

Electroweak Corrections

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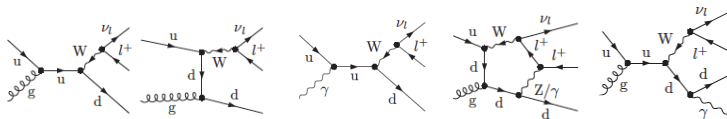
Buffalo, NY, July 18, 2017

Electroweak corrections at hadron colliders: a prelude

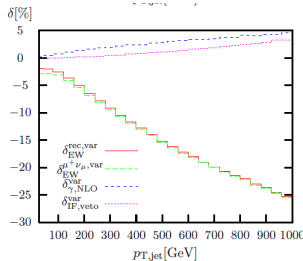
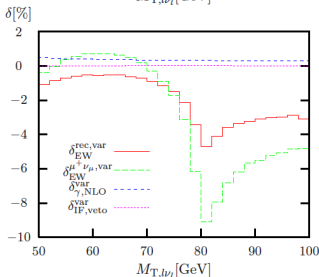
Fixed order prediction for $pp \rightarrow W(l\nu) + 1\text{jet} + X$ up to $\mathcal{O}(\alpha_s\alpha^3)$: A.Denner, S.Dittmaier,

T.Kasprzik, A. Mück, arXiv:0906.1656

$$d\sigma_{NLO} = d\sigma_{LO}(\alpha_s\alpha^2) + d\sigma_{LO,\gamma}(\alpha^3) + (d\sigma_{virtual} + d\sigma_{real})_{EW}(\alpha_s\alpha^3) + d\sigma_{NLOQCD,\gamma}(\alpha_s\alpha^3)$$



Impact on observables usually shown as relative correction: $\delta(\%) = \frac{d\sigma_{NLO}}{d\sigma_{LO}}$



- Recent Les Houches workshop reports
 - **Dictionary of electroweak (EW) corrections:** S.Dittmaier
EW input schemes, EW Sudakov logs, mass-singular logs, QED corrections in PDFs, photon-induced processes, treatment of unstable W, Z bosons, photon-jet separation, combination of QCD and EW corrections  arXiv:1405.1067
 - **High-precision wishlist and status of automation:**  arXiv:1605.04692
- **Precision studies of observables in $pp \rightarrow W \rightarrow l_l$ and $pp \rightarrow \gamma, Z \rightarrow l^+l$ processes at the LHC:**  arXiv:1606.02330
and
Precision Measurement of the W -Boson Mass: Theoretical Contributions and Uncertainties C. M. Carloni Calame, M. Chiesa, H. Martinez, G. Montagna, O. Nicrosini, F. Piccinini, A. Vicini  arXiv:1612.02841
- **Precise predictions for V +jets dark matter backgrounds** J. M. Lindert, S. Pozzorini, R. Boughezal, J. M. Campbell, A. Denner, S. Dittmaier, A. Gehrmann-De Ridder, T. Gehrmann, N. Glover, A. Huss, S. Kallweit, P. Maierhöfer, M. L. Mangano, T.A. Morgan, A. Mück, F. Petriello, G.P. Salam, M. Schönherr, C. Williams  arXiv:1705.04664

EW corrections are needed for

- precisely extracting SM input parameters from observables: M_W , $\sin^2 \theta_{eff}$, m_t , $y_{b,t}$, \dots , Example: M_W extracted from transverse lepton-pair mass or lepton momentum in single W production via Drell-Yan:

$$M_T(l\nu_l) = \sqrt{p_T^l p_T^\nu (1 - \cos(\Phi_l - \Phi_\nu))}$$

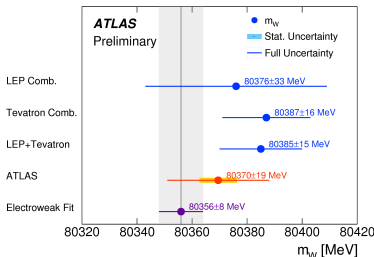
- precisely predicting SM input parameters, e.g., $M_W(\text{theory})$ from

$$\frac{G_\mu}{\sqrt{2}} = \frac{\pi\alpha(0)M_Z^2}{2(M_Z^2 - M_W^2)M_W^2} [1 + \Delta r]$$

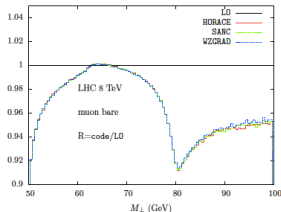
$\Delta r(M_W, m_t, M_H, \dots)$ describes the loop corrections to muon decay. For a review of the role of radiative corrections in EW precision physics see, e. g., [A.FerrogliA, A.Sirlin \(2013\)](#).

