

# SEARCHING FOR DARK PARTON SHOWERS AT THE LHC

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University of Oregon

*based on*

Phys. Rev. Lett. 115, 171804 (2015)

[arXiv: 1503.00009]

with M. Lisanti, H. K. Lou

*and in progress*

with M. Lisanti, H. K. Lou, S. Mishra-Sharma

**CERN-CKC Workshop:  
“What’s Going on at the Weak Scale?”  
Jeju Island, South Korea  
June 2, 2015**

# WHAT IS DARK MATTER?!?

What  
comprises  
the dark  
sector?



# WHAT IS DARK MATTER?!?

What  
comprises  
the dark  
sector?

Weakly coupled?  
(e.g. neutralinos)



# WHAT IS DARK MATTER?!?



# WHAT IS DARK MATTER?!?

How do we ensure  
discovery????

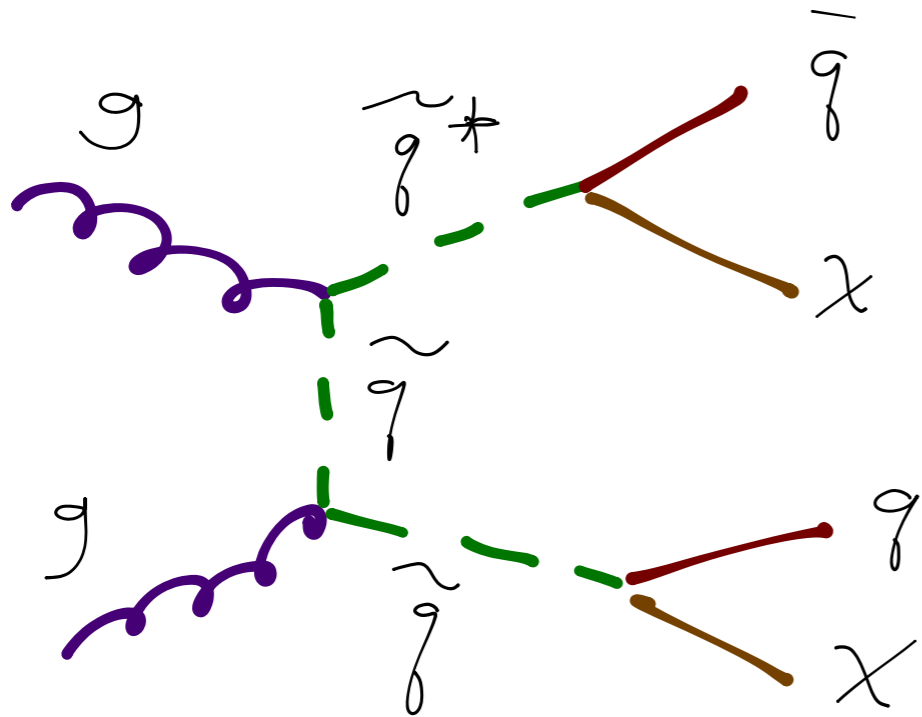
Can we better optimize  
searches????



# "WIMP" SIGNALS

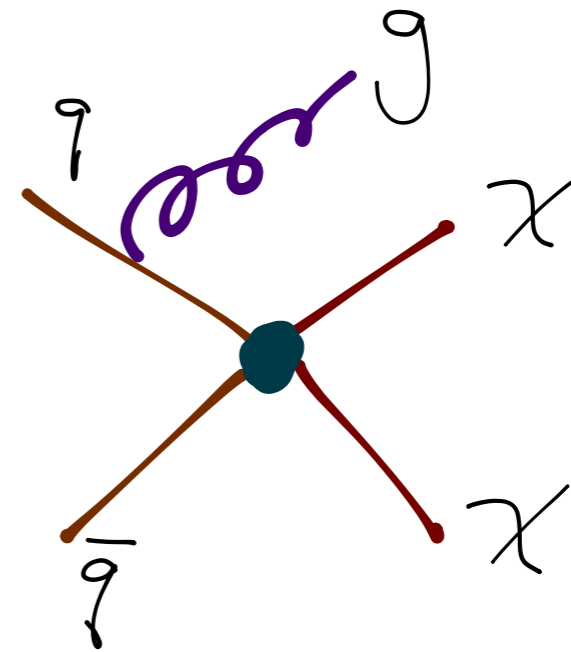
New physics in Jets +  $\cancel{E}_T$ .

## Simplified Model



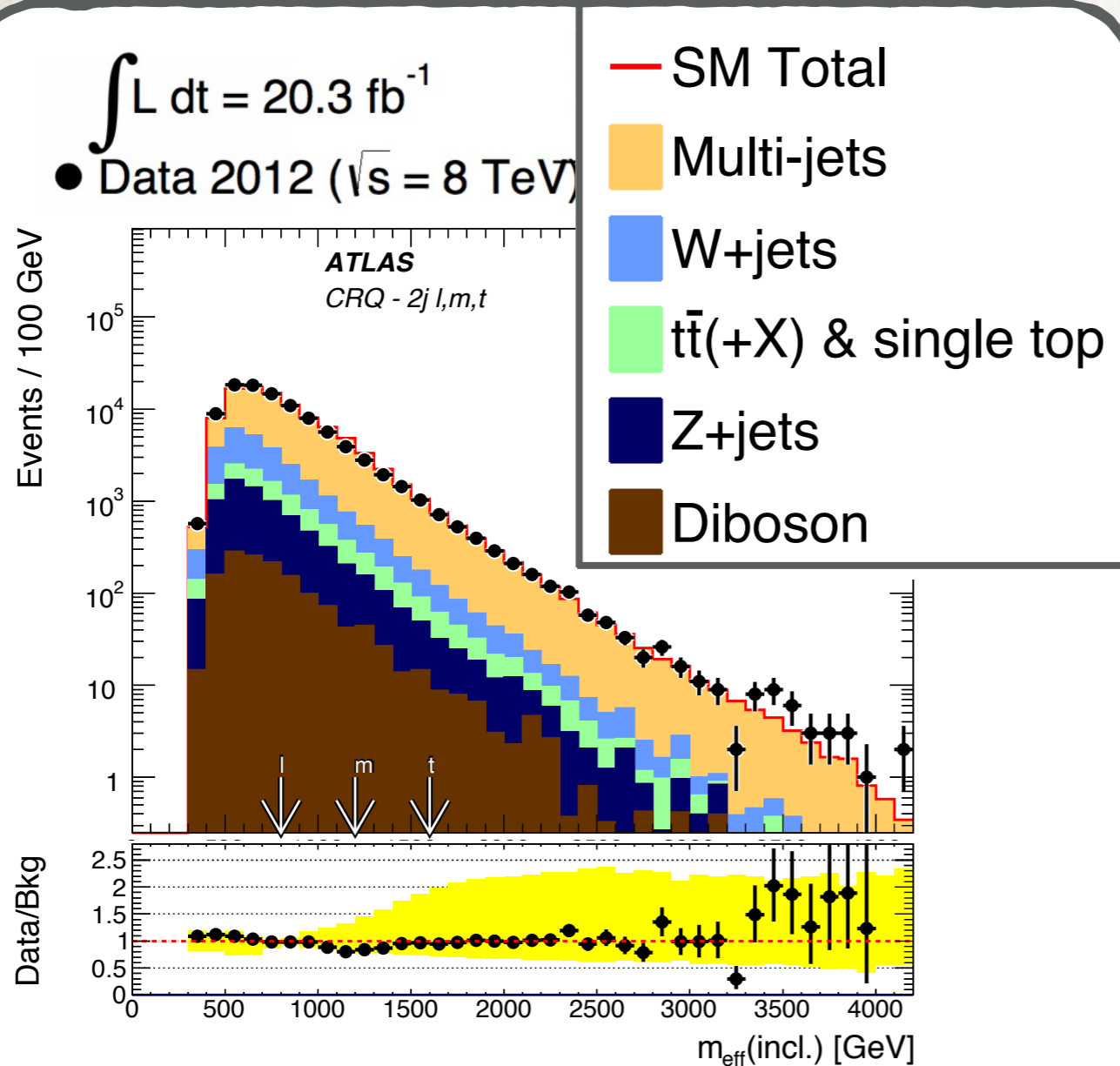
Alwall, Le, Lisanti, Wacker [arXiv:0809.3294]  
Alwall, Schuster, Toro [arXiv:0810.3921]  
LHC New Physics Working Group [arXiv:1105.2838]

## Contact operator



Fox, Harnik, Kopp, Tsai [arXiv:1109.4398]  
Rajaraman, Shepherd, Tait, Wijangco [arXiv:1108.1196]

# BACKGROUNDS



Control region:  $\Delta\phi(\text{jets}, \cancel{E}_T)$  cut reversed.

ATLAS Collaboration [arXiv:1405.7875]

QCD background:

$\cancel{E}_T$  from jet  
mismeasurement.

Background  $\cancel{E}_T$  aligned  
with jets.

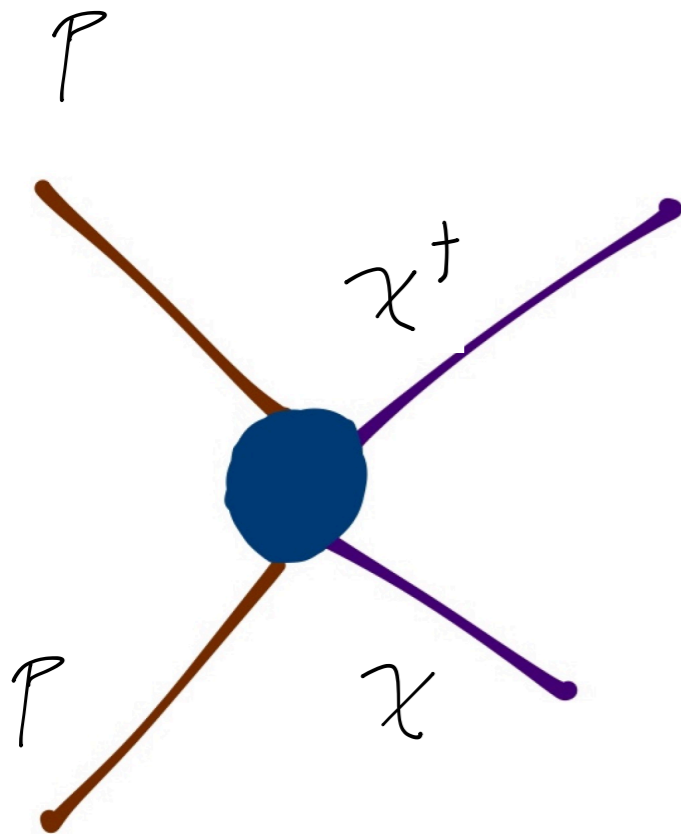
Cut on

$$\min \left[ \Delta\phi(\text{jets}, \cancel{E}_T) \right] \gtrsim 0.4$$

# DARK SHOWER

Signature made classic by "Hidden Valley" models.

Strassler, Zurek [arXiv:hep-ph/0604261]

