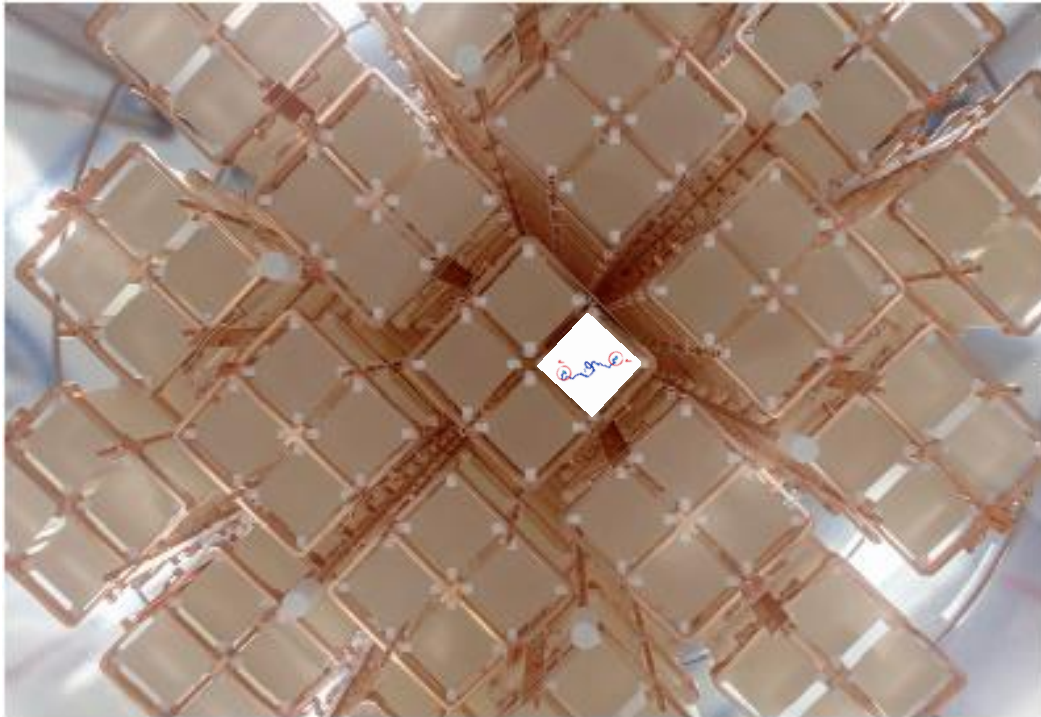


# The search for neutrino-less double beta decay

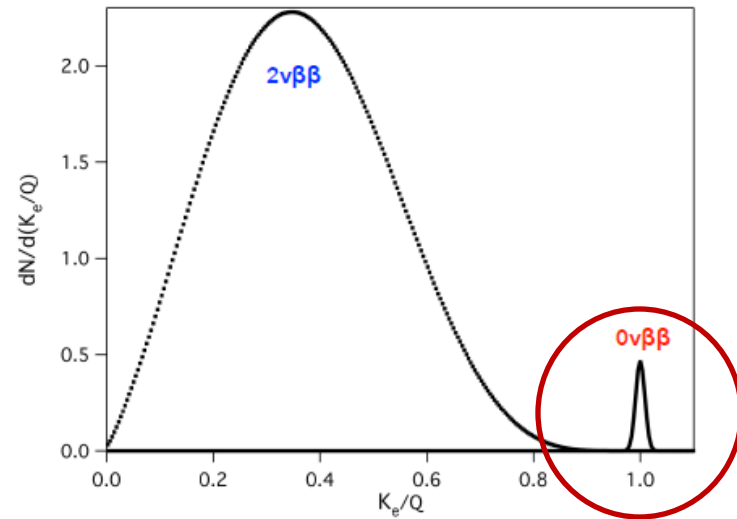
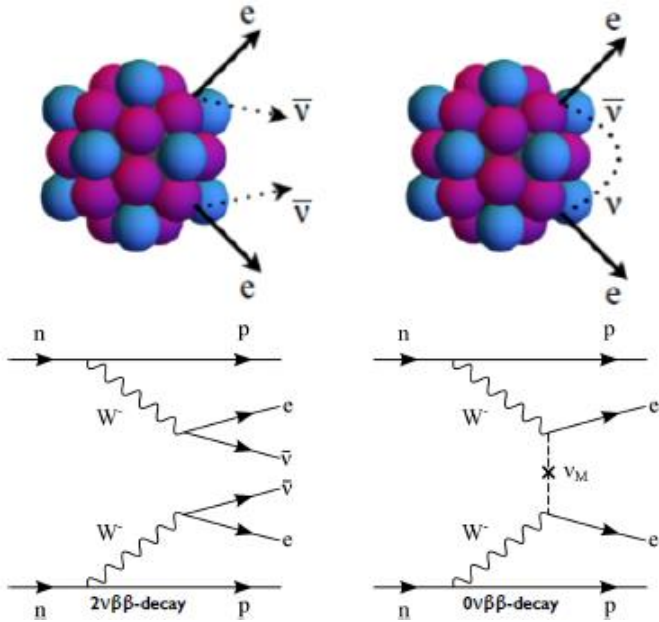
A (Europe-focused) review of current **experiments** and of the impressive results achieved in the last two years

Experimental strategies for the **next generation**

Why such a growth of interest in this field, **right now?**



# Neutrino-less double beta decay (NLDBD)



To date, NLDBD is the only viable option to show that neutrinos are Majorana particles ( $\nu = \bar{\nu}$ )

## Experimental observation of neutrino-less double beta decay

It will establish **violation of lepton number** in particle physics

Shed light on **mass generation mechanisms** and the smallness of neutrino masses

Open a window to understand **matter dominance** in the universe

Provide information on the **size and pattern** of neutrino masses

# NLDBD and neutrino masses

The decay half-life is connected in a non trivial manner to the neutrino masses

Half-life      Phase space factor

$$\frac{1}{T_{1/2}^{0\nu}} = G^{0\nu} |M^{0\nu}|^2 |\langle m_{\beta\beta} \rangle|^2$$

Nuclear matrix element

Effective Majorana neutrino mass:  
(exact form depends on the lepton flavor violating mechanism!)

[See also talk from [E. Lisi](#)]

$$\langle m_{\beta\beta} \rangle = |U_{e1}^2 m_1 + U_{e2}^2 m_2 e^{i(\alpha_1 - \alpha_2)} + U_{e3}^2 m_3 e^{i(-\alpha_1 - 2\delta)}|$$

$$m_{\beta\beta} \equiv \sum_{i=1}^3 U_{ei}^2 m_i$$

Phase space: known with good precision  
Nuclear matrix element: major challenge  
in nuclear physics

$$M_{(0\nu)} = g_A^2 M^{(0\nu)}$$

$$\begin{aligned} g_A^{\text{quark}} &= 1 \\ g_A^{\text{nucleon}} &= 1.27 \\ g_A^{2\nu\beta\beta} &= 1.27 \cdot A^{-0.18} \\ g_A^{0\nu\beta\beta} &= ?? \end{aligned}$$

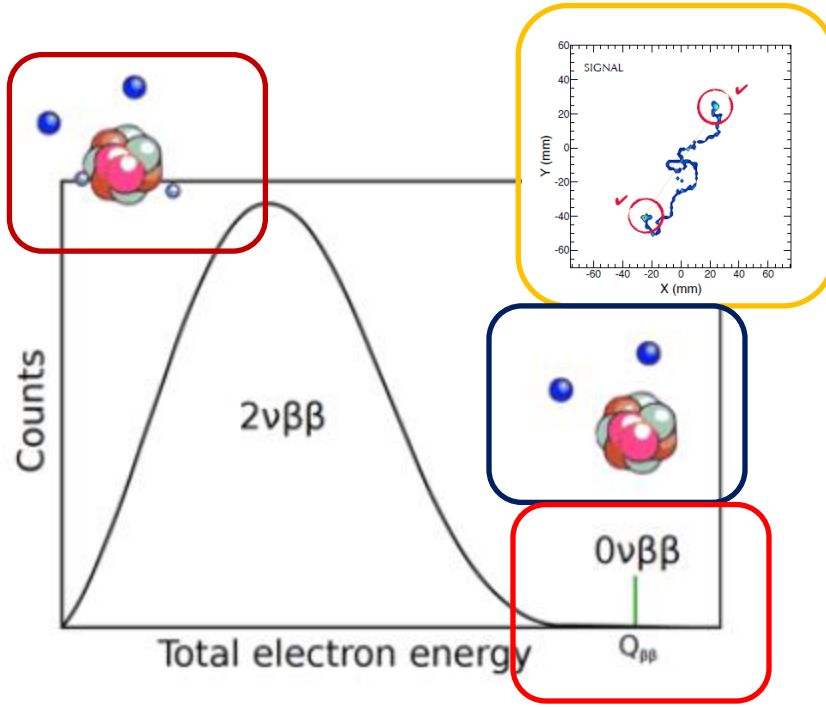
$m_{\beta\beta}$

Linked to neutrino eigenstate (as cosmology), mixing angles (as oscillations) and Majorana CP violating phases (unique!)

