



ASTROPARTICLE PHYSICS AT ATLAS

Fermi-LAT collaboration meeting

ATLAS Astroparticle Forum Conveners

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WHAT IS THE ATLAS ASTROPARTICLE FORUM

The ATLAS Astroparticle Forum, established in 2012, pools the expertise and interest of the ATLAS community on **astroparticle physics related issues, across physics working groups.**

Mandate of Astroparticle Forum:

- communication of ATLAS results to astrophysics/cosmology communities
- provide suggestions on production of results in complementary parameter space to astroparticle experiments
- coordination of astroparticle physics aspects of ATLAS analyses
- organisation of meetings for discussion of astroparticle topics

For questions and future collaborations:
atlas-phys-astro-forum-conveners@cern.ch

WHAT WE LOOK FOR IN ATLAS

WIMPs

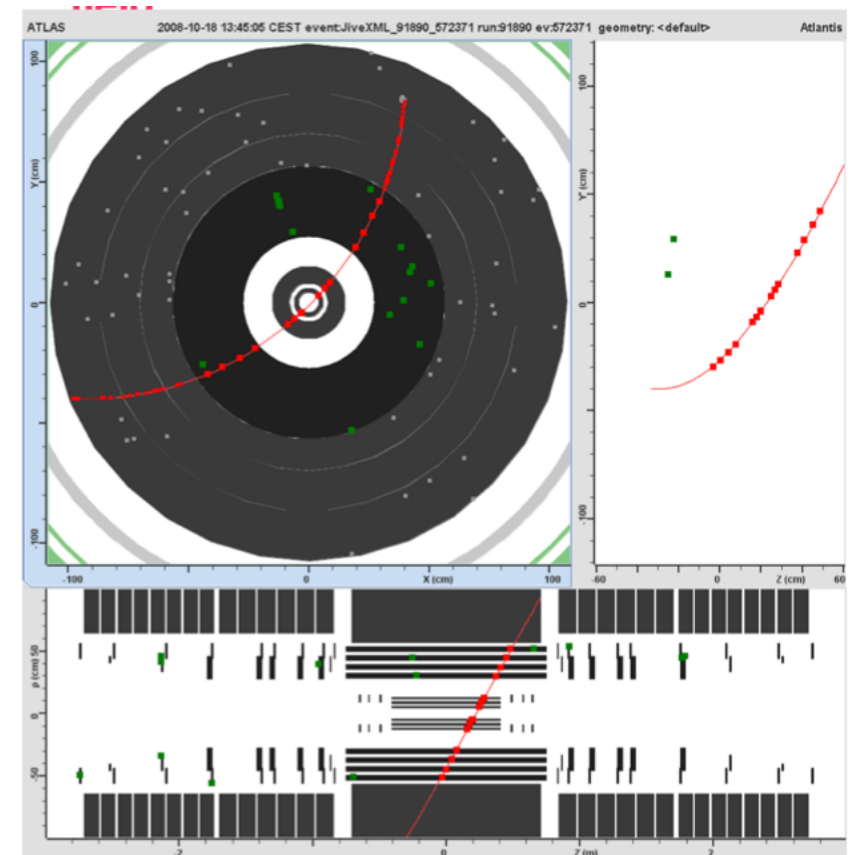
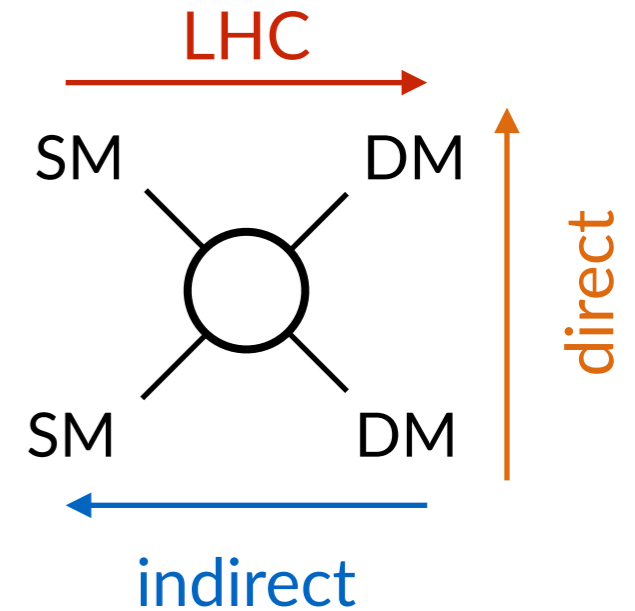
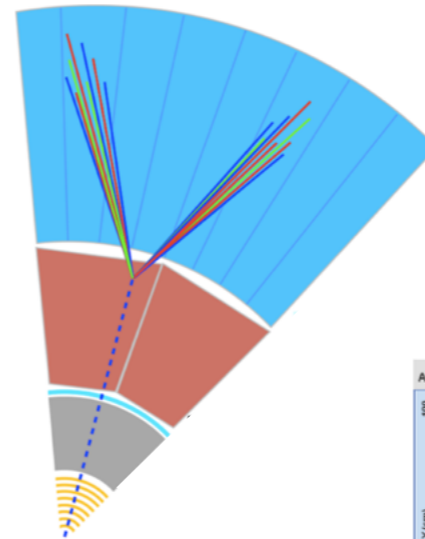
- simplified DM models / effective theories
- supersymmetry

exotic signatures

- long-lived particles

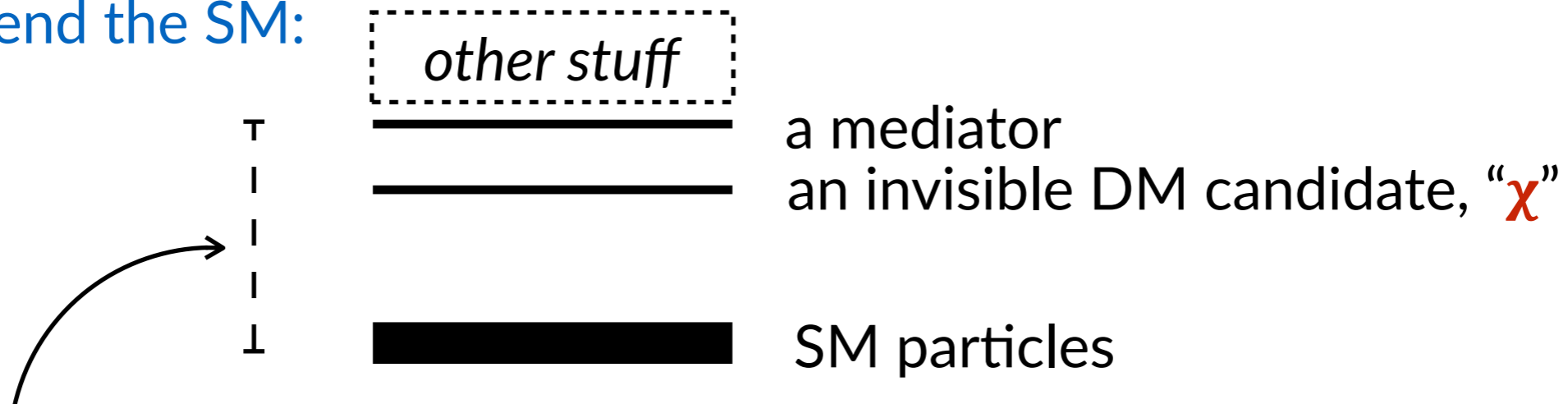
cosmic rays

- w/LHCf: ultra-high-energy showers
- multiplicity measurements



WIMPs: SIMPLIFIED MODEL APPROACH

how we extend the SM:



$\Delta m \gg q^2$: **effective field theory** (like in the case of direct detection)

$\Delta m \lesssim q^2$: use **simplified models**

(simplified Lagrangian w.r.t. UV-complete models like SUSY)

↓
LHC can probe the interaction in detail!

