

### Outline

General **motivations** and **visions** of particle-astroparticle-nuclear physics

#### Concrete examples:

#### **New physics discoveries (dark matter)**

- How to discover new particles
- Complementarity of particle and astroparticle physics experiments
  - Weakly Interacting Massive Particles
  - The case of axion(-like) particles

#### Handling large, heterogeneous amounts of data

- LHC: direct and indirect searches for new physics
- Multimessenger astronomy (nuclear astrophysics)

Examples of ongoing synergistics initiatives

Much inspiration from EPPSU Granada talks, but also some (necessary) personal selection of topics

#### **Conclusions and path forward**



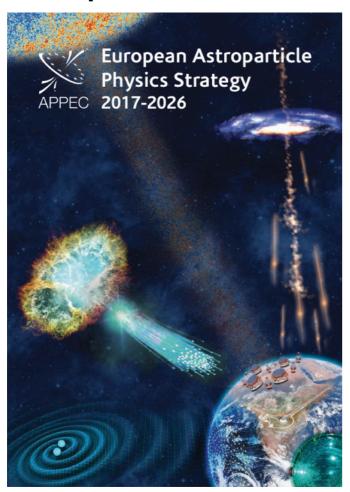


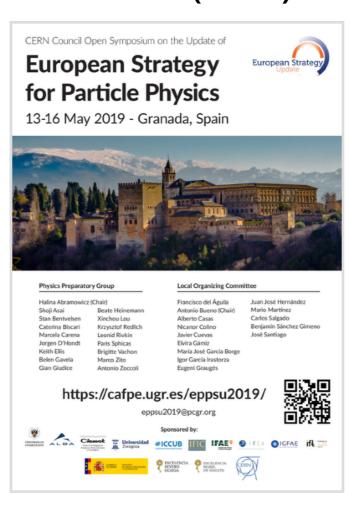


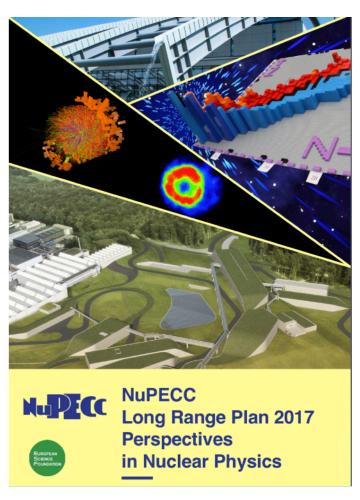
## Visions: APPEC, ECFA, NuPECC

#### Astroparticle (APPEC)

#### Particle (ECFA) Nuclear physics (NuPECC)







Astroparticle, particle and nuclear physics in Europe have **strategies and plans** that **recognize the importance of synergies** between the different fields

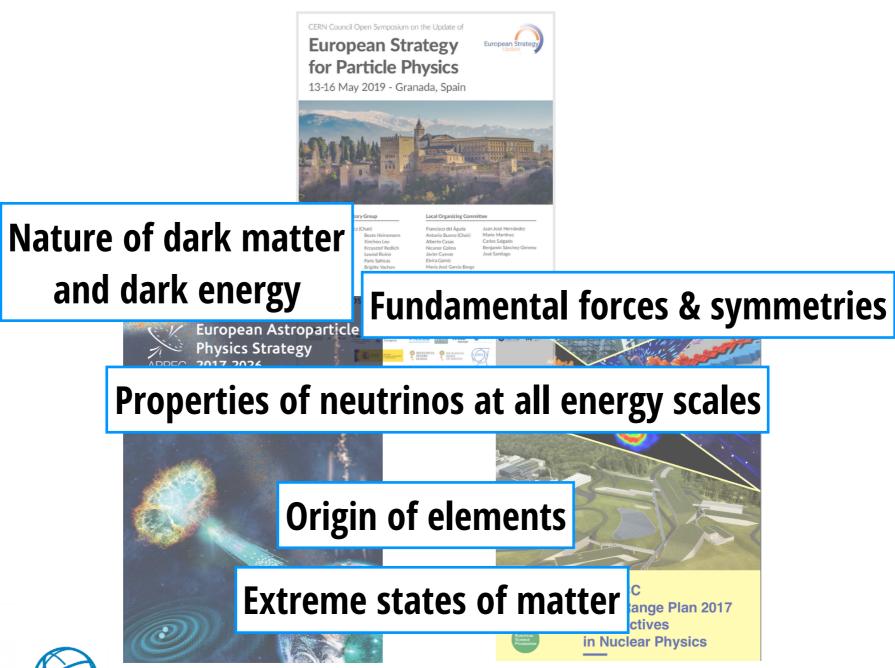






## Visions: APPEC, ECFA, NuPECC

Some of the **common scientific goals** in the strategy documents:

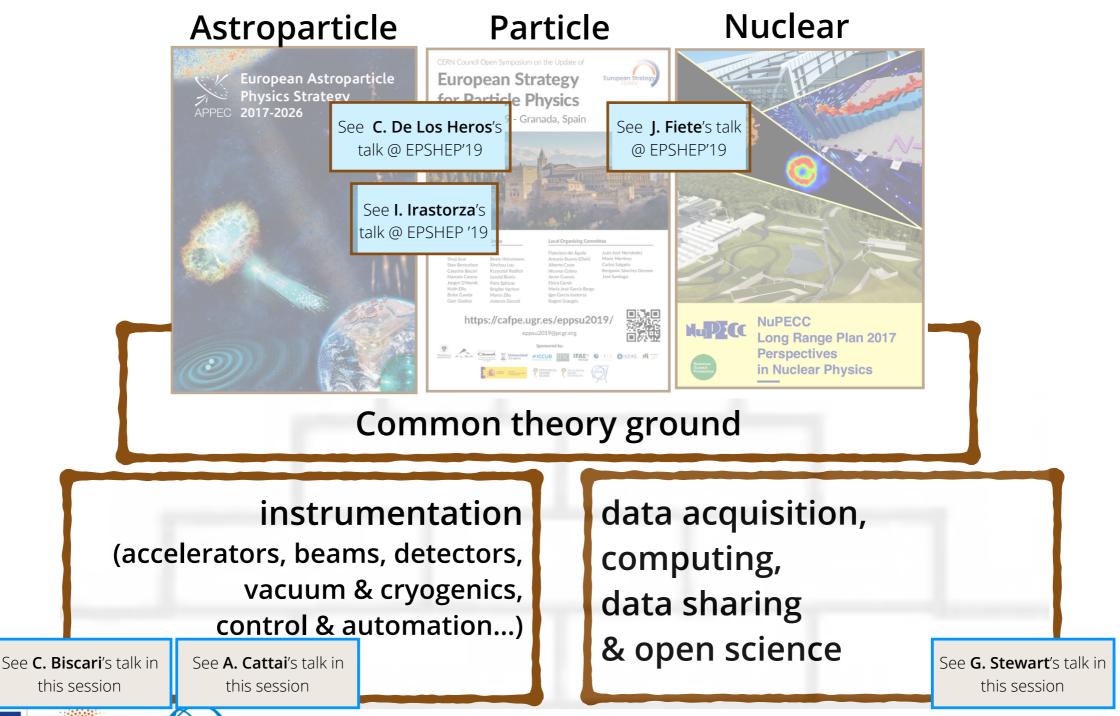








## More synergies: "foundations" for common challenges





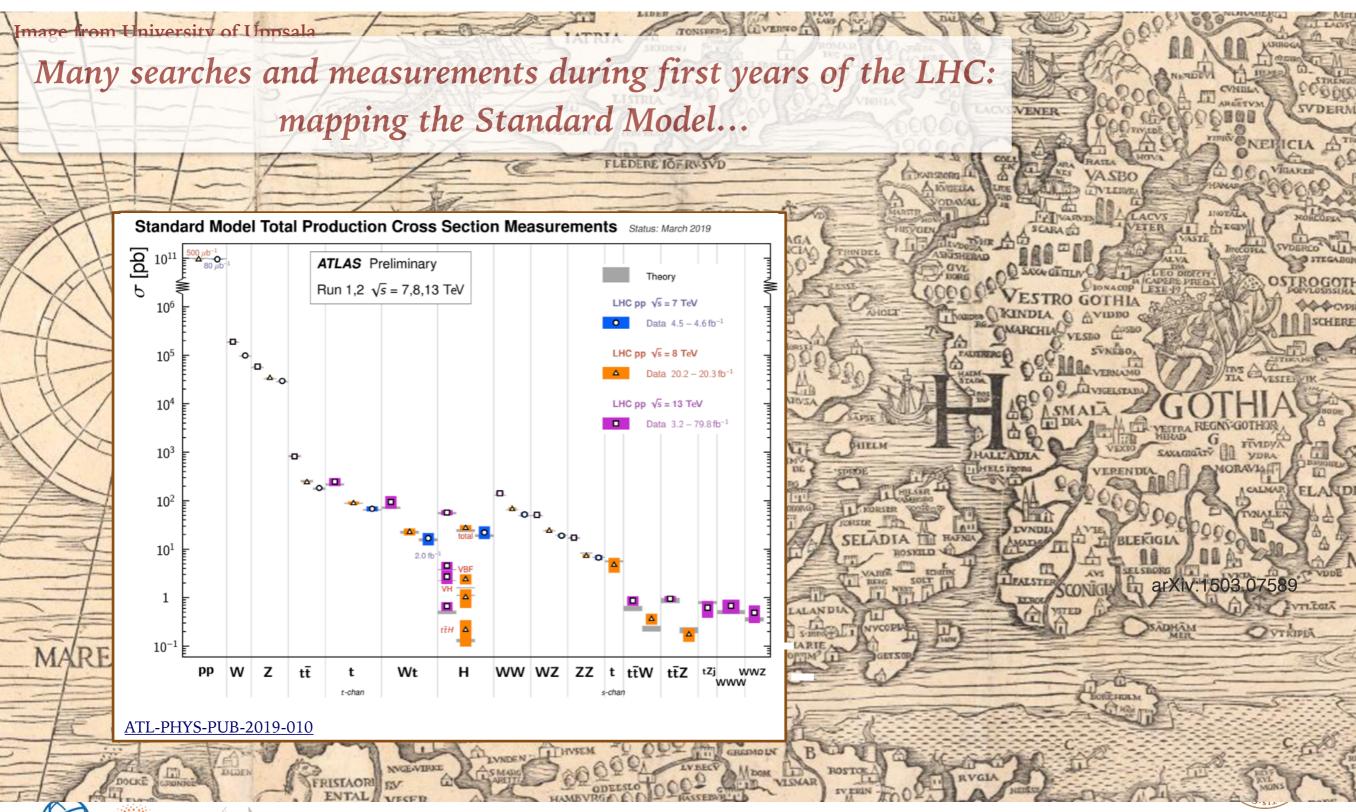


this session



Example of a physics synergy: new physics (of Dark Matter)

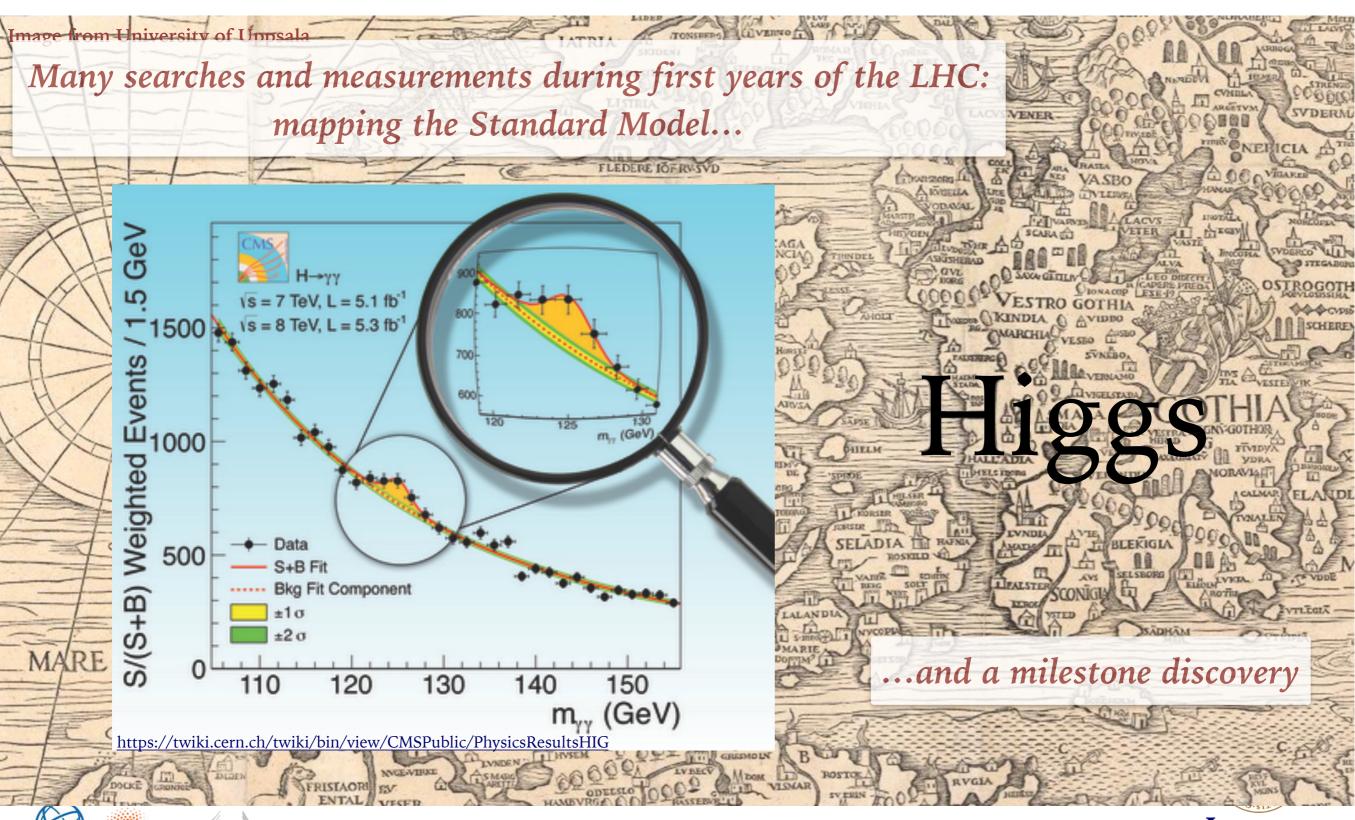
## A chart of measurements (and discoveries)







