

# Higgs doublet decay as the origin of the baryon asymmetry

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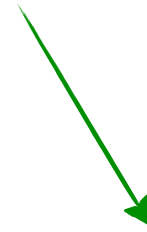
Based on: TH, Daniele Teresi, PRL 2016, arXiv:1606.00017

# Baryogenesis via leptogenesis


 very highly motivated: same origin as neutrino masses



very natural at high scale: a series of numerical coincidences which make it particularly efficient but very difficult to test



clearly possible at low scale: if seesaw seesaw states have a quasi-degenerate mass spectrum and/or if large cancellation among Yukawa couplings

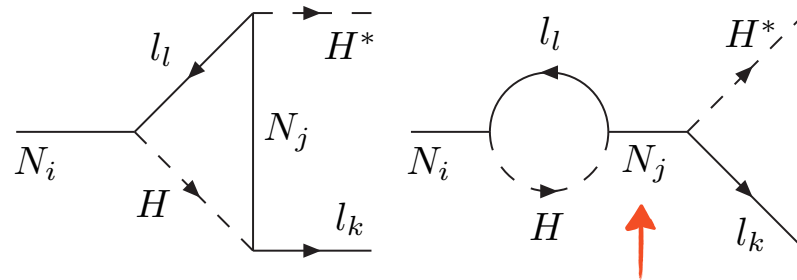
 this talk: new way at low scale: total lepton number violating Higgs doublet decay into  $\sim 0.1$ -100 GeV right-handed neutrinos

# Leptogenesis relevant scales for low $m_N$

$$T_{Sphaler.} \sim 135 \text{ GeV}$$

usual leptogenesis:  $m_N \gg T_{Sphaler.} > m_{H,L}$  : leptogenesis from  $N \rightarrow LH$  decay

↪ creation of L asymmetry at  $T \sim m_N \gg T_{Sphaler.} \Rightarrow$  B asymmetry



resonant propagator if  $m_{N_j} \sim m_{N_i} \Rightarrow$   $\sim$ TeV scale leptogenesis

very low scale leptogenesis:  $T_{Sphaler.} > m_H \gg m_{N,L}$

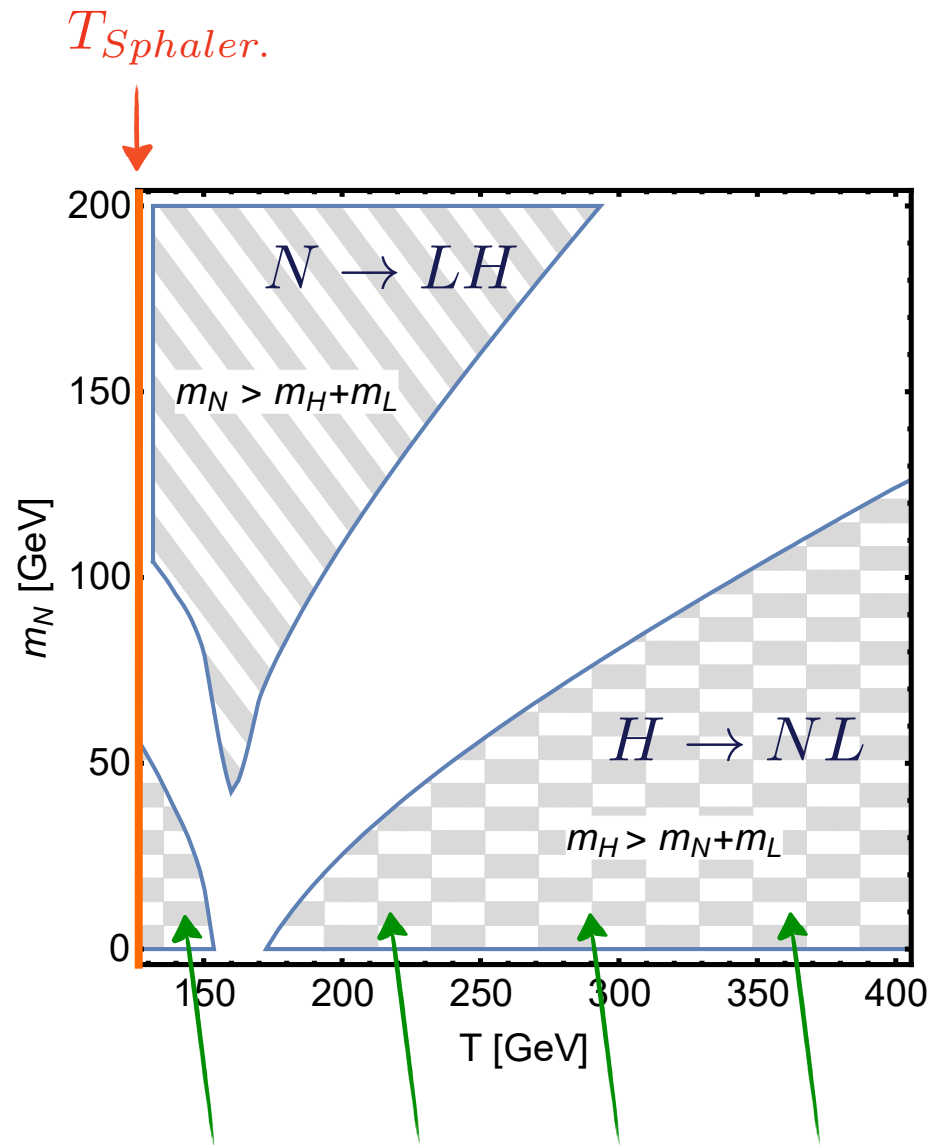
↪ creation of L asymmetry at  $T > T_{Sphaler.} \gg m_N \Rightarrow \neq$  regime

