

# Precision neutrino physics

European Strategy for Future Neutrino Physics

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# Neutrino are optimal windows into the exotic -dark- sectors

- \* Can mix with new neutral fermions, heavy or light
- \* Interactions not obscured by strong and e.m. ones

# What are the main physics goals in $\nu$ physics?

- To determine the absolute scale of masses
  - To determine whether they are Majorana
- \* To discover Leptonic CP-violation

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What is the relation of these putative discoveries to the matter-antimatter asymmetry of the universe?

Can leptogenesis be “proved”?

The short, and rather accurate answer

NO

Nevertheless, a positive discovery of both  
2 last points

would constitute a very compelling argument  
in favour of leptogenesis

# What are the main physics goals in $\nu$ physics?

- To determine the absolute scale of masses  
(Tritium...., cosmo?)
- To determine whether they are Dirac Majorana  
(neutrinoless  $\beta\beta$  decay, degenerate or inverse hierarchy)
- To discover Leptonic CP-violation  
(in  $\nu_\mu$ - $\nu_e$  oscillations at superbeams, betabeams....  
neutrino factory)

Go for those discoveries!

# Entering the era of precision neutrino oscillation physics

~ % level

$\nu_\mu \longleftrightarrow \nu_e$  golden channel...

Neutrino masses indicate new  
physics beyond the SM

Maybe new physics could also  
appear in neutrino couplings ?

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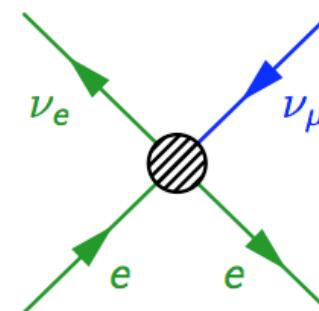
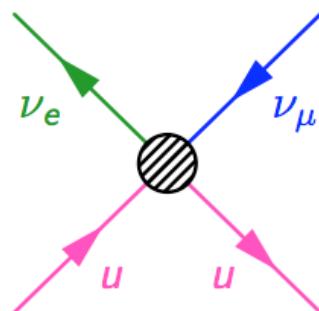
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# Neutrino masses indicate new physics beyond the SM



Maybe new physics could appear in  
**Non Standard neutrino Interactions?**

i.e. NSI



Neutrino masses indicate new  
physics beyond the SM



Maybe new physics could appear in  
**Non Standard neutrino Interactions?**

Or other exotic neutrino couplings....

Or rare lepton decays....

Two cases:

Heavy new scales ( $M > v$ )

Light new scales ( $M \ll v$ )

Heavy new scale     $M > v$

$$\mathcal{L} = \mathcal{L}_{SM} + C^{d=5} \frac{O^{d=5}}{M} + C^{d=6} \frac{O^{d=6}}{M^2} + \dots$$

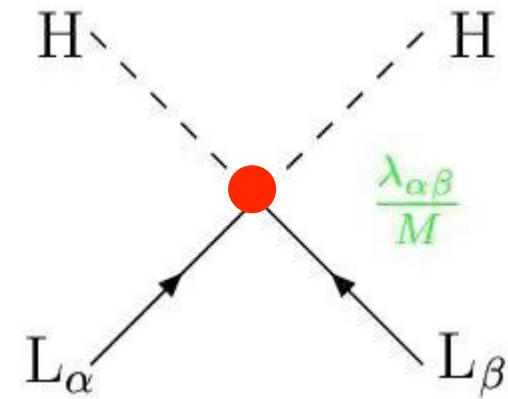
# $v$ masses beyond the SM

## The Weinberg operator

Dimension 5 operator:

$$\frac{\lambda}{M} (L L H H) \rightarrow \frac{\lambda v^2}{M} (v v)$$

$O^{d=5}$

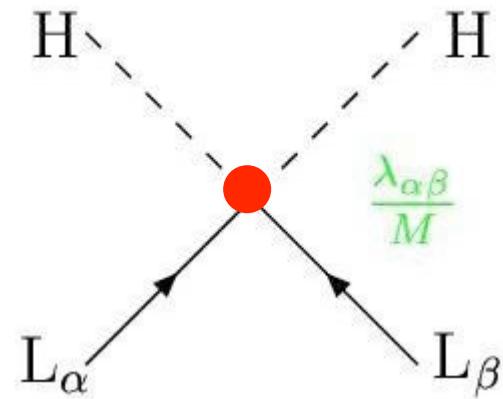


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$$\frac{\lambda}{M} (L L H H) \xrightarrow{O^{d=5}} \frac{\lambda v^2}{M} (vv)$$



It's unique  $\rightarrow$  very special role of  $\nu$  masses:  
lowest-order effect of higher energy physics

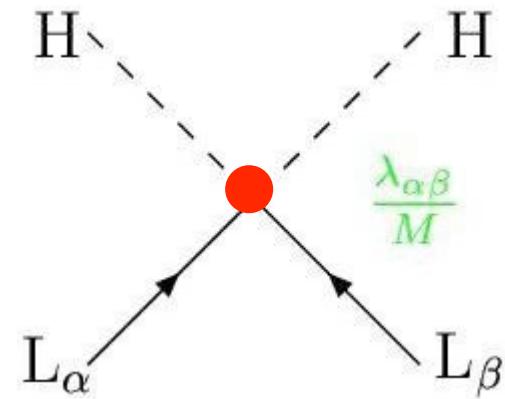
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This mass term **violates lepton number (B-L)**  
 $\rightarrow$  Majorana neutrinos

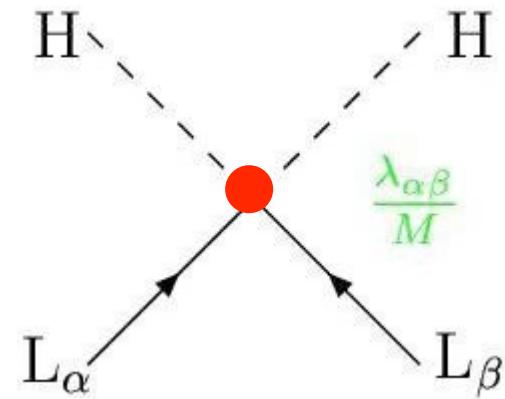
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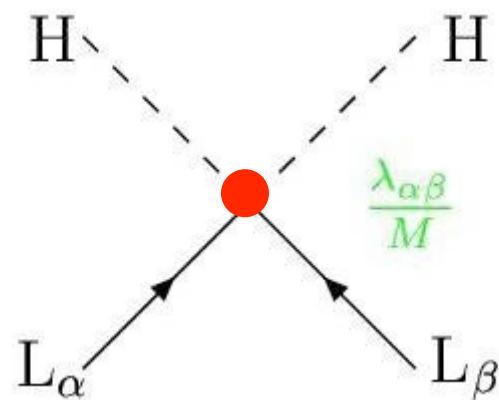
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$O^{d=5}$  *is common to all models of Majorana  $\nu$ s*

# New Standard Model $\textcolor{red}{v}\textbf{SM}$ ?

$$\mathcal{L}_{\textcolor{red}{v}SM} = \mathcal{L}_{SM} + C^{d=5} \frac{\underline{O^{d=5}}}{\Lambda_{LN}} + \dots$$

# $\nu$ masses beyond the SM : tree level



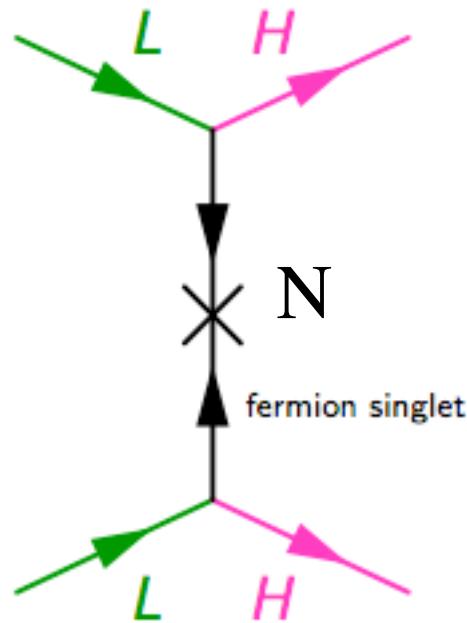
$$2 \times 2 = 1 + 3$$

$$\delta L = C^{d=5} O^{d=5}$$

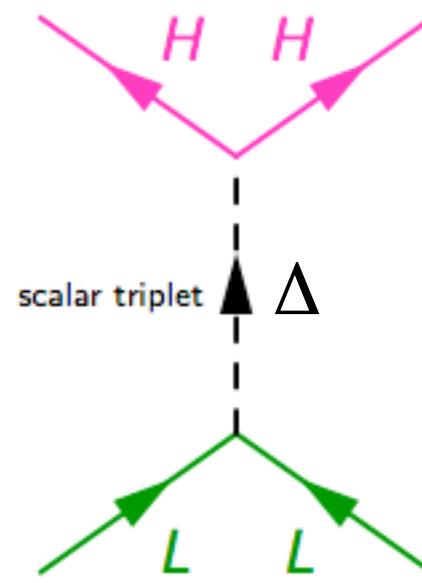
3 generic types (Ma)

# The Seesaw models

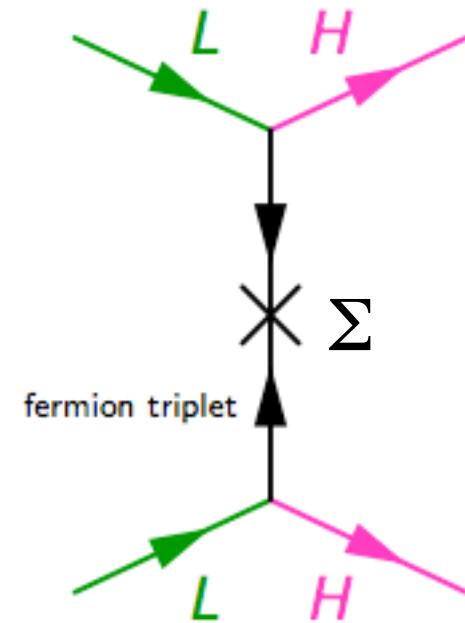
- Three types of models yield the Weinberg operator at tree level