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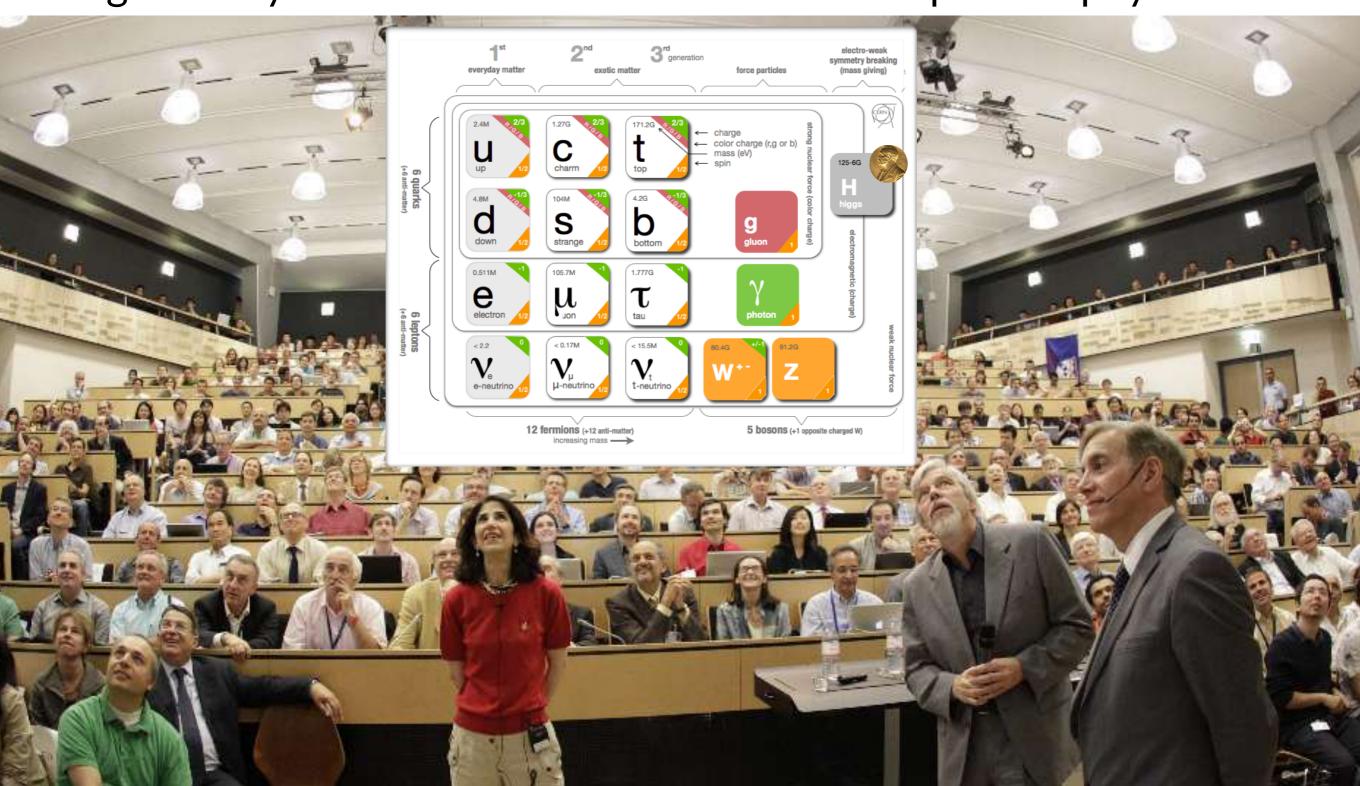




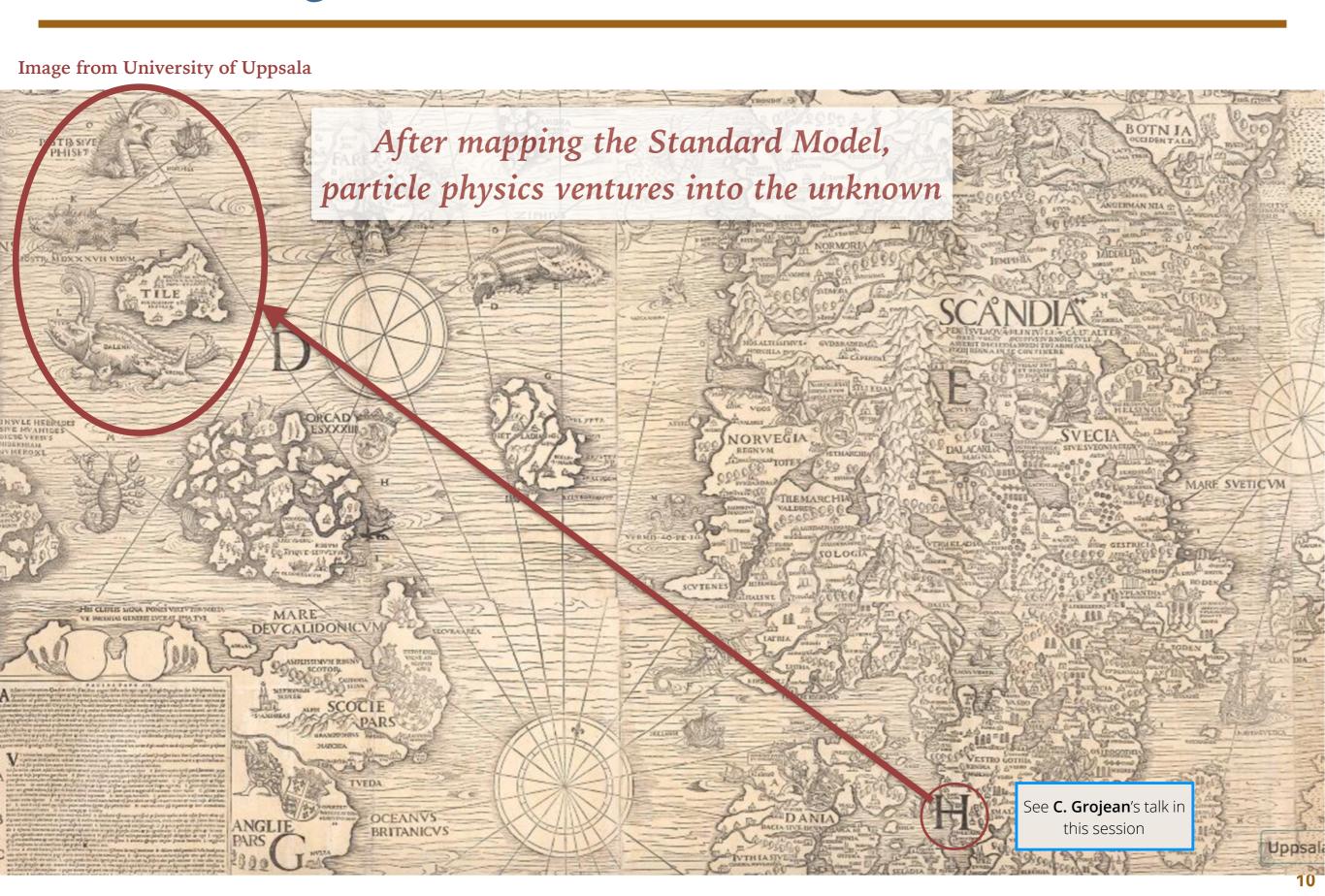
## A chart of searches (and discoveries)

