

Raffaele Tito D'Agnolo

What's going on at the weak scale? - Jeju - June 2017

ROTATING BLACK HOLES AND CONFORMAL QUANTUM MECHANICS

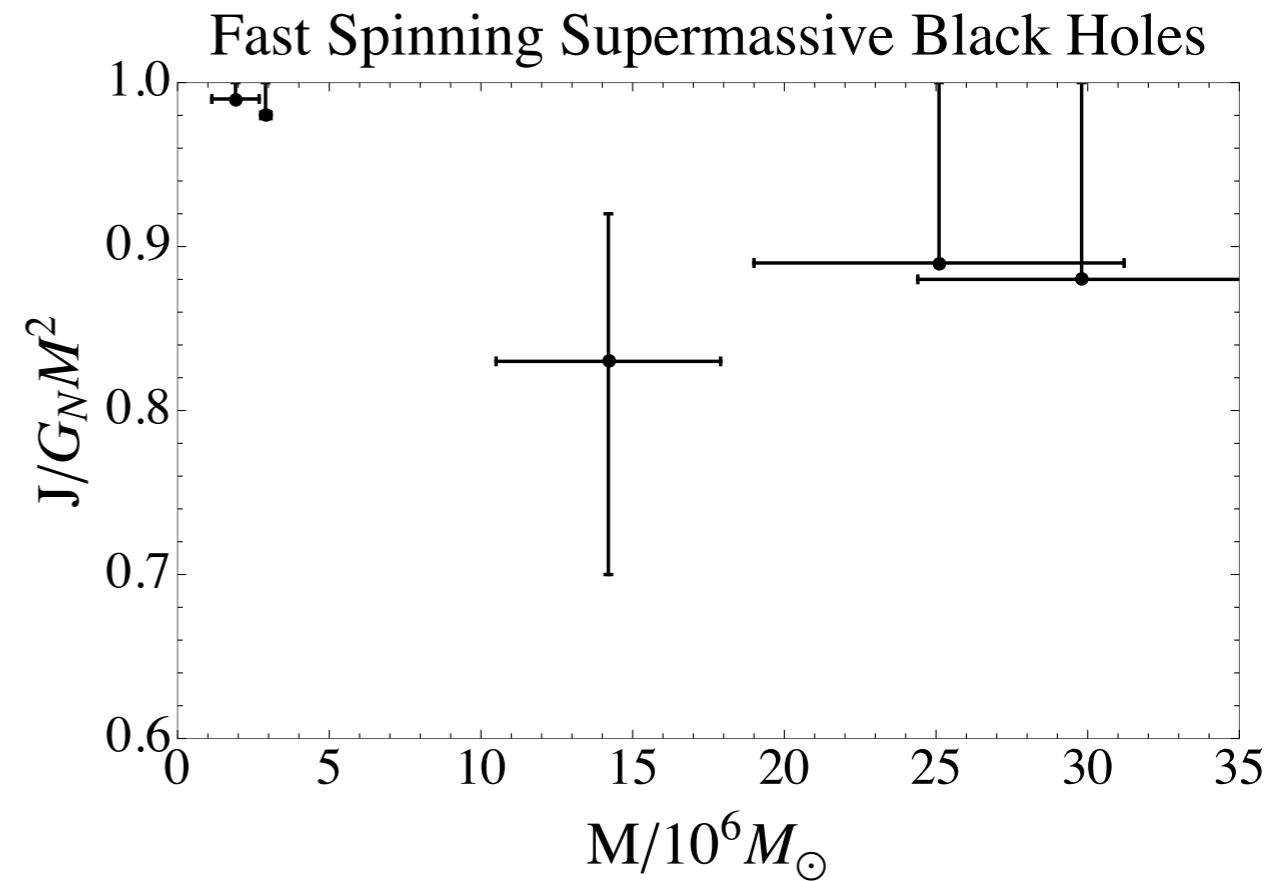
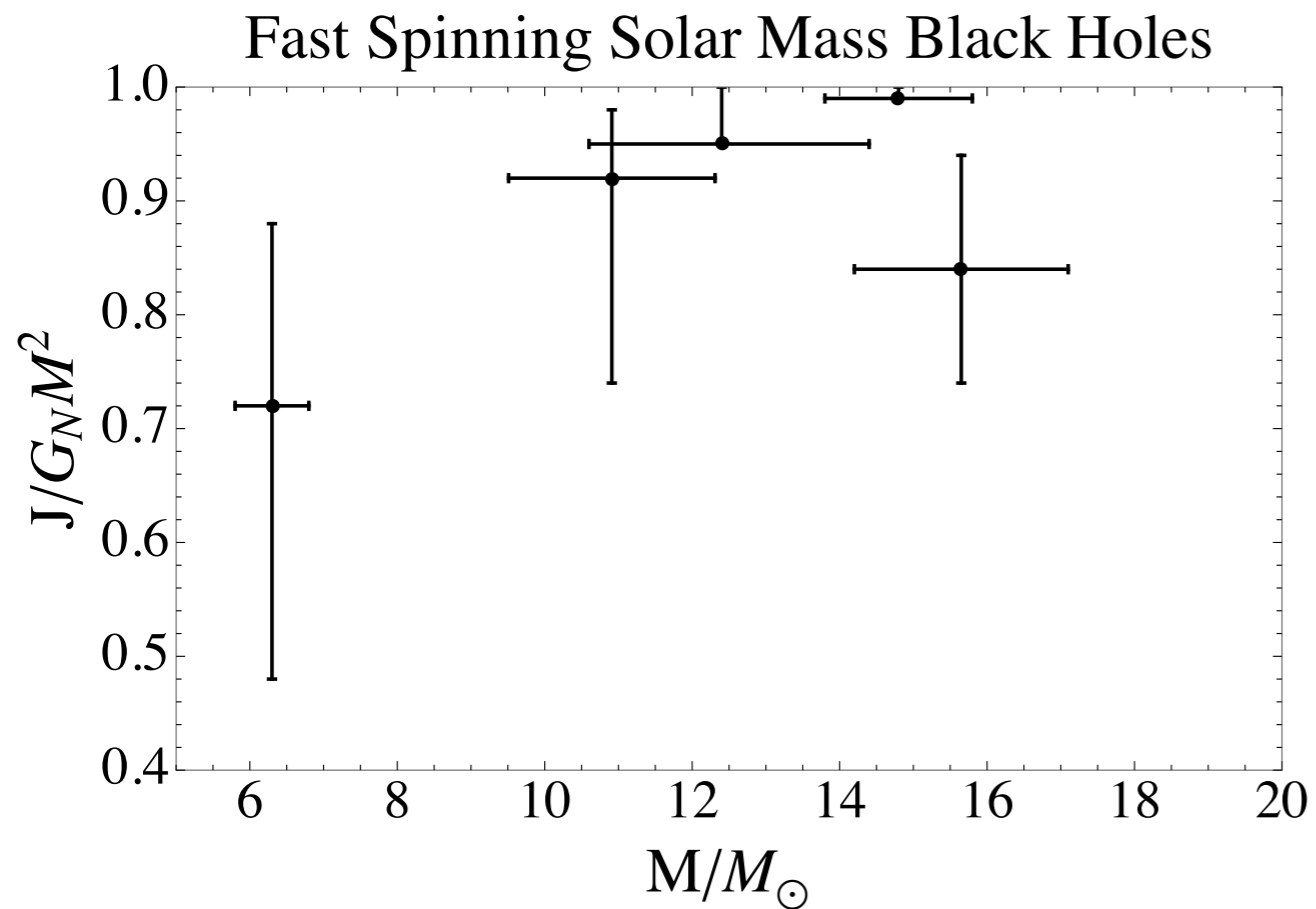