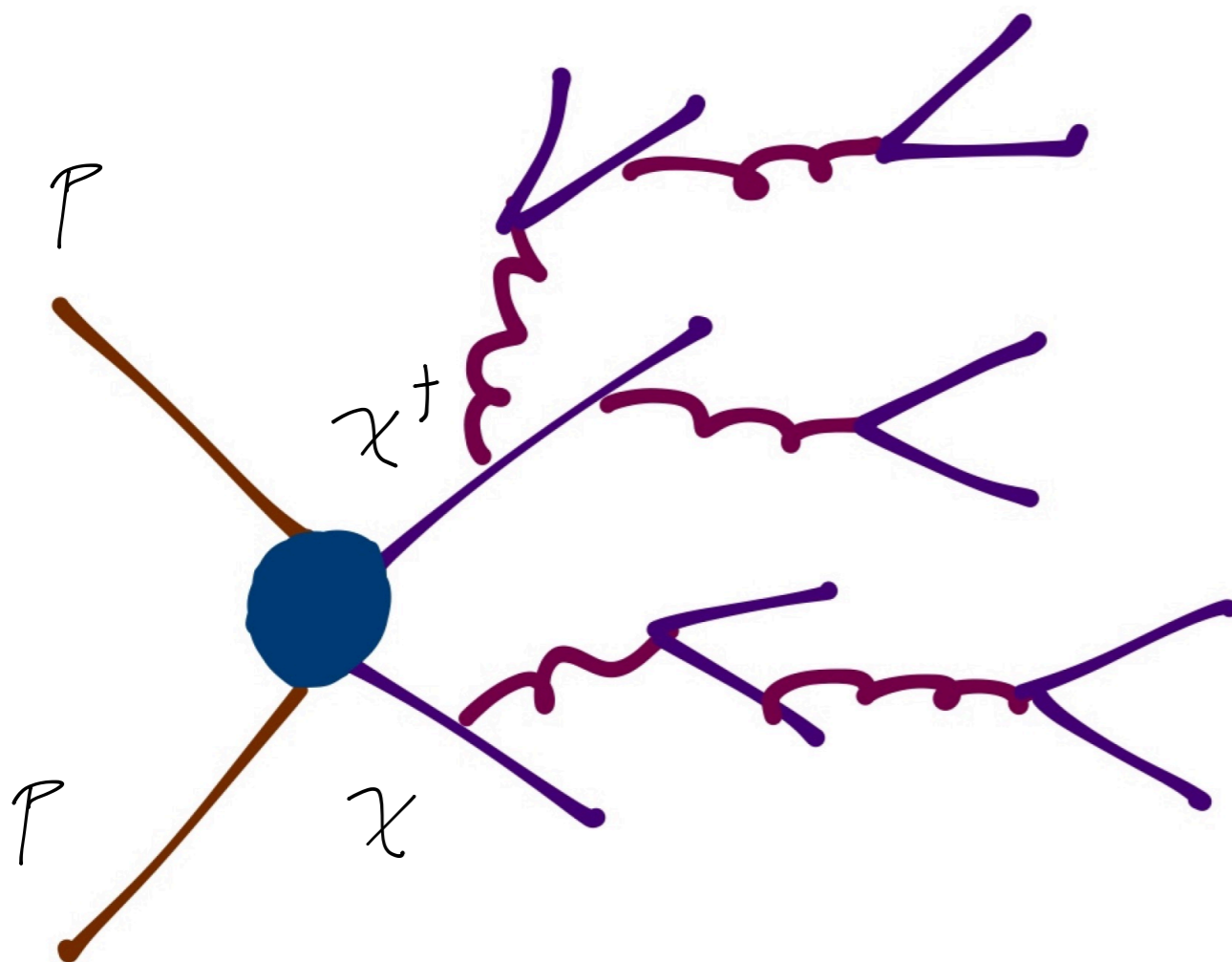




# DARK SHOWER

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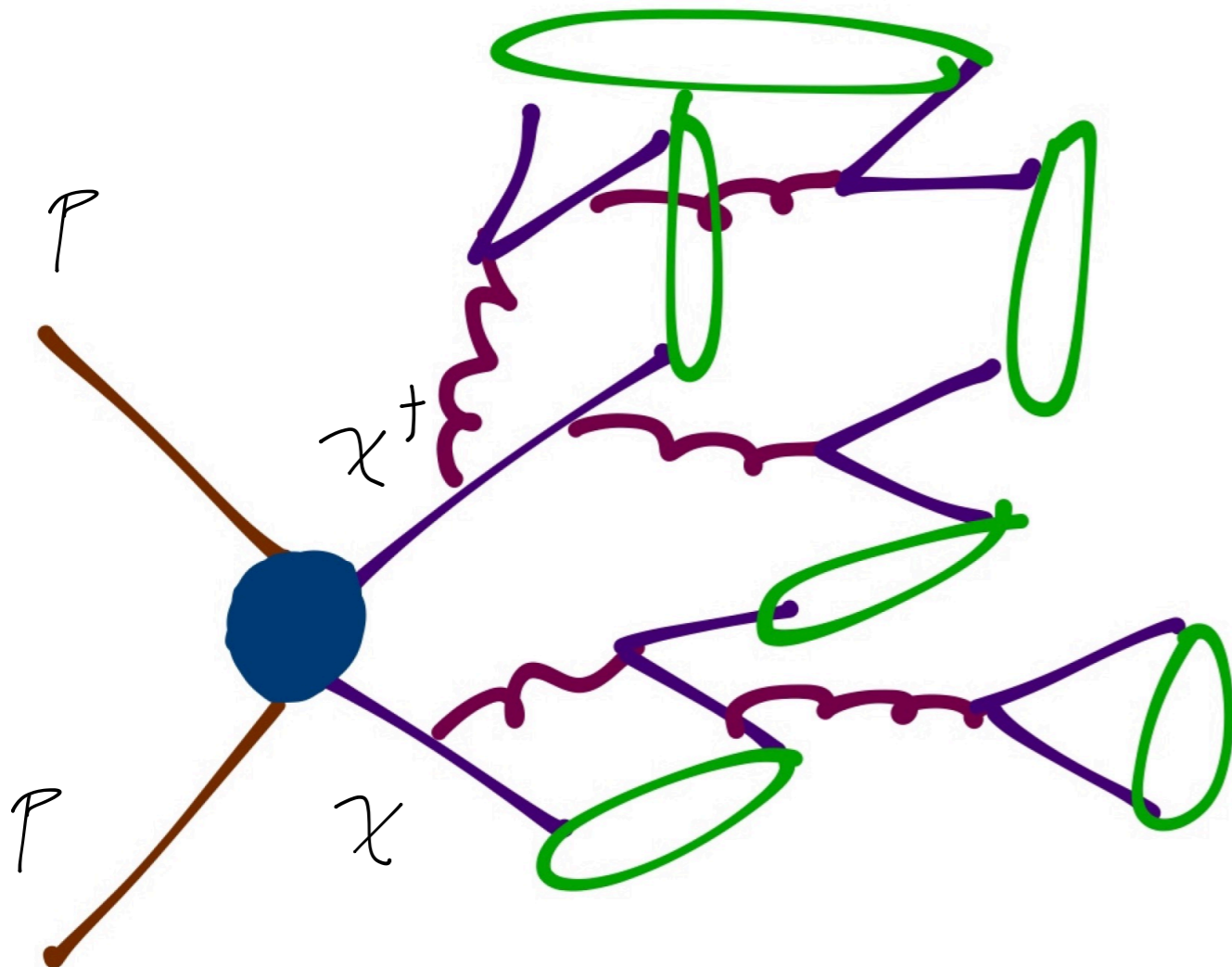
Strassler, Zurek [arXiv:hep-ph/0604261]



# DARK SHOWER

Signature made classic by "Hidden Valley" models.

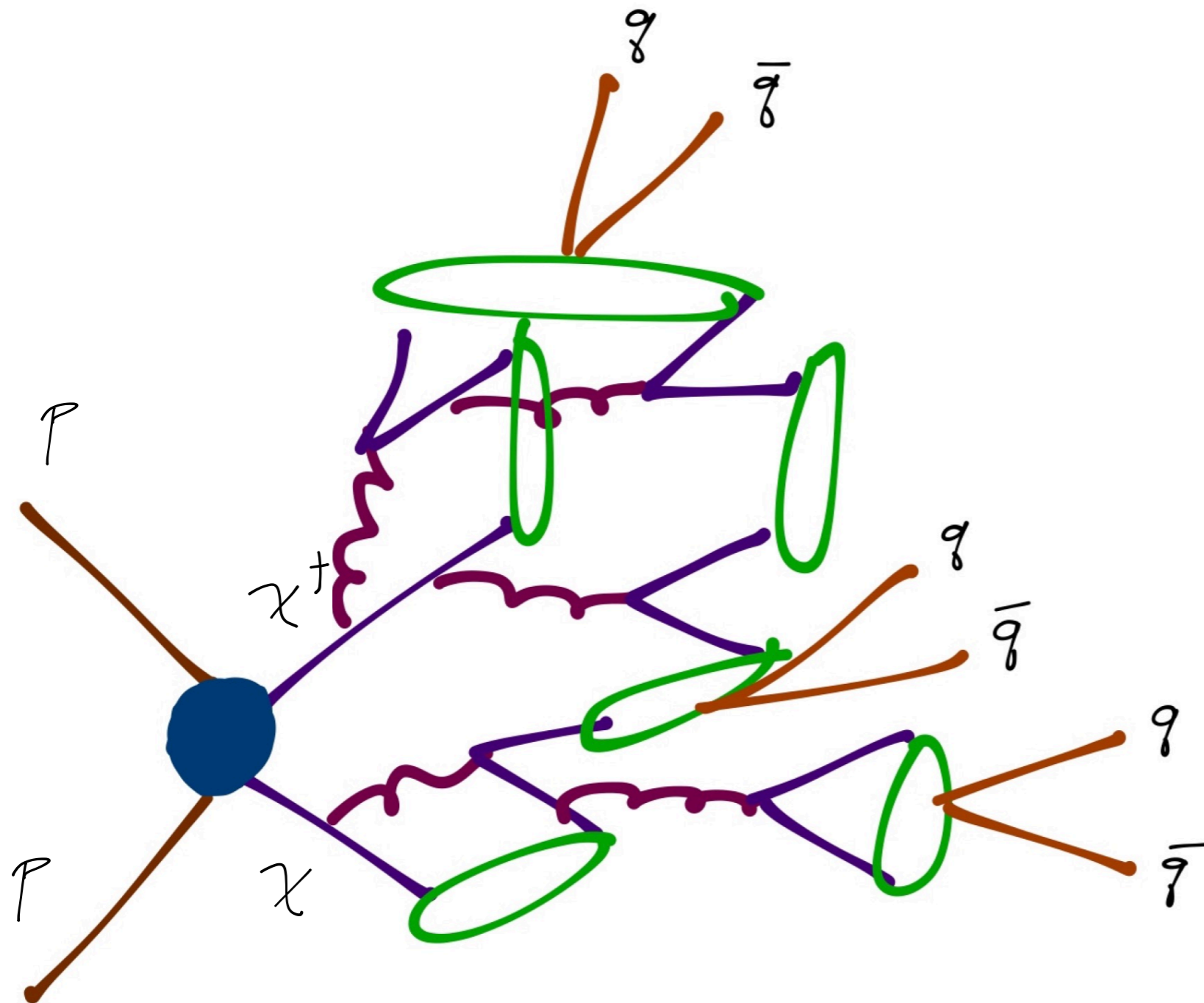
Strassler, Zurek [arXiv:hep-ph/0604261]



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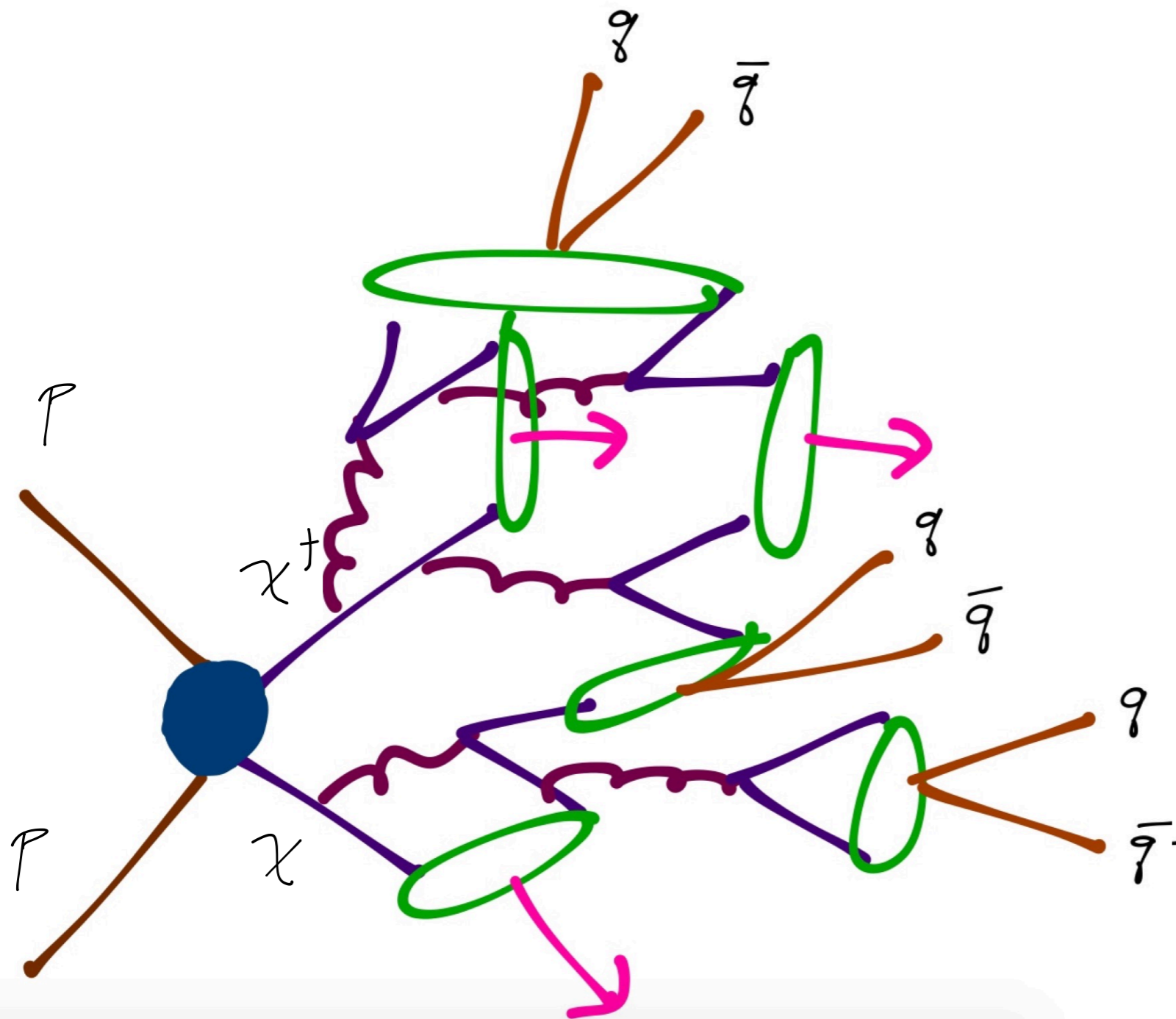
Strassler, Zurek [arXiv:hep-ph/0604261]



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Signature made classic by "Hidden Valley" models.

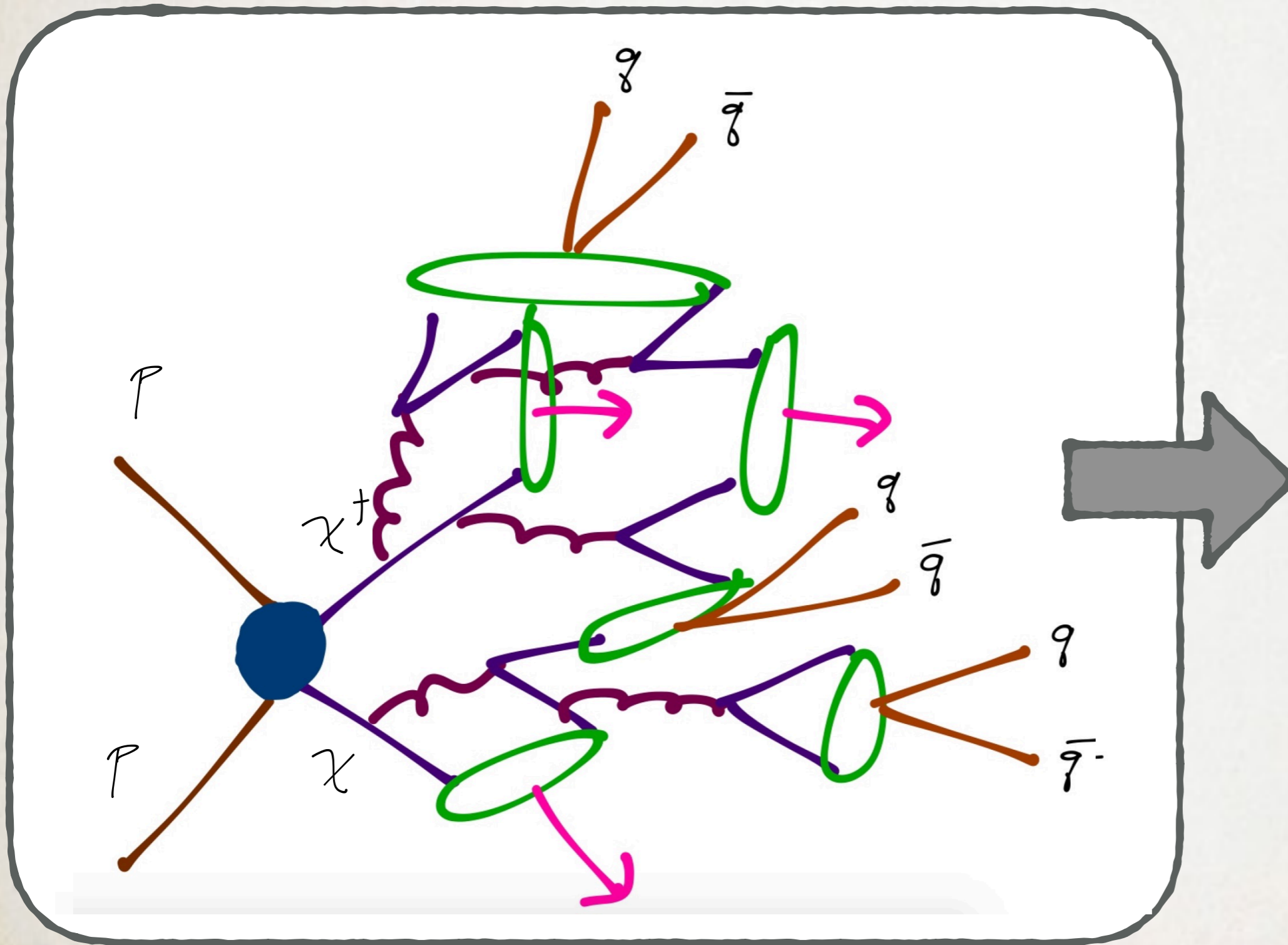
Strassler, Zurek [arXiv:hep-ph/0604261]



# DARK SHOWER

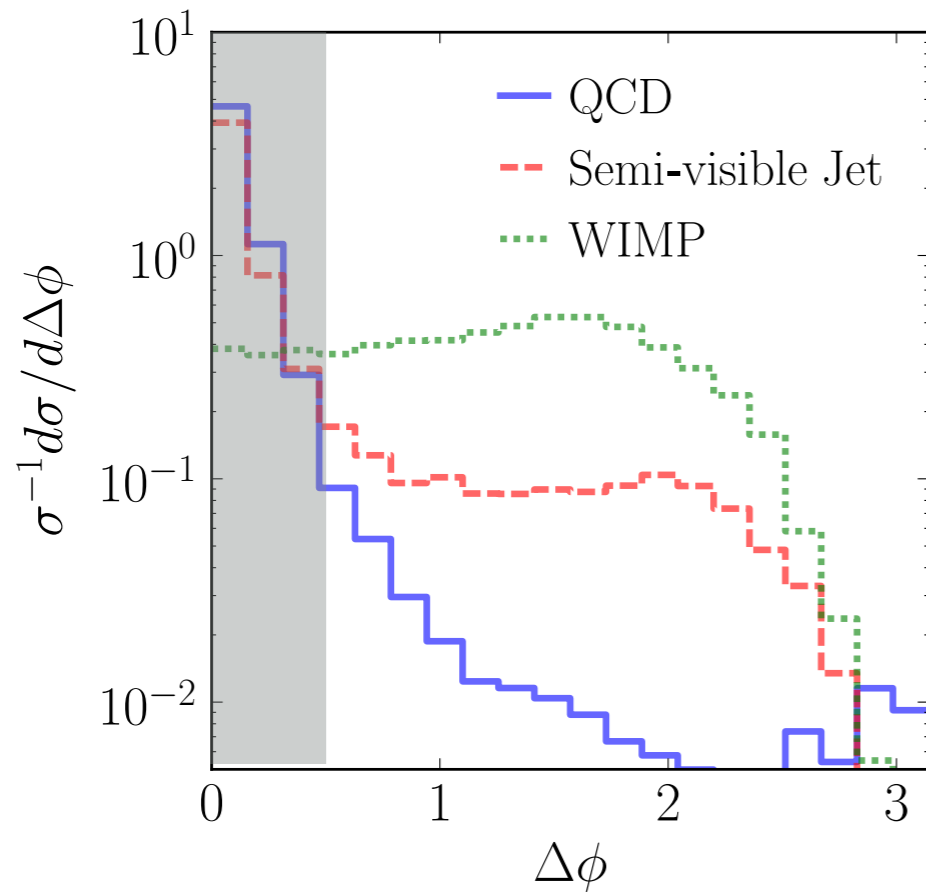
Signature made classic by "Hidden Valley" models.