Type I



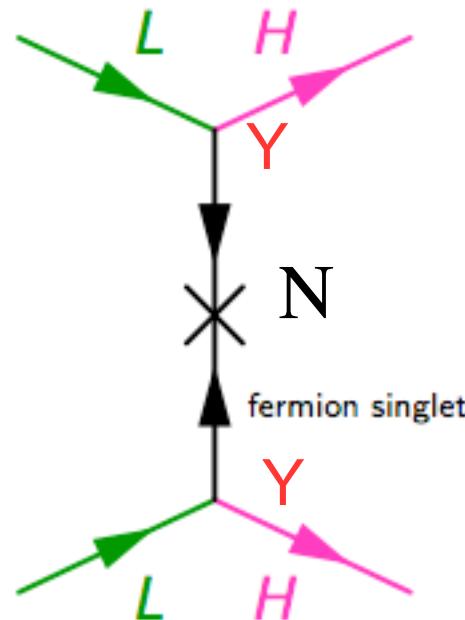
Type II



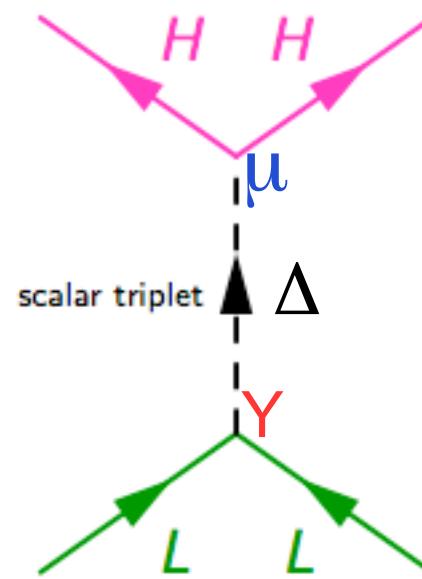
Type III

# The Seesaw models

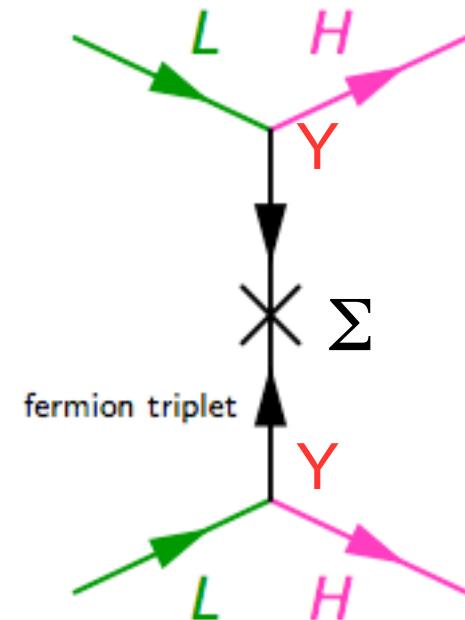
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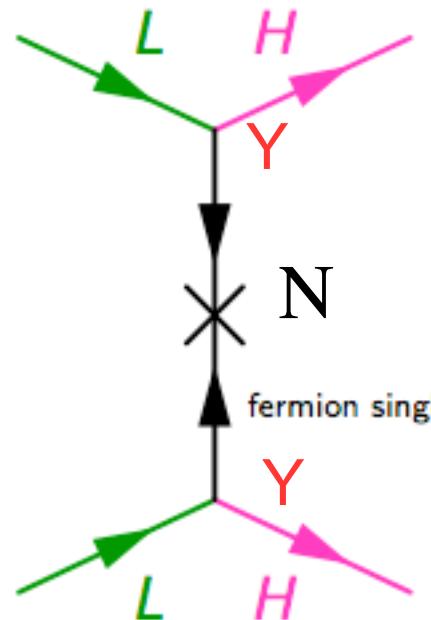
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Type III

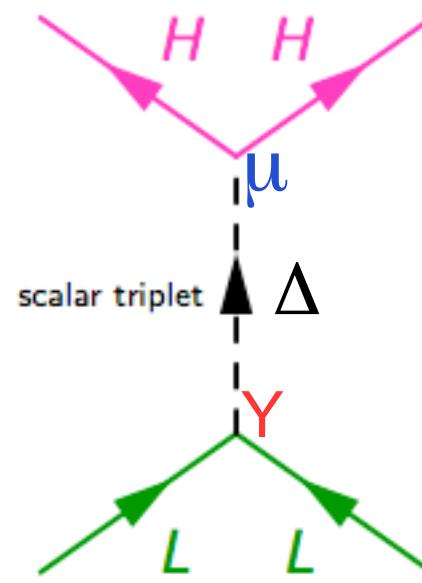
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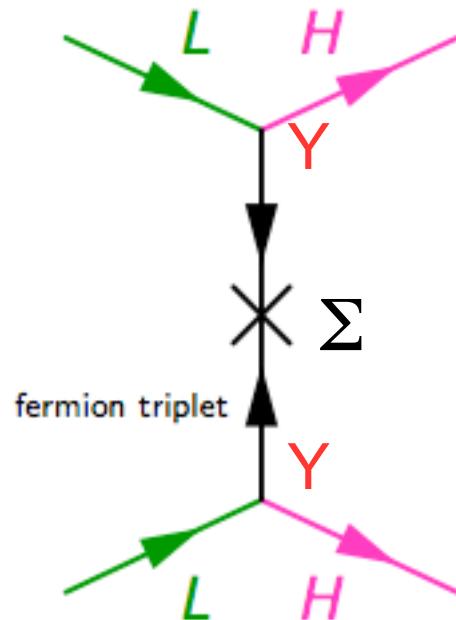
Type I

$$m_v \sim v^2 \frac{Y_N^T}{M_N} \frac{1}{M_N} Y_N$$



Type II

$$m_v \sim v^2 \frac{Y_\Delta^T}{M_\Delta^2} \frac{\mu}{M_\Delta^2}$$

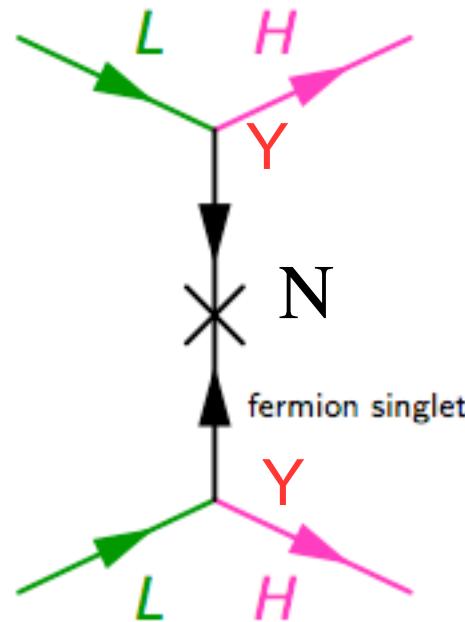


Type III

$$m_v \sim v^2 \frac{Y_\Sigma^T}{M_\Sigma} \frac{1}{M_\Sigma} Y_\Sigma$$

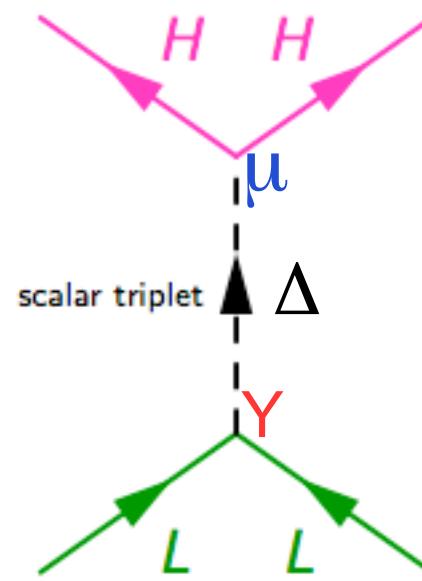
# The Seesaw models

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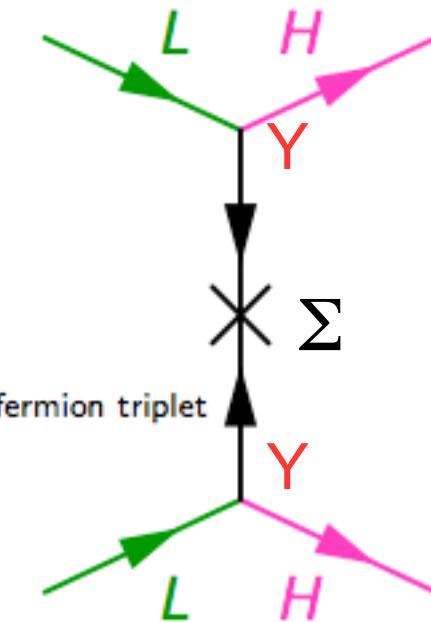
Type I

Heavy fermion singlet  $N_R$   
Minkowski, Gell-Mann, Ramond,  
Slansky, Yanagida, Glashow,  
Mohapatra, Senjanovic



Type II

Heavy scalar triplet  $\Delta$   
Magg, Wetterich, Lazarides,  
Shafi, Mohapatra,  
Senjanovic, Schechter, Valle



Type III

Heavy fermion triplet  $\Sigma_R$   
Ma, Roy, Senjanovic, Hambye et al.,

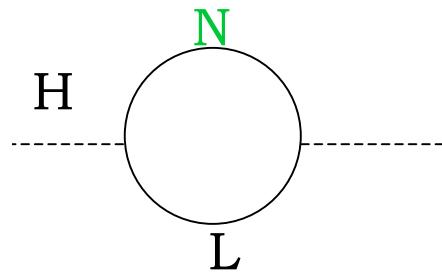
Those fields,  $N_R$ ,  $\Delta$ ,  $\Sigma_R$ , would mediate other processes too....

Which are the new exotic couplings,  
that is, d=6 operators, in Seesaws?