work in progress with
D. Anninos (IAS) and T. Anous (UBC)

OUTLINE

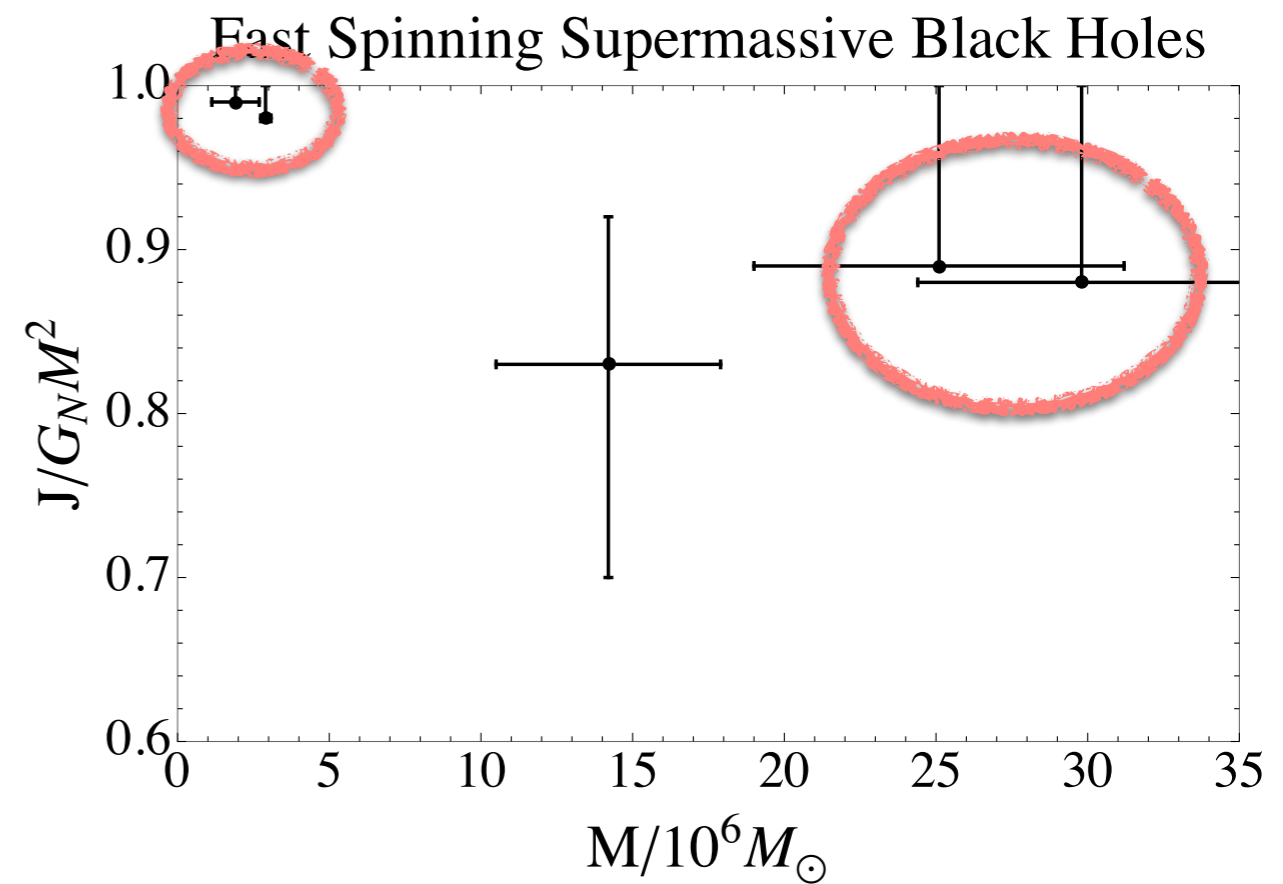