#### Discovery of the Higgs boson:

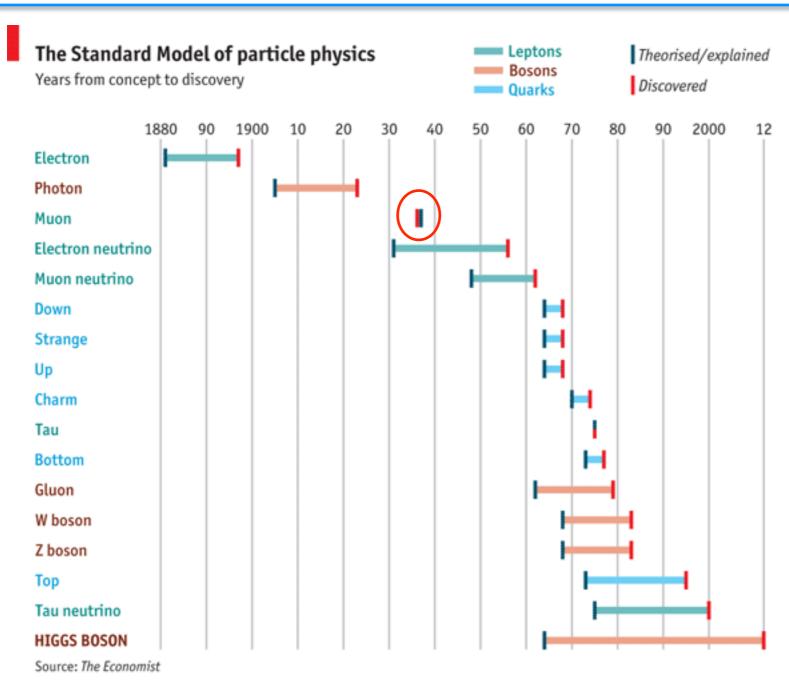
guided by clues from the Standard Model of particle physics