- reducing systematic uncertainties, e.g., W, Z observables are important for constraining quark PDFs.

First W mass measurement at the LHC (ATLAS) arXiv:1701.07240

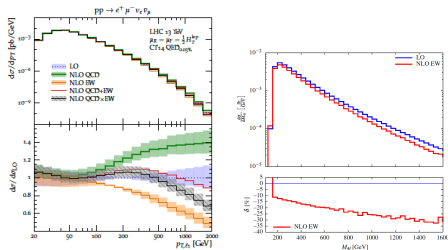


Impact of EW corrections on $M_T(\mu\nu_\mu)$ in $pp \rightarrow W^+ \rightarrow \mu^+ \nu_\mu$: arXiv:1606.02330

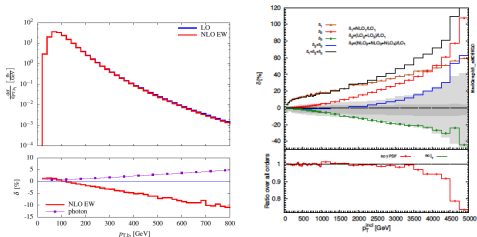


- EW corrections are relevant for modeling signal and background processes for searches of signals of new physics, either due to direct production, higher-dimensional operators, or the virtual presence of new particles in SM observables.
- EW corrections play an especially important role in EW gauge boson production processes: V, VV, VVV (+jets) with $V = \gamma, Z, W^\pm$.
- EW corrections can be numerically at least as important as NNLO QCD corrections and for certain processes and in certain kinematic regions they may even be the dominant corrections.

$pp \rightarrow e^+ \mu^- \nu_e \bar{\nu}_\mu$: S.Kallweit, J.Lindert, S. Pozzorini, M.Schönherr, arXiv:1705.00598; $pp \rightarrow \mu^+ \nu_\mu e^+ \nu_e jj$: B. Biedermann, A.Denner, M.Pellen, arXiv:1611.02951



$pp \rightarrow e^+ \nu_e \mu^- \nu_\mu b \bar{b}$: A.Denner, M.Pellen, arXiv:1607.05571;
 $pp \rightarrow jj$: R.Frederix *et al*, arXiv:1612.06548



Characteristics of EW corrections

Naive estimate of relative size of EW and QCD corrections:

$$\frac{\alpha(M_Z)}{\pi} \approx 0.0025 \text{ vs. } \frac{\alpha_s(M_Z)}{\pi} \approx 0.037 \text{ and } \left(\frac{\alpha_s(M_Z)}{\pi}\right)^2 \approx 0.0014$$

Possible enhancements:

$$\text{QED corrections: } \frac{\alpha(0)}{\pi} \log\left(\frac{m_f^2}{Q^2}\right) \approx -0.024 \text{ for } Q = M_W, f = \mu$$

Origin: Soft/collinear FS photon radiation

In sufficiently inclusive observables these mass singularities completely cancel. [Kinoshita, Lee, Nauenberg \(1962,1964\)](#)

Depending on the experimental lepton identification cuts they can significantly affect the shape of distributions.


IS mass singularities are factorized into PDFs which introduces a QED factorization scheme; PDFs with QED corrections and photon PDF [A.Manohar, 1607.04266 \(LUXqed\)](#)

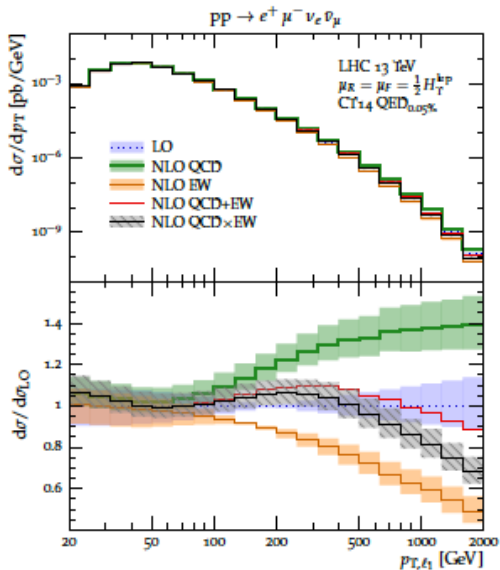
$$\text{Weak Sudakov corrections, e.g., at LL: } -\frac{\alpha}{\pi s_w^2} \log^2\left(\frac{M_V^2}{Q^2}\right) \approx -0.052 \text{ for } Q=2 \text{ TeV}$$

