

The Origin of Cosmic Baryons *and* Displaced Vertices at the LHC

Yanou Cui

UC Riverside

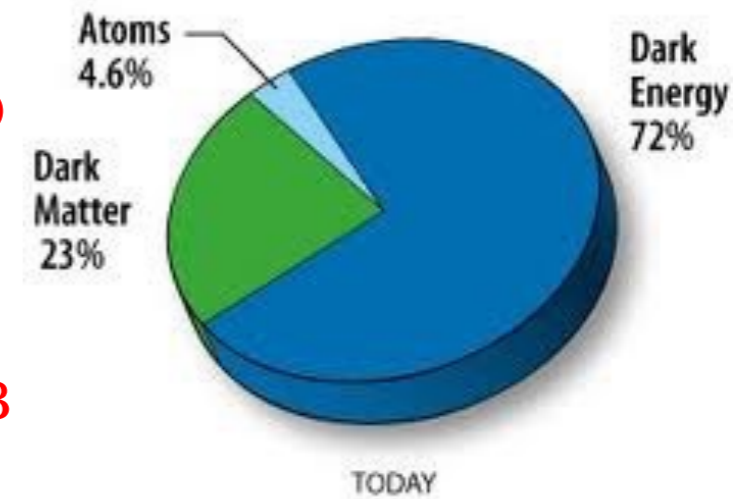


PITT PACC workshop
Feb 23 2017

Probing the Origin of Matter with the LHC?



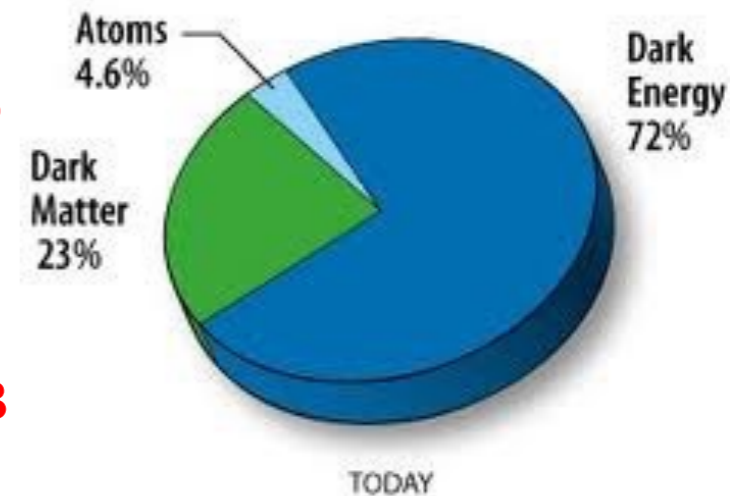
- **Baryon (atomic matter):** $\Omega_B \approx 4\%$
- **Dark Matter:** $\Omega_{DM} \approx 23\%$
- **Coincidence/Similarity:** $\Omega_{DM} \sim \Omega_B$



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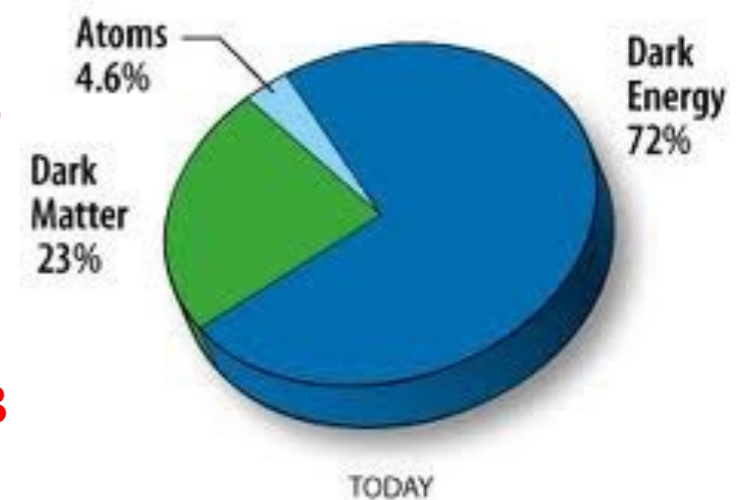


- Familiar/well-studied case: WIMP dark matter (Ω_{DM})
 - Stable, mass $\sim O(10-100)$ GeV, can be produced within $E_{LHC} = 14$ TeV
 - Pair produced (Z_2),
 - Invisible, MET + X

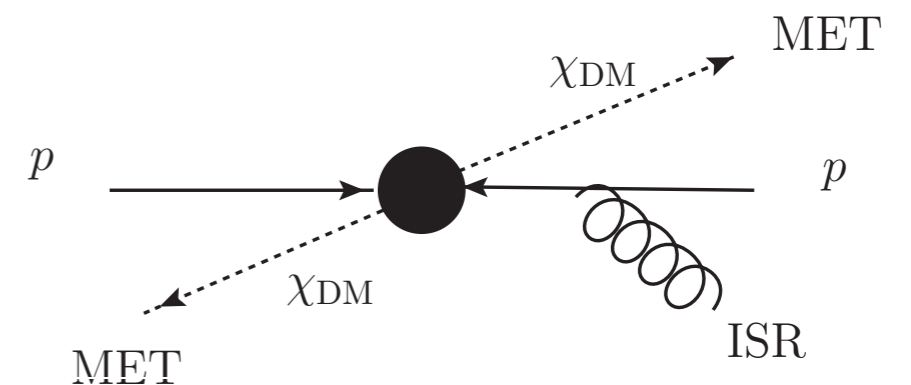
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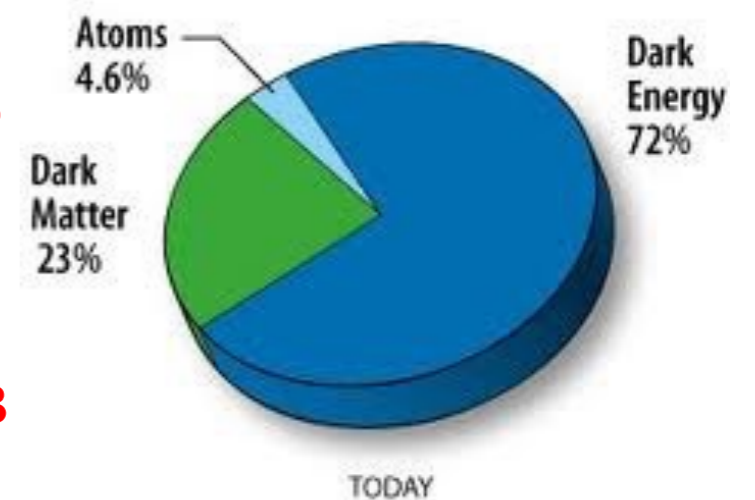
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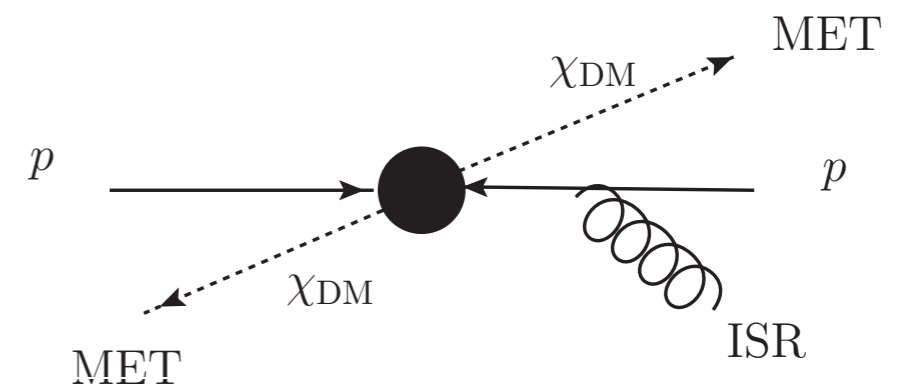
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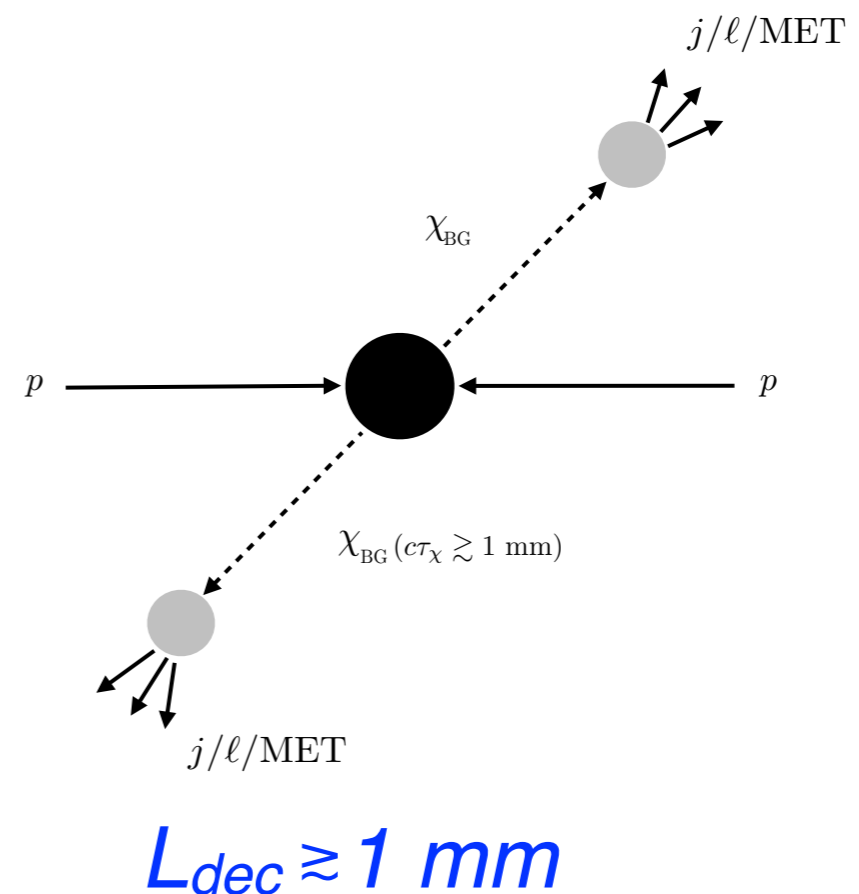
• Direct test for Ω_B , $\Omega_B \sim \Omega_{DM}$ @LHC?

(challenges: high mass, high T (EWBG) involved)

Indirect tests for EWBG: see David Morrissey's talk

Probing the Cosmic Origin of Baryons with Displaced Vertices at the LHC

- **New opportunity: baryogenesis**
(address Ω_B , possibly $+\Omega_B \sim \Omega_{DM}$)
 - New weak scale metastable particle (e.g. long-lived WIMP) as baryon parent
 - Pair produced (approx. Z_2)
 - **Displaced decay** to $j/\ell/MET$ by cosmological conditions!
Generic event topology
(analogy to WIMP DM search!)



Displaced Vertices at the LHC


- Nearly all SM particles decay **promptly**
 $\approx 100 \mu\text{m} - 1 \text{ mm}$ (= *prompt*)
- Ubiquitous predictions from motivated new physics:
long-lived particles, **displaced decay vertices** from all part
of the detector ($L_{dec} \gtrsim 1 \text{ mm}$) (SUSY, twin-Higgs, hidden valley, sterile ν ...)

 **◆ Spectacular signal!** low SM background, sensitive to
rare signal events 

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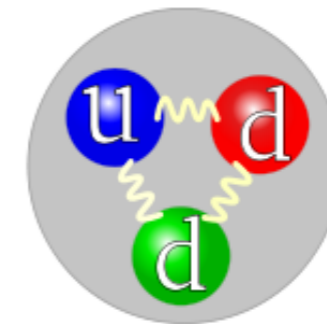
**Impressive developments, dedicated studies in the
past a few years (experimentalists + theorists)!**

Baryogenesis 101

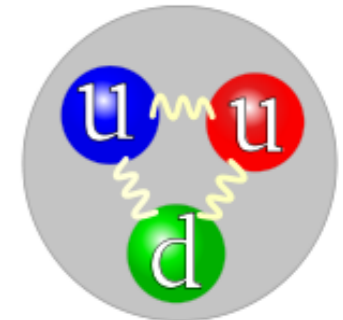
Baryon $\Omega_B \approx 5\%$

— The Unknown Aspects of the Known

- **Baryon**: proton, neutron \Rightarrow atoms, stars, ourselves!
- **Where does Ω_B come from?**
= **Where do we ourselves come from?**



NEUTRON
Quark structure



PROTON
Quark structure

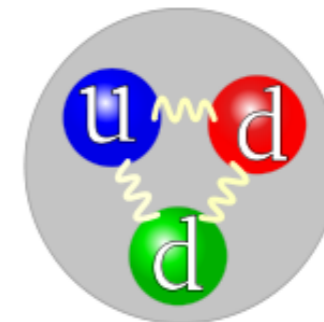
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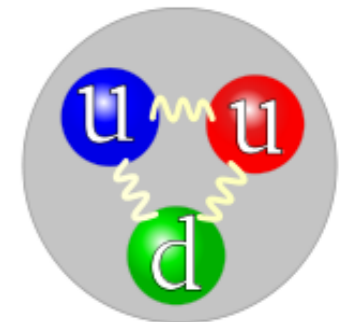
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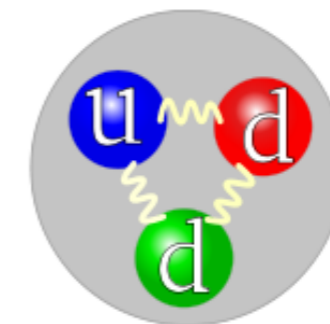
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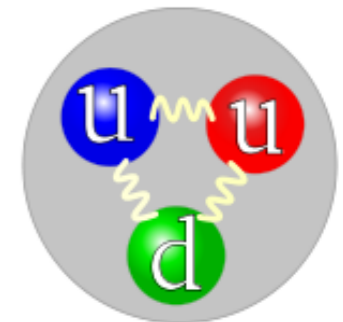
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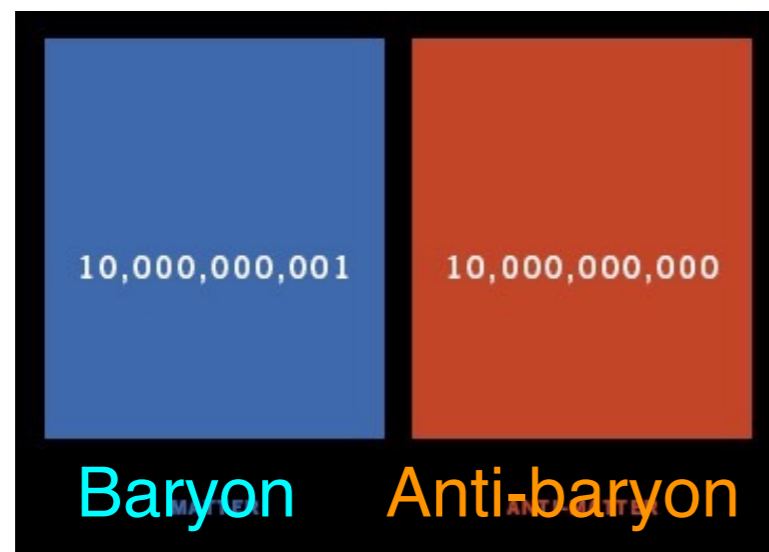
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Initial $B - \bar{B}$ asymmetry