↪ thermal effects are fully relevant:  $T > T_{Sphaler.} > m_H \gg m_{N,L}$

$$m_H^2(T) = m_H^2 + c_H \cdot T^2 \quad m_L^2(T) = m_L^2 + c_L \cdot T^2 \quad m_N^2(T) = m_N^2 + c_N \cdot T^2$$

↪  $N \rightarrow LH$  forbidden but  $H \rightarrow NL$  allowed

# Temperatures allowing the $N \rightarrow LH$ and $H \rightarrow NL$ decays



$H \rightarrow NL$  leptogenesis from this region?

# *L asymmetry production from $H \rightarrow NL$ decay*



 2 issues at first sight:

1) out-of-equilibrium decay?  3rd Sakharov condition

  $H$  decaying particle is in deep thermal equilibrium at  $T > T_{Sphaler.}$

but  $N$  in decay product is not necessarily in thermal equilibr.

$$\frac{dn_N}{dt} \propto (n_N^{eq} - n_N) \cdot \Gamma_{H \rightarrow NL}$$

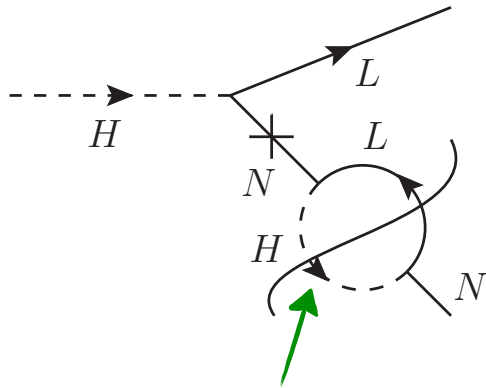
 

$$H \rightarrow NL \quad NL \rightarrow H$$

# *L asymmetry production from $H \rightarrow NL$ decay*

↪ 2 issues at first sight:

2) Absorptive part for CP violation?



$m_H + m_L > m_N \Rightarrow$  no absorptive part?