## Observable effects?

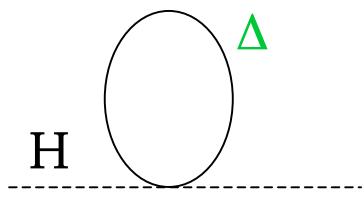
Obviously requires scale near the TeV

$M \sim 1$  TeV is suggested by electroweak hierarchy problem

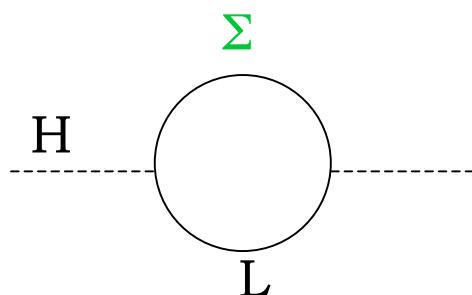


$$\delta m_H^2 = -\frac{Y_N^\dagger Y_N}{16\pi^2} \left[ 2\Lambda^2 + 2M_N^2 \log \frac{M_N^2}{\Lambda^2} \right]$$

(Vissani)



$$\delta m_H^2 = -3 \frac{\lambda_3}{16\pi^2} \left[ \Lambda^2 + M_\Delta^2 \left( \log \frac{M_\Delta^2}{\Lambda^2} - 1 \right) \right] - \frac{\mu_\Delta^2}{2\pi^2} \log \left( \left| \frac{M_\Delta^2 - \Lambda^2}{M_\Delta^2} \right| \right)$$



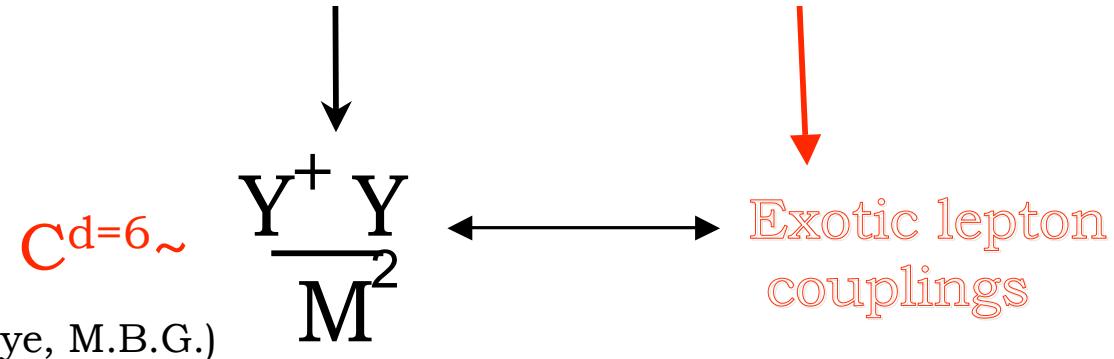
(Abada, Biggio, Bonnet, Hambye, M.B.G.)

$$\delta m_H^2 = -3 \frac{Y_\Sigma^\dagger Y_\Sigma}{16\pi^2} \left[ 2\Lambda^2 + 2M_\Sigma^2 \log \frac{M_\Sigma^2}{\Lambda^2} \right]$$

Model	Effective Lagrangian $\mathcal{L}_{\text{eff}} = c_i \mathcal{O}_i$		
	$c^{d=5}$	$c_i^{d=6}$	$\mathcal{O}_i^{d=6}$
Fermionic Singlet (type I)	$Y_N^T \frac{1}{M_N} Y_N$	$Y_N^\dagger \frac{1}{ M_N ^2} Y_N$	$(\bar{L} \tilde{H}) i \not{\partial} (\tilde{H}^\dagger L)$
Fermionic Triplet (type III)	$Y_\Sigma^T \frac{1}{M_\Sigma} Y_\Sigma$	$Y_\Sigma^\dagger \frac{1}{ M_\Sigma ^2} Y_\Sigma$	$(\bar{L} \vec{\tau} \tilde{H}) i \not{\partial} (\tilde{H}^\dagger \vec{\tau} L)$
Scalar Triplet (type II)	$4Y_\Delta \frac{\mu_\Delta}{ M_\Delta ^2}$	$Y_\Delta^\dagger \frac{1}{2 M_\Delta ^2} Y_\Delta$	$(\bar{\tilde{L}} \vec{\tau} L) (\bar{L} \vec{\tau} \tilde{L})$
		$\frac{ \mu_\Delta ^2}{ M_\Delta ^4}$	$(H^\dagger \vec{\tau} \tilde{H}) (\overleftarrow{D}_\mu \overrightarrow{D}^\mu) (\tilde{H}^\dagger \vec{\tau} H)$
		$-2(\lambda_3 + \lambda_5) \frac{ \mu_\Delta ^2}{ M_\Delta ^4}$	$(H^\dagger H)^3$

(Abada, Biggio, Bonnet, Hambye, M.B.G.)

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		$\frac{ \mu_\Delta ^2}{ M_\Delta ^4}$	$(H^\dagger \vec{\tau} \tilde{H}) (\bar{D}_\mu D^\mu) (\tilde{H}^\dagger \vec{\tau} H)$
		$-2(\lambda_3 + \lambda_5) \frac{ \mu_\Delta ^2}{ M_\Delta ^4}$	$(H^\dagger H)^3$



For all scalar and fermionic  
Seesaw models, present bounds:

$$\frac{v^2}{2} |c^{d=6}|_{\alpha\beta} = \frac{v^2}{2} |Y^\dagger \frac{1}{M^2} Y|_{\alpha\beta} \lesssim 10^{-2}$$



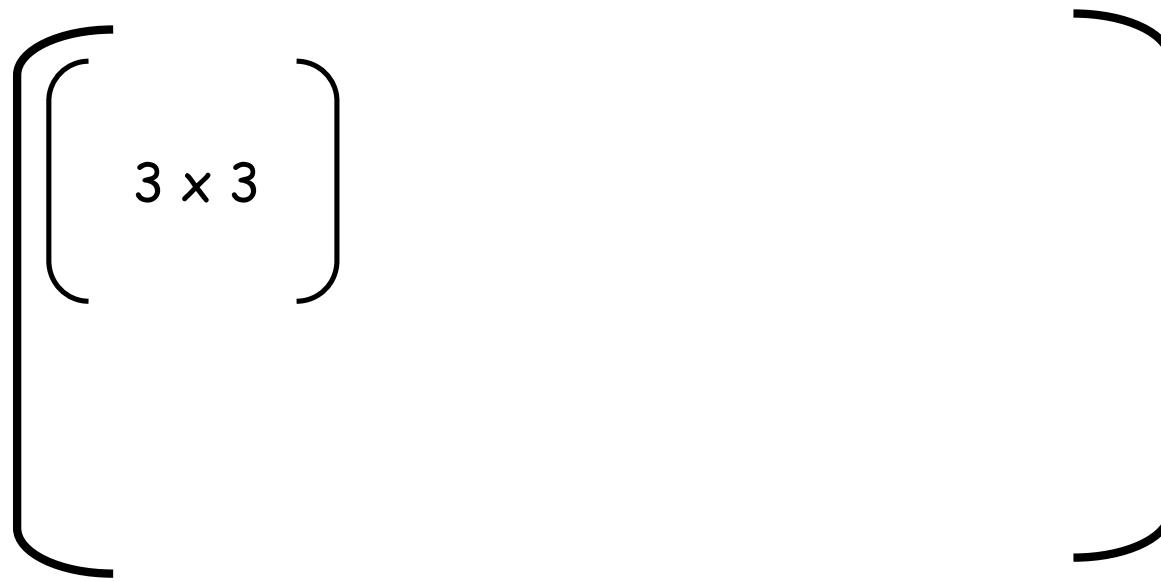
$$|Y| \lesssim 10^{-1} \frac{M}{1 \text{TeV}}$$

or stronger

Fermionic seesaws ---> Non unitarity

The complete theory of  $\nu$  masses is unitary.

i.e, a neutrino mass matrix larger than  $3 \times 3$

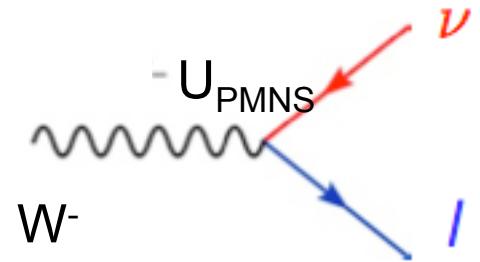


All fermionic Seesaws exhibit non-unitary mixing

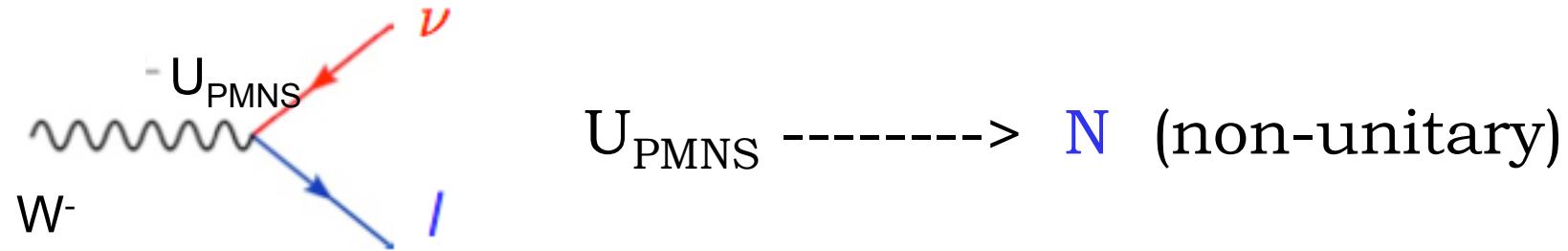
In fact, it is a quite general characteristic of new physics, well beyond the Seesaw scenarios:

*A non-unitary mixing matrix arises when leptons mix with heavy fermions*

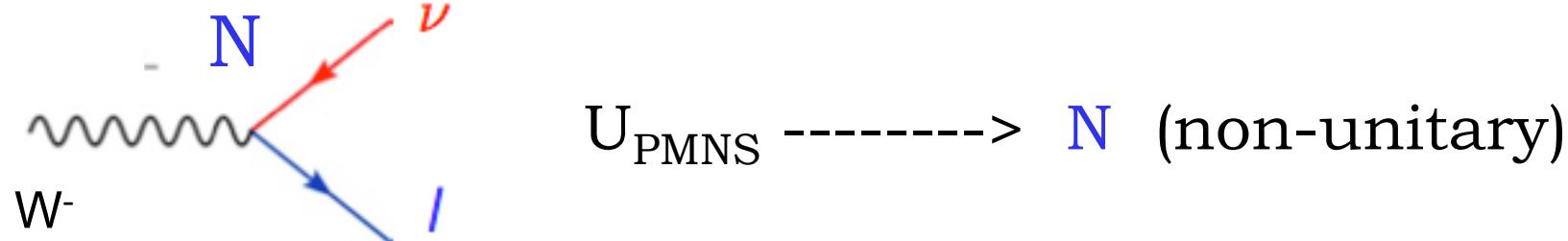
i.e. In charged currents



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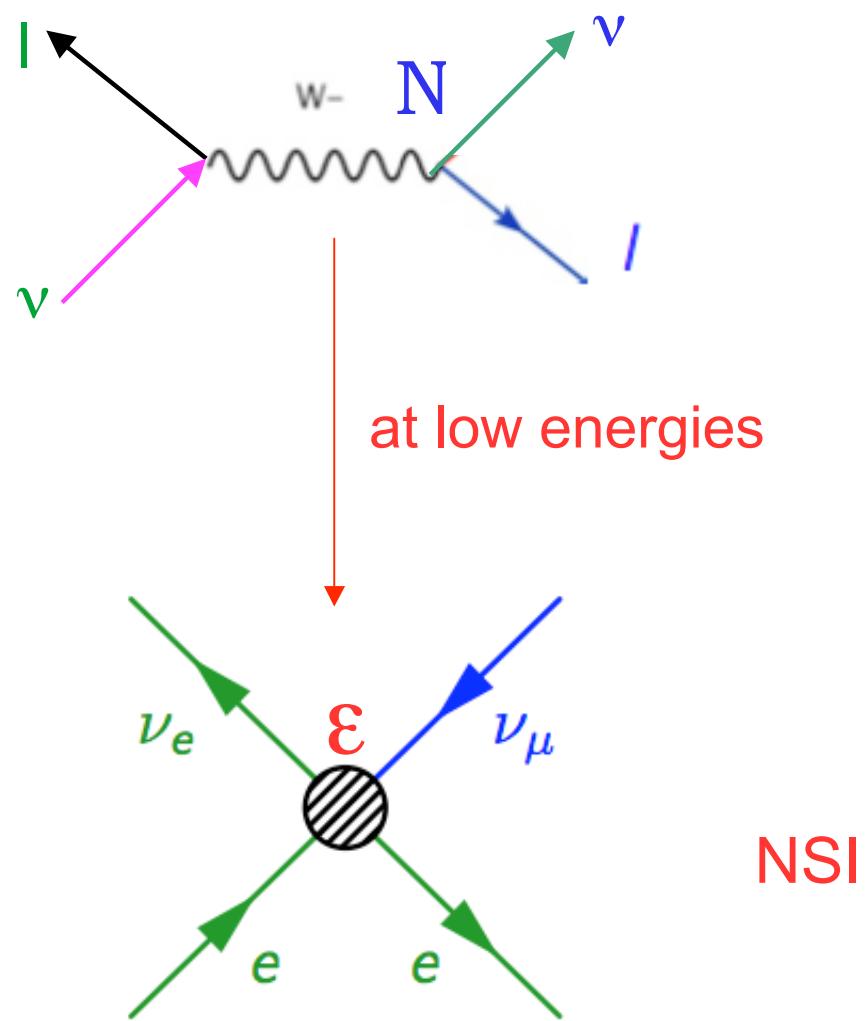
i.e. In charged currents



$$N \approx (1 + \varepsilon) U_{PMNS}$$

$$(|NN^\dagger| - 1)_{\alpha\beta} = |\varepsilon|_{\alpha\beta} = -v^2 \left| \frac{Y^+ Y}{M^2} \right|_{\alpha\beta}$$

i.e.



All fermionic Seesaws exhibit non-unitary mixing

# Unitarity constraints on $(NN^+)$ from:

- \* All  $\nu$  oscillation + near detector data
    - +
  - \* Weak decays...
  - \* W decays
    - \* Invisible Z width
    - \* Universality tests
    - \* Rare lepton decays
- $|N|$  is unitary at the % level

i.e. for heavy singlet neutrinos

$$|\epsilon_{\alpha\beta}| = \frac{v^2}{2} |c_{\alpha\beta}^{d=6,kin}| < \begin{pmatrix} 4.0 \cdot 10^{-3} & 1.2 \cdot 10^{-4} & 3.2 \cdot 10^{-3} \\ 1.2 \cdot 10^{-4} & 1.6 \cdot 10^{-3} & 2.1 \cdot 10^{-3} \\ 3.2 \cdot 10^{-3} & 2.1 \cdot 10^{-3} & 5.3 \cdot 10^{-3} \end{pmatrix}$$

Antusch, Baumann, Fdez-Martinez 08

# Future bounds on $\varepsilon_{\alpha\beta}$

$\nu_e \leftrightarrow \nu_\mu$ :  $\mu \rightarrow eee$ ,  $\mu \rightarrow e\gamma$  ( $\sim 10^{-5}$ MEG),  $\mu$ -e conversion  
(PRISM/PRIME)

$\nu_e \leftrightarrow \nu_\tau$   
 $\nu_\mu \leftrightarrow \nu_\tau$

}  $\tau$  channels at  $\nu$  factory  $\sim 10^{-3}....$

... or simpler?:

$\nu_e \leftrightarrow \nu_\tau$  with betabeams ?  
(Agarwalla,Huber,Link)  
disappearance

$\nu_\mu \leftrightarrow \nu_\tau$  even easier??

A simple proposal for  $\nu_\mu \leftrightarrow \nu_\tau$  :

# MINYSIS

Adam Para

(*Main Injector Non Standard Interactions Search*)  
<http://www-off-axis.fnal.gov/MINYSIS/>

- Emulsion near detector at the NuMI beam

$\nu_\tau$  appearance

can achieve  $\epsilon_{\mu\tau} < 10^{-3}$  in  $\sim 5$  years

(over Chorus-Nomad  $\epsilon_{\mu\tau} < 10^{-2}$ )

PRELIMINARY

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Adam Para +..... A. de Gouvea, E. Fernandez-Martinez, M.B.G., P.  
Hernandez, J. Kopp, O. Mena, A. Nelson, S. Parke, T.Ota, JJ.

PRELIMINARY

## Can we measure the phases of $\mathbf{N}$ ?

$$\mathbf{N} = \begin{pmatrix} N_{e1} & N_{e2} & N_{e3} \\ N_{\mu 1} & N_{\mu 2} & N_{\mu 3} \\ N_{\tau 1} & N_{\tau 2} & N_{\tau 3} \end{pmatrix}$$

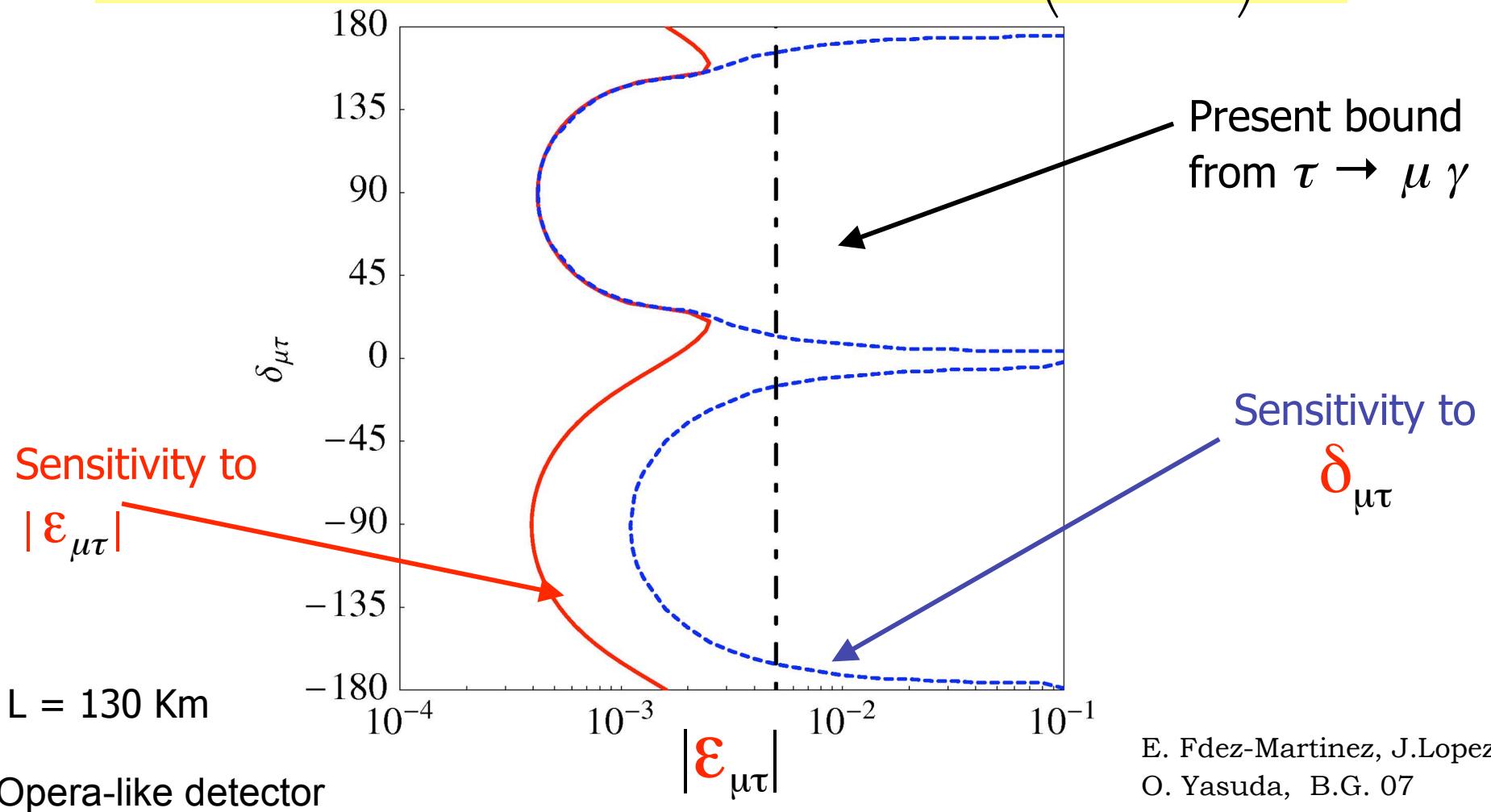
- New CP-violation signals  
even in the two-family approximation