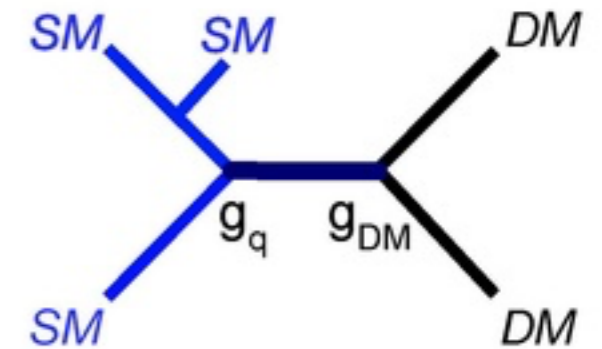
typical degrees of freedom:

- mediator mass and type (vector, axial-vector...)
- DM mass and type (Dirac fermion...)
- mediator couplings

OUR TOOLS

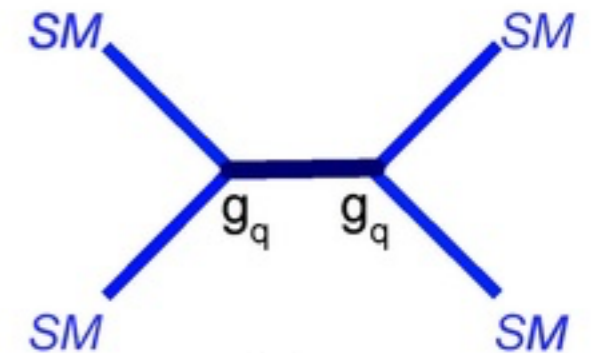
mono- X searches

- produce DM, tag event using a ISR SM particle
- highest sensitivity from mono-jet searches (α_s)
- X can also come from effective DM- X interaction vertex

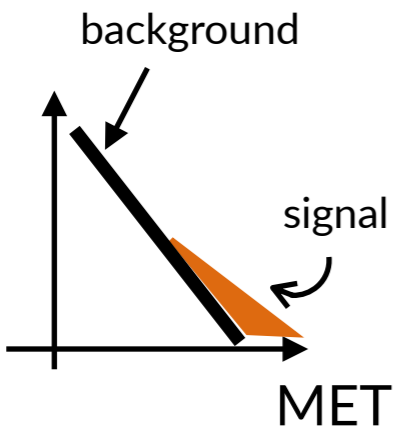
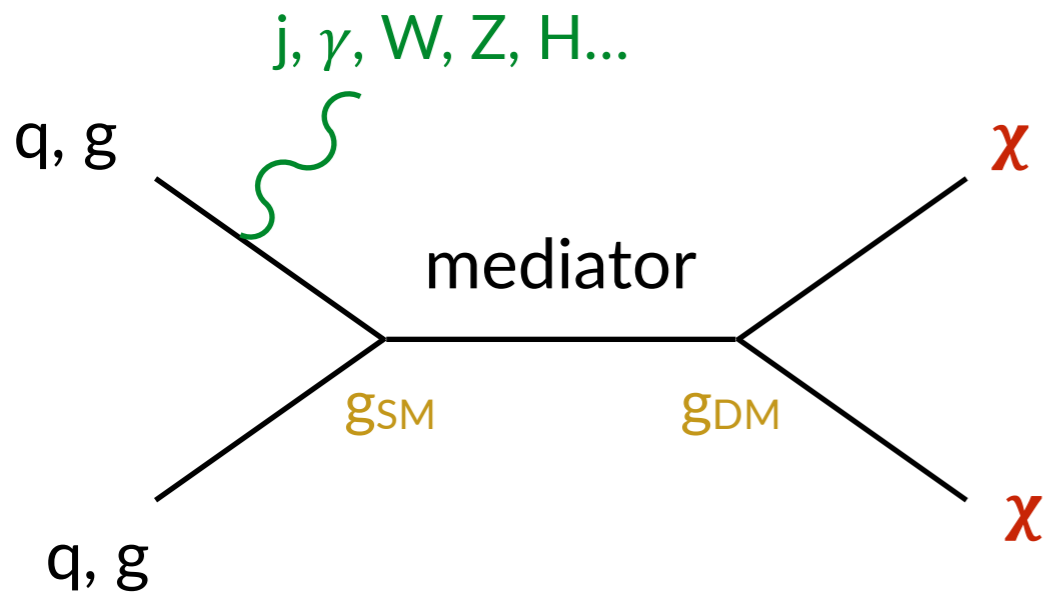


di- X searches

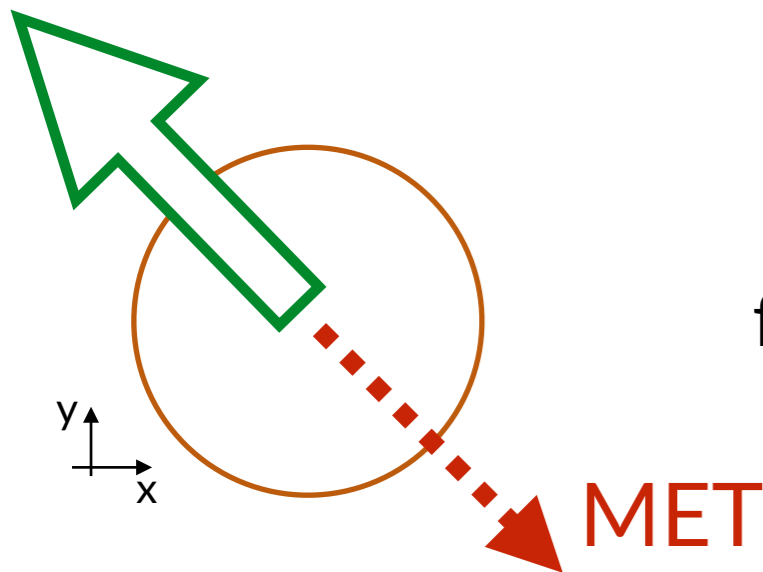
- complementary approach: look for the mediator (resonance), sensitivity mostly independent on m_{DM}
- di-jet searches (limited at low mass by trigger rate)
- di-lepton searches if mediator couples with charged leptons (excellent lepton resolution)



MONO-MANIA



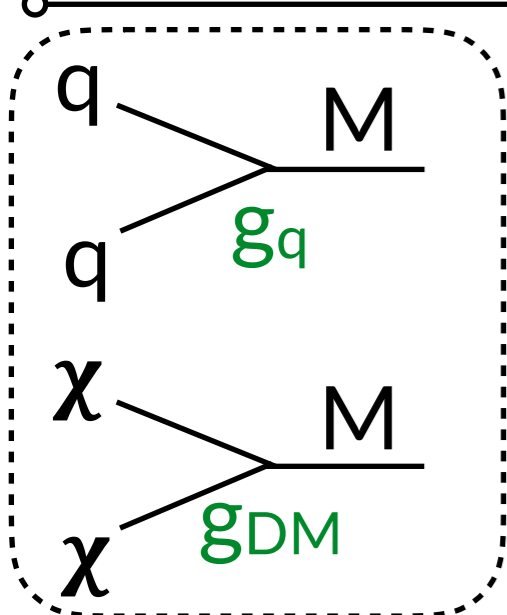
final states with b/t-quarks or Higgs bosons
(mediator couples à la Yukawa with quark masses)



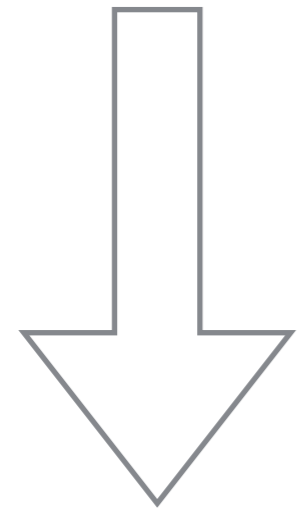
final states with jets

	LHC	DD	ID
scalar	low xsec, soft MET	:	
pseudo-scalar	low xsec, soft MET	: '((velocity suppressed)	:)
vector	large xsec	:) (spin independent)	
axial-vector	large xsec	: ((spin-dependent: experimental issue)	:)

