

Cosmic Frontier and machine learning

C. Tunnell (U. Chicago and Rice University)

July 4th, 2018

Caveat: all data either shown publicly or
tweaked to make non-experiment specific

~~Cosmic Frontier and
machine learning~~

12 minute on manifold learning

C. Tunnell (U. Chicago and Rice University)

July 4th, 2018

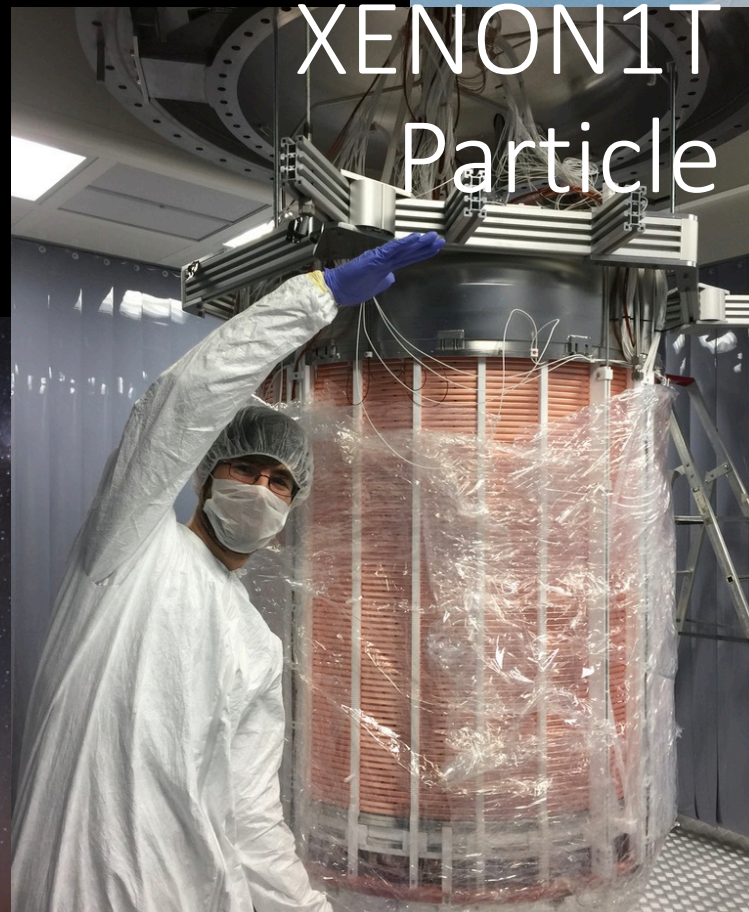
Caveat: all data either shown publicly or
tweaked to make non-experiment specific

Deep Skies Collaboration

Informal astrophysics group exploring
useful ML applications in
astrophysics

SPT
CMB

DES
Optical



Cutting through the hype

Astronomy	Particle physics
Image data (i.e. same as classifying cat/dog)	Tree-like data (e.g. graph convolution applications limited)
Public data releases with docs	Limited data release, often requires membership
Toolchain more friendly to ML	ROOT

What new measurement does ML open?

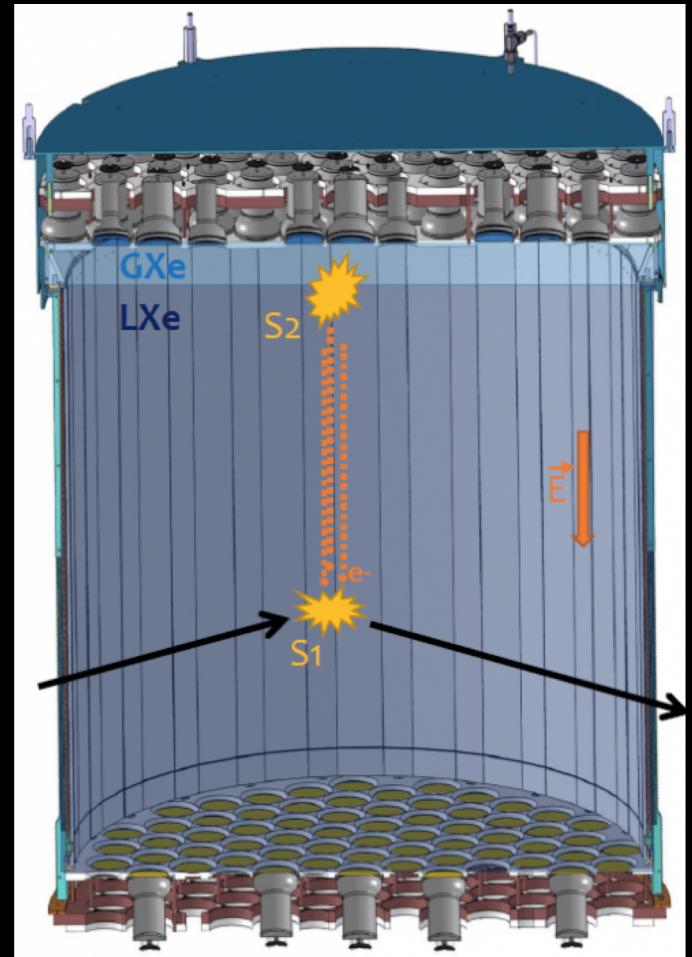
Within direct-detection DM, limited impact.