Origin: Remnants of UV singularities after renormalization and soft/collinear IS and FS emission of virtual and real W and Z bosons.

In contrast to QED and QCD, also in inclusive observables these corrections do not completely cancel. [M.Ciafaloni, P.Ciafaloni, D.Comelli \(2000,2001\)](#) see, e.g., [K.Mishra et al, 1308.1430](#); [J.H.Kühn, Acta Phys.Polon.B39 \(2008\)](#) for examples and a brief review

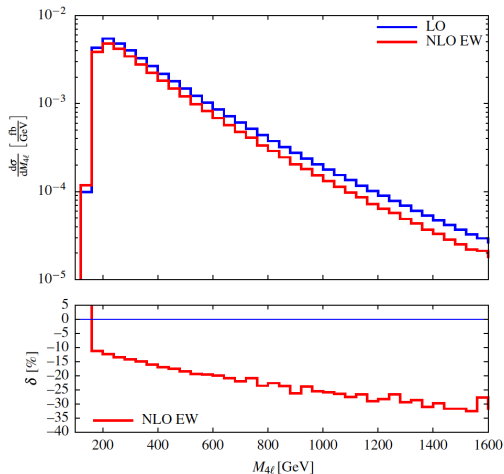
$pp \rightarrow e^+ \mu^- \nu_e \bar{\nu}_\mu$ at NLO(EW+QCD)

S.Kallweit, J.Lindert, S. Pozzorini, M.Schönherr;  arXiv:1705. 00598



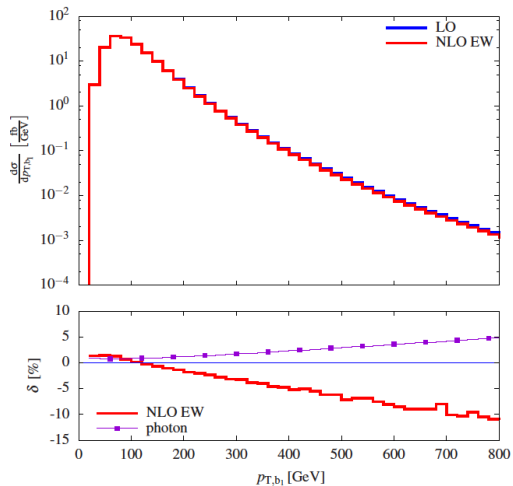
$pp \rightarrow \mu^+ \nu_\mu e^+ \nu_e jj$ at NLO EW (VBS)

B. Biedermann, A. Denner, M. Pellen  arXiv:1611.02951



$pp \rightarrow e^+ \nu_e \mu^- \nu_\mu b \bar{b}$ at NLO EW

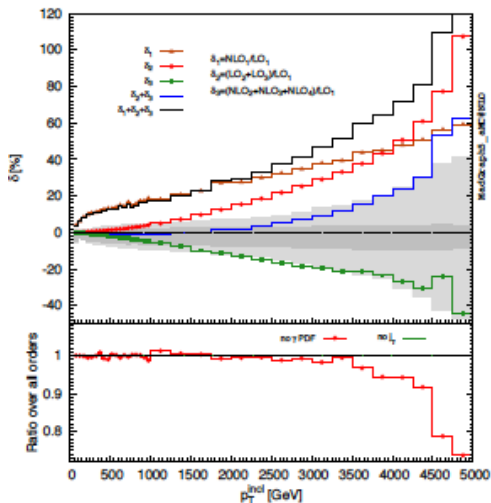
A.Denner, M.Pellen  arXiv:1607.05571;



R.Frederix et al  arXiv:1612.06548

LO: $\mathcal{O}(\alpha_s^2)$, $\mathcal{O}(\alpha^2)$, $\mathcal{O}(\alpha\alpha_s)$