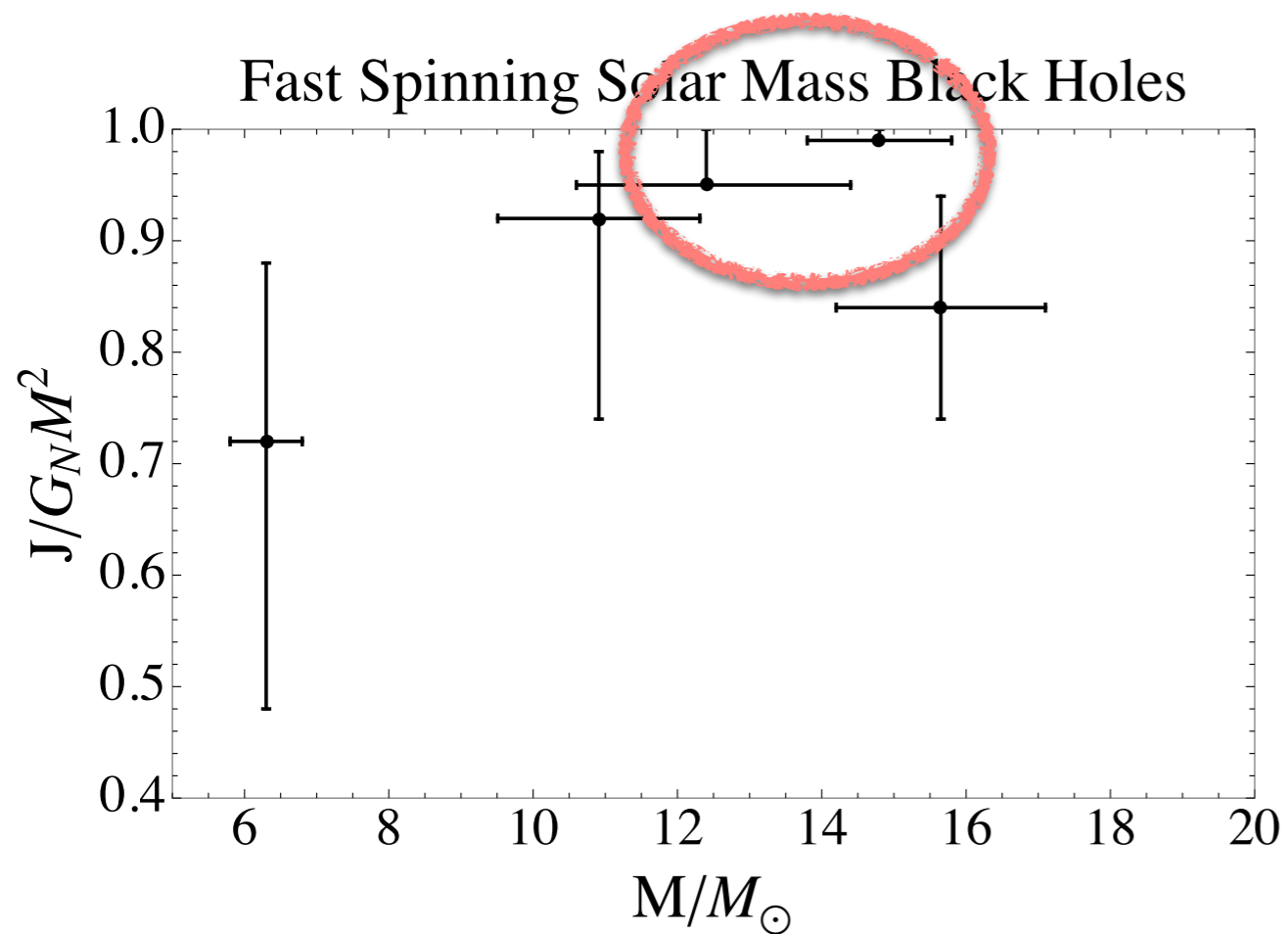
- MOTIVATION
- NEAR THE HORIZON OF EXTREMAL BLACK HOLES
- DISORDERED LARGE-N QUANTUM MECHANICS

