Future Neutrino Experiments

Peter Dornan

Imperial College London

Why? Non-Oscillation Expts Oscillation Expts Active Near Future More distant future

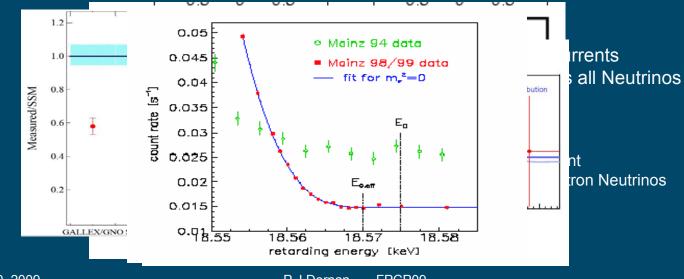
Not included Neutrino astroparticle Expts

Why do Neutrino Physics?

- Least understood particle
 - Yet after photon most common
- Beyond the Standard Model
 - A trivial Addition?
 - The Window on the Fundamental Theory?
- Understanding the Flavour Problem
 - Understanding lepton flavour as fundamental as quark flavour (and no QCD complications)
- New Source(s) of CP Violation
 - Contributor (solution) to the anti-matter asymmetry
- Ultimate theory must relate quarks & leptons
 - Cannot do this without a full understanding of the neutrino sector
 - Cannot do this with LHC or ILC

What do we know?

- 1. There are three light active neutrino species (LEP/SLC)
- 2. Mass of electron neutrino $<2.2 \text{ eV/c}^2$ from H³ decay spectrum (Mainz)
- 3. Muon Neutrinos produced in the upper atmosphere disappear (SuperK)
- 4. Muon Neutrinos produced in π decay disappear (K2K, MINOS
- 5. Electron Neutrinos produced in the sun disappear (Many experiments)
- 6. The total flux of neutrinos from the sun stays constant (SNO)
- 7. Electron antineutrinos from reactors disappear and then reappear (Kamland)



Other Results - Require Confirmation

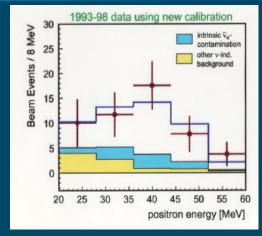
- 1. The appearance of electron antineutrinos in a muon antineutrino beam (LSND)
 - If true a 3-Neutrino Scenario is excluded. Either
 - > 3 Flavours (with heavy lh neutrinos)
 - > 1 or more sterile neutrinos
- 2. The observation that the decay
- $\square \qquad ^{76} Ge \rightarrow ^{76} Se + e^- + e^-$

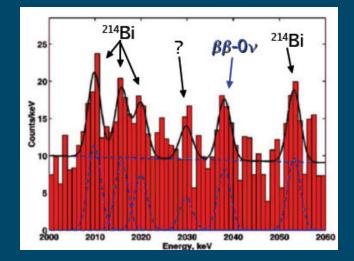
has a half life of

(H.V.Klapdor-

- $T_{1/2}^{0\nu} = 2.3_{-0.31}^{+0.44}.10^{25}$ years Kleingrothaus *et al.*,)
 - o If True
 - Neutrinos are Majorana
 - $M(v_e) = ~0.44 \text{ eV}$

The accepted results are well explained assuming a three family neutrino oscillation





Oscillation Probability

Depends upon Mixing Angle, Mass Squared Difference, L/E

$$\begin{array}{c} \mathbf{2} \text{ state } (\mathbf{e}, \mu) \\ \begin{pmatrix} v_e \\ v_\mu \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} v_1 \\ v_2 \end{pmatrix} \end{array} \quad \text{Prob}_{v_e \Rightarrow v_\mu} = \sin^2 2\theta \sin^2 \frac{\Delta m^2 L}{E} \\ \Delta m^2 = m_2^2 - m_1^2 \end{aligned}$$

- but ambiguities

$$\theta \rightarrow \frac{\pi}{4} - \theta \qquad \Delta m^2 \rightarrow -\Delta m^2$$

$$\begin{pmatrix} v_e \\ v_\mu \\ v_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-id} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{-id} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} e^{i\alpha/2} & 0 & 0 \\ 0 & e^{i\beta/2} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} v_1 \\ v_2 \\ v_3 \end{pmatrix}$$

3 mixing angles, θ_{12} , θ_{13} , θ_{23} , 1 Dirac phase, δ 2 Majorana phases, α , β (=0 if Neutrinos are Dirac)

Vacuum Oscillations $\theta_{12}, \theta_{13}, \theta_{23}$ δ $\left|m_{2}^{2} - m_{1}^{2}\right|, \left|m_{3}^{2} - m_{1}^{2}\right|$ $\sin \Delta_{ij}, \Delta_{ij} = \left(\frac{1.27 \Delta m_{ij}^{2}L}{E}\right)$ May 30, 2009

In matter

$$\begin{aligned} \sin\Delta_{13} \Rightarrow \frac{\sin(\Delta_{13} - AI)}{(\Delta_{13} - AI)} \Delta_{13} & \sin\Delta_{12} \Rightarrow \frac{\sin(AI)}{(AI)} \Delta_{12} \\ with \ A = \pm \frac{G_F N_e}{\sqrt{2}} \end{aligned}$$

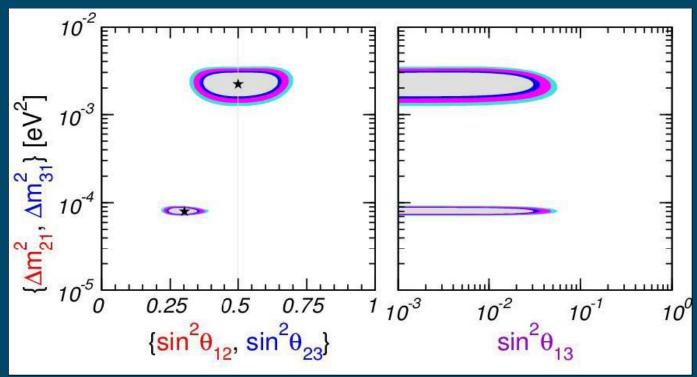
Sensitive to Sign (Δm^2)

- depends upon *AL* i.e. density and baseline

P J Dornan - FPCP09

Where are we?

from: Maltoni, Schwetz, Tortola, Valle ('04)



Also know $m_2 > m_1$ from matter effects in the sun

Future - Non-Oscillation Experiments

• Majorana/Dirac

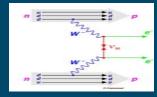
Absolute Mass

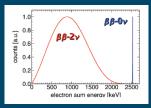
Dirac or Majorana? - $0\nu\beta\beta$

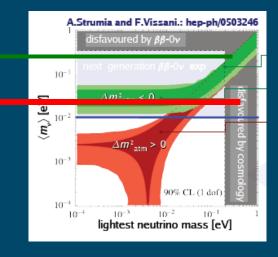
- If Dirac lepton no. is conserved. If so what is it?
- o If Majorana
 - $\Delta L = 2$ BSM processes, More CP violation, Seesaws possible, leptogenesis, solution to antimatter asymmetry?
- Only approach $0v\beta\beta$ Rate depends on mass

$$m_{ee} = \left|\sum U_{ei}m_i\right|^2$$

- Positive Result
 - o v's are Majorana
 - A measurement of the absolute Mass (subject to NME)
- Negative result
 - Nothing cannot even confirm Dirac nature
- Many Experiments planned
 - CUORE, GERDA, MAJORANA, SuperNEMO, EXA, XMASS, SNO+....
 - Range of techniques:Bolometers, Scintillators, Tracking, Combinations
 - Many Isotopes: Te, Ge, Se, Xe, Mo, Cd, Ca, Gd...
 - Aim `for mass values down to ~ 0.02 eV







HM

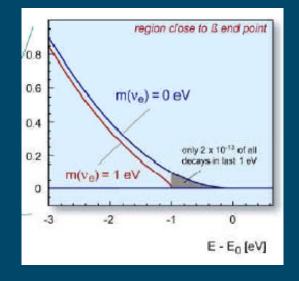
Future

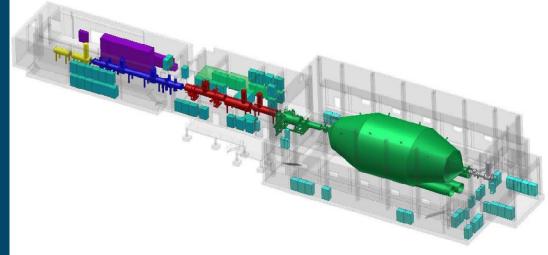
The Absolute Mass?

- Cosmological data
- Neutrinoless double-beta decay
- Tritium Decay Spectrum

Model Dependent Currently $\Sigma m < 0.7 \text{ eV}$ Potential $\Sigma m < 0.07 \text{ eV}$?

Katrin aims for $M(v_e) < 0.2 eV$





'Standard' Oscillation Goals

- Observe Flavour of an Oscillated v
- How small is θ_{13} ?
- Mass hierarchy?
- Leptonic CP Violation?
- Is θ_{23} maximal? is it > or < $\pi/4$?
- Is the MNS approach correct?

Are there sterile neutrinos?

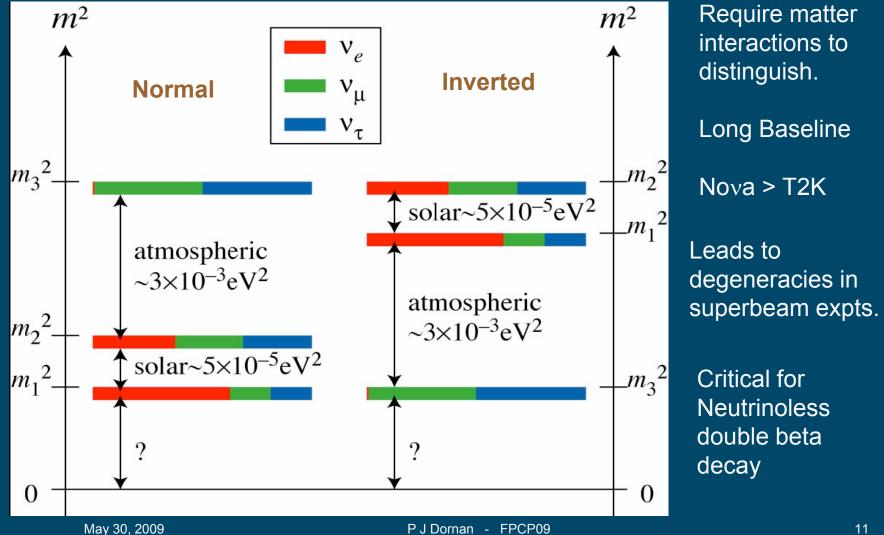
CPT violation?