## Where to go next? Uncharted territories



## Expected and unexpected particle discoveries



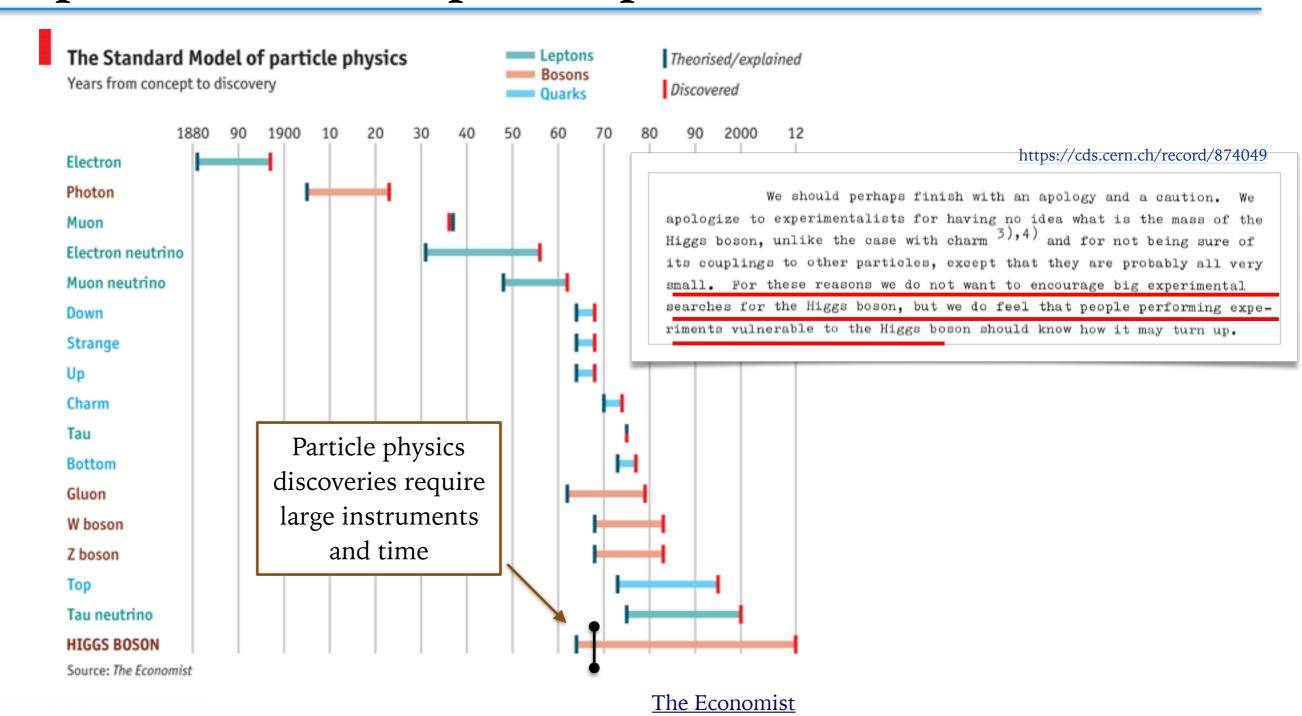
The Economist







## Expected and unexpected particle discoveries

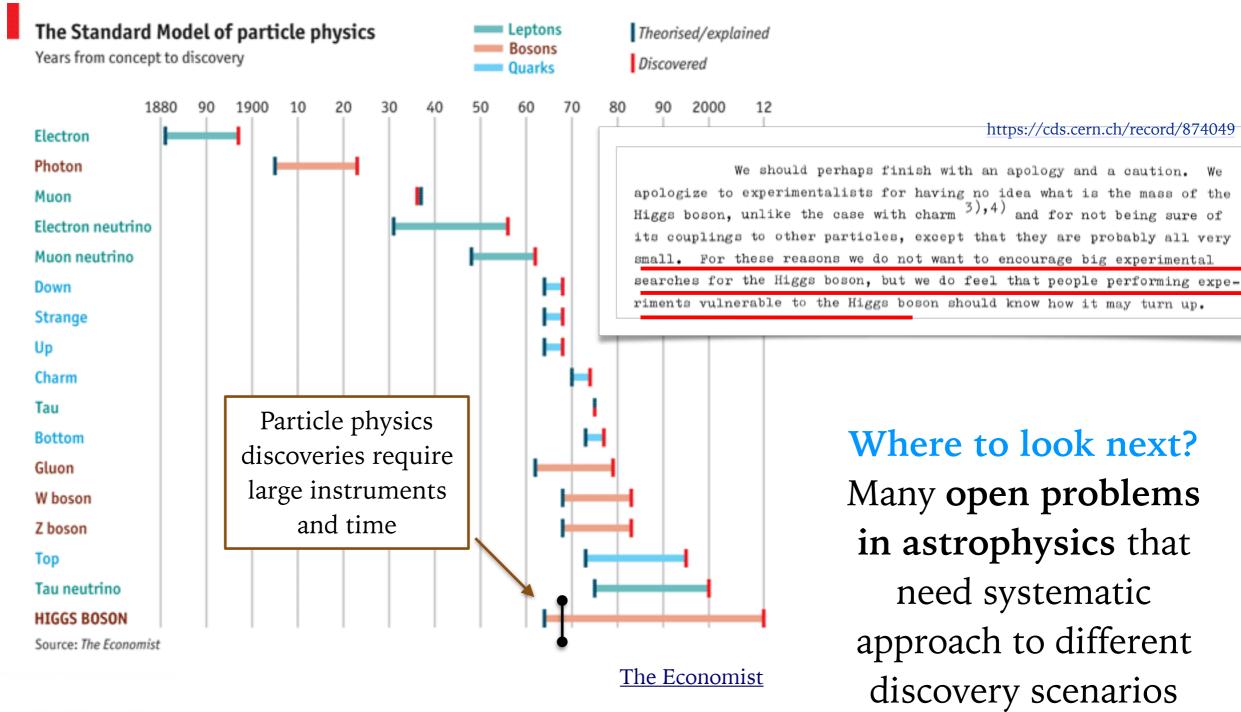








## Expected and unexpected particle discoveries





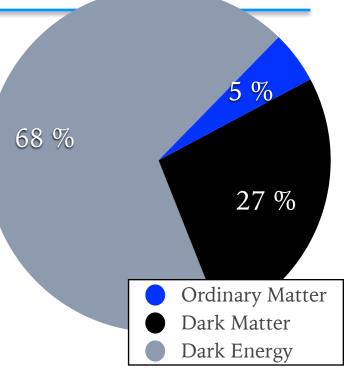




Guidance from astrophysics: dark matter

Empirical **problem** in the Standard Model of Particle Physics: no explanation for **Dark Matter** 

One of the possible **solutions**, guided by **relic density**: invisible **Dark Matter particles** at the **TeV scale** (Weakly Interacting Massive Particles)









68 %

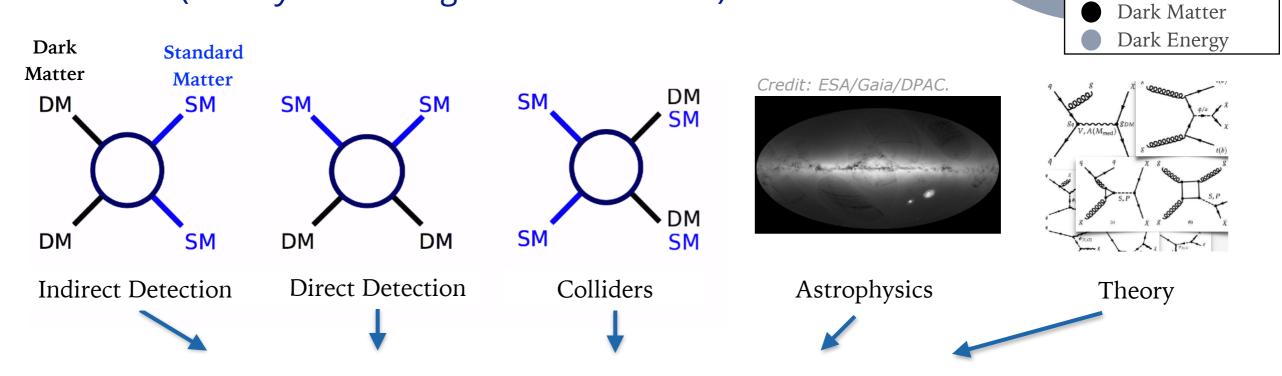
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(Weakly Interacting Massive Particles)



Complementary experimental strategies & inputs







27 %

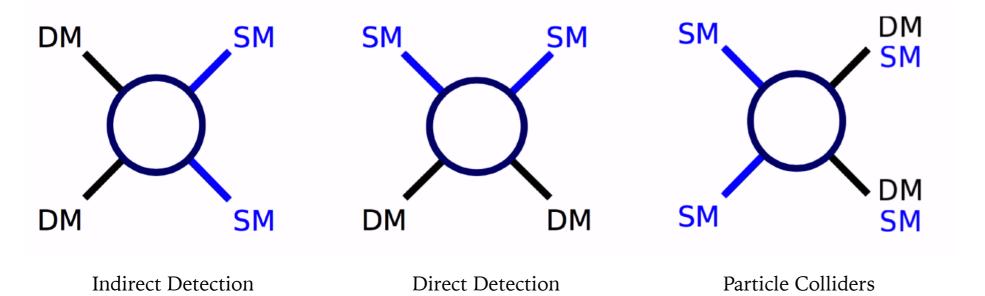
**Ordinary Matter** 

## Collider, direct and indirect detection

Big Question at Granada symposium:

How will Direct and Indirect DM Detection experiments inform/guide accelerator searches and vice-versa?

- Why we need complementarity:
  - DD/ID can discover DM with cosmological origin







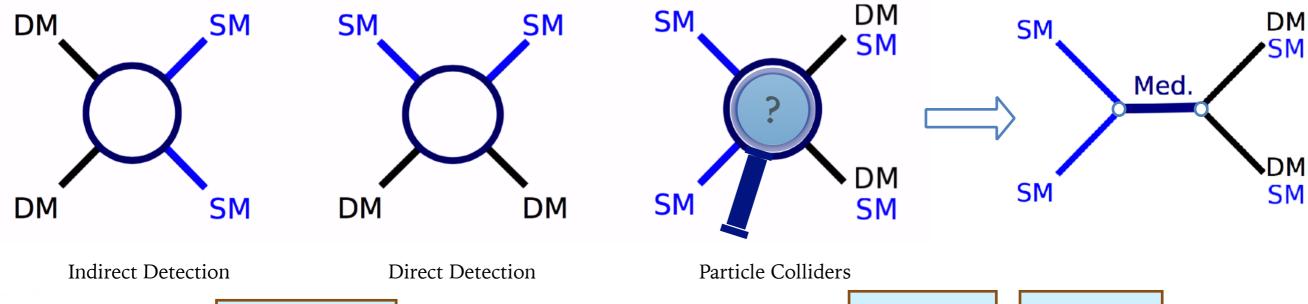


## Collider, direct and indirect detection

Big Question at Granada symposium:

How will Direct and Indirect DM Detection experiments inform/guide accelerator searches and vice-versa?