EXTREMAL BLACK HOLES



<https://arxiv.org/pdf/1704.05081.pdf>

EXTREMAL BLACK HOLES



THE MAIN CHARACTERS

4D ASYMPTOTICALLY FLAT

M, Q, J

$$\hbar = c = G_N = 1$$

'30-'60

Birkhoff

Israel

Carter-Robin

Hawking, Wald

THE MAIN CHARACTERS

M, Q, J

	$Q = 0$	$Q \neq 0$
STATIC $J = 0$	Schwarzschild $r_+ = 2M$	
STATIONARY $J \neq 0$		


THE MAIN CHARACTERS

M, Q, J

	$Q = 0$	$Q \neq 0$
STATIC $J = 0$	Schwarzschild $r_+ = 2M$	Reissner-Nordström $r_+ = M + \sqrt{M^2 - Q^2}$
STATIONARY $J \neq 0$	Kerr $r_+ = M + \sqrt{M^2 - J^2/M^2}$	Kerr-Newman $r_+ = M + \sqrt{M^2 - Q^2 - J^2/M^2}$

NEAR HORIZON LIMIT

EXTREMAL RN BLACK HOLE $M = Q$

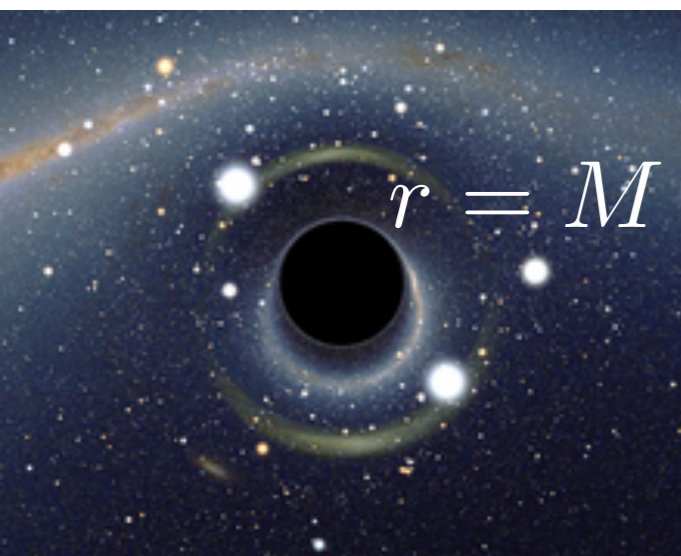
$$ds^2 = \left(1 - \frac{M}{r}\right)^2 dt^2 - \frac{dr^2}{\left(1 - \frac{M}{r}\right)^2} - r^2 d\Omega_{(2)}^2$$


S_2
 $SU(2)$

NEAR HORIZON LIMIT

EXTREMAL RN BLACK HOLE $M = Q$

$$ds^2 = \left(1 - \frac{M}{r}\right)^2 dt^2 - \frac{dr^2}{\left(1 - \frac{M}{r}\right)^2} - r^2 d\Omega_{(2)}^2$$

