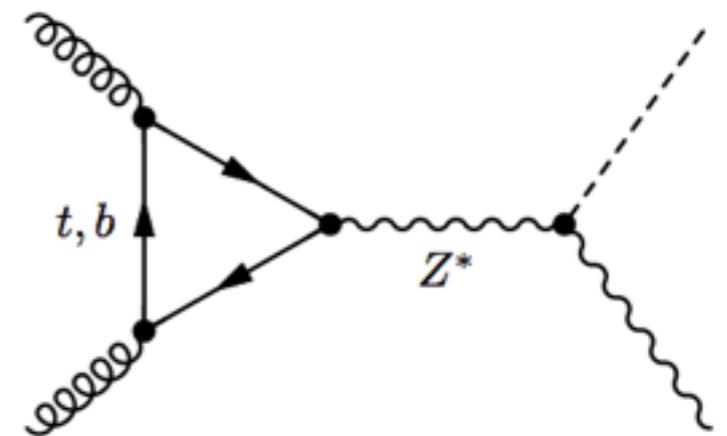
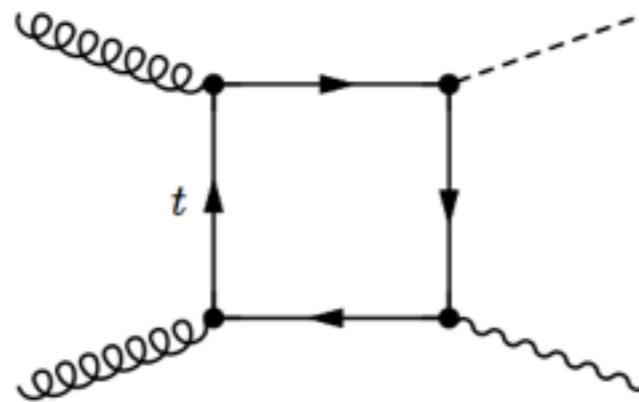
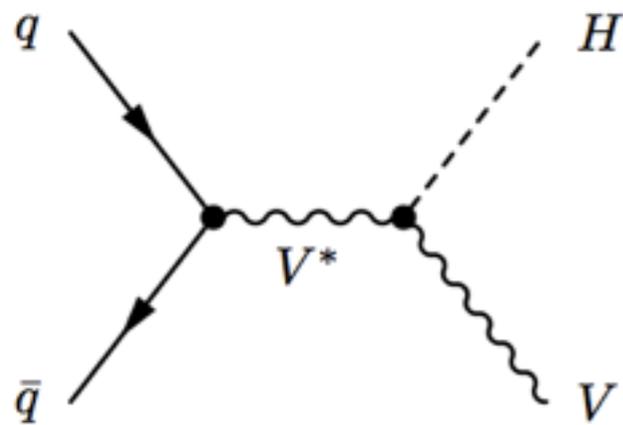


# WG1:VH experimental view

John Campbell, Francesco Tramontano, Luca Perrozzi, CP

13th Workshop of the LHC Higgs Cross Section Working Group

Cern, 13-14 July 2017



UNIVERSITÉ  
DE GENÈVE

# Organization of the VH subgroup

**Re-organization post-YR4: VH and VBF WG1 subgroups are now splitted**  
(but still in strict collaboration on common issues and general topics...!)

**VH twiki:** <https://twiki.cern.ch/twiki/bin/view/LHCPhysics/LHCHXSWGvH> (work in progress)

## Mailing lists

- ▶ [lhc-higgs-xsbr@cern.ch](mailto:lhc-higgs-xsbr@cern.ch)  
[general WG1 thread - for discussions / meeting advertisement]
- ▶ [lhc-higgs-vh-convener@cern.ch](mailto:lhc-higgs-vh-convener@cern.ch)  
[convener mailing list - for direct communication]

## Indico page for VH WG1 meetings:

<https://indico.cern.ch/category/5847/>

## First meeting on June 29, 2017:

several contributions and very fruitful discussions  
(many of which covered in Francesco's talk)

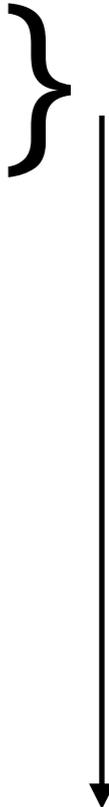
The screenshot shows the Indico meeting agenda for the WG1 - VH subgroup on Thursday, June 29, 2017, from 14:00 to 16:00. The agenda is as follows:

Time	Topic	Speaker	Duration
14:00 - 14:10	News and general overview	Speakers: Carlo Enrico Pandini (Università di Genova (IIG)), Francesco Tramontano (Università di INFN, Napoli (IT)), John Campbell (FNU), Luca Perrotti (Eidgenössische Technische Hochschule Zürich (CH))	10m
14:10 - 14:30	Towards ggZH at NLO	Speaker: Sophia Carola Borowka (University of Oxford (UK))	20m
14:30 - 14:50	VH(+bb) at NNLO QCD	Speakers: Giancarlo Ferrera (University of Milan), Giancarlo Ferrera (Università degli Studi e INFN Milano (IT))	20m
14:50 - 15:10	EWK NLO corrections for VH using POWHEG+MINLO	Speaker: Carlo Oleari (Università e INFN, Milano-Brescia (IT))	20m
15:10 - 15:30	Updates on NNLOPS for VH	Speakers: Mr. Wojciech Skoczko (University of Oxford (UK)), Wojciech Jozef Skoczko	20m
15:30 - 15:50	Simplified Template Cross Sections for VH	Speaker: Frank Tackmann (Deutsches Elektronen-Synchrotron (DE))	20m

# Outline and Introduction

## Swift start of activities in the VH WG1 subgroup

first topics considered to design a roadmap for the future (short / medium / long term):

- ▶ precise modeling of VH processes @NNLO[QCD] and @NLO[EW]
  - ▶ treatment of loop-induced  $gg \rightarrow ZH$
  - ▶ latest experimental VH results
    - state of VH predictions / tools used by experimental collaborations
    - main theory limitations & “whishlist”, possible improvements
  - ▶ Simplified Template Cross Section STXS approach
  - ▶ open point: treatment of main SM backgrounds for VH measurements
- 

**This talk:** experimental take on VH matters, in light of recent results and towards the full Run-2 analyses

**Francesco's talk:** overview from the theory side - several contributions from the first VH subgroup meeting

# VH signal model in ATLAS & CMS

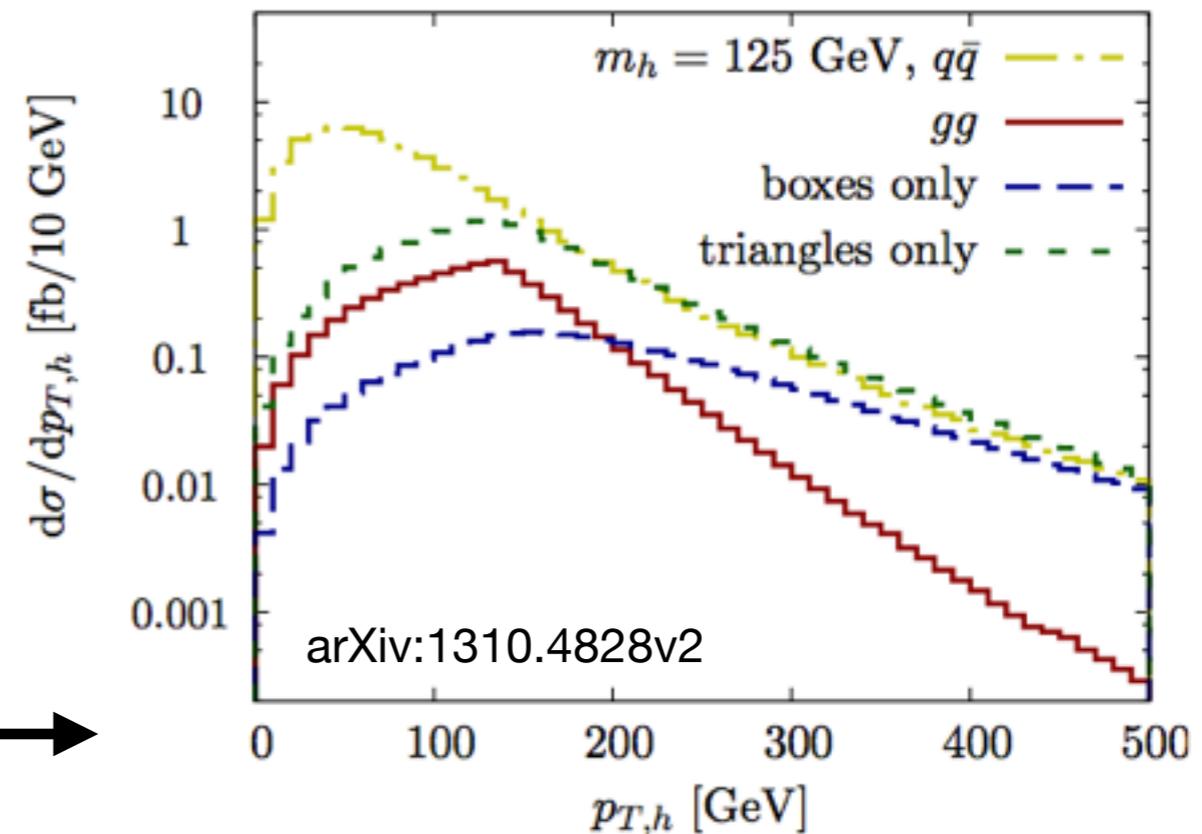
Monte Carlo events for VH processes

qqZH, WH	Powheg-MiNLO + Pythia8	MiNLO[QCD]
loop-induced $gg \rightarrow ZH$	Powheg + Pythia8	LO[QCD]

- ▶ PDF4LHC15 set for matrix-element
- ▶ dedicated ATLAS / CMS PS tunes

Cross-section predictions from YR4:

- ▶ qqZH and WH: **NNLO[QCD]+NLO[EW]** including photon-induced (3% in WH, 1% ZH) and top-loop induced (1%) contributions
- ▶ **loop-induced  $gg \rightarrow ZH$ : NLO(approx)+NLL[QCD]** kNLO $\sim 2$  from ( $m_{top} \rightarrow \infty$ ) calculation