$$\eta_B = (n_B - n_{\bar{B}})/n_\gamma \sim 10^{-10}$$



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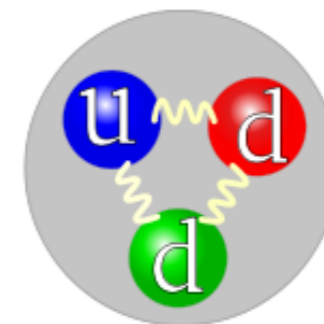
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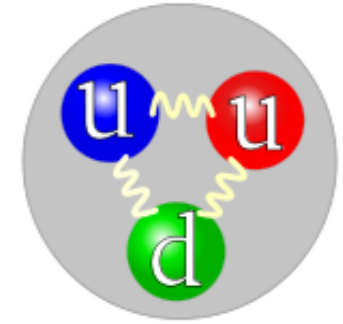
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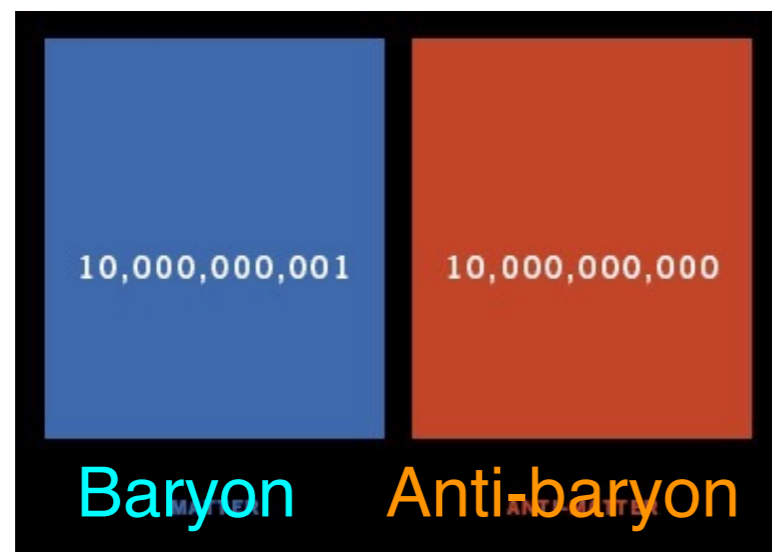
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symmetric
annihilation



Baryogenesis

- the Origin of the Baryon Asymmetry

The Universe starts with $B = 0$, \rightarrow ? $B \neq 0$

$B - \bar{B}$ asymmetry



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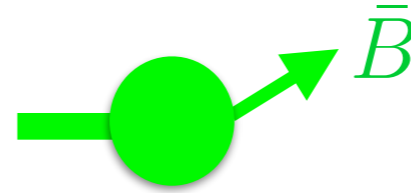
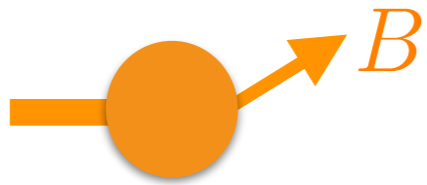
Sakharov Conditions (1967):



Baryogenesis

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- Require departure from equilibrium!

Thermal equilibrium + CPT symmetry

→ $n_B^{\text{eq}} = n_{\bar{B}}^{\text{eq}}, \langle B \rangle_{\text{eq}} = 0$  = 

Baryogenesis

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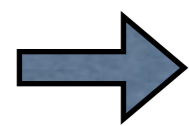


$\Omega_B \approx 5\%$:

Need beyond the
Standard Model
Particle Physics!

- Require departure from equilibrium!

Thermal equilibrium + CPT symmetry



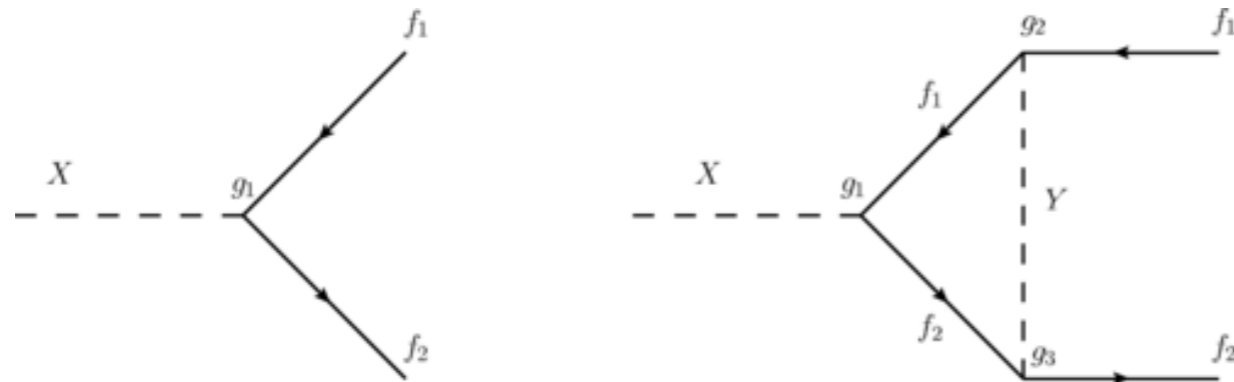
$$n_B^{\text{eq}} = n_{\bar{B}}^{\text{eq}}, \quad \langle B \rangle_{\text{eq}} = 0$$



Baryogenesis from Out-of-Equilibrium Decay

A general class of baryogenesis models (e.g. leptogenesis)

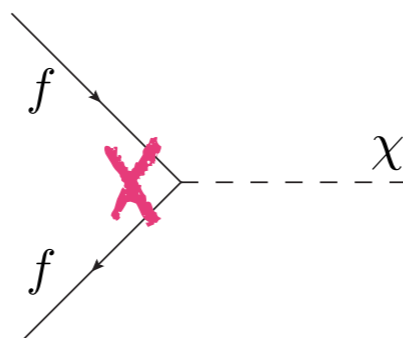
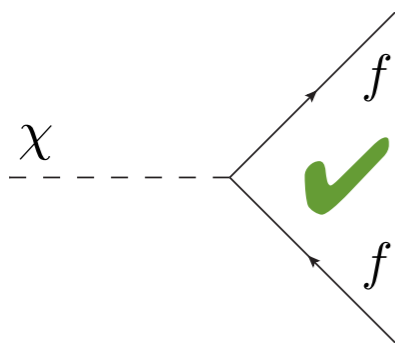
- Consider an unstable massive neutral particle χ
- Baryon asymmetry produced in its decay (**B-, C-, CP-violating**)



$$\Gamma(\chi \rightarrow f) \neq \Gamma(\chi \rightarrow \bar{f})$$

$$n_f - n_{\bar{f}} \neq 0$$

- Typically, the inverse processes efficiently erase the asymmetry
- But, if χ is **long-lived**, and **decays only after $T_f < M_\chi$** :



Inverse decay:
Boltzmann suppressed

$$e^{-M_\chi/T_{\text{decay}}}$$

Baryogenesis from Out-of-Equilibrium Decay

👉 **Out-of-equilibrium decay** → Sakharov conditions ✓

An intriguing observation (YC, Sundrum; YC, Shuve):

If χ has **weak scale** mass,

$$\Gamma_\chi < H(T = M_\chi) \quad \longleftrightarrow \quad c\tau_\chi \gtrsim \text{mm}$$

- A **generic connection** between **cosmological slow rates at $T \sim 100 \text{ GeV}$** and **displaced vertices at colliders!**

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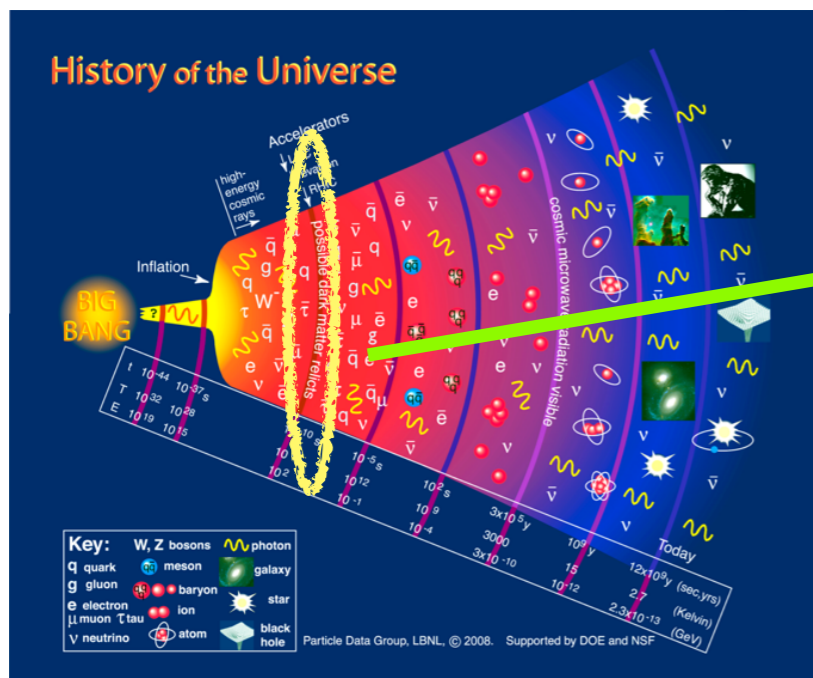
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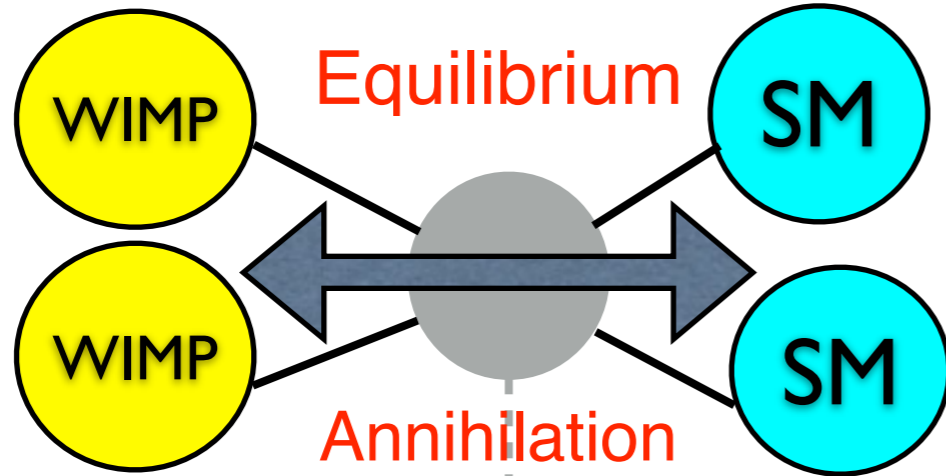
Our universe around EW phase transition was just slightly bigger than LHC tracking resolution!

Baryogenesis from WIMPs

— A New Proposal to Address $\Omega_B, \Omega_B \sim \Omega_{DM}$

- YC and Raman Sundrum, Phys.Rev. D87 (2013) 11
- YC, JHEP 1312 (2013) 067

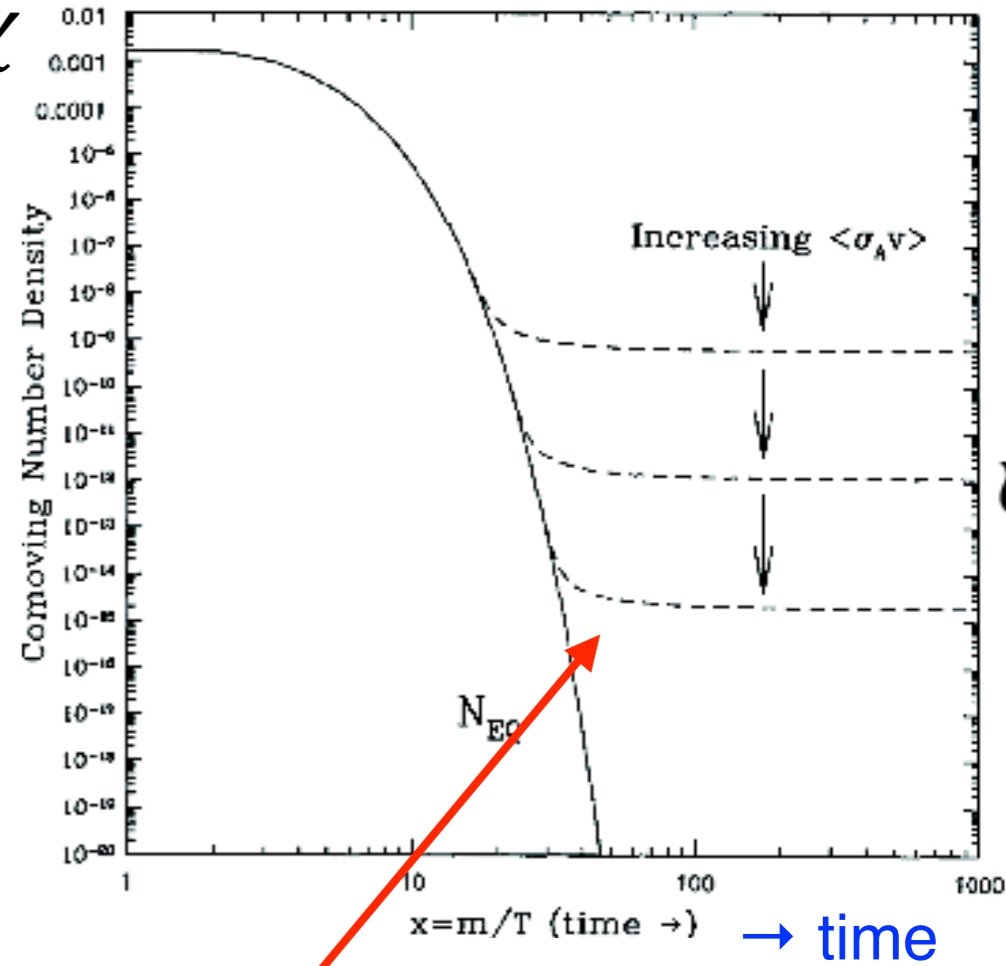
• **Cosmic Evolution of a stable WIMP χ**