Do we have Tribimaximal mixing?

Identify a v_{τ} or v_{e} in a v_{u} Beam Critical Parameter to define future programme $m_3 < or > m_2$, requires $\theta_{13} \neq 0$ Requires $\theta_{13} \neq 0$, $\delta \neq 0$ Improve accuracy of atmospheric parameters Measure parameters in as many ways – and as precisely - as possible Precision measurement of parameters. Repeat LSND? (MiniBooNE does not disprove LSND in exotic scenarios – different E) Measure parameters in as many ways – and as precisely - as possible

If yes presumably some underlying symmetry. Needs precision meaurements

Mass Hierarchy?



CP Violation

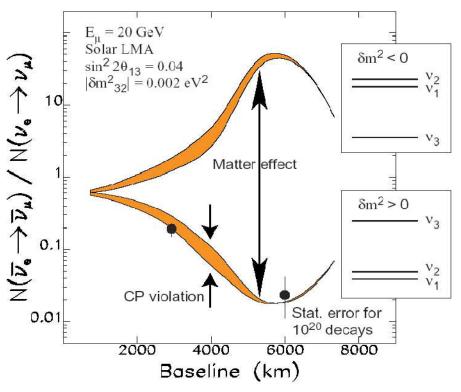
 Vitally important for our understanding of nature and the universe

0

Arises from phases which flip sign on the change particle $\leftarrow \rightarrow$ antiparticle

Standard MNS matrix has one phase, δ , (like CKM)

 BUT also matter effects flip sign on the change particle ←→ antiparticle Wrong-Sign Muon Measurements



AND, if neutrinos are Majorana additional phases and potential for substantial CP violation - but Majorana phases appear virtually unmeasurable

May 30, 2009

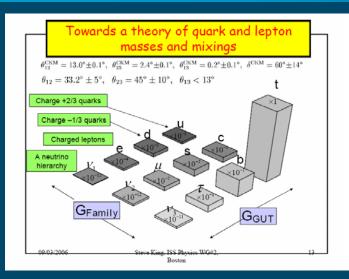
The Need for Precision Measurements

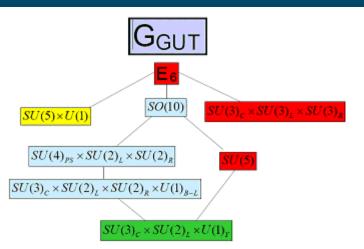
- No understanding of the underlying physics without precision measurements - as in the quark sector
 - CP violation in the neutrino sector?
 - Interesting to demonstrate this at 3σ level
 - Much more useful to know δ precisely
- Quark Lepton Complementarity
 - Motivated by GUTs
 - Also by intriguing relations such as

$$\theta_{12}^{MNS} + \theta_{12}^{CKM} \cong \frac{\pi}{4}$$
$$\theta_{23}^{MNS} + \theta_{23}^{CKM} \cong \frac{\pi}{4}$$

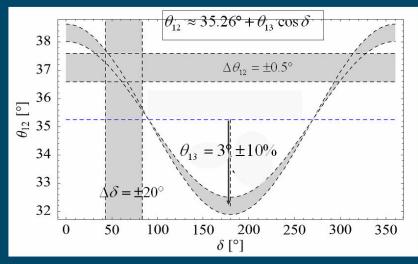
Improved precision of solar and atmospheric angles needed

Ideas/Speculation





A Neutrino Sum Rule? (S King)



Tribimaximal Mixing? (Harrison, Perkins, Scott)

$$\begin{pmatrix} \sqrt{\frac{2}{3}} & \sqrt{\frac{1}{3}} & 0 \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & -\sqrt{\frac{1}{2}} \\ -\sqrt{\frac{1}{6}} & \sqrt{\frac{1}{3}} & \sqrt{\frac{1}{2}} \end{pmatrix}$$

The Imminent Future

Two long base line experiments using a v_{μ} beam now taking data

MINOS

- Numi Beam from FNAL to Soudan 735 km
 - Two Detectors near and far magnetized Fe-scintillator
 - Look for v_{μ} disappearance $\rightarrow \theta_{23}$, Δm_{23}^2
 - v_e Appearance $\rightarrow \theta_{13}$ (~factor 2 better than Chooz)

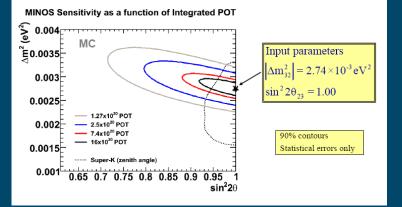
OPERA

- CNGS Beam from CERN to Gran Sasso 732 km
 - One Far Detector Emulsion
 - Look for v_{τ} Appearance, now expect ~10 observed v_{τ} events in 5 years

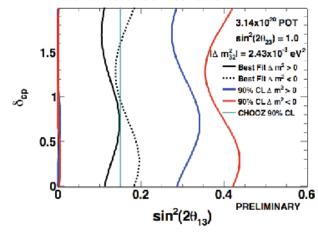
MINOS

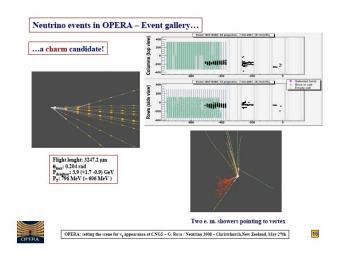
OPERA

Projected sensitivity of MINOS



Feldman-Cousins C.L. contours for ANN

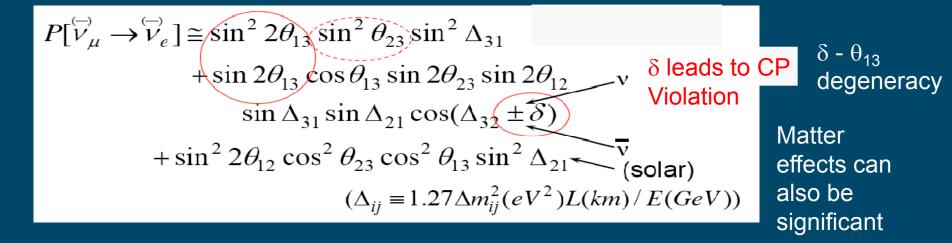




τ- decaγ	Signal ÷ ∆m² (Full mixing)			
channels	2.5 x 10-3 (eV ²)	3.0 x 10-3 (eV ²)	Background	
τ-→µ-	2.9	4.2	0.17	
τ· → e·	3.5	5.0	0.17	
τ· → h·	3.1	4.4	0.24	
$\tau \rightarrow 3h$	0.9	1.3	0.17	
ALL	10.4	15.0	0.76	

In the next ~4 Years - θ_{13}

ν_e Appearance in a ν_µ Beam - SuperBeam



v_e Disappearance in a v_e Beam - Reactor

 $+\cos^4\theta_{13}\sin^22\theta_{12}\sin^2\Delta_{21}$

 $P[v_e \rightarrow \operatorname{Not} v_e] \cong \sin^2 2\theta_{13} \sin^2 \Delta_{31}$

No δ term, no matterr effects

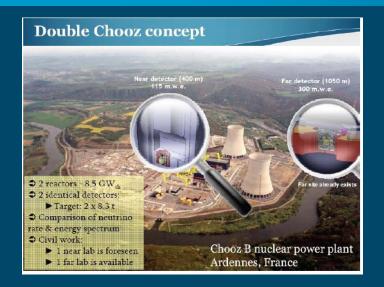
For θ_{13}

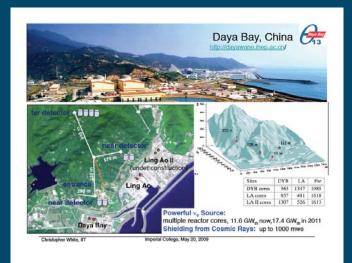
- Three Approved Reactor Experiments
 - o Double Chooz (France)
 - o Daya Bay (China)
 - o RENO (Korea)

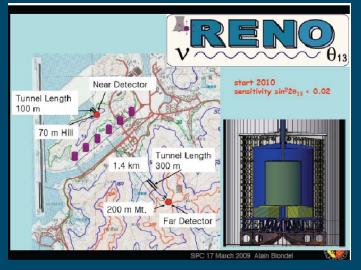
Two off-axis Superbeam Experiments

- Main aim is θ_{13} but will also improve other parameters
- o T2K
- o Nova maybe also the mass hierarchy

Reactor Expts for θ_{13}

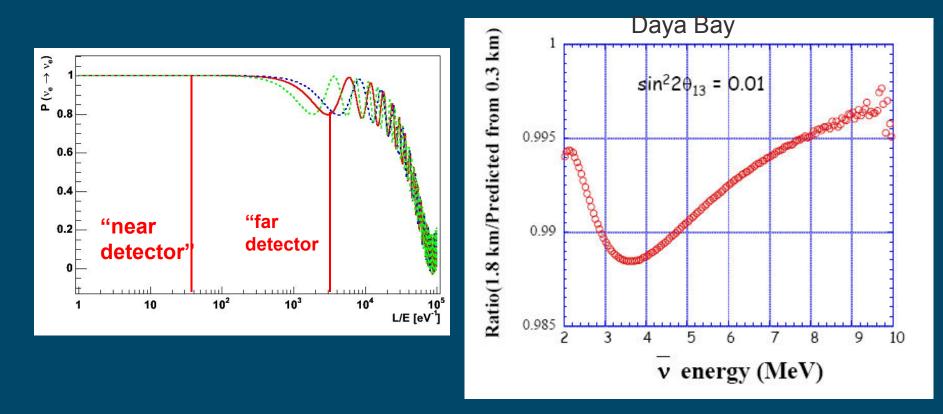






	Start	$sin^2 2\theta_{13}$
Double Chooz	2009- 10	>~0.02
Daya Bay	2011	>~0.01
RENO	2011	>~0.02

Reactor Measurement of θ_{13}



No Dependence on δ , No matter effects But small effect and control of systematics vital.