EW Corrections from YR4:

- ▶ NLO EW differential reweighting (generally applied as  $f(p_T^V)$ ) from HAWK

$X_S(gg \rightarrow ZH)$  O(14%) of total  $X_S(pp \rightarrow ZH)$   
enhanced contribution at medium-high  $p_T^V$

Future EW NLO accurate MC generators (Francesco's talk)

# VH signal model in ATLAS & CMS

Monte Carlo events for VH processes

qqZH, WH

Powheg-MiNLO + Pythia8

MiNLO[QCD]

loop-induced  $gg \rightarrow ZH$

Powheg + Pythia8

LO[QCD]

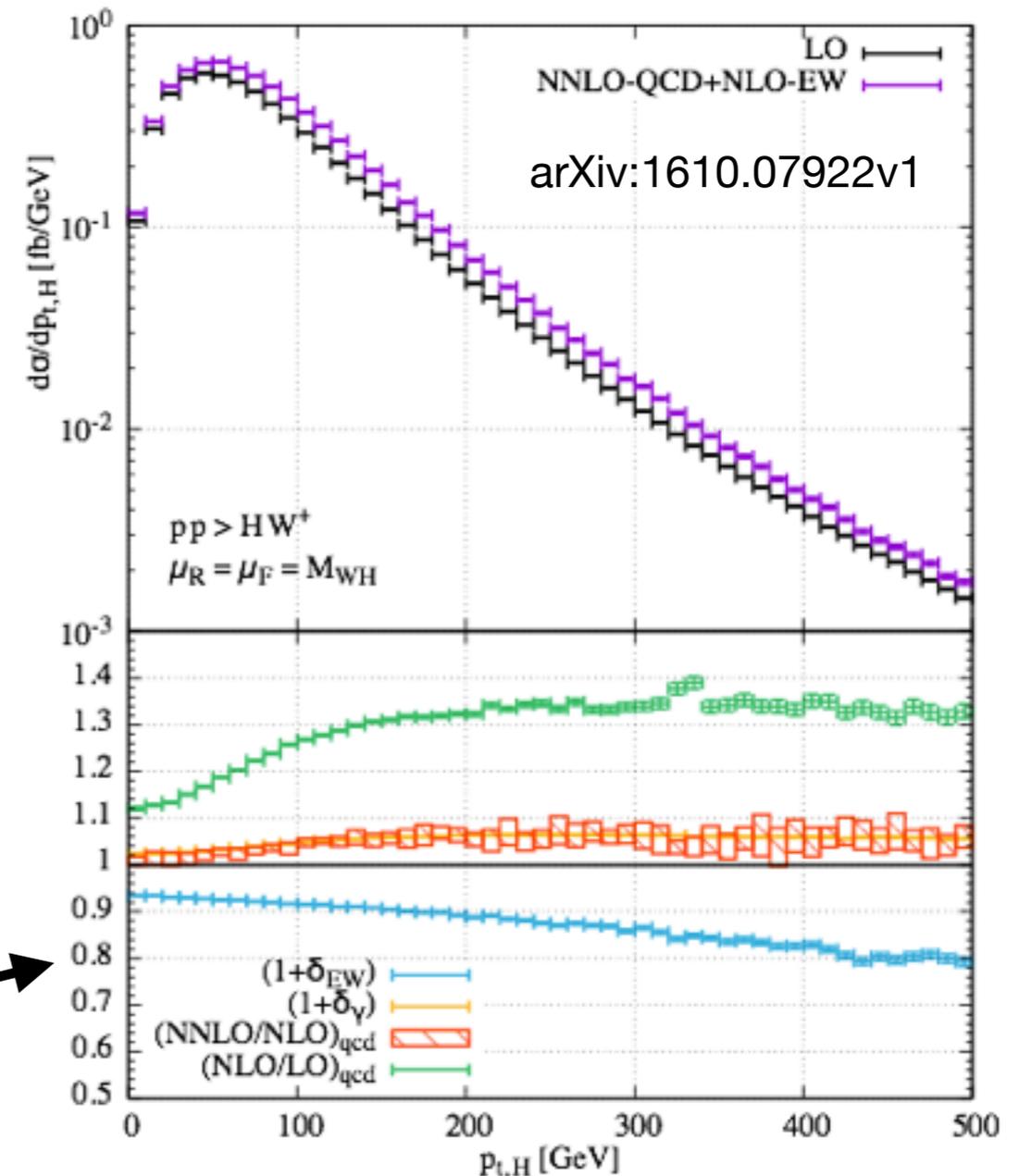
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Future EW NLO accurate MC generators (Francesco's talk)

# VH production @ LHC

**Most relevant decay channel for VH production:  $H(\rightarrow bb)$**

(other channels are sensitive to VH -  $H(\rightarrow bb)$  relies on  $V(\text{lept.})H$  production **almost** entirely)

**From YR4:** main selections and cuts coming from typical  $VH(\rightarrow bb)$  requirement (+case study)

<b>VH(bb)</b>	significance obs(exp) [ $\sigma$ ]	signal strength $\mu = \sigma/\sigma_{SM}$
<b>ATLAS Run-1</b>	1.4 (2.6)	$0.5 \pm 0.4$ ( $\pm 0.31$ stat $\pm 0.24$ syst)
<b>CMS Run-1</b>	2.1 (2.5)	$0.9 \pm 0.4$
<b>ATLAS Run-2</b>	3.5 (3.0)	$1.2 \pm 0.4$

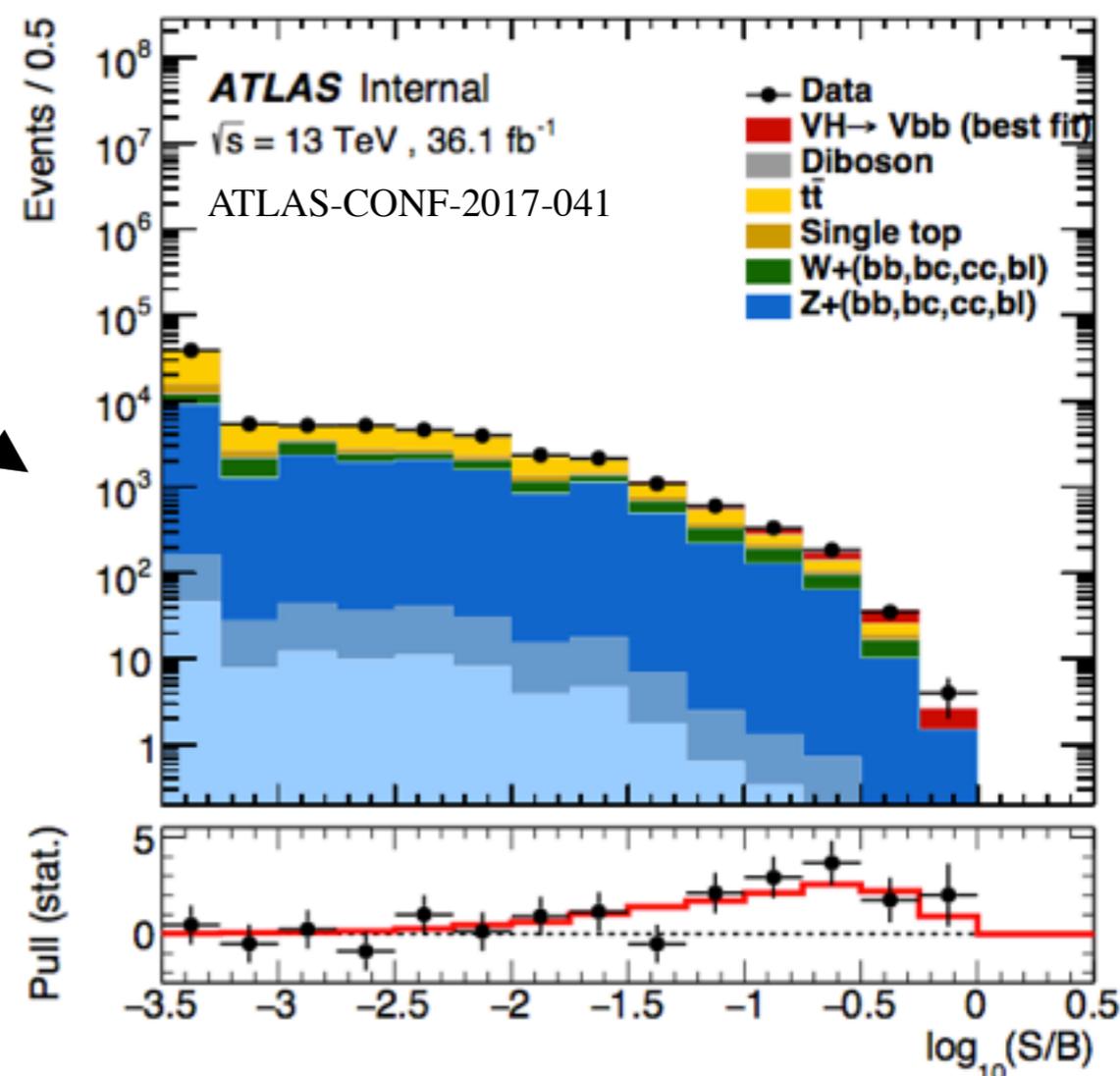
**Main channels:**  $V(\rightarrow \text{leptons: } ll, lv, vv) + H(bb)$

**Standard selection cuts:**

- ▶ 2 central jets ( $|\eta| < 2.5$ )  $p_T > 20\text{GeV}$   
reconstruct  $H(bb)$  candidate
- ▶  $n_{\text{Jet}} \leq 3$  for  $Z(vv)H$  and  $W(lv)H$  channels
- ▶  $p_T^V > 150\text{GeV}$  [ $+75\text{-}150\text{GeV}$  for  $Z(ll)H$ ]
- ▶ multivariate approach MVA

Sensitive region relatively boosted:  $O(150\text{-}250)\text{GeV}$

## VH(bb) evidence @ 13 TeV



*VH(bb) signal strength @ 13 TeV [ $L_{int} = 36/\text{fb}$ ] =  $1.2 \pm 0.24(\text{stat}) \pm 0.31(\text{syst})$*