Strassler, Zurek [arXiv:hep-ph/0604261]



Jet of visible matter and dark matter: "semi-visible" jet

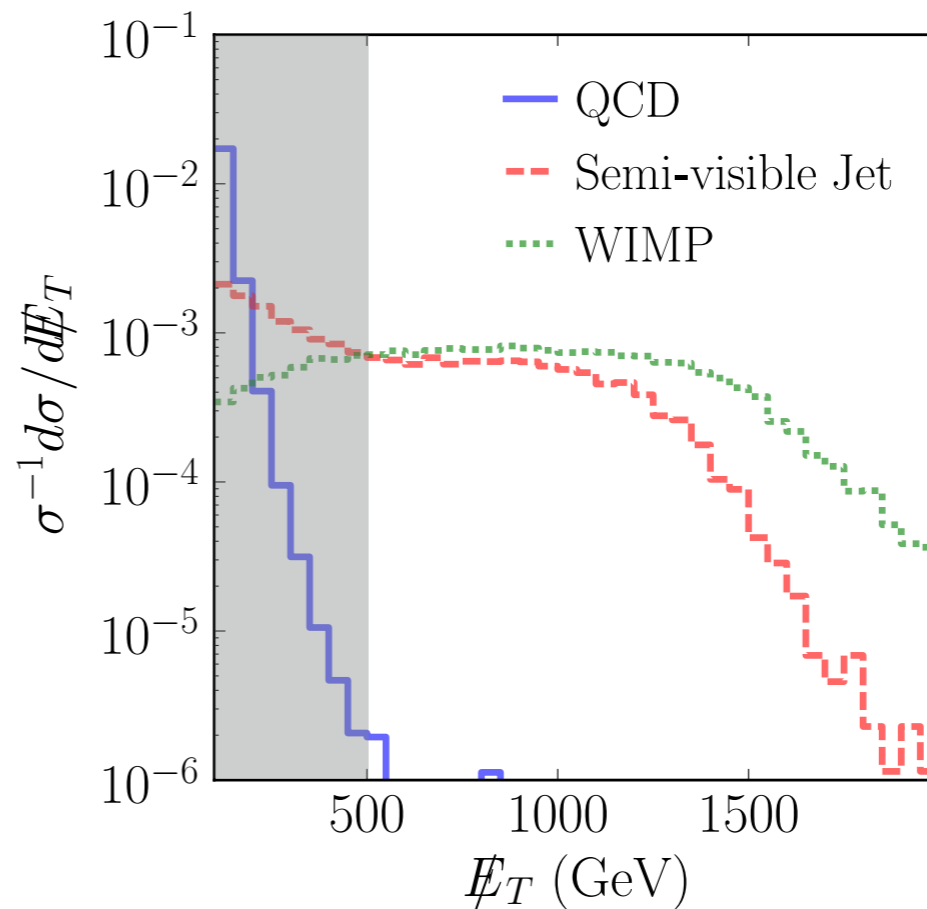
# DISTRIBUTIONS



“Standard” cuts:

$$\cancel{E}_T \gtrsim 500 \text{ GeV}$$

$$\Delta\phi \gtrsim 0.4$$



Efficiencies:

$\sim 40\%$  WIMP

$\sim 1\%$  Semi-visible jets

# NO STONE UNTURNED

Want to ensure discovery of new physics.

Nothing stopping the dark sector from being complicated.

**Goal:** minimal parametrization to cover broad class of models.

# OUTLINE

MODELS

SEARCHES

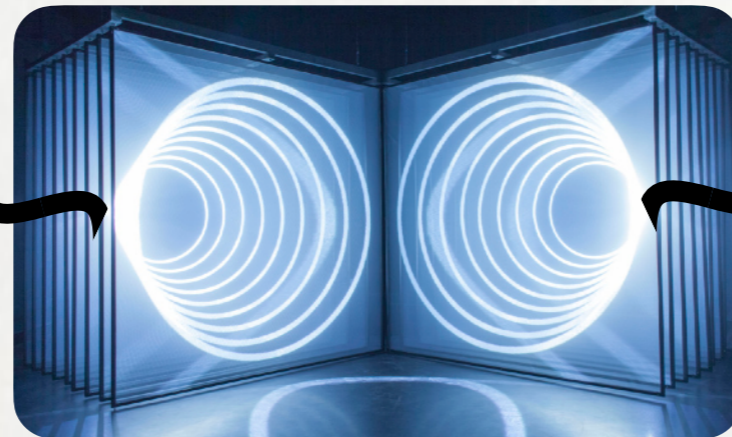
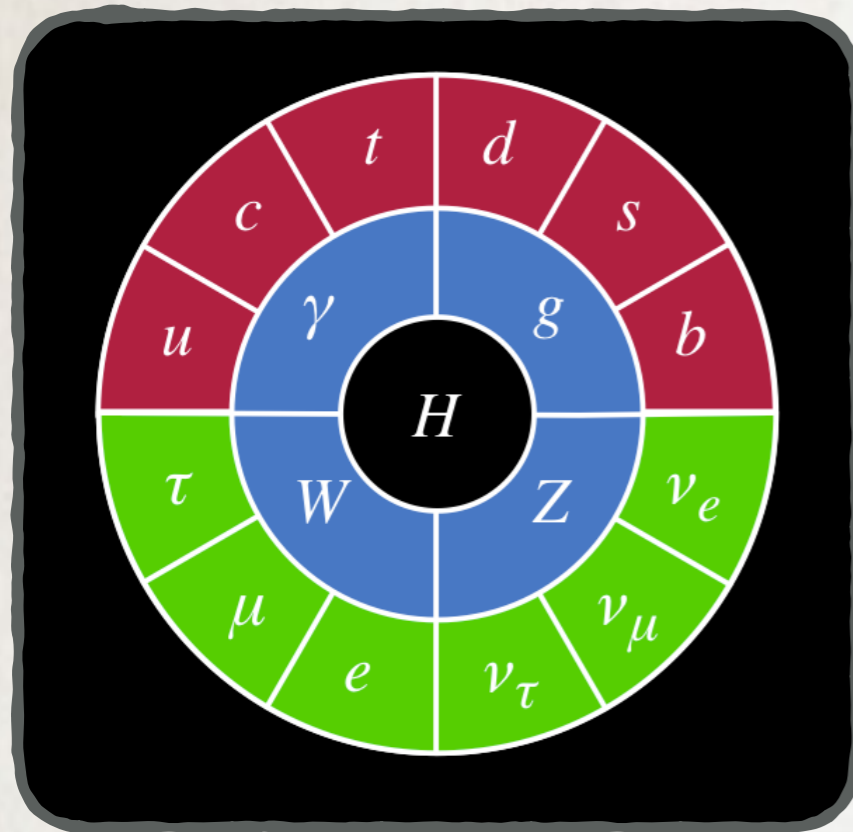
OUTLOOK



# MODELS

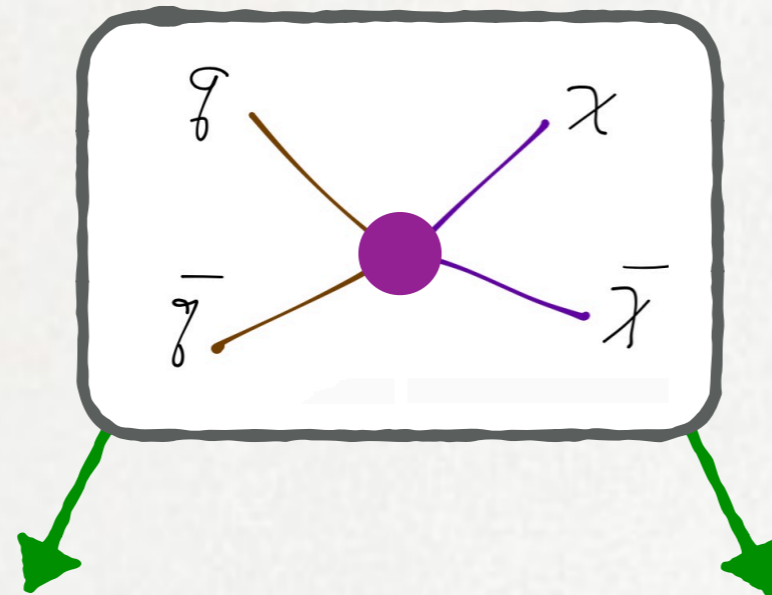
**Simulation:** pythia8 hidden valley module (modified to include running couplings) piped through DELPHES (CMS card) at 14 TeV.