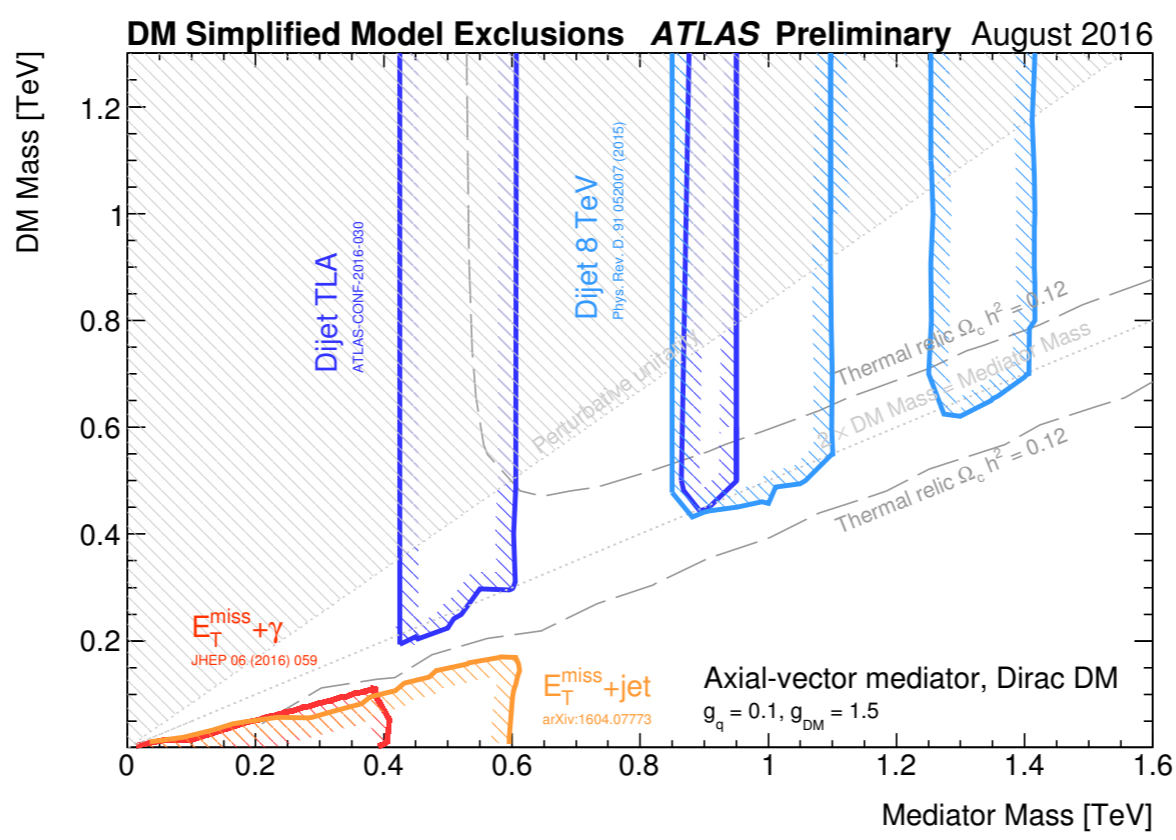
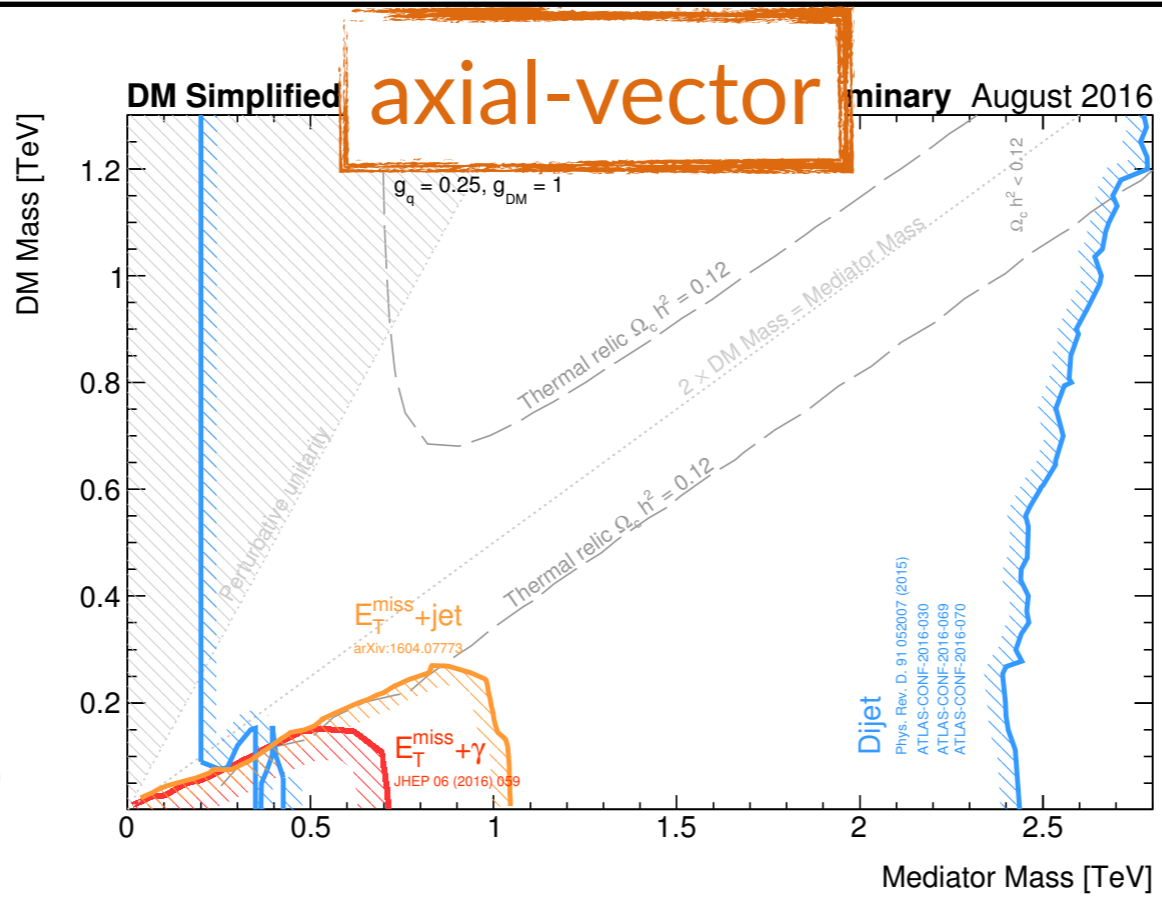
OUR RESULTS



$g_q = 0.25$
 $g_{DM} = 1.0$



$g_q = 0.10$
 $g_{DM} = 1.5$



$$\sigma_{\text{monojet}} \sim g_q g_{DM}$$

$$\sigma_{\text{dijet}} \sim g_q^2$$

coupling assumption strongly influences sensitivity & compatibility with Ω_h^2 :

disadvantage need to explore the parameter space (~2 more degrees of freedom other than masses) -> “re-interpretation”

advantage multi-signature: could characterise a discovery and fully probe SM extensions