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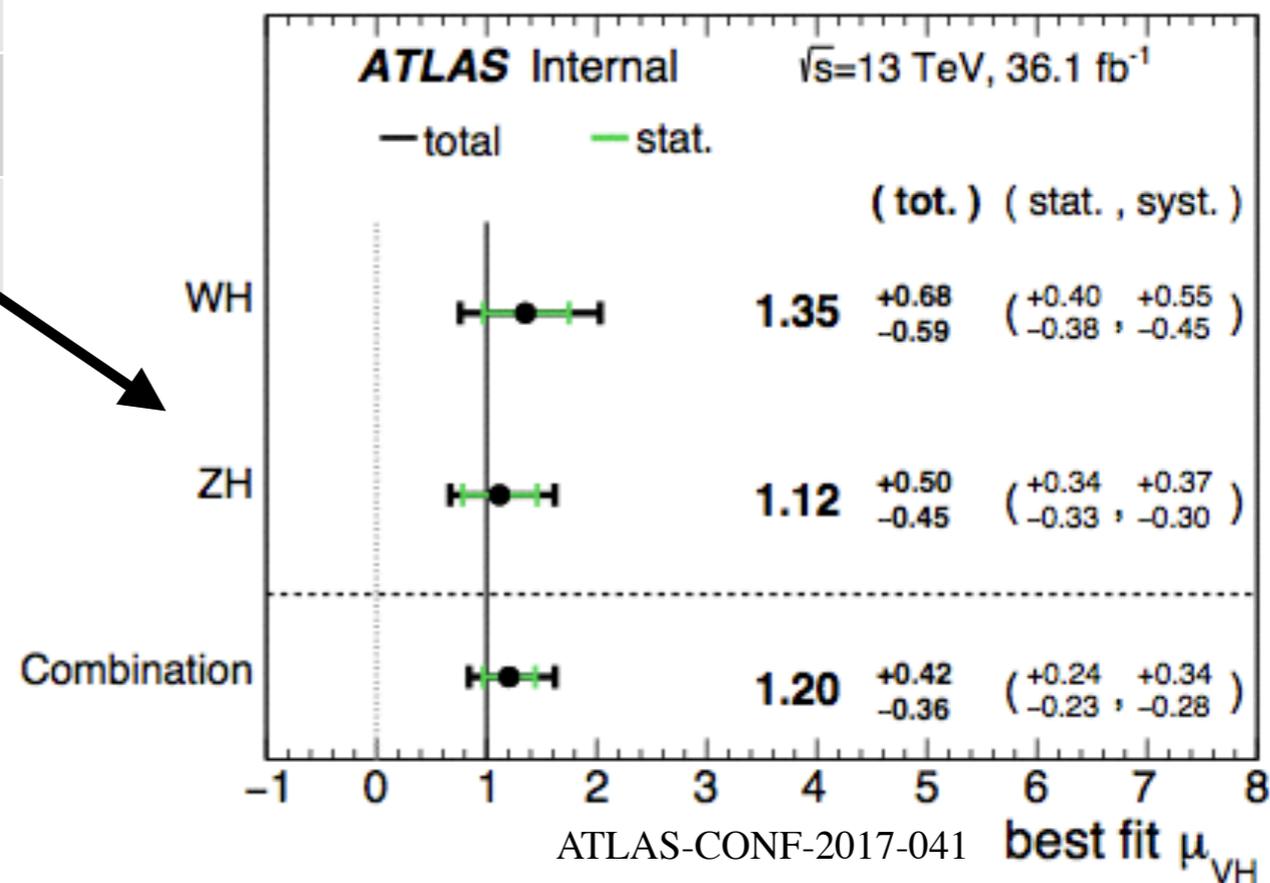
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# VH production @ 13TeV - signal modeling

VH(bb) signal strength @ 13 TeV [ $L_{int} = 36/\text{fb}$ ] =  $1.2 \pm 0.24(\text{stat}) \pm 0.31(\text{syst})$

## Important role of VH signal systematic uncertainties in VH(bb) analyses:

- UEPS acceptance uncertainties**  
 “experimental recipe”:  
 Pythia8 vs Herwig7 +  
 PS eigentune variations
- QCD scale variations**  
 (total XS from YR4, acceptance from  
 Powheg-MiNLO variations)

Good understanding of the VH signal model becomes even more critical with additional data:

$L_{int}$	stat. error on $\mu = \sigma/\sigma_{SM}$	syst. error on $\mu = \sigma/\sigma_{SM}$
36/fb	0.24	0.31
150/fb	0.12	?

Source of uncertainty	$\sigma_\mu$
Total	0.39
Statistical	0.24
<b>Systematic</b>	<b>0.31</b>
Experimental uncertainties	
Jets	0.03
$E_T^{\text{miss}}$	0.03
Leptons	0.01
$b$ -tagging	0.09
<b><math>b</math>-jets</b>	0.09
<b><math>c</math>-jets</b>	0.04
light jets	0.04
extrapolation	0.01
Pile-up	0.01
Luminosity	0.04
Theoretical and modelling uncertainties	
<b>Signal</b>	<b>0.17</b>
Floating normalizations	0.07
Z+jets	0.07
W+jets	0.07
$t\bar{t}$	0.07
Single-top	0.08
Diboson	0.02
Multijet	0.02
MC statistics	0.13

# VH production @ 13TeV - signal modeling

VH(bb) signal strength @ 13 TeV [ $L_{int} = 36/\text{fb}$ ] =  $1.2 \pm 0.24(\text{stat}) \pm 0.31(\text{syst})$

## Important role of VH signal systematic uncertainties in VH(bb) analyses:

1. **UEPS acceptance uncertainties**  
“experimental recipe”:  
Pythia8 vs Herwig7 +  
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2. **QCD scale variations**  
(total XS from YR4, acceptance from  
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Good understanding of the VH signal model becomes even more critical with additional data:

$L_{int}$	stat. error on $\mu = \sigma/\sigma_{SM}$	syst. error on $\mu = \sigma/\sigma_{SM}$
36/fb	0.24	0.31
150/fb	0.12	?



### “PS uncertainties”

#### Long-standing topic

Possible improvements in the treatment of UE/PS/MPI/had. uncertainties?

### QCD scale variations

possible improvements from NNLOPS applied to Powheg-MiNLO & improved prediction for  $gg \rightarrow ZH$

(Francesco's talk)

# VH production @ Run-1 - signal modeling

## Snapshot from Run-1 - main VH theory limitations for VH searches

- ▶ VH signal modeling:  $O(5\%)$  on  $\mu = \sigma/\sigma_{SM}$  - not limiting but still relevant
- ▶ dominated by UEPS uncertainties and QCD scale variations

VH(bb)	significance obs(exp) [ $\sigma$ ]	signal strength $\mu = \sigma/\sigma_{SM}$
ATLAS Run-1	1.4 (2.6)	$0.5 \pm 0.4$ ( $\pm 0.31$ stat $\pm 0.24$ syst)
CMS Run-1	2.1 (2.5)	$0.9 \pm 0.4$

Signal	
Cross section (scale)(3:ggZH)	1% (qq), 50% (gg)
Cross section (PDF)	2.4% (qq), 17% (gg)
Branching Ratio (5)	3.3 %
Acceptance (scale)	1.5–3.3%
3-jet acceptance (scale)	3.3–4.2%
$p_T^V$ shape (scale)(2)	S
Acceptance (PDF) (4)	2–5%
$p_T^V$ shape (NLO EW correction)	S
Acceptance (parton shower)(1)	8–13%

Source	Type	Event yield uncertainty range (%)	Individual contribution to $\mu$ uncertainty (%)	Effect of removal on $\mu$ uncertainty (%)
Signal cross section (scale and PDF)	norm.	4	3.9	0.3
Signal cross section ( $p_T$ boost, EW/QCD)	norm.	2/5	3.9	0.3

### Total uncertainty on VH(bb) XS of $O(3-5\%)$

- ▶ significant effect when considering acceptance variations across **nJet** or  $p_T^V$  regions: feedback on current techniques (e.g. Stewart-Tackmann), possible improvements
- ▶ parton-shower effects can be large and difficult to assess properly

→ Possible to provide guidelines to address these points?

# Background studies for VH analyses

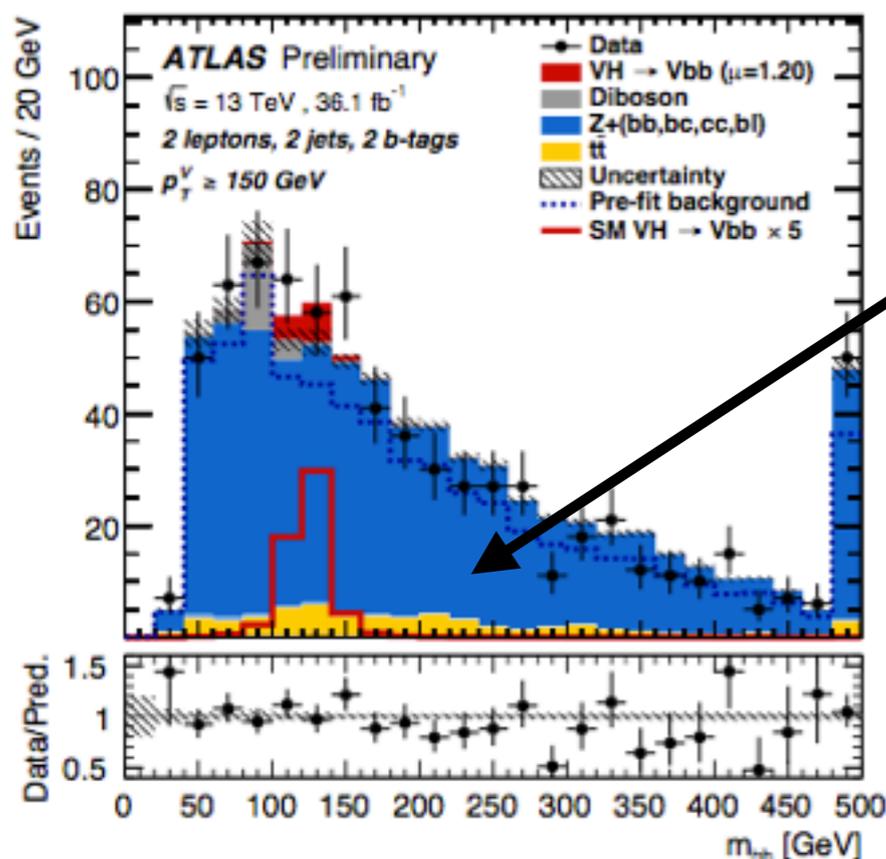
**Systematic** bkg studies, for all VH searches are **beyond the scope of the LHC HXSWG**