Universe expands, cools, $T \downarrow$



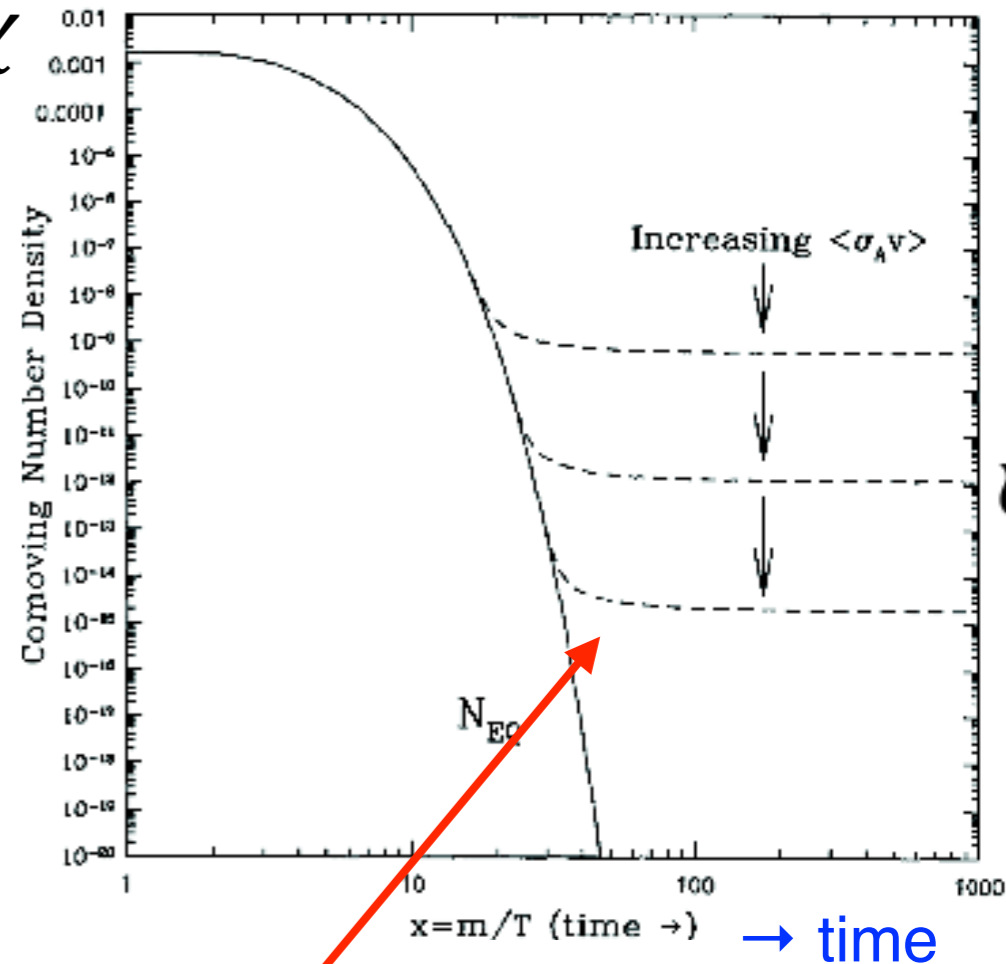
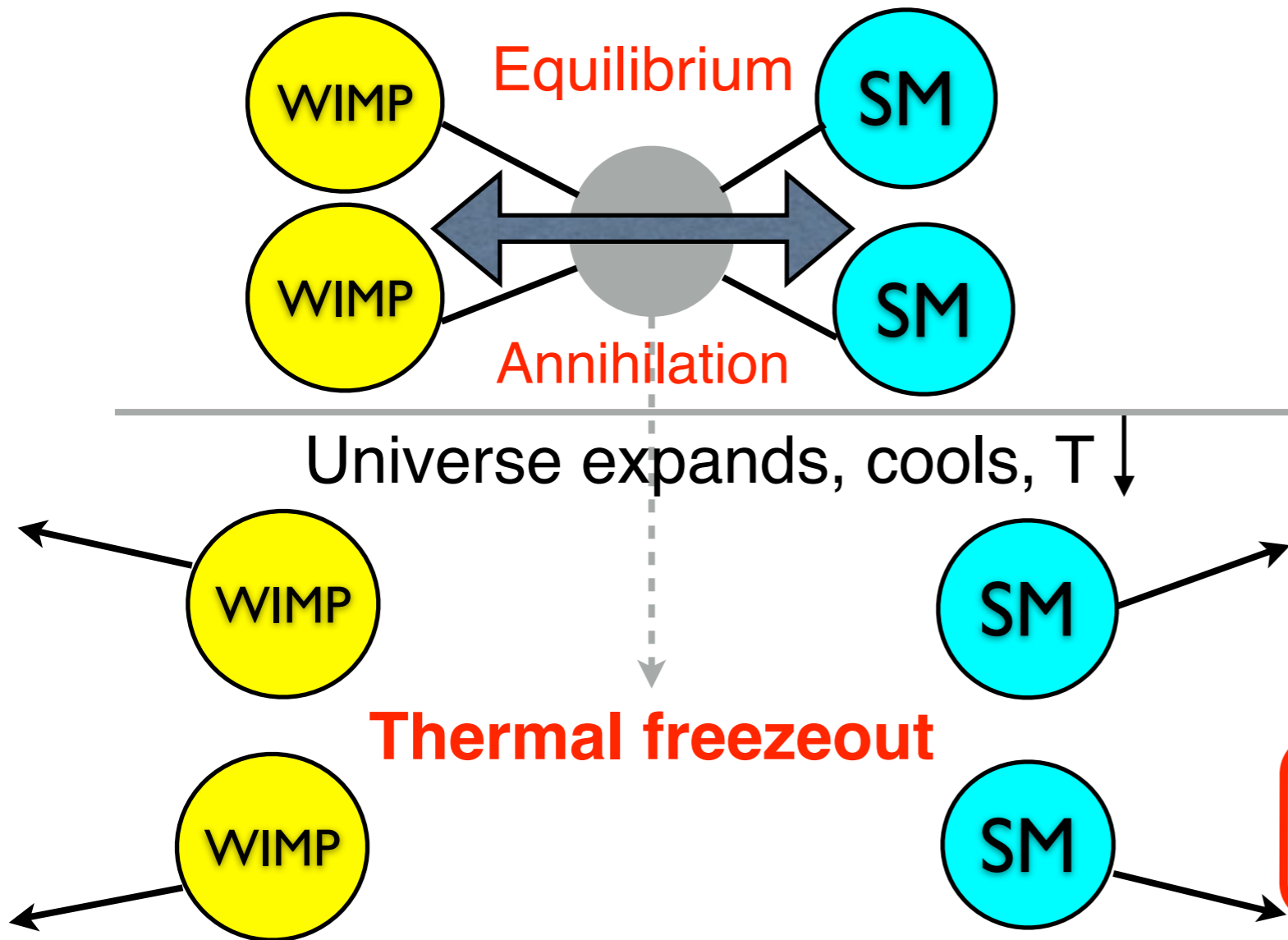
Thermal freezeout



**Departure from equilibrium:
key to Ω_{WIMP} !**

(recall Ω_B ?)

• Cosmic Evolution of a stable WIMP χ



**Departure from equilibrium:
key to Ω_{WIMP} !**

(recall Ω_B ?)

• Relic abundance:

$$\Omega_\chi \propto \langle \sigma_{ann} v \rangle^{-1}$$

$$\sim 0.1 \left(\frac{G_{Fermi}}{G_\chi} \right)^2 \left(\frac{M_{weak}}{m_\chi} \right)^2$$

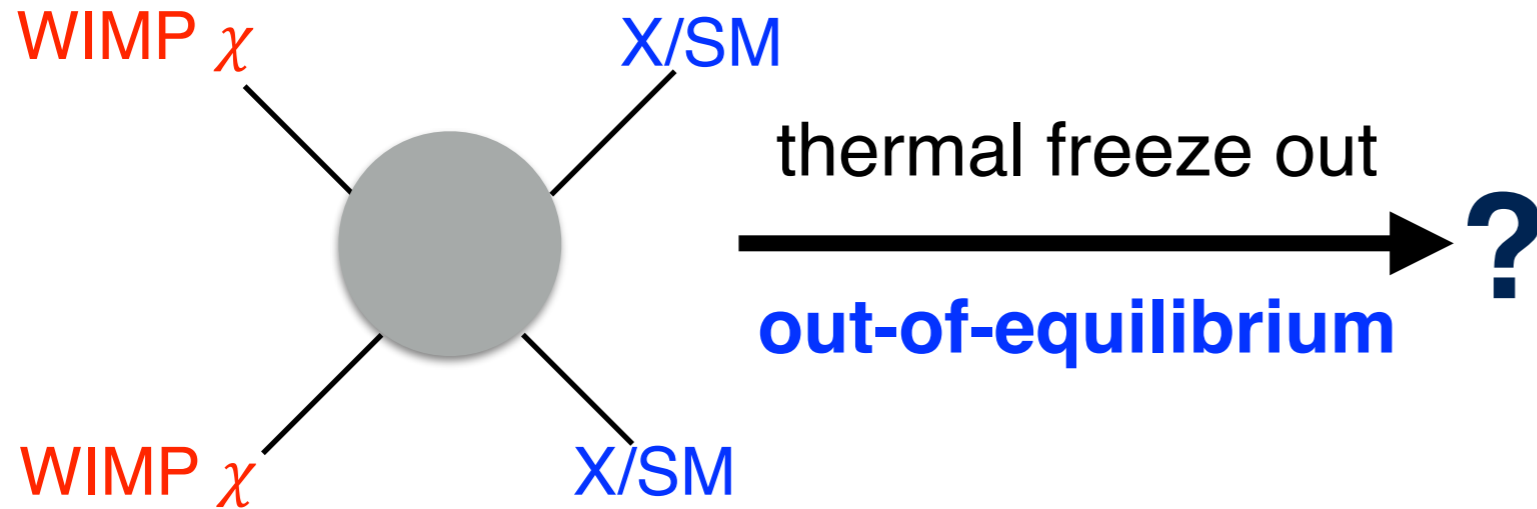


WIMP Miracle!

WIMP Miracle for Baryons?