# PORTAL INTO THE DARK SECTOR

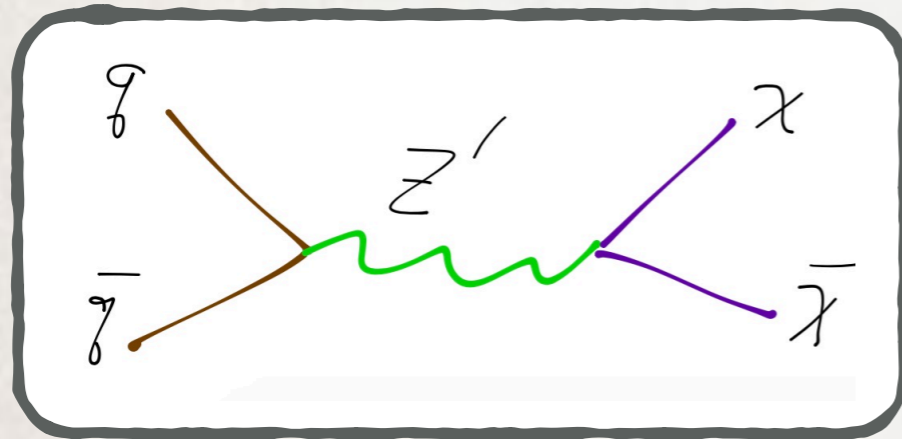


# PORTALS

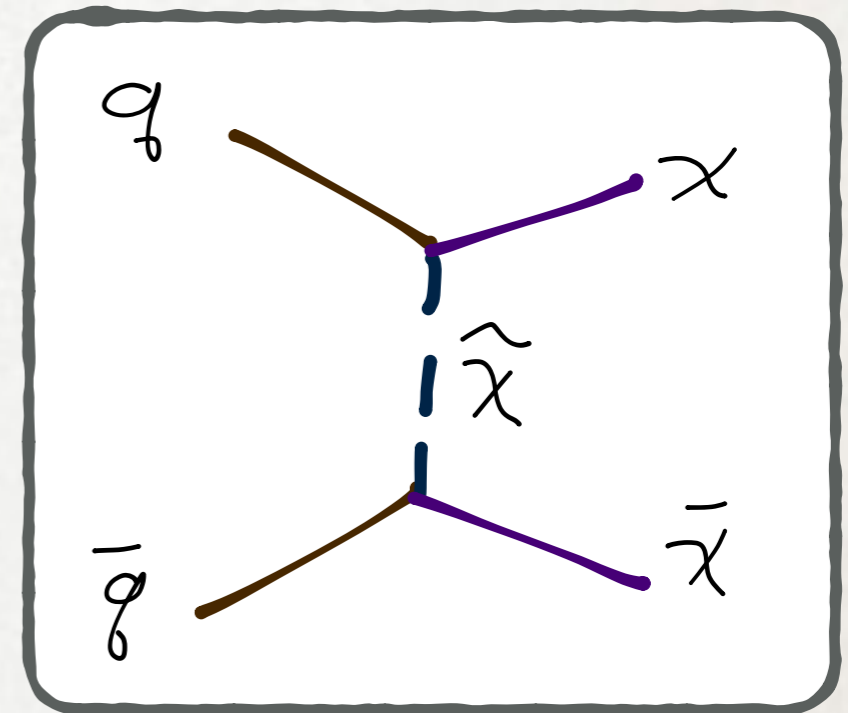
Contact operator



s-channel

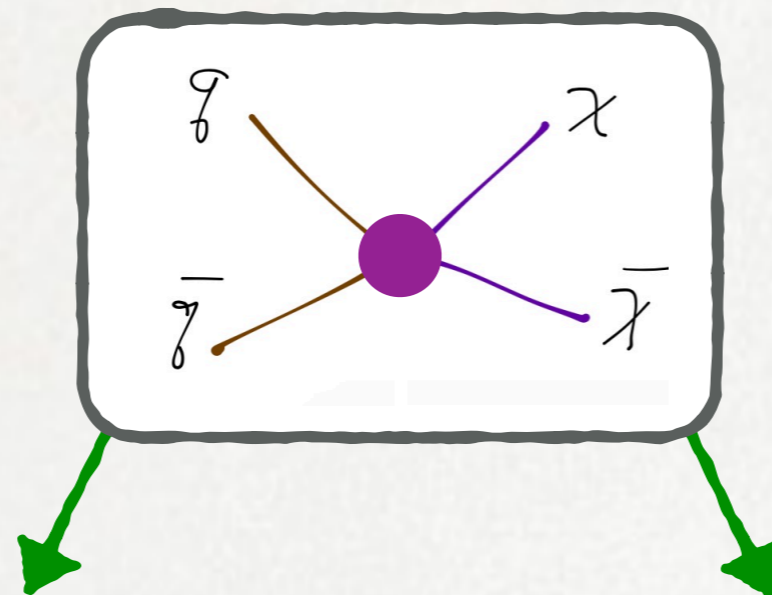


t-channel

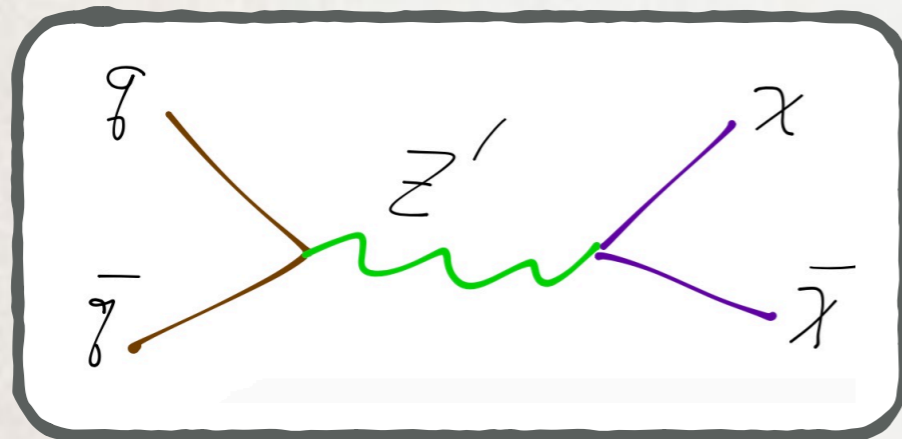


# PORTALS

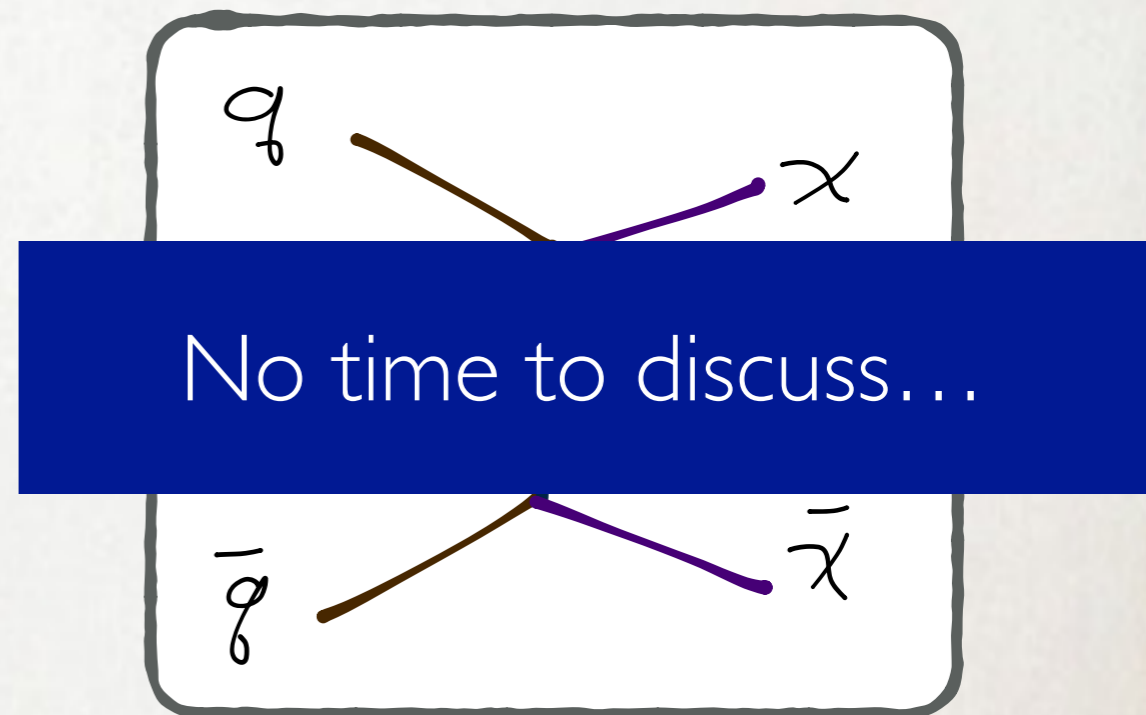
Contact operator



s-channel



t-channel



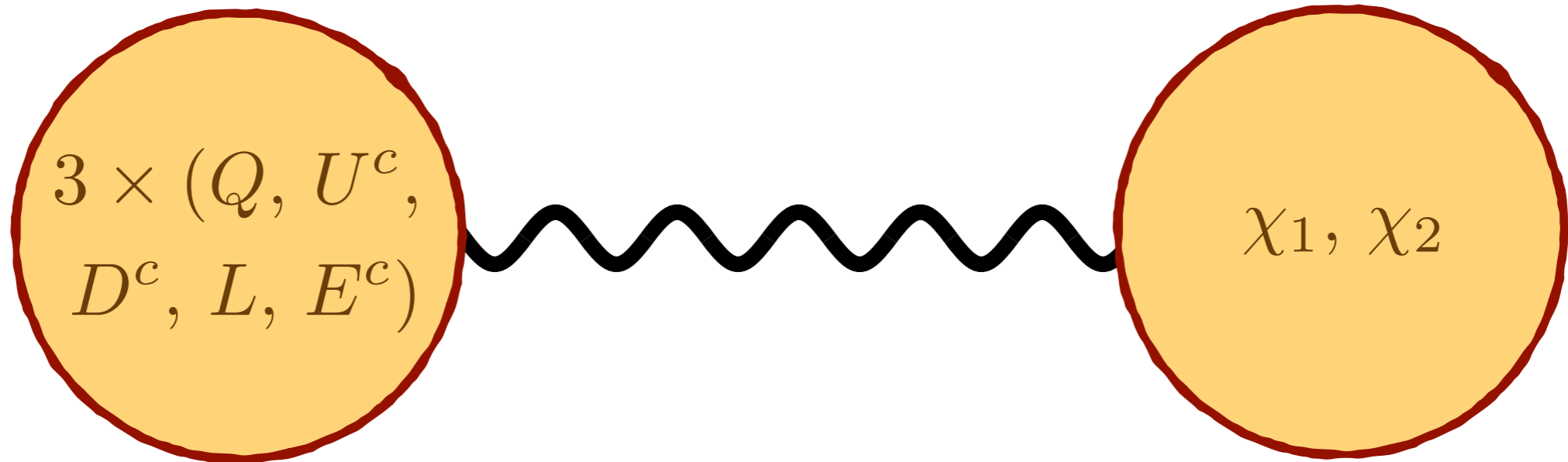
# DARK SECTOR

Dark force:  $SU(2)_d \implies \alpha_d$  or equivalently  $\Lambda_d$ .

Dark quarks:  $\chi_1, \chi_2$ .

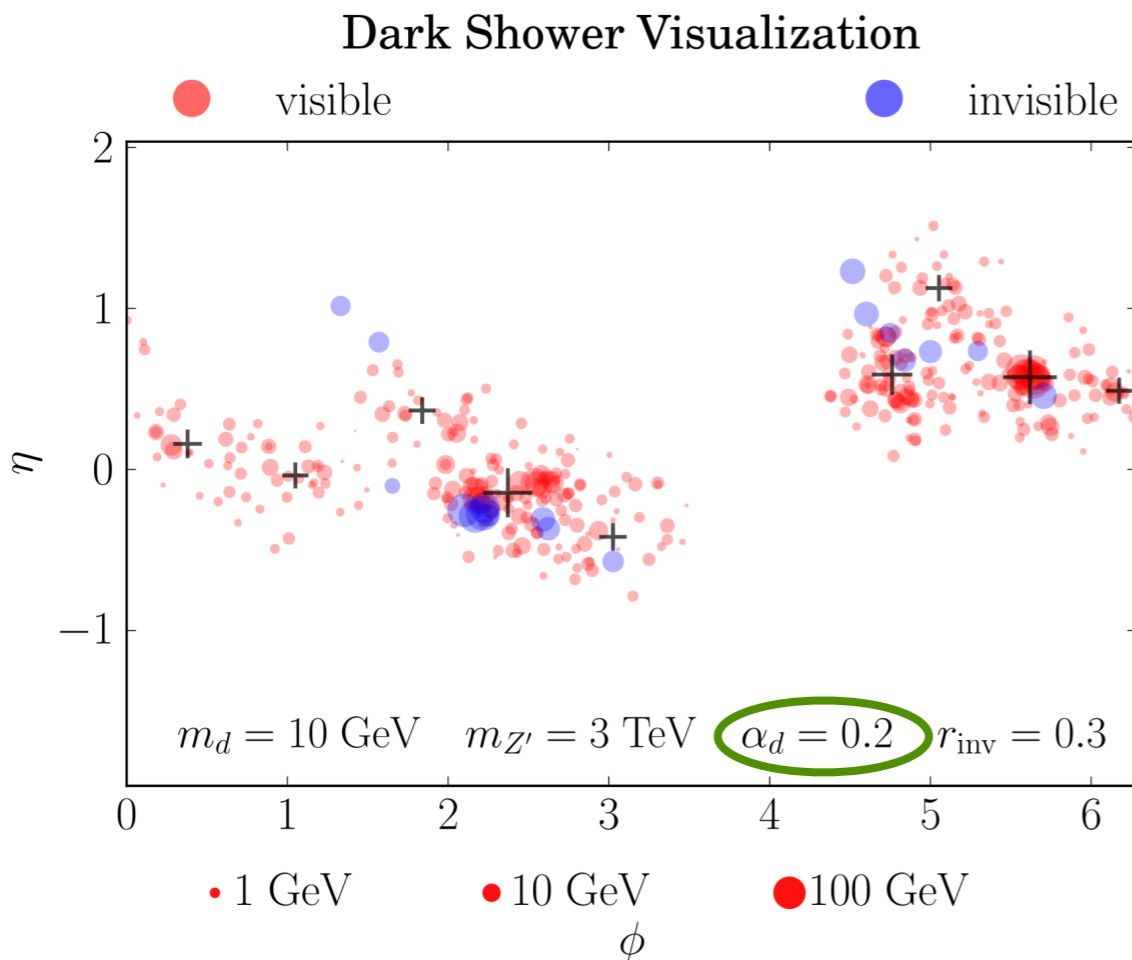
$SU(3)_C \times SU(2)_W \times U(1)_Y$

$SU(2)_d$

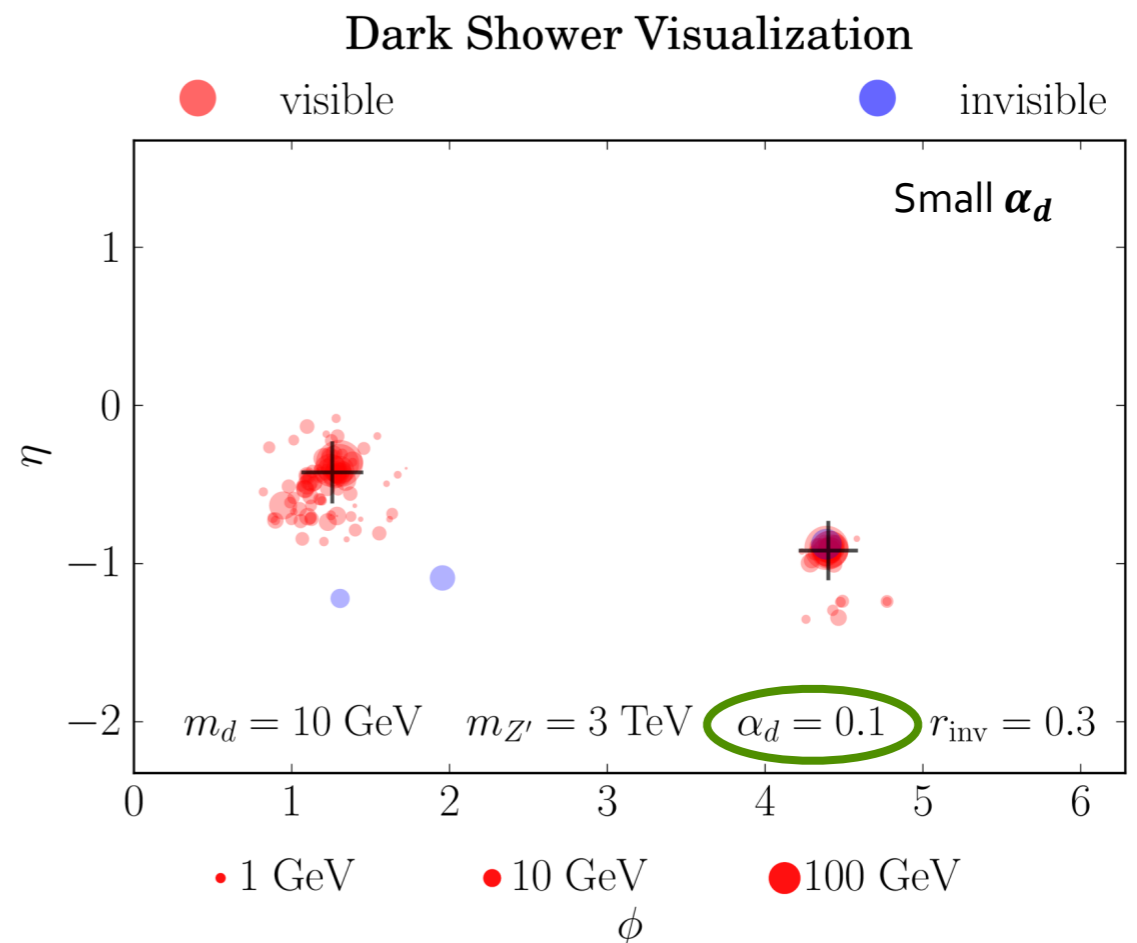


$$\Lambda_D \ll \Lambda_{\text{contact}}$$

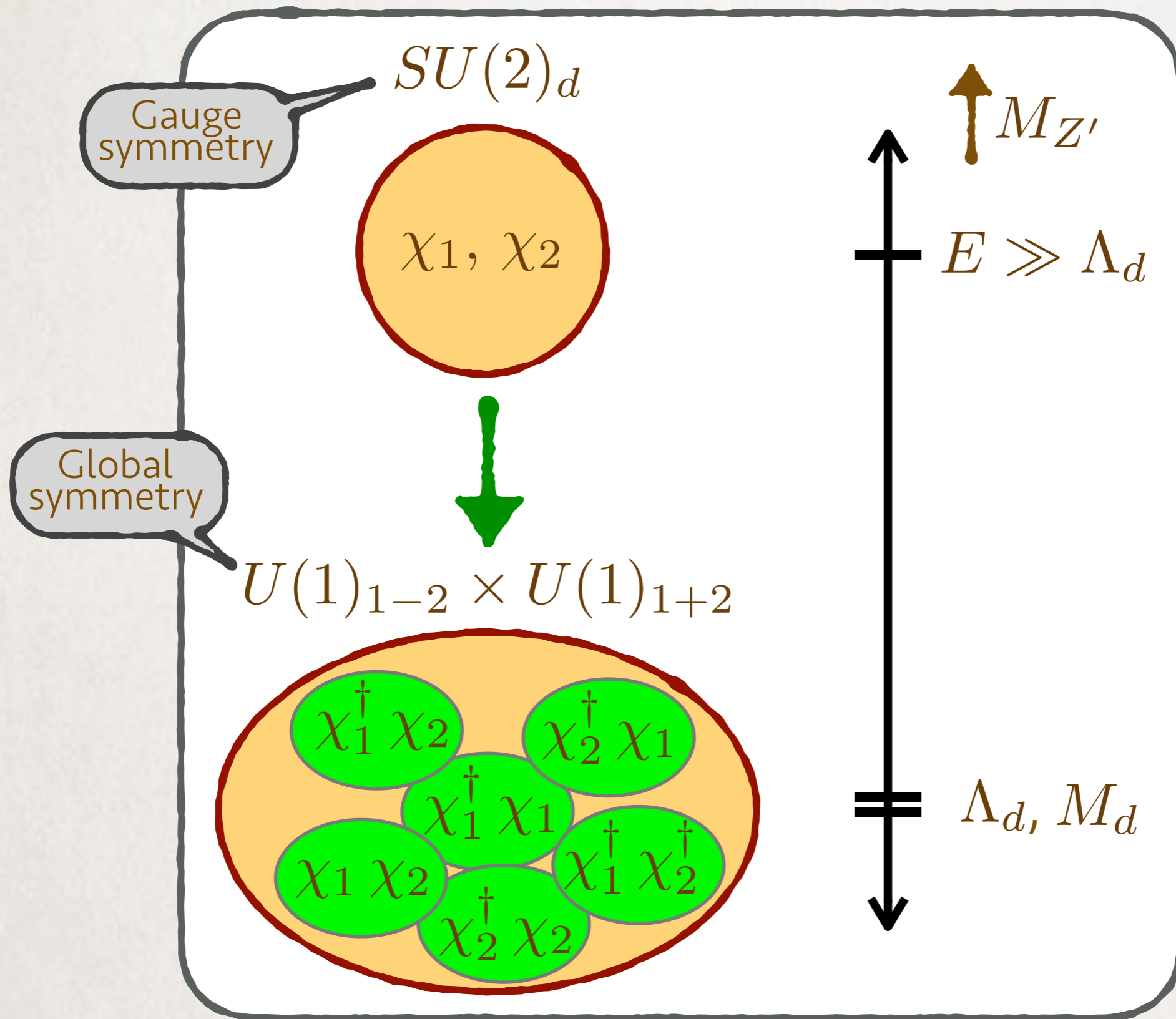
# DARK SHOWER STRENGTH



Dark gauge coupling,  $\alpha_d$ ,  
unknown parameter.