If neutrino are Majorana particles, oscillation data provide guidance for the NLDBD half time. Still, the constraints are loose and a lot of work is needed to reduce the uncertainties on M and  $g_A$ .

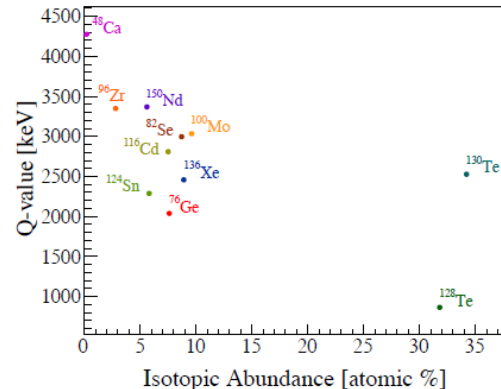
# The perfect experiment

The experimental signature of NLDBD is extremely simple: two electrons whose energy sum (Q-value) is known in advance with very high precision (Penning traps). “Intrinsic” background:  $(A,Z) \rightarrow (A,Z+2) 2e^- 2\nu$



Track the out-coming electrons to separate  $\alpha$ ,  $\beta$ ,  $\gamma$  from  $\beta\beta$

Choose the detector with the best energy resolution (<1% FWHM) to reduce the “intrinsic”  $2\nu$  background



Select isotopes with high Q-value against natural radioactivity and high isotopic abundance

$\beta\beta$ Decay Reaction	Isotopic Abundance [atomic %]	Q-value [keV]
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	0.2	4274
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	7.6	2039
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	8.7	2996
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	2.8	3348
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	9.6	3034
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	7.5	2814
$^{124}\text{Sn} \rightarrow ^{124}\text{Te}$	5.8	2288
$^{128}\text{Te} \rightarrow ^{128}\text{Xe}$	31.8	866
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	34.2	2528
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	8.9	2458
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	5.6	3368

# Back to earth....

The number of candidate isotopes is rather small and we often end up choosing among alternative options:

## Superior energy resolution (0.1%)

Germanium detectors      **Gerda, Majorana**

Bolometers      **CUORE, CUPID-0, AMORE**

Compromise:  

## Superior tracking capability

Source not in detector

TPC with source in detector

**SuperNemo**  





EXO-200, **Next**, PandaX-III  

Compromise:

## Huge isotope masses (1 ton)

Loaded scintillators      **Kamland-Zen, SNO+**

Compromise:   

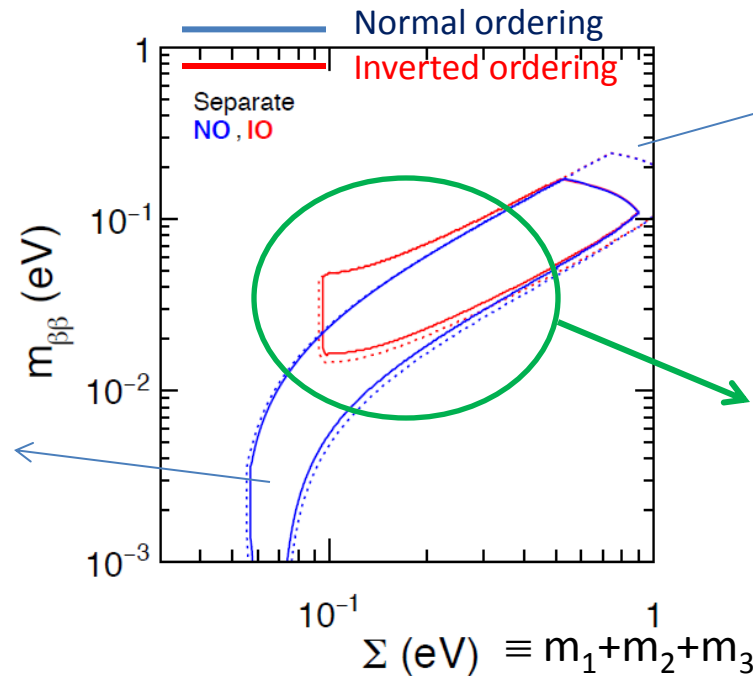
-  energy resolution
-  restricted isotope list
-  tracking capability
-  mass scalability

# From the “perfect” to the “best” experiment

After 2012 (high precision measurement of all mixing angles),  $m_{\beta\beta}$  is well constrained from neutrino (and cosmology) data. Now we know where to look for.

See also talks of E. Lisi and S. Schoenert

Accidental cancellation among CP phases. Worst case scenario: possible but unlikely



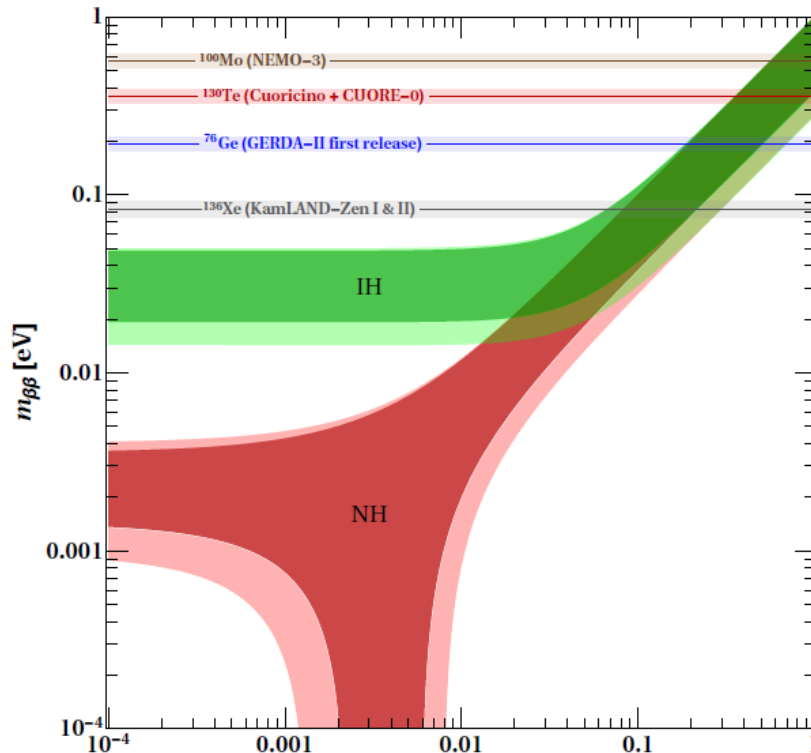
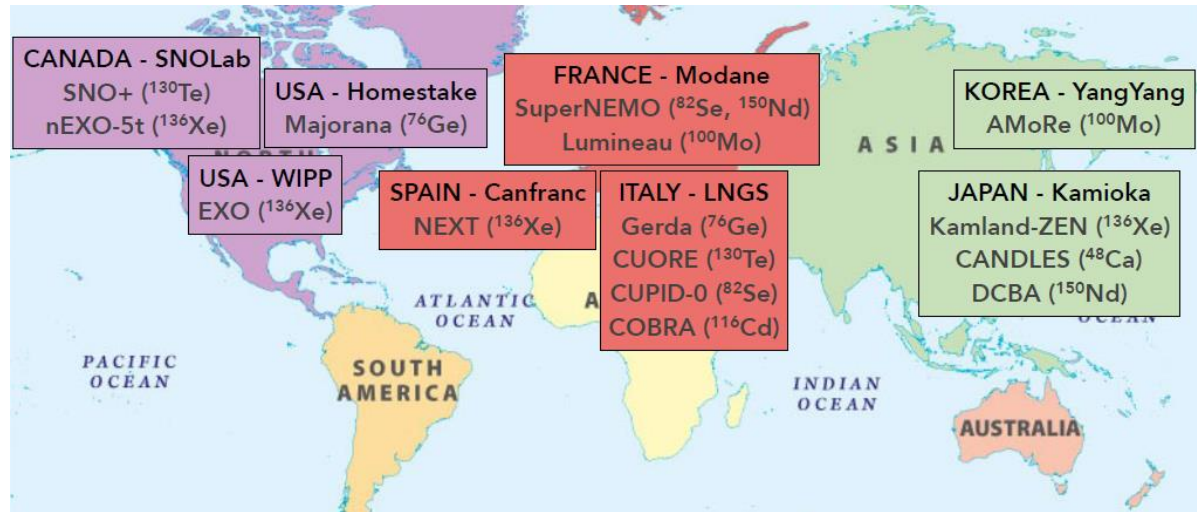
Nearly degenerate mass eigenstate. Best case scenario: possible but unlikely (cosmology)