**However:** main VH *user/contributor* [**VH(bb)**] can be quite sensitive to background modeling

## V+heavy-flavor background:

The WG1:VH subgroup can be the forum to gather discussions on modeling issues for this major background, specifically in the VH(bb) phase-space

- ▶ guidelines on the impact of MC generator choices on experimental analyses: LO vs NLO, 4F vs 5F, matching options, alternative MC generators, ...
- ▶ impact of higher-order prediction in the specific phase-space
- ▶ estimate of theoretical uncertainties related to VH(bb) selections



*Large Z(ll)+heavy-flavor contribution  $m(bb)$*   
background prediction is improved by fitting to data

However, accurate MC modeling is beneficial when training complex MVA analyses

Possible to extend to a future discussion on **VZ(bb) diboson background** modeling: standard candle used by experimental analysis to validate VH(bb) searches / measurements

# Simplified Template Cross Sections - VH

## Moving towards STXS approach for VH measurements:

- experiments will provide results in STXS bins (stage0 → stage1)
- theorists will provide estimates of theory uncertainties in STXS bins (and afterwards, of course, interpretations!)

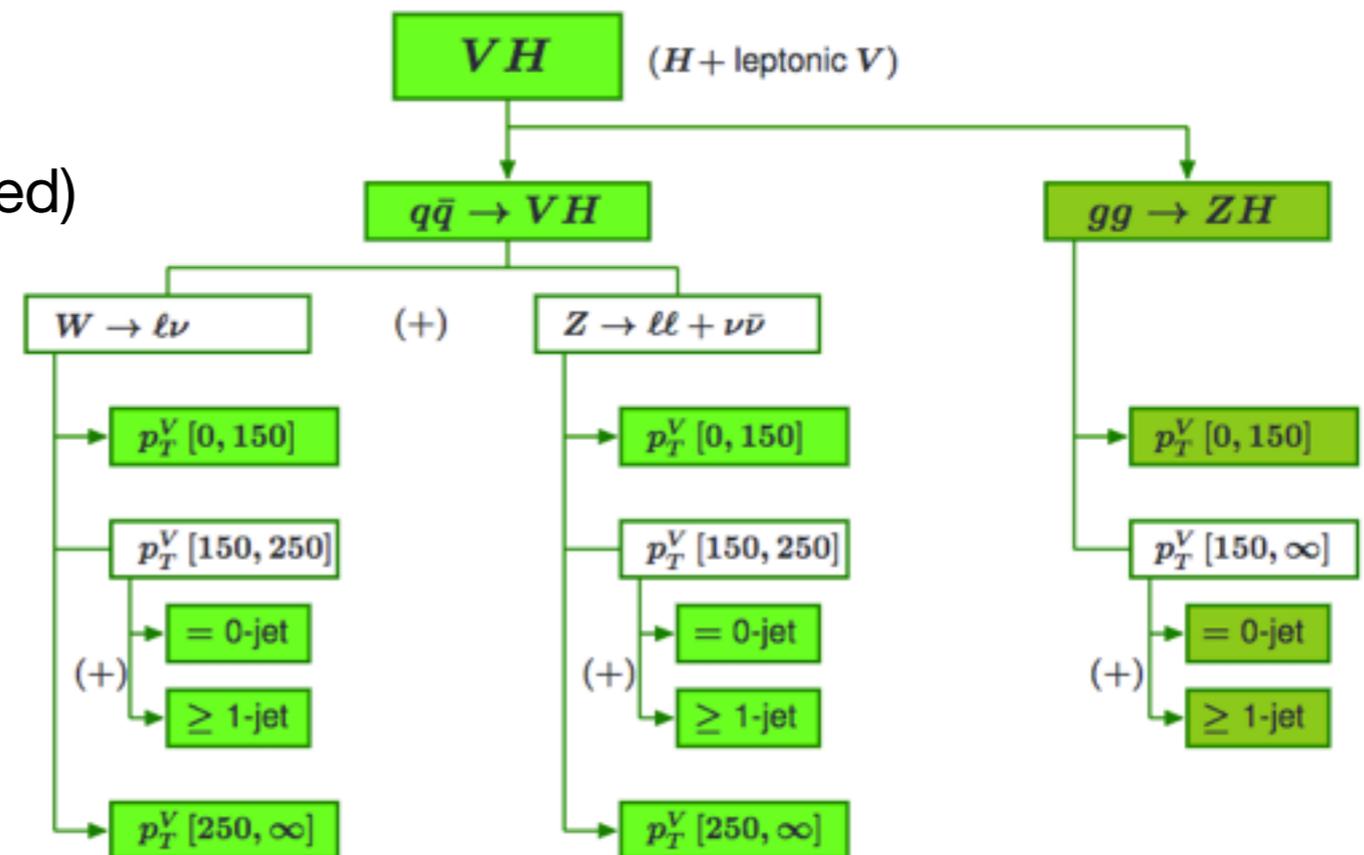
*Strong theory-experiment collaboration!*

## VH:STXS categorization

- ▶ “VH” bins include leptonic VH (H undecayed)
- ▶  $qq \rightarrow V(qq)H$  as part of “VBF” bins
- ▶  $gg \rightarrow Z(qq)H$  as part of “ggF”

## VH:STXS uncertainties sources

- ▶ **QCD:** nJet and  $p_T^V$  migration effects correlating WH and ZH prod. modes
- ▶ **EW:** Sudakow effects (corr. WH and ZH)  
non-Sudakow uncertainties (uncorr. WH and ZH)
- ▶ **note:** separate/uncorrelated uncertainties for Drell-Yan-like ZH, and loop-induced  $gg \rightarrow ZH$  correlation with other STXS bins to be studied



# Simplified Template Cross Sections - VH

## WG2:STXS

- ▶ main guidelines / recommendations on the STXS approach (e.g. baseline STXS bin split, 1st uncertainty scheme, ...)
- ▶ contact point among Higgs “production modes” for STXS matters (e.g. VH contributions to VBF and ggF bins, consistent unc. treatment)

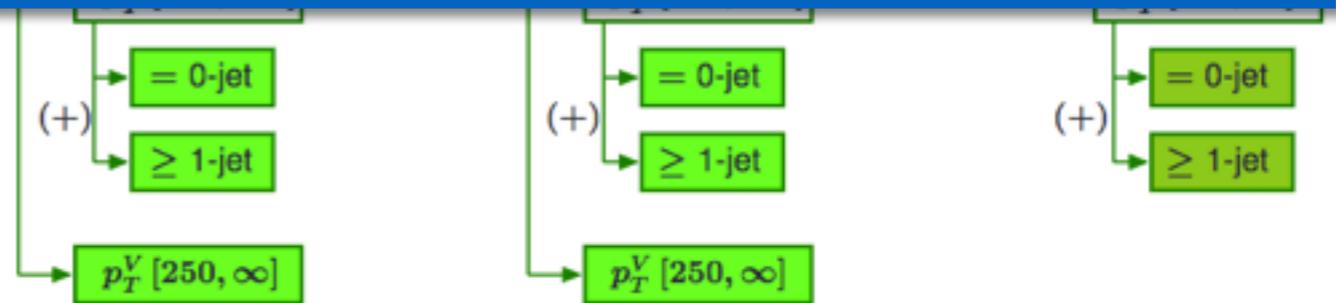
## WG1:VH

- ▶ discussion on finer details of STXS bin split / uncertainty scheme: gather feedback from VH theory experts
- ▶ **estimate** of theory uncertainty for the STXS framework
- ▶ focused experimental point of view on feasibility/practicality