- Another variation of WIMP miracle

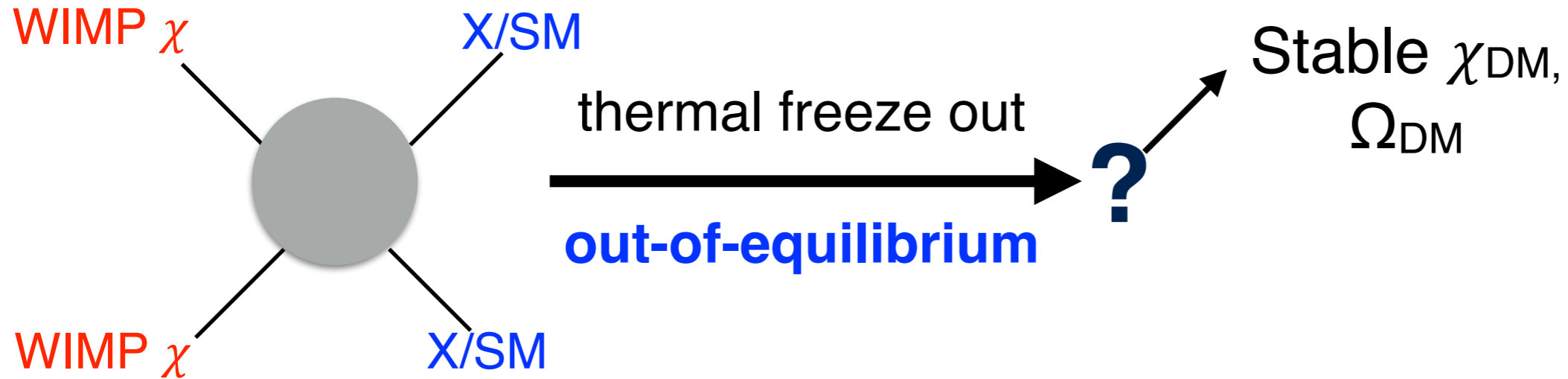
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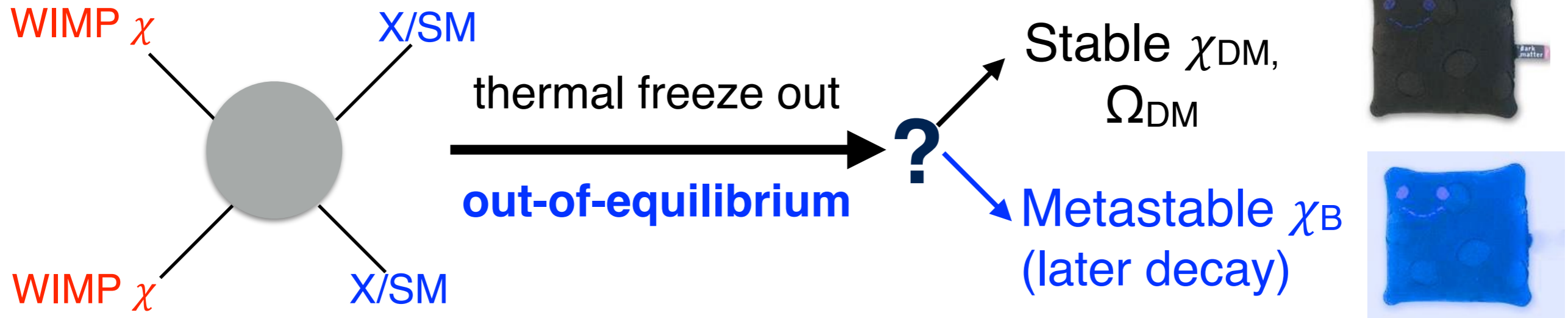
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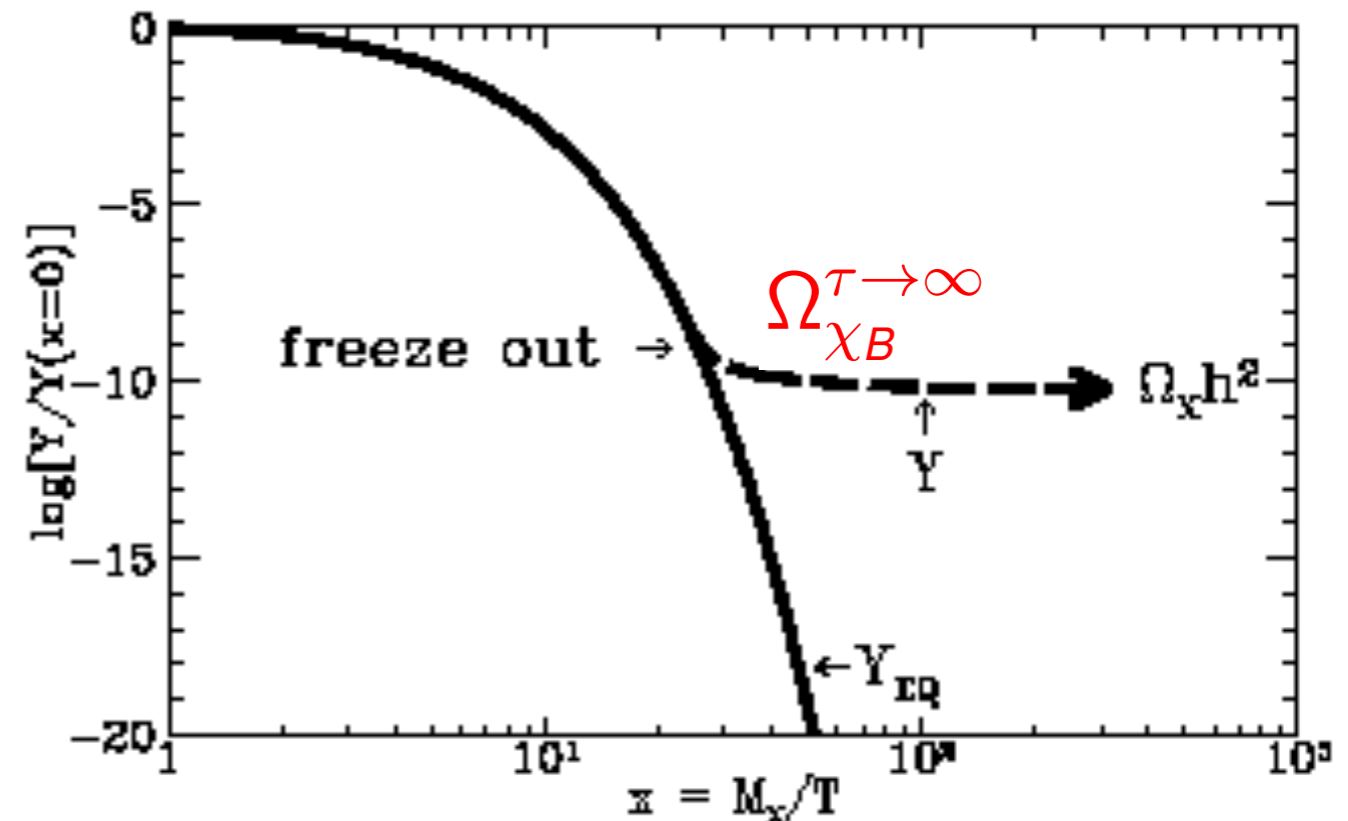
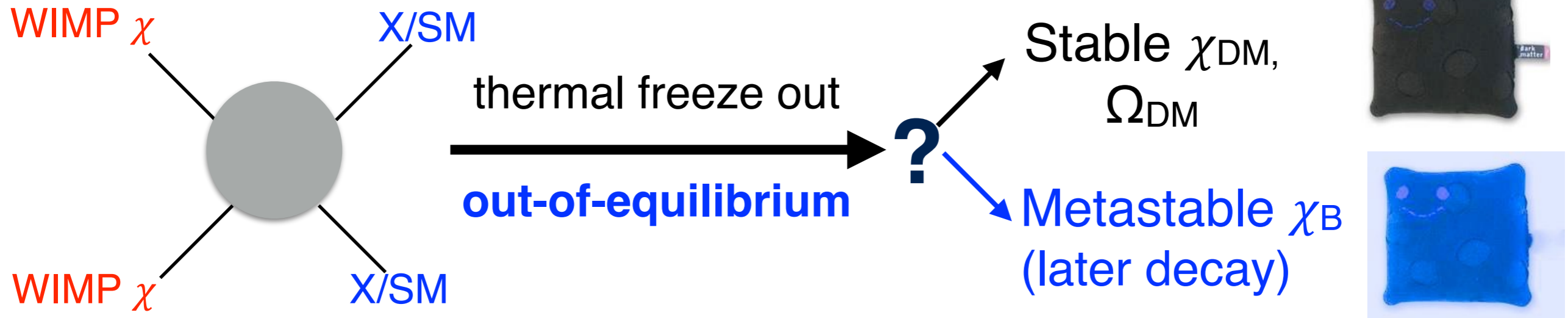
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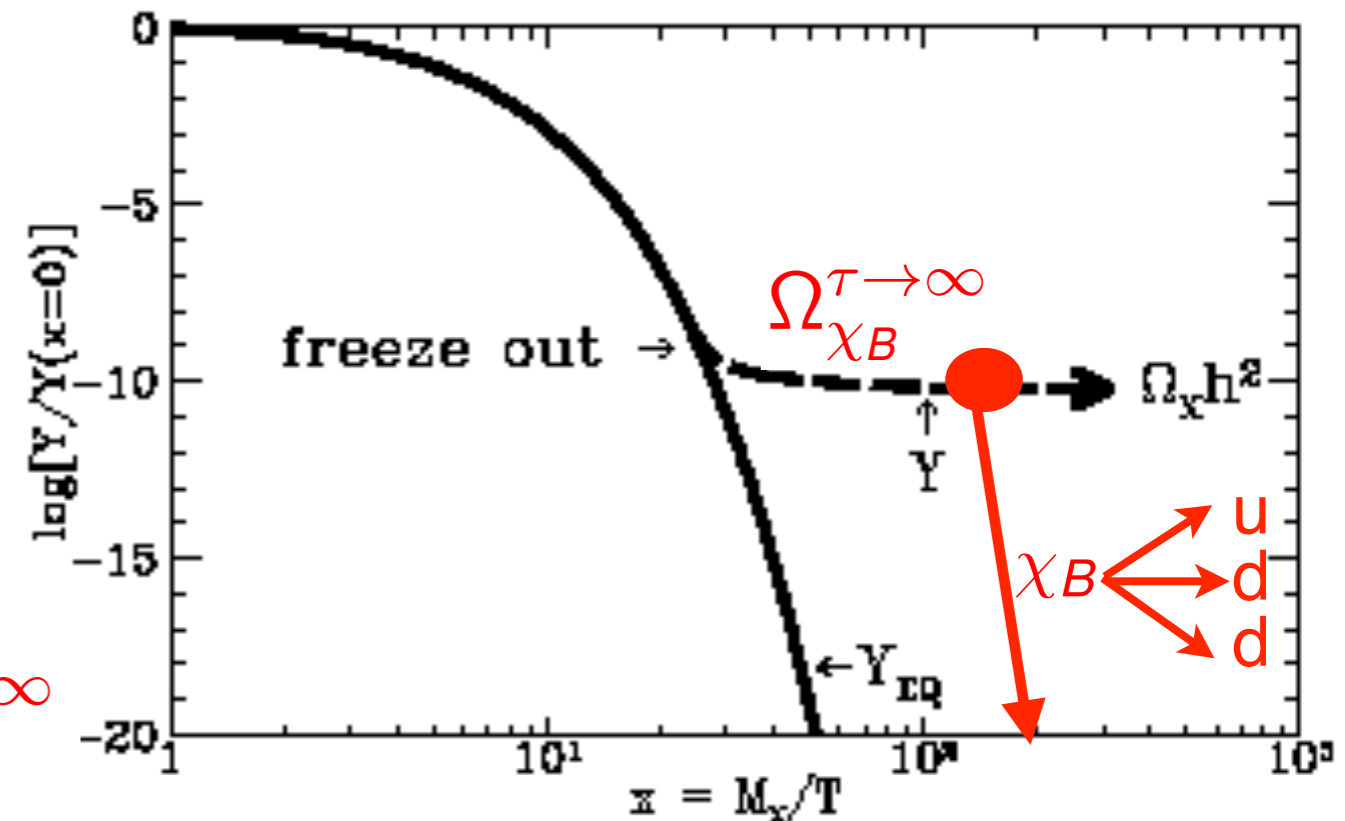
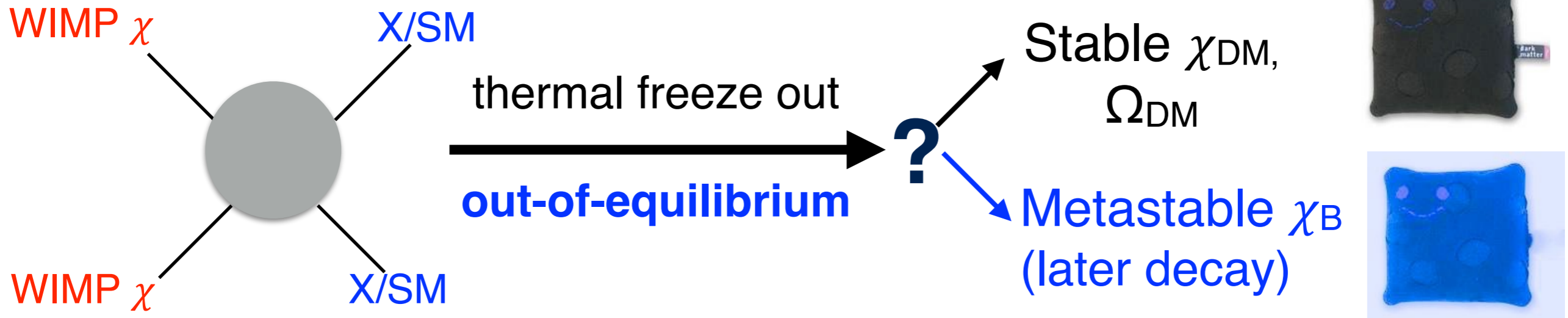
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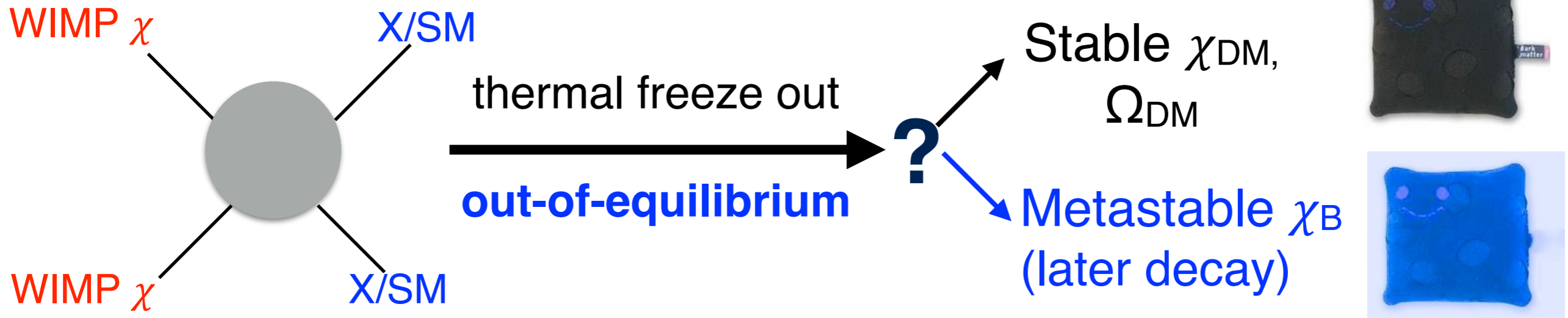


$$\Omega_B = \epsilon_{CP} \frac{m_p}{m_{\chi_B}} \Omega_{\chi_B}^{\tau \rightarrow \infty}$$

WIMP Miracle for Baryons?

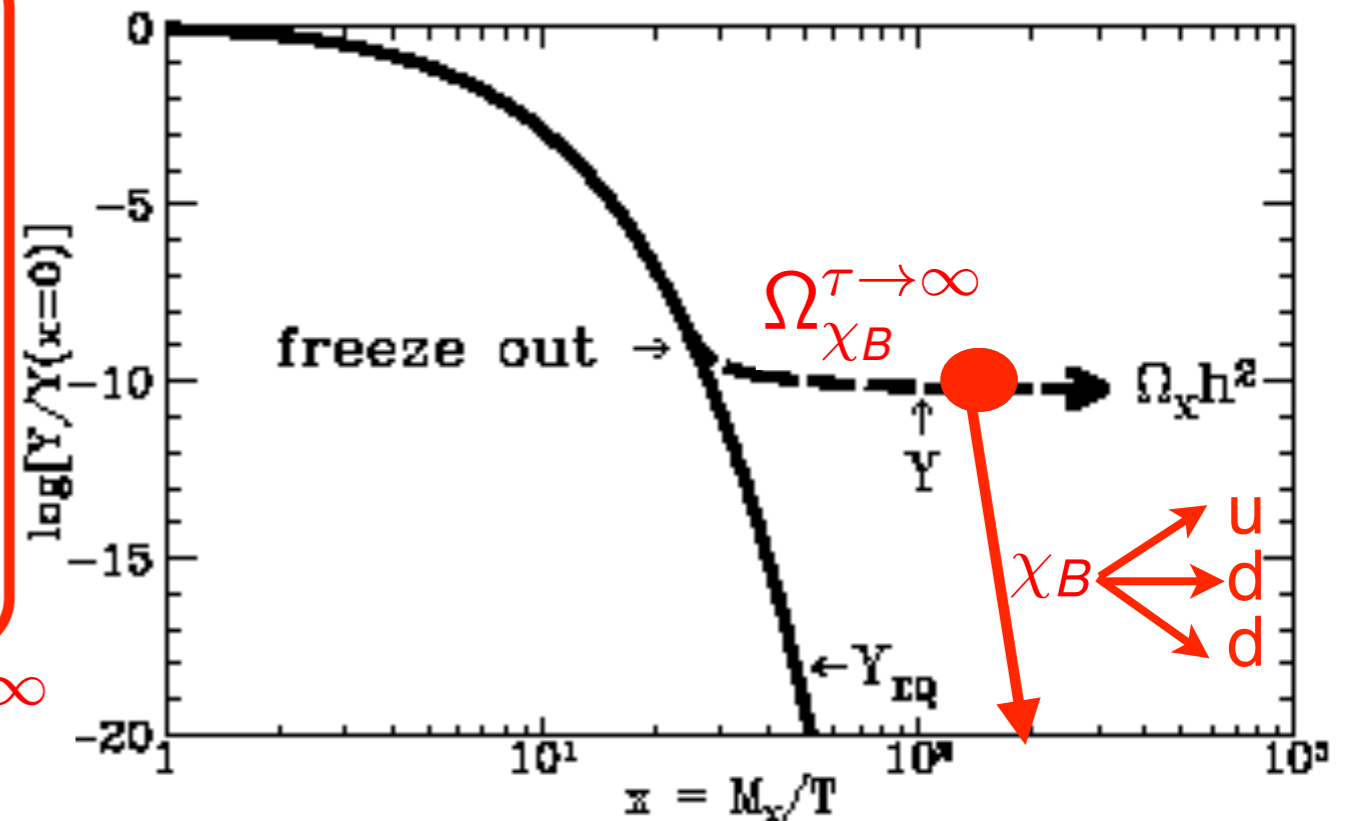
- Another variation of WIMP miracle

(YC, w/Sundrum)