We want to be here: possible and likely

We are developing technologies that are able to explore NLDBD half-lives at the level of  $10^{26}$  y and that are potentially scalable up to  $10^{28}$  y.

For “zero background experiments”, sensitivity  $\sim$  exposure  
For finite background: sensitivity  $\sim$  (exposure) $^{1/2}$

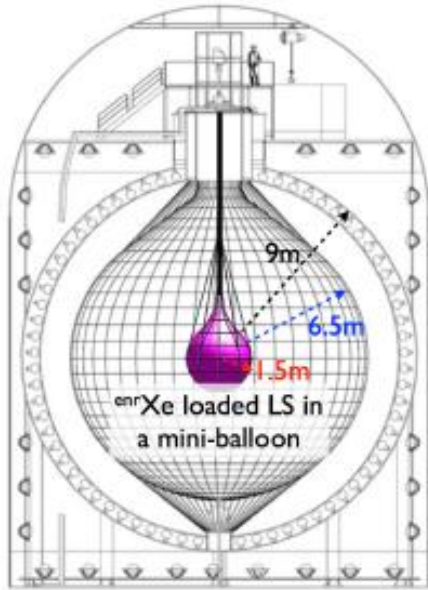
# The present generation of NLDBD experiment



Osc. par.: F. Capozzi *et al.*, *Nucl. Phys. B* 908, 218 (2016)  
 NMEs (IBM-2): J. Barea *et al.*, *Phys. Rev. C* 91, 034304 (2015)  
 PSFs: J. Kotila, F. Iachello, *Phys. Rev. C* 85, 034316 (2012)  
 $g_A = 1.269$   
 Experiment sensitivities:  
 $^{100}\text{Mo}$ : R. Arnold *et al.*, *Phys. Rev. D* 92, 072011 (2015)  
 $^{130}\text{Te}$ : K. Alfonso *et al.*, *Phys. Rev. Lett.* 115, 102502 (2015)  
 $^{76}\text{Ge}$ : M. Agostini, Presentation at Neutrino 2016  
 $^{136}\text{Xe}$ : A. Gando *et al.*, *Phys. Rev. Lett.* 117, 082503 (2016)

# Kamland-Zen (Japan)

The world leader in the field and the paradigm of a scalable, finite-background experiment

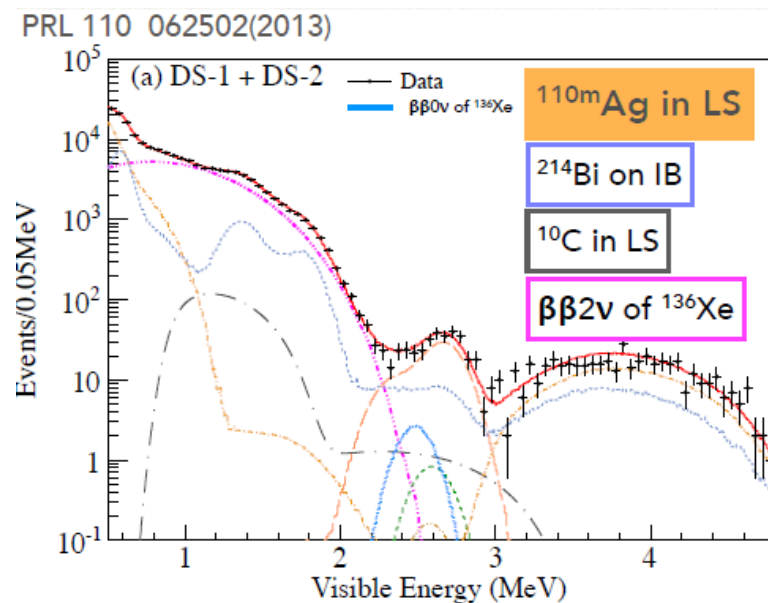


Isotope:  $^{136}\text{Xe}$ ,  $Q\text{-value} = 2458 \text{ keV}$

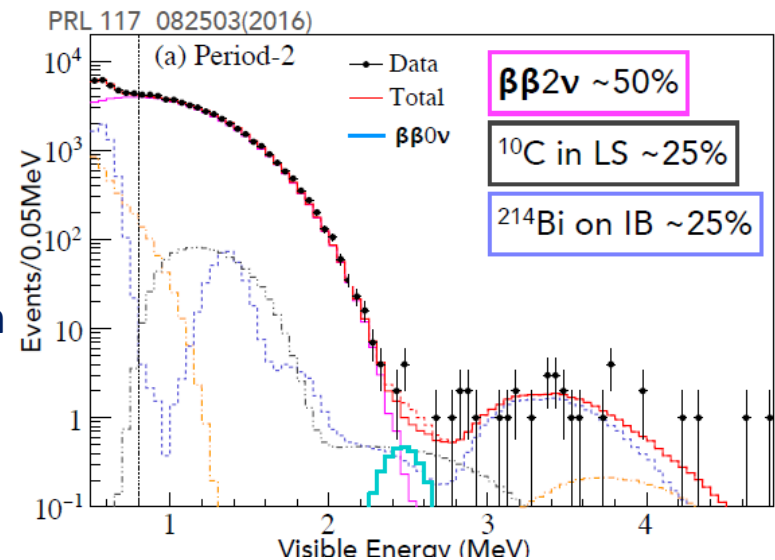
$T_{1/2} > 1.07 \times 10^{26} \text{ y}$  (90 % CL)

- ✓ Active target 350 kg
- ✓ Energy resolution 4% at  $Q\text{-value}$
- ✓ Position reconstruction: bkg rejection
- ✓ Bi-Po for  $^{214}\text{Bi}$ , 3-fold coincidence for  $^{10}\text{C}$

Unexpected background:  $^{110\text{m}}\text{Ag}$



→  
18 months  
purification  
campaign



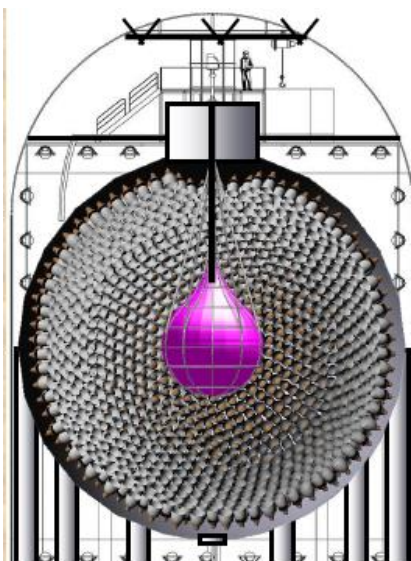


# The scalability of Kamland-Zen

Phase III (Kamland-Zen 800) : 750 kg of enriched Xe.