E. Fdez-Martinez, J.Lopez, O. Yasuda, B.G. 07

i.e.  $P(\nu_\mu \rightarrow \nu_\tau) \neq P(\bar{\nu}_\mu \rightarrow \bar{\nu}_\tau)$

- Increased sensitivity to the moduli  $|N|$   
in future Neutrino Factories

$$P_{\mu\tau} - P_{\bar{\mu}\bar{\tau}} = -4 \operatorname{Im}(\varepsilon_{\mu\tau}) \sin(2\theta_{23}) \sin\left(\frac{\Delta m_{23}^2 L}{2E}\right)$$



E. Fdez-Martinez, J.Lopez,  
O. Yasuda, B.G. 07

For non-trivial  $\delta_{\mu\tau}$ , one order of magnitude improvement for  $|N|$



Good prospect for  $\nu_\mu - \nu_\tau$  channel at near/short distance detector - $O(100 \text{ km})$

- \* Recently: Goswami+ Ota; Altarelli+Meloni, Tang+Winter at nufact,
- \* Antusch et al.-->  $\varepsilon_{e\tau}$  from the golden channel  $\nu_e - \nu_\mu$  at  $\sim 1000 \text{ km}$

# Could d=6 be stronger than d=5 ?

- \* Two independent scales in d=5, d=6 from a symmetry principle: lepton number

Cirigliano et al; Kersten,Smirnov; Abada et al

- \* d=5 requires to violate lepton number
- \* d=6 does not violate any symmetry

$$\Lambda_5 \sim \Lambda_{LN} \gg \Lambda_6 \sim \Lambda_{LFV} \sim \text{TeV}$$

$\Lambda_{LN} \gg \Lambda_{fl} \sim \text{TeV}$  ?

$$\mathcal{L} = \mathcal{L}_{SM} + \sum_i \frac{\alpha}{\Lambda_{LN}} O_i^{d=5} + \sum_i \frac{\beta_i}{\Lambda^2} \underset{\text{flavour}}{O_i^{d=6}} + \dots$$

Cirigliano, et al

There is a sensible physics motivation:

- Origin of lepton/quark flavour violation linked/close to the EW scale
- (Effective) Lepton number breaking scale higher and responsible for the gap between  $\nu$  and other fermion

Seesaw mechanism

vs

Minimal Flavour Violation

# Minimal Flavour Violation

The global Flavour symmetry of the SM: without Yukawas

$$G = SU(3)_{Q_L} \times SU(3)_u \times SU(3)_d \times SU(3)_L \times SU(3)_e$$

$$Q_u \rightarrow L_u Q_u, \quad d_R \rightarrow R_d d_R, \dots$$

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**MFV Hypothesis**  $\equiv$  The Yukawas are the only sources (*irreducible*) of flavour violation.

R. S. Chivukula and H. Georgi, Phys. Lett. B 188, 99 (1987).



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Quark sector  $\mathcal{L} = \mathcal{L}_{SM} + C^{d=6} \frac{O^{d=6}}{\Lambda_{fl}^2} + \dots$

(D'Ambrosio, Cirigliano, Isidori, Grinstein, Wise....Buras....)

Predictive !

# Minimal Flavour Violation

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(D'Ambrosio, Cirigliano, Isidori, Grinstein, Wise....Buras....)

i.e.  $C^{d=6} \sim \frac{Y_{\alpha\beta}^+ Y_{\gamma\delta}}{M^2}$        $O^{d=6} \sim \bar{Q}_\alpha Q_\beta \bar{Q}_\gamma Q_\delta$

## A rationale for the MFV ansatz?

- Flavour data (i.e. B physics) consistent with all flavour physics coming from Yukawa
- Inspired in “condensate” flavour physics a la Froggat-Nielsen ( $\text{Yukawas} \sim \langle \Psi \bar{\Psi} \rangle^n / \Lambda_{\text{fl}}$ ), rather than in susy-like options
- It makes you think on the relation between scales:  
electroweak vs. flavour vs lepton number scales

# *What happens in the presence of neutrino masses?*

Cirigliano, Isidori, Grinstein, Wise

In the lepton sector

$$\mathcal{L} = \dots + Y_e \bar{L} \phi e_R \dots + \sum_i c_{d=5}^i \mathcal{O}_{d=5}^i + c_{d=6}^i \mathcal{O}_{d=6}^i \dots$$

$\xleftarrow{\mathcal{L}_{SM}}$

↑  
 $\Lambda_{LN}$   
↑  
 $\Lambda_{flavour}$

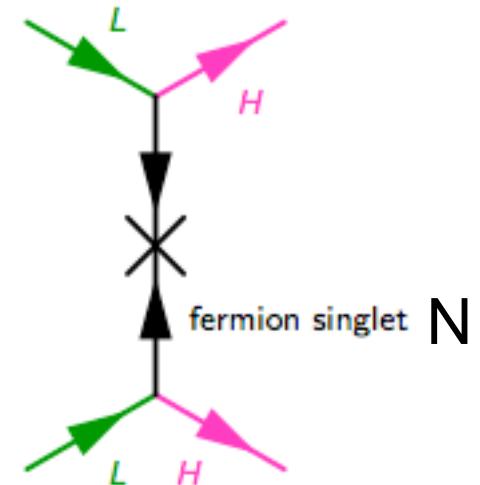
Delicate:

- \* Majorana masses are model dependent :  $c^{d=5}(Y_e, ?)$ ,  $c^{d=6}(Y_e, ?)$
- \* Requires to separate lepton number from flavour origin

# An unsuccessful model: simplest type I

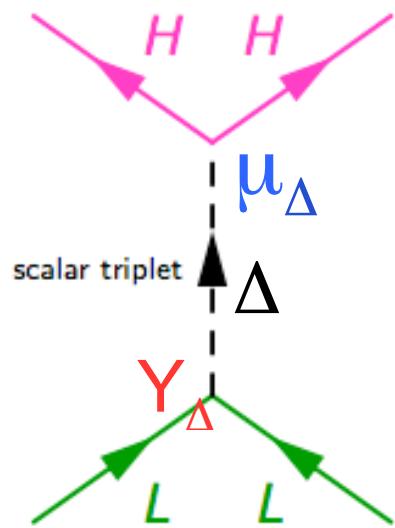
Standard Seesaw (Type I) doesn't work

$$\mathcal{L} = \dots - Y_N \bar{N} \phi^\dagger L_L - \Lambda_{LN} \bar{N}^c N \dots$$



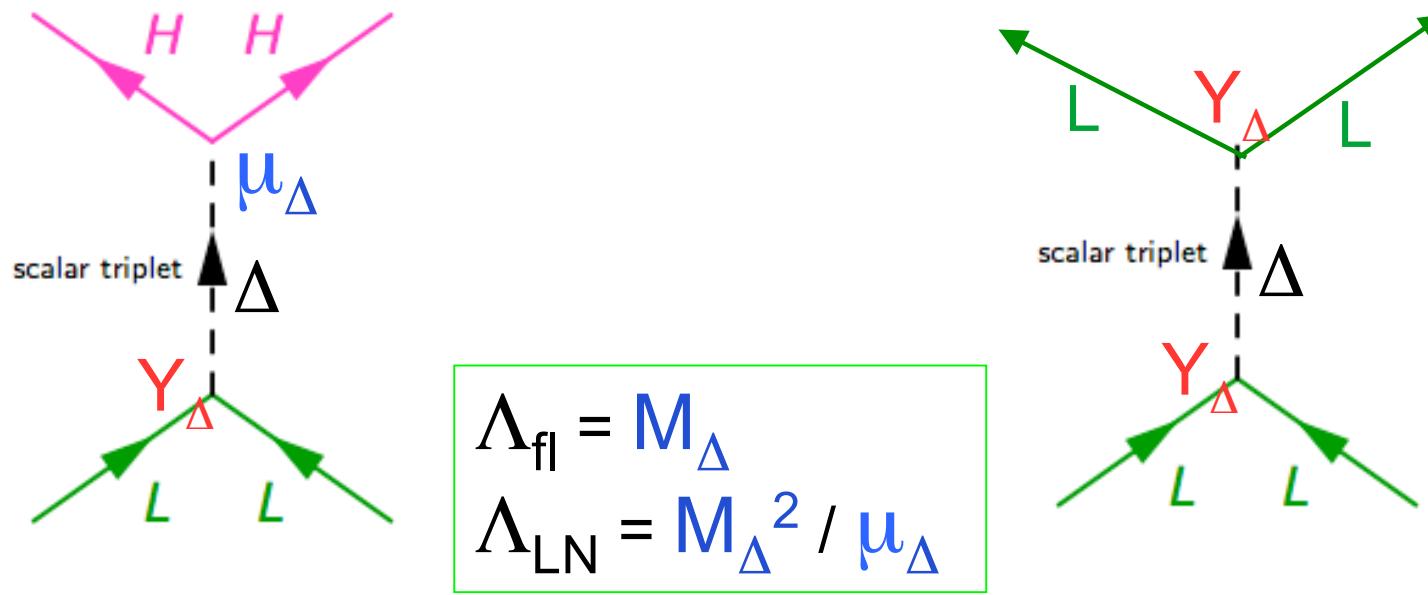
- **Neutrino masses:** Ok.  $m_\nu \propto Y_N^T \frac{1}{\Lambda_{LN}} Y_N$
- **Measurable flavour:** NOT OK!.  $\Lambda_{fl} \equiv \Lambda_{LN}$
- **Predictivity:** More or less Ok.  $c_{d=5} \propto c_{d=6}$  if no CP

