



A new way to search for QCD Axion Dark Matter with a Dielectric Haloscope

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for the MADMAX collaboration

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RWTH Aachen University, Aachen, Germany

DESY, Hamburg, Germany

University of Zaragoza, Spain



MAX-PLANCK-GESELLSCHAFT



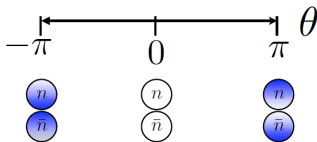
Max-Planck-Institut für Physik
(Werner-Heisenberg-Institut)

The Strong CP-Problem

QCD allows for a term

$$\mathcal{L} = -\theta \frac{g_s}{32\pi^2} G_{\mu\nu}^a \tilde{G}_a^{\mu\nu}, \quad \theta = -\pi \dots \pi$$

but experimentally: $|\theta| < 10^{-10}$ (neutron electric dipole moment)

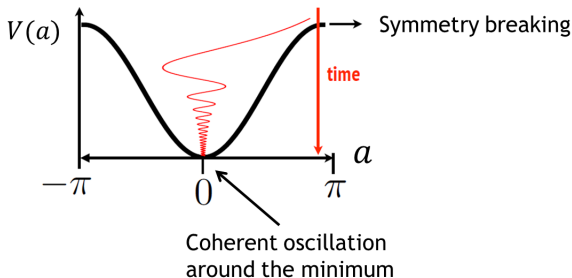


The Strong CP-Problem

make θ a dynamic field: $\theta \rightarrow a(t; \mathbf{x})$ (Peccei-Quinn 1977)

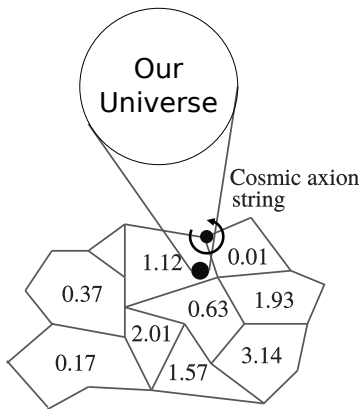
$$\mathcal{L} = - a \frac{g_S}{32\pi^2} G_{\mu\nu}^a \tilde{G}_a^{\mu\nu} + \frac{1}{2} \partial_\mu a \partial^\mu a$$

rolldown to CP conserving limit:



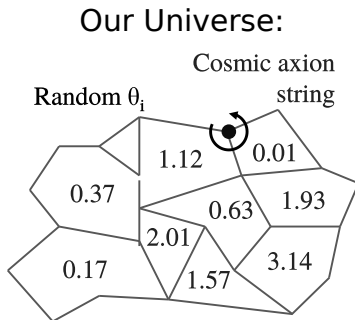
Peccei-Quinn Symmetry Breaking...

before inflation:



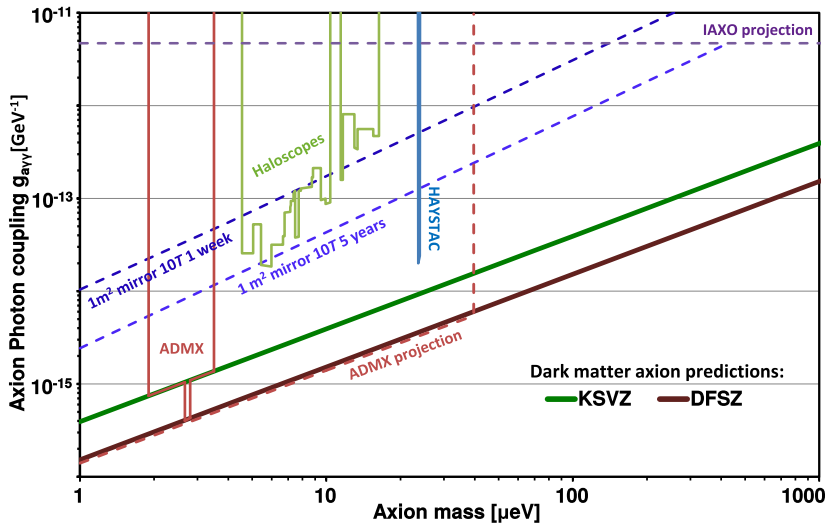
$$m_A \lesssim \text{meV}$$

after inflation:

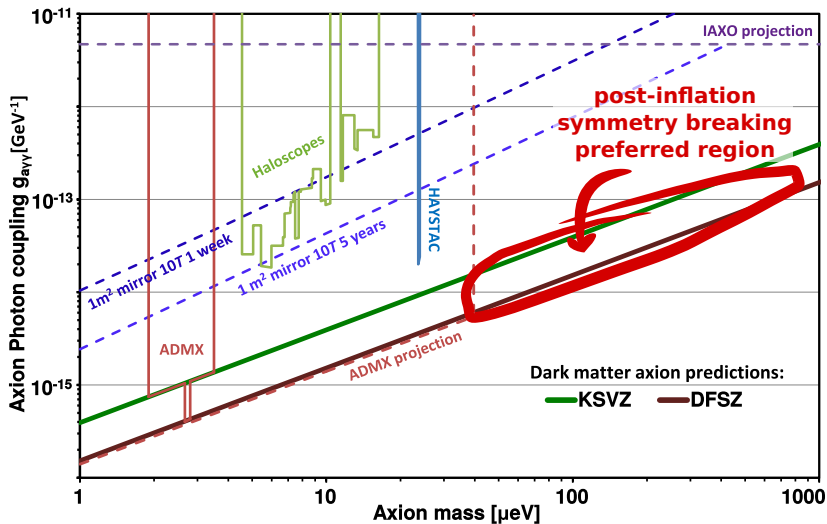


$$m_A \sim 100 \mu\text{eV}$$

The Axion - Parameterspace



The Axion - Parameterspace



Axion Electrodynamics

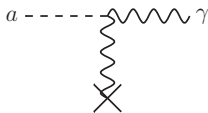
$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} - j^\mu A_\mu + \frac{1}{2}\partial_\mu a \partial^\mu a - \frac{1}{2}m_a^2 a^2 - \frac{g_{a\gamma}}{4}a F_{\mu\nu}\tilde{F}^{\mu\nu}$$