- ✓ New clean Inner Balloon to reduce  $^{214}\text{Bi}$
- ✓ Improve  $^{10}\text{C}$  tag

In progress. Slowed down by a leak in the balloon (see Y. Efremenko at MEDEX 2017)



Improvement in energy resolution steers the long-term plans of Kamland2-Zen

Winston cone	light collection $\times 1.8$
high q.e. PMT $17''\phi \rightarrow 20''\phi$ $\epsilon = 22 \rightarrow 30\%$	light collection $\times 1.9$
New LAB LS (better transparency)	light collection $\times 1.4$

expected  $\sigma(2.6\text{MeV}) = 4\% \rightarrow \sim 2\%$   
target sensitivity 20 meV

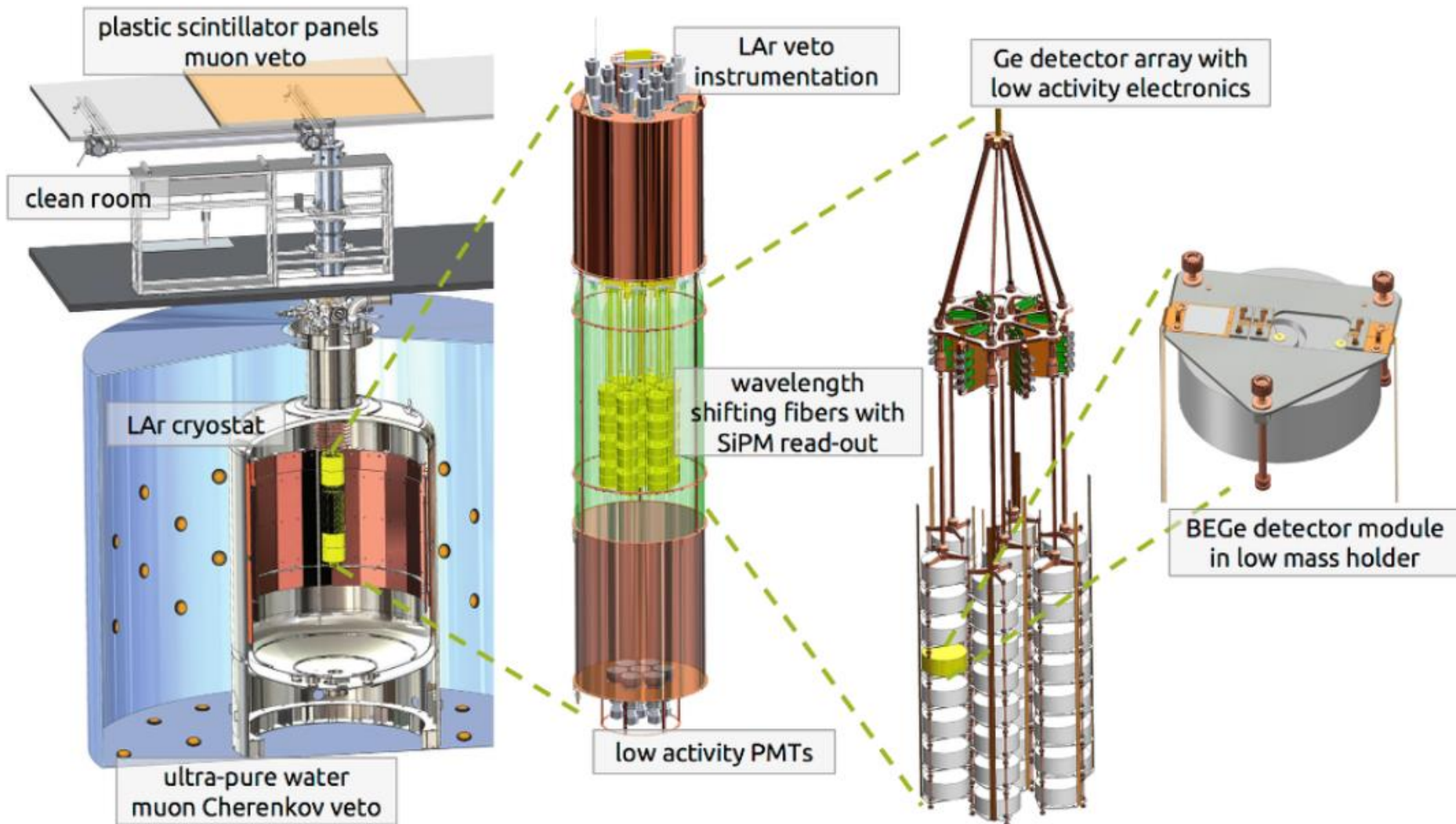
Ultimate facility:

SuperKamland-ZEN in parallel with HyperKamiokande ?

Similar considerations apply to SNO+ (Canada)  
Not covered in this review

# GERDA (Gran Sasso, Italy)

The paradigm of a scalable, zero background experiment

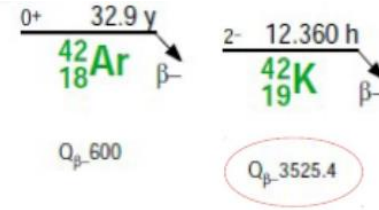


Isotope:  $^{76}\text{Ge}$ , Q-value = 2039 keV  
 $T_{1/2} = 5.3 \cdot 10^{25}$  y (90 % CL)

Both **GERDA** and **Majorana (US)** employ high-purity Germanium detector with 3 keV energy resolution at the  $^{76}\text{Ge}$  Q-value. The  $2\nu$  “intrinsic” background is not an issue neither for the current nor for the next generation of experiment

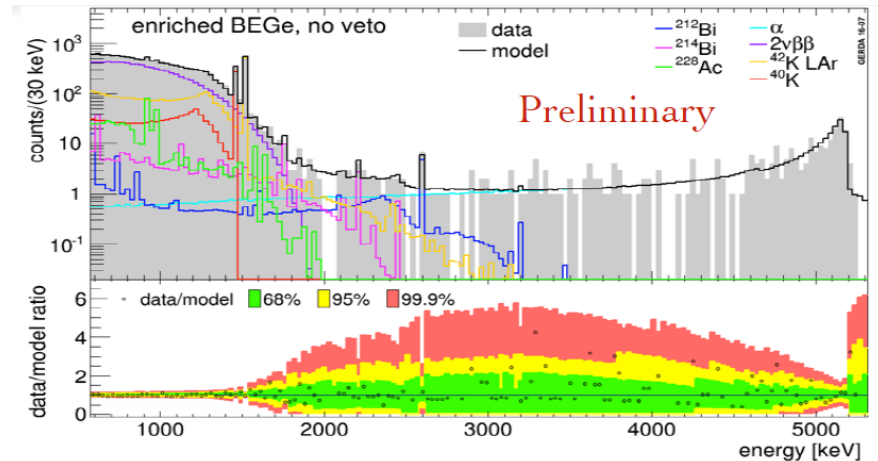
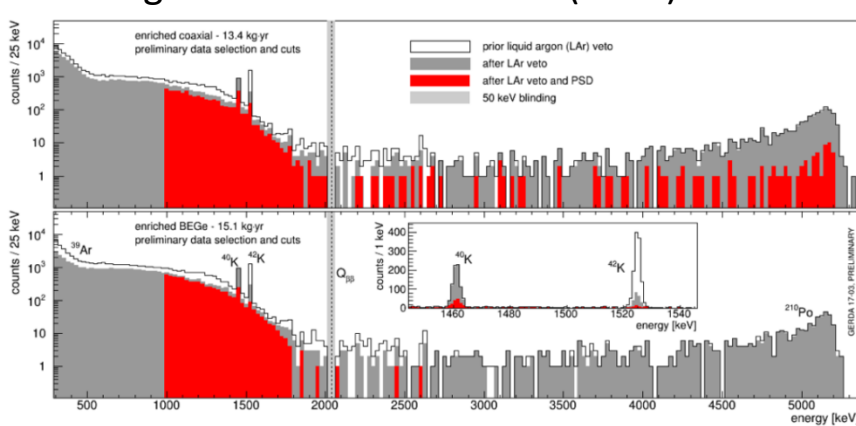
“Zero background” has been achieved in Gerda employing a liquid argon veto system and pulse shape analysis