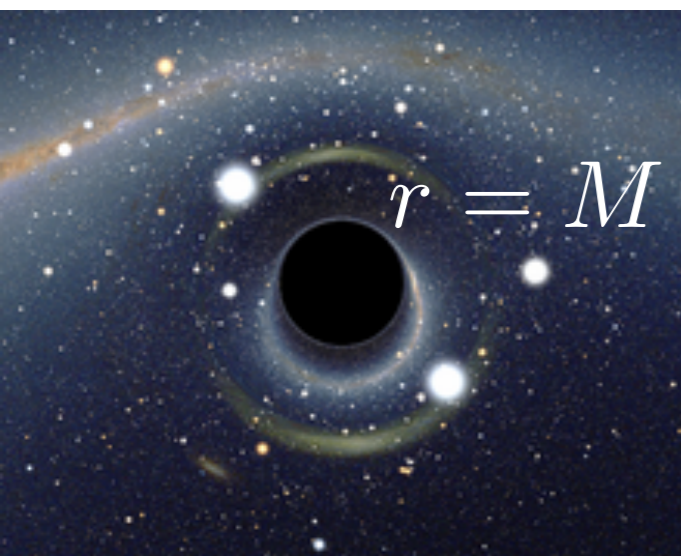


$$r = M (1 + \epsilon/z), \quad t = MT/\epsilon$$

NEAR HORIZON LIMIT

EXTREMAL RN BLACK HOLE $M = Q$

$$ds^2 = \frac{M^2}{z^2} (dT^2 - dz^2) - M^2 d\Omega_{(2)}^2 + \mathcal{O}(\epsilon)$$



$$r = M (1 + \epsilon/z), \quad t = MT/\epsilon$$

NEAR HORIZON LIMIT

EXTREMAL RN BLACK HOLE $M = Q$

$$ds^2 = \frac{M^2}{z^2} (dT^2 - dz^2) - M^2 d\Omega_{(2)}^2 + \mathcal{O}(\epsilon)$$