# A successful model: Scalar-triplet Seesaw (type II)



$$\begin{aligned}\mathcal{L}_\Delta = & \cdots + (D_\mu \Delta)^\dagger (D^\mu \Delta) - M_\Delta^2 \Delta^\dagger \Delta + + \\ & + Y_\Delta^{\alpha\beta} \overline{\tilde{L}} (\tau \cdot \Delta) L + \mu_\Delta \tilde{\phi}^\dagger (\tau \cdot \Delta)^\dagger \phi + \dots\end{aligned}$$

# A successful model: Scalar-triplet Seesaw (type II)

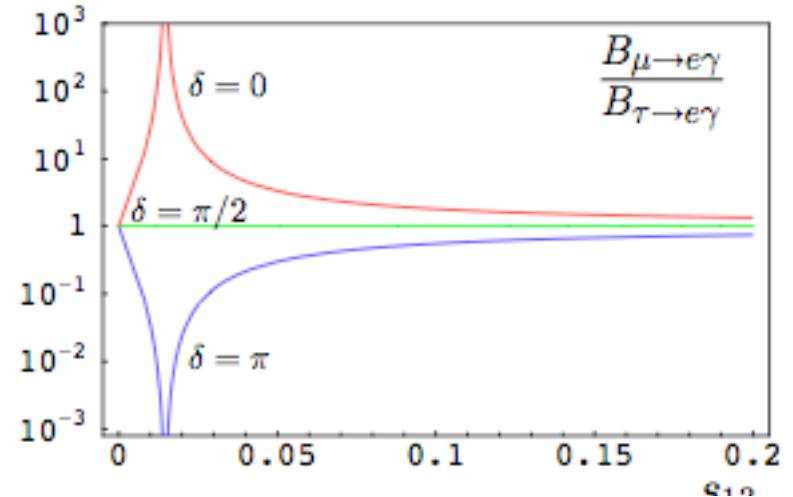
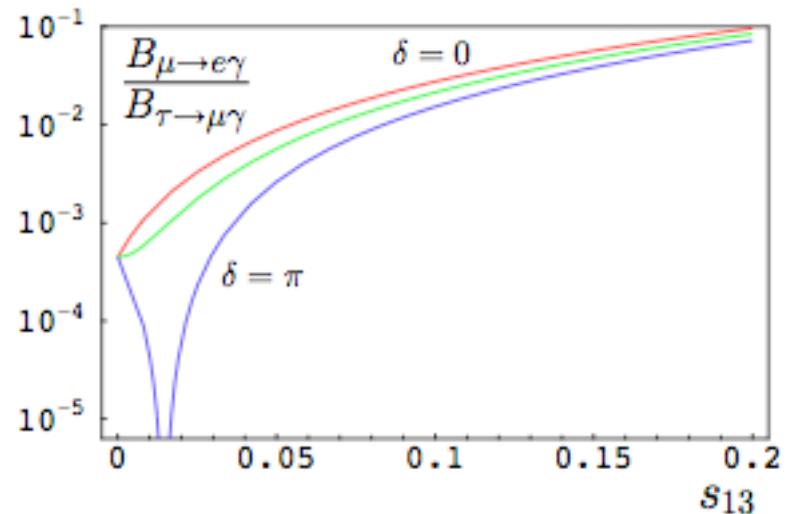


$$\begin{aligned}
 \mathcal{L}_\Delta = & \cdots + (D_\mu \Delta)^\dagger (D^\mu \Delta) - M_\Delta^2 \Delta^\dagger \Delta + + \\
 & + Y_\Delta^{\alpha\beta} \overline{\tilde{\mathbf{L}}} (\tau \cdot \Delta) \mathbf{L} + \mu_\Delta \tilde{\phi}^\dagger (\tau \cdot \Delta)^\dagger \phi + \dots
 \end{aligned}$$

## Correlations among weak processes, i.e.

$\mu \rightarrow e\gamma / \tau \rightarrow e\gamma / \tau \rightarrow \mu\gamma$

- \* Neutrino masses OK
- \* Measurable flavour OK
- \* Predictivity OK



V. Cirigliano, B. Grinstein, G. Isidori, M. Wise, hep-ph/0507001.  
 M. B. Gavela, T. Hambye, P. Hernández, D.H., 0906.1461

# Successful fermionic-mediated Seesaws:

One more mediator, one more scale.... i.e. Inverse seesaws

Instead of  $\mathcal{L}_m = \begin{pmatrix} 0 & Y_N^T v \\ Y_N v & M_N \end{pmatrix}$

# Successful fermionic-mediated Seesaws:

One more mediator, one more scale.... i.e. Inverse seesaws

$$\mathcal{L}_{M_\nu} = \begin{pmatrix} \bar{L}_i & \bar{N}^c & \bar{N}'^c \\ 0 & Y_N^T v & 0 \\ Y_N v & 0 & \Lambda^T \\ 0 & \Lambda & 0 \end{pmatrix}$$

# Successful fermionic-mediated Seesaws:

One more mediator, one more scale.... i.e. Inverse seesaws

$$\mathcal{L}_{M_\nu} = \begin{pmatrix} \bar{L}_i & \bar{N}^c & \bar{N}'^c \\ 0 & Y_N^T v & 0 \\ Y_N v & 0 & \Lambda^T \\ 0 & \Lambda & 0 \end{pmatrix}$$

*Lepton number conserved*

U(1)

$$\boxed{\begin{aligned} \Lambda_{\text{fl}} &= \Lambda \\ \Lambda_{\text{LN}} &= \infty \end{aligned}}$$

Wyler+Wolfenstein; Mohapatra+Valle; Branco+Grimus+Lavoura....

# Successful fermionic-mediated Seesaw:

One more mediator, one more scale.... i.e. Inverse seesaws

$$\mathcal{L}_{M_\nu} = \begin{pmatrix} \bar{L}_i & \bar{N}^c & \bar{N}'^c \\ 0 & Y_N^T v & \epsilon Y_N'^T v \\ Y_N v & \mu' & \Lambda^T \\ \epsilon Y_N' v & \Lambda & \mu \end{pmatrix}$$

*Lepton number violated  
by any of those 3 entries*

# Successful fermionic-mediated Seesaw:

One more mediator, one more scale.... i.e. Inverse seesaws

$$\mathcal{L}_{M_\nu} = \begin{pmatrix} \bar{L}_i & \bar{N}^c & \bar{N}'^c \\ 0 & Y_N^T v & \epsilon Y_N'^T v \\ Y_N v & \mu' & \Lambda^T \\ \epsilon Y_N' v & \Lambda & \mu \end{pmatrix}$$

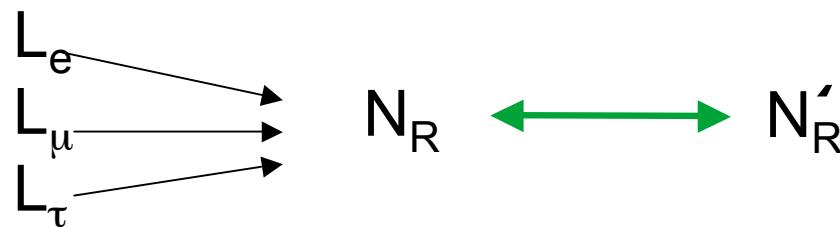
*Lepton number violated  
by any of those 3 entries*

$\Lambda$  may be  $\sim$  TeV and  $Y$ s  $\sim 1$ , and be ok with  $m_v$

Case: Three light active families + one  $N_R$  + one  $N'_R$

$$\mathcal{L}_{M_\nu} = \begin{pmatrix} \bar{L}_i & \bar{N}^c & \bar{N}'^c \\ 0 & Y_N^T v & \epsilon Y_N'^T v \\ Y_N v & \mu' & \Lambda^T \\ \epsilon Y_N' v & \Lambda & \mu \end{pmatrix}$$

$\mu'$  is irrelevant (at tree-level)

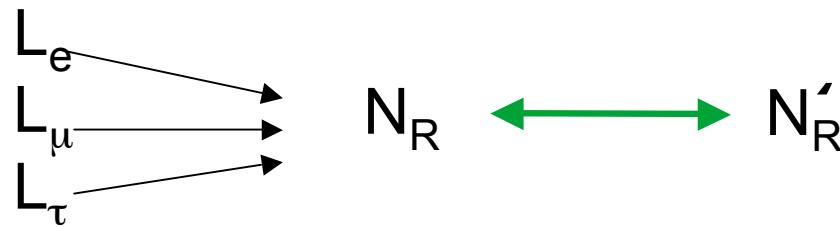


- one massless neutrino
- just one low-energy Majorana phase

Case: Three light active families + one  $N_R$  + one  $N'_R$

$$\mathcal{L}_{M_\nu} = \begin{pmatrix} \bar{L}_i & \bar{N}^c & \bar{N}'^c \\ 0 & Y_N^T v & \epsilon Y_N'^T v \\ Y_N v & \mu' & \Lambda^T \\ \epsilon Y_N' v & \Lambda & \mu \end{pmatrix}$$

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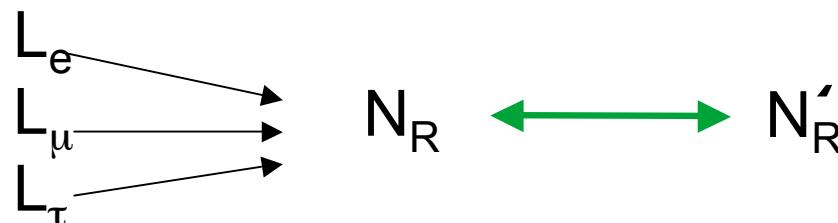
- one massless neutrino
- just one low-energy Majorana phase

*arguably the simplest model of neutrino mass*

Case: Three light active families + one  $N_R$  + one  $N'_R$

$$\mathcal{L}_{M_\nu} = \begin{pmatrix} \bar{L}_i & \bar{N}^c & \bar{N}'^c \\ 0 & Y_N^T v & \epsilon Y_N'^T v \\ Y_N v & \mu' & \Lambda^T \\ \epsilon Y_N' v & \Lambda & \mu \end{pmatrix}$$

$\mu'$  is irrelevant (at tree-level)



### FUNDAMENTAL

	moduli	phases
$Y_N$	3	3
$Y'_N$	3	3
$\Lambda$	1	1

vs

### LOW ENERGY

- 3 angles and 2 phases in the  $U_{PMNS}$
- 2 masses and 0 phases in  $M_\nu$
- 2 overall factors and 5 phases absorbed.