I will discuss one neat example (which coincidentally first plenary said used widely elsewhere but not in HEP)



What is XENON?

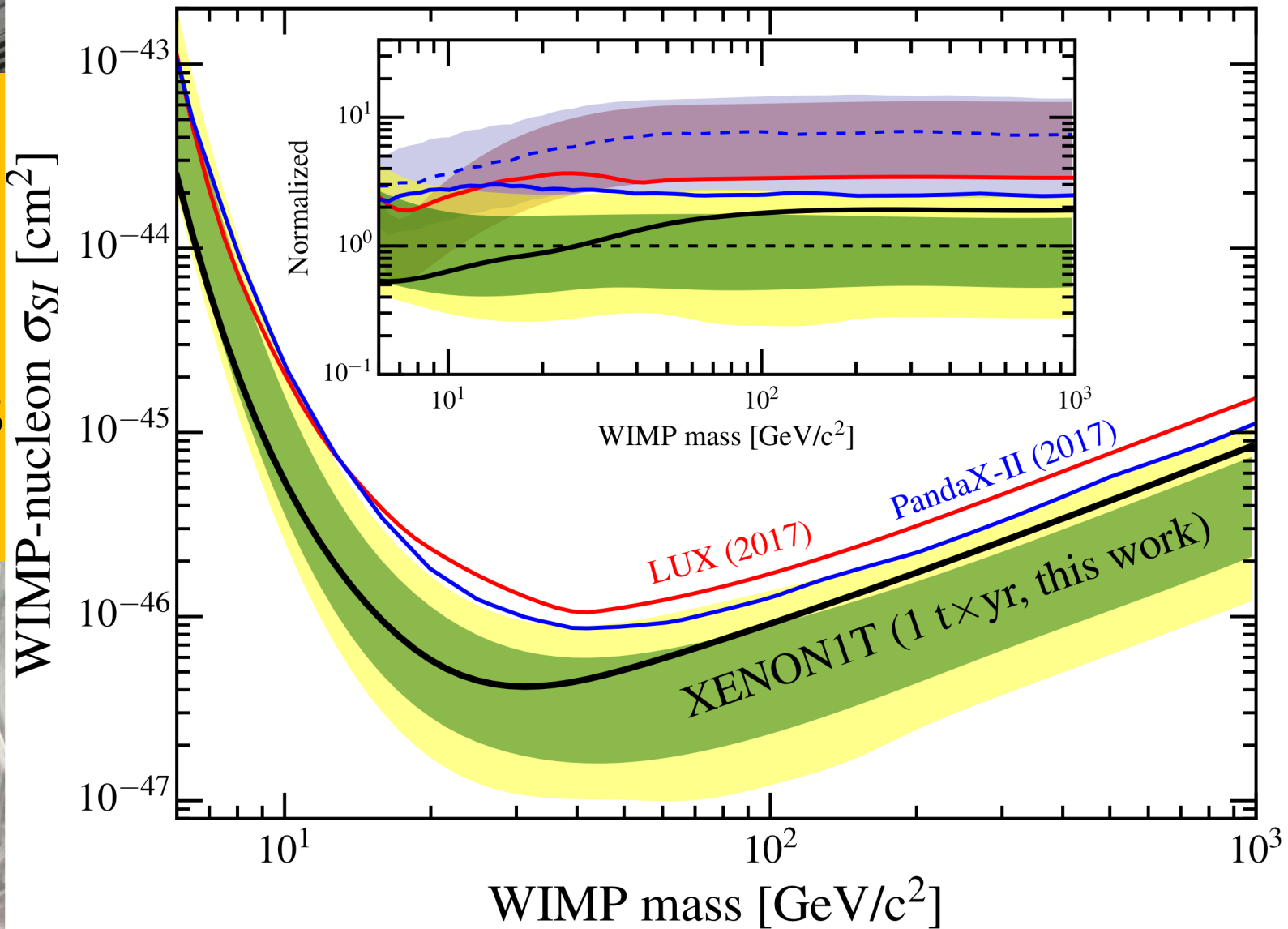
Liquid XENON dark matter detector instrumented with 248 photomultipliers and 10-ns flash ADCs. We make a world-leading new experiment every few years.