$\text{AdS}_2 \times S_2$

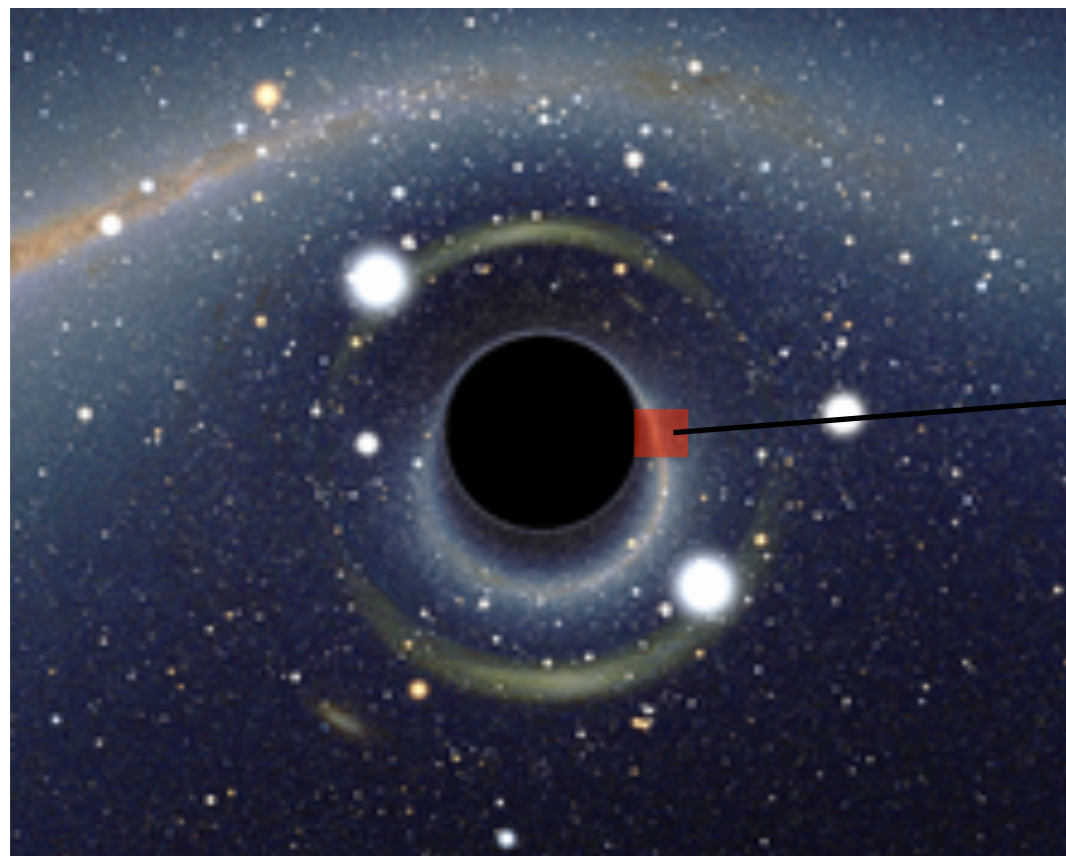
NEAR HORIZON LIMIT

FOR ANY EXTREMAL M, Q, J

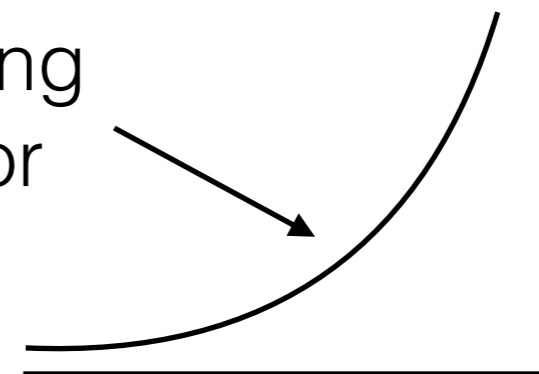
SAME SYMMETRIES AS

AdS_2

NEAR HORIZON LIMIT



warping
factor



AdS₂

AdS boundary $z = 0$
BH Horizon $z = \infty$

N.B.

$$g_{tt} = \left(1 - \frac{M}{r}\right)^2 \rightarrow 0$$

$r \rightarrow r_+$

SYMMETRIES

AdS₂

TIME
REPARAMETRIZATIONS

$$T \rightarrow f(T)$$

BOUNDARY

$SL(2, R)$

$$\partial_T, z\partial_z + T\partial_T$$
$$\frac{1}{2} (T^2 + z^2) + Tz\partial_z$$

BULK

SYMMETRIES

AdS₂

VIRASORO

TIME
REPARAMETRIZATIONS

$$T \rightarrow f(T) \quad f(T)\partial_T = \sum_n f_n T^n \partial_T$$

$SL(2, R)$

$$\partial_T, z\partial_z + T\partial_T \\ \frac{1}{2} (T^2 + z^2) + Tz\partial_z$$

1D CONFORMAL

$$\partial_T, T\partial_T, T^2\partial_T \\ z \rightarrow 0$$



THROUGH THE ¹⁶LOOKING GLASS

A SIMPLE QUANTUM MECHANICAL MODEL

$$S_E = \int d\tau \left[\psi_\alpha^{\dagger, \dot{a}} \dot{\psi}_\alpha^a + \left(\frac{\omega_{\alpha\beta}}{2} \psi_\alpha^a \epsilon_{ab} \psi_\beta^b + \text{h.c.} \right) \right]$$

$$\alpha, \beta = 1, \dots, N$$

$$\dot{a}, a = 1, 2 \text{ global } SU(2)$$

$\omega_{\alpha\beta}$

Gaussian variables with

$$\langle \omega \rangle = 0$$

$$\langle \omega^2 \rangle = \frac{\Omega^2}{N}$$

REPLICA TRICK

$$S_E = \int d\tau \left[\psi_\alpha^\dagger \dot{\psi}_\alpha + \left(\frac{\omega_{\alpha\beta}}{2} \psi_\alpha \psi_\beta + h.c. \right) \right]$$

At large N we can average over the noise

$$\overline{F} = -T \overline{\log Z} = -T \lim_{n \rightarrow 0} \frac{\overline{Z^n} - 1}{n}$$

$$\overline{Z^n} = \int d\omega d\bar{\omega} P(\omega, \bar{\omega}) Z^n$$

NEW VARIABLES

$$S_E = \int d\tau \left[\psi_\alpha^\dagger \dot{\psi}_\alpha + \left(\frac{\omega_{\alpha\beta}}{2} \psi_\alpha \psi_\beta + h.c. \right) \right]$$

For convenience

$$S_{AB}^a(\tau, \tau') = \delta_{AB} S^a(\tau, \tau') = \frac{1}{N} \langle \bar{\psi}_\alpha^{\dot{a}}(\tau) \psi_\alpha^a(\tau') \rangle$$

$$1 = \prod_{A,B} \int \mathcal{D}S_{AB} \mathcal{D}\Lambda_{AB} e^{i \int d\tau d\tau' \Lambda_{AB}(\tau, \tau') [S_{AB}(\tau, \tau') - \psi_{\alpha A}^\dagger(\tau) \psi_{\alpha B}(\tau')]}$$

THE EFFECTIVE LAGRANGIAN

$$S_E = \int d\tau \left[\psi_\alpha^\dagger \dot{\psi}_\alpha + \left(\frac{\omega_{\alpha\beta}}{2} \psi_\alpha \psi_\beta + h.c. \right) \right]$$