Solve EOM under external magnetic field \mathbf{B}_e :

$$\epsilon \nabla \cdot \mathbf{E} = \rho - g_{a\gamma} \mathbf{B}_e \cdot \nabla a$$

$$\nabla \times \mathbf{H} - \dot{\mathbf{E}} = \mathbf{J} + g_{a\gamma} \mathbf{B}_e \dot{a}$$

$$\ddot{a} - \nabla^2 a + m_a^2 a = g_{a\gamma} \mathbf{E} \cdot \mathbf{B}_e$$

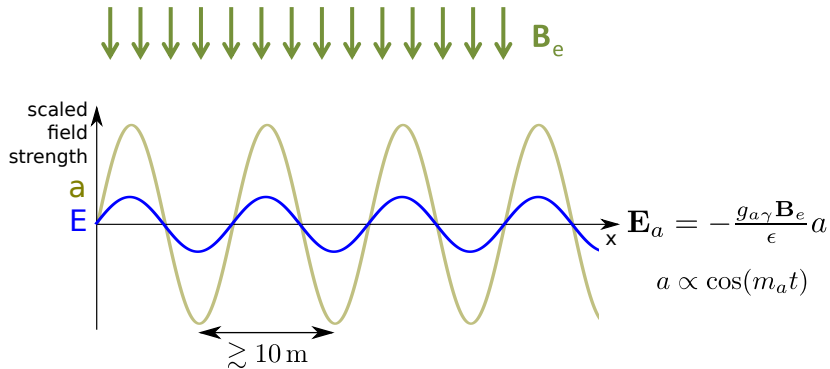


Primakoff
process

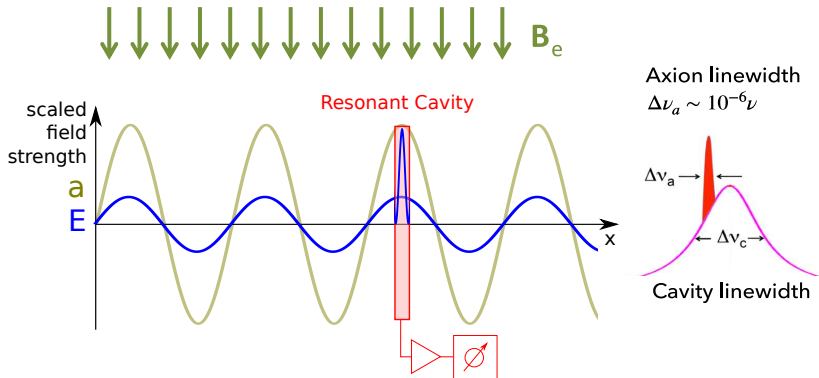
Axion induced electric field:

$$\mathbf{E}_a = -\frac{g_{a\gamma} \mathbf{B}_e}{\epsilon} a = 1.3 \times 10^{-12} \text{ V m}^{-1} \times \left(\frac{B_e}{10 \text{ T}} \right) \frac{C_{a\gamma} f_{DM}^{1/2}}{\epsilon}$$

Axion - Photon Mixing



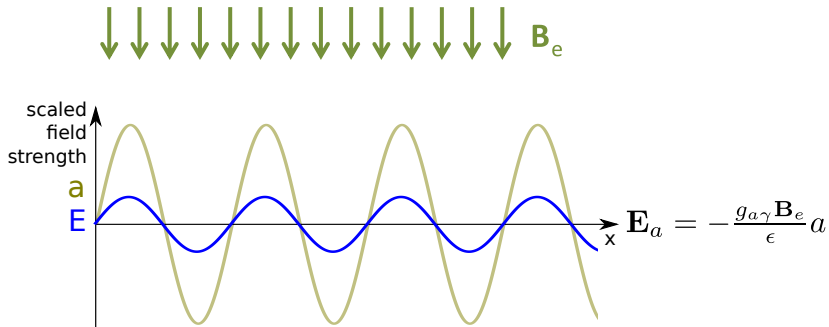
Axion - Photon Mixing



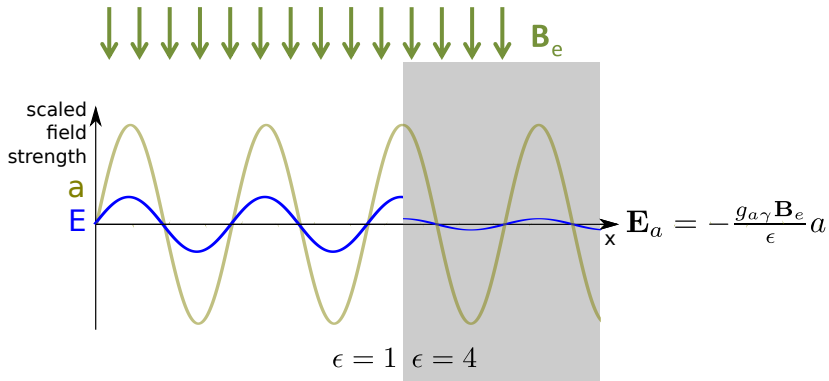
$$P_{\text{sig}} = (B^2 Q V C_{nml}) (g_{a\gamma\gamma}^2 m_a \rho_a)$$