Unexpected background:  $^{42}\text{K}$



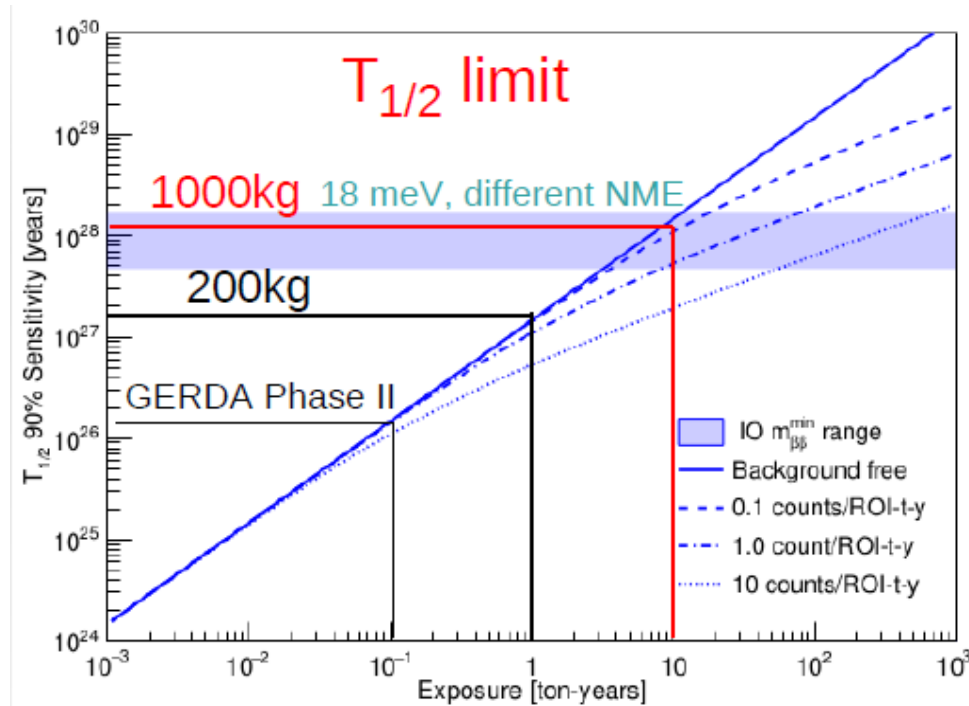
(charged)  $^{42}\text{K}$  drift in field of Ge detectors  $^{42}_{20}\text{Ca}$

M. Agostini et al. Nature 544 (2017) 47



Background index:  $0.6 \cdot 10^{-3}$  counts/keV kg y

# The scalability of GERDA and Majorana



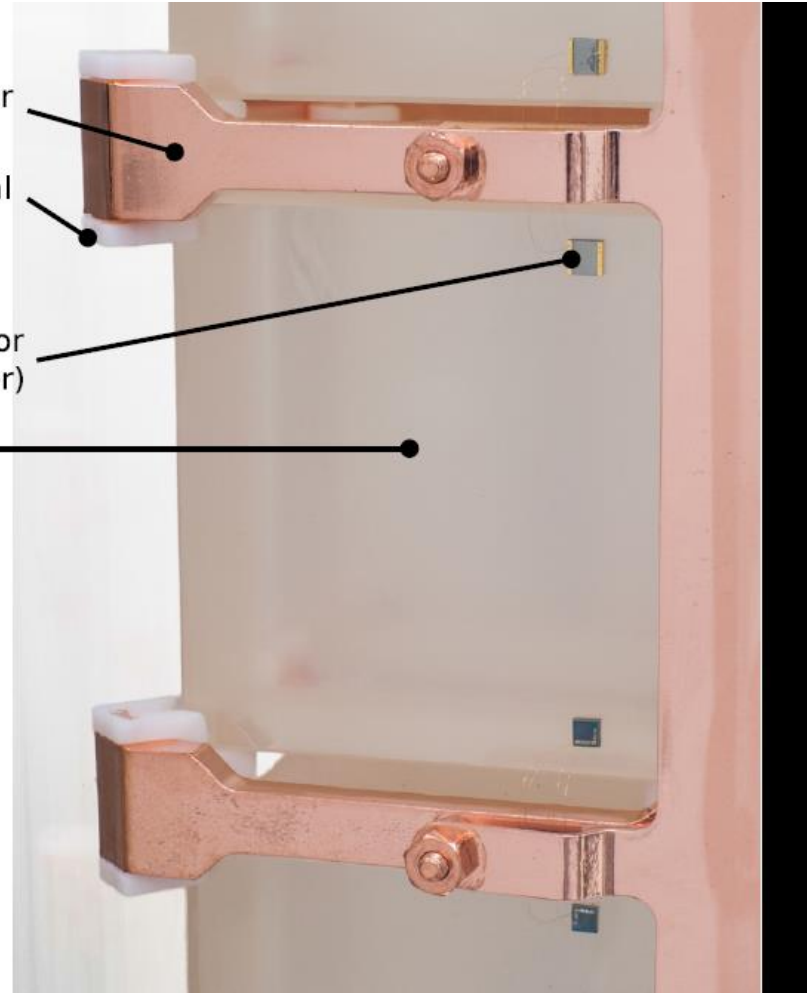
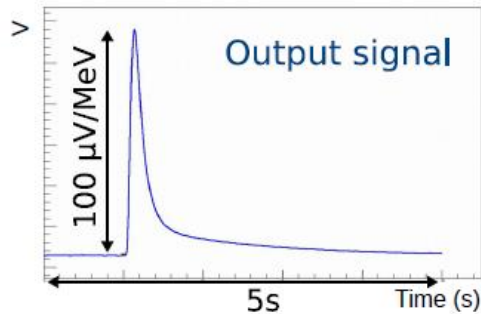
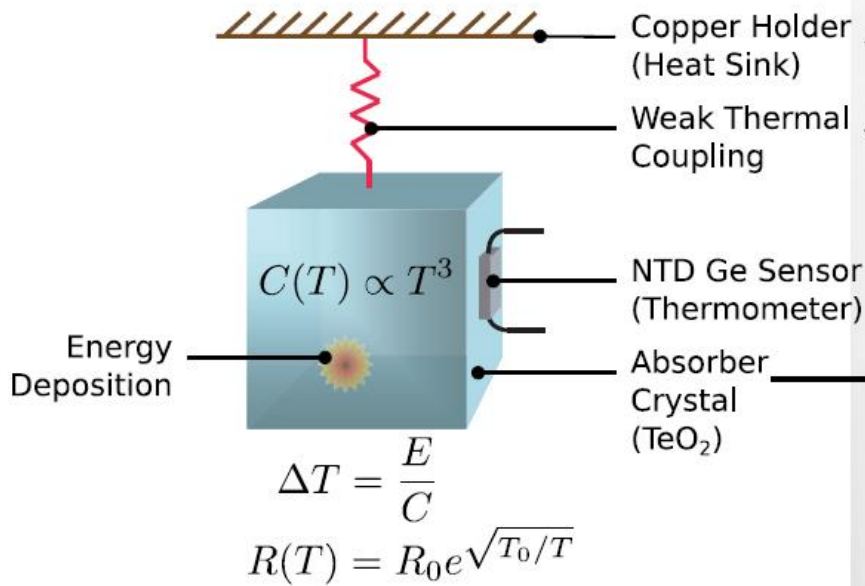
LEGEND (a new collaboration merging the expertise of GERDA and Majorana):

- up to 200 kg of  $^{76}\text{Ge}$  in the present LNGS cryostat. Assume a reduction of 5 in background with respect to GERDA
- 1000 kg of  $^{76}\text{Ge}$  in a new installation if Ge is chosen by the US down-select process. Further background reduction (1/30 of GERDA to stay in “zero background mode”) is needed to fully exploit this facility.