## VH:STXS uncertainties sources

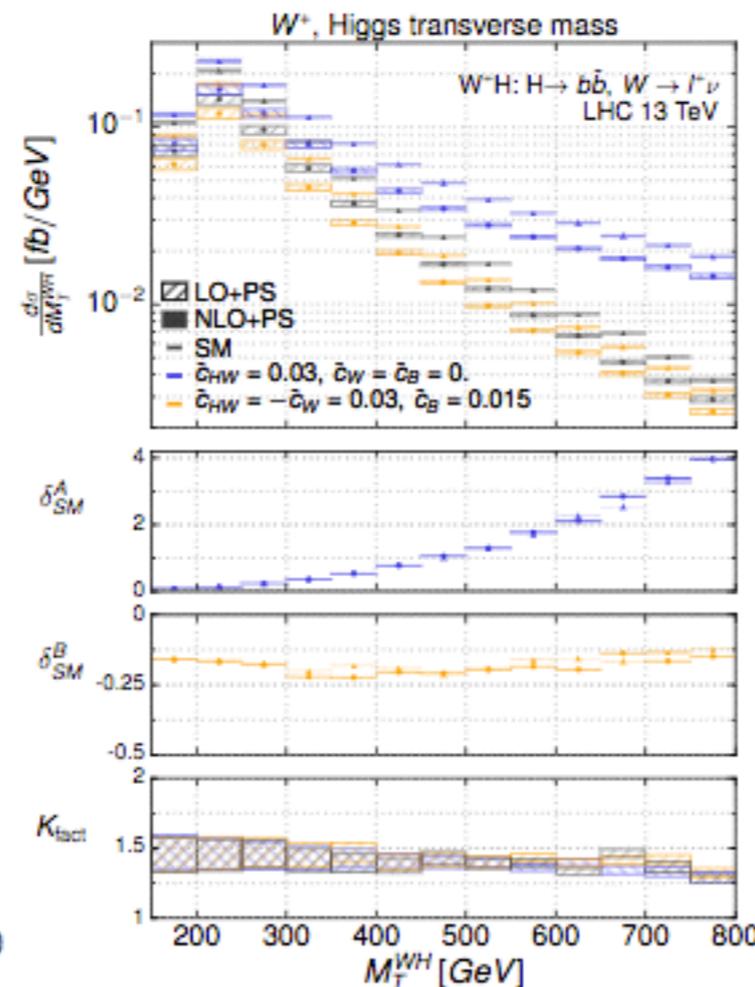
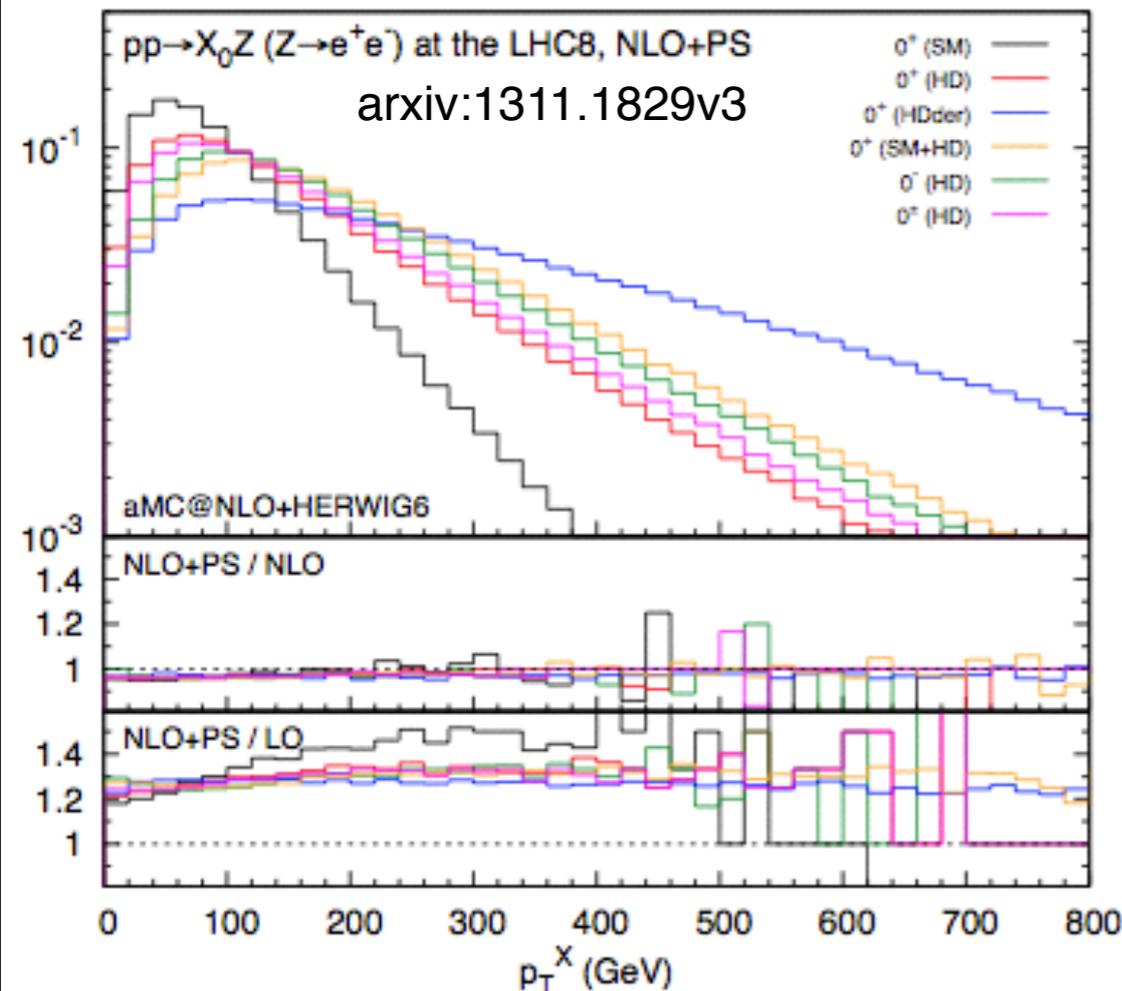
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# Interaction with other WGs: BSM VH

VH production can play a relevant role both in specific BSM models and in EFTs

- ▶ possible synergy with WG(2,3) on EFT anomalous coupling studies and model-specific BSM VH production ?
- ▶ SM VH predictions at high precision: can we benefit from same/similar level of precision for BSM VH production? could shared experience & tool help?



Some examples already from [Francesco Riva's talk](#) this morning

- ▶ EFT / PO tools at NLO QCD already available for VH
- ▶ common issues (e.g. higher-order QCD, EW corrections - treatment of ggZH, treatment of uncertainties) may benefit from WG2-WG1 interaction

# Conclusions and Outlook

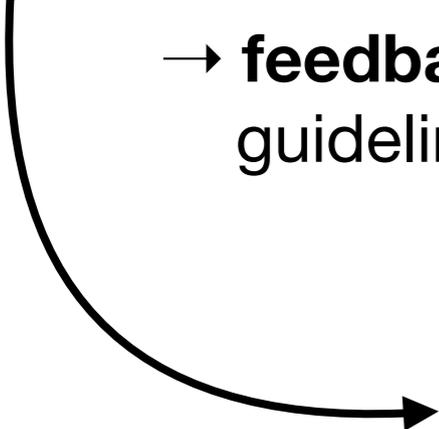
▶ **WG1:VH activities ongoing** - many different contributions already discussed in the first meeting → lots of material to define an interesting roadmap for the future

▶ **VH experimental results:**

main channel VH(bb) can reach level of evidence with current dataset

→ **very** important to have a clear plan of improvements already for the short/medium-term (full LHC Run-2 dataset)

→ **feedback** from experimental collaborations crucial to provide **practical** guidelines to address the most sensitive issues



**STXS** approach could become very interesting with full Run-2 dataset [potentially  $5\sigma$  -  $O(20\%)$  precision on  $\mu$ ]

▶ Potential opening to background studies related specifically to VH(bb) phase-space

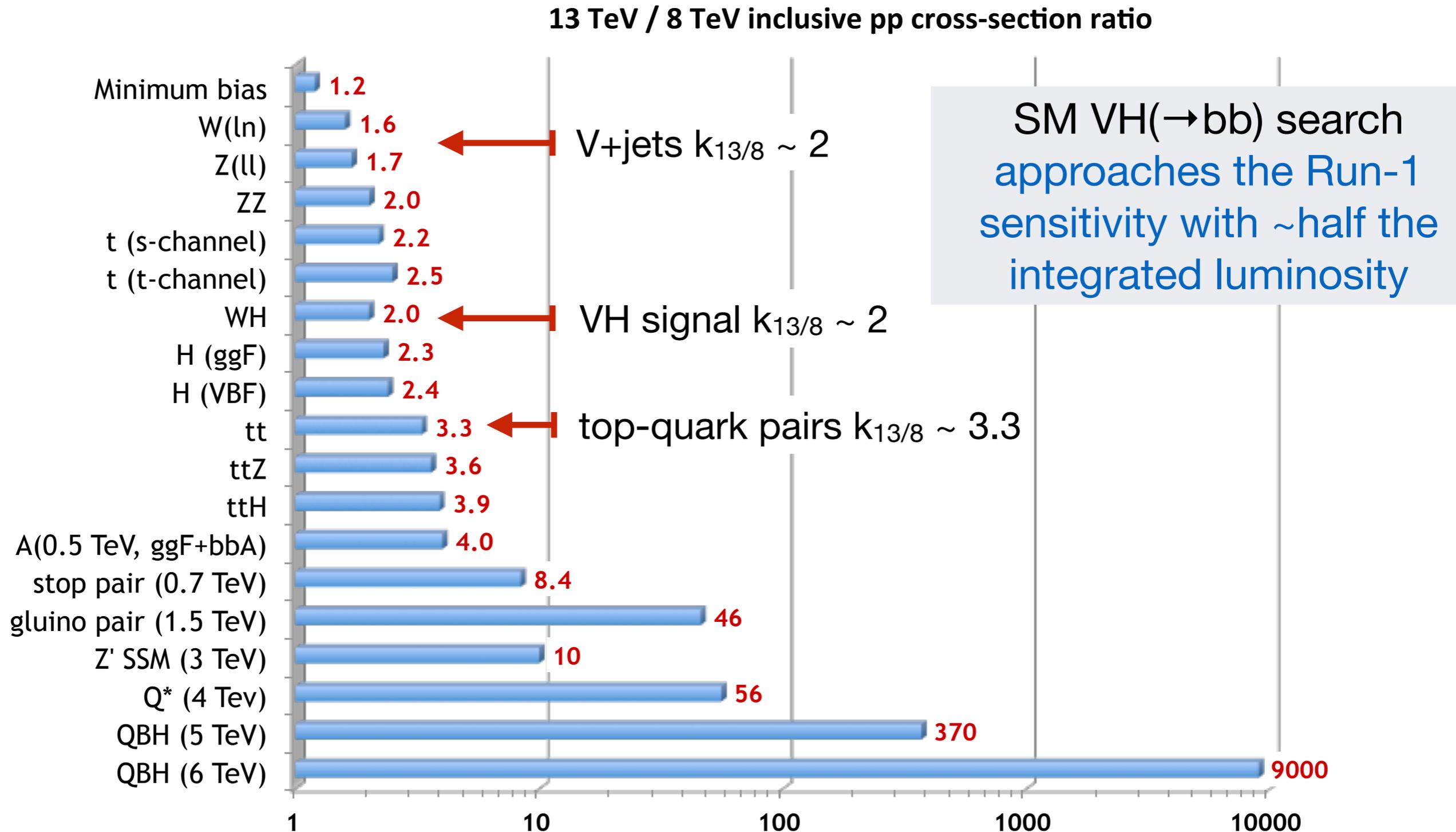
▶ Several theory developments on a medium-long timescale (see *Francesco's talk*): feedback on the “experimental sensitivity” to these improvements

▶ BSM VH - possible interaction with other WG\* ?