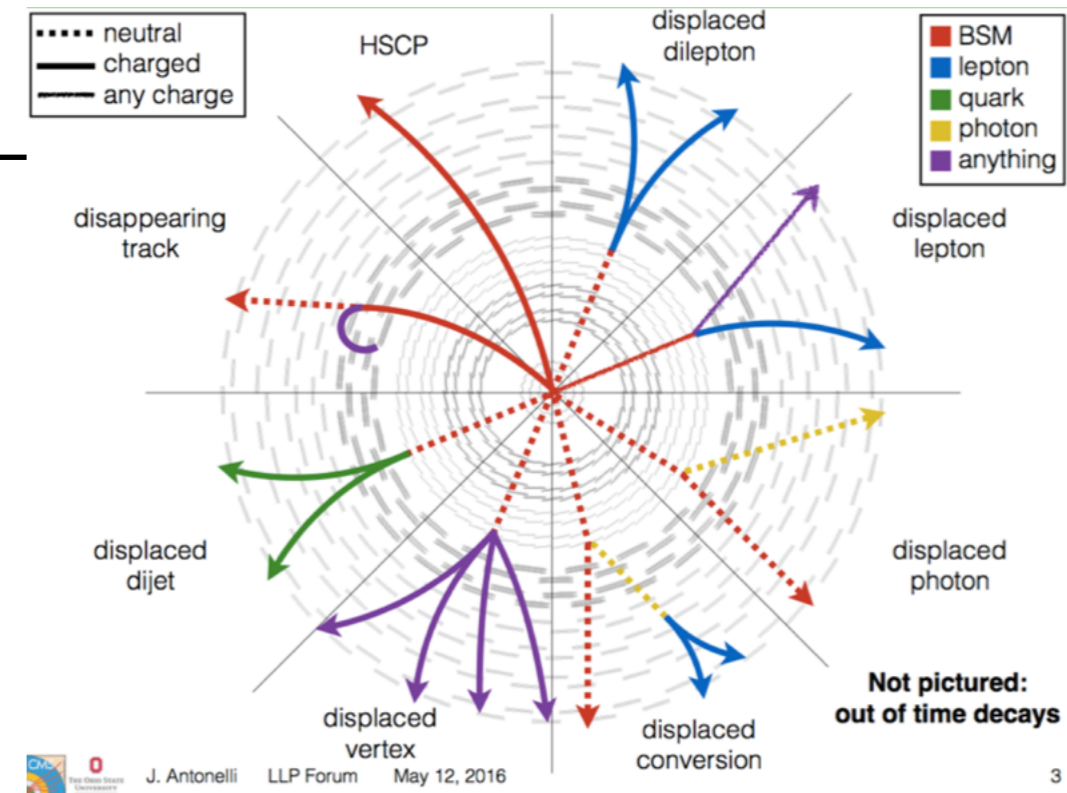
LONG-LIVED PARTICLES

many SM extensions predict long-lived particles

- e.g. RPV supersymmetry

an experimental challenge

- high-luminosity searches (100 fb^{-1} and beyond)
- push trigger, tracking and jet reconstruction at their limit (hits don't start where you expect them to!)



ATLAS Long-lived Particle Searches* - 95% CL Exclusion
 Status: July 2015 ATLAS Preliminary
 $\int \mathcal{L} dt = (18.4 - 20.3) \text{ fb}^{-1}$ $\sqrt{s} = 8 \text{ TeV}$

Model	Signature	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Lifetime limit	Reference	
SUSY	RPV $\chi_1^0 \rightarrow e\nu/\mu\nu/\mu\mu\nu$	20.3	χ_1^0 lifetime 7-740 mm	$m(\tilde{g}) = 1.3 \text{ TeV}, m(\chi_1^0) = 1.0 \text{ TeV}$ 1504.05162	
	GGM $\chi_1^0 \rightarrow Z\tilde{G}$	20.3	χ_1^0 lifetime 6-480 mm	$m(\tilde{g}) = 1.1 \text{ TeV}, m(\chi_1^0) = 1.0 \text{ TeV}$ 1504.05162	
	AMSB $pp \rightarrow \chi_1^+ \chi_1^0, \chi_1^+ \chi_1^-$	20.3	χ_1^\pm lifetime 0.22-3.0 m	$m(\chi_1^\pm) = 450 \text{ GeV}$ 1310.3675	
	AMSB $pp \rightarrow \chi_1^+ \chi_1^0, \chi_1^+ \chi_1^-$	18.4	χ_1^\pm lifetime 1.31-9.0 m	$m(\chi_1^\pm) = 450 \text{ GeV}$ 1506.05332	
	GMSB	non-pointing or delayed γ	20.3	χ_1^0 lifetime 0.08-5.4 m	SPS8 with $\Lambda = 200 \text{ TeV}$ 1409.5542
	Stealth SUSY	2 ID/MS vertices	19.5	\tilde{S} lifetime 0.12-90.6 m	$m(\tilde{g}) = 500 \text{ GeV}$ 1504.03634
Higgs BR = 10%	Hidden Valley $H \rightarrow \pi, \pi_\nu$	20.3	π_ν lifetime 0.41-7.57 m	$m(\pi_\nu) = 25 \text{ GeV}$ 1501.04020	
	Hidden Valley $H \rightarrow \pi, \pi_\nu$	19.5	π_ν lifetime 0.31-25.4 m	$m(\pi_\nu) = 25 \text{ GeV}$ 1504.03634	
	FRVZ $H \rightarrow 2\gamma_d + X$	20.3	γ_d lifetime 14-140 mm	$H \rightarrow 2\gamma_d + X, m(\gamma_d) = 400 \text{ MeV}$ 1409.0746	
	FRVZ $H \rightarrow 4\gamma_d + X$	20.3	γ_d lifetime 15-260 mm	$H \rightarrow 4\gamma_d + X, m(\gamma_d) = 400 \text{ MeV}$ 1409.0746	
Higgs BR = 5%	Hidden Valley $H \rightarrow \pi, \pi_\nu$	20.3	π_ν lifetime 0.6-5.0 m	$m(\pi_\nu) = 25 \text{ GeV}$ 1501.04020	
	Hidden Valley $H \rightarrow \pi, \pi_\nu$	19.5	π_ν lifetime 0.43-18.1 m	$m(\pi_\nu) = 25 \text{ GeV}$ 1504.03634	
	FRVZ $H \rightarrow 4\gamma_d + X$	20.3	γ_d lifetime 26-160 mm	$H \rightarrow 4\gamma_d + X, m(\gamma_d) = 400 \text{ MeV}$ 1409.0746	
300 GeV scalar	Hidden Valley $\Phi \rightarrow \pi, \pi_\nu$	20.3	π_ν lifetime 0.29-7.9 m	$\sigma \times \text{BR} = 1 \text{ pb}, m(\pi_\nu) = 50 \text{ GeV}$ 1501.04020	
	Hidden Valley $\Phi \rightarrow \pi, \pi_\nu$	19.5	π_ν lifetime 0.19-31.9 m	$\sigma \times \text{BR} = 1 \text{ pb}, m(\pi_\nu) = 50 \text{ GeV}$ 1504.03634	
900 GeV scalar	Hidden Valley $\Phi \rightarrow \pi, \pi_\nu$	20.3	π_ν lifetime 0.15-4.1 m	$\sigma \times \text{BR} = 1 \text{ pb}, m(\pi_\nu) = 50 \text{ GeV}$ 1501.04020	
	Hidden Valley $\Phi \rightarrow \pi, \pi_\nu$	19.5	π_ν lifetime 0.11-16.3 m	$\sigma \times \text{BR} = 1 \text{ pb}, m(\pi_\nu) = 50 \text{ GeV}$ 1504.03634	
Other	HV $Z'(1 \text{ TeV}) \rightarrow q, q_\nu$	20.3	π_ν lifetime 0.1-4.9 m	$\sigma \times \text{BR} = 1 \text{ pb}, m(\pi_\nu) = 50 \text{ GeV}$ 1504.03634	
	HV $Z'(2 \text{ TeV}) \rightarrow q, q_\nu$	20.3	π_ν lifetime 0.1-10.1 m	$\sigma \times \text{BR} = 1 \text{ pb}, m(\pi_\nu) = 50 \text{ GeV}$ 1504.03634	

$\sqrt{s} = 8 \text{ TeV}$

*Only a selection of the available lifetime limits on new states is shown.