What is XENON?

Liquid XENON dark matter detector instrumented with 248 photomultipliers and 10-ns flash ADCs. We make a world-

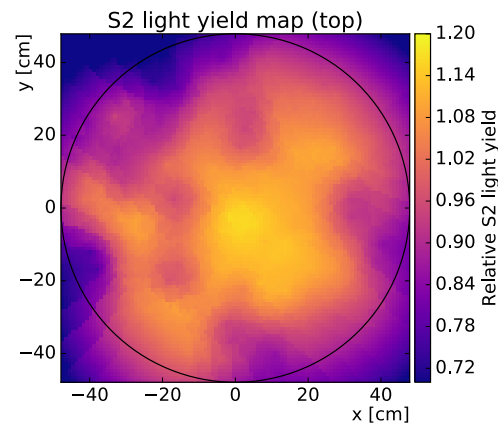
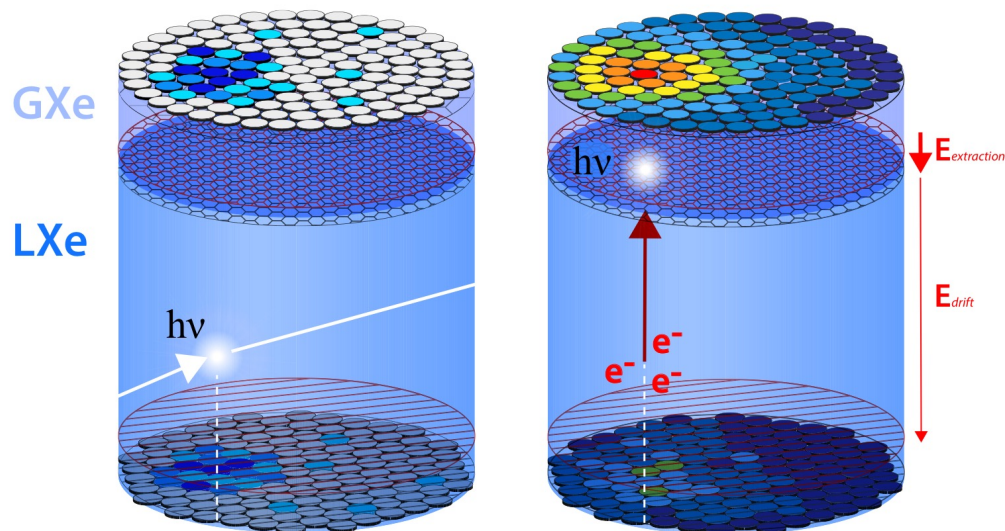
We are
 current
 front
 runners
 ... but



Dark Matter Search Results from a One Tonne×Year Exposure of XENON1T

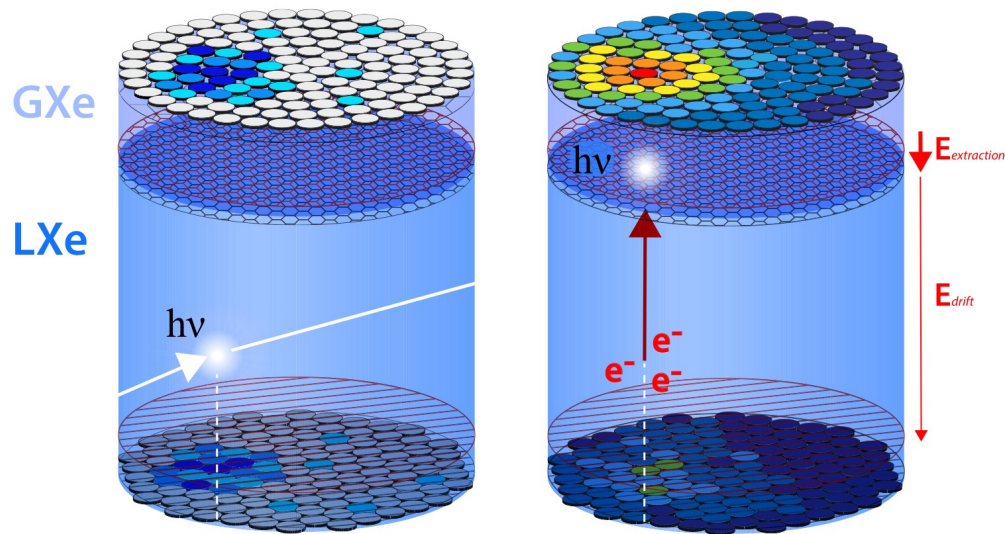
TABLE I: Best-fit expected event rates with 278.8 days live-time in the 1.3 t fiducial mass, 0.9 t reference mass, and 0.65 t core mass, for the full (cS1, cS2_b) ROI and, for illustration, in the NR signal reference region. The table lists each background (BG) component separately and in total, the observed data, and the expectation for a 200 GeV/c² WIMP prediction assuming the best-fit $\sigma_{SI} = 4.7 \times 10^{-47}$ cm².