# DARK HADRONS



# DARK HADRONS

Two global symmetries:

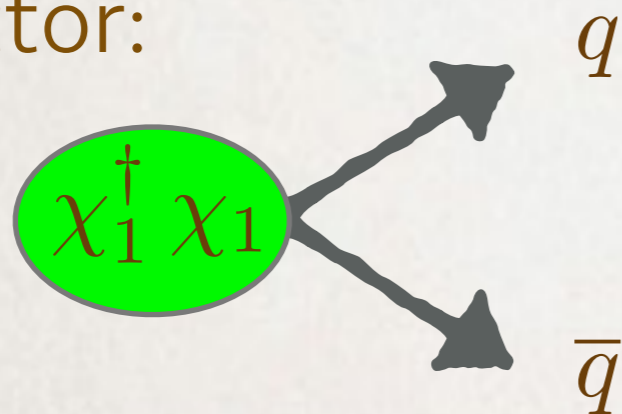
$$U(1)_{1-2} \quad \text{and} \quad U(1)_{1+2}$$

Dark isospin

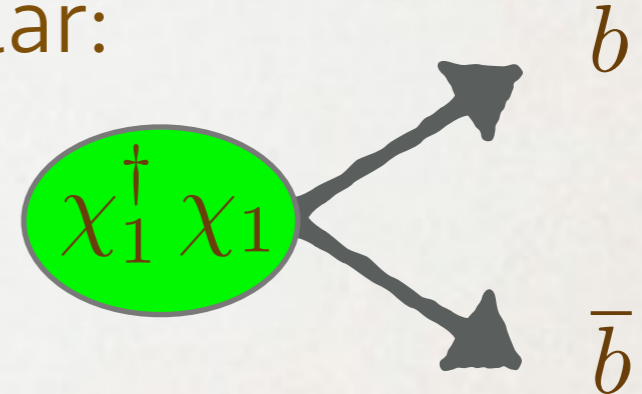
Dark baryon number

	Hadrons	$U(1)_{1-2}$	$U(1)_{1+2}$
unstable meson	$\chi_1^\dagger \chi_1$ $\chi_2^\dagger \chi_2$	0	0
stable meson	$\chi_1 \chi_2^\dagger$ $\chi_1^\dagger \chi_2$	+2, -2	0
baryon	$\chi_1 \chi_2$ $\chi_1^\dagger \chi_2^\dagger$	0	+2, -2

Vector:



Scalar:





# INVISIBLE BR PARAMETER

$$r_{\text{inv}} \equiv \left\langle \frac{\# \text{ stable}}{\text{total}} \right\rangle$$

$$\text{Fragmentation rate: } T \sim \exp \left( - \frac{4 \pi |M_2^2 - M_1^2|}{\Lambda_d^2} \right)$$

Andersson, Gustafson, Ingelman, Sjostrand [Phys. Rept. 97 (1983)]

$$2 \text{ flavors with } M_1^2 = M_2^2 \longrightarrow r_{\text{inv}} \simeq 0.5$$

$$2 \text{ flavors with } M_1^2 \ll M_2^2 \longrightarrow r_{\text{inv}} \rightarrow 0$$

(ignoring baryons)

# INVISIBLE BR PARAMETER

$$r_{\text{inv}} \equiv \left\langle \frac{\# \text{ stable}}{\text{total}} \right\rangle$$

***Naively counting:***

$N_F$  singlet combinations  $\chi_i^\dagger \chi_i$ .

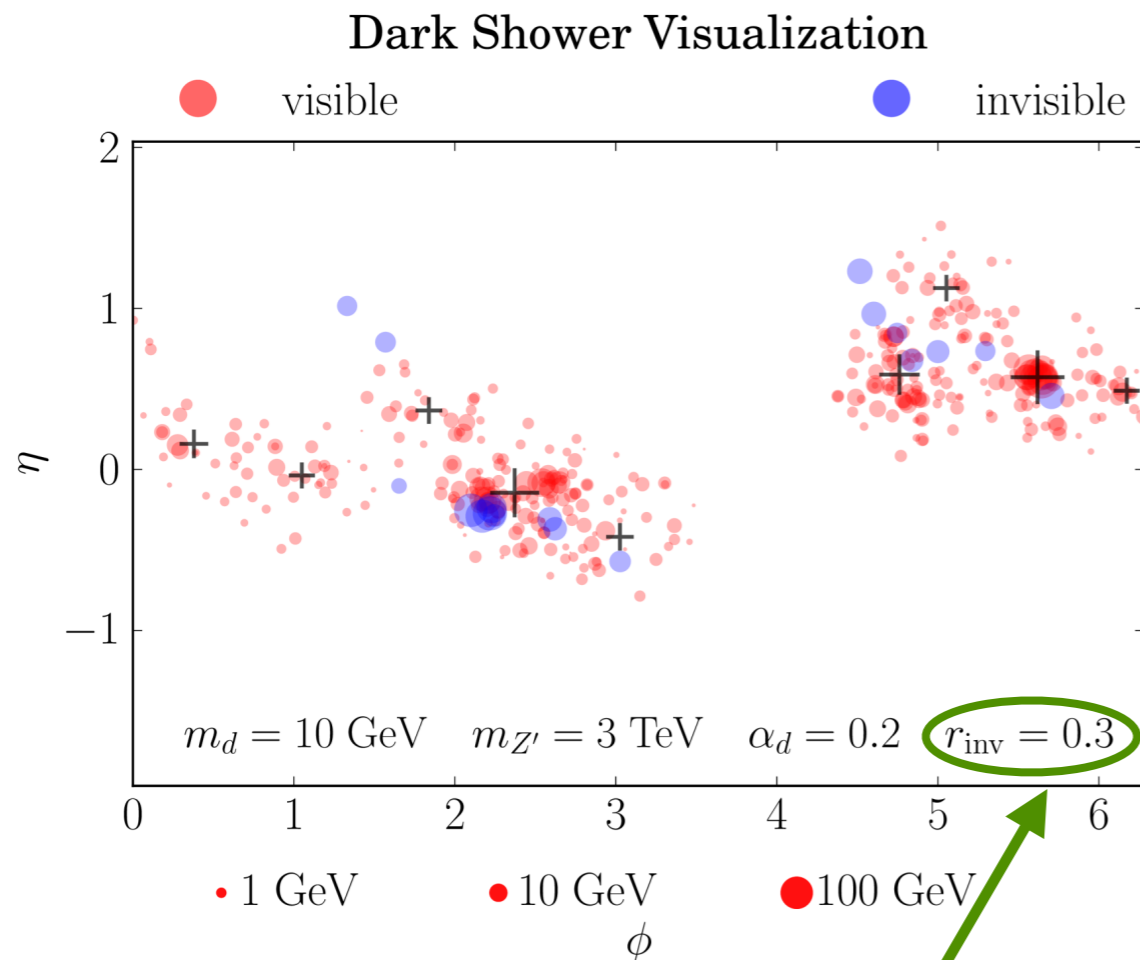
$\frac{N_F(N_F - 1)}{2}$  "charged" combinations  $\chi_i^\dagger \chi_j$  with  $i \neq j$ .

Number of (possibly) stable particles grows as  $N_F \rightarrow \infty$

$$\Rightarrow r_{\text{inv}} \rightarrow 1$$

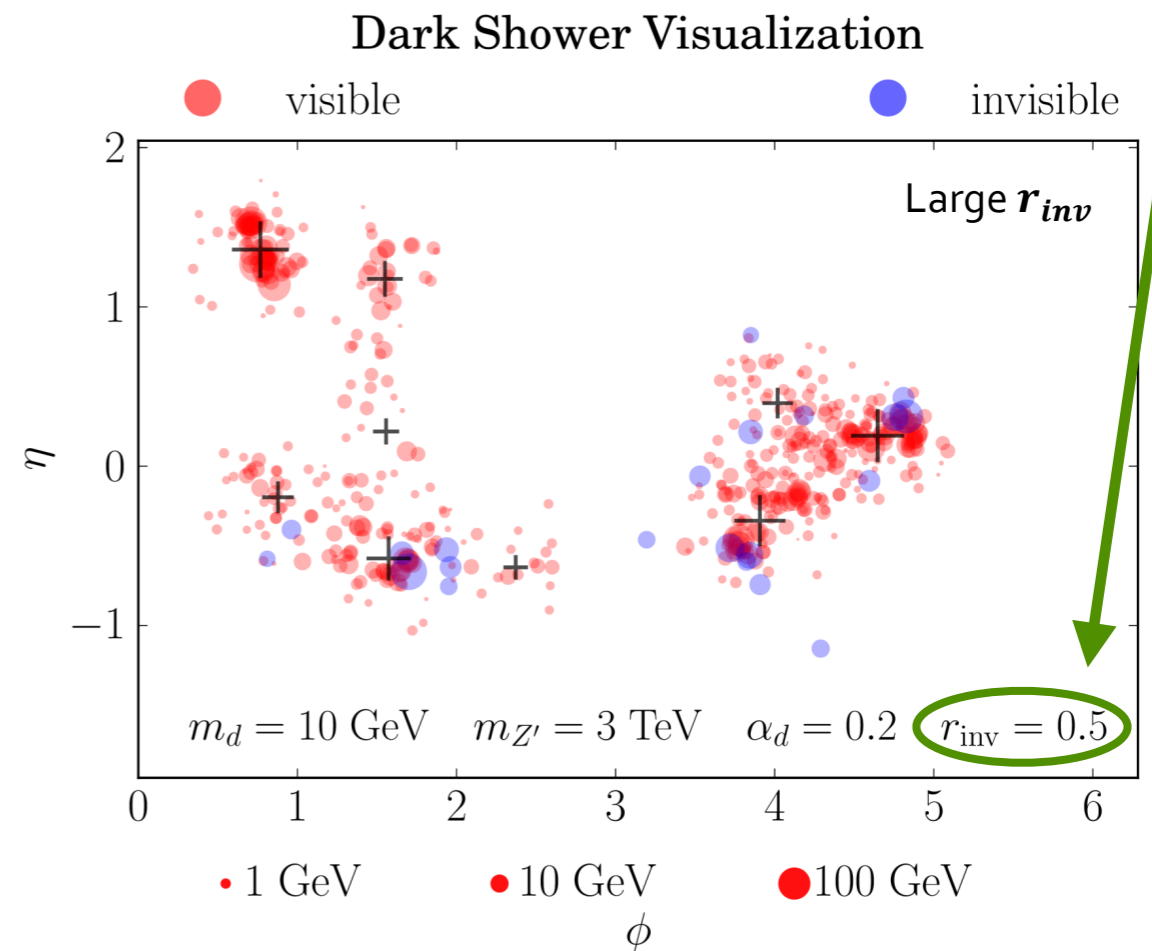
$$0 < r_{\text{inv}} < 1$$

# LEGO PLOTS

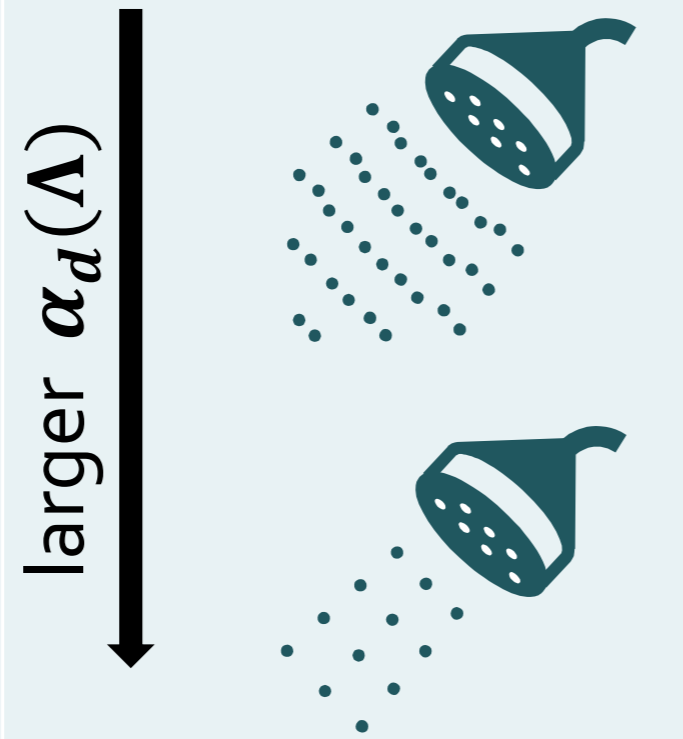
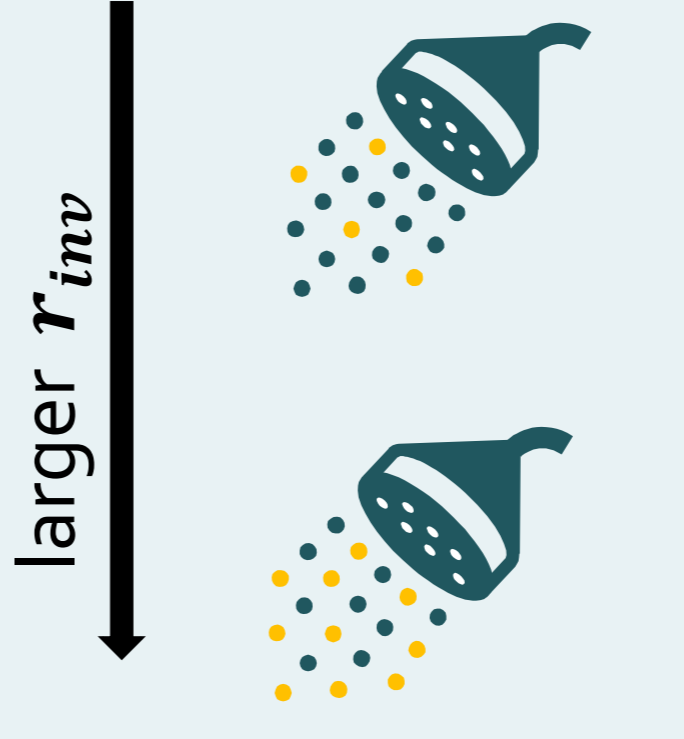


$r_{inv} = 0.3$

$r_{inv} = 0.5$



# PARAMETER SUMMARY

Shower Strength	Invisible ratio
$\alpha_d(1\text{TeV})$	$r_{inv}$
<p>larger <math>\alpha_d(\Lambda)</math></p> 	<p>larger <math>r_{inv}</math></p> 

Dark meson mass,  $m_d$ , and production rate:  $\sigma \times \text{BR}$ .

# SEARCHES

# CONTACT OPERATOR SEARCH

Follow “mono- $X$ ” strategy.