SUPERSYMMETRY

the ultimate multi-signature tool

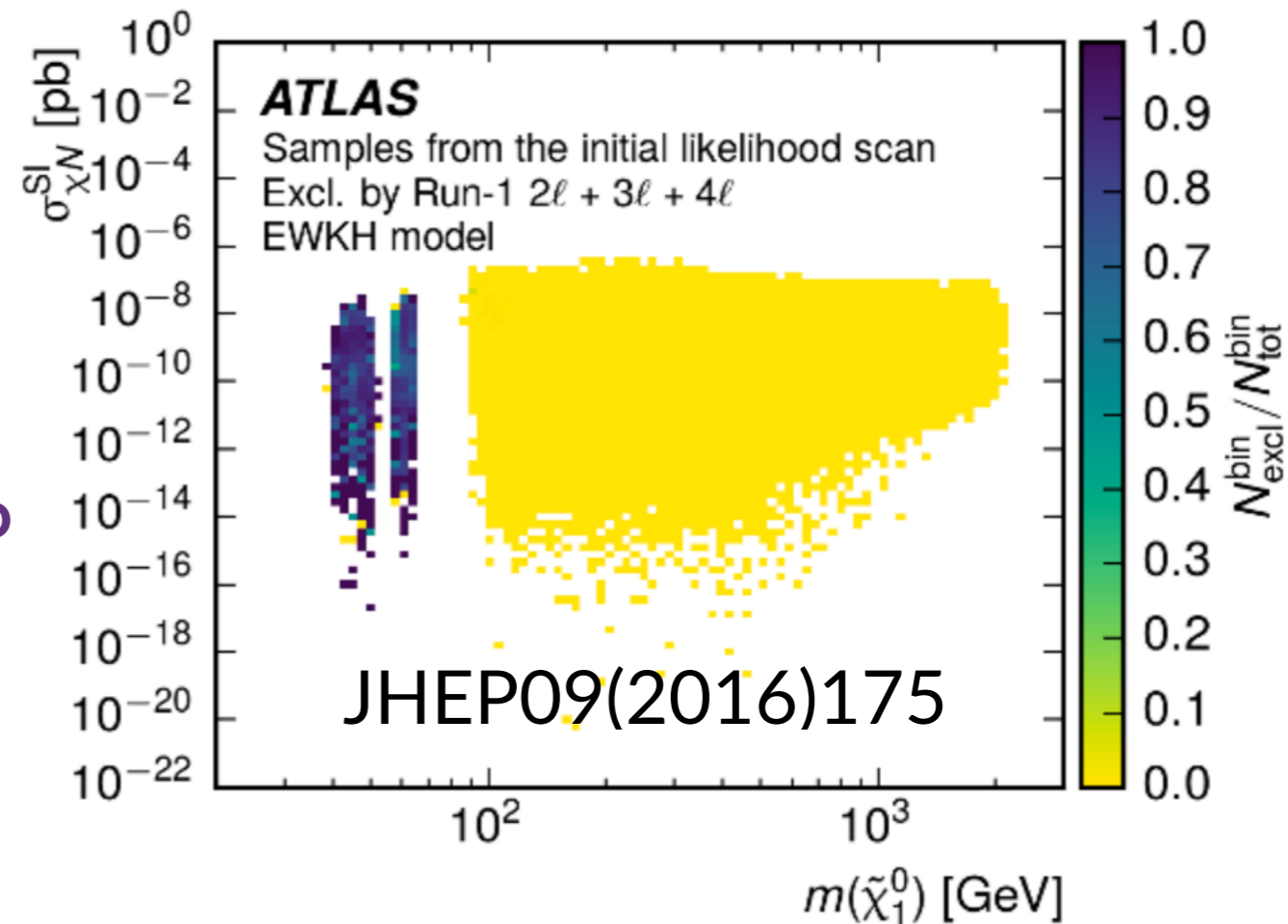
- UV-complete model

if R-parity conserved, WIMP=LSP

- can combine EW neutralino searches in terms of DM parameters
- with more luminosity, exploit Higgsino scenarios (still need >2 years to beat LEP sensitivity ~ 100 GeV!)