Mass (cS1, cS2 _b)	1.3 t Full	1.3 t Reference	0.9 t Reference	0.65 t Reference
ER	627±18	1.62±0.30	1.12±0.21	0.60±0.13
neutron	1.43±0.66	0.77±0.35	0.41±0.19	0.14±0.07
CEνNS	0.05±0.01	0.03±0.01	0.02	0.01
AC	0.47 ^{+0.27} _{-0.00}	0.10 ^{+0.06} _{-0.00}	0.06 ^{+0.03} _{-0.00}	0.04 ^{+0.02} _{-0.00}
Surface	106±8	4.84±0.40	0.02	0.01
Total BG	735±20	7.36±0.61	1.62±0.28	0.80±0.14
WIMP _{best-fit}	3.56	1.70	1.16	0.83
Data	739	14	2	2



THE CHALLENGE

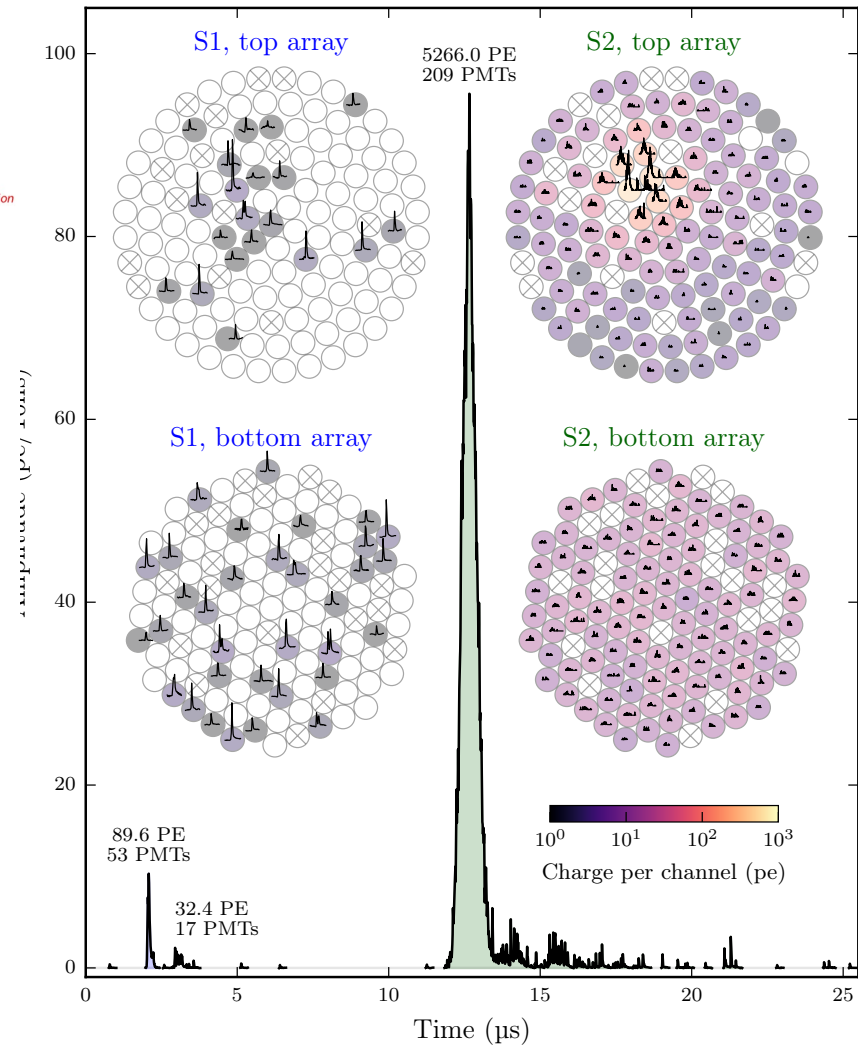
1. The optics and response isn't uniform for S2s
 1. Hard to measure reflectivity Teflon ex situ
2. There is no internal calibration
3. When doing ML, including what we do know is not straightforward