- Why we need complementarity:
  - DD/ID can discover DM with cosmological origin
  - Colliders can produce DM and probe the dark interaction





See I. Pellmann's

talk @ EPSHEP'19

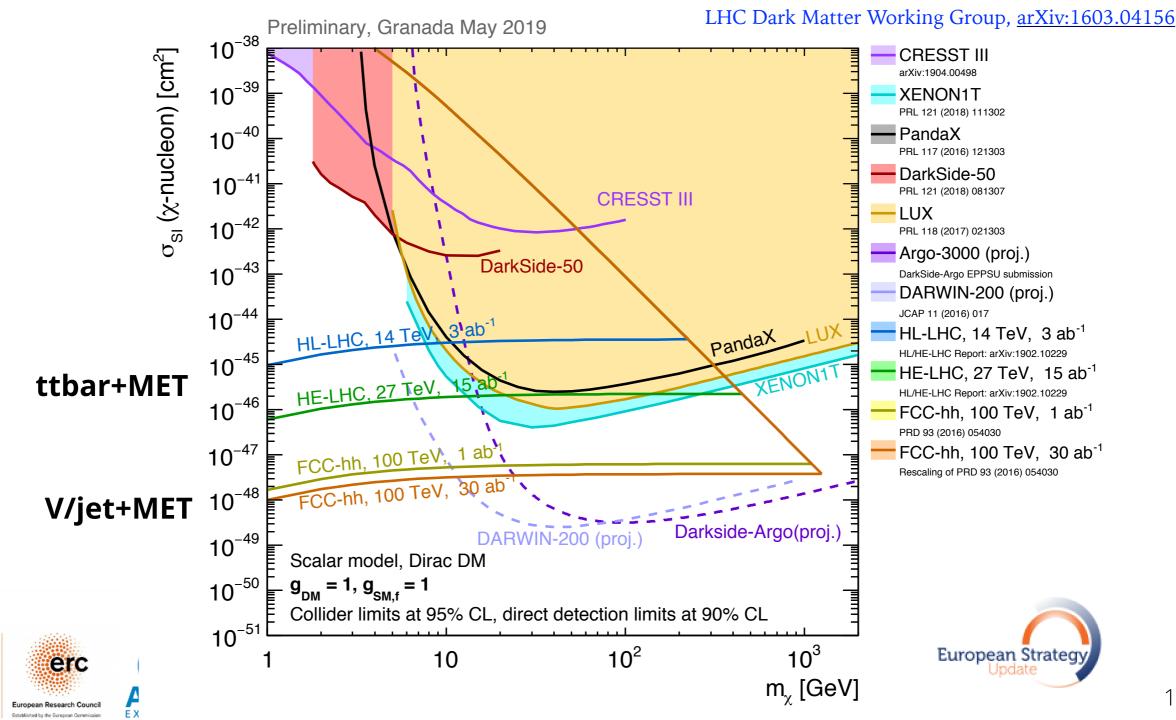
See M. Genest's

talk @ EPSHEP'19

### A simple example: scalar mediator between SM and DM

 Collider constraints on simple models of DM can be shown in terms of DM-nucleon interactions

 $\sigma_{\rm SI} \simeq 6.9 \times 10^{-43}~{\rm cm}^2 \cdot \left(\frac{g_q g_{\rm DM}}{1}\right)^2 \left(\frac{125\,{\rm GeV}}{M_{\rm med}}\right)^4 \left(\frac{\mu_{n\chi}}{1\,{\rm GeV}}\right)^2$ 



### A simple example: scalar mediator between SM and DM

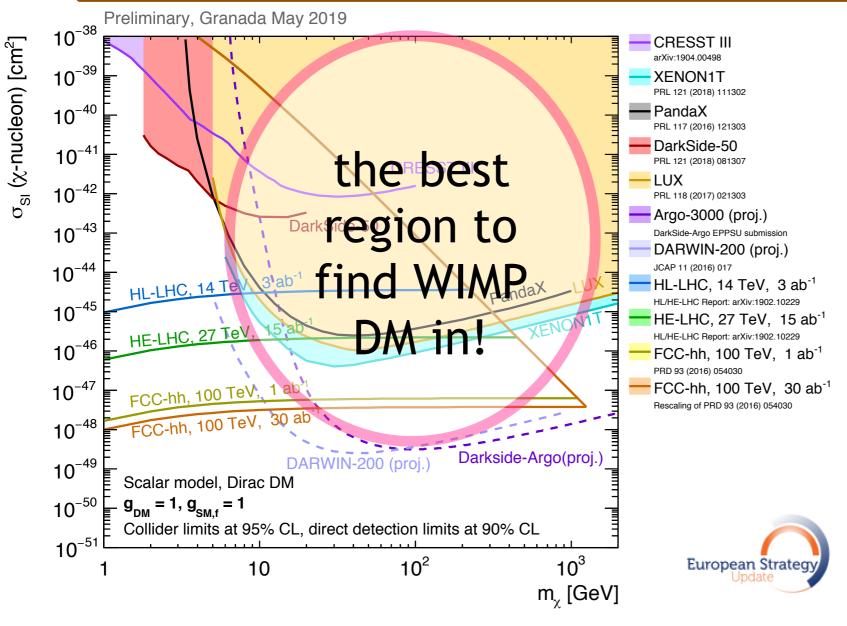
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LHC Dark Matter Working Group, arXiv:1603.04156 Preliminary, Granada May 2019 10<sup>-38</sup>  $\sigma_{\rm SI}$  ( $\chi$ -nucleon) [cm<sup>2</sup>]  $10^{-39}$  $10^{-40}$ PRL 117 (2016) 121303 DarkSide-50  $10^{-41}$ PRL 121 (2018) 081307 **CRESST III**  $10^{-42}$ Argo-3000 (proj.) DarkSide-50  $10^{-43}$ DARWIN-200 (proj.)  $10^{-44}$ LHC, 14 TeV, 3 ab<sup>-1</sup>  $10^{-45}$ JHC, 27 TeV, 15 ab<sup>-1</sup> ttbar+MET  $10^{-46}$ FCC-hh, 100 TeV, 1 ab<sup>-1</sup> FCC-hh, 100 TeV, 1 ab FCC-hh, 100 TeV, 30 ab<sup>-1</sup> of PRD 93 (2016) 054030 FCC-hh, 100 TeV, V/jet+MET **Keep in mind:** these plots are only valid for the couplings specified, Darkside-Argo(proj 10<sup>-49</sup> in the **limited space of a** Scalar model, Dirac DM benchmark model!  $g_{DM} = 1, g_{SM,f} = 1$ Collider limits at 95% CL, direct detection limits at 90% CL  $10^{-51}$ European Strategy 10 Caterina Doglioni - 2019/07/13 - John EPS-ECFA session

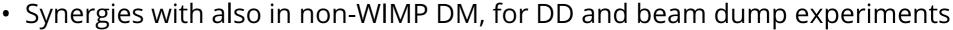
## A simple example: scalar mediator between SM and DM

#### Synergy: complementary reach for future colliders and direct detection



- Collider discovery of invisible particle needs confirmation of cosmological origin from DD/ID
- DD/ID discovery needs collider understanding of nature of interaction
- A future collider program
   that increases sensitivity to
   invisible particles coherently
   with DD/ID serves these
   purposes