**Back-up**

# VH(bb) searches @ 13TeV

First search for the SM VH(bb) search with 13 TeV data from the LHC Run-2



# Simplified Template Cross Sections - VH

## STXS: separating measurements from interpretations

- ▶ maximize measurements sensitivity
- ▶ minimize theory dependence (models&systematics)
- ▶ combine **all decay channels**
- ▶ measure **XS instead of signal strengths**
- ▶ measure XS separately for **production modes**
- ▶ measure XS in **simplified fiducial volumes**
- ▶ allow for advanced analysis techniques (**MVAs**)

## Exclusive phase space regions (“bins”) defined to

- ▶ maximize experimental sensitivity
- ▶ minimize dependence on theory uncertainties directly folded into the measurements
- ▶ provide sensitivity to BSM scenarios

$$\sigma^{\text{meas}} = A^{\text{ggH}} \times \mu_{\text{ggH}} \times \sigma_{\text{ggH}}^{\text{SM}} + A^{\text{VBF}} \times \mu_{\text{VBF}} \times \sigma_{\text{VBF}}^{\text{SM}}$$

$$\downarrow$$

$$= A^{\text{ggH}} \times \sigma_{\text{ggH}} + A^{\text{VBF}} \times \sigma_{\text{VBF}}$$

$A^{\text{ggH}}$  Signal acceptance  
 $A^{\text{VBF}}$  theory dependent

$$\sigma^{\text{meas}} = A_a^{\text{ggH}} \times \sigma_a^{\text{ggH}}^{\text{SM}} + A_b^{\text{ggH}} \times \sigma_b^{\text{ggH}}^{\text{SM}} + A_c^{\text{VBF}} \times \sigma_c^{\text{VBF}}^{\text{SM}} \quad \text{a,b,c = “bins” of STXS}$$

$A_i^{\text{ggH}}$  Signal acceptance dependent on SM signal kinematic only within the given bin “i”

$A_i^{\text{VBF}}$  [reduce theory dependence]

# VH Signal Model @ 13TeV

$m_H = 125 \text{ GeV}$ at $\sqrt{s} = 13\text{TeV}$				
Process	Cross section $\times$ BR [fb]	Acceptance [%]		
		0-lepton	1-lepton	2-lepton
$q\bar{q} \rightarrow (Z \rightarrow \ell\ell)(H \rightarrow b\bar{b})$	29.9	< 0.1	< 0.1	7.0
$gg \rightarrow (Z \rightarrow \ell\ell)(H \rightarrow b\bar{b})$	4.8	< 0.1	< 0.1	15.7
$q\bar{q} \rightarrow (W \rightarrow \ell\nu)(H \rightarrow b\bar{b})$	269.0	0.2	1.0	–
$q\bar{q} \rightarrow (Z \rightarrow \nu\nu)(H \rightarrow b\bar{b})$	89.1	1.9	–	–
$gg \rightarrow (Z \rightarrow \nu\nu)(H \rightarrow b\bar{b})$	14.3	3.5	–	–

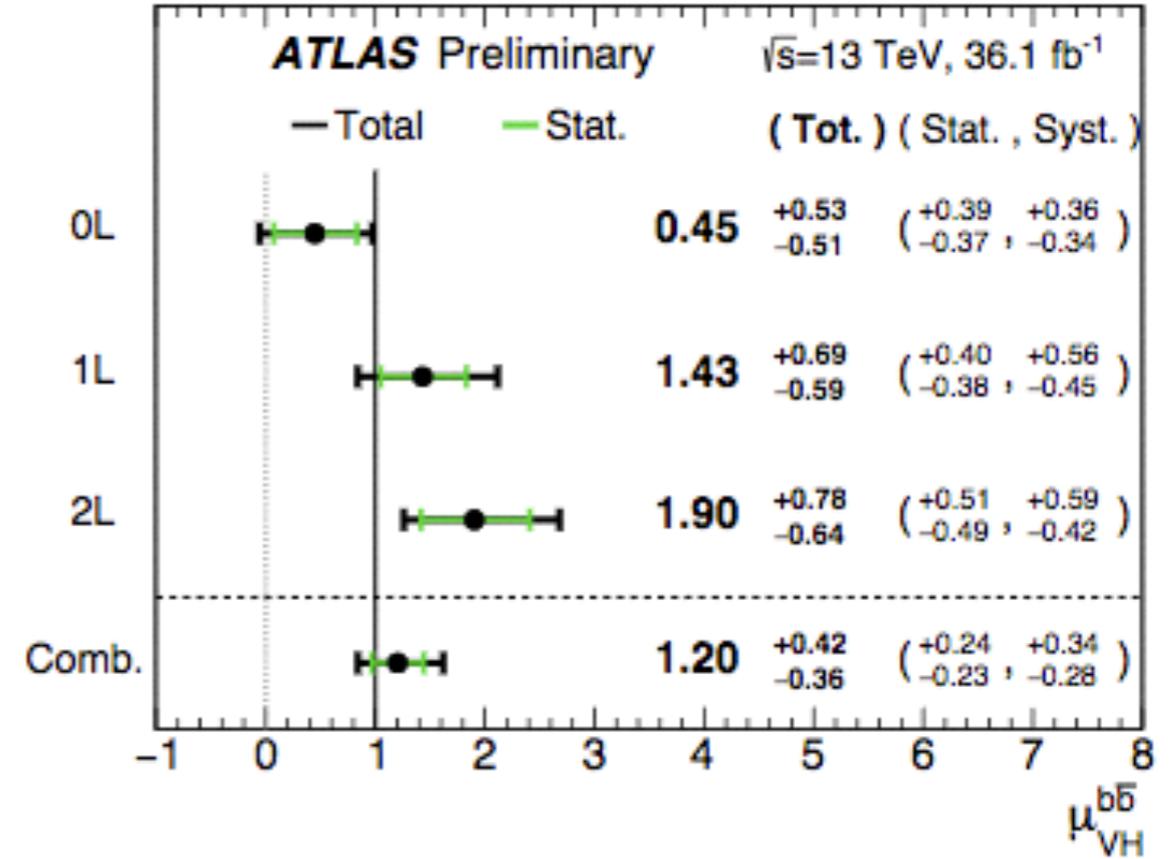
Table 8: Summary of the systematic uncertainties in the signal modelling. “PS/UE” indicates parton shower / underlying event. An “S” symbol is used when only a shape uncertainty is assessed.

Signal	
Cross-section (scale)	0.7% ( $qq$ ), 27% ( $gg$ )
Cross-section (PDF)	1.9% ( $qq \rightarrow WH$ ), 1.6% ( $qq \rightarrow ZH$ ), 5% ( $gg$ )
Branching ratio	1.7 %
Acceptance from scale variations (var.)	2.5% – 8.8% (Stewart-Tackmann jet binning method)
Acceptance from PS/UE var. for 2 or more jets	10.0% – 13.9% (depending on lepton channel)
Acceptance from PS/UE var. for 3 jets	12.9%–13.4% (depending on lepton channel)
Acceptance from PDF+ $\alpha_s$ var.	0.5%–1.3%
$m_{bb}, p_T^V$ , from scale var.	S
$m_{bb}, p_T^V$ , from PS/UE var.	S
$m_{bb}, p_T^V$ , from PDF+ $\alpha_s$ var.	S
$p_T^V$ from NLO EW correction	S

# VH(bb) @ 13TeV [ATLAS]

Evidence of VH(bb) at 3.5 obs. sigma (3.5 exp.)

Dataset	$p_0$		Significance	
	Exp.	Obs.	Exp.	Obs.
0-lepton	4.2%	30%	1.7	0.5
1-lepton	3.5%	1.1%	1.8	2.3
2-lepton	3.1%	0.019%	1.9	3.6
Combined	0.12%	0.019%	3.0	3.5



- ▶ 3-lepton channels compatible at  $\sim 10\%$
- ▶  $XS(VH \rightarrow Vbb) = 1.57^{(+0.54, -0.47)} \text{ pb}$   
[from a simple signal strength scaling]
- ▶  $XS_{\text{expSM}}(VH \rightarrow Vbb) = 1.31 \text{ pb}$
- ▶ provide sensitivity to BSM scenarios

Channel	SR/CR	Categories			
		2 $b$ -tagged jets			
		$75 < p_T^V < 150 \text{ GeV}$		$p_T^V > 150 \text{ GeV}$	
		2 jets	3 jets	2 jets	3 jets
0 lepton	SR	-	-	BDT	BDT
1 lepton	SR	-	-	BDT	BDT
2 lepton	SR	BDT	BDT	BDT	BDT
1 lepton	$W+HF$ CR	-	-	yield	yield
2 lepton	$e\mu$ CR	$m_{b\bar{b}}$	$m_{b\bar{b}}$	$m_{b\bar{b}}$	$m_{b\bar{b}}$