NLO: $\mathcal{O}(\alpha_s^3)$, $\mathcal{O}(\alpha_s^2\alpha)$, $\mathcal{O}(\alpha_s\alpha^2)$, $\mathcal{O}(\alpha^3)$



Mass-singular logarithms of QED origin

Multiple FS photon radiation and exponentiation at LL, $L = \log\left(\frac{Q^2}{m_f^2}\right)$:

- Exponentiation of YFS form factor [Yennie, Frautschi, Suura \(1961\)](#):

$$Y(m \ll Q) = \frac{\alpha}{\pi} \left\{ 2(L-1) \ln\left(\frac{2\Delta E_\gamma}{Q}\right) + \frac{1}{2}L - \frac{1}{2} - \frac{\pi^2}{6} \right\}$$

Implemented in WINHAC for W production [Placzek et al \(2003\)](#), matched to NLO EW of SANC [Bardin et al \(2008\)](#); and in Sherpa [M. Schönherr, F. Krauss \(2008\)](#).

- QED parton shower: emission of n photons ($I_+ = \int_0^{1-\epsilon} dz P(z)$)

$$d\sigma = \exp\left[-\frac{\alpha}{2\pi} I_+ L\right] \sum_n^\infty |M_n^{LL}|^2 d\Phi_n$$

Implemented in HORACE [Carloni-Calame et al \(2003,2004,2006\)](#), matched to full NLO EW.

- QED structure function [Kuraev, Fadin \(1985\)](#):

$$d\sigma = d\sigma_{LO} \int dz \Gamma(z) \theta_{cut}(z p_I); \beta_I = \frac{2\alpha(0)}{\pi} (L-1)$$

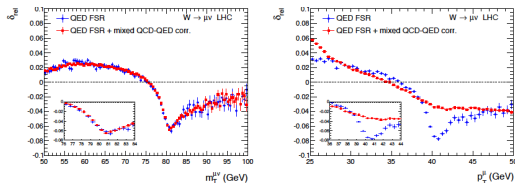
$$\Gamma(z, Q^2) = \frac{\exp[-\beta_I/2\gamma_E + \frac{3}{8}\beta_I]}{\Gamma(1 + \beta_I/2)} \frac{\beta_I}{2} (1-z)^{\beta_I/2-1} + \dots + \mathcal{O}(\beta_I^4)$$

Implemented in W production [Breusing, Dittmaier, Krämer, Mück \(2008\)](#) and Z production [Dittmaier, Huber \(2009\)](#), matched to full NLO EW.

- POWHEG(NLO QCD+EW) \otimes (QCD+QED) PS; QED PS with PHOTOS ([Golonka, Was \(2005,2006\)](#)) or with PYTHIA 8 for W production [Carloni Calame et al, 1612.02841](#).

The implementation of EW corrections in the POWHEG BOX Barze *et al.*, 1202.0465; Carloni Calame *et al.*, 1612.02841

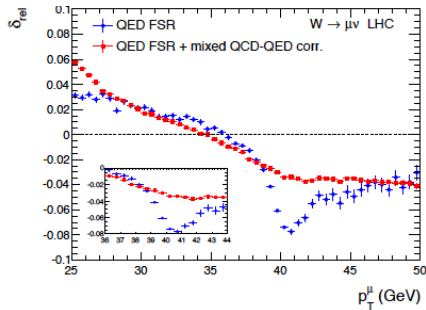
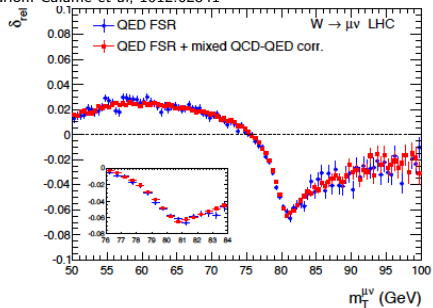
- ensures normalization with NLO QCD + EW accuracy,
- combines the complete SM NLO corrections with a mixed QCD \otimes QED parton cascade, where the particles present in the shower are coloured particles or photons, and
- consequently, incorporates mixed $\mathcal{O}(\alpha\alpha_s)$ contributions and allows to study consistently the interplay between QCD and EW radiation, *e.g.*, the link between a photon emitted after QCD radiation and viceversa.