See **V. Dutta**'s talk @ EPSHEP'19







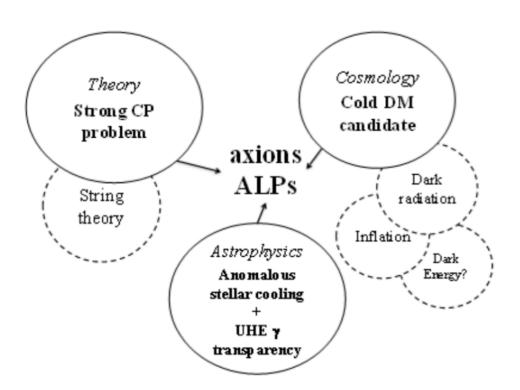


### What if DM isn't a WIMP?

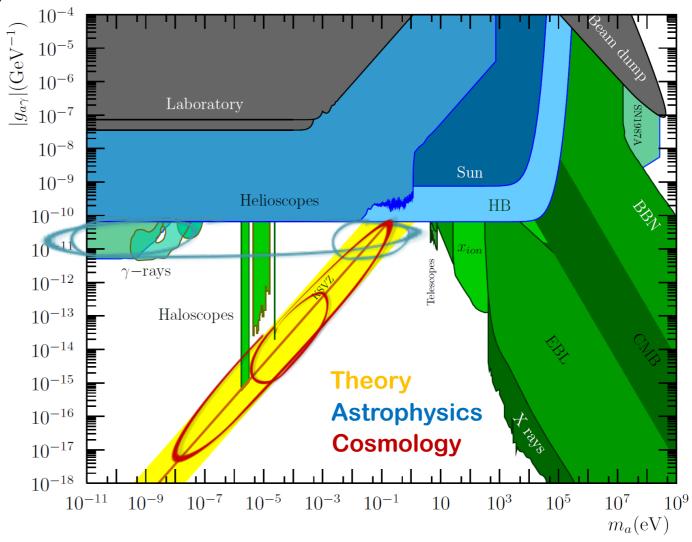
See C. Vallee's talk in this session

**Axions/Axion-Like Particles** (ALPs):

example of new particle with inter-field connections



Figures taken From I. Irastorza's talk @ EPSHEP '19



also using nuclear physics experiments (EDM rings)

Synergies beneficial for many smaller experiments: from joint expertise and common discussion platforms





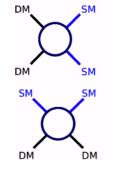


## Synergies in dark matter searches

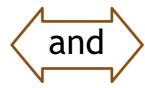
Huge progress planned for direct and indirect detection for WIMP DM

Future colliders and experiments can follow:

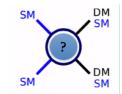
essential complementarity between



cosmological origin astrophysics



nature of the DM-SM interaction particle physics



Similarly, **combination of complementary experiments + theory** needed to identify nature of DM in case of **non-WIMP DM** 

How to strengthen common foundations?

Many common challenges, e.g.

particle detectors and instrumentation,

strategies to handle large amounts of data

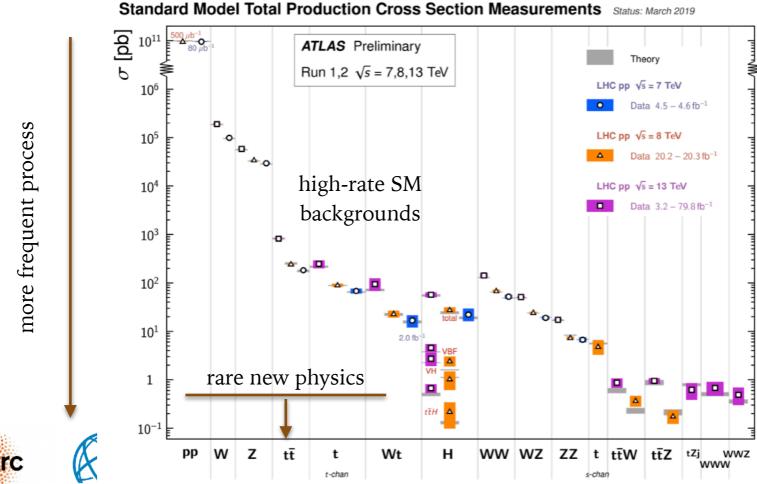




Example of a common challenge: analysis of large, complex datasets

# Enabling discoveries in particle physics

- Many different theories can explain particle physics shortcomings
  - None of these theories is yet favored by data
  - Very different signatures in the detector
    - Some signals buried in high-rate backgrounds

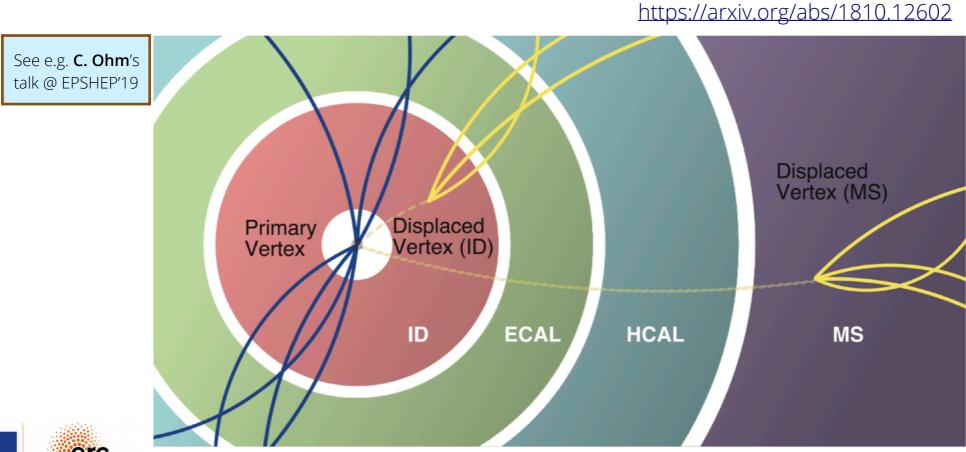






# Enabling discoveries in particle physics

- Many different theories can explain particle physics shortcomings
  - None of these theories is yet favored by data
  - Very different signatures in the detector
    - Some signals buried in high-rate backgrounds
    - Some signals very unusual but rare







# Enabling discoveries in particle physics

- Many different theories can explain particle physics shortcomings
  - None of these theories is yet favored by data
  - Very different signatures in the detector

A key challenge: within millions p-p collisions/second, select/analyze the interesting ones in real time

