics



new particles



$\sum_{0}^{5000} \sum_{10^{10}}^{10^{10}} \frac{1}{10^{10}} \frac{1}{10$

new physics

loophole-free searches

Conclusion

new particle that couple prefentially to leptons

→ chirality of new process
 can be tested

The ILC has the potential to find ...



precision measurements



polarisation: control of systematics



new particles





new physics

loophole-free searches

Lepton colliders cover unique parameter space -> complementary to the LHC

Polarisation of both beams is crucial for many searches -> ILC

Contact

DESY. Deutsches Elektronen-Synchrotron Moritz Habermehl FLC moritz.habermehl@desy.de

www.desy.de

SUSY: Testing Gaugino Mass Unification

Natural SUSY with light Higgsinos

Running gaugino masses

- gaugino masses are fitted (at 1 TeV)
 - using Fittino (arxiv:hep-ph/0412012)
 - with Higgsino masses and polarised cross-sections as input
- evolving gaugino masses to higher scales
 - using renormalisation group equations
 - bands show 1σ confidence levels

Case 2: unification below GUT scale

- unification at GUT scale can be excluded from fit
- good precision with 10% measurement of gluino mass from HL-LHC





talk by S. Sasikumar



Case 1: unification at GUT scale

- running of M1 and M2
- unification at GUT scale can be fitted
- running of M3 from unification to lower energies
 - matches model point at 1 TeV
 - could be cross-checked with LHC results
 Page 26

BACKUP: Dark photon search I

Limits below 250 GeV dominated by 250 GeV run

approach

- muon momentum resolution taken from full detector simulation → propagated to mass peak width
- energy dep. BG cross-section from generator level
- around Z mass
 - shape taken from CEPC study (arxiv:1503.07209)





BACKUP: Dark photon search II

Comparison to other experiments



Page 28

BACKUP: Additional Higgs Bosons

Search for extra light scalars

arxiv:1902.06118, arxiv:1801.08164

 black curve: update with 500 GeV and IDR detector model





BACKUP: Supersymmetry

UV-complete BSM models

Assumption: discovery of several new particles

- nature of new physics can be tested .
- conclusions about underlying parameters •

Detector simulations of the complete model

- complete particle spectrum .
- model branching ratios (instead of just . assuming 100% for dominant channel)

Loop-hole free searches

arxiv:1308.1461

[GeV]

- next-to-lightest SUSY particle searches .
- in kinematic reach: • discovery guaranteed or immutable limits
- in case of a discovery: accurate measurements .



DESY. | Particle Discovery Opportunities at the ILC | M. Habermehl, 12 July 2019