↪ but only for  $T = 0!$

finite T corrections: thermal cut: if  $H$  or  $L$  comes from the thermal bath the cut is kinematically allowed

Giudice, Notari, Raidal, Riotto, Strumia 03'

Frossard, Garny, Hohenegger, Kartavtsev, Mitrouskas 12'

↪ absorptive part  $\Gamma_N(T)$  (calculated in Kadanoff Baym formalism)

# Total $L$ number violating $CP$ asymmetry

$$\epsilon_{CP} = \frac{\text{Im}[(Y_N Y_N^\dagger)_{12}^2]}{(Y_N Y_N^\dagger)_{11} (Y_N Y_N^\dagger)_{22}} \cdot \frac{2 \Delta m_N^0 \Gamma_N(T)}{4 \Delta m_N(T)^2 + \Gamma_N(T)^2}$$

↪ with thermal mass splitting:  $\Delta m_N(T) \simeq \Delta m_N^0 + \frac{\pi T^2}{4 m_N^2} \Gamma_{22} \sqrt{\left(1 - \frac{\Gamma_{11}}{\Gamma_{22}}\right)^2 + 4 \frac{|\Gamma_{12}|^2}{\Gamma_{22}^2}}$

$$\Gamma_{ij} \equiv m_N (Y_N Y_N^\dagger)_{ij} / (8\pi)$$

Boltzmann equations:

$$\frac{n_\gamma H_N}{z} \frac{d\eta_N}{dz} = \left(1 - \frac{\eta^N}{\eta_N^{\text{eq}}}\right) \left[ \gamma_D + 2(\gamma_{Hs} + \gamma_{As}) + 4(\gamma_{Ht} + \gamma_{At}) \right],$$

$$\frac{n_\gamma H_N}{z} \frac{d\eta_L}{dz} = \gamma_D \left[ \left( \frac{\eta^N}{\eta_N^{\text{eq}}} - 1 \right) \epsilon_{CP}(z) - \frac{2}{3} \eta_L \right] - \frac{4}{3} \eta_L \left[ 2(\gamma_{Ht} + \gamma_{At}) + \frac{\eta^N}{\eta_N^{\text{eq}}} (\gamma_{Hs} + \gamma_{As}) \right]$$