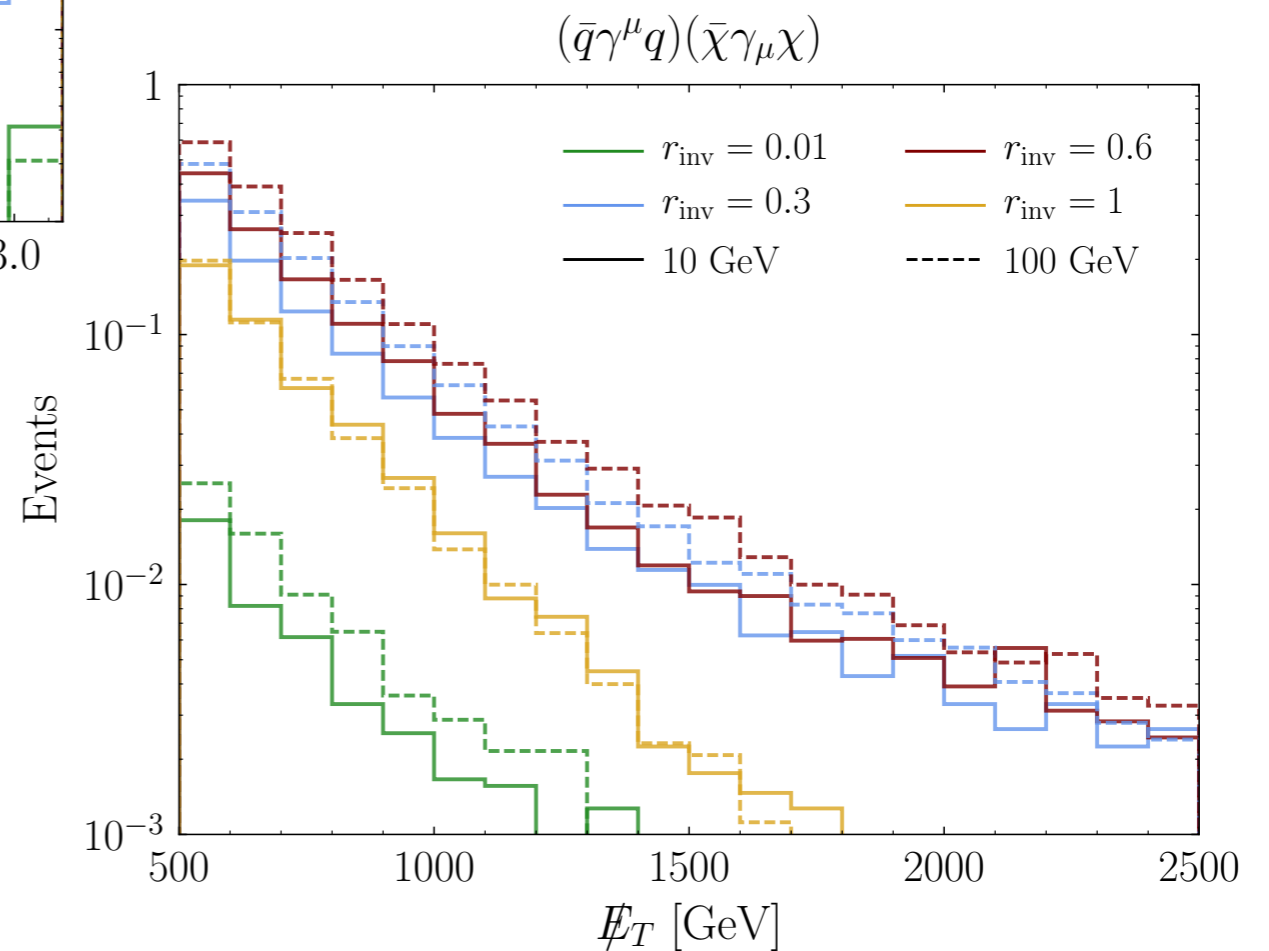
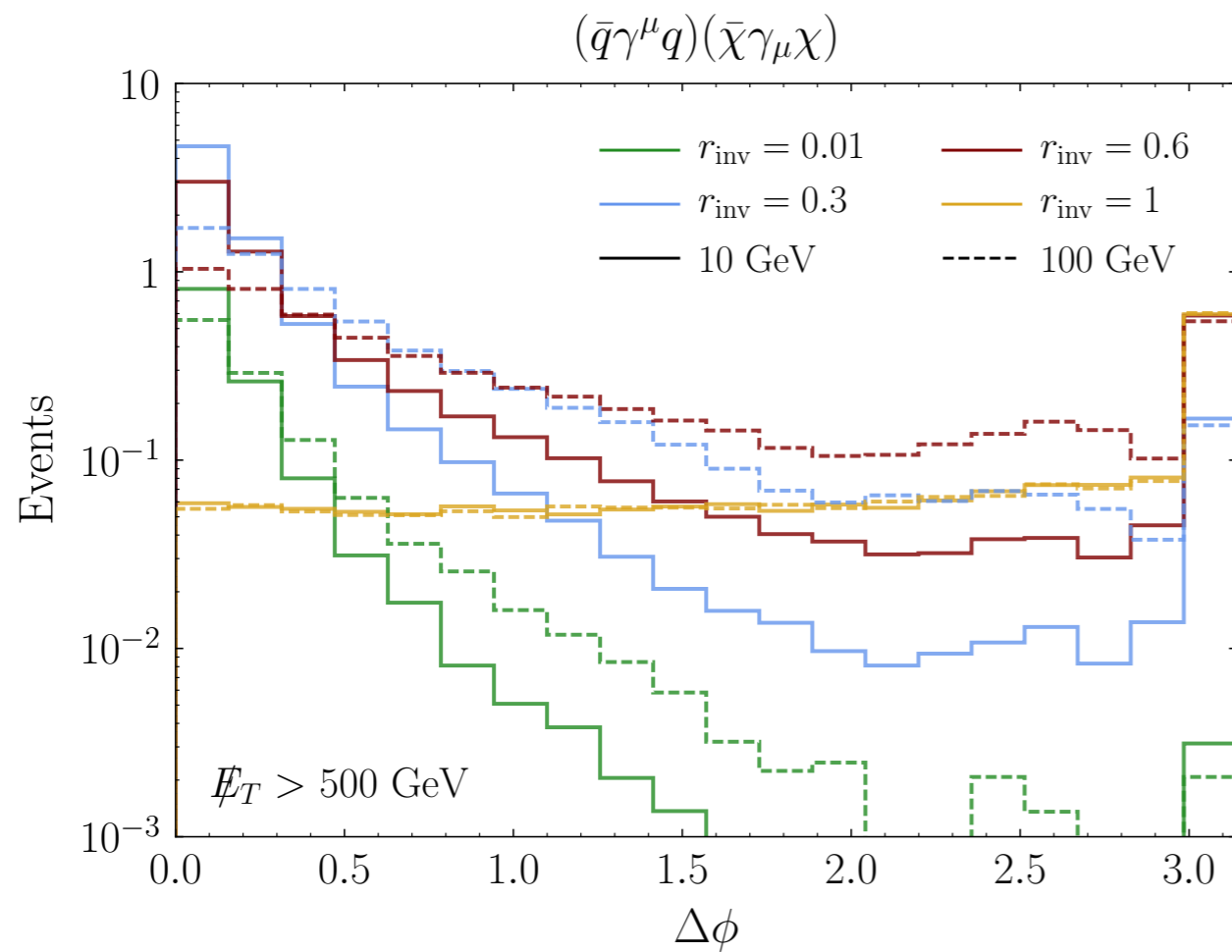
Focus on the vector-vector contact operator:

$$\frac{1}{\Lambda^2} (\bar{q}\gamma^\mu q) (\bar{\chi}\gamma_\mu \chi)$$

Then UV complete into model with  $s$ -channel exchange.

(For  $t$ -channel model, see paper.)

# DISTRIBUTIONS



# CUTS

## Trigger:

- At least one jet with  $|\eta| < 2.5$  and  $p_T > 250$  GeV.
- Veto leptons.

## Signal Regions:

- Require  $\cancel{E}_T > 800, 900, 1000$  GeV.
- Require either  $\Delta\phi > 0.4$  or  $\Delta\phi < 0.4$ .



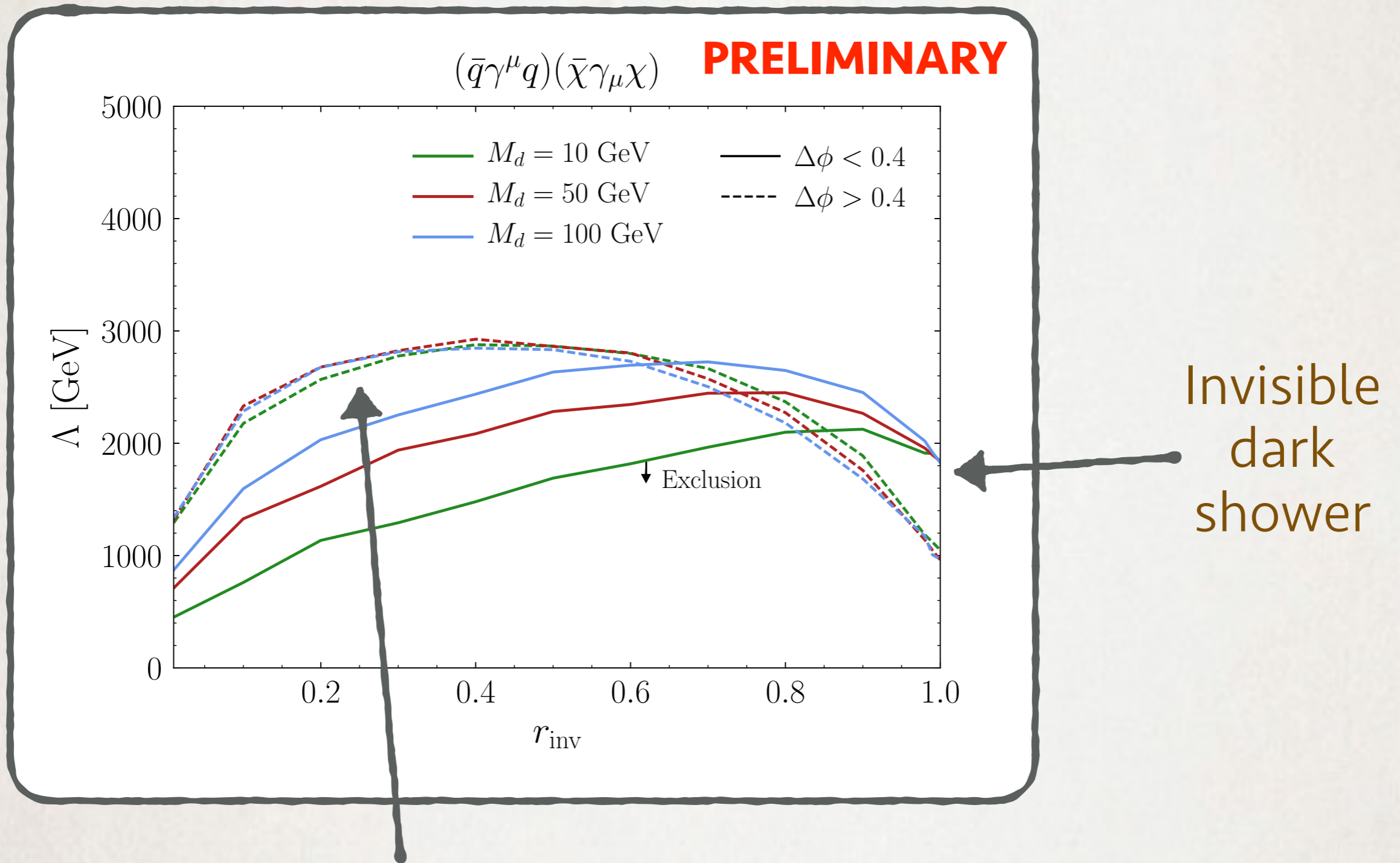
Standard cut.



"Semi-visible" cut.



# PROJECTED REACH



Semi-visible jet  
search dominates

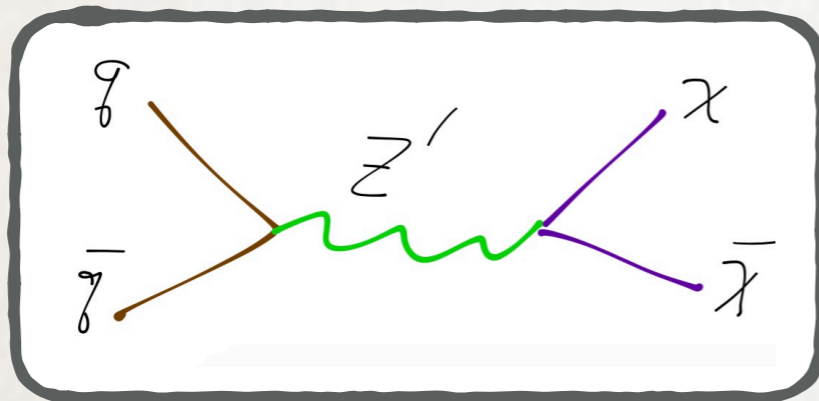
# S-CHANNEL UV COMPLETION: HOW DOES SENSITIVITY CHANGE?

# Z' SEARCH

Couple a  $Z'$  to quarks.

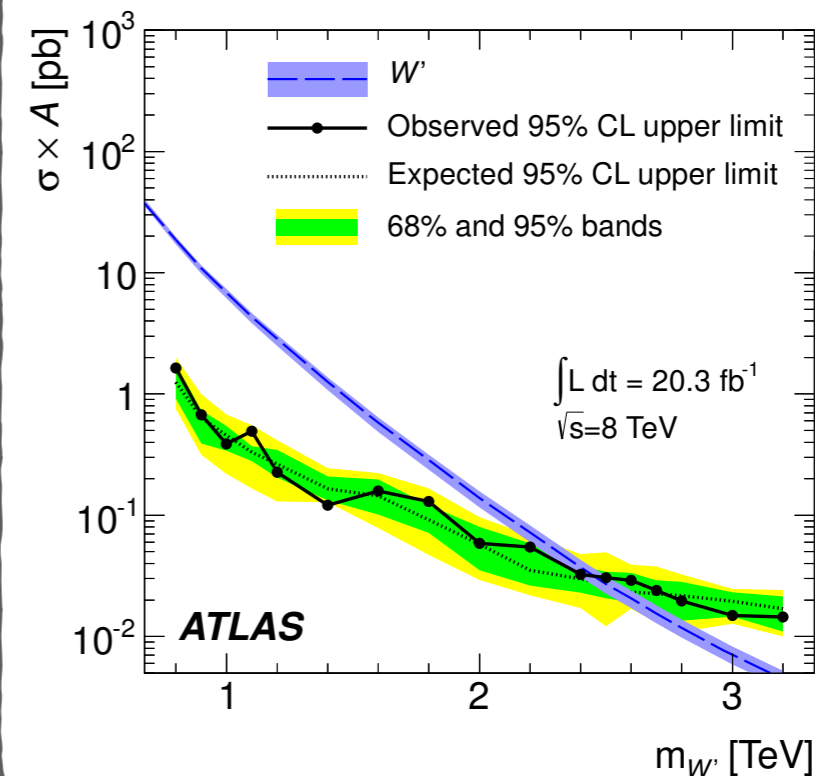
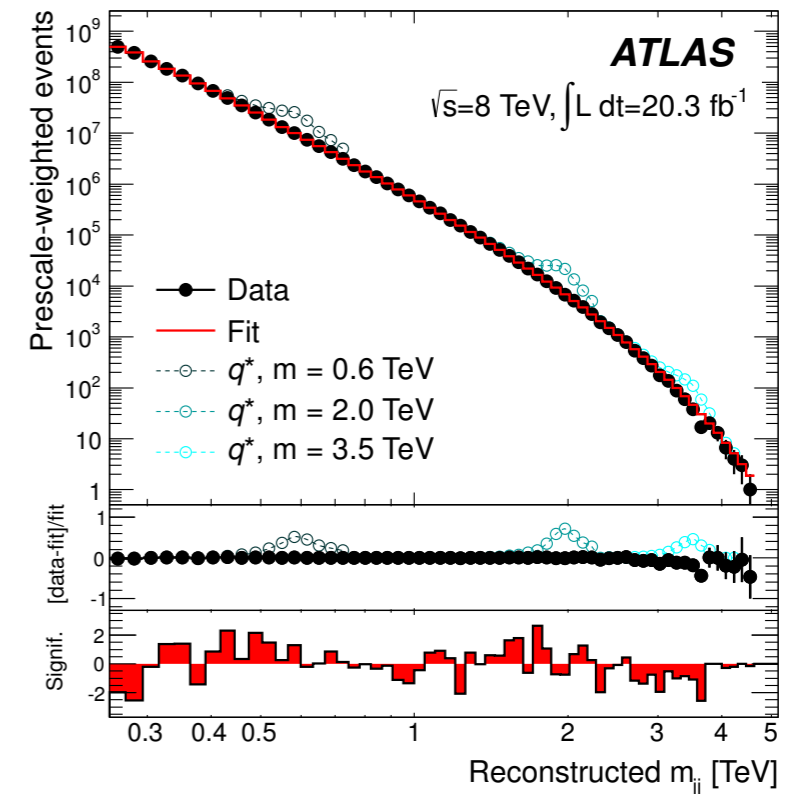
Produced resonantly at LHC.