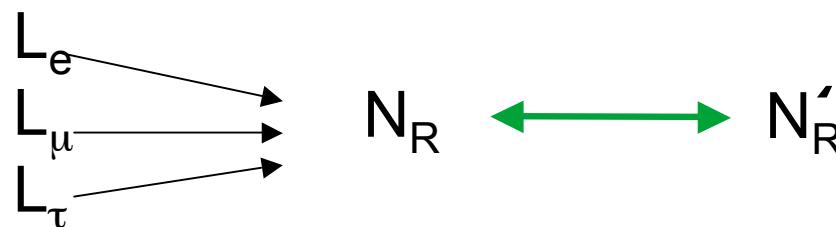
- A normalization factor apart, Yukawas are determined from the  $U_{PMNS}$  and neutrino masses!

Hambye, Hernandez<sup>2</sup>, B.G. 09

Case: Three light active families + one  $N_R$  + one  $N'_R$

$$\mathcal{L}_{M_\nu} = \begin{pmatrix} \bar{L}_i & \bar{N}^c & \bar{N}'^c \\ 0 & Y_N^T v & \epsilon Y_N'^T v \\ Y_N v & \mu' & \Lambda^T \\ \epsilon Y_N' v & \Lambda & \mu \end{pmatrix}$$

$\mu'$  is irrelevant (at tree-level)



i.e.

$$Y_N^T \simeq y \begin{pmatrix} e^{i\delta} s_{13} + e^{-i\alpha} s_{12} r^{1/4} \\ s_{23} \left( 1 - \frac{\sqrt{r}}{2} \right) + e^{-i\alpha} r^{1/4} c_{12} c_{23} \\ c_{23} \left( 1 - \frac{\sqrt{r}}{2} \right) - e^{-i\alpha} r^{1/4} c_{12} s_{23} \end{pmatrix} \quad r = \frac{|\Delta m_{12}^2|}{|\Delta m_{13}^2|}$$

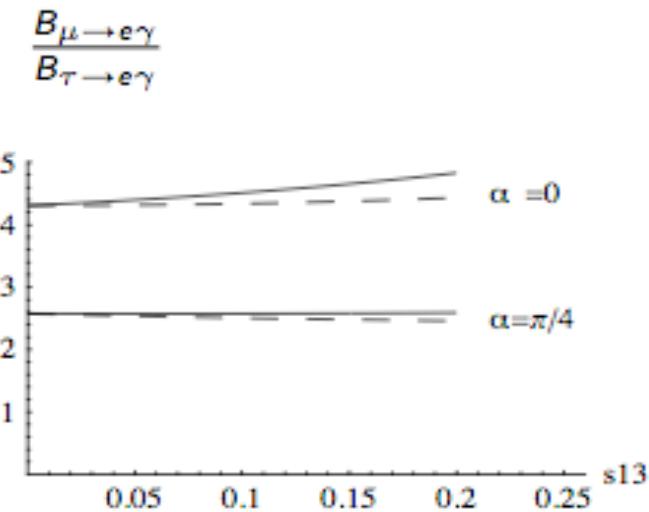
Normal hierarchy

- A normalization factor apart, Yukawas are determined from the  $U_{PMNS}$  and neutrino masses!

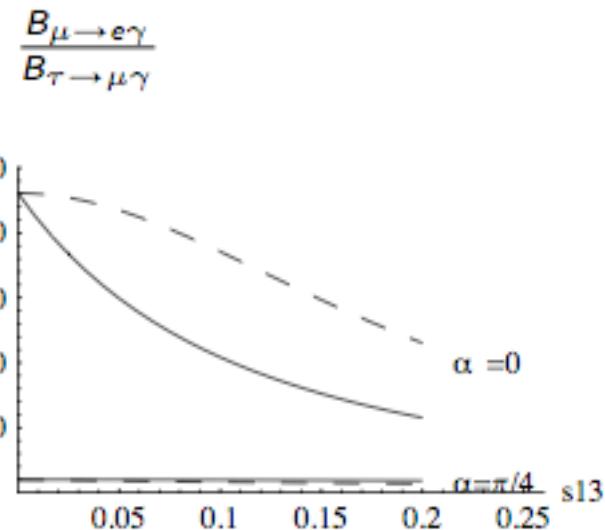
Hambye, Hernandez<sup>2</sup>, B.G. 09



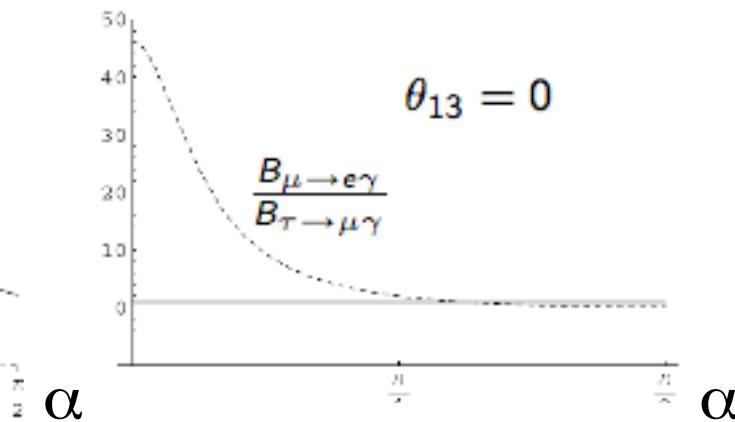
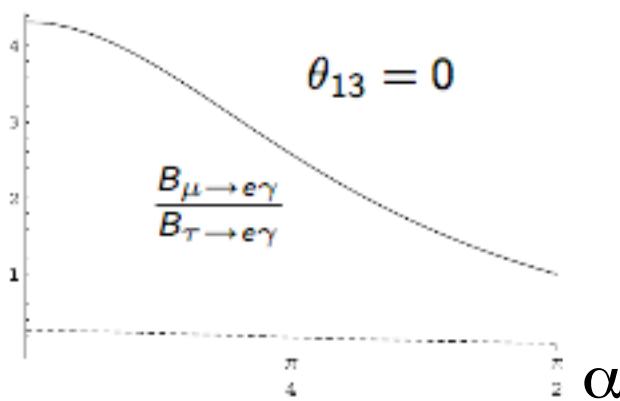
## NORMAL HIERARCHY



## INVERTED HIERARCHY



Strong dependence on the Majorana phase!



## Light new scales    $M \ll v$

Most bounds on non-unitarity do NOT apply (i.e.  $M < \text{GeV}$ )

Can oscillation experiments detect/bound light sterile  $\nu$ s?

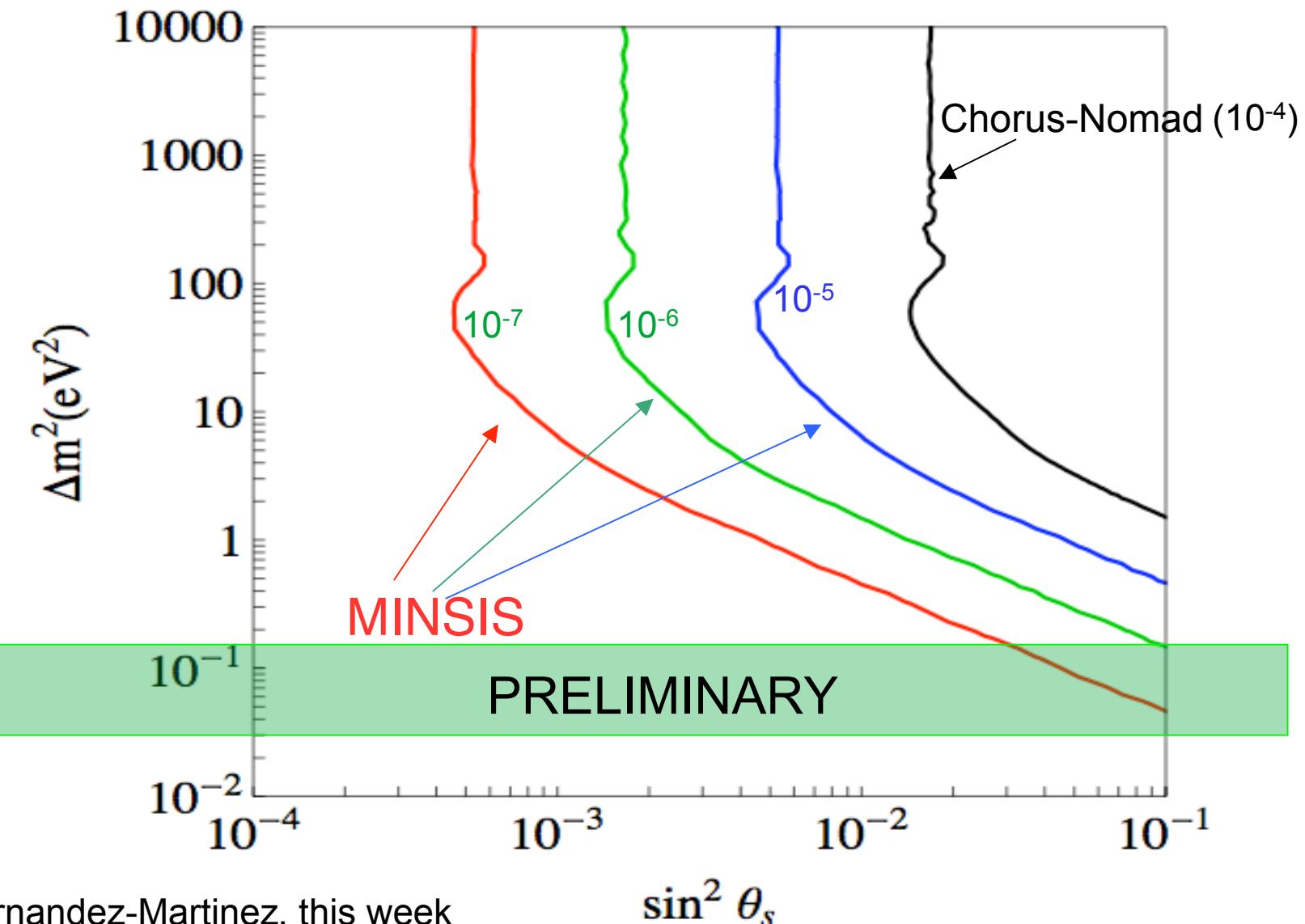
Yes, mainly with near detectors

\* Disappearance i.e. MINOS

\* Appearance i.e. MINSIS

i.e for a mixing 3-sterile  $\theta_s$

$$P(\nu_\mu - \nu_\tau) = \cos^2(\theta_{12}) \cdot \sin^4(\theta_s) \cdot \sin^2(\Delta m^2 L / 4E)$$



# Conclusions

- \* The guideline is clear:  $m_\nu + \text{Majorana} + \theta_{13} \rightarrow \text{CP}$
- The **seesaw mechanism** induces:  
non-unitary mixing matrix; new CP channels; rare decays.