Superbeams – near future – T2K, Nova

• Both use a v_{μ} off-axis beam and have both e and μ detection

Aim

- Measurement or much better upper limit on θ_{13} using electron appearance
- Improve accuracy of $\sin^2 2\theta_{23}$, Δm^2_{23} using muon disappearance
- Determine mass hierarchy (Nova) depends on θ_{13}

T2K

- Super-K, 50kt Water Cerenkov Detector, 295 km from J-PARC
- A near detector but not same technique

• J-PARC Beam

- Just started , starts in earnest early 2010 at about 30 50 kW and will then steadily improve to ~1.5 MW, 30 – 50 GeV protons, 3 horns
- Off axis beam $\langle E_v \rangle \sim 0.65$ GeV, better enrgy spread and less contamination

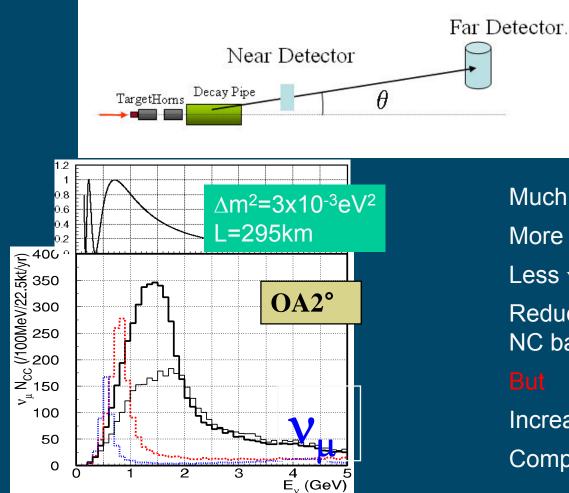
Nova

- Huge totally active scintillating detector (TASD)
- Work on the site has recently started , first events ~2012

Numi Beam

- Already available
- Upgrade to 700 kW

Off axis Superbeams

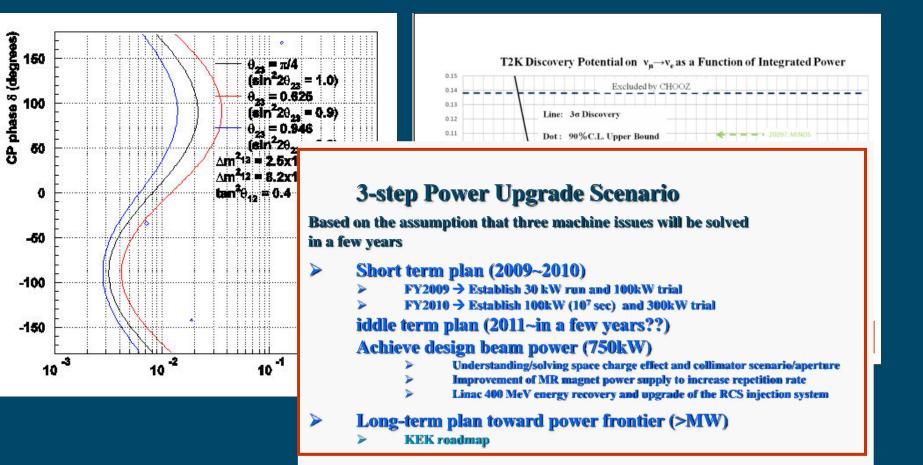


Much better energy resolution More flux at oscillation max Less v_e Conamination Reduces high energy tail and hence NC background

Increases near/far differences Complicates disappearance



T2K Prospects



Search for $v_{\mu} \rightarrow v_{e}$ appearance T2K

Look for excess events in 1-ring e-like sample at SK

sin²2θ ₁₃	Ba	Cianal		
	ν_{μ} induced	Beam v _e	Total	Signal
0.1	10	13	23	103
0.01				10

Expected number of events at SK (0.75kW beam x 5yr)

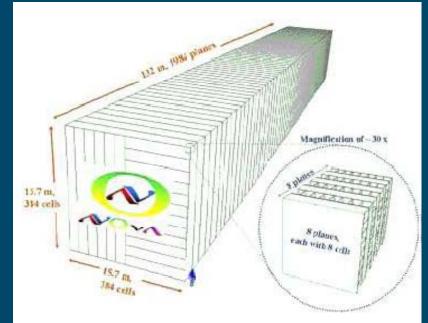
- Dominant background sources:
 - Beam v_e contamination
 - Irreducible, but different energy spectrum from oscillated ve
 - ν-induced NC1π⁰ events
 - one of 2γ from π⁰ is missed
 - Reducible, needs knowledge of NC1π⁰ interaction

→ To be studies/estimated at near detector

Nova

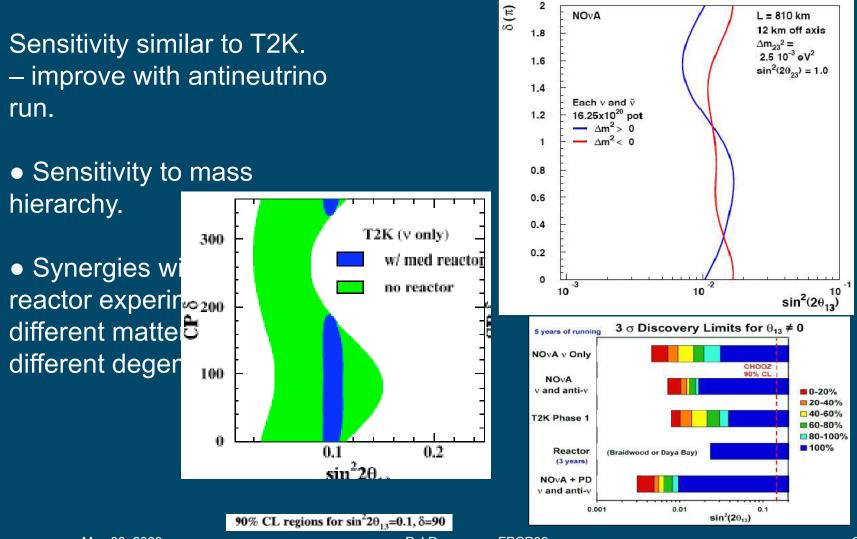
Upgrade of FNAL NuMi program.

- Off-axis configuration, and larger proton intensity (6.5x1020 proton/year).
- Very Long Baseline (810km) and sizeable matter effects
- Complementary to T2K program (mass hierarchy).
- <E_V> ~ 2.22GeV
- 30Kton "fully" active detector.
- Liquid scintillator.
- Data taking Expected to start 2012



May 30, 2009

Nova



May 30, 2009

P J Dornan - FPCP09

The Precision Era - after T2K and Nova

- Around 2012 2015
- We shall have good measurements of
 - ο $θ_{12}, θ_{23}, \Delta m^2_{12}, \Delta m^2_{23}$
- Probably have a measurement of θ_{13}
- Possibly know the mass hierarchy

So can now plan for the ultimate neutrino measurements

- o Refine all parameters
- Check consistency
- o Measure CP Violation

Long Baseline Experiments

What can be measured?

o v_e Beam

 $v_e \rightarrow v_e$ $\overline{v}_e \rightarrow \overline{v}_e$ 'disappearance' $v_e \rightarrow v_{\mu}$ $\overline{v}_e \rightarrow \overline{v}_{\mu}$ 'golden channel' $v_e \rightarrow v_{\tau}$ $\overline{v}_e \rightarrow \overline{v}_{\tau}$ 'silver channel'

o v_{μ} Beam

 $v_{\mu} \rightarrow v_{e}$ $\overline{v}_{\mu} \rightarrow \overline{v}_{e}$ 'platinumchannel $v_{\mu} \rightarrow v_{\mu}$ $\overline{v}_{\mu} \rightarrow \overline{v}_{\mu}$ 'disappearance' $v_{\mu} \rightarrow v_{\tau}$ $\overline{v}_{\mu} \rightarrow \overline{v}_{\tau}$ 'silver channel

To test the Unitarity of the PMNS matrix and CPT invariance ideally measure all of these

Determining Factors

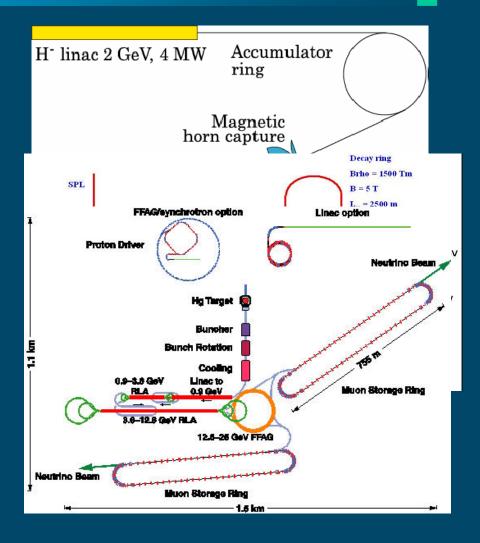
Beam

- Intensity, Purity, Divergence, Energy, Energy Spread?
- Detector
 - Size, energy resolution, for μ 's, for e's, **Detection threshold**, Pictorial (τ 's and e's), Magnetisable
- Near Detector
 - Crucial to understand systematics, same technology as far?
- Experiment
 - **Baseline**, L/E (optimum determined by Δm_{ij}^2 , now reasonably known, Removal of degeneracies- **More than 1 L/E, Backgrounds**
- Systematics
 - Cross sections, neutrino cross sections are still not well known
 - Particle/antiparticle
 - Near/Far
- Cost

Neutrino source – options:

- Second generation super-beam
 - CERN, FNAL, BNL, J-PARC II
- Beta-beam

Neutrino Factory



v_e, \overline{v}_e Beams - Golden Channel

Reactors

- Can only measure v_e disappearance
- It is hard to see reactors playing a major role after Double Chooz, Daya Bay and RENO

Beta beams

- High purity, Low Divergence beam possible
- Intensity questionable, depends primarily on ion source. particularly for v
- Wide range of possible detectors

Neutrino Factory

- High Purity, Low Divergence beam possible, can have high energy & very long baselines
- Intensity dependent on proton driver
- Can have alternate nu, nubar pulses
- Requires Magnetisable Detector

$v_{\mu}, \overline{v}_{\mu}$ Beams - Platinum Channel

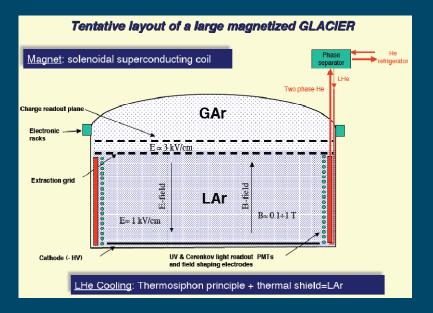
- Superbeams
 - v, anti-v in separate experiments. Significant v_e contamination
 - On axis wide band beam, Off axis narrow band. Wide band beam can give 1st and 2nd max at same L
 - Intensity dependent on proton driver, target and horn system
- Neutrino Factory
 - High Purity, Low Divergence beam possible, can have very long baselines
 - Intensity dependent on proton driver, target
 - Can have alternate nu, nubar pulses
 - Requires Magnetisable Detector not ideal for electron detection

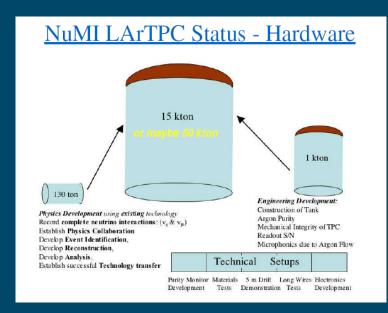
Detectors

- o Water Cerenkov
 - Good for low energy SB, BB, well established, also p-decay
- Liquid Argon TPC
 - SB, BB, (NF if magnetisable), prob also p-decay
- o MIND Iron / Scintillator sandwich (MINOS like)
 - NF baseline
- o TASD Totally Active Scintillating Detector (Nova, Minerva like)
 - SB, BB, (NF if magnetisable), poss also p-decay
- Hybrid Emulsion Detectors
 - NF for silver channel
- o Beam Diagnostic Devices
- o Near Detector
 - Vital for precision

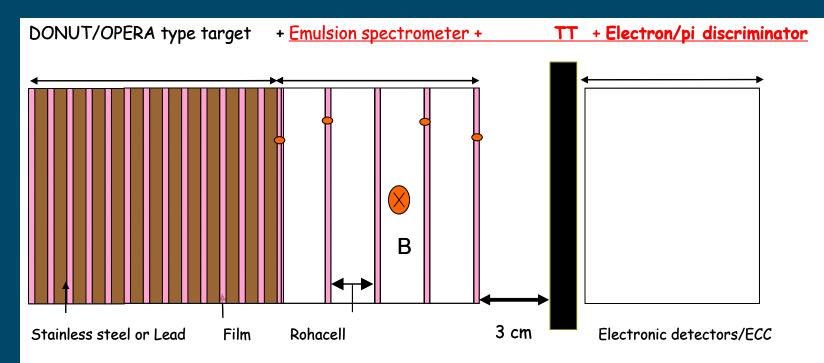
Liquid argon

Detector concepts





Emulsion detector – MECC



Assumption: accuracy of film by film alignment = 10 micron (conservative)

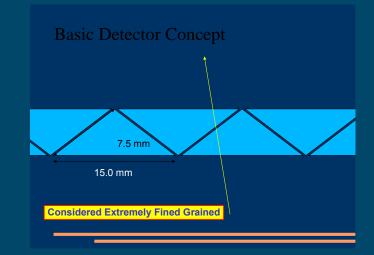
13 lead plates (~2.5 X_0) + 4 spacers (2 cm gap) (NB in the future we plan to study stainless steel as well. May be it will be the baseline solution: lighter target)

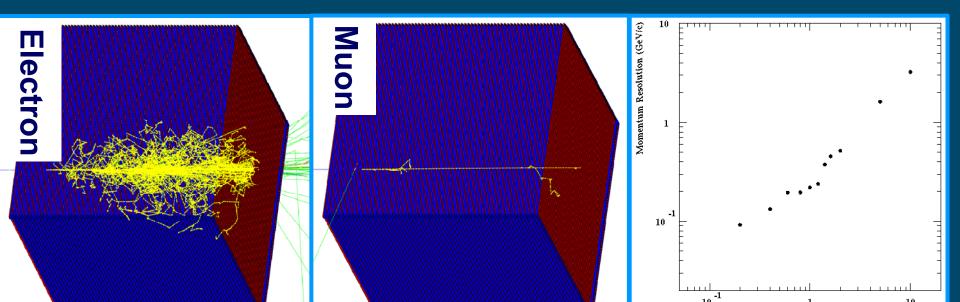
The geometry of the MECC is being optimized

Magnetic sampling calorimeter

Concept:

- Magnetised iron?
 - Sampling fraction?
- Air toroid
- o Cost
 - ~ \$300M



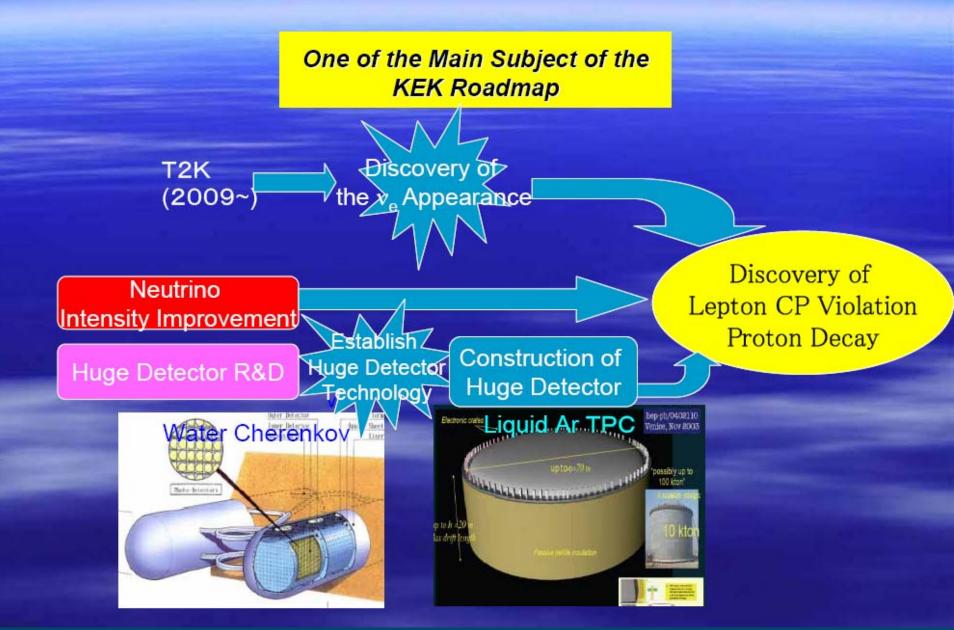


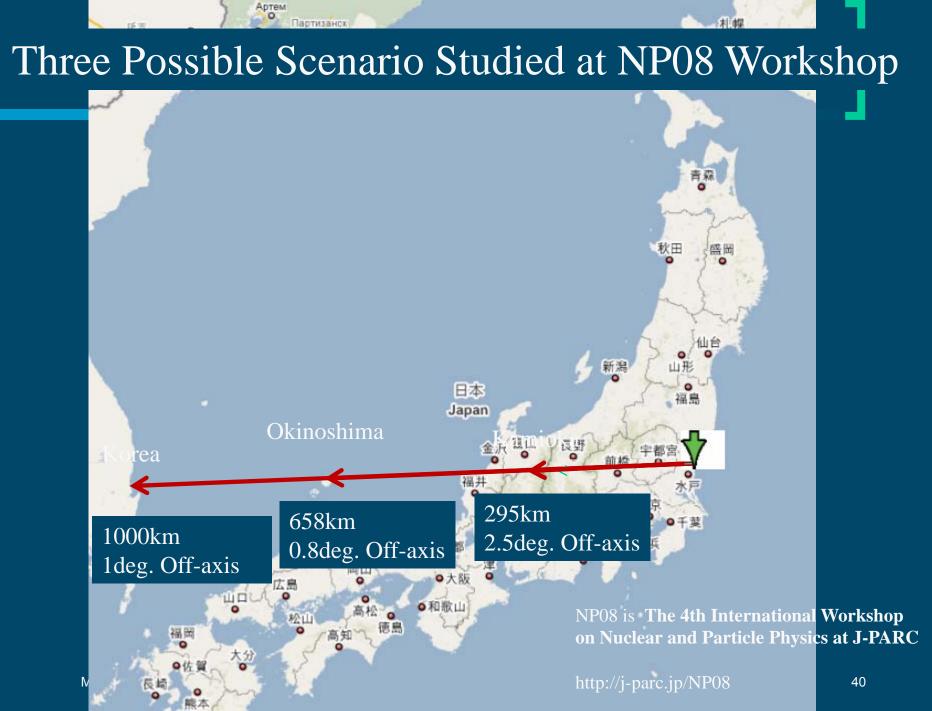
Possible Next Generation Superbeams

Typical parameters

- Proton driver 2 4 Mw
- Long Baselines
- At least two L/E
- Far detector very large which could also search for proton decay
- o Japan
 - Boost power at JPARC to ~4 MW
 - T2HK ~1 Mton water Cerenkov or very large Liquid Argon
 - T2KK split detector half in Japan, half in Korea
 - T2K to Okinoshima