- **Novel baryogenesis**
 Ω_B ✓
- **Generalized WIMP miracle**
 (+ stable WIMP DM Ω_{DM} ✓)
 $\Omega_B \sim \Omega_{DM}$ ✓

$$\Omega_B = \epsilon_{CP} \frac{m_p}{m_{\chi_B}} \Omega_{\chi_B}^{\tau \rightarrow \infty}$$



A Minimal Model Example

(Phys.Rev. D87 (2013) 11, **YC** w/Sundrum)

- BSM CP-, B-violating Lagrangian:

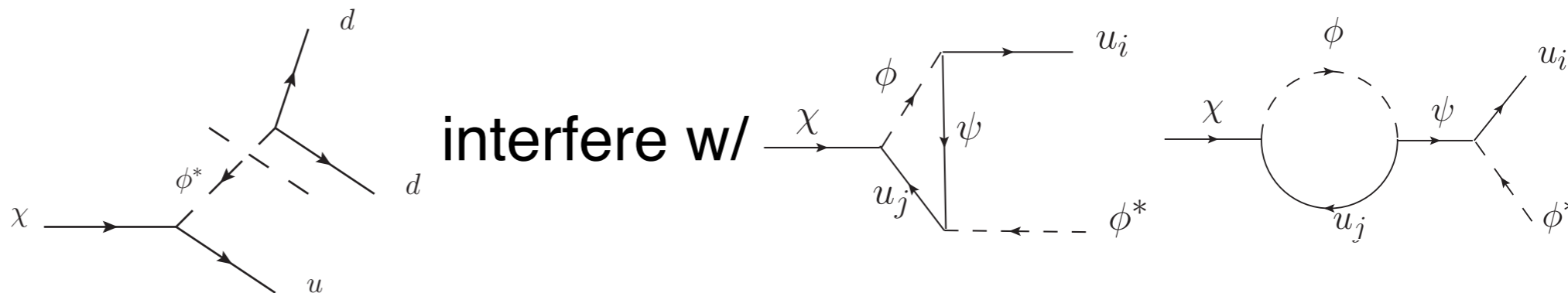
$$\begin{aligned} \Delta\mathcal{L} = & \lambda_{ij}\phi d_i d_j + \varepsilon_i \chi \bar{u}_i \phi + M_\chi^2 \chi^2 + y_i \psi \bar{u}_i \phi + M_\psi^2 \psi^2 \\ & + \alpha \chi^2 S + \beta |H|^2 S + M_S^2 S^2 + \text{h.c.} \end{aligned}$$

ϕ : di-quark scalar w/same charges as SM u-quark;

χ, ψ : SM singlet Majorana fermions;

ε : small breaking of a χ -parity \rightarrow long-lived χ (baryon parent)

CP asymmetry
from:



- **Easy embedding in RPV natural SUSY (+ singlet)! ($\phi \rightarrow \tilde{t}$)**
- **Late-time baryogenesis**; a **remedy** for a potential cosmological crisis with RPV SUSY: RPV washout of existing Ω_B (e.g. Barry, Graham, Rajendran 2013)

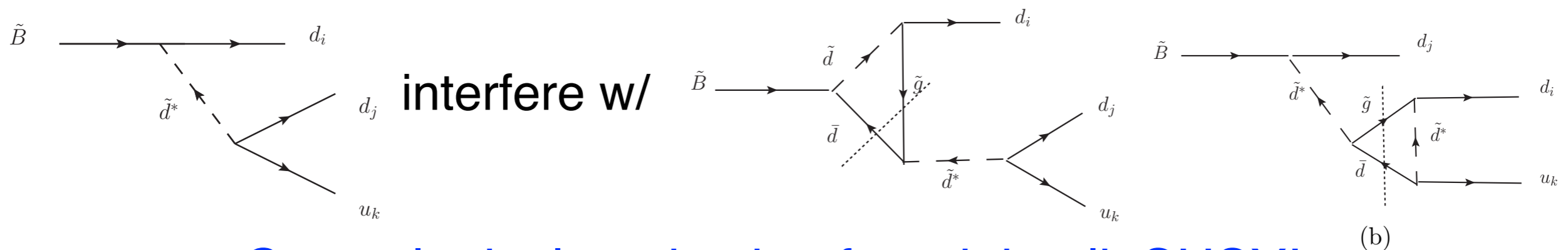
Embedding in Mini-split SUSY

(YC, JHEP 1312 (2013) 067)

Baryogenesis with **Minimal** SUSY model (MSSM+RPV)!

Bino $\tilde{B} \rightarrow \Delta B$!

- Sakharov#1: out-of equilibrium ✓
Split spectrum \rightarrow automatic late decay of gauginos
- Sakharov#2, 3: CP-, B-(L-)violations, rich sources in RPV SUSY ✓
- Bino as baryon parent (“would-be” over-abundance desirable for Ω_B)
- Nanopoulos-Weinberg theorem \rightarrow extra BV source in loop: \tilde{W} , \tilde{g} !



Cosmological motivation for mini-split SUSY!

Baryogenesis from Out-of-equilibrium Decays

— Collider Phenomenology

YC and Shuve, JHEP 1502 (2015) 049

(YC and Okui, Yunesi, Phys.Rev. D94 (2016))



★ *Strategies/results generally applicable to other new physics searches via displaced vertices*

Reproduce Baryogenesis at the LHC!

(YC w/Sundrum; w/Shuve)

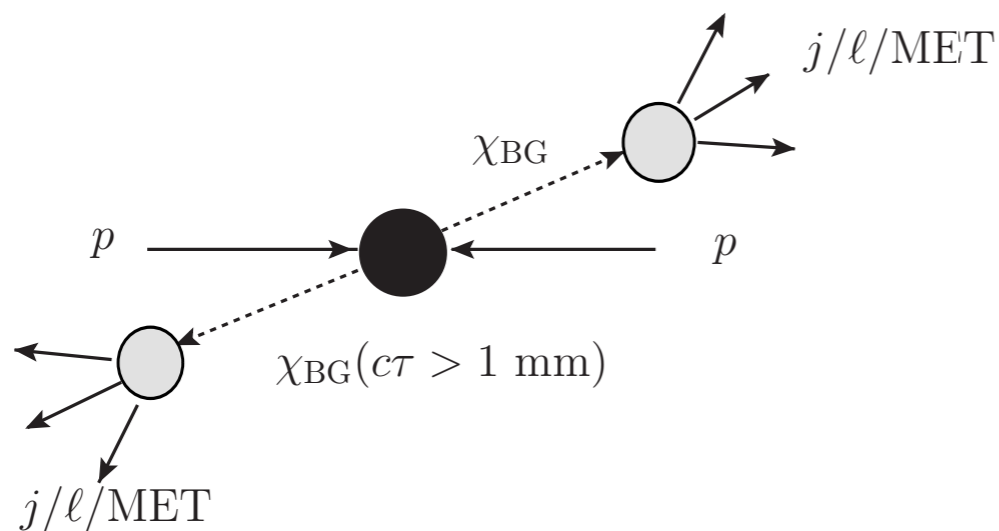
- WIMP $m_\chi \sim O(100 \text{ GeV})$ can be produced within $E_{\text{LHC}}=14 \text{ TeV}$!
- **Cosmological condition for baryogenesis:**

χ lives beyond its thermal freeze out time

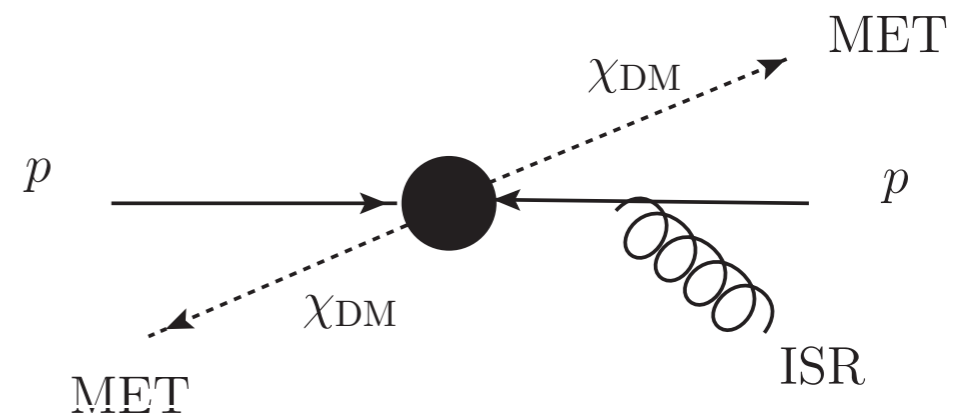
$$\Gamma_\chi < H_{\text{fo}} \longleftrightarrow c\tau_\chi \gtrsim \text{mm}$$

LHC tracking resolution!

- **Distinctive signal: displaced decay vertex inside detectors**
—not well-covered, low bkg search channel, rising interest!



Metastable WIMP baryon parent@LHC:
displaced vertex



Stable WIMP DM@LHC:
missing energy (analogy)

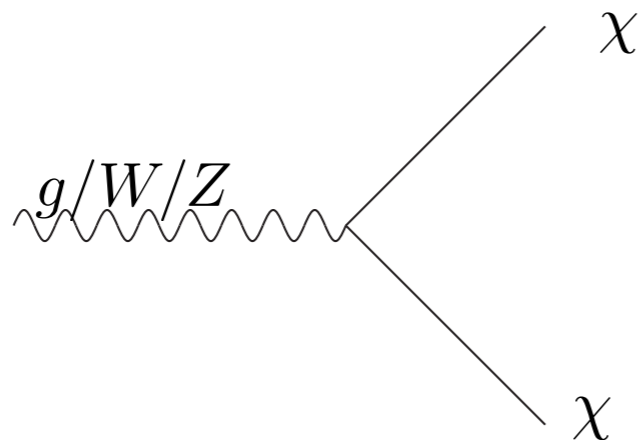
Simplified Model Approach for LHC Pheno

(YC and Shuve arxiv:1409.6729, JHEP)

- Classify production modes (analogy to DM search @LHC!)

Charged under SM gauge interactions:

wino/gluino-like (state in interference loop)



- Classify decay modes (unlike DM search)

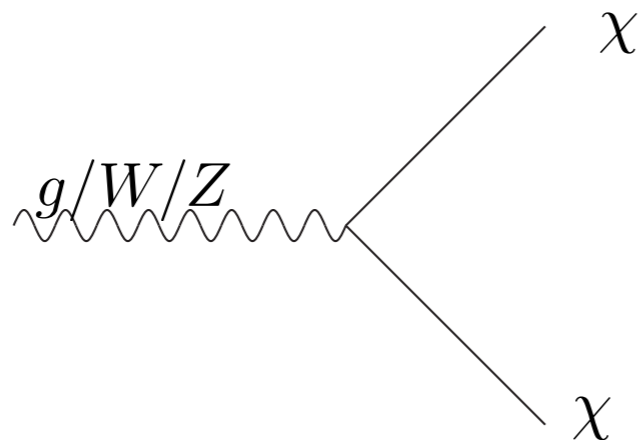
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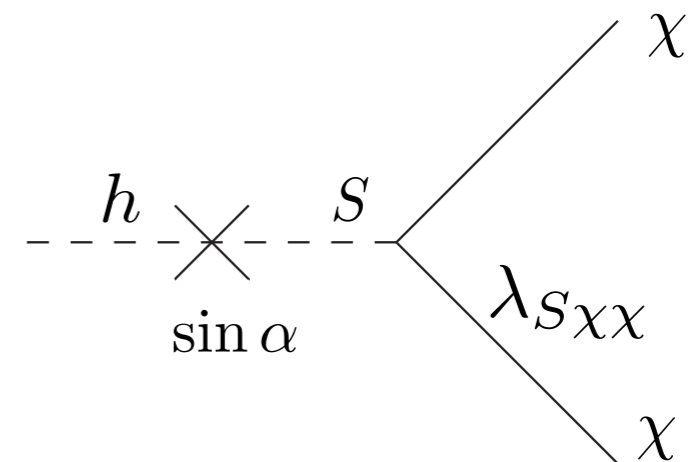
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Higgs portal:

singlet-like (e.g. $M_\chi = 150$ GeV)



- Classify decay modes (unlike DM search)

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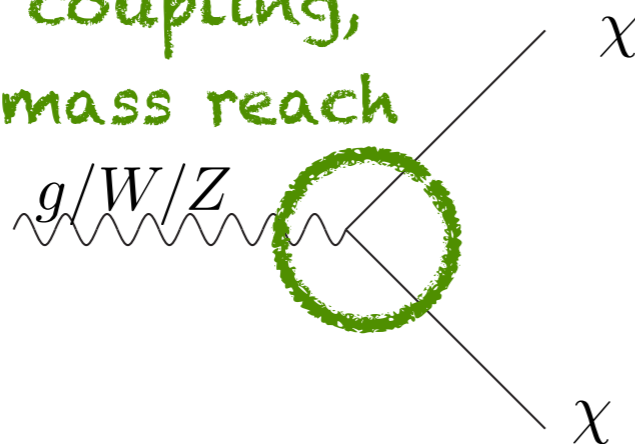
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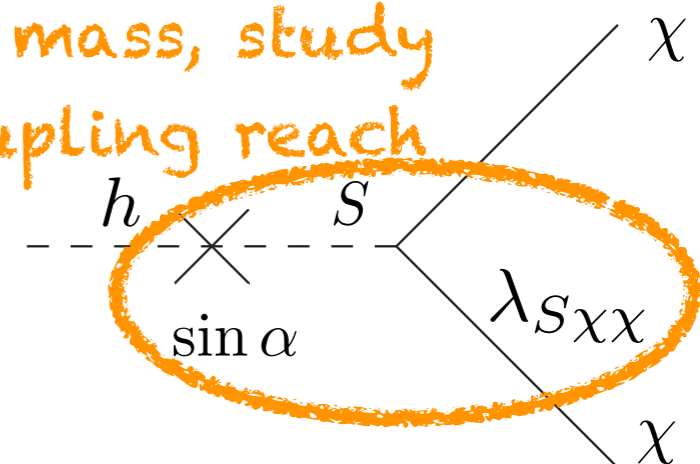
fixed coupling,
study mass reach