if R-parity is violated, meta-stable particle can be DM candidate

- gravitino searches



ATLAS SUSY Searches* - 95% CL Lower Limits

Status: March 2017

ATLAS Preliminary

$\sqrt{s} = 7, 8, 13$ TeV

Model	e, μ, τ, γ	Jets	E_T^{miss}	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Mass limit	$\sqrt{s} = 7, 8$ TeV	$\sqrt{s} = 13$ TeV	Reference	
Inclusive Searches	MSUGRA/CMSSM	0-3 $e, \mu/1-2 \tau$	2-10 jets/3 b	Yes	20.3	\tilde{q}, \tilde{g}	1.85 TeV	$m(\tilde{g})=m(\tilde{q})$	1507.05525
	$\tilde{q}\tilde{q}, \tilde{q}\tilde{q} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{q}	1.57 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(1^{\text{st}} \text{ gen. } \tilde{q})=m(2^{\text{nd}} \text{ gen. } \tilde{q})$	ATLAS-CONF-2017-022
	$\tilde{q}\tilde{q}, \tilde{q}\tilde{q} \rightarrow q\tilde{\chi}_1^0$ (compressed)	mono-jet	1-3 jets	Yes	3.2	\tilde{q}	608 GeV	$m(\tilde{q})-m(\tilde{\chi}_1^0) < 5$ GeV	1604.07773
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow q\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{g}	2.02 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV	ATLAS-CONF-2017-022
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow qq\tilde{\chi}_1^{\pm} \rightarrow qqW^{\pm}\tilde{\chi}_1^0$	0	2-6 jets	Yes	36.1	\tilde{g}	2.01 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV, $m(\tilde{\chi}^{\pm})=0.5(m(\tilde{\chi}_1^0)+m(\tilde{g}))$	ATLAS-CONF-2017-022
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow qq(\ell\ell/\nu\nu)\tilde{\chi}_1^0$	3 e, μ	4 jets	-	13.2	\tilde{g}	1.7 TeV	$m(\tilde{\chi}_1^0) < 400$ GeV	ATLAS-CONF-2016-037
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow qqWZ\tilde{\chi}_1^0$	2 e, μ (SS)	0-3 jets	Yes	13.2	\tilde{g}	1.6 TeV	$m(\tilde{\chi}_1^0) < 500$ GeV	ATLAS-CONF-2016-037
	GMSB ($\tilde{\ell}$ NLSP)	1-2 $\tau + 0-1 \ell$	0-2 jets	Yes	3.2	\tilde{g}	2.0 TeV		1607.05979
	GGM (bino NLSP)	2 γ	-	Yes	3.2	\tilde{g}	1.65 TeV	$c\tau(\text{NLSP}) < 0.1$ mm	1606.09150
	GGM (higgsino-bino NLSP)	γ	1 b	Yes	20.3	\tilde{g}	1.37 TeV	$m(\tilde{\chi}_1^0) < 950$ GeV, $c\tau(\text{NLSP}) < 0.1$ mm, $\mu < 0$	1507.05493
	GGM (higgsino-bino NLSP)	γ	2 jets	Yes	13.3	\tilde{g}	1.8 TeV	$m(\tilde{\chi}_1^0) > 680$ GeV, $c\tau(\text{NLSP}) < 0.1$ mm, $\mu > 0$	ATLAS-CONF-2016-066
	GGM (higgsino NLSP)	2 e, μ (Z)	2 jets	Yes	20.3	\tilde{g}	900 GeV	$m(\text{NLSP}) > 430$ GeV	1503.03290
Gravitino LSP	0	mono-jet	Yes	20.3	$F^{1/2}$ scale	865 GeV	$m(\tilde{G}) > 1.8 \times 10^{-4}$ eV, $m(\tilde{g})=m(\tilde{q})=1.5$ TeV	1502.01518	
3rd gen. \tilde{g} med.	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$	0	3 b	Yes	36.1	\tilde{g}	1.92 TeV	$m(\tilde{\chi}_1^0) < 600$ GeV	ATLAS-CONF-2017-021
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0-1 e, μ	3 b	Yes	36.1	\tilde{g}	1.97 TeV	$m(\tilde{\chi}_1^0) < 200$ GeV	ATLAS-CONF-2017-021
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow b\tilde{t}\tilde{\chi}_1^+$	0-1 e, μ	3 b	Yes	20.1	\tilde{g}	1.37 TeV	$m(\tilde{\chi}_1^0) < 300$ GeV	1407.0600
3rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1\tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$	0	2 b	Yes	3.2	\tilde{b}_1	840 GeV	$m(\tilde{\chi}_1^0) < 100$ GeV	1606.08772
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1\tilde{b}_1 \rightarrow t\tilde{\chi}_1^{\pm}$	2 e, μ (SS)	1 b	Yes	13.2	\tilde{b}_1	325-685 GeV	$m(\tilde{\chi}_1^0) < 150$ GeV, $m(\tilde{\chi}_1^{\pm}) = m(\tilde{\chi}_1^0) + 100$ GeV	ATLAS-CONF-2016-037
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1\tilde{t}_1 \rightarrow b\tilde{\chi}_1^{\pm}$	0-2 e, μ	1-2 b	Yes	4.7/13.3	\tilde{t}_1	117-170 GeV	$m(\tilde{\chi}_1^{\pm}) = 2m(\tilde{\chi}_1^0)$, $m(\tilde{\chi}_1^0) = 55$ GeV	1209.2102, ATLAS-CONF-2016-077
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1\tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$ or $t\tilde{\chi}_1^0$	0-2 e, μ	0-2 jets/1-2 b	Yes	20.3	\tilde{t}_1	90-198 GeV	$m(\tilde{\chi}_1^0) = 1$ GeV	1506.08616, ATLAS-CONF-2017-020
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1\tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$	0	mono-jet	Yes	3.2	\tilde{t}_1	90-323 GeV	$m(\tilde{t}_1)-m(\tilde{\chi}_1^0) = 5$ GeV	1604.07773
	$\tilde{t}_1\tilde{t}_1$ (natural GMSB)	2 e, μ (Z)	1 b	Yes	20.3	\tilde{t}_1	150-600 GeV	$m(\tilde{\chi}_1^0) > 150$ GeV	1403.5222
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2\tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 e, μ (Z)	1 b	Yes	36.1	\tilde{t}_2	290-790 GeV	$m(\tilde{\chi}_1^0) = 0$ GeV	ATLAS-CONF-2017-019
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2\tilde{t}_2 \rightarrow \tilde{t}_1 + h$	1-2 e, μ	4 b	Yes	36.1	\tilde{t}_2	320-880 GeV	$m(\tilde{\chi}_1^0) = 0$ GeV	ATLAS-CONF-2017-019
EW direct	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	2 e, μ	0	Yes	20.3	$\tilde{\ell}$	90-335 GeV	$m(\tilde{\chi}_1^0) = 0$ GeV	1403.5294
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\pm}, \tilde{\chi}_1^{\pm} \rightarrow \tilde{\ell}\nu(\tilde{\ell}\bar{\nu})$	2 e, μ	0	Yes	13.3	$\tilde{\chi}_1^{\pm}$	640 GeV	$m(\tilde{\chi}_1^0) = 0$ GeV, $m(\tilde{\ell}, \bar{\nu}) = 0.5(m(\tilde{\chi}_1^{\pm}) + m(\tilde{\chi}_1^0))$	ATLAS-CONF-2016-096
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\pm}, \tilde{\chi}_1^{\pm} \rightarrow \tilde{\tau}\nu(\tilde{\tau}\bar{\nu})$	2 τ	-	Yes	14.8	$\tilde{\chi}_1^{\pm}$	580 GeV	$m(\tilde{\chi}_1^0) = 0$ GeV, $m(\tilde{\tau}, \bar{\nu}) = 0.5(m(\tilde{\chi}_1^{\pm}) + m(\tilde{\chi}_1^0))$	ATLAS-CONF-2016-093
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_2^0 \rightarrow \tilde{\ell}_L\nu_{\tilde{\ell}_L}(\tilde{\ell}\bar{\nu})$, $\tilde{\nu}\tilde{\ell}_L\ell(\tilde{\nu}\bar{\nu})$	3 e, μ	0	Yes	13.3	$\tilde{\chi}_1^{\pm}, \tilde{\chi}_2^0$	1.0 TeV	$m(\tilde{\chi}_1^{\pm}) = m(\tilde{\chi}_2^0)$, $m(\tilde{\chi}_1^0) = 0$, $m(\tilde{\ell}, \bar{\nu}) = 0.5(m(\tilde{\chi}_1^{\pm}) + m(\tilde{\chi}_1^0))$	ATLAS-CONF-2016-096
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 Z\tilde{\chi}_1^0$	2-3 e, μ	0-2 jets	Yes	20.3	$\tilde{\chi}_1^{\pm}, \tilde{\chi}_2^0$	425 GeV	$m(\tilde{\chi}_1^{\pm}) = m(\tilde{\chi}_2^0)$, $m(\tilde{\chi}_1^0) = 0$, $\tilde{\ell}$ decoupled	1403.5294, 1402.7029
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_2^0 \rightarrow W\tilde{\chi}_1^0 h\tilde{\chi}_1^0$, $h \rightarrow b\tilde{b}/WW/\tau\tau/\gamma\gamma$	e, μ, γ	0-2 b	Yes	20.3	$\tilde{\chi}_1^{\pm}, \tilde{\chi}_2^0$	270 GeV	$m(\tilde{\chi}_1^{\pm}) = m(\tilde{\chi}_2^0)$, $m(\tilde{\chi}_1^0) = 0$, $\tilde{\ell}$ decoupled	1501.07110
	$\tilde{\chi}_2^0\tilde{\chi}_3^0, \tilde{\chi}_2^0\tilde{\chi}_3^0 \rightarrow \tilde{\ell}_R\ell$	4 e, μ	0	Yes	20.3	$\tilde{\chi}_{2,3}^0$	635 GeV	$m(\tilde{\chi}_2^0) = m(\tilde{\chi}_3^0)$, $m(\tilde{\chi}_1^0) = 0$, $m(\tilde{\ell}, \bar{\nu}) = 0.5(m(\tilde{\chi}_2^0) + m(\tilde{\chi}_1^0))$	1405.5086
	GGM (wino NLSP) weak prod.	1 $e, \mu + \gamma$	-	Yes	20.3	\tilde{W}	115-370 GeV	$c\tau < 1$ mm	1507.05493
	GGM (bino NLSP) weak prod.	2 γ	-	Yes	20.3	\tilde{W}	590 GeV	$c\tau < 1$ mm	1507.05493
	Long-lived particles	Direct $\tilde{\chi}_1^+\tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^{\pm}$	Disapp. trk	1 jet	Yes	36.1	$\tilde{\chi}_1^{\pm}$	430 GeV	$m(\tilde{\chi}_1^{\pm}) - m(\tilde{\chi}_1^0) \sim 160$ MeV, $\tau(\tilde{\chi}_1^{\pm}) = 0.2$ ns
Direct $\tilde{\chi}_1^+\tilde{\chi}_1^-$ prod., long-lived $\tilde{\chi}_1^{\pm}$		dE/dx trk	-	Yes	18.4	$\tilde{\chi}_1^{\pm}$	495 GeV	$m(\tilde{\chi}_1^{\pm}) - m(\tilde{\chi}_1^0) \sim 160$ MeV, $\tau(\tilde{\chi}_1^{\pm}) < 15$ ns	1506.05332
Stable, stopped \tilde{g} R-hadron		0	1-5 jets	Yes	27.9	\tilde{g}	850 GeV	$m(\tilde{\chi}_1^0) = 100$ GeV, $10 \mu\text{s} < \tau(\tilde{g}) < 1000$ s	1310.6584
Stable \tilde{g} R-hadron		trk	-	-	3.2	\tilde{g}	1.58 TeV		1606.05129
Metastable \tilde{g} R-hadron		dE/dx trk	-	-	3.2	\tilde{g}	1.57 TeV	$m(\tilde{\chi}_1^0) = 100$ GeV, $\tau > 10$ ns	1604.04520
GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu}) + \tau(e, \mu)$		1-2 μ	-	-	19.1	$\tilde{\chi}_1^0$	537 GeV	$10 < \tan\beta < 50$	1411.6795
GMSB, $\tilde{\chi}_1^0 \rightarrow \gamma\tilde{G}$, long-lived $\tilde{\chi}_1^0$		2 γ	-	Yes	20.3	$\tilde{\chi}_1^0$	440 GeV	$1 < \tau(\tilde{\chi}_1^0) < 3$ ns, SPS8 model	1409.5542
$\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow e\tilde{\nu}/e\mu\nu/\mu\mu\nu$		displ. $ee/e\mu/\mu\mu$	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$7 < c\tau(\tilde{\chi}_1^0) < 740$ mm, $m(\tilde{g}) = 1.3$ TeV	1504.05162
GGM $\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow Z\tilde{G}$		displ. vtx + jets	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$6 < c\tau(\tilde{\chi}_1^0) < 480$ mm, $m(\tilde{g}) = 1.1$ TeV	1504.05162
RPV	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e\mu/\tau\mu/\mu\tau$	$e\mu, e\tau, \mu\tau$	-	-	3.2	$\tilde{\nu}_\tau$	1.9 TeV	$\lambda'_{311} = 0.11, \lambda'_{132/133/233} = 0.07$	1607.08079
	Bilinear RPV CMSSM	2 e, μ (SS)	0-3 b	Yes	20.3	\tilde{q}, \tilde{g}	1.45 TeV	$m(\tilde{q}) = m(\tilde{g}), c\tau_{\text{LSP}} < 1$ mm	1404.2500
	$\tilde{\chi}_1^+\tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow e\tilde{\nu}, e\mu\nu, \mu\mu\nu$	4 e, μ	-	Yes	13.3	$\tilde{\chi}_1^{\pm}$	1.14 TeV	$m(\tilde{\chi}_1^0) > 400$ GeV, $\lambda_{12k} \neq 0$ ($k = 1, 2$)	ATLAS-CONF-2016-075
	$\tilde{\chi}_1^+\tilde{\chi}_1^-, \tilde{\chi}_1^+ \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \tau\tau\nu, e\tau\nu$	3 $e, \mu + \tau$	-	Yes	20.3	$\tilde{\chi}_1^{\pm}$	450 GeV	$m(\tilde{\chi}_1^0) > 0.2 \times m(\tilde{\chi}_1^{\pm}), \lambda_{133} \neq 0$	1405.5086
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow qq\tilde{\chi}_1^0$	0	4-5 large- R jets	-	14.8	\tilde{g}	1.08 TeV	$\text{BR}(t) = \text{BR}(b) = \text{BR}(c) = 0\%$	ATLAS-CONF-2016-057
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow qq\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qq\tilde{\chi}_1^0$	0	4-5 large- R jets	-	14.8	\tilde{g}	1.55 TeV	$m(\tilde{\chi}_1^0) = 800$ GeV	ATLAS-CONF-2016-057
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qq\tilde{\chi}_1^0$	1 e, μ	8-10 jets/0-4 b	-	36.1	\tilde{g}	2.1 TeV	$m(\tilde{\chi}_1^0) = 1$ TeV, $\lambda_{112} \neq 0$	ATLAS-CONF-2017-013
	$\tilde{g}\tilde{g}, \tilde{g}\tilde{g} \rightarrow \tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	1 e, μ	8-10 jets/0-4 b	-	36.1	\tilde{g}	1.65 TeV	$m(\tilde{t}_1) = 1$ TeV, $\lambda_{323} \neq 0$	ATLAS-CONF-2017-013
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	0	2 jets + 2 b	-	15.4	\tilde{t}_1	410 GeV		ATLAS-CONF-2016-022, ATLAS-CONF-2016-084
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\ell$	2 e, μ	2 b	-	20.3	\tilde{t}_1	0.4-1.0 TeV	$\text{BR}(\tilde{t}_1 \rightarrow b\ell/\mu) > 20\%$	ATLAS-CONF-2015-015
Other	Scalar charm, $\tilde{c} \rightarrow c\tilde{\chi}_1^0$	0	2 c	Yes	20.3	\tilde{c}	510 GeV	$m(\tilde{\chi}_1^0) < 200$ GeV	1501.01325