ML Challenge: Photosensors locations in 2D that aren't grid

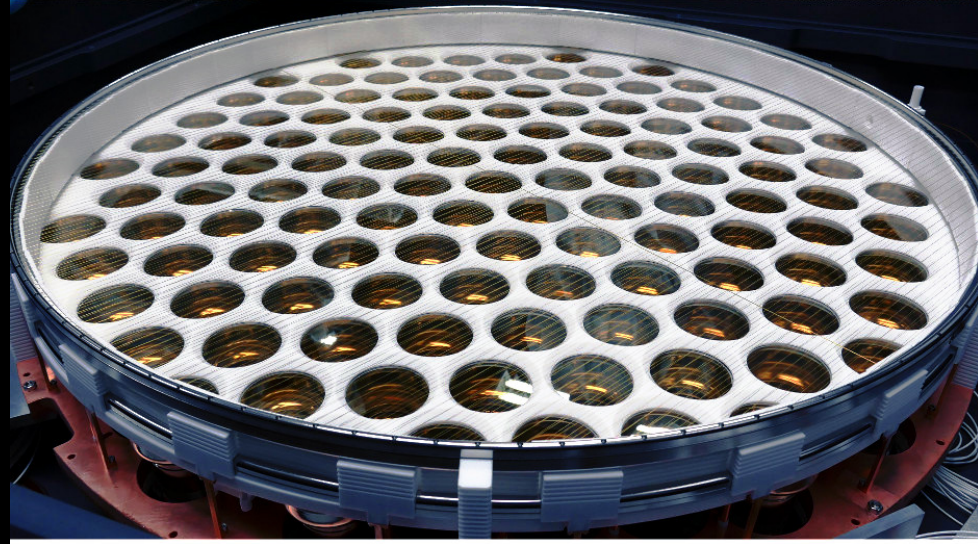
→ Sporadic spacing hard

→ Image not 2D but rather 248D



Each photosensor represents
input dimension

Looking for embedding
 $\mathbb{R}^{128} \rightarrow \mathbb{R}^2$

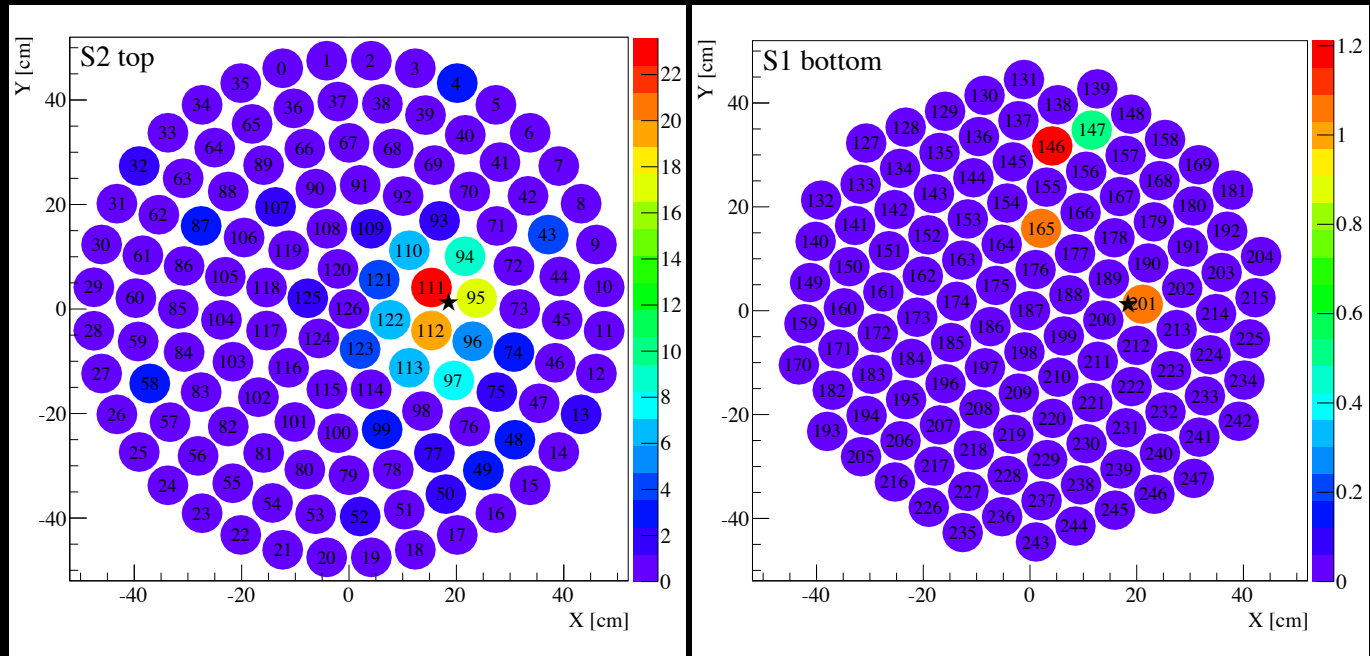


Each photosensor represents input dimension

Looking for embedding