Higgs portal:

singlet-like (e.g. $M_\chi = 150$ GeV)

fix mass, study
coupling reach



- Classify decay modes (unlike DM search)

Simplified Model Approach for LHC Pheno

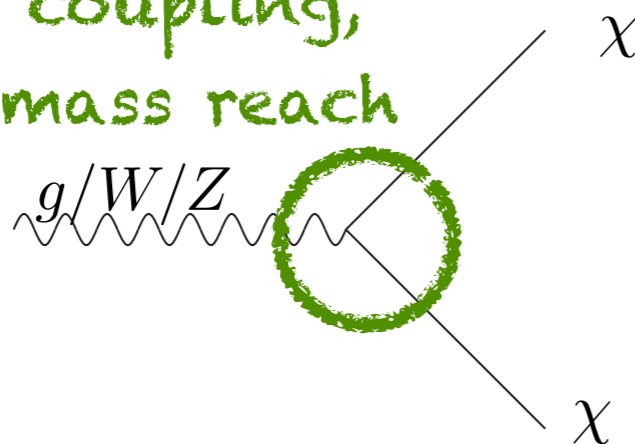
(YC and Shuve arxiv:1409.6729, JHEP)

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Charged under SM gauge interactions:

wino/gluino-like (state in interference loop)

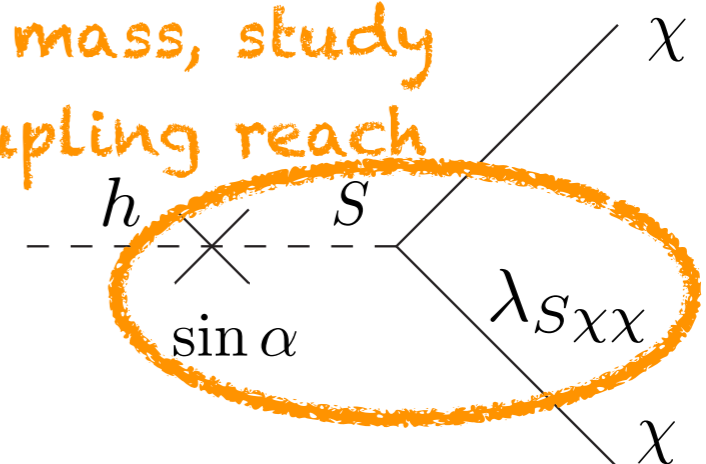
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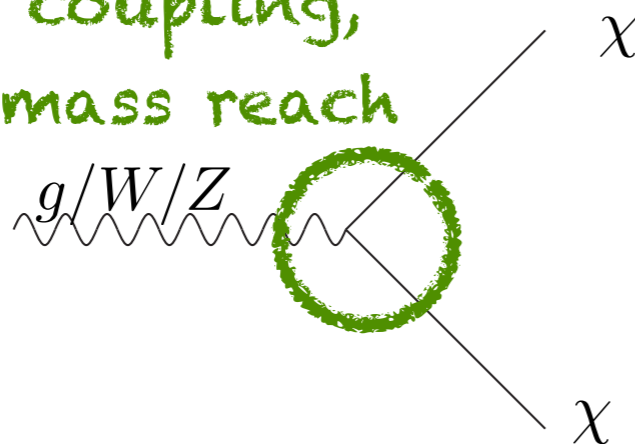
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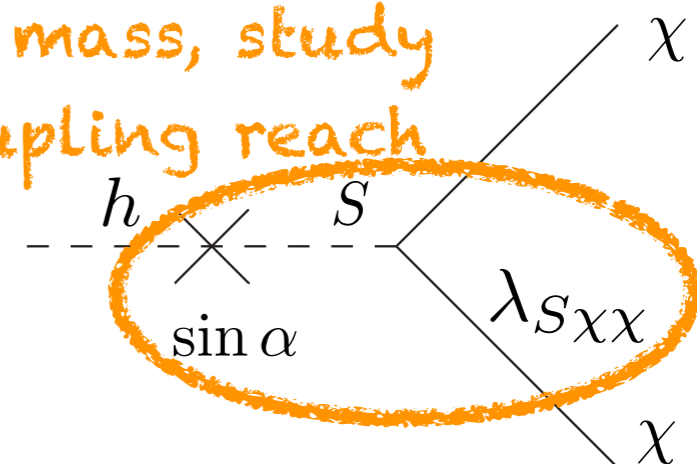
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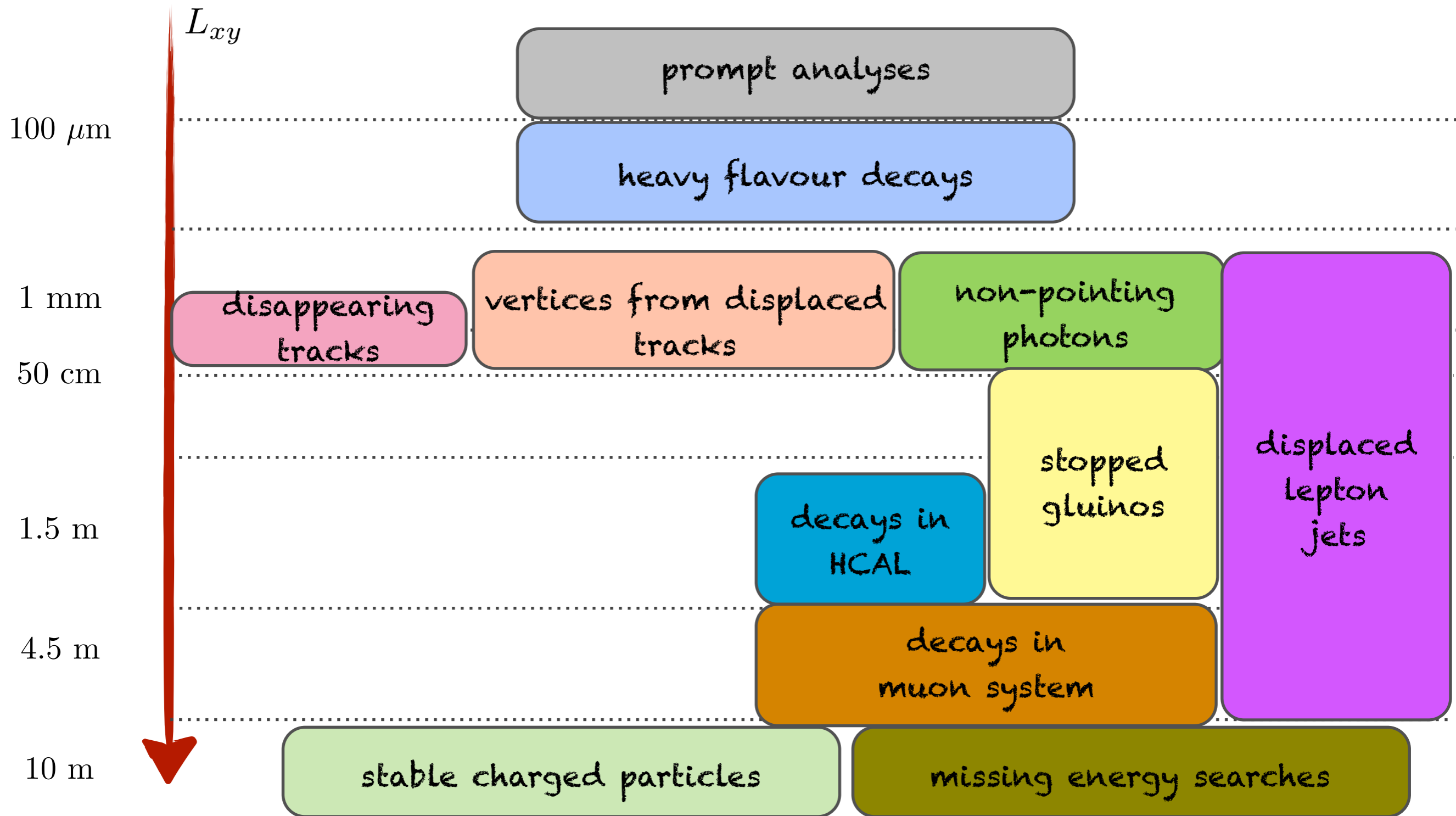
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Lepton number violating:

$$\chi \rightarrow L_i Q_j \bar{d}_k$$

$$\chi \rightarrow L_i L_j \bar{E}_k$$

LHC DV Search Possibilities



New detector: MATHUSLA? (Chou, Curtin, Lubatti +others)

Recast Existing LHC Searches

- Focus on displaced decay in tracking volume

Near lower bound $c\tau_\chi \gtrsim \text{mm}$, better sensitivity to wide lifetime range, easier to model with theorists' tools!

(decay in other parts of detector important too!)

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Baryon number violating:

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displaced jets (all-hadronic)

CMS, arXiv:1411.6530

Lepton number violating:

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displaced muon + tracks

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- Goal of our analysis:

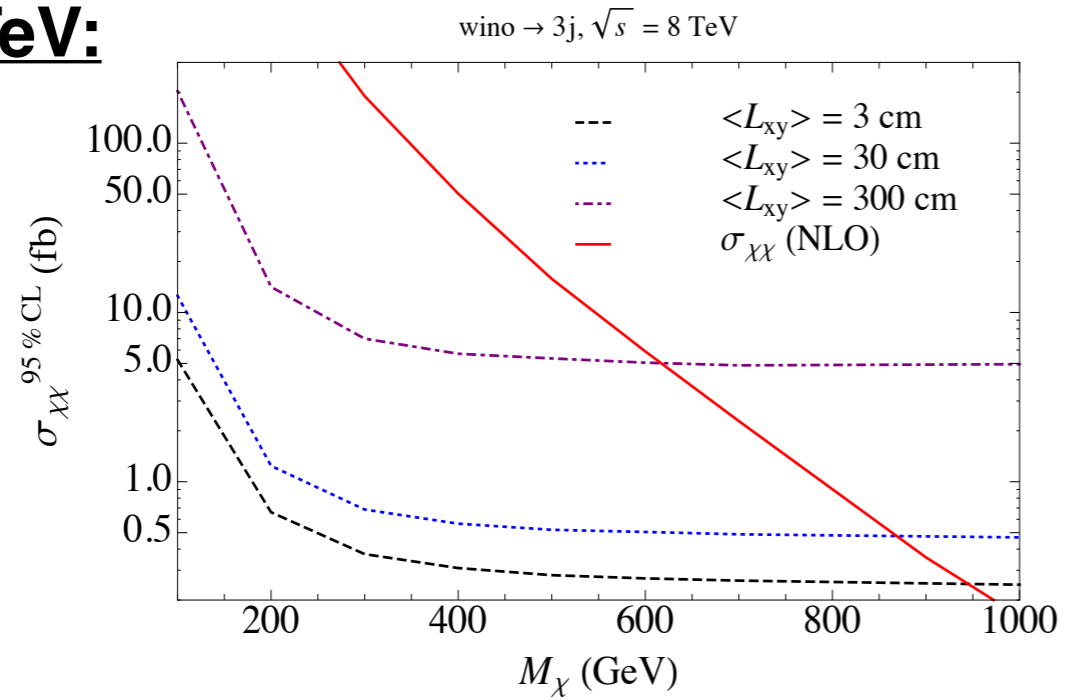
- What is the coverage for our simplified models based on benchmarks chosen by the collaborations?
- What advice can we provide for general experimental improvement?

Fully hadronic displaced vertices

CMS displaced dijet, arXiv:1411.6530

wino

8 TeV:

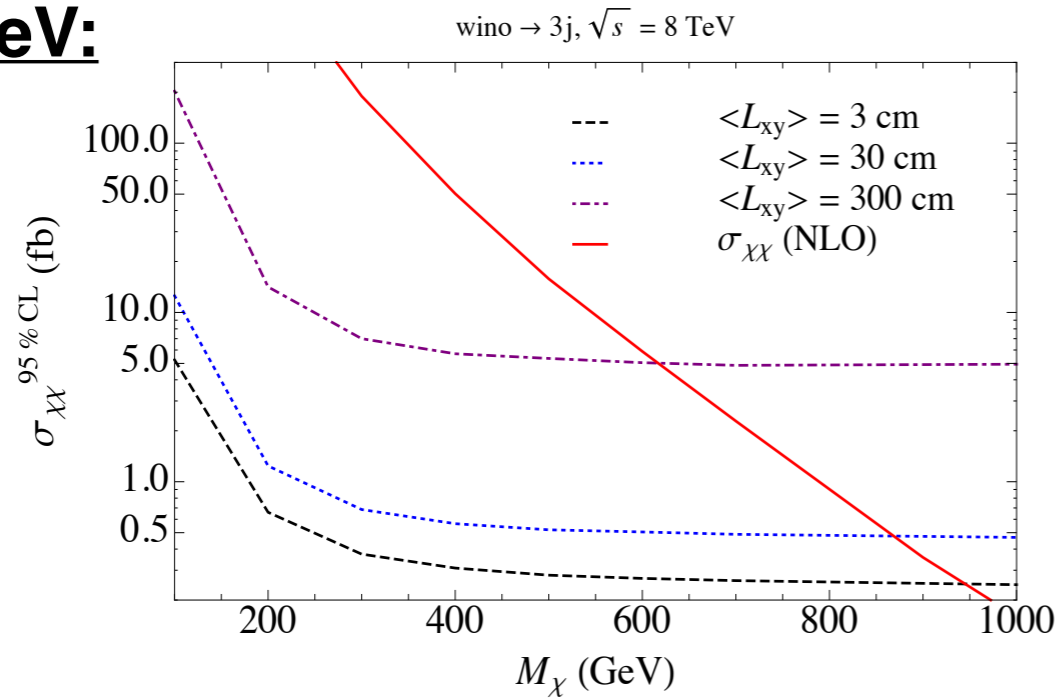


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singlet-like (Higgs portal)

We studied a challenging case:

$M_{\chi} = 150$ GeV, moderately off-shell!

No bound @ 8 TeV 20 fb⁻¹!