*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

10⁻¹

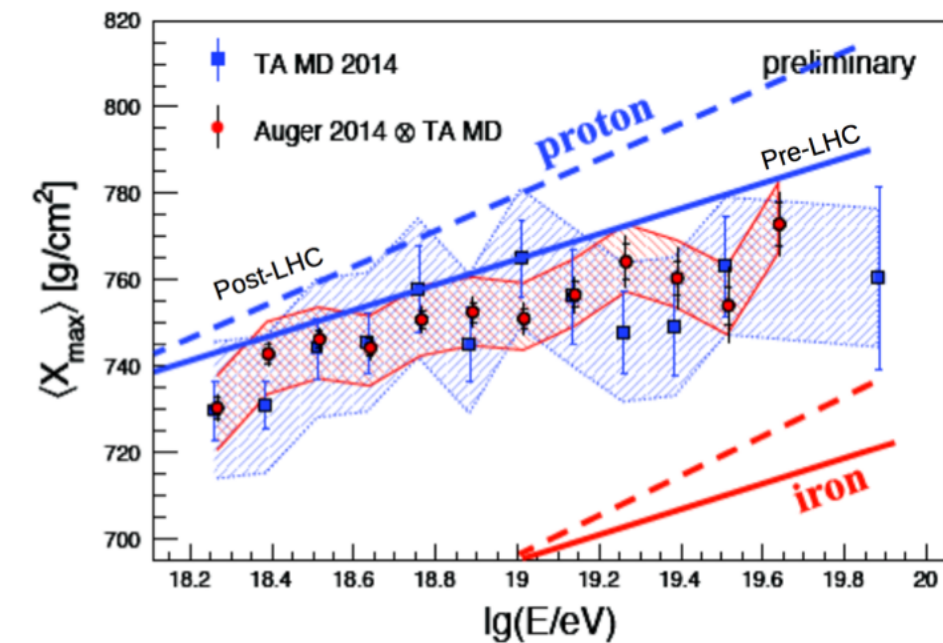
1

Mass scale [TeV]

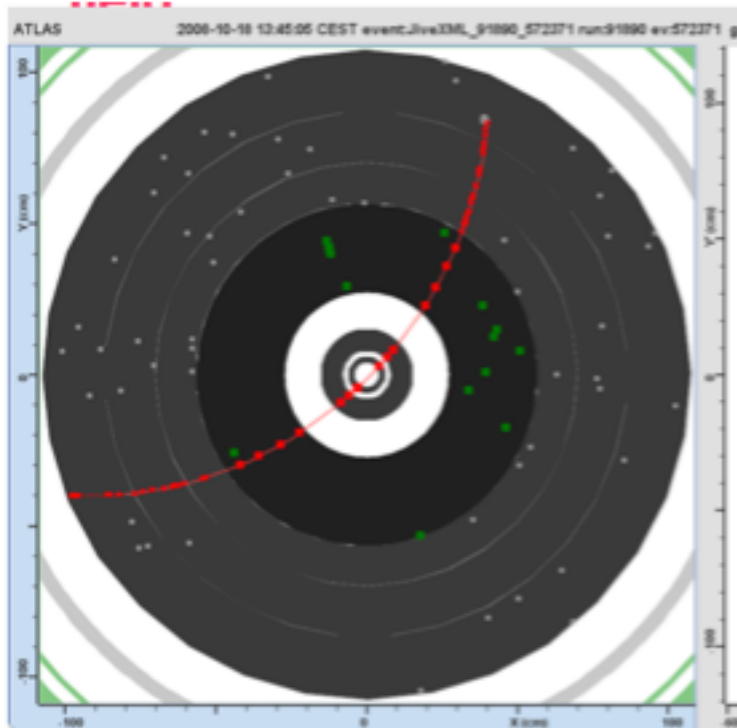
COSMIC RAY PHYSICS

use forward physics to better understand cosmic ray physics

- joint data analysis with LHCf (ATLAS gives track multiplicity in the central region, LHCf measures photons)
- a few million events already collected, can be used to constrain diffractive models
- ATL-PHYS-PUB-2015-038



From Roberto Aloiso talk (2015 working group) at UHECR 2016



ATLAS is a cosmic-ray detector per-se!

- 100 m underground
- dedicated cosmic runs, trigger rate ~ 100 Hz, sensitive to muons with $p > \sim 5$ GeV
- already have a few million events
- - so far used for performance so far (e.g. arXiv: 1011.6665) but could perform dedicated measurements as well (e.g. energy spectrum, cosmic charge ratio...). **Ideas?**

CONCLUSIONS

- ▶ Dark Matter search is a key ingredient in the ATLAS physics programme
 - ❖ start with simplified extensions of the SM
 - ❖ the rest of the BSM sector plays a crucial role
 - nothing is as hard as looking for a black cat in a dark room, especially if there is no cat
 - ◆ fully decoupled (mediator mass > TeV)?
 - ◆ accessible at the LHC? (e.g. SUSY)
 - ◆ particle masses and coupling values determine exclusion contours and predicted values of $\sigma_{\chi N}$ and relic density
 - ❖ access the $m_{\text{DM}} \sim \text{few } 100 \text{ GeV}$ region in the next ~ 6 years
 - ❖ current: mono-X and resonance searches, seek complementarity with DD/ID
 - ❖ work in progress: Higgs-related and more exotic scenarios
- ▶ lots of potential for cosmic ray physics (with LHCf and using ATLAS as a cosmic detector!)