Quest for the Origin of Matter Dominated Universe





Possible Next Generation Superbeams

o USA

- Project X
 - o Initially 150 kW at 8 GeV, 2 MW at 120 GeV
 - First to boost beam to Nova
 - Later to produce a new wide band beam to a new detector, Water Cerenkov or Liquid Argon at DUSEL
- o CERN
 - SPL
 - Replacement for present PS booster, part of the LHC luminosity upgrade
 - Low power and high power versions
 - High Power version, ~4 MW at 5 GeV for neutrino and ISOLDE programmes
 - Could produce a new neutrino beam to Frejus (Memphis Water Cerenkov) or possibly to Gran Sasso

Possible Next Generation Superbeams

o USA

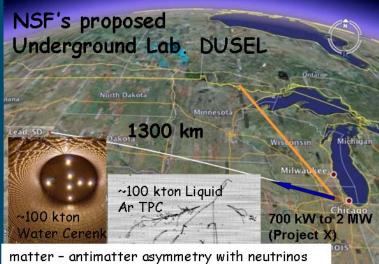
Project X

- Initially 150 kW at 8 GeV, 2 MW at 120 GeV
- First to boost beam to Nova
- Later to produce a new wide band beam to a new detector, Water Cerenkov or Liquid Argon at DUSEL

o CERN

SPL

- Replacement for present PS booster, part of the LHC luminosity upgrade
- Low power and high power versions
- High Power version, ~4 MW at 5 GeV for neutrino and ISOLDE programmes
- Could produce a new neutrino beam to Frejus (Memphis Water Cerenkov) or possibly to Gran Sasso



matter – antimatter asymmetry with neutrinos Proton decay – Supernovae

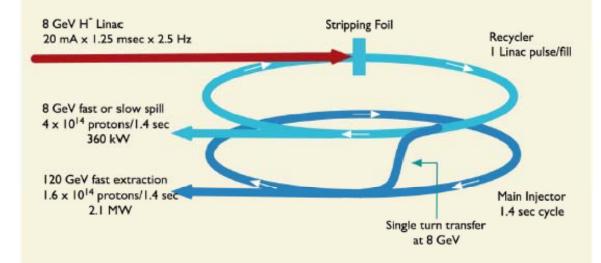


FNAL – Project X



Initial Configuration

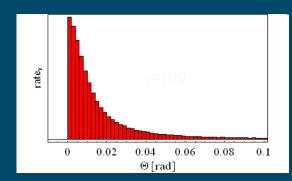
- A multi-megawatt proton source is considered a central and cohesive element for future accelerator developments needed in the exploration of both Energy and Intensity Frontiers at Fermilab
- Project X Design Criteria
 - 2 MW of beam power over the range 60 120 GeV;
 - Simultaneous with at least 150 kW of beam power at 8 GeV;
 - Compatibility with future upgrades to 2-4 MW at 8 GeV



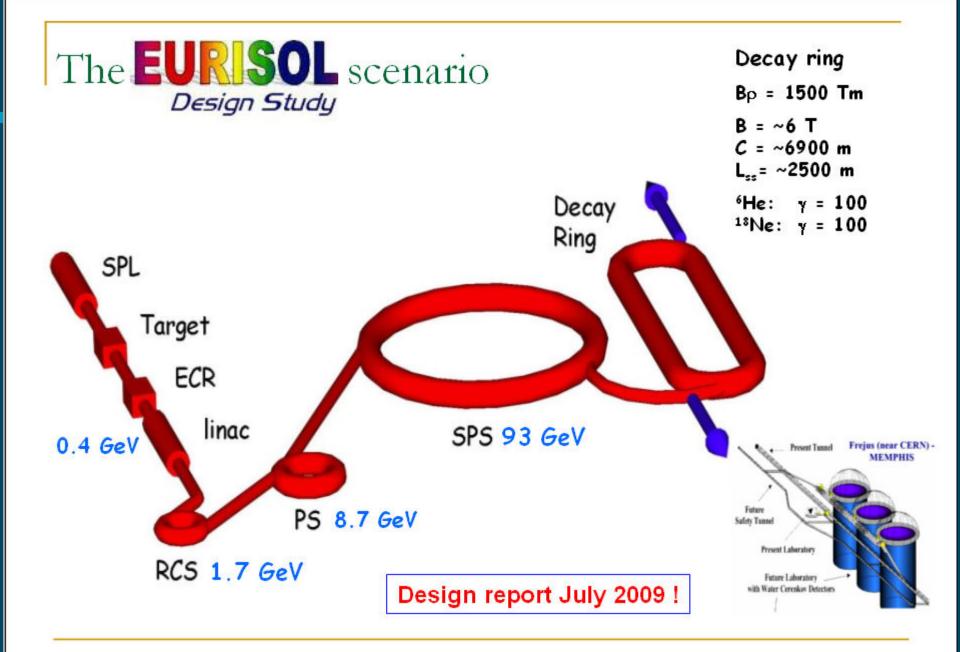
Beta Beams

- Storage ring with β-decaying ions (P Zucchelli)
 - β^- decay => anti- v_{e_i} β^+ decay => v_{e_i} Very pure beam ideal for golden channel
- Choice of lons
 - Lifetime ~1 sec, shorter, cannot store, longer too few v's
 - Low Z better lower mass/charge/v
 - High E desired, $\sigma \alpha E_{\nu}$, $E_{\nu} \alpha Q_{decay} E_{ion}$, Divergence $\alpha 1/E_{ion}$
 - o Choices
 - Original γ_{ion} = 100 (feasible at CERN with current SPS)
 - ${}^{6}\text{He}_2 \text{ for } v, Q = 3.5 \text{ MeV}, < E_v > ~ 350 \text{ MeV}$
 - ¹⁸Ne₁₀ for v-bar, Q = 3.0 MeV, $\langle E_v \rangle \sim 300$ MeV
 - Improved γ = 350 (needs now PS and SPS at CERN)
- High Q versions (C Rubbia)
 - Increase energy using high-Q decays rather than high γ
 - ⁸Li₃ for ν, Q = 13.0 MeV
 - ⁸B₅ for v-bar, Q = 13.9 MeV

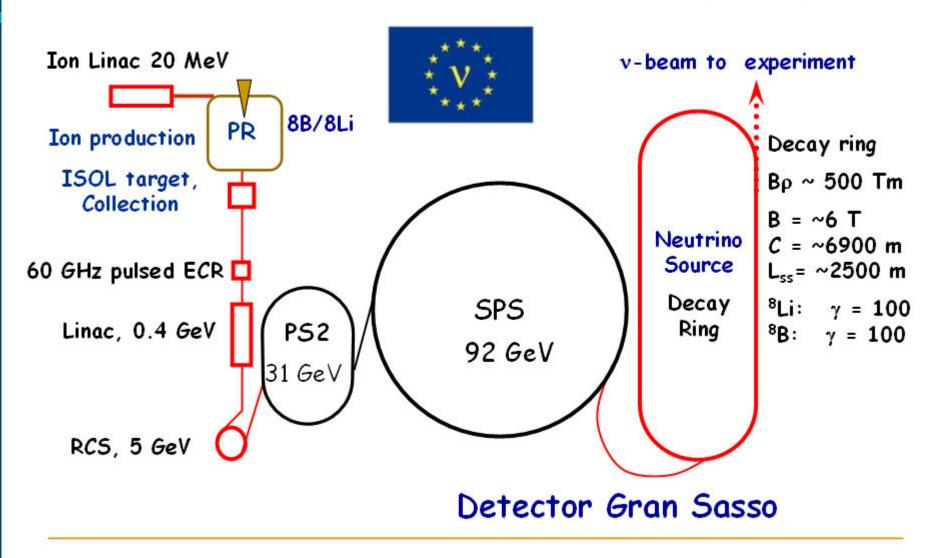
Solution 0 0.5 1 1.5 2 2.5 3 3.5 E_v [MeV]



Production of sufficient ions is a major challenge, ⁶He₂ OK, ¹⁸Ne₁₀, ⁸Li₃ hard, ⁸B₅ appears very difficult.



Beta Beam scenario EUROv (FP7)



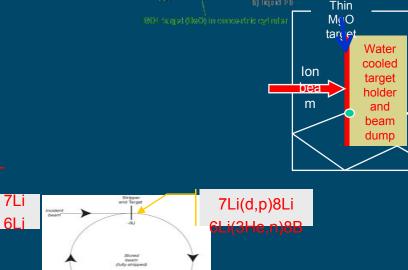
Ion production (from Elena Wildner)

ISOL method at 1-2 GeV (200 kW)

- >1 10¹³ ⁶He per second
- <8 10^{11 18}Ne per second
- o Studied within EURISOL
- Direct production
 - >1 10¹³ (?) ⁶He per second
 - 1 10¹³ ¹⁸Ne per second
 - o ⁸Li ?
 - o Studied at LLN, Soreq, WI and GANIL

Production ring

- o 10¹⁴ (?) ⁸Li
- o >10¹³ (?) ⁸B



^BHe and ⁴He

Transfer line 🦷

N.B. Nuclear Physics has limited interest in those elements ->> Production rates not pushed!