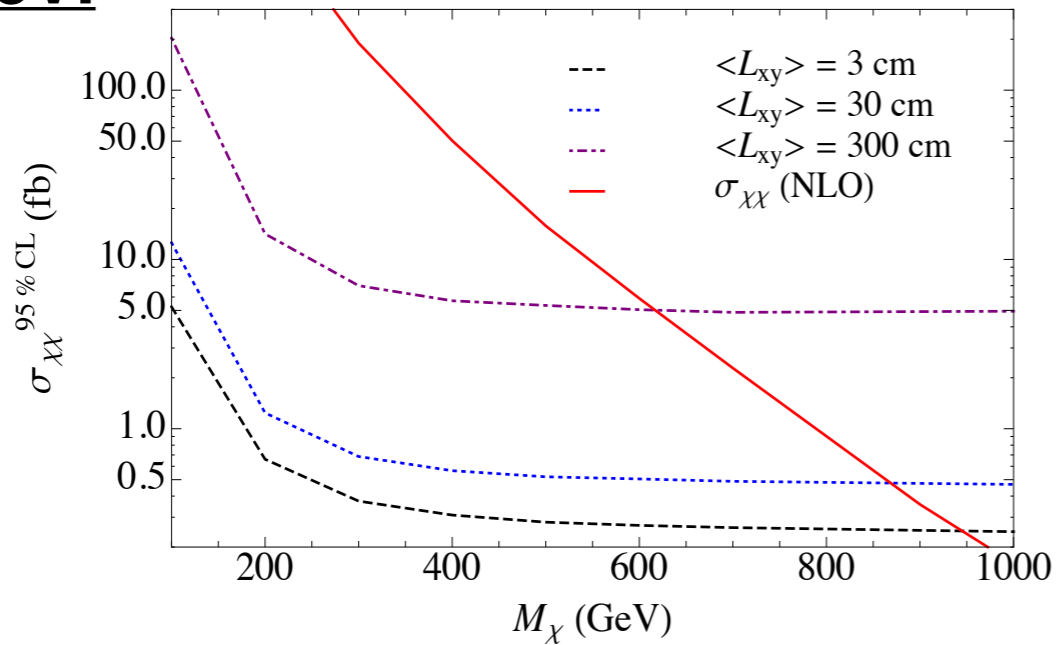
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wino \rightarrow 3j, $\sqrt{s} = 8$ TeV



singlet-like (Higgs portal)

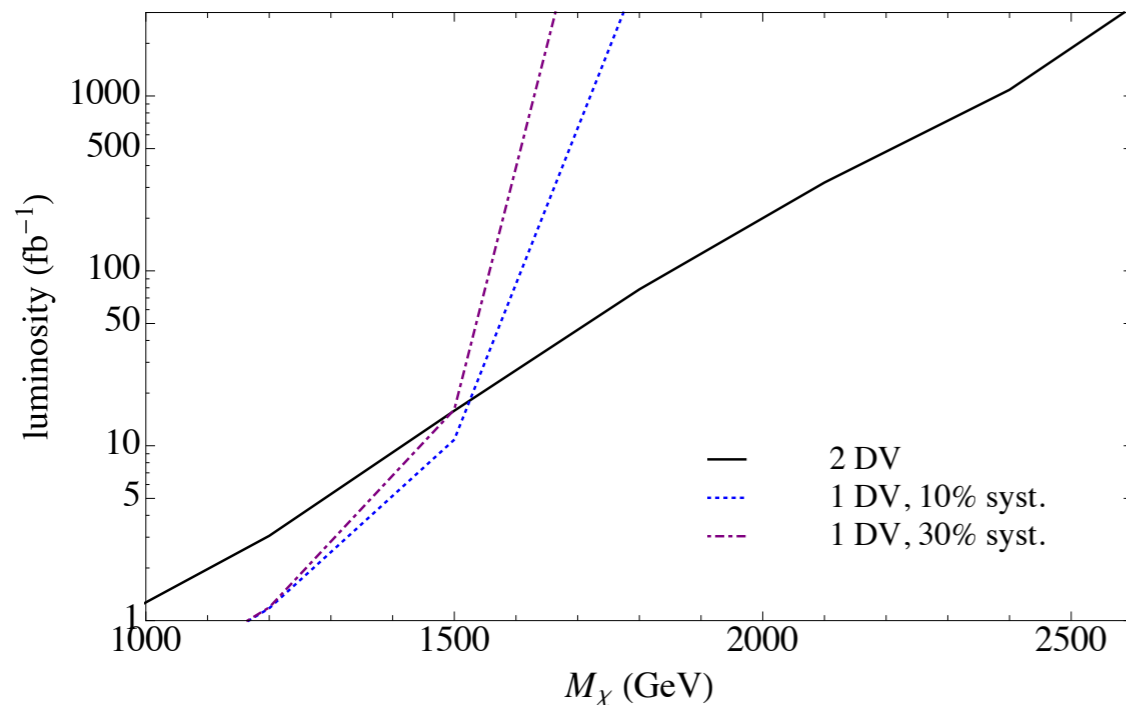
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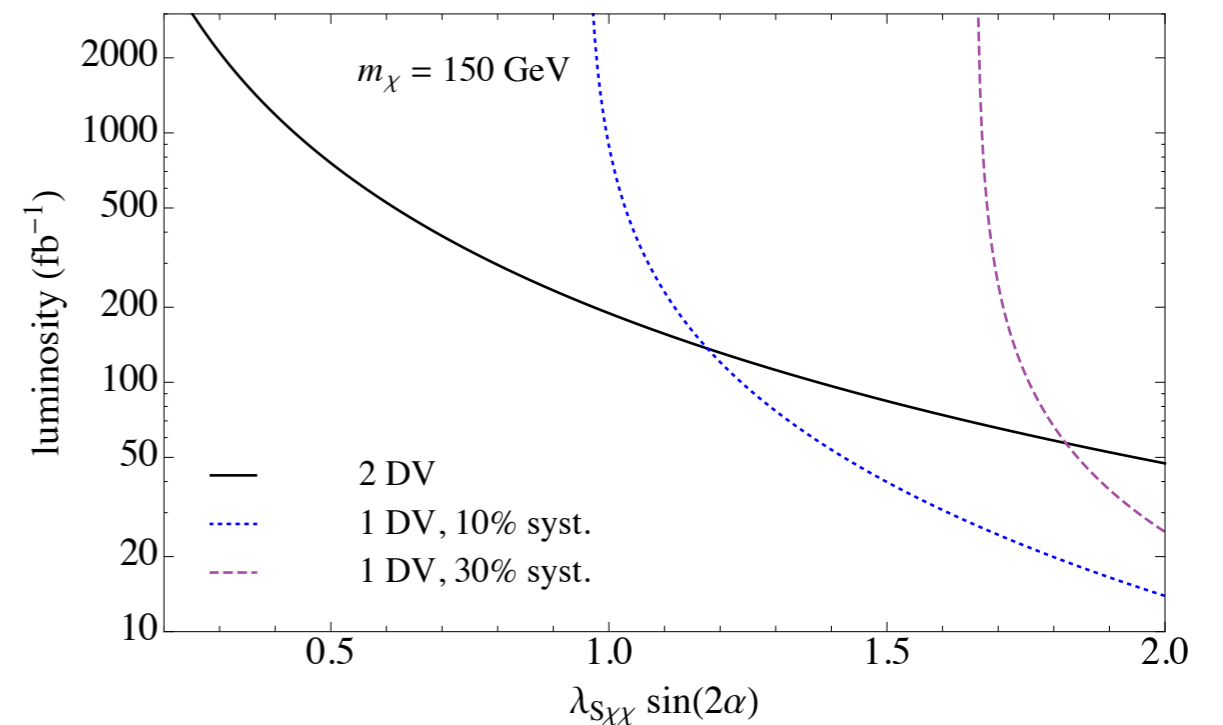
13 TeV:

wino \rightarrow 3j, 2 DV, luminosity for 3 events, $\sqrt{s} = 13$ TeV

$L_{xy} = 3$ cm



Higgs portal $\chi \rightarrow$ 3j, 1DV vs. 2DV comparison $\sqrt{s} = 13$ TeV

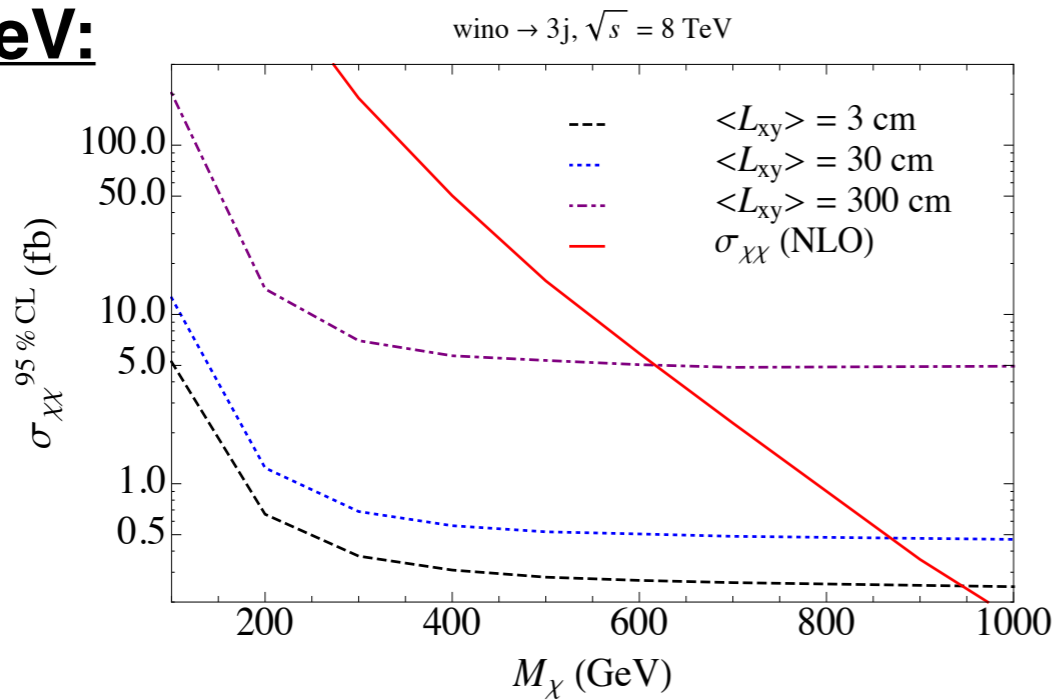


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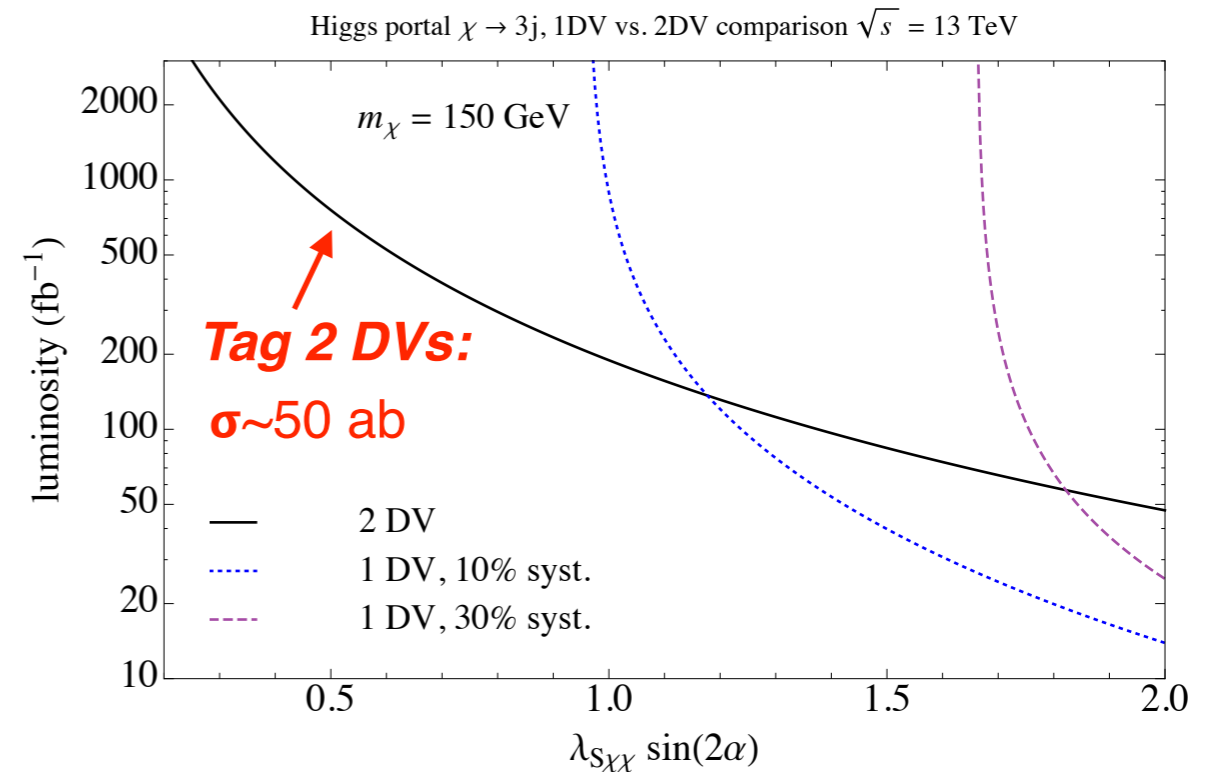
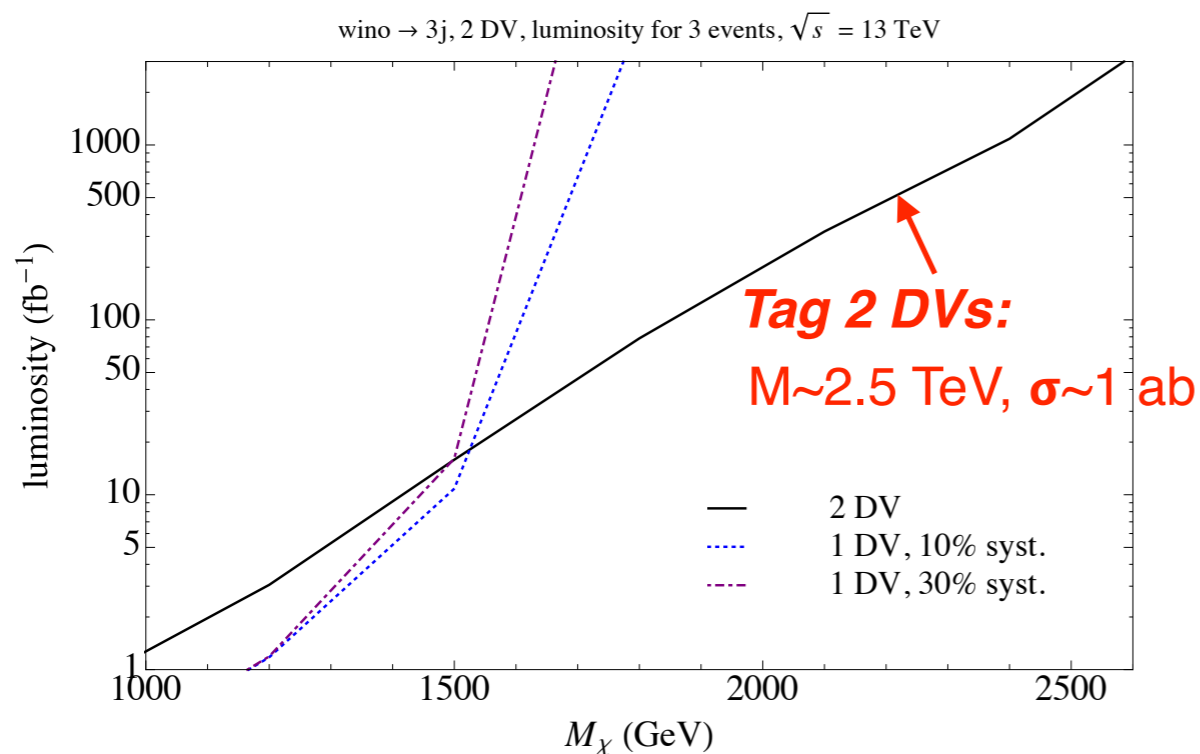
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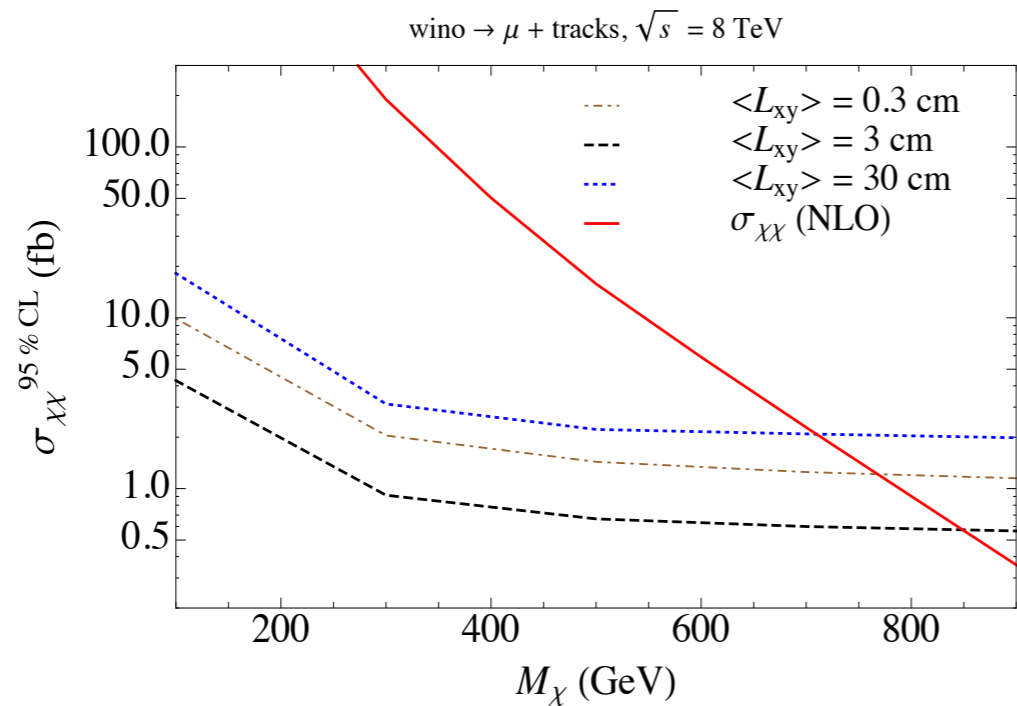