# ZH XS update

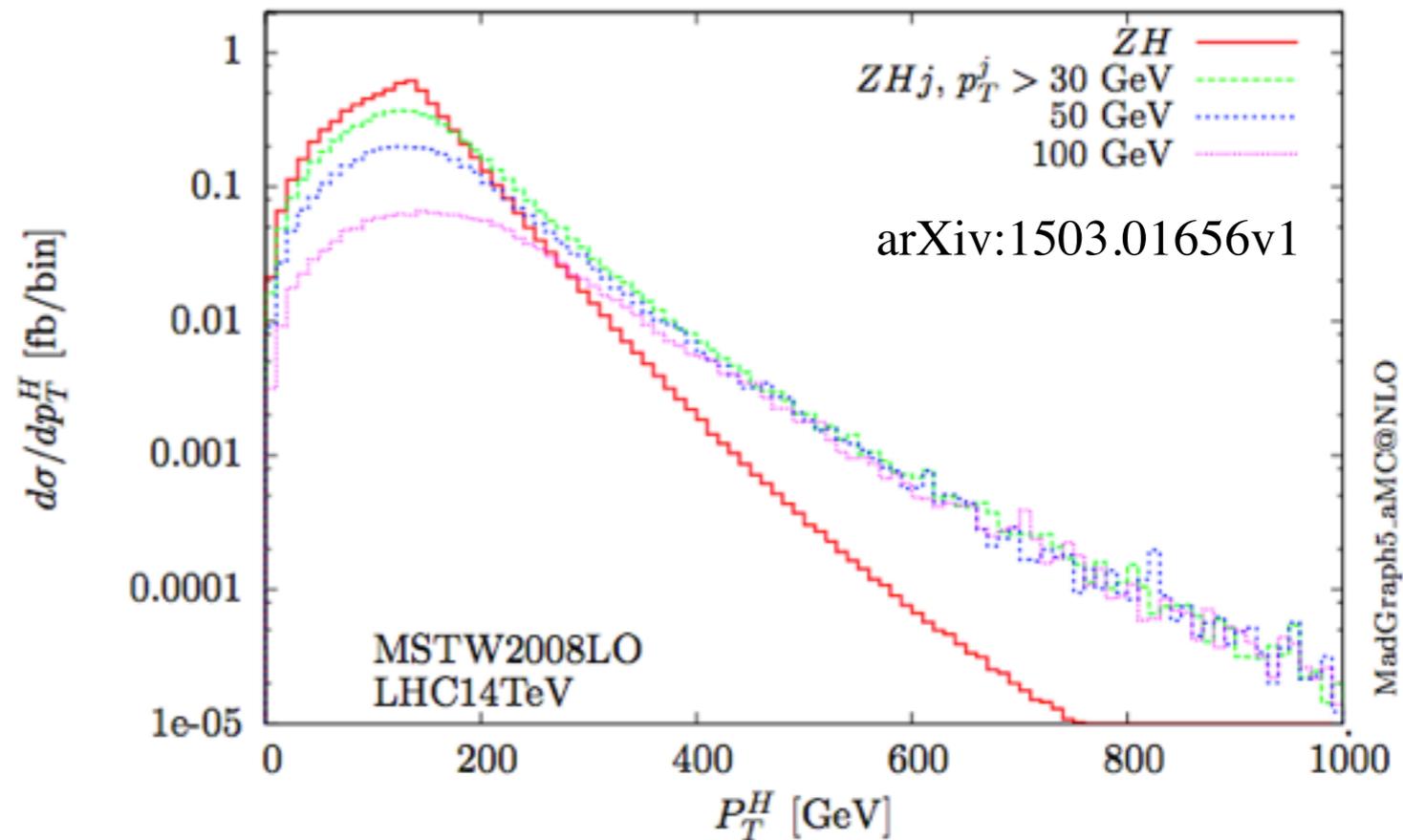
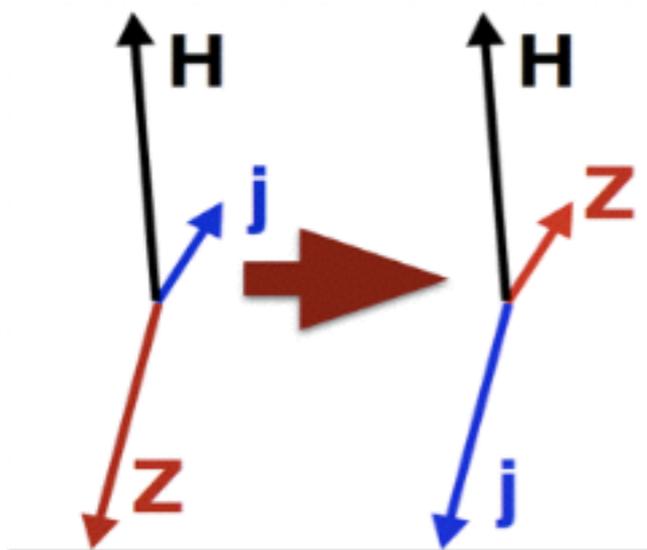
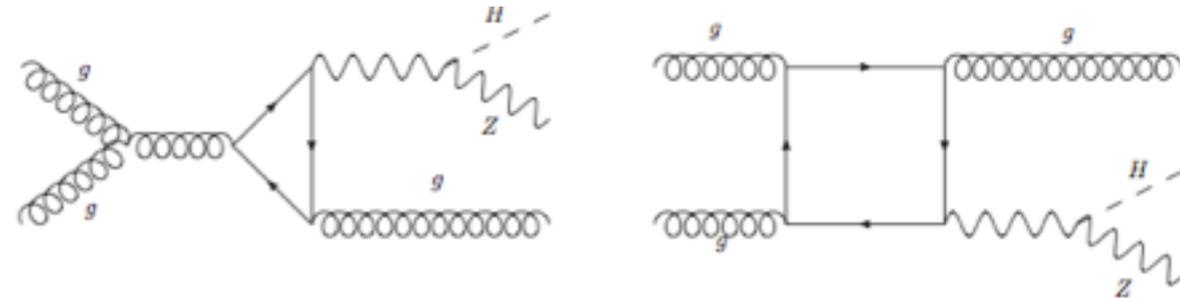
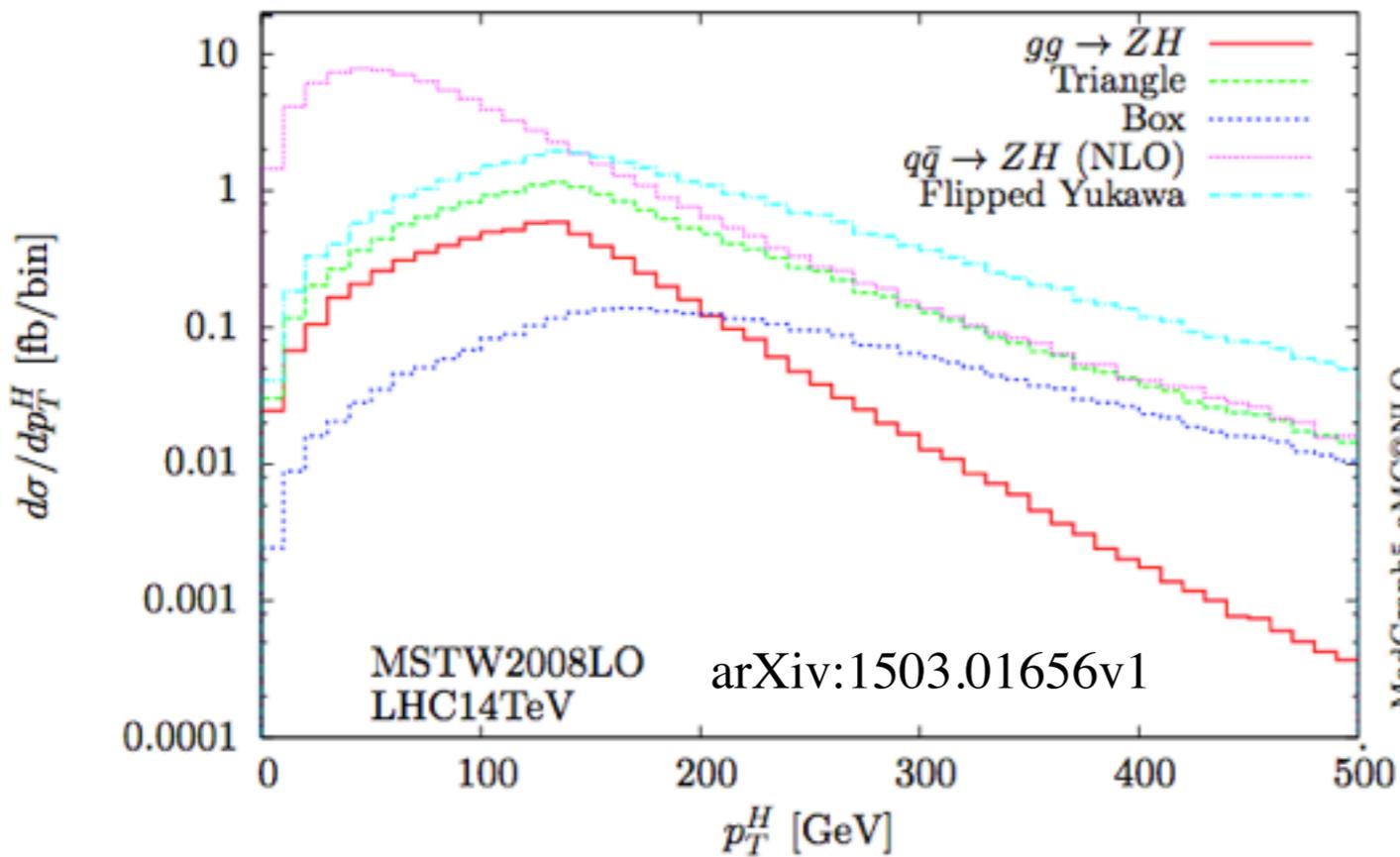
post-YR4 update: separate uncertainties for qqZH and ggZH (many thanks to Robert!)

## gg→ZH Cross Section

- ZH production has two distinct sources of gg→ZH:
  1. a genuine NNLO contribution to what called "Drell-Yan-like", where ZH is accompanied by two-parton radiation, gg→HZ+qqbar.
  2. top- and bottom-loop induced contribution without any additional partons in the final state.
- What is usually meant by gg→HZ below is 2) above.
- The statement that "all but gg→HZ" is the same as "qq- and qg-initiated" is correct only through NLO QCD.
- For separate cross sections and associated QCD scale uncertainties in qq/qg→ZH(+gg→HZ+qqbar) and gg→ZH for NLO/LO MC normalization, use
  - $\sigma(\text{all but gg}\rightarrow\text{ZH}) = \sigma(\text{pp}\rightarrow\text{ZH})@(\text{NNLO QCD} + \text{NLO EW, NLO+NLL QCD gg}\rightarrow\text{ZH}) - \sigma(\text{gg}\rightarrow\text{ZH})@(\text{NLO+NLL QCD})$ ,
  - Separate QCD scale uncertainties are  $\sigma(\text{all but gg}\rightarrow\text{ZH})$  or on  $\sigma(\text{gg}\rightarrow\text{ZH})$  are calculated with VH@NNLO program.
- For  $M_H=125.0$  GeV and at  $\sqrt{s}=13$  TeV,

Process	Cross Section (pb)	+QCD Scale %	-QCD Scale %	±(PDF+ $\alpha_s$ ) %	±PDF %	± $\alpha_s$ %
pp→ZH	0.8839	+3.8%	-3.1%	±1.6%	±1.3%	±0.9%
qq/qg→ZH, gg→HZ+qqbar (all but gg→ZH)	0.7612	+0.5%	-0.6%	±1.9%	±1.7%	±0.9%
gg→ZH	0.1227	+25.1%	-18.9%	±2.4%	±1.8%	±1.6%

# ggZH (loop-induced) diagrams



# Latest PO development: NLO in QCD

**HiggsPO** code can be found at <http://www.physik.uzh.ch/data/HiggsPO/>

Allows to simulate EW Higgs production (VH, VBF) in the PO formalism @ NLO in QCD.