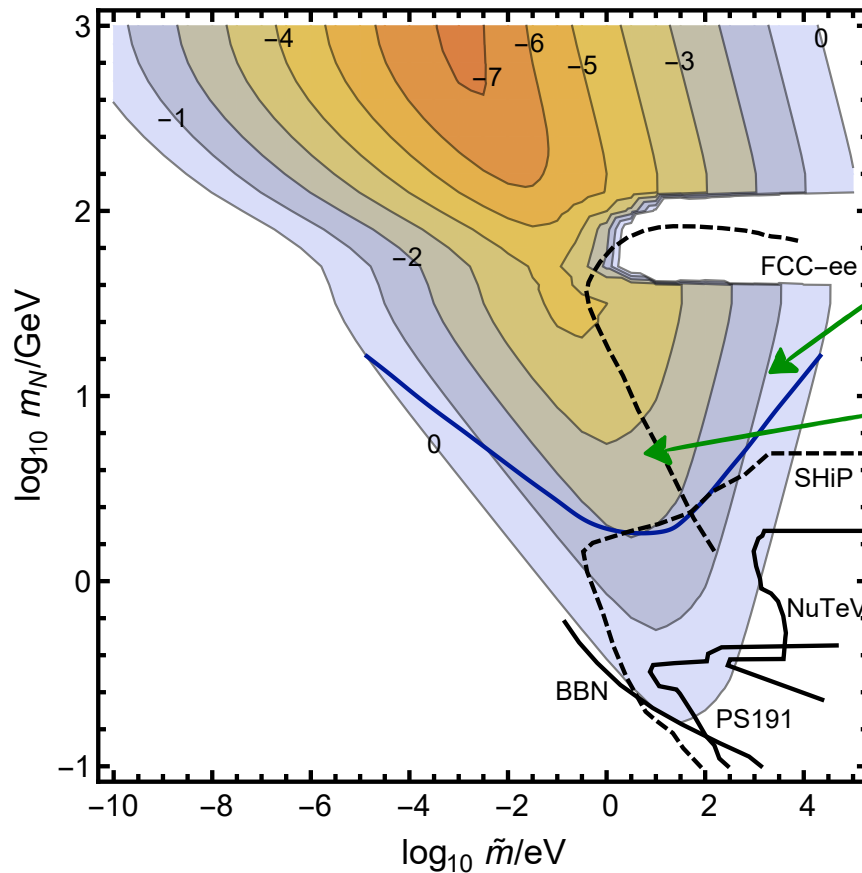
$$\eta_N \equiv n_N / n_\gamma$$

$$z \equiv m_N / T$$

# Results for the case where the $N$ have thermalized

if  $N$  thermalized by large  $Y_N$  Yukawas or other interaction (e.g. a  $W_R$ ) before an asymmetry is produced

CP-asymmetry needed for successful leptog.



the lower is  $m_N$ , the later it goes out-of-equilibrium, the more it will be in equilibrium at  $T > T_{Sphaler}$ .



lower bound on  $m_N$

$$m_N > 2.2 \text{ GeV}$$

if only  $N \rightarrow LH$  decay we get:  $m_N > 50 \text{ GeV}$

$$\tilde{m} \equiv \frac{Y_N Y_N^\dagger v^2}{2m_N}$$

requires that at least 2 of the  $N$  have quasi-degenerate masses



# Results for the case where the $N$ have not thermalized

- if no extra interaction thermalizing  $N$ , no thermalization is much more natural than in ordinary leptogenesis: thermalization at  $T > T_{Sphaler.} \gg m_N$  requires much larger  $Y_N$  Yukawas than in ordinary leptogenesis at  $T \sim m_N$

$$\tilde{m} \equiv \frac{Y_N Y_N^\dagger v^2}{2m_N}$$

$$\tilde{m} \gg 10^{-3} \text{ eV}$$

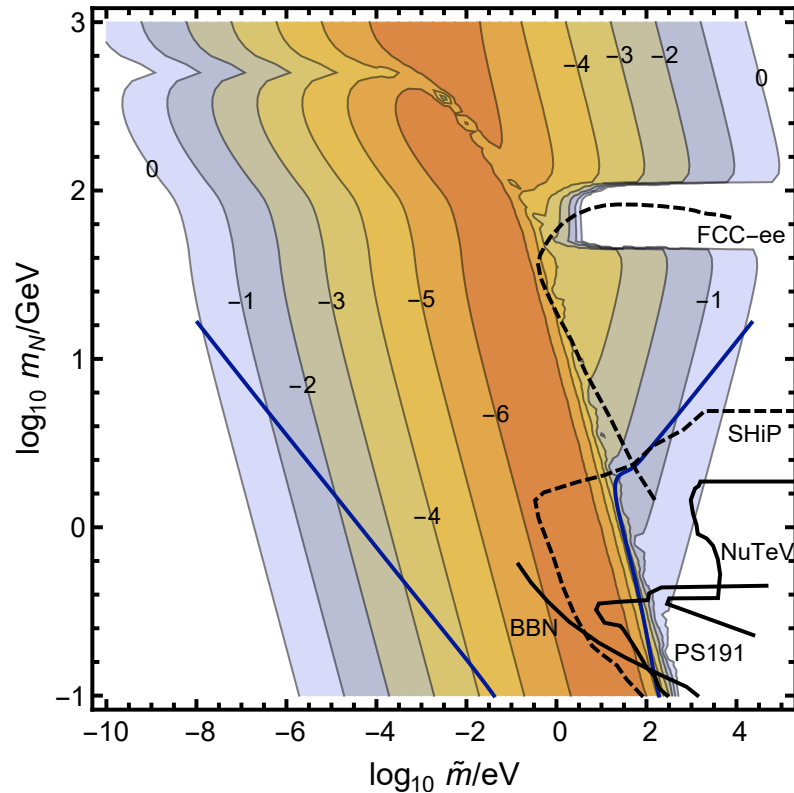
$$\tilde{m} \gtrsim 10^{-3} \text{ eV}$$