Displaced muon + Tracks

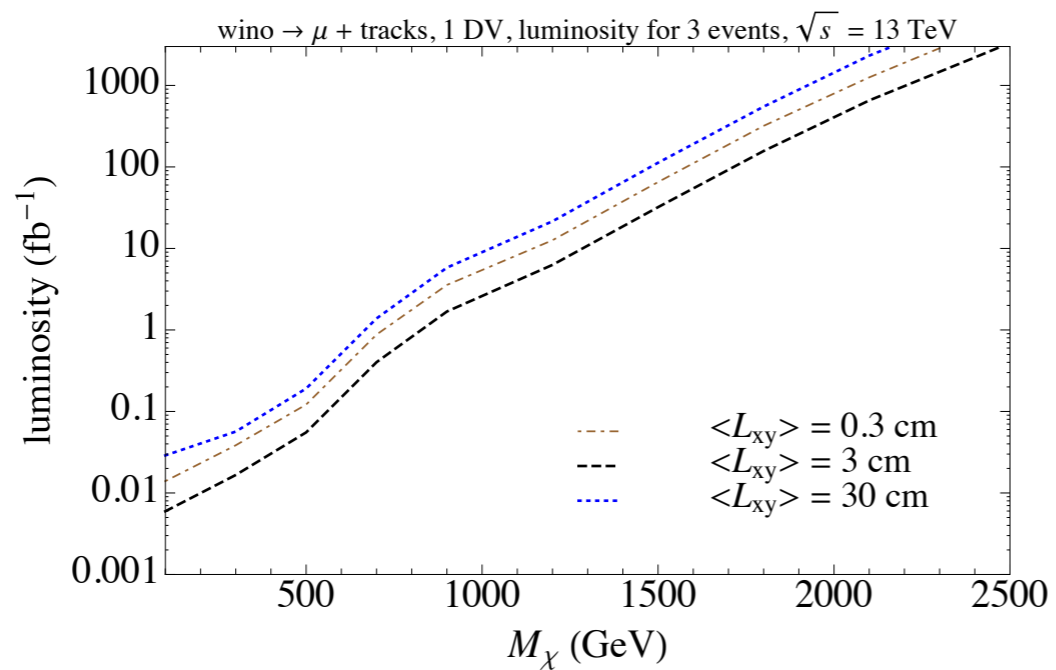
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wino

8 TeV



13 TeV:
Tag 1 DV
 $M \sim 2.5 \text{ TeV}$
(lower bkg
than all-
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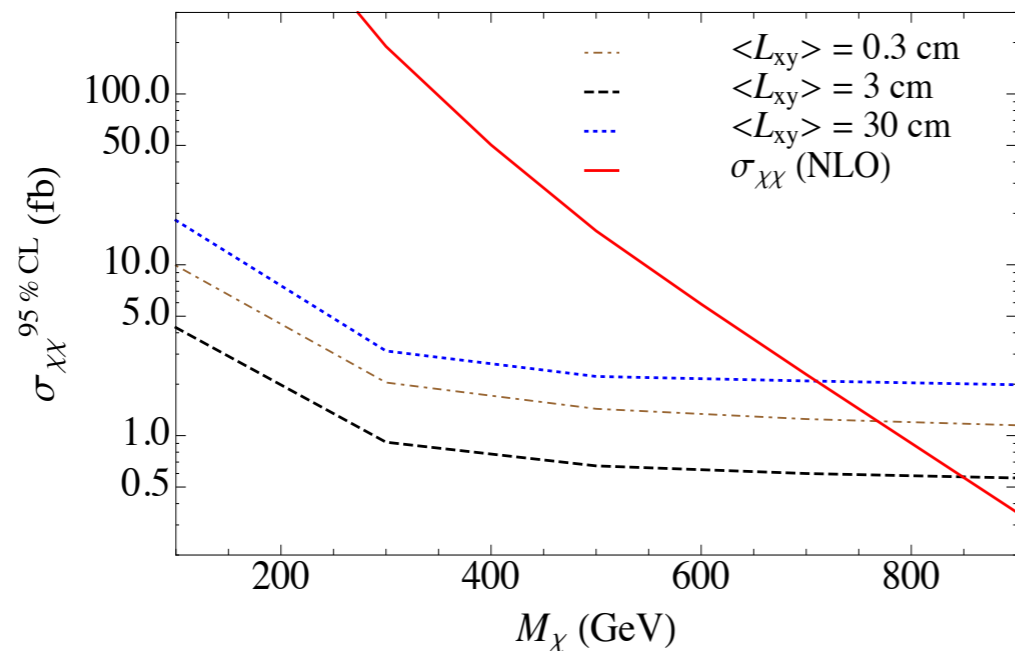


Displaced muon + Tracks

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wino

wino $\rightarrow \mu + \text{tracks}$, $\sqrt{s} = 8 \text{ TeV}$



8 TeV

singlet (Higgs portal)

(singlet-like, $M_\chi = 150 \text{ GeV}$)

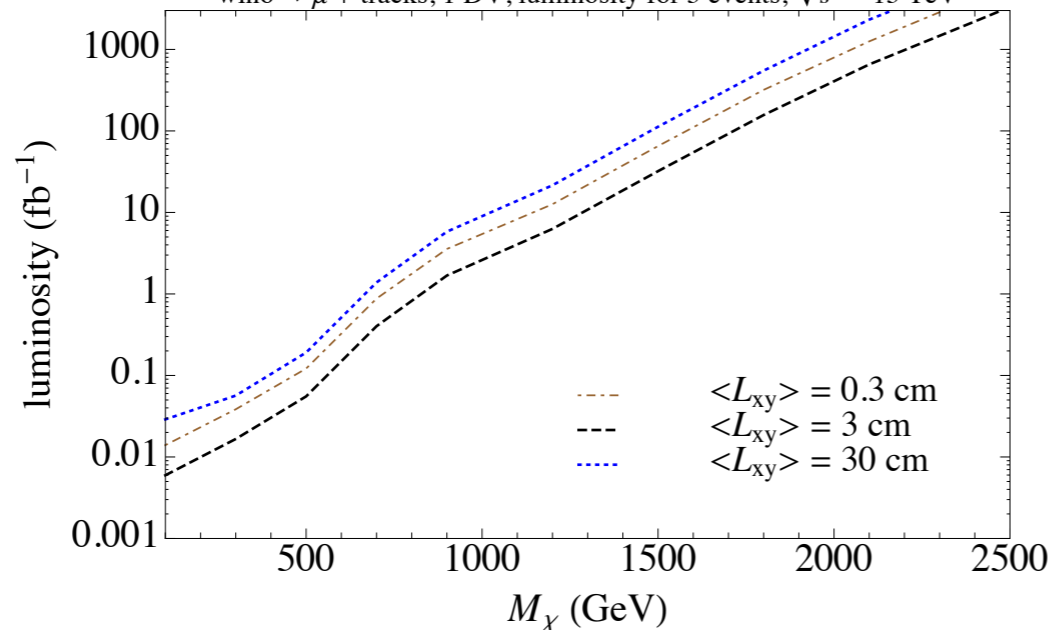
No bound @ 8 TeV 20 fb⁻¹

- 13 TeV: $\sigma_S \sim 50 \text{ ab}$ for $L_{xy} \sim 1 \text{ cm}$
(Tag 1 DV)

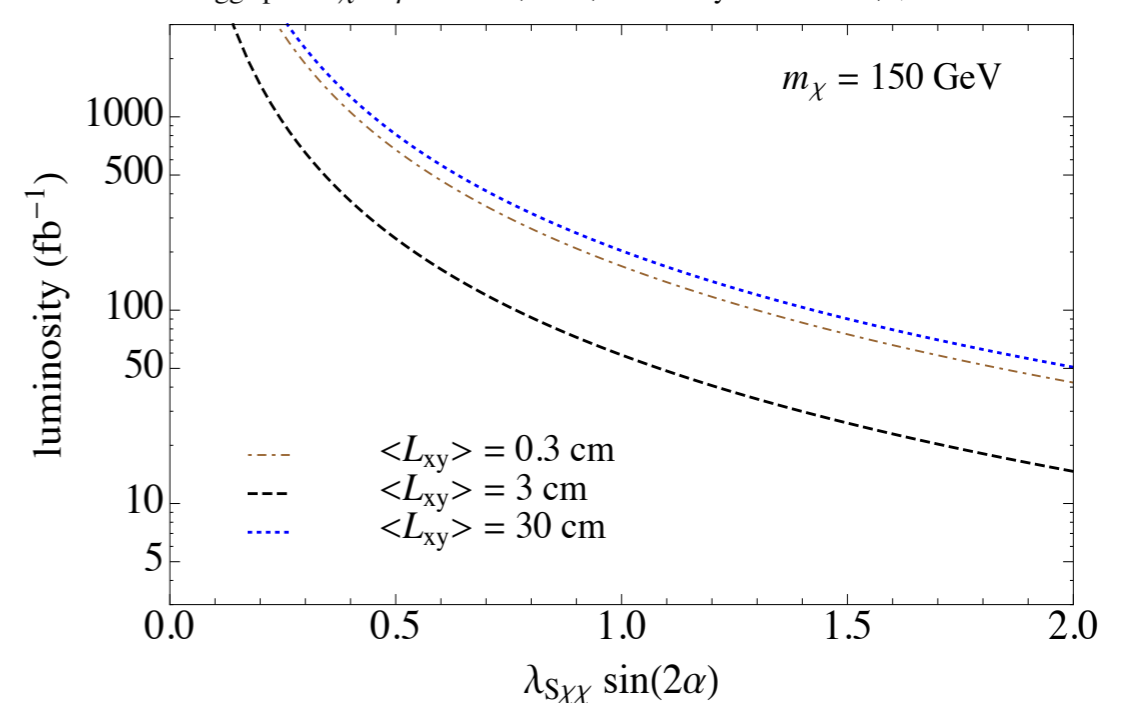
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Tag 1 DV
M ~ 2.5 TeV
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wino $\rightarrow \mu + \text{tracks}$, 1 DV, luminosity for 3 events, $\sqrt{s} = 13 \text{ TeV}$



Higgs portal $\chi \rightarrow \mu + \text{tracks}$, 1DV, luminosity for 3 events, $\sqrt{s} = 13 \text{ TeV}$



Summary/Outlook

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 - A **robust cosmological motivation** for DV searches
 - Reproduce/study early universe BG @ LHC! (cf. WIMP DM search)
- **WIMP baryogenesis:** a motivated example, new mechanism addressing $\Omega_B (+) \Omega_B \sim \Omega_{DM}$, natural embedding in SUSY

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- **WIMP baryogenesis:** a motivated example, new mechanism addressing $\Omega_B (+) \Omega_B \sim \Omega_{DM}$, natural embedding in SUSY
- **Simplified models for LHC pheno:** signal generator for **general** DV searches (*cooperation with ATLAS displaced jets group, officially approved as a **new benchmark** for Run-2 analysis*)
- **Further pheno explorations:**
 - Other signal channels: diphoton resonance, multi-b/t events (YC and Okui, Yunesi arxiv:1605.08736, JHEP)
 - Challenging case for DV search: Light WIMP ($\lesssim 100 \text{ GeV}$), longer lifetime, hadronic decay (sphaleron turned off)... ?