$$\mathbb{R}^{128} \rightarrow \mathbb{R}^2$$

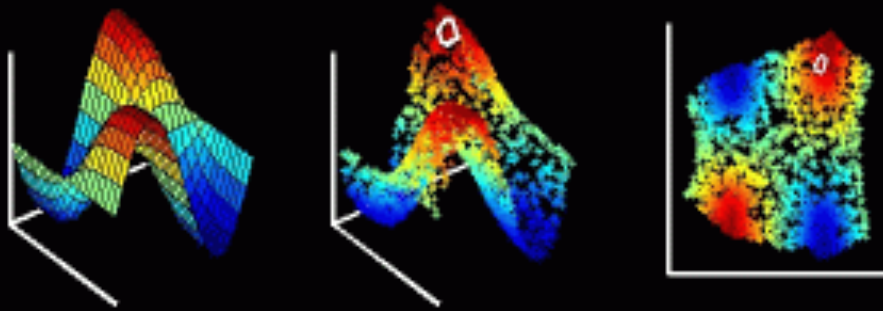
Key insight: sensors near in 2D see similar
signals, so sensor dimensions correlated



Dimensionality reduction

Often have higher dimensional data that we know has a low-dimensional representation

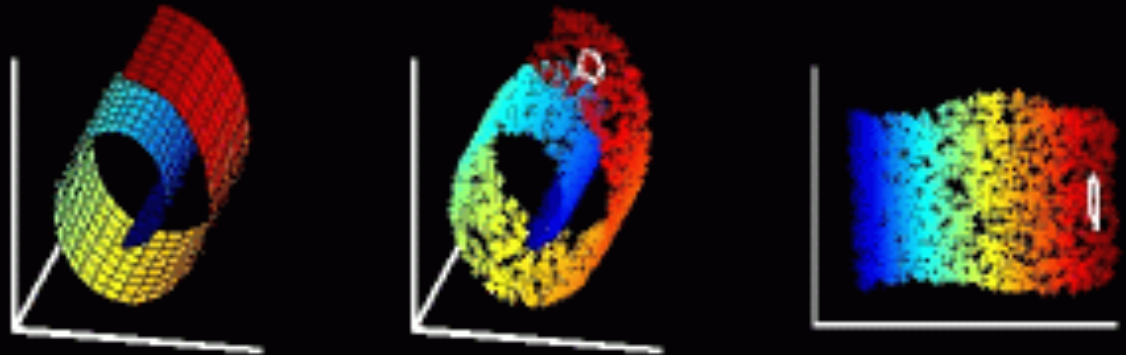
3D \rightarrow 2D: how to flatten? i.e. preserve neighbors