For questions and future collaborations:
atlas-phys-astro-forum-conveners@cern.ch

BACKUP

THE FUTURE: CHALLENGES & COMPLEMENTARITY

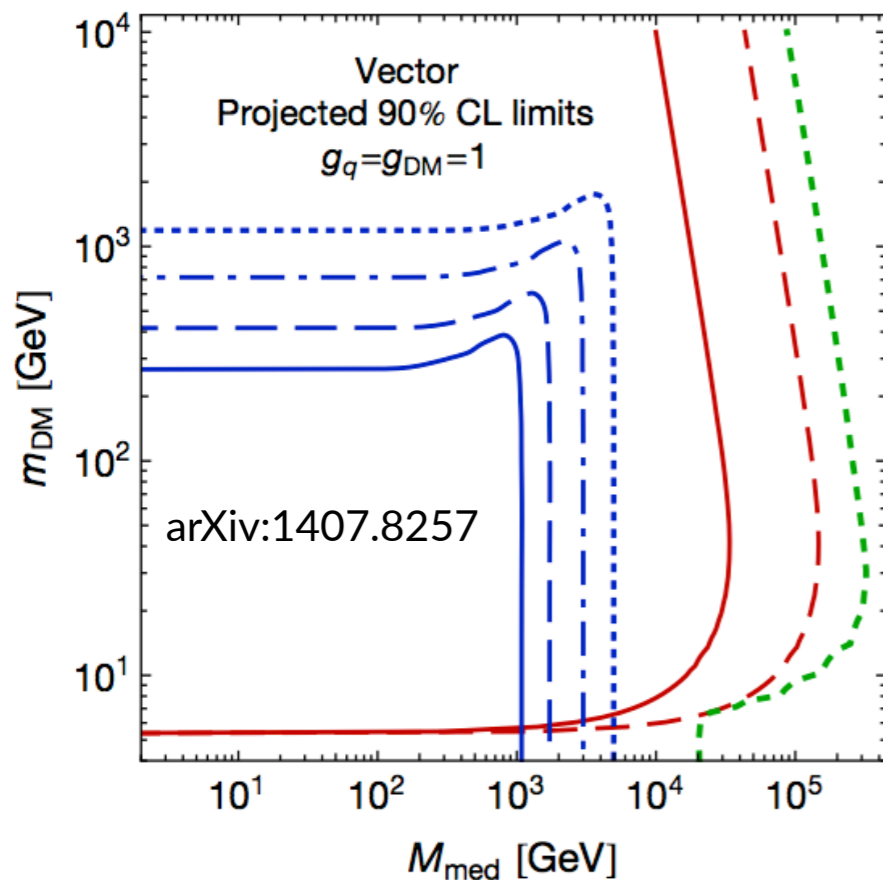
expected luminosity

now: 36 fb⁻¹
 end of 2018: 100 fb⁻¹
 end of 2023: 300 fb⁻¹
 HL-LHC (~2035): 3000 fb⁻¹

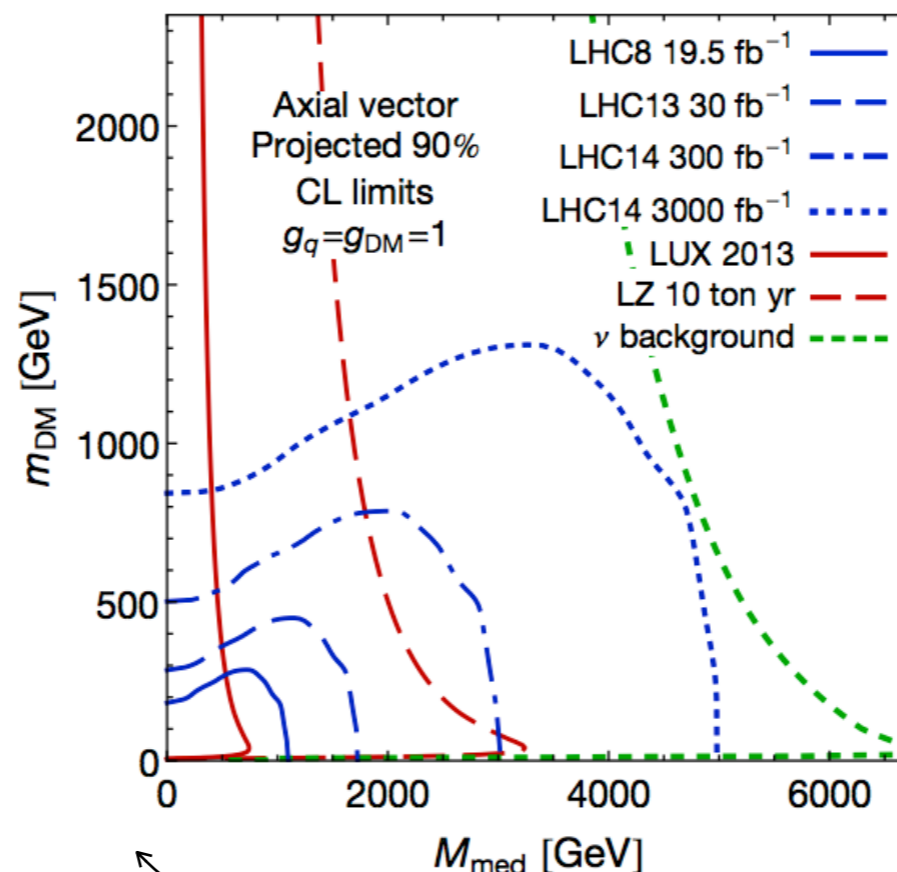
see also <https://indico.cern.ch/event/539266>

- balance between sensitivity to low-momentum signals (e.g. spin-zero) and robustness at very high energy
 - trigger & detector performance are crucial!
- explore lower-cross-section extensions of the SM (SUSY, long-lived particles...)

vector



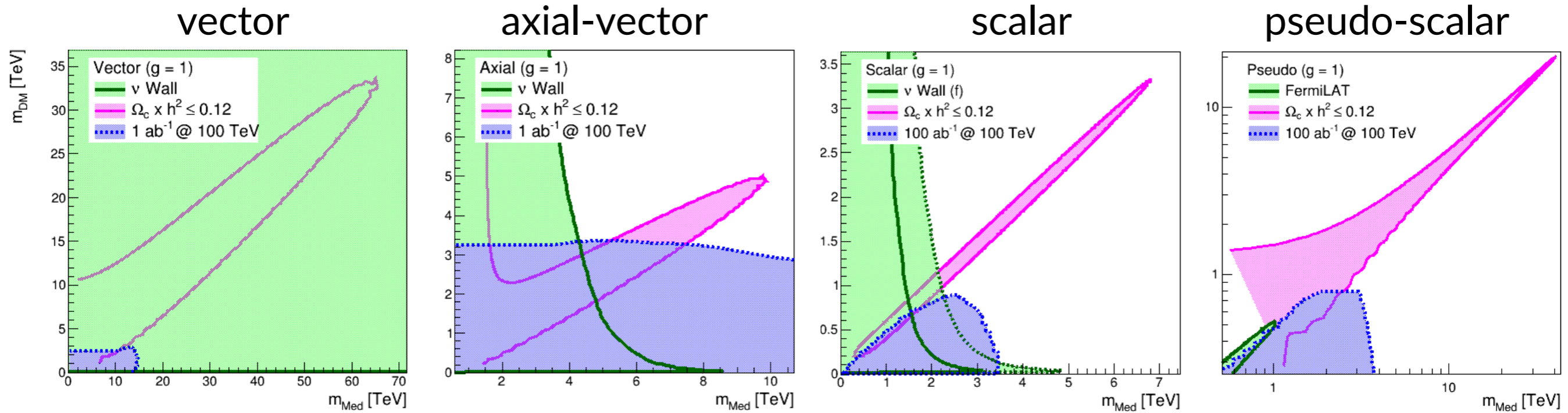
axial-vector



could extend the m_{DM} sensitivity up to 0.5 TeV in ~6 years (mind the couplings!)

region to the left of each curve is expected exclusion; LHC := "mono-jet"

BEYOND THE LHC



green: $x_{sec} \leq \text{neutrino bkg}$
blue: $1000 \text{ fb}^{-1} @ 100 \text{ TeV}$
red: compatible with measured relic density

(for some choice of the couplings)

a higher-energy circular collider may push sensitivity to the TeV scale