Q : Quality Factor, C_{nml} : mode factor

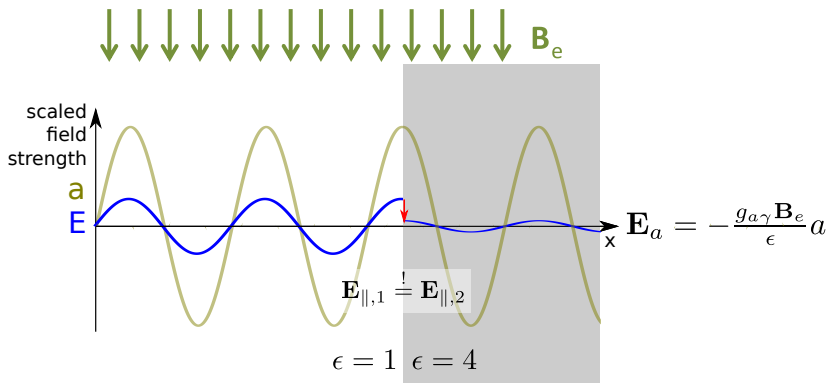
Axion - Photon Mixing



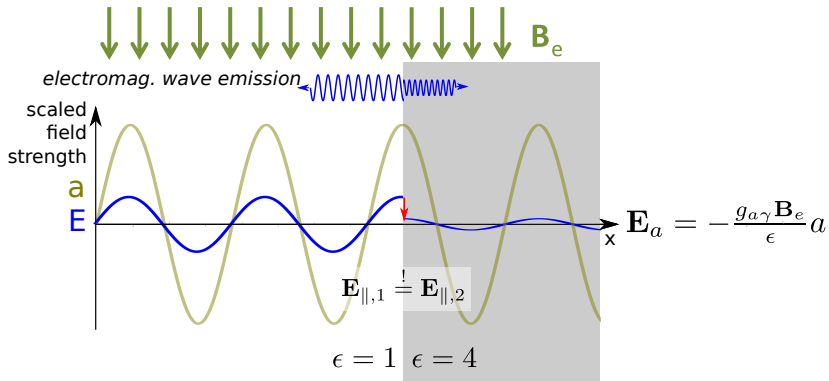
Axion - Photon Mixing



Axion - Photon Mixing

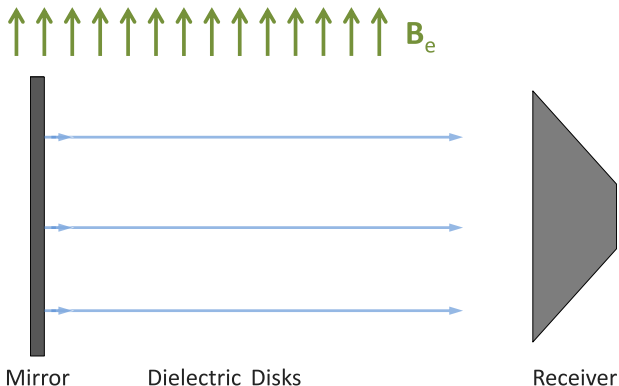


Axion - Photon Mixing



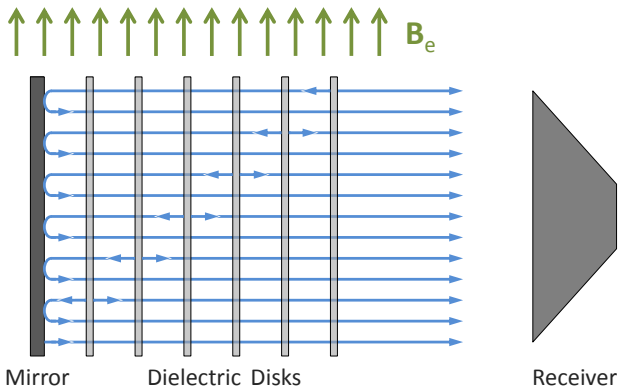
$$P/A = 2.2 \times 10^{-27} \text{ W m}^{-2} \left(\frac{B_e}{10 \text{ T}} \right)^2 C_{a\gamma}^2 \cdot f(\epsilon_1, \epsilon_2)$$

The MADMAX Idea



$$P/A = 2.2 \times 10^{-27} \text{ W m}^{-2} \left(\frac{B_e}{10 \text{ T}} \right)^2 C_{a\gamma}^2 \cdot 1$$

The MADMAX Idea

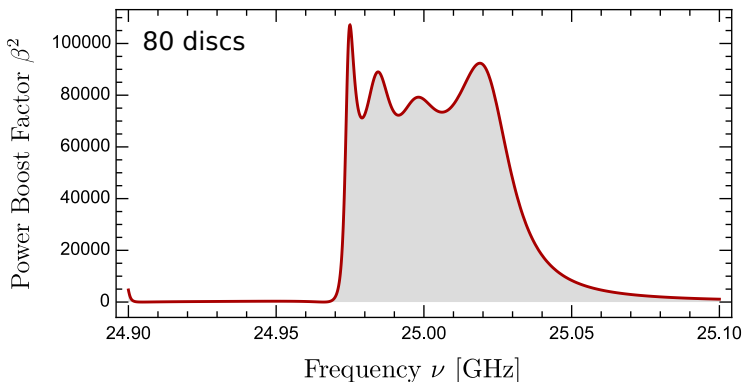


$$P/A = 2.2 \times 10^{-27} \text{ W m}^{-2} \left(\frac{B_e}{10 \text{ T}} \right)^2 C_{a\gamma}^2 \cdot \beta^2$$

β^2 : power emitted by booster / power emitted by single mirror ($\epsilon = \infty$)

Power Boost Factor β^2

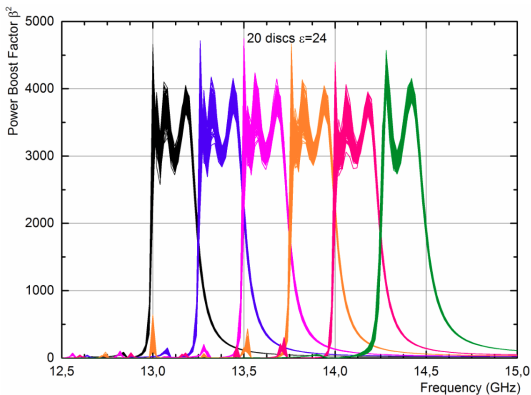
adjust disc spacings:



Wide Bandwidth Boost Factor of $10^4 - 10^5$ Possible

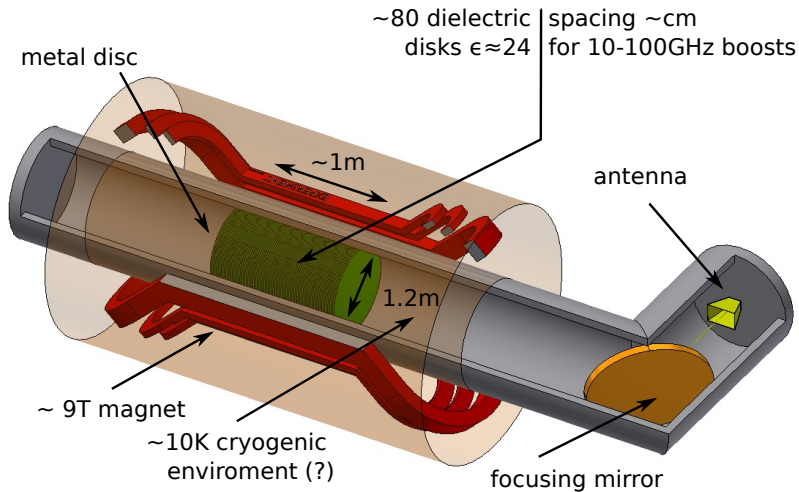
Power Boost Factor β^2

adjust disc spacings:

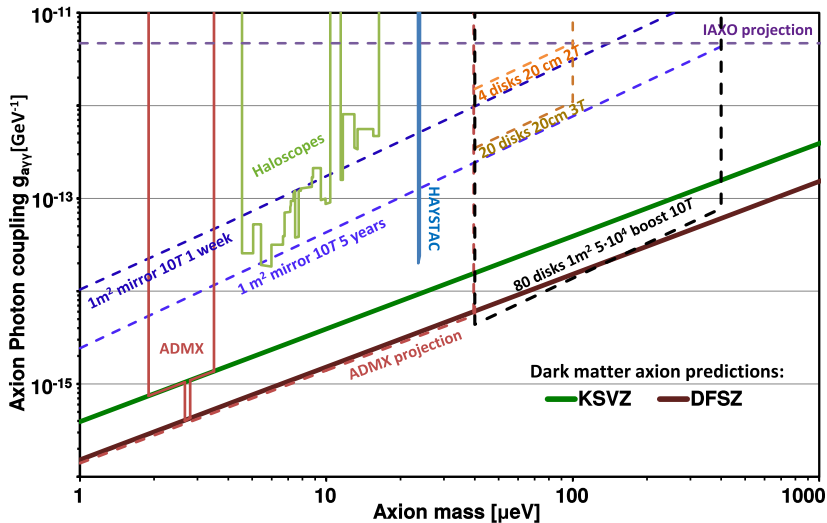


Frequency Band Tunable

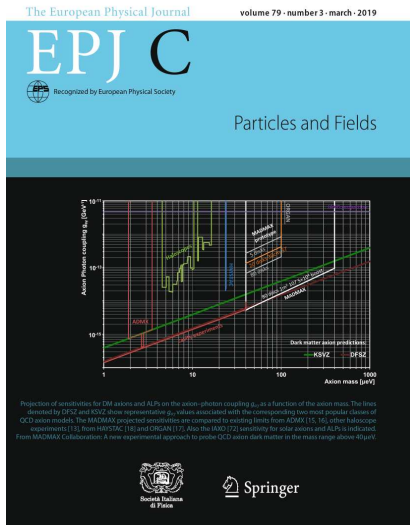
The Vision



Sensitivity

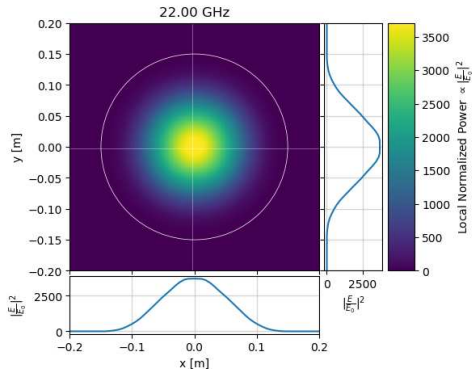
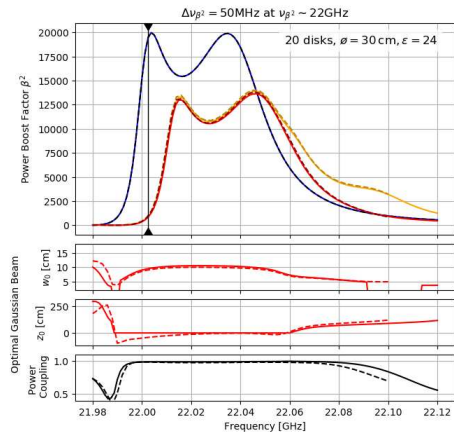


Sensitivity



3D Simulations – Beam Shape

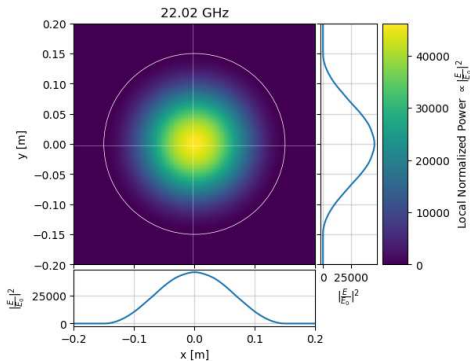
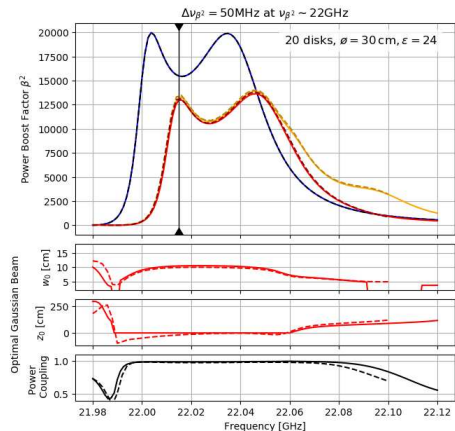
methods see arXiv:1906.02677



well matched to Gaussian beam $w_0 \approx 10\text{ cm}$

3D Simulations – Beam Shape

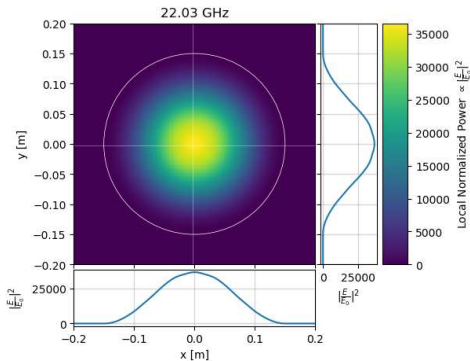
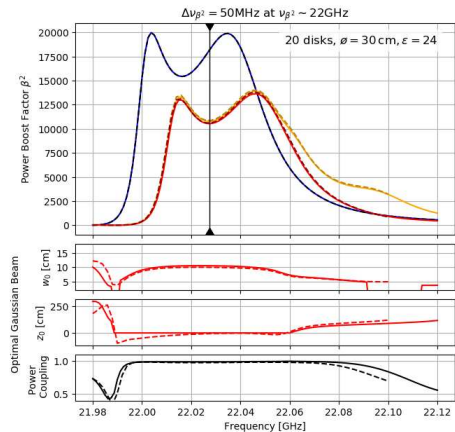
methods see arXiv:1906.02677



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3D Simulations – Beam Shape

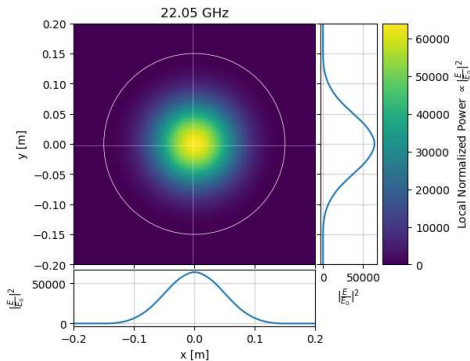
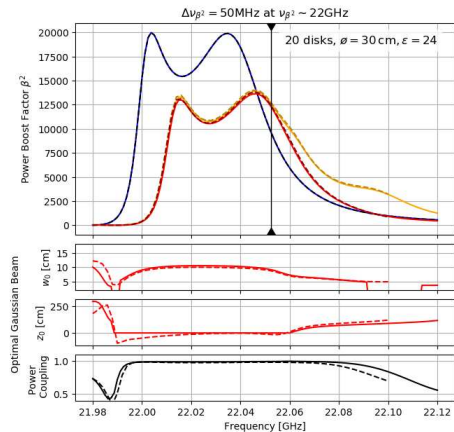
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3D Simulations – Beam Shape

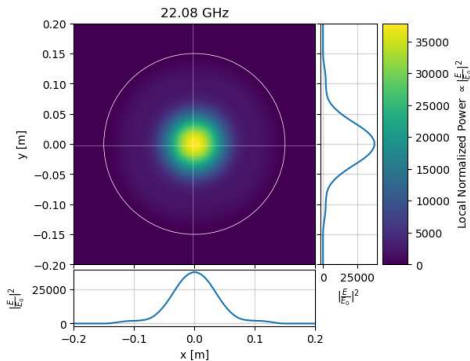
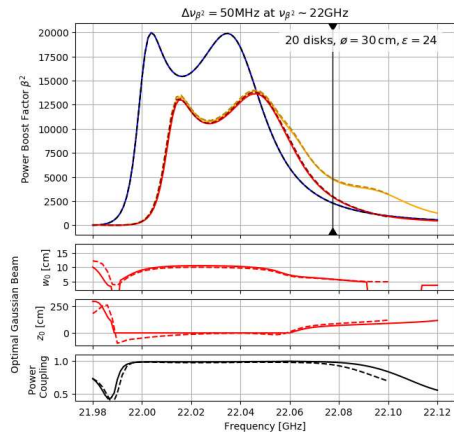
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well matched to Gaussian beam $w_0 \approx 10\text{ cm}$