Large N saddle point

$$\frac{S_E}{Nn} = \sum_a \text{Tr} \log S^a + \int d\tau d\tau' \left[-\frac{\Omega^2}{2} S^1 S^2 + \delta(\tau - \tau') \sum_a \partial_\tau S^a \right]$$

SYMMETRIES

$$\frac{S_E}{Nn} = \sum_a \text{Tr} \log S^a + \int d\tau d\tau' \left[-\frac{\Omega^2}{2} S^1 S^2 + \delta(\tau - \tau') \sum_a \partial_\tau S^a \right]$$

AT LOW ENERGIES

TIME REPARAMETRIZATIONS $\times SU(2)$

$$\tau \rightarrow f(\tau) , \quad S^a(\tau, \tau') \rightarrow f'(\tau)^{1/2} S^a(f(\tau), f(\tau')) f'(\tau')^{1/2}$$

SPONTANEOUSLY BROKEN TO $SL(2, R) \times SU(2)$

$$S^a(\tau, \tau') \approx \frac{1}{\pi\Omega(\tau - \tau')}$$

SYMMETRIES

$$S_E = \int d\tau \left[\psi_\alpha^\dagger \dot{\psi}_\alpha + \left(\frac{\omega_{\alpha\beta}}{2} \psi_\alpha \psi_\beta + h.c. \right) \right]$$

NEAR HORIZON BH

LOW ENERGY

$$\tau \rightarrow f(\tau)$$

BOUNDARY

$$T \rightarrow f(T)$$

SPONTANEOUSLY BROKEN TO

$$SL(2, R)$$

BULK

$$SL(2, R)$$

N.B. True also in SYK

$$H = (i)^{q/2} \sum J_{i_1 i_2 \dots i_q} \psi_{i_1} \psi_{i_2} \dots \psi_{i_q}$$

SYMMETRIES

$$S_E = \int d\tau \left[\psi_\alpha^\dagger \dot{\psi}_\alpha + \left(\frac{\omega_{\alpha\beta}}{2} \psi_\alpha \psi_\beta + h.c. \right) \right]$$

NEAR HORIZON BH

EXPLICIT BREAKING

UV Effects

$$\partial_\tau S^a$$

Deviations from
extremality and leading
near horizon corrections

THERMODYNAMICS

$$S_E = \int d\tau \left[\psi_\alpha^\dagger \dot{\psi}_\alpha + \left(\frac{\omega_{\alpha\beta}}{2} \psi_\alpha \psi_\beta + h.c. \right) \right] \quad \text{NEAR HORIZON BH}$$

$$C \sim T$$

$$C \sim T$$

Massless (large N) modes
from the spontaneous breaking of

$$SL(2, R)$$

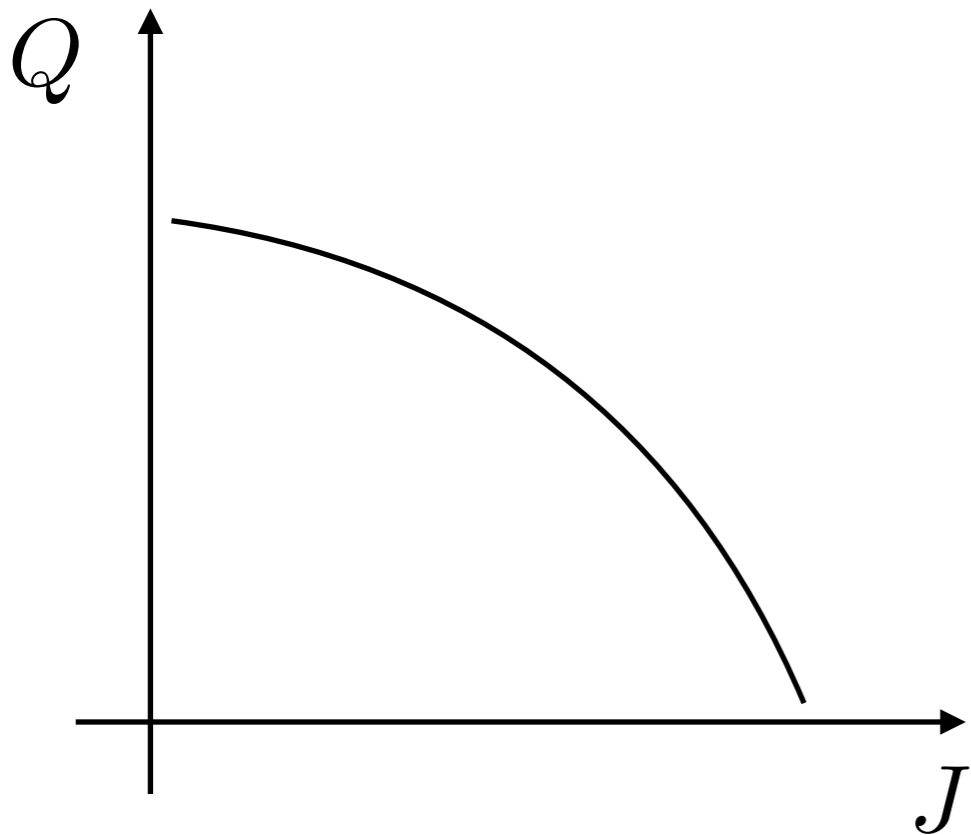
$$S \propto \int d\tau \{f(\tau), \tau\} \quad \{f(\tau), \tau\} = \frac{f'''}{f'} - \frac{3}{2} \left(\frac{f''}{f'} \right)^2$$