See **G. Stewart**'s talk in this session



LHC data volumes after selection of "interesting" data





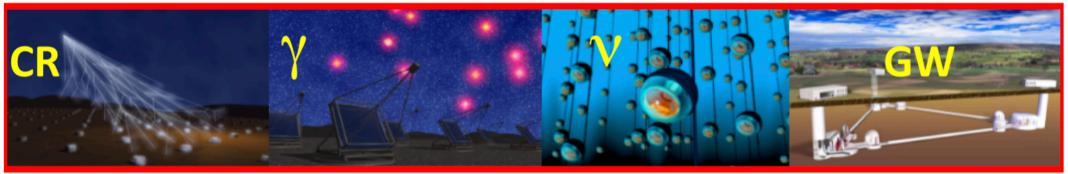


# The advent of Multimessenger astronomy

- Revolutionary combination of information on the cosmos
- Simultaneous detection of astrophysics events
  - "highly heterogeneous, high-volume, high-velocity datasets" arXiv:180

arXiv:1807.04780.pdf







See **E. Bernardini**'s talk @ EPSHEP'19





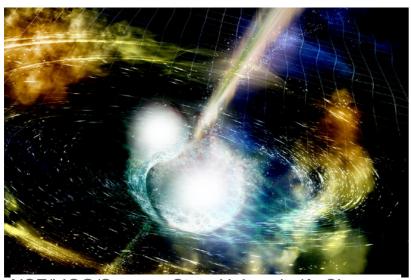


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A key challenge: fast follow-up of interesting events with higher resolution instruments



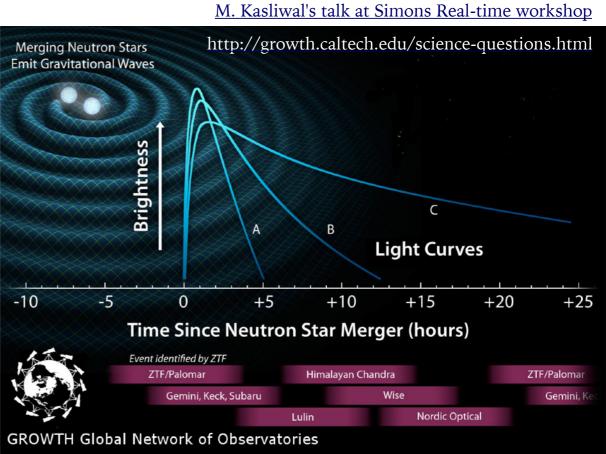
NSF/LIGO/Sonoma State University/A. Simonnet

Light from neutron star mergers can shed light on cosmic origin of heavy elements







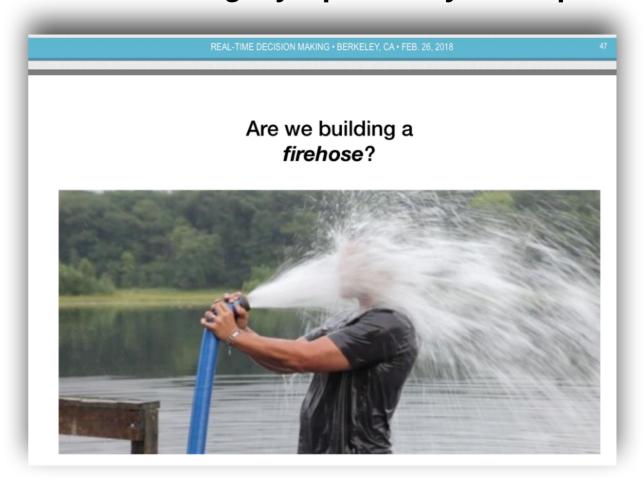


## Extremely large datasets, in different contexts

C. Fitzpatrick, **LHCb** 



E. Bellm, Large Synoptic Survey Telescope



The LHC and modern astrophysics surveys are data firehoses



Can benefit from common techniques and tools for data taking & data reduction

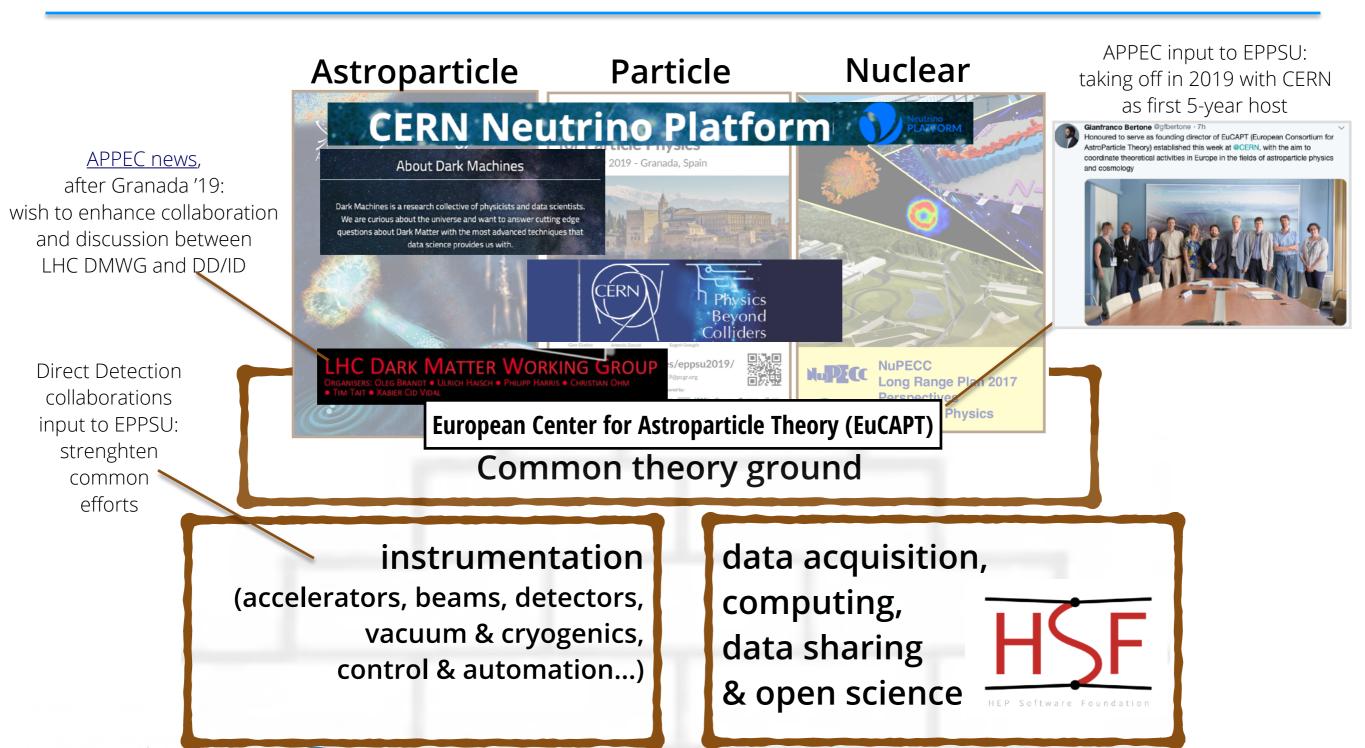
(e.g. on-detector / real-time data analysis, machine learning) with applications beyond physics research





Synergy initiatives and outlook

### A constellation of activities and initiatives









### Conclusions and outlook

- Answering fundamental physics questions requires **concerted work** from **particle, astroparticle and nuclear physics** 
  - Examples: dark matter (in this talk), neutrino physics...
  - Common challenges in terms of foundations (detector, computing...)
- A number of synergistic initiatives exist, many hosted by CERN
  - What is the best way forward? Discussion started at Granada meeting
- More discussion at the APPEC-ECFA-NuPECC meeting in Orsay this October