See Talk of S. Schoenert

# CUORE (Gran Sasso, Italy)

The largest bolometric experiment ever



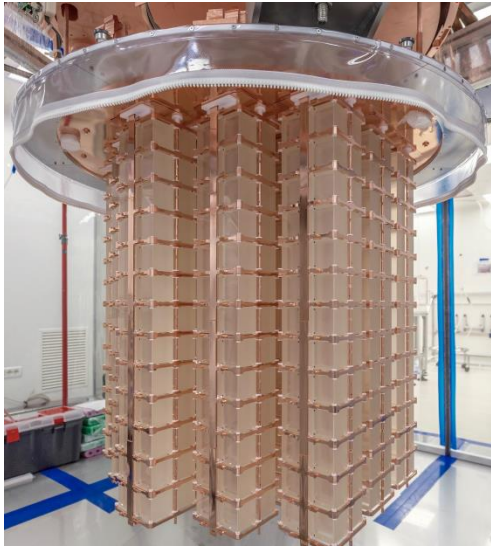
Isotope: <sup>130</sup>Te, Q-value = 2528 keV

Mass: 206 kg of <sup>130</sup>Te (742 kg of TeO<sub>2</sub>)  
Detectors: 988 TeO<sub>2</sub> bolometers (5x5x5 cm<sup>3</sup>)  
working at a **temperature of about 10 mK**

$T_{1/2} > 4 \sim 10^{24}$  y (90 % CL)  
(combined CUORE-0 + Cuoricino)

K. Alfonso et al., Phys. Rev. Lett. 115, 102502 (2015)

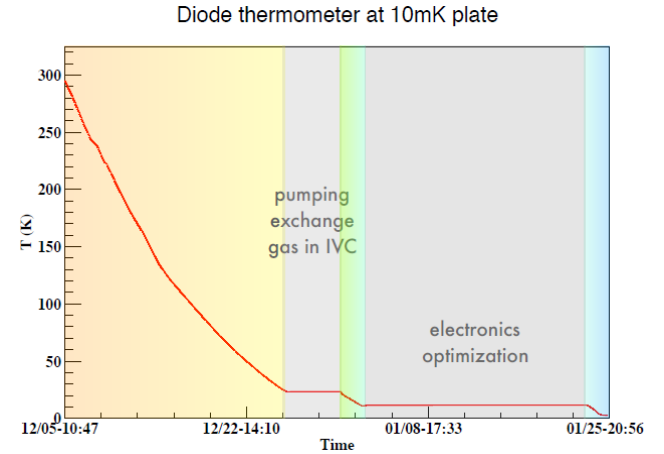
# A very special year for CUORE



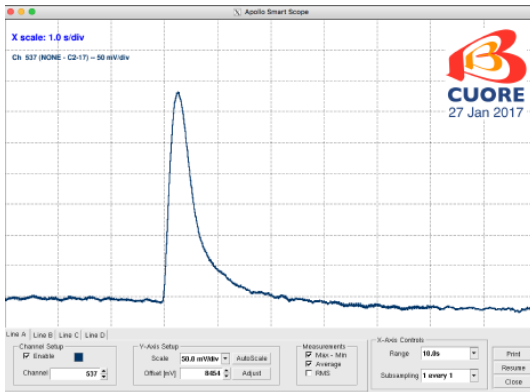
End of detector installation: Aug 2016



Shields, and electronics installed: Sep – Nov 2016



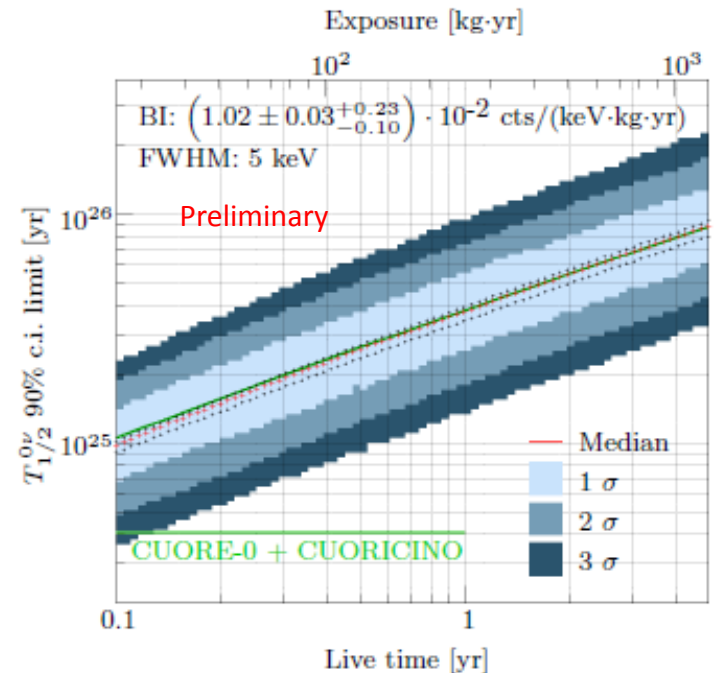
Detector cooled down to base temperature (7 mK)



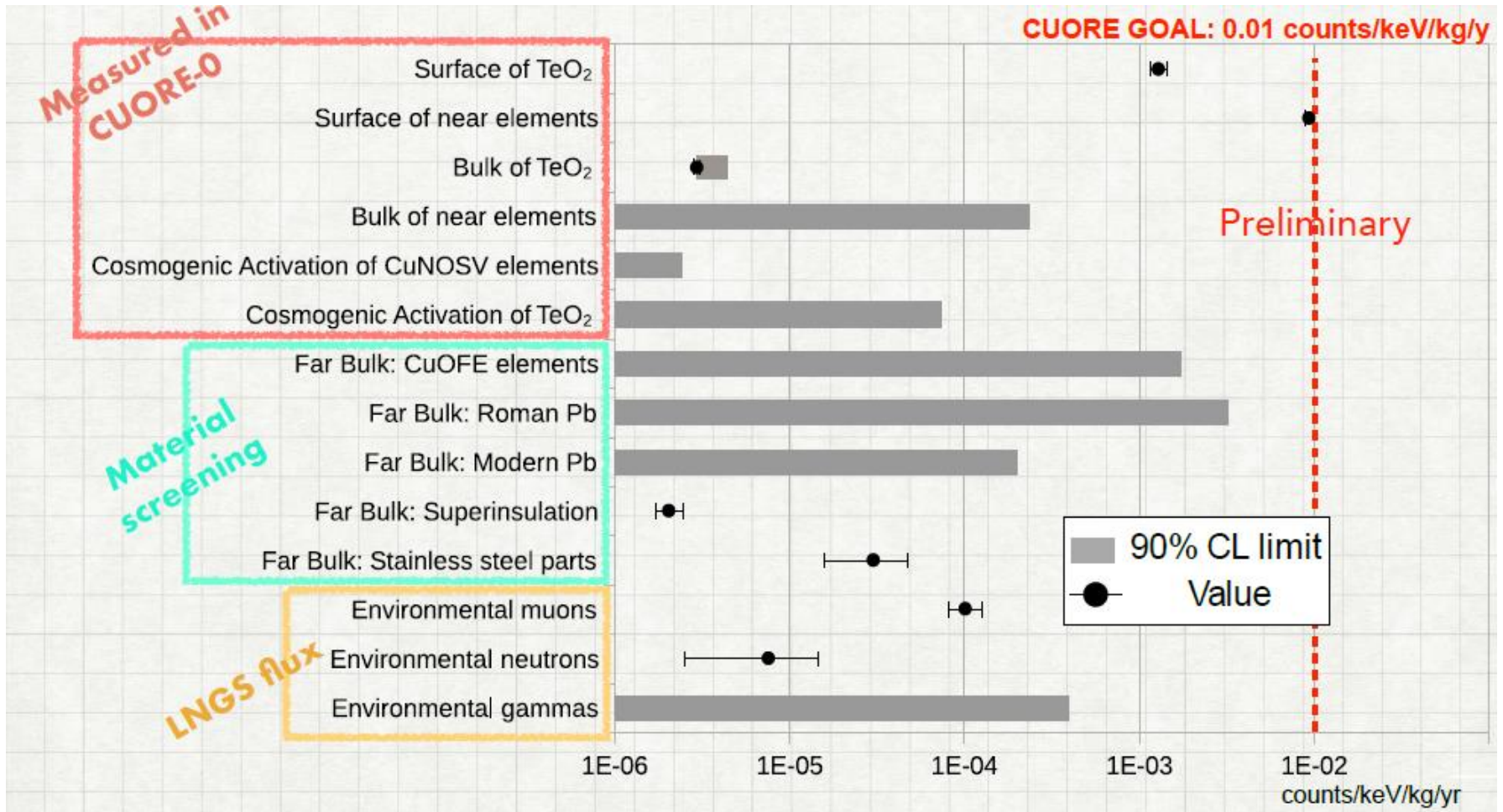
The first CUORE pulse (27 Jan)