$$\sigma \times \text{BR}$$

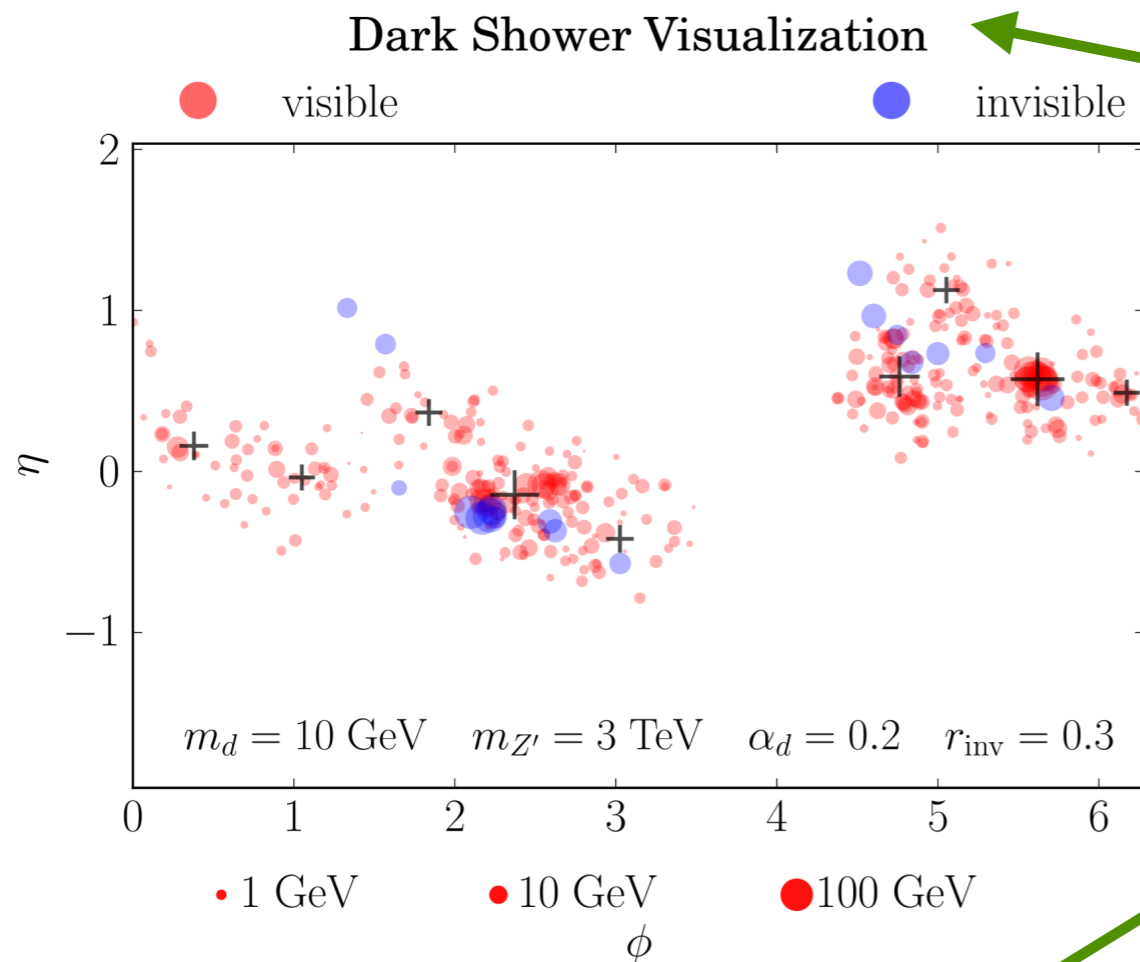


At partonic center of mass energy

$$\sqrt{\hat{s}} \simeq M_{Z'}$$

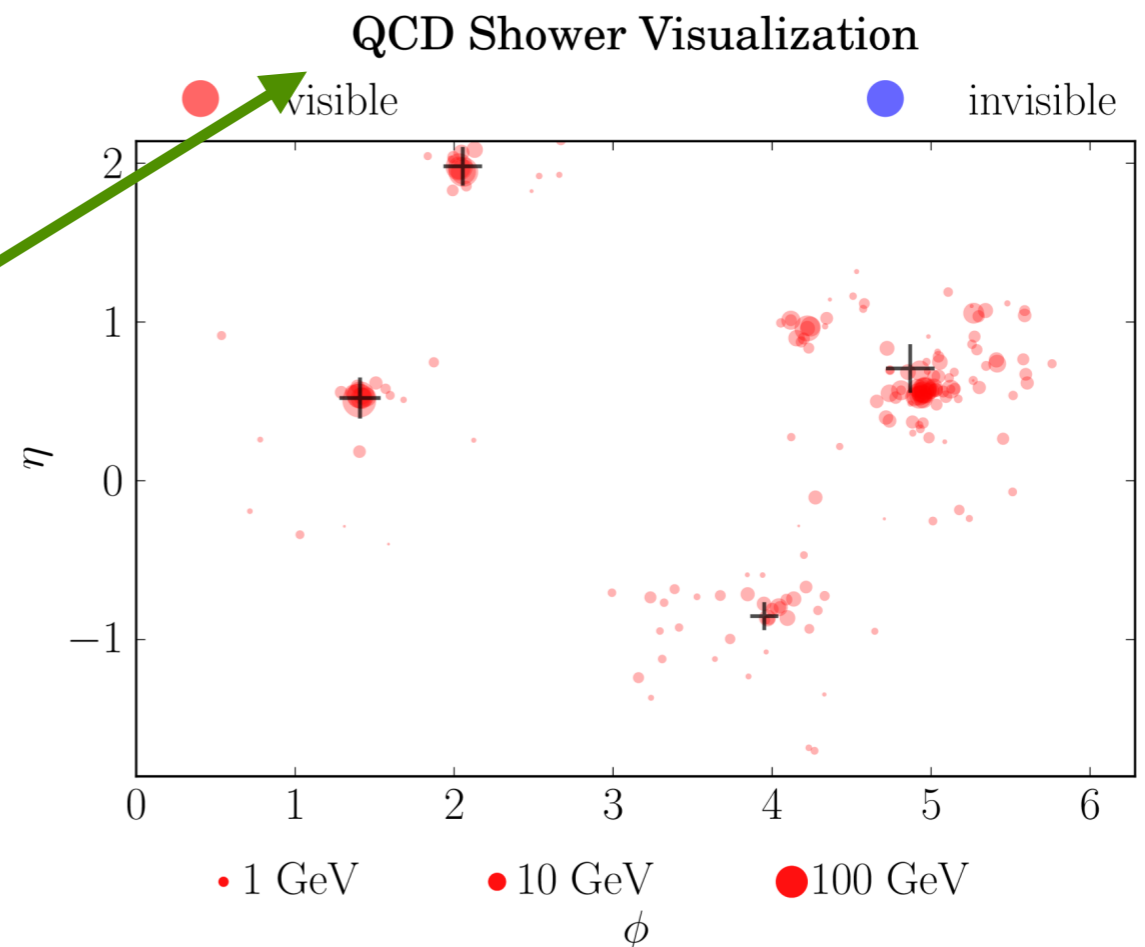


# SIGNAL VERSUS BACKGROUND



Dark shower

QCD shower



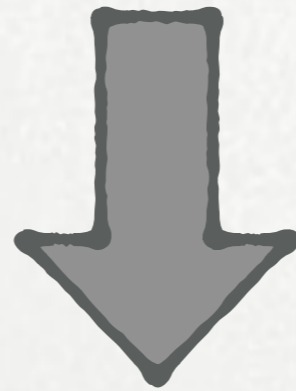
# BUMP HUNT

Looking for a resonance.

Di-jet like.

Jets are fatter than QCD jets.

Non-zero  $\cancel{E}_T$  (also washes out resonance).



Try the transverse mass:

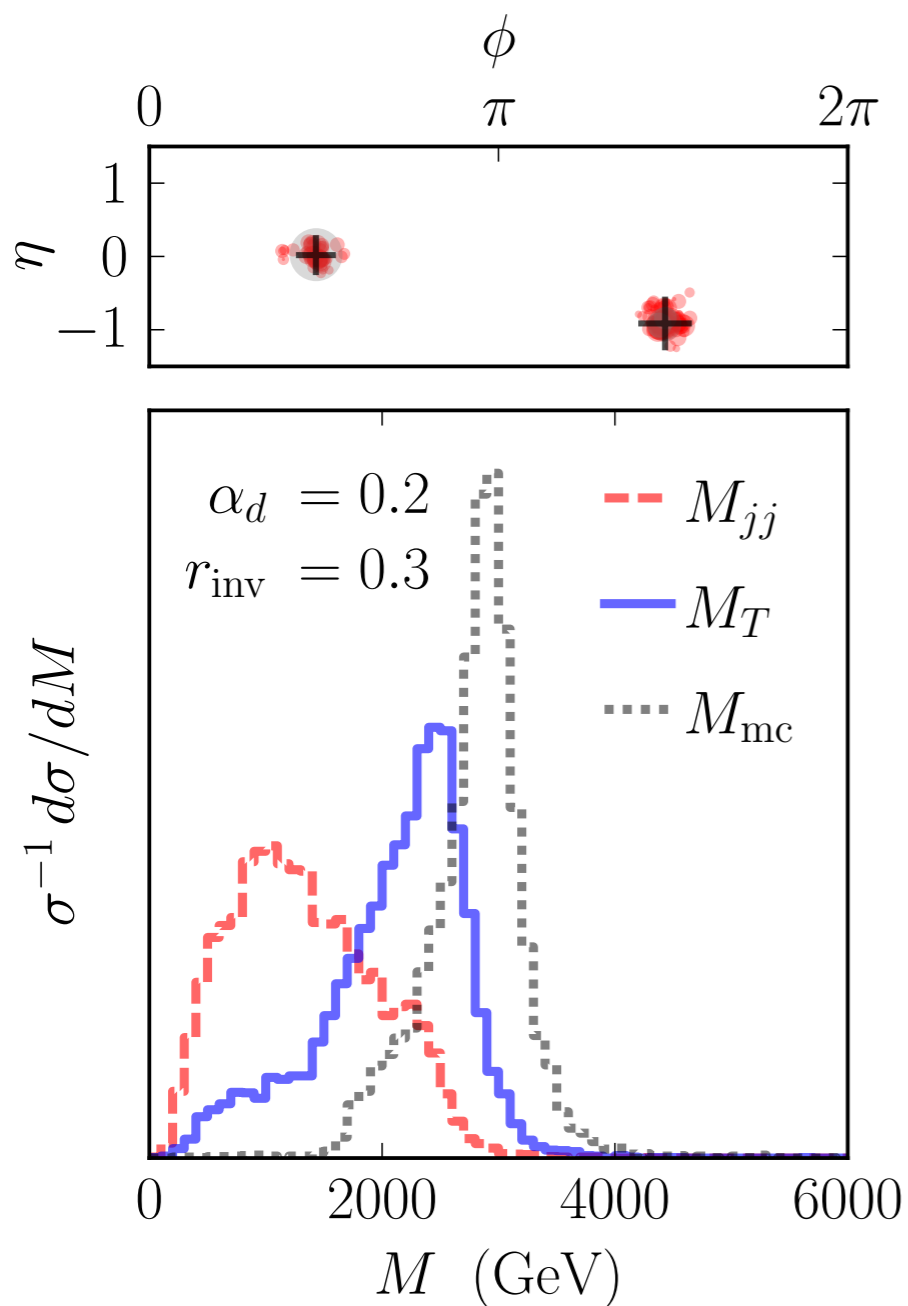
$$M_T^2 = M_{jj}^2 + 2 \left( \sqrt{M_{jj}^2 + p_{Tjj}^2} \cancel{E}_T - \vec{p}_{Tjj} \cdot \vec{\cancel{E}}_T \right)$$

# BUMP HUNT

Cluster into  $R = 0.5$  anti- $k_T$  jets.

Re-cluster into  $R = 1.1$  C/A jets.

Compute  $M_T$  using the fat jets.



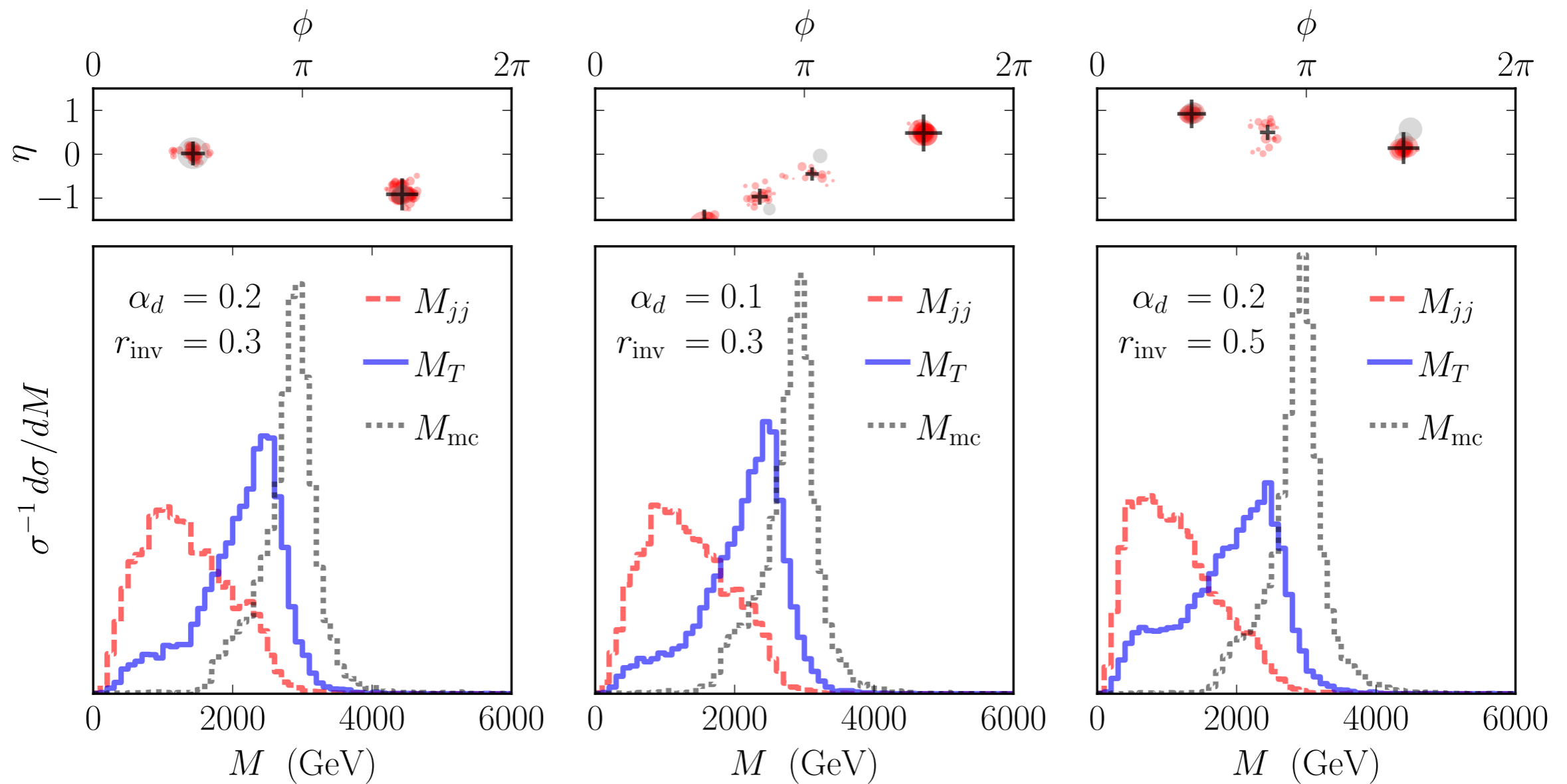
$$M_{jj} \leq M_T \leq M_{Z'}$$

Di-fat jet  
invariant mass

True invariant  
mass. Mocked up  
by  $M_{mc}$ : computed  
using visible and  
invisible 4-vectors.