- for  $H \rightarrow NL$  decay, to start from no  $N$  in the thermal bath boosts the asymmetry production, unlike for ordinary  $N \rightarrow LH$  leptogenesis

$H \rightarrow NL$ : many  $H$  to decay and produce the asymmetry but few  $N$  to  $NL \rightarrow H$  inverse decay

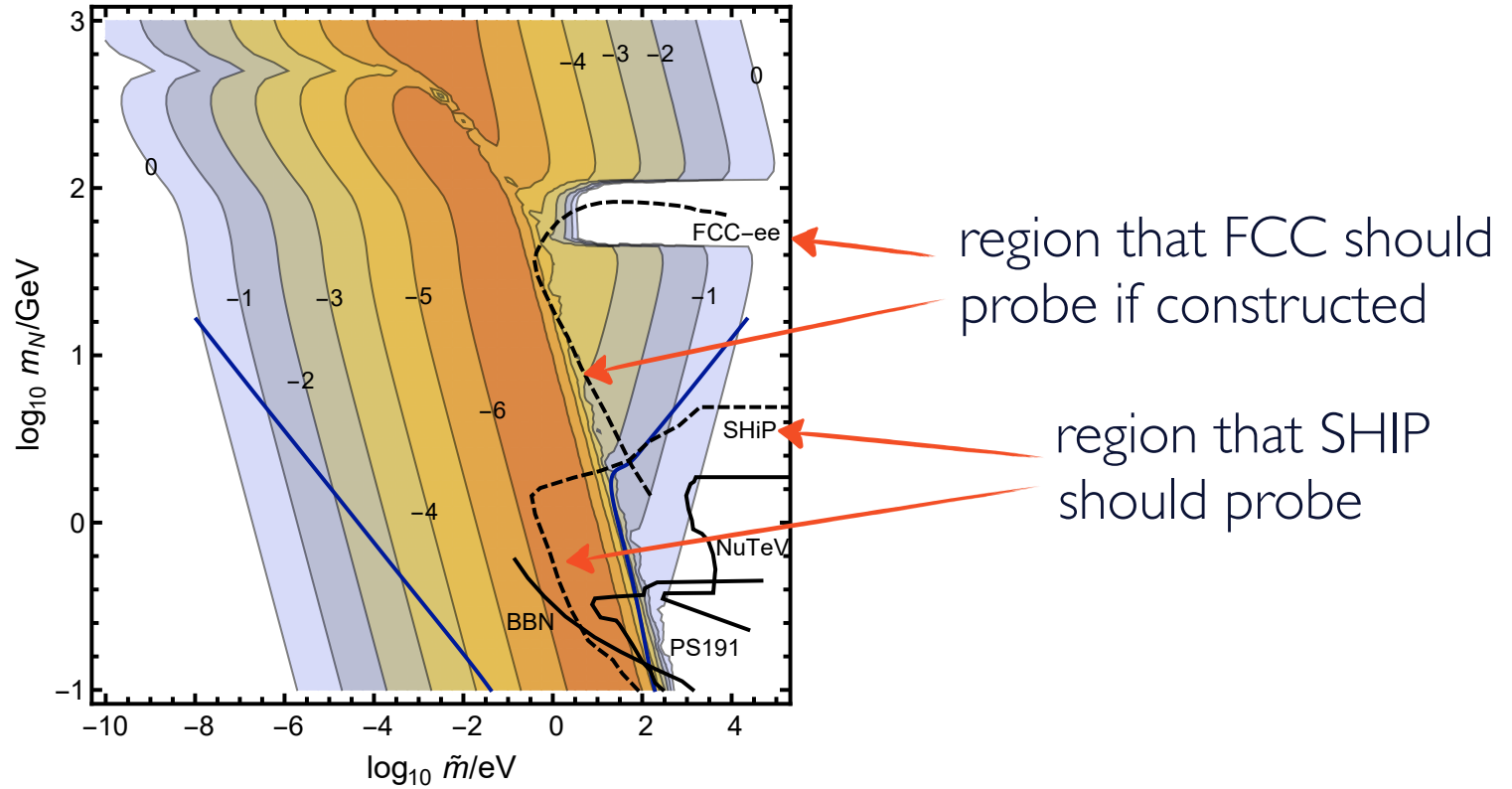
$$n_N^{eq} - n_N \sim n_N^{eq} \gg n_N$$