**End of commissioning and start of data taking: April 2017**

Expected sensitivity:  
C. Alduino et al.  
arXiv:1705.10816



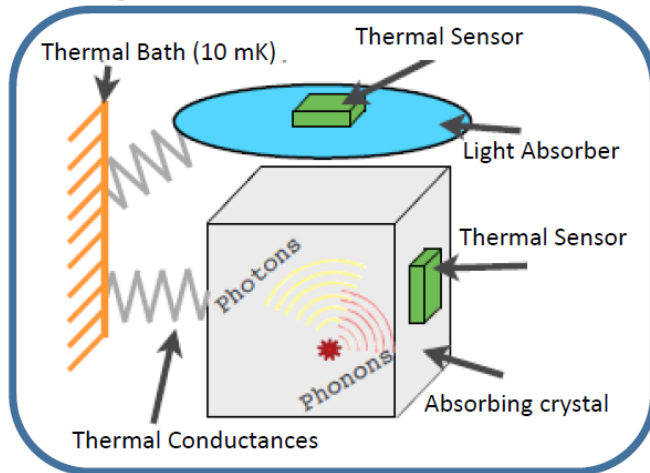
# Background budget of CUORE



Background rejection in CUORE is achieved by the measurement of the energy deposited in the crystals, pulse-shape analysis and detector coincidences but...

# Scalability of the bolometric technique

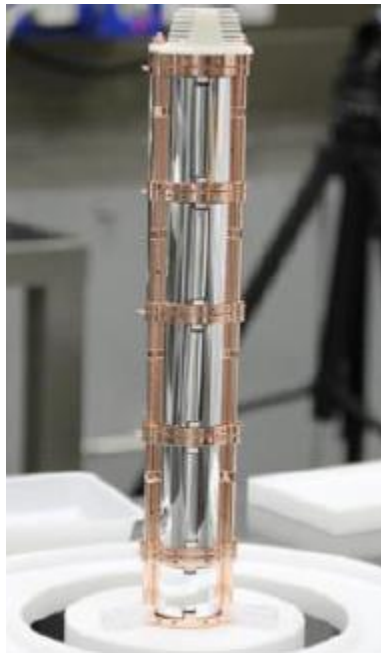
Bolometers offer additional opportunities to reach the  $10^{-4}$  counts/kg keV y regime



Observation of photons in scintillating bolometers (CUPID-0)

Observation of Cherenkov photons or surface effects in  $\text{TeO}_2$  crystals

G. Wang et al. [CUPID Coll.] arXiv:1504.03599



CUPID-0 ( $^{82}\text{Se}$ , SeZn bolometers)  
expected background

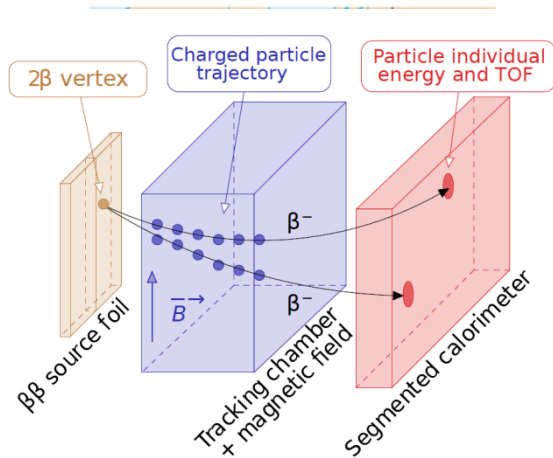
Eur. Phys. J. C76 (2016) 7, 364.

Background at $^{82}\text{Se}$ $Q_{\beta\beta}$ (counts/keV/kg/y)	
after $\alpha$ discrimination	$4 \times 10^{-3}$
coincidences rejection	$2.3 \times 10^{-3}$
$^{208}\text{Tl} - ^{212}\text{Bi}$ time delay rejection	$1 \times 10^{-3}$
+ cryostat $\gamma$ contamination	$< 1.5 \times 10^{-3}$



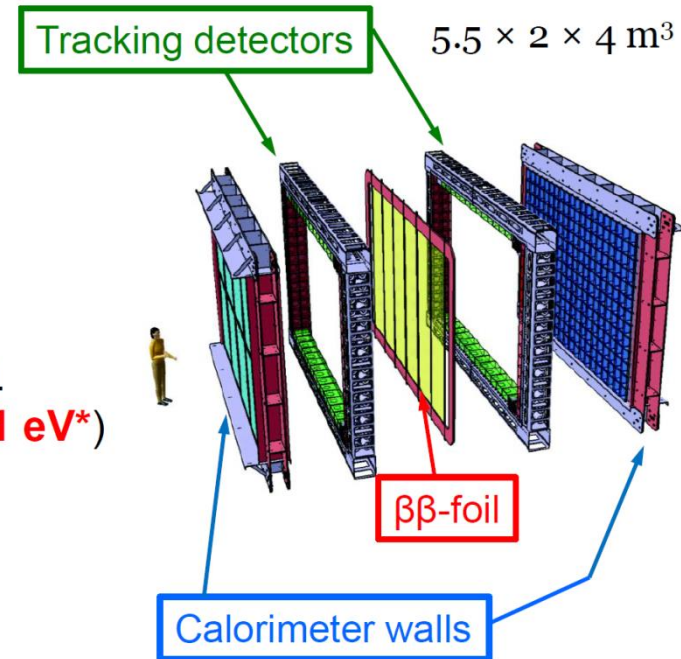
# SuperNEMO (Modane, France)

- Modular geometry (20 modules)
- Planned start: **2017**
- Placed in **LSM** (Modane, FRA)
- Studied isotope:  **$^{82}\text{Se}$**



7 kg of isotope (**100+ kg\***)  
 $0\nu\beta\beta$ :  $T_{1/2} > 6 \times 10^{24}$  yr ( **$10^{26}$  yr\***).  
 Limit  $m_{\beta\beta}$ : 0,2-0,4 eV (**0,04 -0,11 eV\***)

\* Full SuperNEMO design = 20 modules



Maximum flexibility in the choice of the isotope: **source outside detector** ( $^{82}\text{Se}$ ,  $^{150}\text{Nd}$ ,  $^{100}\text{Mo}$ ) and outstanding precision in the **topology reconstruction**.

It comes at the expenses of scalability and energy resolution

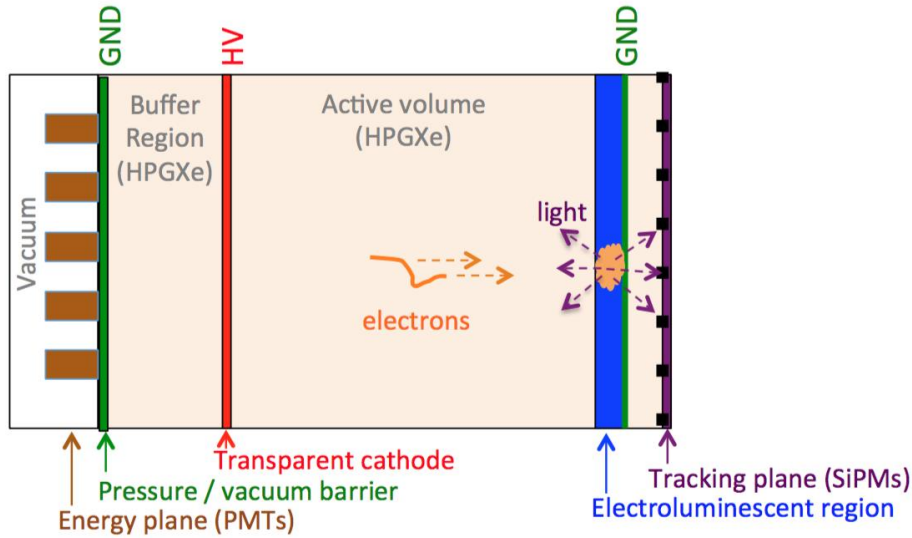
NEMO 2,3

Isotope	$T_{1/2}$ (yr)
$^{82}\text{Se}$	$2.1 \times 10^{23}$
$^{100}\text{Mo}$	$1.1 \times 10^{24}$
$^{116}\text{Cd}$	$1.6 \times 10^{22}$
$^{96}\text{Zr}$	$8.6 \times 10^{21}$
$^{150}\text{Nd}$	$1.8 \times 10^{22}$
$^{48}\text{Ca}$	$1.3 \times 10^{22}$

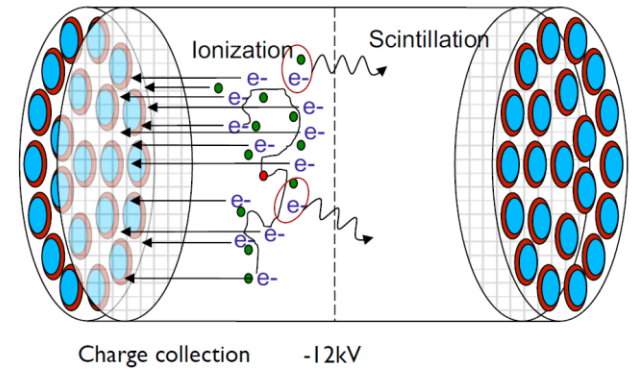
# NEXT (Canfranc, Spain) and (n)EXO (US)

It is fortunate that  $^{136}\text{Xe}$  is both a NLDBD candidate and a noble gas!