Difficult signals, but they should be there if:

- $M \sim \text{TeV}$  (also at LHC)
- Yukawas near  $\sim O(1)$

- \* Near/short distance detectors are called for

$\nu_e \leftrightarrow \nu_\tau$  : from  $\nu$  factory (*and/or betabeams ? ...*)

$\nu_\mu \leftrightarrow \nu_\tau$  : from present beams ? (*and/or  $\nu$  factory...*)

and may shed light on TeV seesaws... or light sterile  $\nu$ s

# Back-up slides

# $N$ elements from oscillations only

- **KamLAND+CHOOZ+Atmospheric+K2K**
- + Near detectors: MINOS, NOMAD, BUGEY, KARMEN

without unitarity  
OSCILLATIONS  $|\varepsilon| = \begin{cases} < 0.025 & < 0.05 & < .34 \\ < 0.05 & < -0.05 & < 0.013 \\ < 0.09 & < 0.013 & ? \end{cases}$

# Scalar triplet seesaw

Bounds on  $c^{d=6}$

Process	Constraint on	Bound ( $\times (\frac{M_\Delta}{1\text{TeV}})^2$ )
$M_W$	$ Y_{\Delta \mu e} ^2$	$< 7.3 \times 10^{-2}$
$\mu^- \rightarrow e^+ e^- e^-$	$ Y_{\Delta \mu e}   Y_{\Delta ee} $	$< 1.2 \times 10^{-5}$
$\tau^- \rightarrow e^+ e^- e^-$	$ Y_{\Delta \tau e}   Y_{\Delta ee} $	$< 1.3 \times 10^{-2}$
$\tau^- \rightarrow \mu^+ \mu^- \mu^-$	$ Y_{\Delta \tau \mu}   Y_{\Delta \mu \mu} $	$< 1.2 \times 10^{-2}$
$\tau^- \rightarrow \mu^+ e^- e^-$	$ Y_{\Delta \tau \mu}   Y_{\Delta ee} $	$< 9.3 \times 10^{-3}$
$\tau^- \rightarrow e^+ \mu^- \mu^-$	$ Y_{\Delta \tau e}   Y_{\Delta \mu \mu} $	$< 1.0 \times 10^{-2}$
$\tau^- \rightarrow \mu^+ \mu^- e^-$	$ Y_{\Delta \tau \mu}   Y_{\Delta \mu e} $	$< 1.8 \times 10^{-2}$
$\tau^- \rightarrow e^+ e^- \mu^-$	$ Y_{\Delta \tau e}   Y_{\Delta \mu e} $	$< 1.7 \times 10^{-2}$
$\mu \rightarrow e \gamma$	$ \Sigma_{l=e,\mu,\tau} Y_{\Delta l \mu}^\dagger Y_{\Delta e l} $	$< 4.7 \times 10^{-3}$
$\tau \rightarrow e \gamma$	$ \Sigma_{l=e,\mu,\tau} Y_{\Delta l \tau}^\dagger Y_{\Delta e l} $	$< 1.05$
$\tau \rightarrow \mu \gamma$	$ \Sigma_{l=e,\mu,\tau} Y_{\Delta l \tau}^\dagger Y_{\Delta \mu l} $	$< 8.4 \times 10^{-1}$

# Scalar triplet seesaw

Combined bounds on  $c^{d=6}$

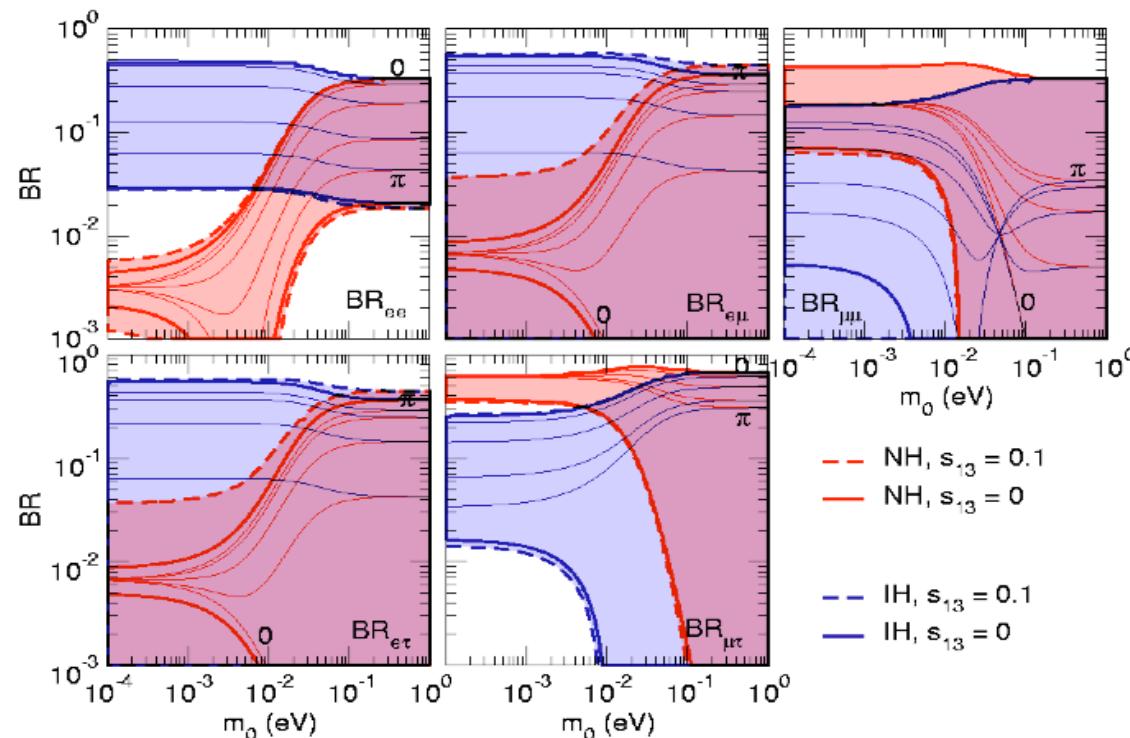
Combined bounds		
Process	Yukawa	Bound $\left( \times \left( \frac{M_\Delta}{1 \text{ TeV}} \right)^4 \right)$
$\mu \rightarrow e\gamma$	$ Y_{\Delta_{\mu\mu}}^\dagger Y_{\Delta_{\mu e}} + Y_{\Delta_{\tau\mu}}^\dagger Y_{\Delta_{\tau e}} $	$< 4.7 \times 10^{-3}$
$\tau \rightarrow e\gamma$	$ Y_{\Delta_{\tau\tau}}^\dagger Y_{\Delta_{\tau e}} $	$< 1.05$
$\tau \rightarrow \mu\gamma$	$ Y_{\Delta_{\tau\tau}}^\dagger Y_{\Delta_{\tau\mu}} $	$< 8.4 \times 10^{-1}$

# $M_\Delta \sim \text{TeV}$ : direct searches at LHC ?

See-saw II: Pair-production of charged triplet scalars

$$pp \rightarrow \Delta^{++} \Delta^{--} \rightarrow l^+ l^+ l^- l^-$$

Flavour structure one-to-one to  $m_\nu$  !  $\text{BR}(\Delta^{++} \rightarrow l_\alpha^+ l_\beta^+) \sim |M_{\alpha\beta}|^2$



Han et al; Garayoa, Schwetz; Kadastik, et al ; Akeroyd, et al; Fileviez et al

## Normal hierarchy:

Up to terms of  $\mathcal{O}(\sqrt{r}, s_{13})$ , we find

$$Y_N^T \simeq y \begin{pmatrix} e^{i\delta} s_{13} + e^{-i\alpha} s_{12} r^{1/4} \\ s_{23} \left(1 - \frac{\sqrt{r}}{2}\right) + e^{-i\alpha} r^{1/4} c_{12} c_{23} \\ c_{23} \left(1 - \frac{\sqrt{r}}{2}\right) - e^{-i\alpha} r^{1/4} c_{12} s_{23} \end{pmatrix} \quad r = \frac{|\Delta m_{12}^2|}{|\Delta m_{13}^2|}$$

## Inverted hierarchy:

$$Y_N^T \simeq \frac{y}{\sqrt{2}} \begin{pmatrix} c_{12} e^{i\alpha} + s_{12} e^{-i\alpha} \\ c_{12} (c_{23} e^{-i\alpha} - s_{23} s_{13} e^{i(\alpha-\delta)}) - s_{12} (c_{23} e^{i\alpha} + s_{23} s_{13} e^{-i(\alpha+\delta)}) \\ -c_{12} (s_{23} e^{-i\alpha} + c_{23} s_{13} e^{i(\alpha-\delta)}) + s_{12} (s_{23} e^{i\alpha} - c_{23} s_{13} e^{-i(\alpha+\delta)}) \end{pmatrix}$$

Hence, measuring  $U_{PMNS}$ , and  $\nu$  masses determines  
**EVERYTHING!!**

$$B_{\mu \rightarrow e\gamma} \propto |Y_{N_e} Y_{N_\mu}|^2, \quad |m_{ee}|_{IH} \simeq |s_{12}^2 e^{-2i\alpha} - c_{12}^2 e^{2i\alpha}|$$