Twin peaks

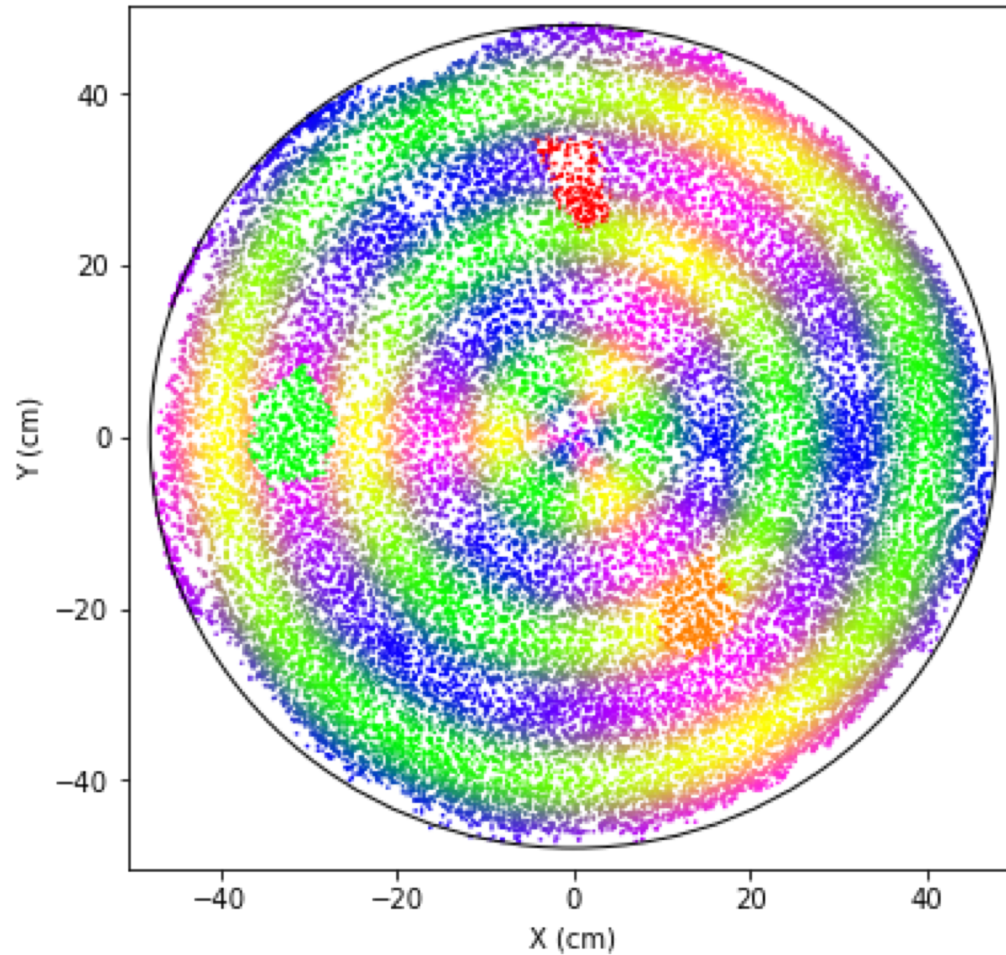
Swiss roll

“Manifold learned”



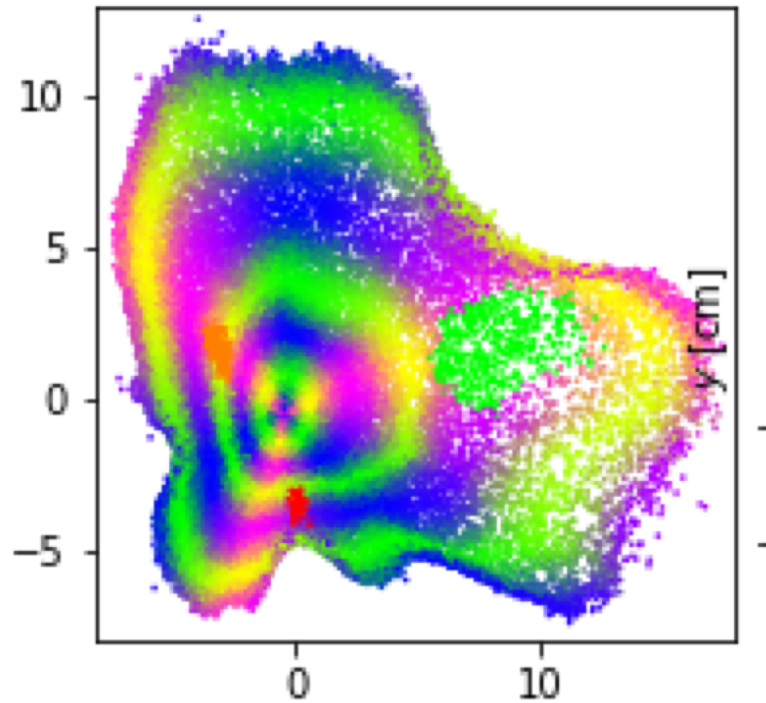
Baseline

Fig 3: TPF point field (space **C**)