$$M_T^2 = M_{jj}^2 + 2 \left( \sqrt{M_{jj}^2 + p_{Tjj}^2} E_T - \vec{p}_{Tjj} \cdot \vec{E}_T \right)$$

# VARY PARAMETERS



# CUTS

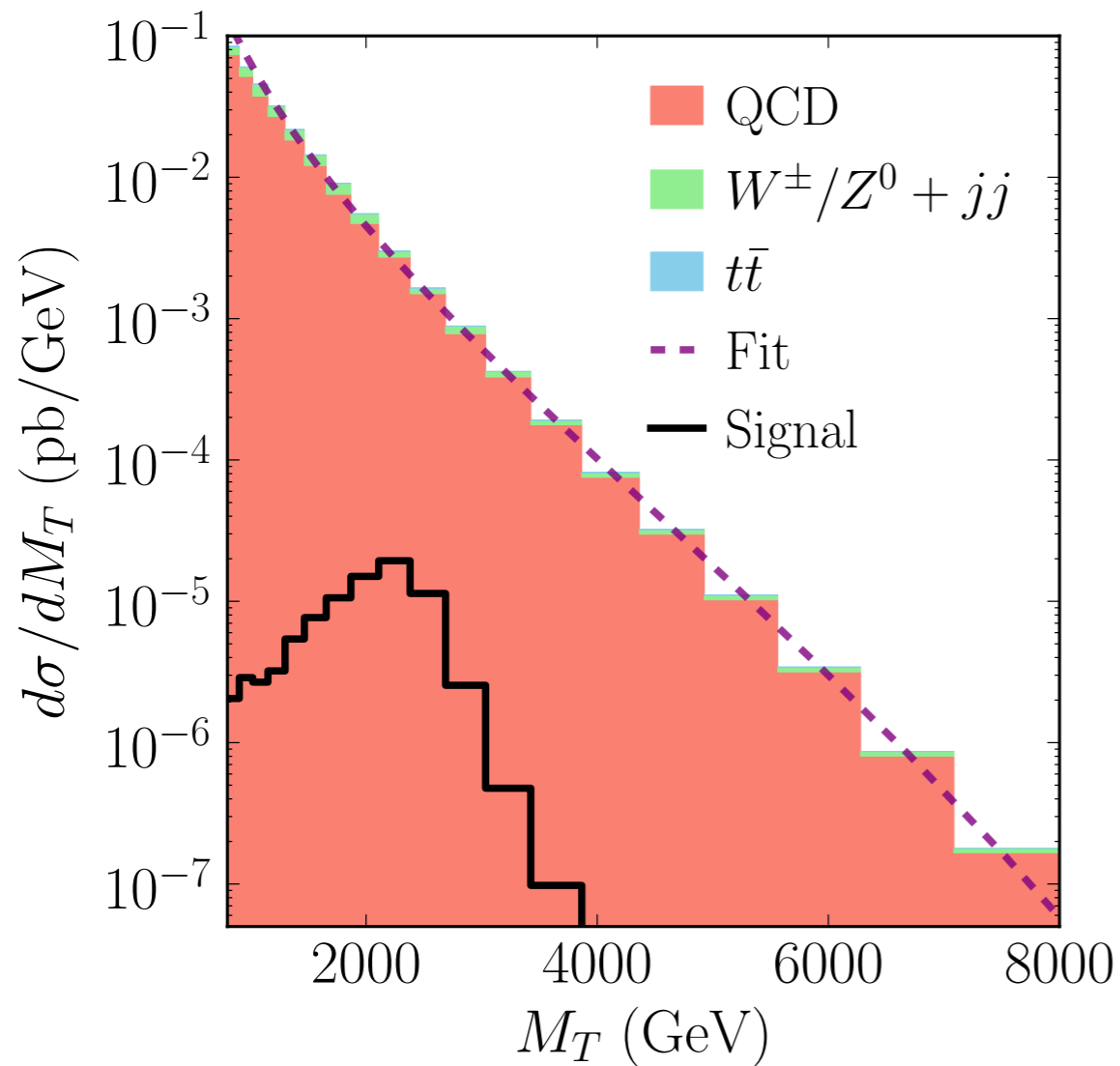
- Reclusters jet into  $R = 1.1$  C/A jets ( $j_1, j_2$ ).
- Require  $|\eta_{j_1} - \eta_{j_2}| < 1.1$ .
- Require  $\Delta\phi < 1$ .
- Veto isolated  $e^\pm/\mu^\pm$  with  $p_T > 20$  GeV and  $|\eta| < 2.4$ .
- Require  $\cancel{E}_T/M_T > 0.15$ .

Then bump hunt in  $M_T$ .

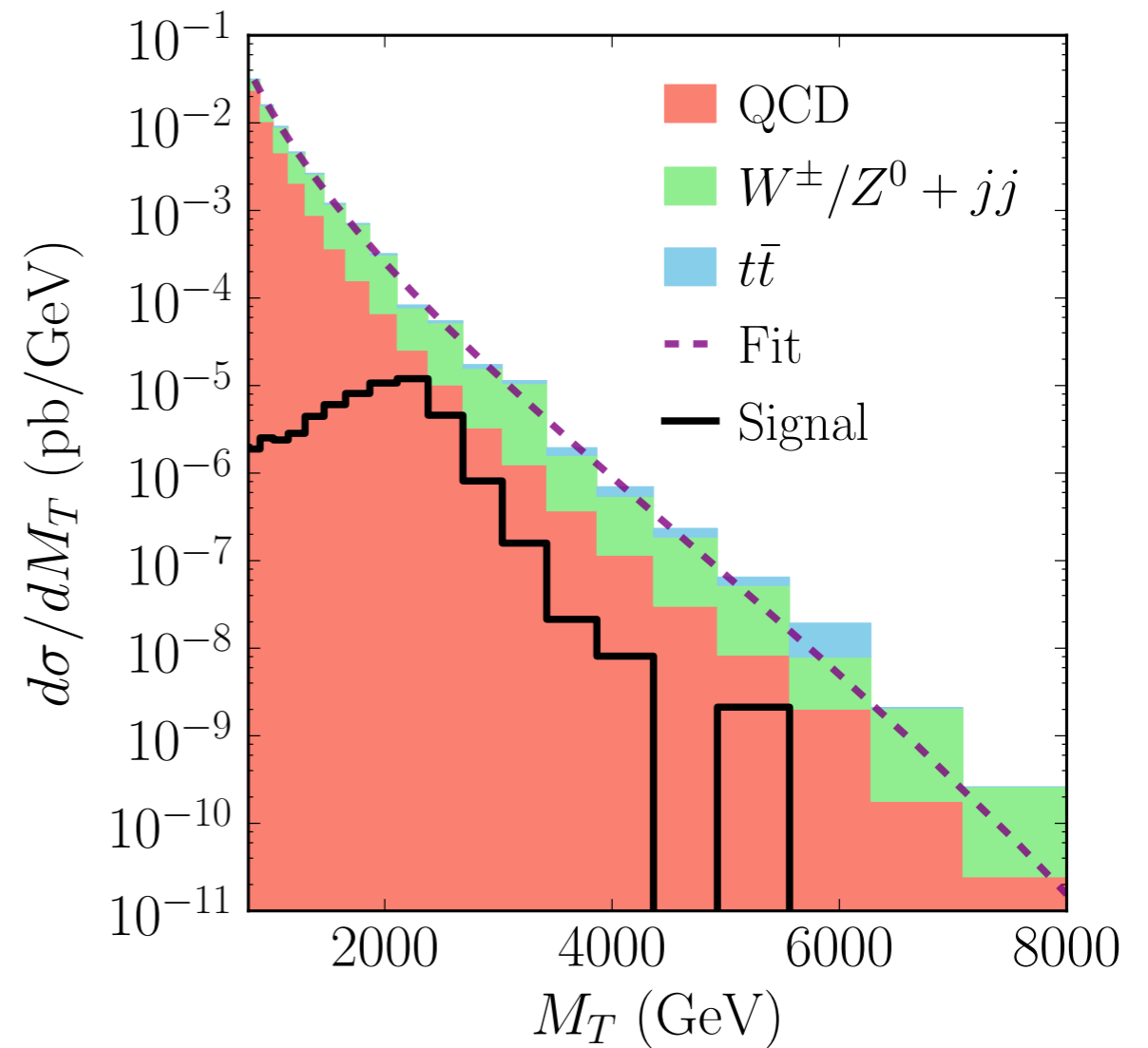


# DISTRIBUTIONS

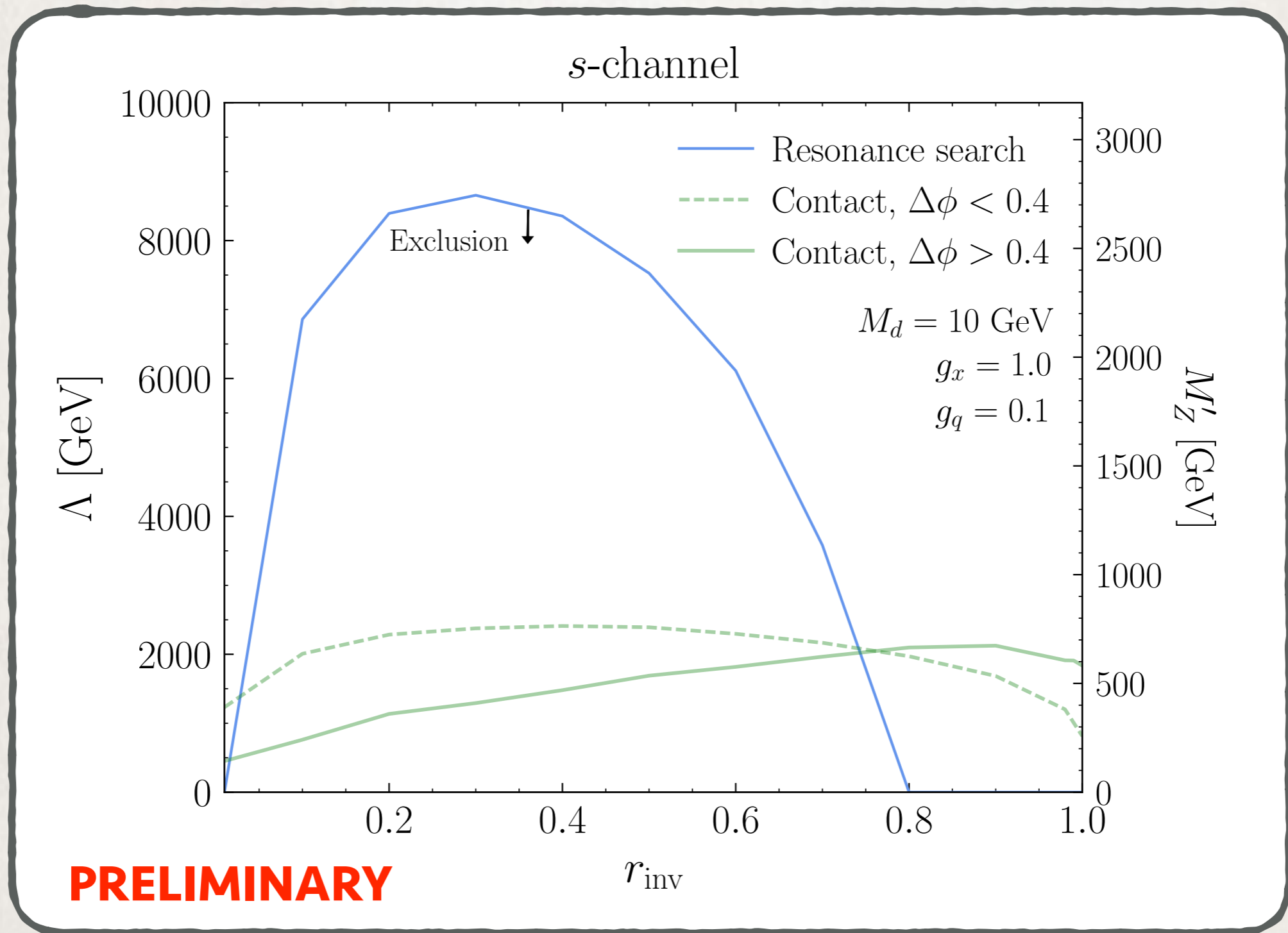
After pre-selection



With  $\cancel{E}_T/M_T > 0.15$



# PROJECTED REACH



# OUTLOOK

# DIRECT DETECTION

## Weakly coupled DM

Direct detection fully correlated with LHC signatures.

$$\begin{array}{ccc} \text{LHC} & \longrightarrow & \text{DD} \\ \frac{1}{\Lambda^2} (\bar{q}\gamma^\mu q) (\bar{\chi}\gamma_\mu \chi) & & \frac{1}{\Lambda^2} (\bar{q}\gamma^\mu q) (\bar{\chi}\gamma_\mu \chi) \end{array}$$

## Composite DM

Direct detection depends on UV completion.

$$\begin{array}{ccc} \text{LHC} & \longrightarrow & \text{DD} \\ \frac{1}{\Lambda^2} (\bar{q}\gamma^\mu q) (\bar{\chi}\gamma_\mu \chi) & & \frac{1}{\Lambda_{\text{DD}}^2} (\bar{q}\gamma^\mu q) (\bar{\pi}_d \gamma_\mu \pi_d) \end{array}$$

$\chi$  are dark quarks.

$\pi_d$  are dark mesons.

Implication: s-channel model has vanishing DD (to leading order)!

# RELIC DENSITY

## ***Symmetric abundance***

Large annihilation cross section yields small relic density.

Some spectra could use “forbidden channel” mechanism.

Correlation between relic density and  $r_{\text{inv}}$ .

Are such spectra possible?

## ***Asymmetric abundance***

Straightforward to implement.

Relate global symmetry charge to baryon or lepton number.

Big symmetric annihilation cross section.

# CONCLUSIONS

Strongly coupled dark sector could yield semi-visible jets.

Useful parametrization of dark sector properties:

$$(\Lambda_d, M_d, r_{\text{inv}})$$

Many portal possibilities:

Contact operator limit

$s$ -channel:  $Z'$  resonance search

( $t$ -channel: will be presented in paper)

Simplified parametrization allows for optimization,  
and useful way to present limits.

Direct Detection is UV completion dependent.