See also C. Bernaciak, D.W., arXiv:1201.4804

$pp \rightarrow l^+l^- X$ and $pp \rightarrow \nu l X$ at $\mathcal{O}(\alpha\alpha_s)$ in pole approximation: S. Dittmaier, A. Huss, C. Schwinn, arXiv:1405.6897; 1403.3216; 1511.08016

Carlioni Calame *et al*, 1612.02841



When the characteristic energies are larger than $M_{W,Z}$ higher-order EW corrections may be approximated as an expansion in EW Sudakov logs:

- Results (fixed order and resummed to all orders) are available for hadronic cross sections for, e.g., $V(+\text{jets})$, VV , $t\bar{t}$, bb , cc , jj , and VBF.

$Z + \leq 3$ jets in ALPGEN [Chiesa et al \(2013\)](#)

$Z, t\bar{t}, jj$ production implemented in MCFM [Campbell, D.W., Zhou \(2016\)](#)

- Best studied so far for four-fermion process $f\bar{f} \rightarrow f'\bar{f}'$:
 - up to $N^3\text{LL}$ for massless fermions ($a = \frac{\alpha}{4\pi s_W^2}$, $L = \log(s/M_W^2)$):

$$\frac{\delta\sigma(e^+e^- \rightarrow q\bar{q})(s)}{\sigma_{LO}} = -2.18aL^2 + 20.94aL - 35.07a + 2.79a^2L^4 - 51.98a^2L^3 + 321.20a^2L^2 - 757.35a^2L$$

[Jantzen, Kühn, Penin, Smirnov, hep-ph/0509157](#)

- up to NNLL for massive fermions [Denner, Jantzen, Pozzorini \(2008\)](#).
- up to NLL for $V + \text{jets}$ [J. H. Kuhn, A. Kulesza, S. Pozzorini, M. Schulze \(2005,2007\)](#); see also [J.Lindert et al, 1705.04664](#)
- Impact of real W, Z radiation [Baur \(2006\)](#); [Bell et al \(2010\)](#); [Manohar et al \(2014\)](#)
- Resummation with SCET [Chiu et al, \(2008,2009\)](#); [Manohar, Trott \(2012\)](#); [Bauer, Ferland \(2017\)](#)

“Amplitude calculators”:

- **Recola** S.Actis *et al*, 1605.01090 + **Collier** A.Denner *et al* 1604.06792
- **OpenLoops** F.Cascioli *et al*, 1111.5206
- **Gosam** M.Chiesa *et al*, 1507.08579 + **MadDipole** T.Gehrmann *et al*, 1011.0321
- **Madgraph5_aMC@NLO** J.Alwall *et al*, 1405.0301

Some recent results for multi-particle processes which consistently include higher-order QCD and EW corrections implemented in PS MCs:

- **Recola+Sherpa** B.Biedermann *et al*, 1704.05783
Examples: $pp \rightarrow V + \text{jets}$, $pp \rightarrow ZZ \rightarrow 4 \text{ leptons}$, $pp \rightarrow t\bar{t}H$
- **OpenLoops+Munich/Sherpa**
Examples: $pp \rightarrow W + 1, 2, 3 \text{ jets}$, S.Kallweit *et al*, 1412.5157, and $V + 1, 2 \text{ jets with } V \rightarrow ll'$ and MEPS@NLO jet merging, S.Kallweit *et al*, 1511.08692; $pp \rightarrow 2\nu 2l$, S.Kallweit *et al*, 1705.00598
- **GOSAM+Sherpa** M.Chiesa *et al*, 1706.09022
Examples: $pp \rightarrow \gamma\gamma + 0, 1, 2 \text{ jets}$
- **Madgraph5_aMC@NLO**
Examples: $pp \rightarrow t\bar{t} + (H, Z, W)$, S.Frixione *et al*, 1504.03446; $pp \rightarrow jj$, Frederix *et al*, 1612.06548

- Electroweak corrections have a reach structure and their numerical impact strongly depend on the kinematic regime, details of the lepton identification/analysis cuts, interplay with QCD radiation, etc.
- Predictions for a selected number of SM processes have reached a high level of sophistication.
- Automated calculations of NLO EW+QCD corrections are becoming available, matched to PS Monte Carlo programs.
- We can look forward to a new era of EW phenomenology.