E Wildner New Opportunities in the Physics Landscape at CERN May 2009

Neutrino Factory

- Neutrinos from muon decay in a storage ring
- Can arrange for alternate bunches to be from μ^+ and μ^-

 v_e and \overline{v}_{μ} from μ^+ \overline{v}_e and v_{μ} from μ^-

Very pure beams but must have charge identification in the detector to establish 'wrong' sign product

The performance of the detector is very important - magnetisable, and ideally

as low a threshold as possible for μ sign determination electron detection sign determination tau detection appears feasible with magnetized emulsion detector (MECC) or liquid Argon

Has the greatest potential but still technical problems requiring solution. These are in common with mu collider R&D May 30, 2009 P J Doman - FPCP09

NF Feasibility/Design Studies

Several studies at the turn of the century

- US Studies I, II, IIa
- ECFA/CERN Study
- NuFact-J Study

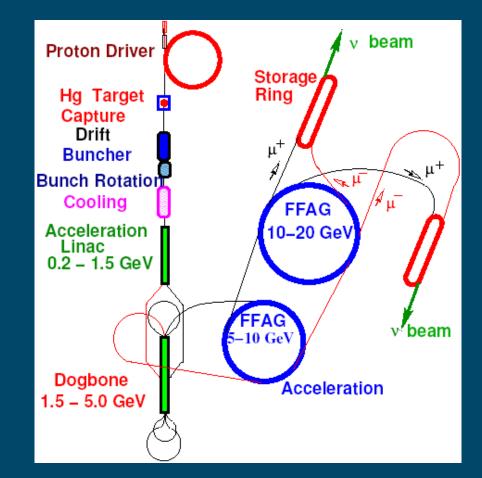
established feasibility & R&D programme

- MUCOOL, MICE, MERIT....
- International Scoping Study (ISS) launched in 2005
 - International integrated accelerator, detector and physics study
 - Completed in 2007 with 3 reports: Accelerator, Detectors & Physics case,(now accepted for publication) `- include a comparison with performance of superbeams and beta-beams
 - Produced baseline design for
- International Design Study (IDS-NF) launched 2008
 - Goal to produce a CDR inc approx costing by 2012
 - Being conducted in conjunction with NFMC Collaboration in US, EuroNu Design Study in Europe, UKNF in UK + others.

Accelerator Subsystems

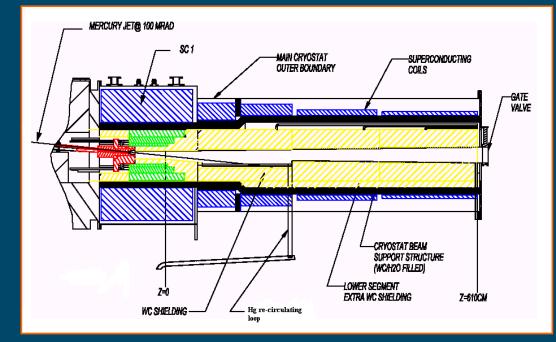


- Target and capture
- Front end
 - Bunching and phase rotation
 - o Cooling
- Acceleration
- Decay ring



Target – Major difficulty at 4 MW

Baseline - Mercury Jet

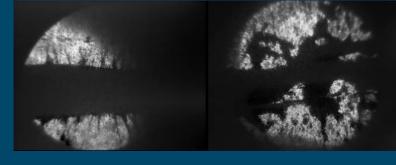


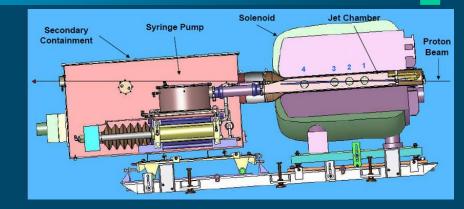
MERIT EXot serves as a satisfactory proofof-principle of Hgjet concept (Ilias Efthymiopoulos)

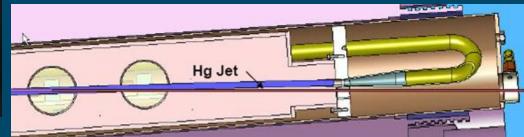
Field tapers from 20 T, 15 cm to 1.75 T, 60 cm over 20 m

MERIT expt at CERN (2008)

 High-power liquidmercury jet target engineering demonstration







CONCLUSIONS

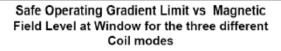
Power handling of target is adequate disruption length of 28 cm \Rightarrow 70 Hz rep. rate at 20 m/s 115 kJ per pulse x 70 Hz gives 8 MW of beam power 4 MW design value seems <u>May 30, 2009</u> P J Dornan - FPCP09

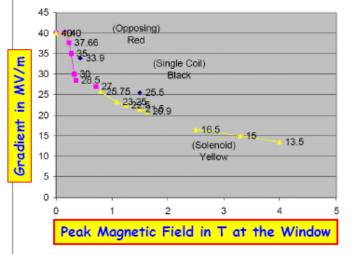


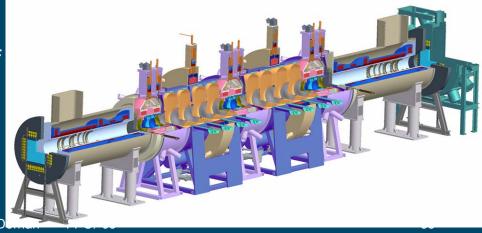
Cooling: R&D programme

- 2 Complementary programmes:
 - MuCool: (FNAL)
 - Design, prototype, and test using an intense proton beam – cooling channel components
 - o MICE: (RAL)
 - Design, construct, commission, and operate – in a muon beam – a section of cooling channel and measure its performance in a variety of modes

MICE results 2012-13

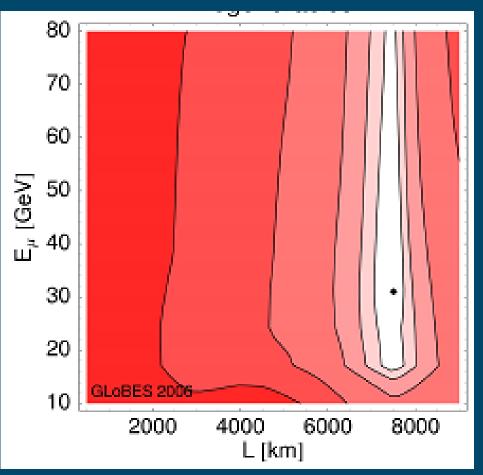






NuFact - sin²20₁₃ sensitivity

ISS Physics Report

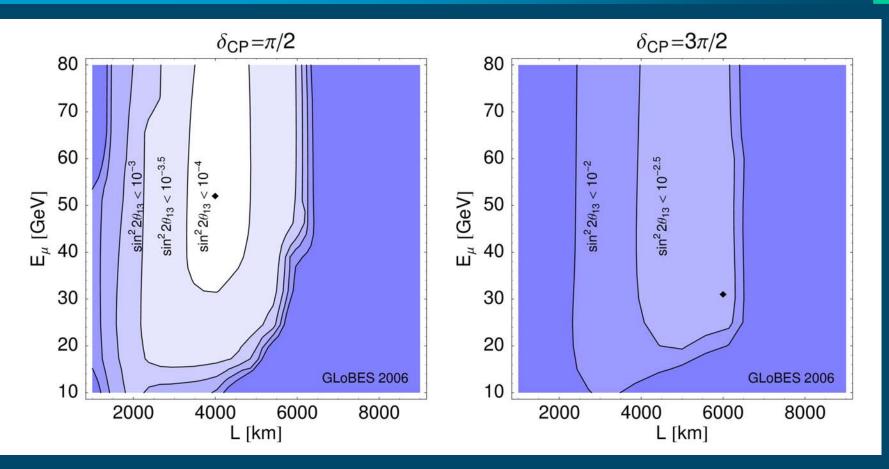


Taking into account statistics, systematics, correlations, degeneracies

Magic baseline (7500 km) good degeneracy solver

Stored muon energy > 20 GeV

NuFact - CP sensitivity



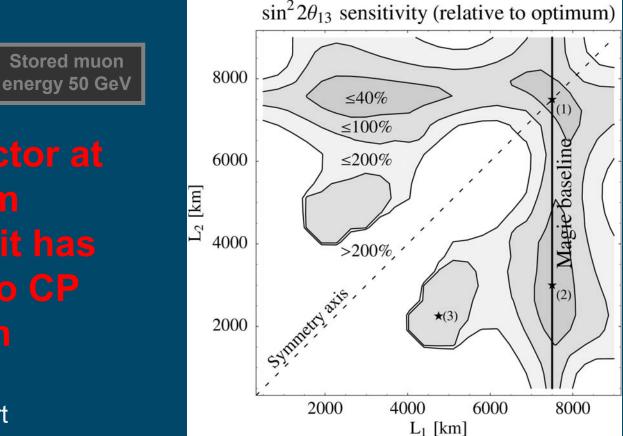
Baseline: 3000 – 5000 km Stored-muon energy > 30 GeV

ISS Physics Report

May 30, 2009

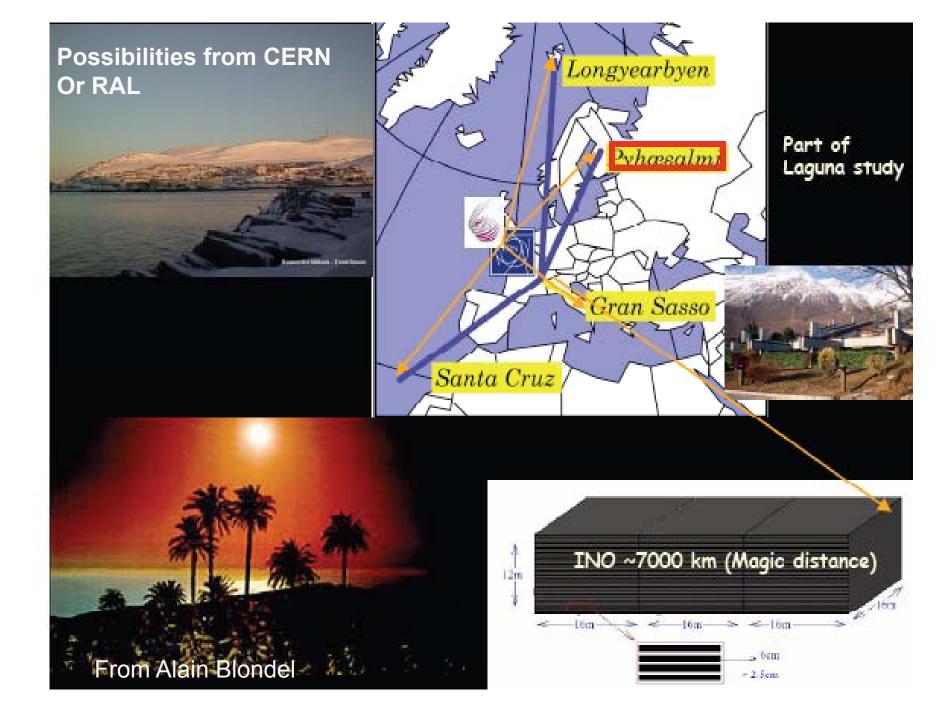
NuFact - Multiple baselines:

Performance for two 25kT detectors relative to the performance for one 50 kT detector at the magic baseline

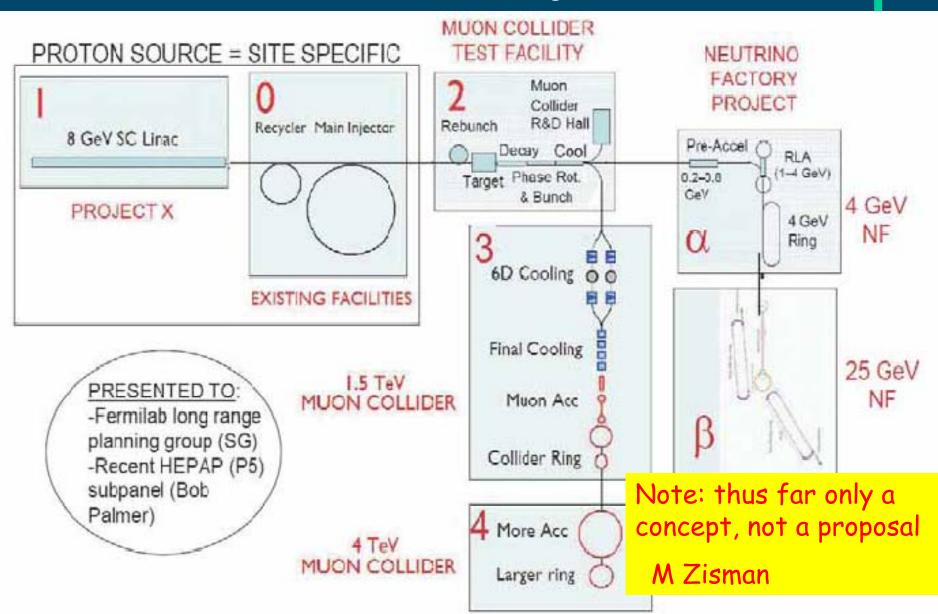


Second detector at ~3000 km preferred as it has sensitivity to CP violation

ISS Physics Report

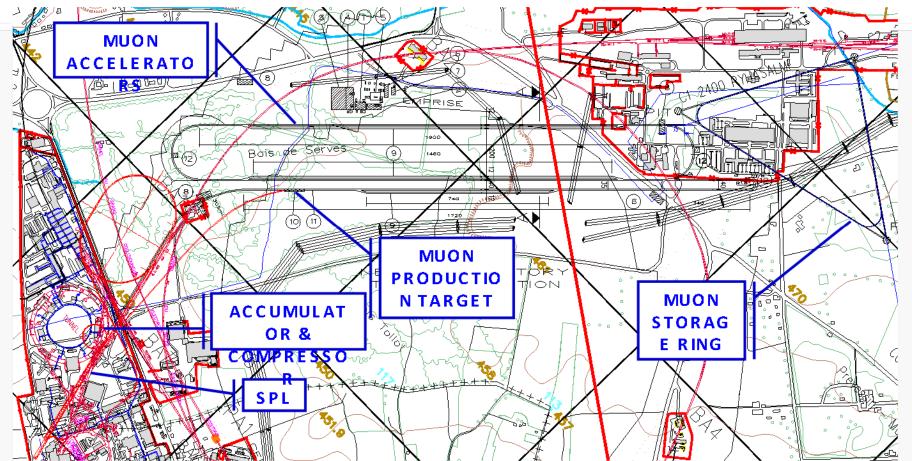


A U.S. Scenario / Project X



<u>v FACTORY</u>

 \checkmark The layout of the future injectors is compatible with a v Factory at CERN.



From Roland Garoby

R.G. @ ETH Zurich 18/11/2008

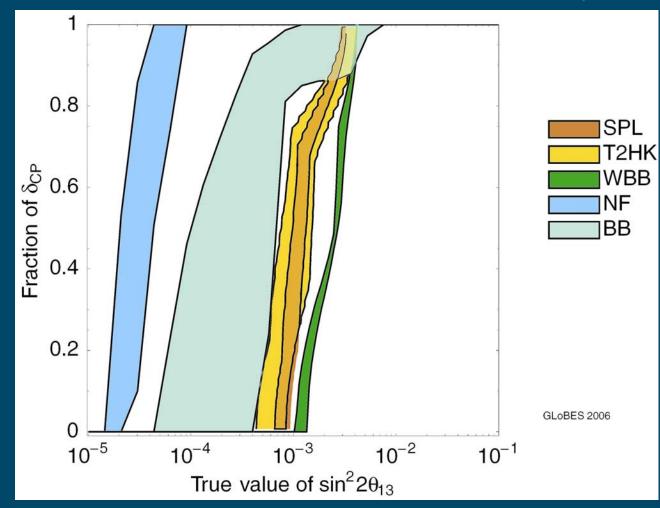


A performance comparison (ISS)

- Super-beam:
 - T2HK, SPL CERN beam to Frejus, A Long baseline Wide-band Beam
- Beta beam:
 - Low γ : γ = 100 and *L* = 130 km
 - Flux (~10¹⁸ decays per year)
 - High γ : γ = 350 and *L* = 700 km
 - Flux (~10¹⁸ decays per year)
 - Also Beta beam abd superbeam combination
- o Neutrino Factory
 - Performance studied as a function of:
 - E and L
 - E_{thresh} and E_{Res}
- The bands correspond to basic/enhanced performance

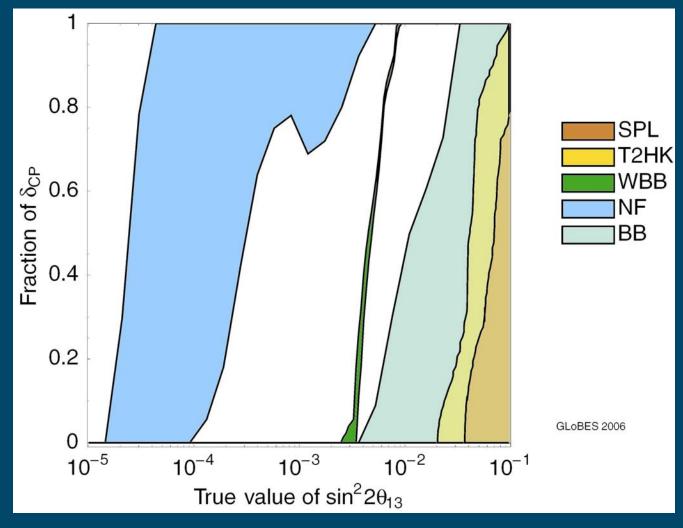
For θ_{13}

ISS Physics Report



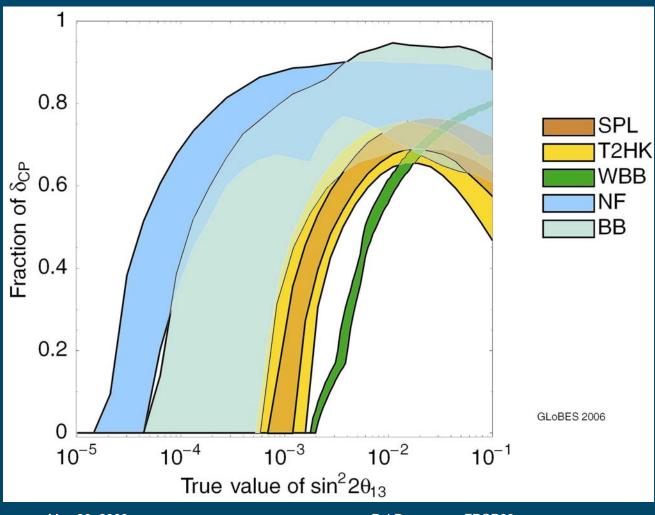
For the Mass Hierarchy

ISS Physics Report



For CP violation

ISS Physics Report



Summary

- The Lepton Sector has shown the first sign of BSM physics and may hold the key to the Flavour problem, CP violation and a viable GUT
- Progress will need precision measurements of the oscillation parameters (comparable with CKM)
- This will demand major construction beyond T2K and Nova to produce intense neutrino beams and very large detectors
- Realistic proposals exist but both beams & detectors still require R&D to show viabiliy. This is taking place now.
- By 2012 2014 there should be enough information, R&D results, cost estmates and crucially the value of θ_{13} to decide the best approach
- With funding, a precision era experiment could take data by 2020-25
- It would be wonderful to know if neutrinos are Majorana and what is m_1 ,



European Strategy for Future Neutrino Physics

A workshop to discuss the possibilities for future neutrino investigations in Europe and the links to CERN

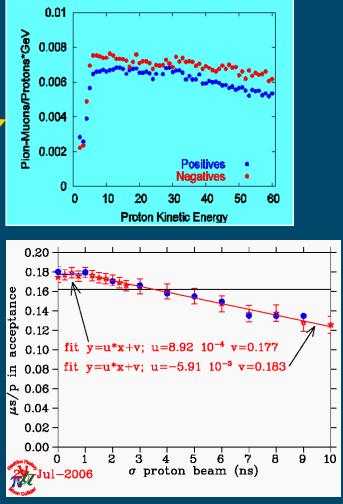
CERN Oct 1-3, 2009

All Welcome

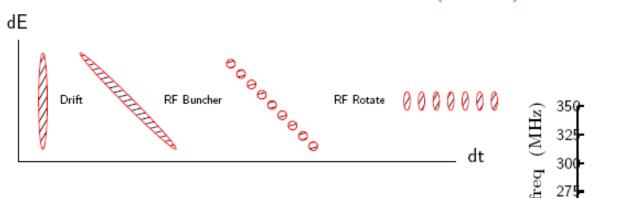


Proton Driver

Baseline Parameters MARS14 0.01 Pion-Muons/Protons*GeV 0.008 0.006 0.004 Parameter <u>Value</u> 0.002 10 ± 5 Energy (GeV) Beam power (MW) 4 10 20 Repetition rate (Hz) ≈50 0.20 No. of bunch trains 3,5^{a)} 0.18 acceptance 0.16 0.14 Bunch length, rms (ns) 2 ± 1 0.12 0.10 -Beam duration^{b)} (µs) ≈40 0.08