# Results for the case where the $N$ have not thermalized



- ↪ for example for  $m_N \sim 10 \text{ GeV}$  and  $\tilde{m} \sim 0.1 \text{ eV}$  one needs  $\Delta m_N^0/m_N \lesssim 10^{-5}$
- ↪ leptogenesis for  $m_N$  as low as  $\sim 20 \text{ MeV}$  is possible (but BBN concerns)
- ↪ in all cases: asymmetry production at  $T$  just above  $T_{Sphaler}$ .  $\Rightarrow$  no dependence on UV physics!

# Testability!



# Links with other very low scale frameworks

Akhmedov, Rubakov, Smirnov 98'  
 Asaka, Shaposhnikov 05'; Shaposhnikov 08'  
 Drewes, Garbrecht 11'  
 Canetti, Drewes, Frossard, Shaposhnikov 13'  
 Hernandez, Kekic, Lopez-Pavon, Racker, Rius 15'  
 .....

## $N$ oscillation frameworks

↪ based on density matrix formalism

↪ based on purely flavour asymmetries  $> <$   $H \rightarrow NL$  total L number violating framework which doesn't require flavour

$$(Y_N Y_N^\dagger)_{ll} > (Y_N Y_N^\dagger)_{l'l'} \leftarrow \text{especially if large } \Delta m_N$$

↪ has been considered in many  $\neq$  regimes: large/small  $\Delta m_N$ , large/small  $\tilde{m}$ , 2 or 3  $N$ , many/few oscillat., ...

↙  
flavour asym.

$$\propto Y_N^2 T^2$$

↪ production at  $T \gg \gg T_{Sphaler.}$

$$\propto \frac{Y_N^6 M_{Planck}^{4/3}}{(\Delta m_N^2)^{2/3}}$$

↘  
L violating asym.

$$\propto m_N^2$$

↪ production at  $T \gtrsim T_{Sphaler.}$

$$\propto \frac{Y_N^4 m_N^2}{\Delta m_N^2}$$

⇒ e.g. dominance or L violating asym. for  $\Delta m_N$  not too large, especially if no big cancellations between the  $Y_N$  in  $m_\nu$  or if flavour hierarchy is small or low reheating temperature

# Links with other very low scale frameworks

- ↪ as a matter of principle the  $L$  violating decay contribution must also appear in density matrix formalism
- ↪ there is a  $L$  violating term showing up in density matrix Boltzmann equations: `` $R_M$ `` term, considered to be negligible for  $T \gg m_N$  and thus neglected for baryon asymmetry production

Shaposhnikov 08'

Canetti, Drewes, Frossard, Shaposhnikov 13'

we are computing and comparing the  $R_M$  contribution in  $N$ -oscillation formalism to our decay formalism contribution

# Summary

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In usual leptogenesis decay formalism the L violating  $H \rightarrow NL$  decay can easily lead to enough baryon asymmetry for  $m_N < m_H$

↳ in type-I seesaw model with nothing else

↳ thanks to thermal effect leading to  $N$  self-energy thermal cut

↳ from total L number violating CP asymmetries: no need for flavour interplay

↳ at electroweak scale temperatures:  $T \gtrsim T_{Sphaler.}$

↳ with boosted production if no  $N$  to begin with

↳ in a testable way (SHIP,...) for part of the parameter space

We are looking at same effect in density matrix formalism...



baryon asymmetries obtained for 3 values of  $\Delta m_N^0/m_N$