3D Simulations – Beam Shape

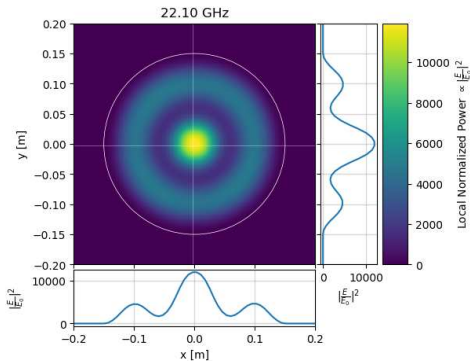
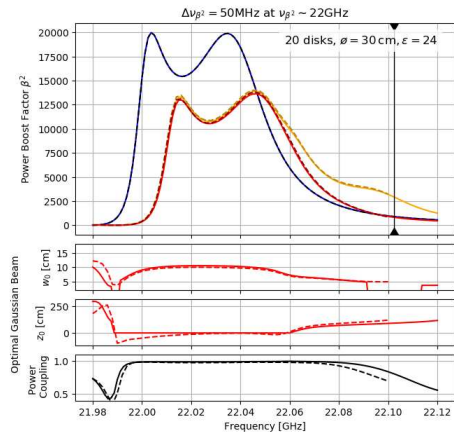
methods see arXiv:1906.02677



well matched to Gaussian beam $w_0 \approx 10\text{ cm}$

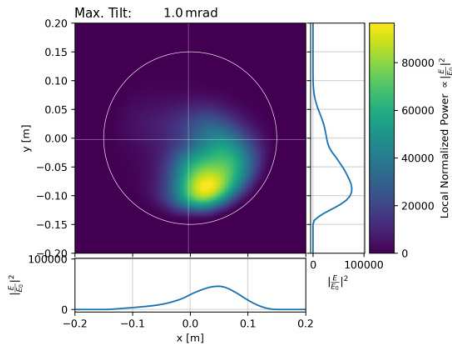
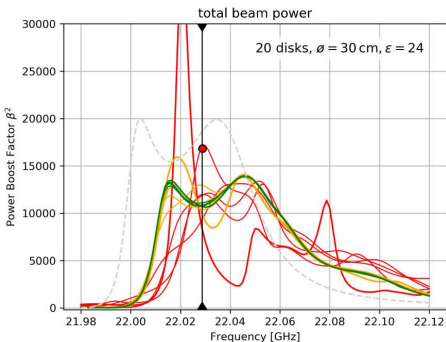
3D Simulations – Beam Shape

methods see arXiv:1906.02677



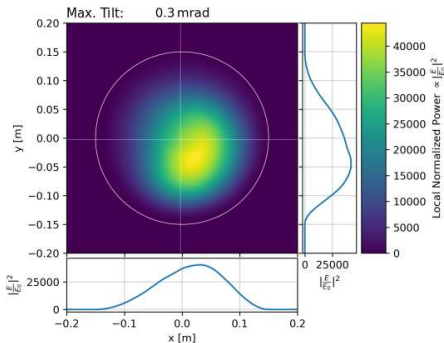
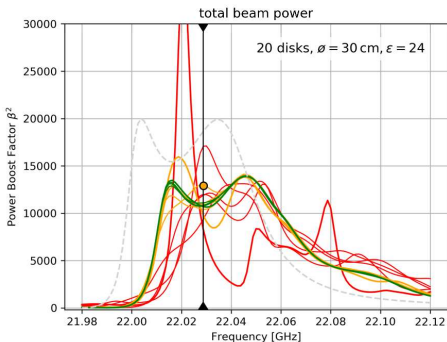
well matched to Gaussian beam $w_0 \approx 10\text{ cm}$

3D Simulations – Tilts



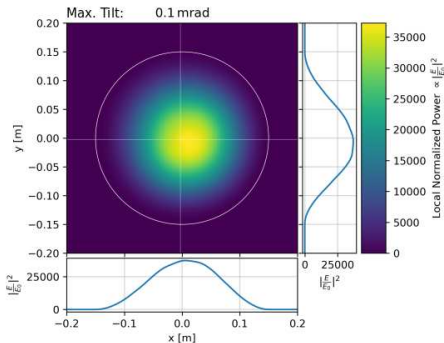
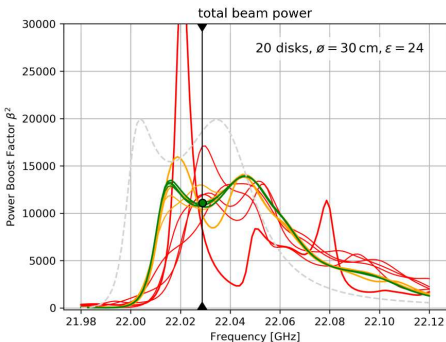
tilts better than 0.1 mrad $\approx 30 \mu\text{m}/30$ cm required

3D Simulations – Tilts



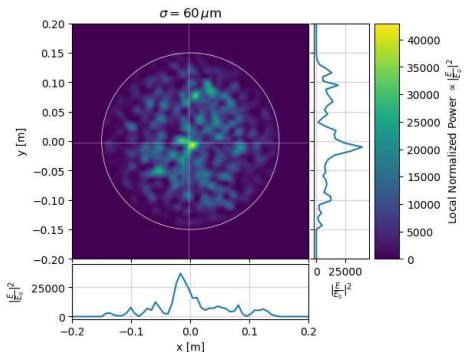
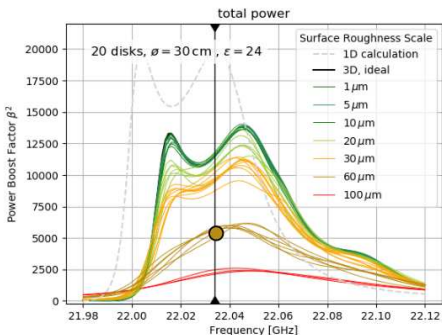
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3D Simulations – Tilts



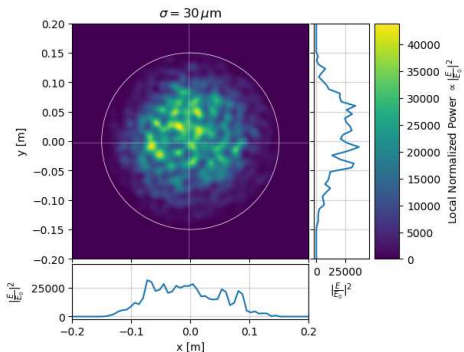
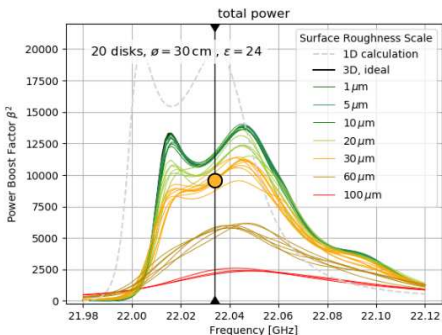
tilts better than 0.1 mrad $\approx 30 \mu\text{m}/30$ cm required