EXTREMALITY CONTINUED

M, Q, J

	$Q = 0$	$Q \neq 0$
STATIC $J = 0$		
STATIONARY $J \neq 0$		Kerr-Newman $M^2 = Q^2 + \frac{J^2}{M^2}$

EXTREMALITY AND MARGINAL DEFORMATIONS



$$SU(2) \rightarrow U(1)$$
$$z \int d\tau \bar{\psi}_{\alpha}^{\dot{a}}(\tau) \boldsymbol{\sigma}_z^{\dot{a}b} \psi_{\alpha}^b(\tau)$$

MARGINAL DEFORMATIONS

$$z \int d\tau \bar{\psi}_{\alpha}^{\dot{a}}(\tau) \sigma_z^{\dot{a}b} \psi_{\alpha}^b(\tau)$$

$$z < 2\Omega$$

$$z > 2\Omega$$

$$S^1(\tau, \tau') = \frac{1}{\Omega} \frac{\sqrt{1 - \left(\frac{z}{2\Omega}\right)^2}}{\pi(\tau - \tau')} + \text{contact}$$

$$S^1(\tau, \tau') \propto \delta(\tau - \tau')$$

EXTREMALITY AND MARGINAL DEFORMATIONS

$$z \int d\tau \bar{\psi}_{\alpha}^{\dot{a}}(\tau) \sigma_z^{\dot{a}b} \psi_{\alpha}^b(\tau)$$

$$z < 2\Omega$$

$$J = M \sqrt{M^2 - Q^2}$$

Continuum of CTFs (N.B. large N) ~
1 parameter family of extremal BHs

$$z > 2\Omega$$

$$J > M \sqrt{M^2 - Q^2}$$

BH solution ceases to exist ~
Gapped fermions that saturate J
(all “spins” are aligned)

CONCLUSION

- SEVERAL ROTATING BLACK HOLES CONSISTENT WITH EXTREMALITY HAVE BEEN OBSERVED IN THE UNIVERSE
- THERE ARE NON TRIVIAL, FASCINATING ANALOGIES BETWEEN EXTREMAL BLACK HOLES IN 4D AND SIMPLE LARGE-N QUANTUM MECHANICAL MODELS
- THE CORRESPONDENCE IS STILL FAR FROM BEING FORMULATED PRECISELY, BUT WE MIGHT BE ON THE RIGHT PATH

BACKUP

NON-ANALYTIC BEHAVIOR AT SMALL TEMPERATURE

$$z < 2\Omega$$

$$\log Z = \frac{N\pi}{6\Omega^2\beta} \sqrt{4\Omega^2 - z^2} + \dots$$

$$J < M\sqrt{M^2 - Q^2}$$

$$F = \frac{Q}{s+1} \left[\sqrt{\frac{(s+1)(s+3)^3}{32}} + \mathcal{O}(T) \right]$$

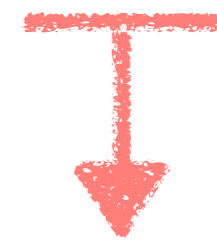
$$s \equiv \sqrt{1 - 8\Omega_H^2 Q^2}$$

NON-ANALYTIC BEHAVIOR AT ZERO TEMPERATURE

$$T = 0 \quad \frac{1}{N} \frac{d^2}{dz^2} \langle \hat{J}_z \rangle_{\beta=\infty} = - \frac{z}{\Omega^2 \sqrt{4\Omega^2 - z^2}}$$

$$J = M \sqrt{M^2 - Q^2}$$

$$\log Z_0 = \frac{Q^2}{2} \sqrt{4\pi^2 - \beta_L^2}$$



Chemical potential for J
at extremality

EXTREMAL BLACK HOLES

$$M, Q, J$$

	$Q = 0$	$Q \neq 0$
STATIC $J = 0$	$M \geq 0$	$M \geq Q$
STATIONARY $J \neq 0$	$M^2 \geq J$	$M^2 \geq Q^2 + \frac{J^2}{M^2}$