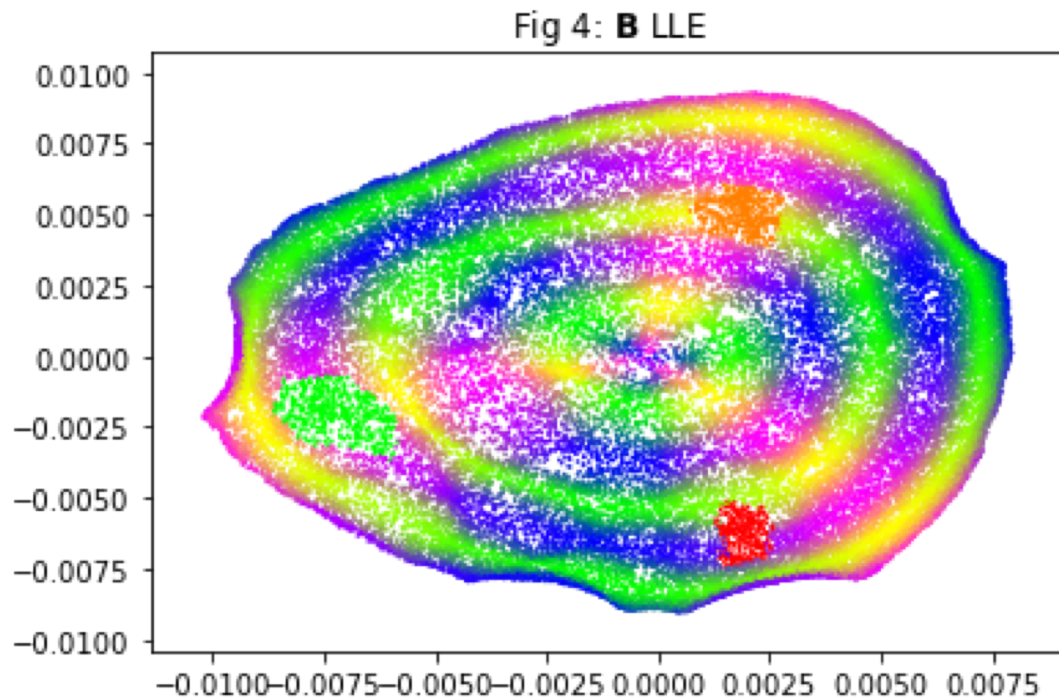


Dimensionality reduction: Principal Component Analysis

Fig 36: pca **B**

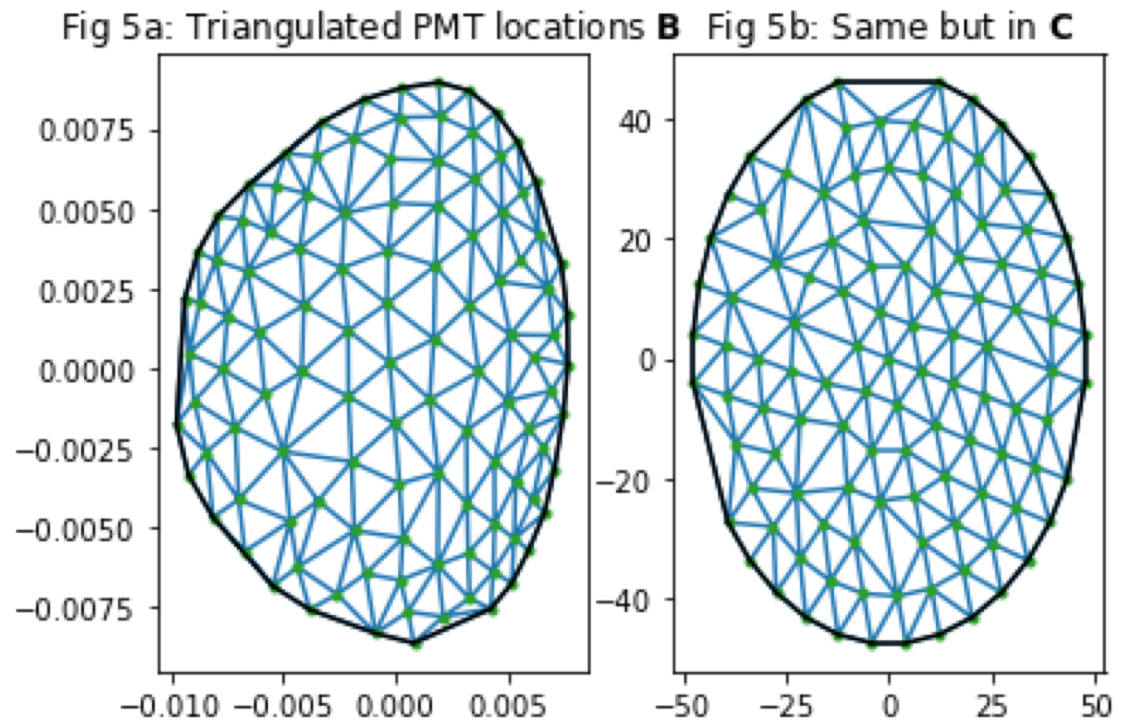
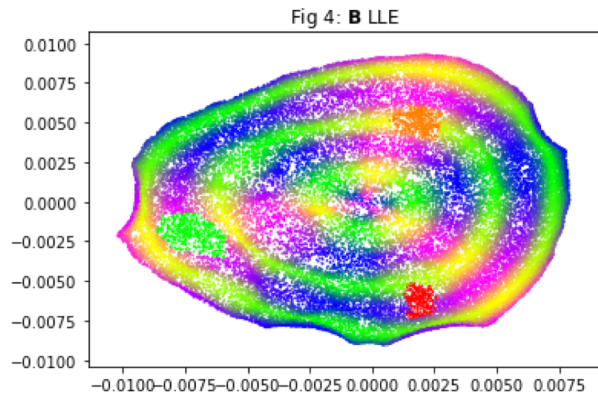


Dimensionality reduction: Local linear embedding



Scale off.... now can use physics knowledge!!

Stretching image



Triangulation interpolation from known positions in learned and lab space

Applications / Conclusion

- DAQ: can find flipped channels
- Can “learn” optics/response of uncalibratable detector
- Can be used for detector alignment
- Only assumption neighborhoods preserved
 - P.S. ask me about spatial statistics and Ripley K functions if you ever want to check your manifold after learning on ‘uniform’ distribution