3D Simulation – Surface Roughness



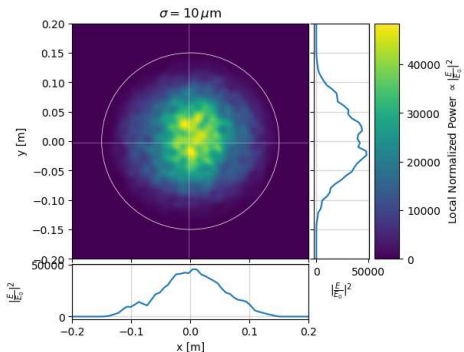
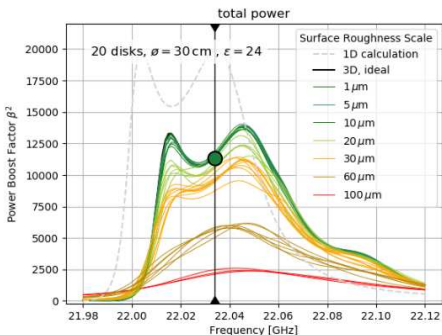
10 μm unproblematic

3D Simulation – Surface Roughness



10 μm unproblematic

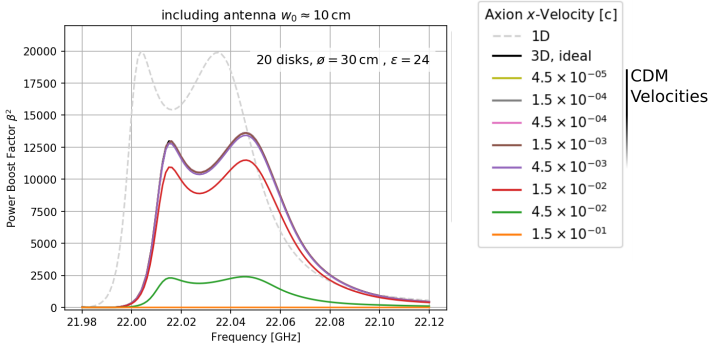
3D Simulation – Surface Roughness



10 μm unproblematic

3D Simulations – Axion Velocity

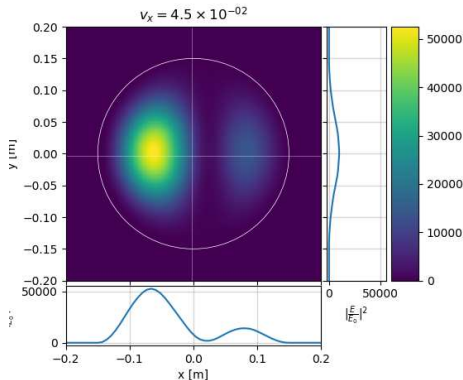
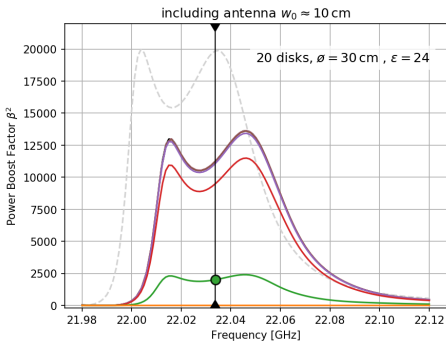
each disk: tilted emission with angle v_x



CDM velocities: **unproblematic**

3D Simulations – Axion Velocity

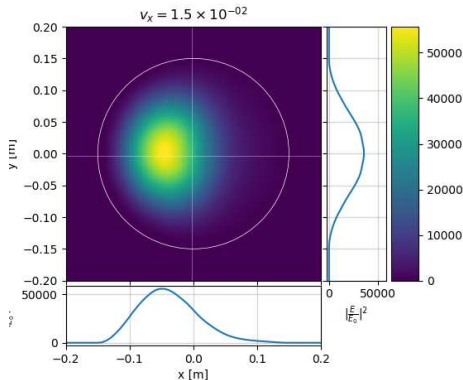
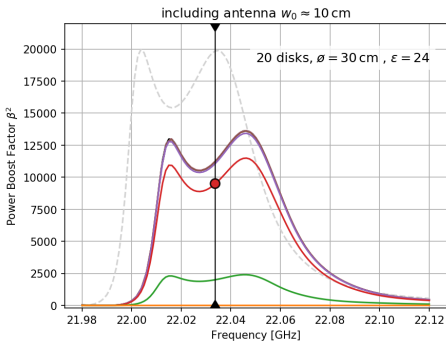
each disk: tilted emission with angle v_x



CDM velocities: **unproblematic**

3D Simulations – Axion Velocity

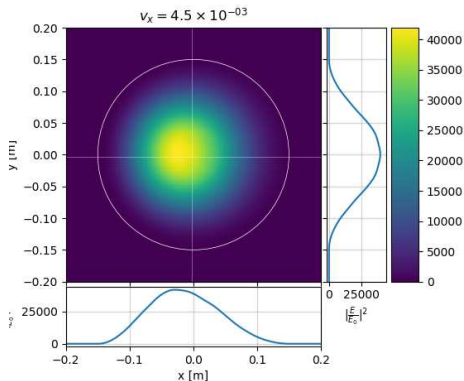
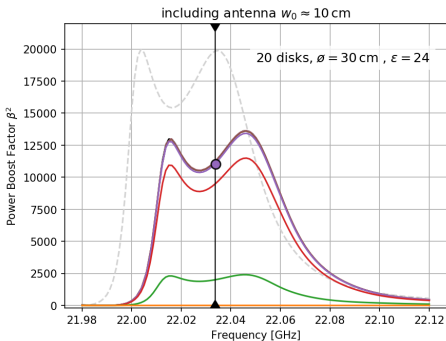
each disk: tilted emission with angle v_x