## Higgs PO

DESCRIPTION

DOWNLOAD

CONTACTS

Download the model

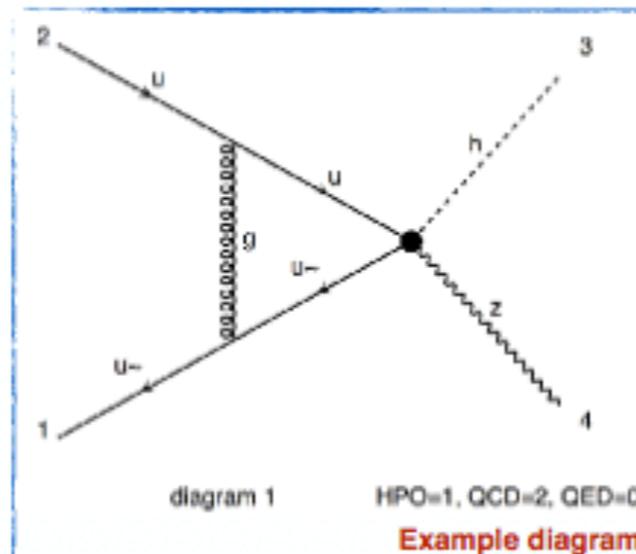
Version 1.2 - NEW

HiggsPO UFO model for electroweak Higgs production at NLO in QCD.

- UFO model: [HPO\\_ewk\\_prod\\_NLO](#) (zip)
- Presentation and examples (slides from the LHC HXSWG meeting of May 8th, 2017): [link](#)

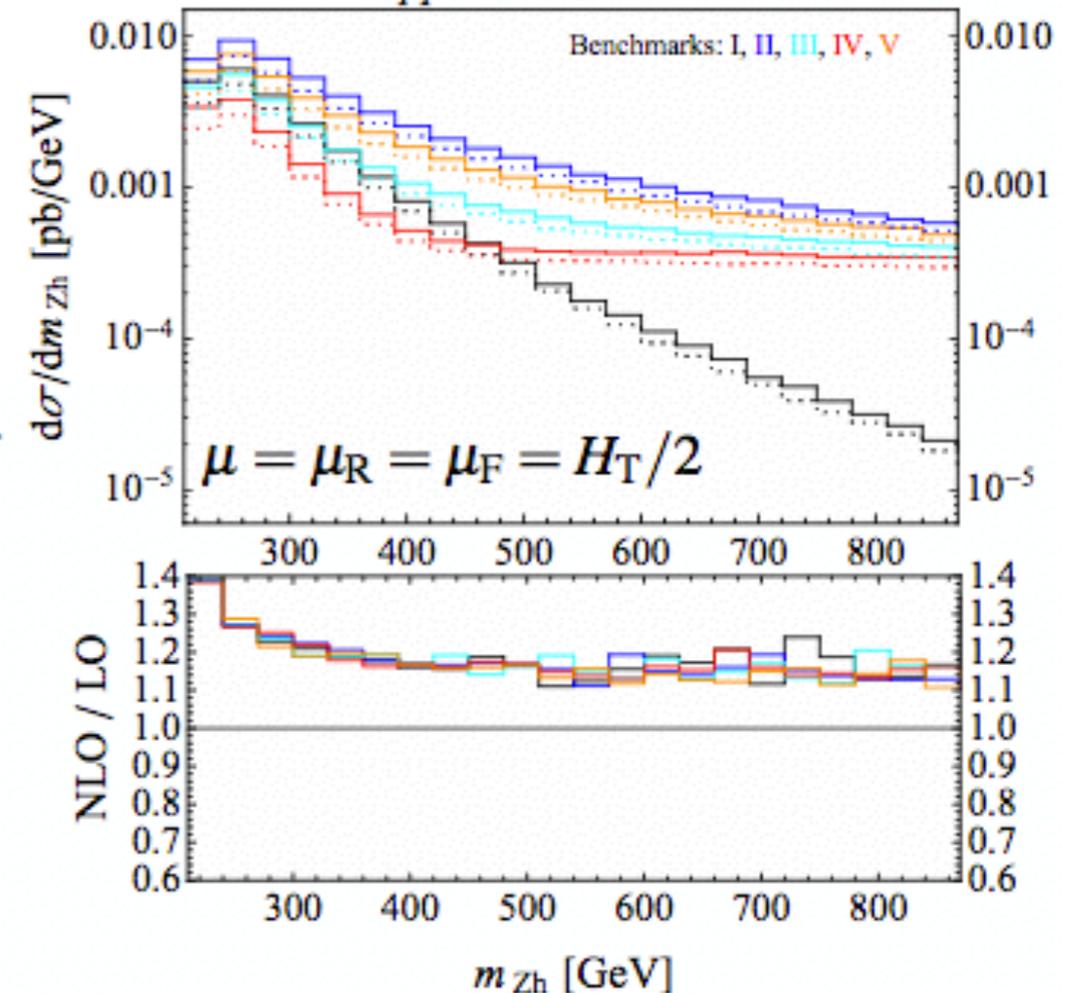
For info and examples see A. Greljo [talk](#) at WG2 kickoff meeting.

Paper coming soon.



Example: Zh at NLO

$pp \rightarrow Zh$  @ 13 TeV



# Parton Shower Models and Variations

**AZNLO**: designed for the Powheg+Pythia8 NLO+PS generator, and provide a very good description of ISR in the low and medium  $p_T$  region

Measurements of the  $Z/\gamma^*$  boson transverse momentum distribution  
(and  $\phi_\eta^*$  angular correlation) in  $pp$  collisions at  $\sqrt{s} = 7$  TeV

JHEP, 09:145, 2014

1211.6899

Strategy for the Powheg+Pythia8 tune → tunes performed for  $p_T(Z) < 26$  GeV and  $\phi_\eta^* < 0.29$   
(best description of the tuning parameters)

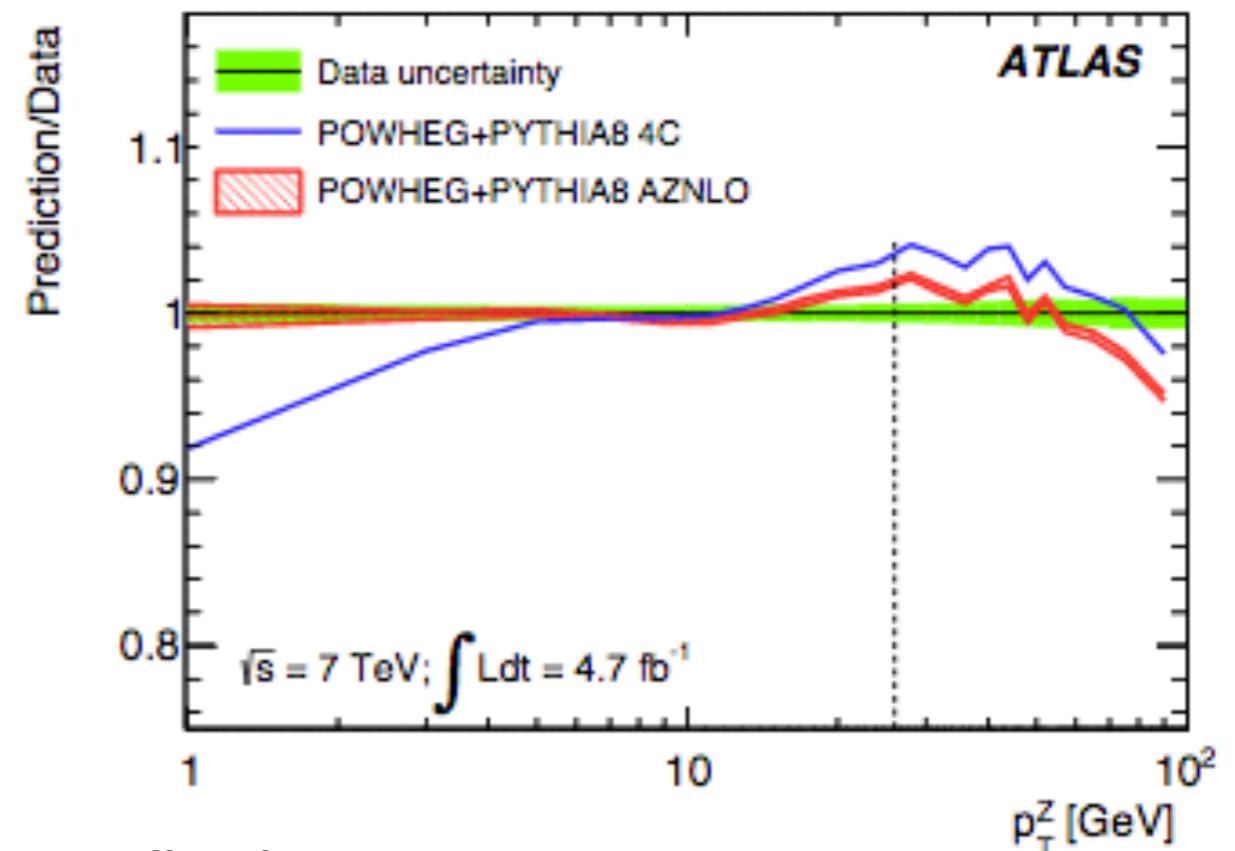
The tuning only varies the ISR shower cut-off  
and the primordial  $k_T$  in Pythia8:  
essentially constrained by data  $p_T(Z) < 12$  GeV  
- not affected by tuning upper bound  
(plus MPI parameters)

Tuned predictions agree with the measured  
XS within 2% for  $p_T(Z) < 50$  GeV

## “Eigentune variations”:

only covering ISR/primordial- $k_T$  variations scale  
variations for FSR and MPI cut-off parameters are  
recommended to cover the full range of UE/PS/MPI  
uncertainties

- ▶ VAR1, VAR2: eigentune diagonalization
- ▶ MPIUp, MPIDown
- ▶ FSRUp, FSRDown



(CTEQ6L1) LO PDF  
Main31 ME+PS matching