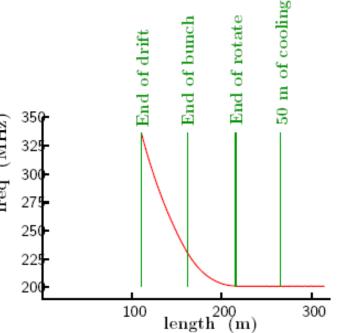
Phase Rotation & Bunching



 RF frequency must vary along bunching channel (high mom. bunches move faster than low)

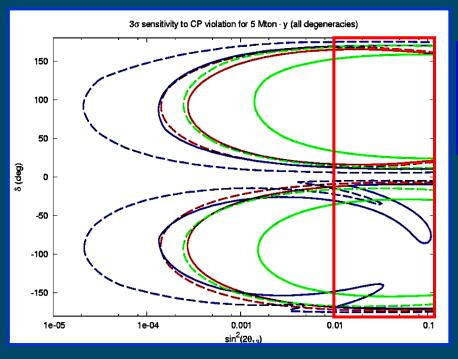
Bunched Beam Rotation with 200 MHz RF (Neuffer)

 Bunched Beam method captures both signs in interleaved bunches



Requires R & D

Beta beam, super beam comparison

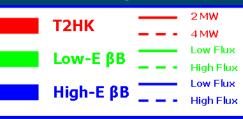


• High θ_{13} :

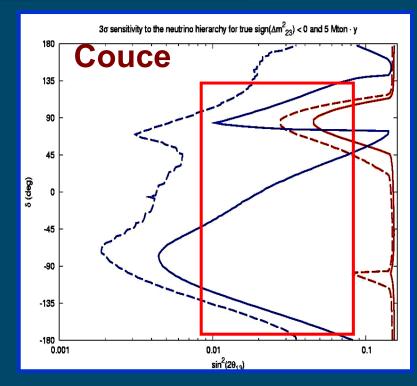
• Beta beam & super beam alone:

- δ sensitivity good
- Poor sign(Δm_{23}^2) sensitivity

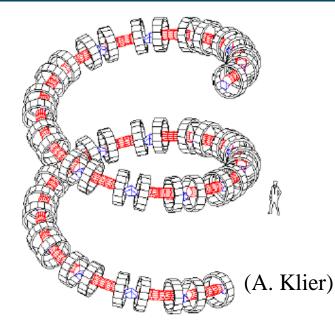
Preliminary! Need to include better



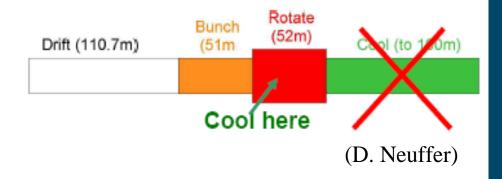
treatment of background and sys. err.



Front end: new configurations



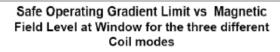
- "Guggenheim" cooling channel
- provides longitudinal cooling
- solves problems with injection, absorber heating
- can taper parameters

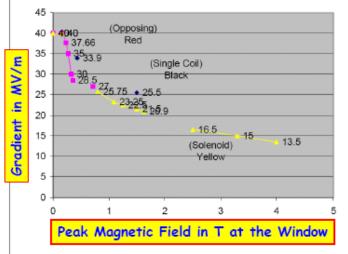


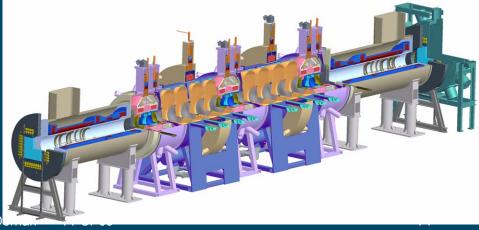
- cool while doing phase rotation
- cost savings
- 150 atm hydrogen (room temp)
- 24 MV/m RF
- performance looks promising

Cooling: hardware R&D programme

- Complementary programmes:
 - MuCool:
 - Design, prototype, and test using an intense proton beam – cooling channel components
 - MICE:
 - Design, construct, commission, and operate – in a muon beam – a section of cooling channel and measure its performance in a variety of modes
- Both programmes well advanced





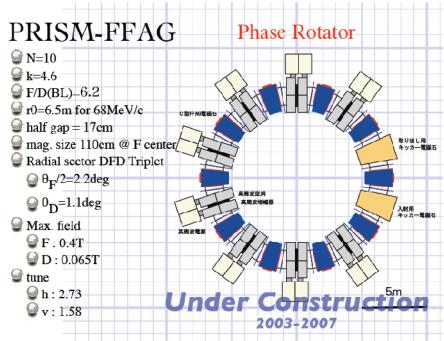


Acceleration: FFAG development

Increasing effort on scaling and non-scaling FFAG

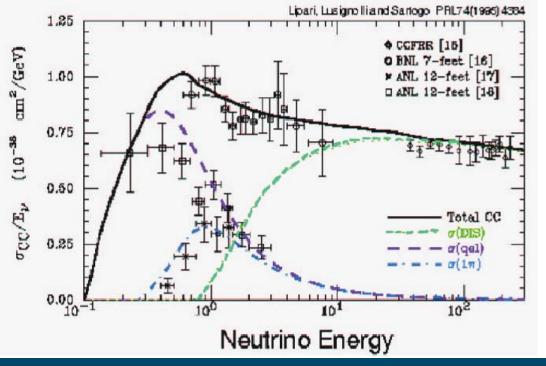
- PRISM: Phase rotated intense muon source
 - Under construction in Osaka Commissioning 2007
 - Proof of principle 'non-scaling' FFAGDecay ring

EMMA :Electron model of muon acceleration



Near detector

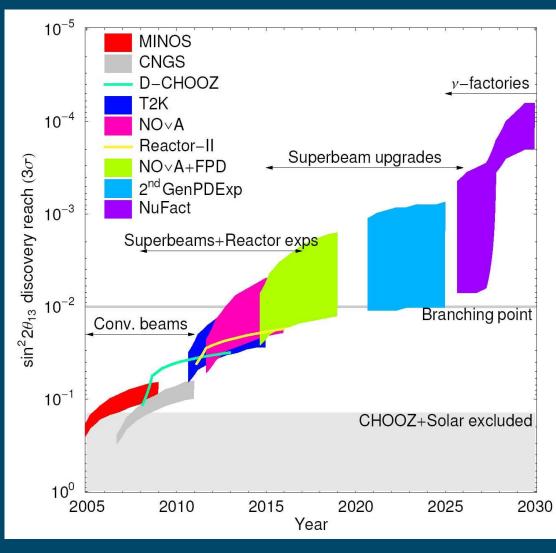
- Measurement of cross sections in DIS, QE and RES.
- **Coherent** π
- Different nuclear targets: H₂, D₂
- Nuclear effects, nuclear shadowing, reinteractions



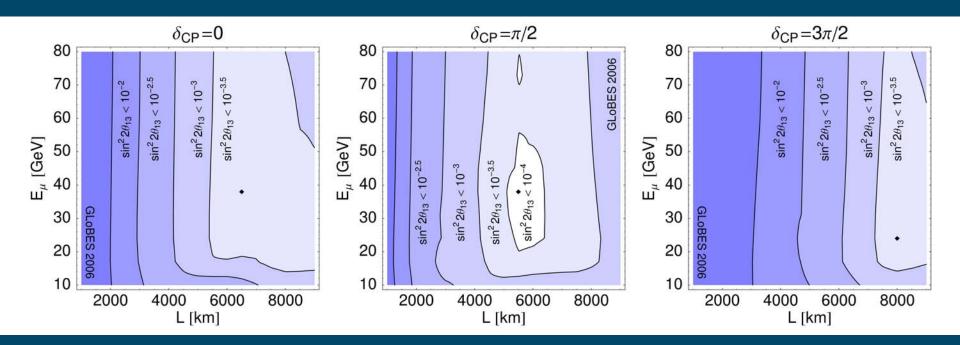
 With modest size targets can obtain very large statistics
 What is lowest energy

what is lowest energy we can achieve? E.g. with LAr can go down to ~MeV

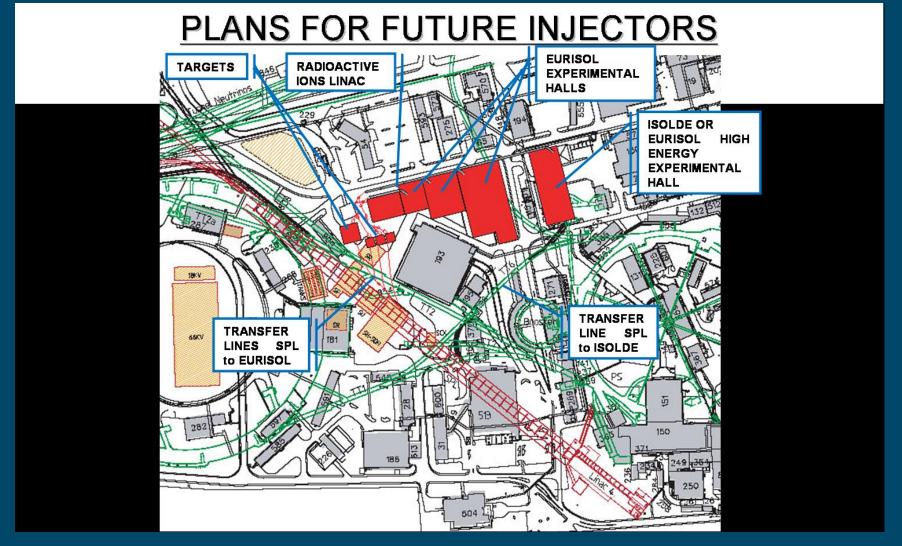
Timescales: the challenge



NuFact - Mass Hierarchy sensitivity



Baseline: ~7500 km Stored muon energy 20 – 50 GeV



May 30, 2009