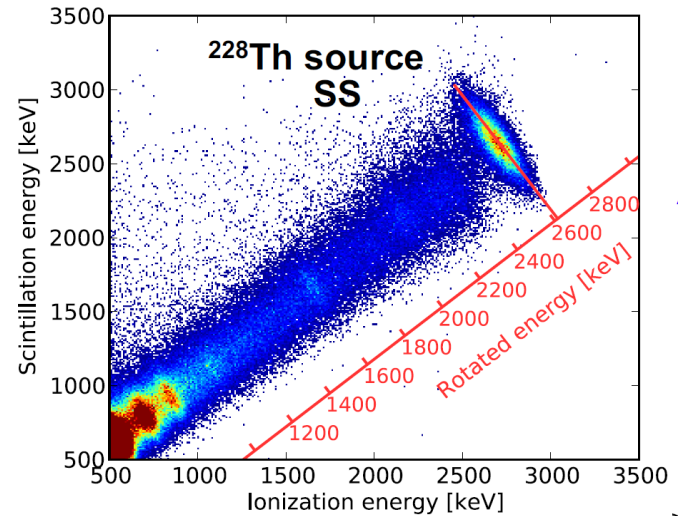
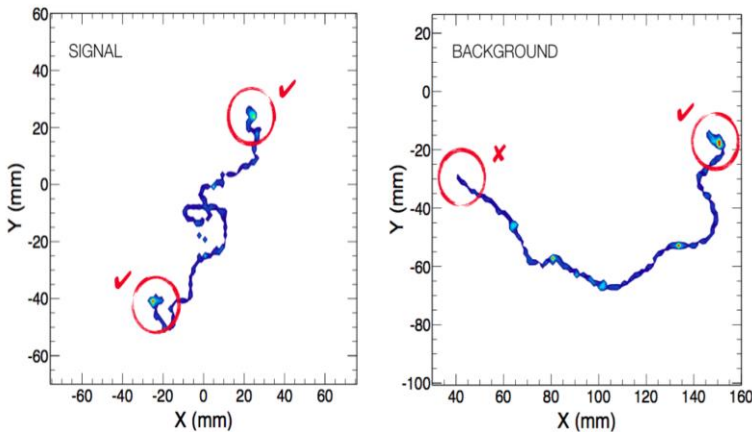
Isotope:  $^{136}\text{Xe}$ , Q-value = 2458 keV  
 $T_{1/2} = 1.1 \cdot 10^{25}$  y (90 % CL) (EXO-200)



## Liquid Xe TPC



## High pressure gas TPC

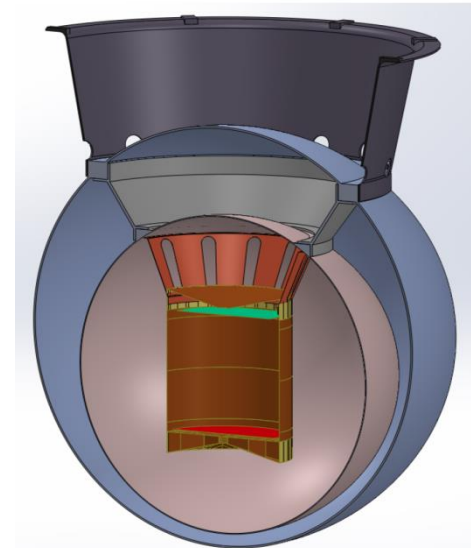
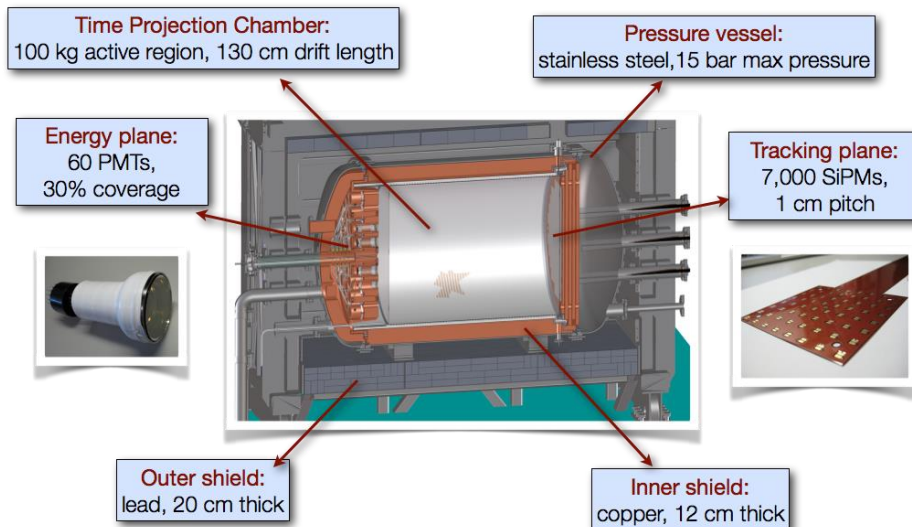


# Scalability of the Xe-TPC techniques

Both options have still large scalability potential using conventional techniques (**self shielding**, radio-purity assessment, background mitigation). Non conventional options (Ba tagging, diffusion reduction with dopants) under study.

nEXO (ton scale) at SNOLAB

NEXT-100 (100 kg scale)



Improvement	Reason
About 30× mass	Sensitivity to cover inverted hierarchy
Remove cathode from middle	Provide larger LXe-only volume in the middle with lower backgrounds.
6× high voltage	Required to enable a longer drift length
> 3× electron lifetime	Required to enable a longer drift length
Increased photo-coverage	Energy resolution (to 1% $\sigma/E$ ), lower scintillation threshold
SiPMs instead of APDs	Higher gain, lower bias, better energy resolution, lower scintillation threshold, less material, lower radioactivity.
In LXe front end electronics	Lower noise, fewer cables, lower threshold to ID Compton
Low outgassing materials	Longer electron lifetime from higher purity LXe
New calibration methods	Required to calibrate “deep” detector
Deeper site	Reduced cosmogenic activation
Charge tiles instead of wires	3mm position resolution, simpler/smaller mechanical supports, lower radioactivity

# Conclusions

Why **interest in NLDBD is growing** even if experimental searches started more than 70 years ago?

- ✓ Because the physics case became very strong after the discovery of neutrino oscillations
- ✓ Because the precise knowledge of mixing parameters drives our efforts toward the  $T_{1/2} = 10^{26} - 10^{28}$  y half life range

Are the experimental techniques **mature enough**?

- ✓  $T = 10^{26}$  y is the present of NLDBD and major results are coming at very high speed:

EXO-200 (Nature, 2014), Kamland-Zen (PRL 2016), Gerda (Nature 2017) .... easy to predict that others will come soon (CUORE, e.g., is in data taking)

- ✓ Most of the technologies are **scalable**. The most promising remains: liquid scintillators, germanium, bolometers, tracking (e.g. with TPC) but we should keep an eye on **ongoing R&D's** (not covered in this talk)

Can a PhD student make a **breakthrough** in this field?

- ✓ Yes. No clear path to reach the  $T = 10^{29}$  y, which remains very interesting from the point of view of parameter spaces.