CDM velocities: **unproblematic**

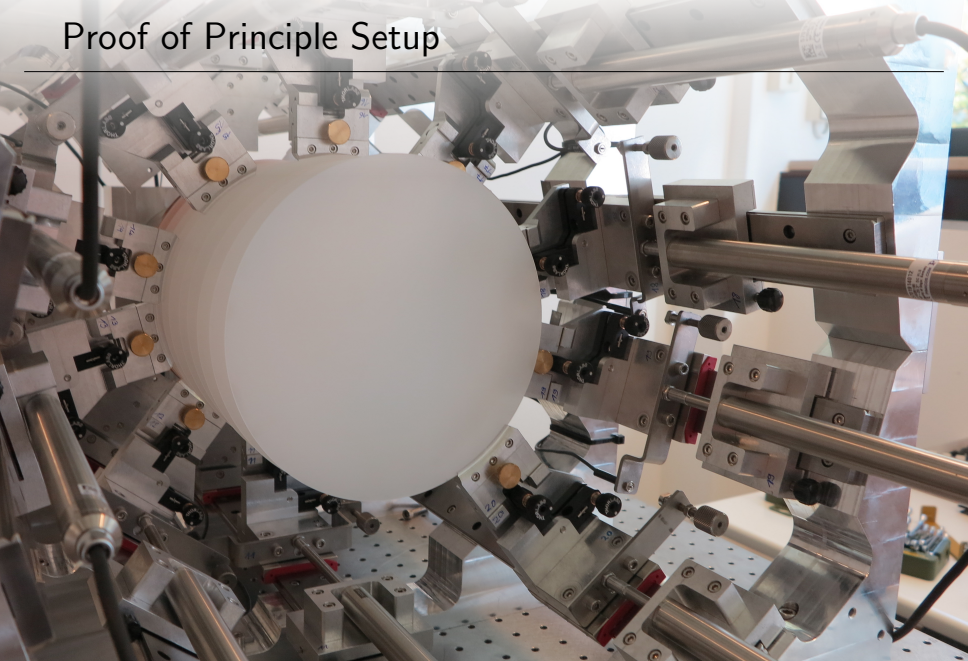
3D Simulations – Axion Velocity

each disk: tilted emission with angle v_x

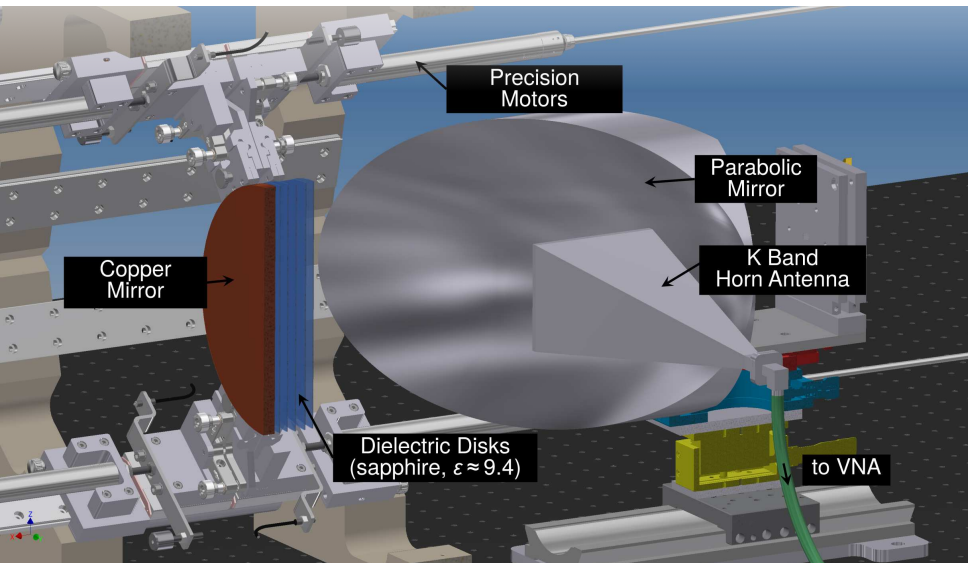


CDM velocities: **unproblematic**

Proof of Principle Setup

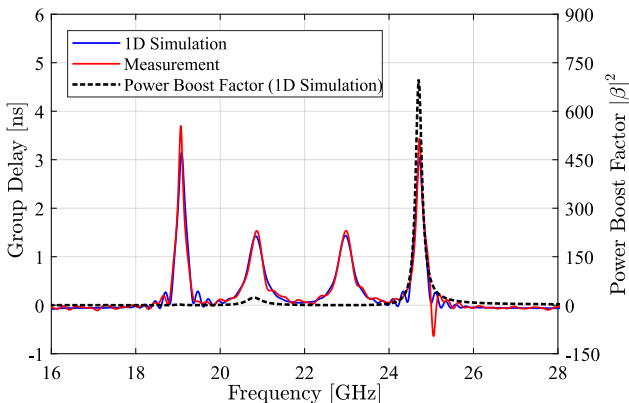


Proof of Principle Setup



Probing the Boost Factor see also Eur. Phys. J. C (2019) 79: 186

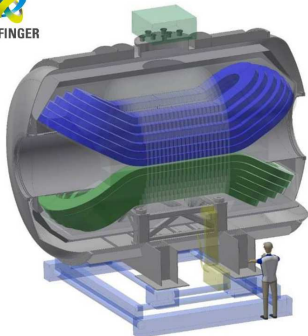
4 equidistant sapphire discs



predicted electromagnetic response demonstrated
ongoing: comparison with 3D simulation, more discs

Further Challenges...

9 T Magnet



similar design also by

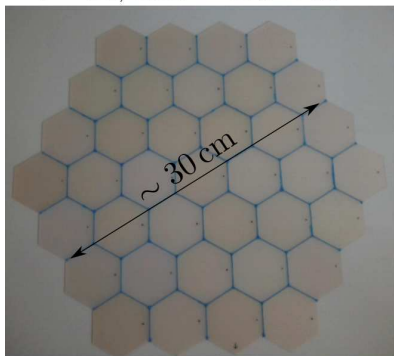


$$B^2 A \sim 100 \text{ T}^2 \text{ m}^2,$$

$$\phi \sim 1.5 \text{ m}, l \sim 1 \text{ m}; \text{ NbTi}$$

Dielectric Discs

$$\epsilon \approx 24, \tan \delta \sim \text{few} \times 10^{-5}$$



U. Hamburg

Motor R&D (e.g. Piezo)

MADMAX collaboration



Universität Hamburg
DER FORSCHUNG | DER LEHRE | DER BILDUNG



Max-Planck-Institut für Physik
(Werner-Heisenberg-Institut)



Max-Planck-Institut
für Radioastronomie



Universidad
de Zaragoza

RWTH AACHEN
UNIVERSITY

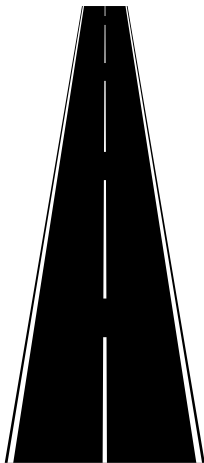
EBERHARD KARLS
UNIVERSITÄT
TÜBINGEN



NEEL
institut



Roadmap



ongoing:

magnet design studies: $B^2 A \approx 100 \text{ T}^2 \text{ m}^2$
(two independent partners)

R&D: on mechanics, LaAlO_3 dielectric plates, noise contribution of booster, receiver

booster studies: 20 disc seed setup, 3D simulations

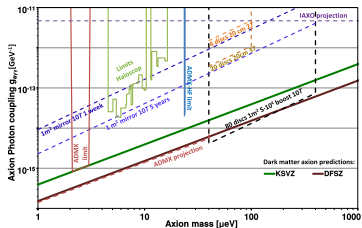
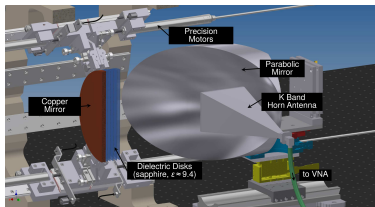
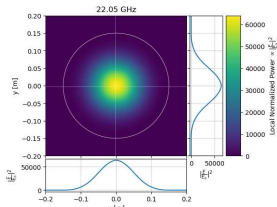
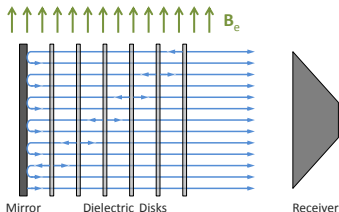
in 1-3 years:

20 disc prototype: $\varnothing_{\text{disc}} \approx 30 \text{ cm}$, $B = 3 - 4 \text{ T}$
 \Rightarrow first physics results

afterwards (2025?):

full scale experiment

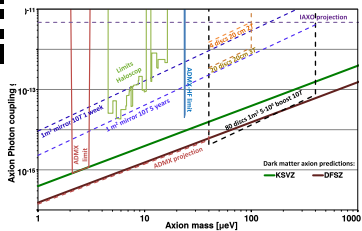
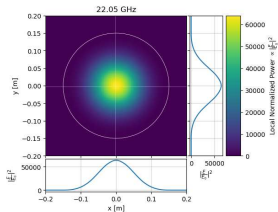
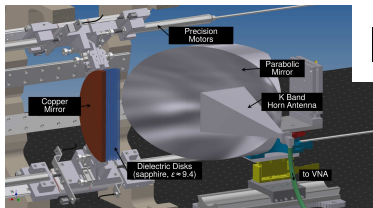
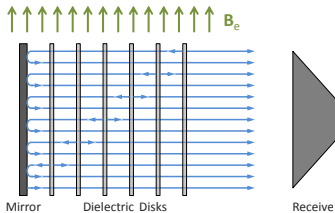
Conclusions



for more information:

[MADMAX collaboration white paper](#),
Phys. Rev. Lett. 118, 091801, **JCAP** 1701:061,2017

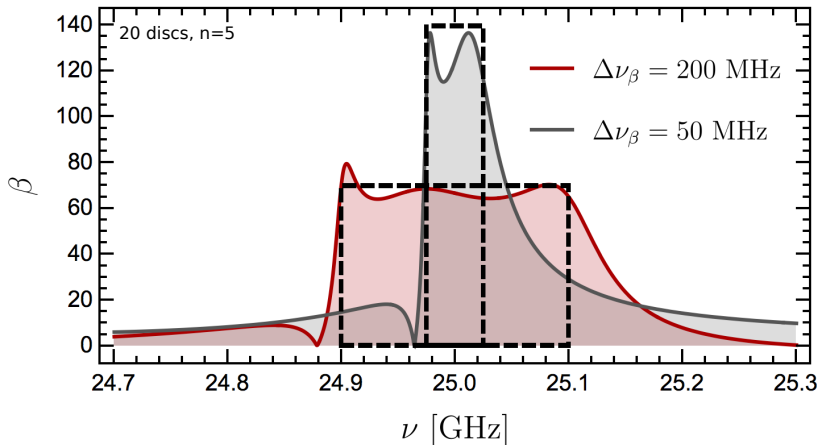
Bedankt voor uw aandacht



for more information:

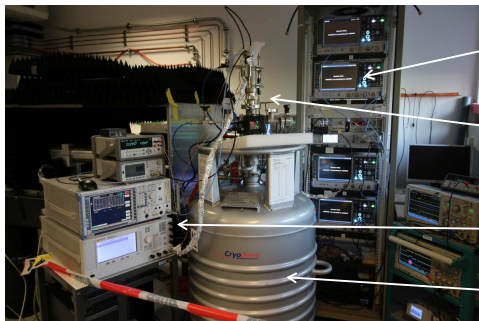
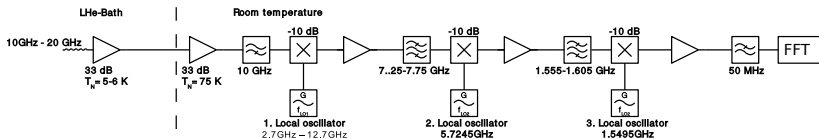
[MADMAX collaboration white paper](#),
Phys. Rev. Lett. 118, 091801, **JCAP** 1701:061,2017

Optimizing the Boost Factor



Area under Boost Factor curve approximately conserved

Receiver System



signal analyzer
(4 samplers, 1.4% dead time)

front end mixers
and amps

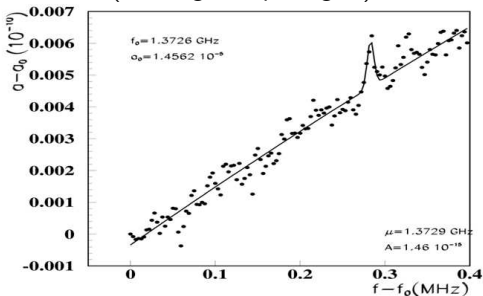
"Fake Axion"

LHe bath
 $\rightarrow 4\text{K } T_{\text{He}} + 5.5\text{K } T_{\text{Amp}} \approx 9.5\text{K } T_{\text{Sys}}$

Receiver System

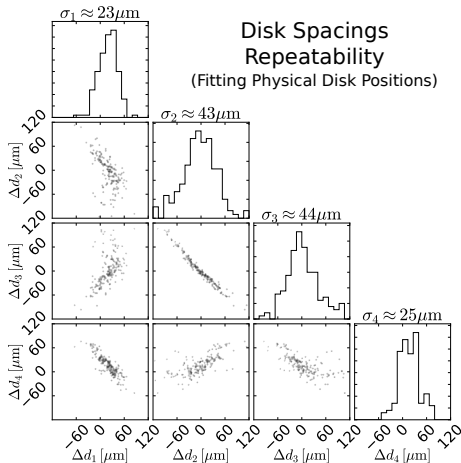
typical one week measurement

(with higher input signal)



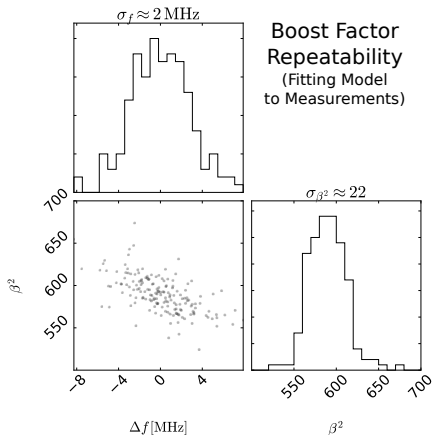
with preamp @ 4K:
signal down to $\sim 10^{-23}$ W detected

Probing the Boost Factor – Results



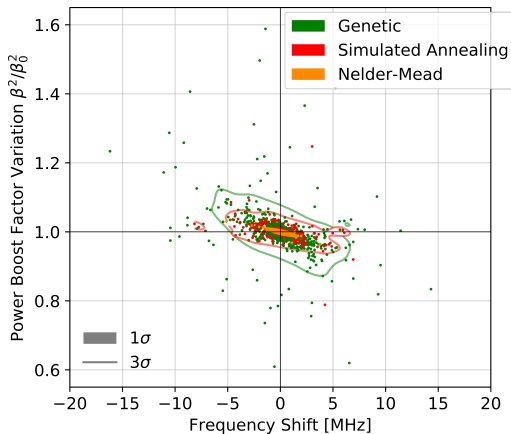
Predicted Electromagnetic Response Demonstrated

Probing the Boost Factor – Results



Predicted Electromagnetic Response Demonstrated

Boost Factor Repeatability



errors on boost factor under control