

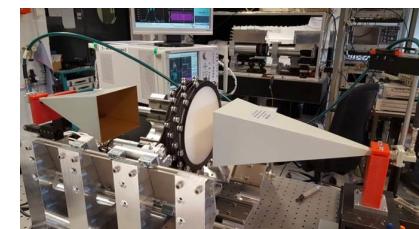
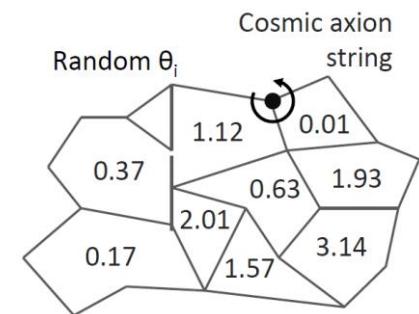
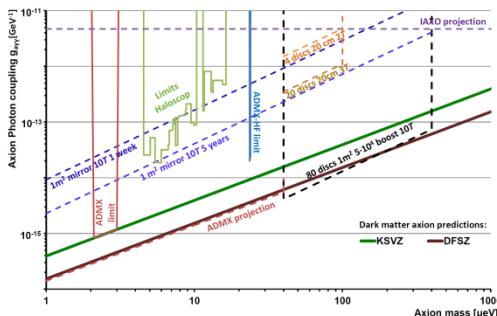
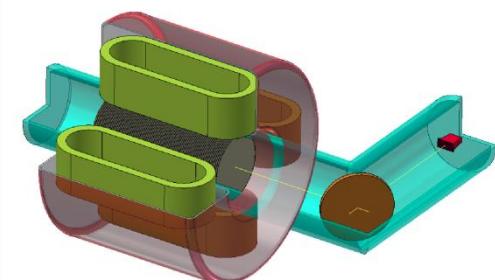


MADMAX: A new road to axion dark matter detection

Béla Majorovits
MPI für Physik, München, Germany
for the MADMAX interest group

OUTLINE:

- Axions as dark matter:
The post inflationary scenario
- Experimental idea
- First proof of principle measurements
- Outlook





MADMAX: A new road to axion dark matter detection

Béla Majorovits

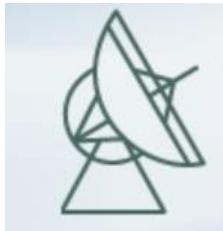
MPI für Physik, München, Germany

for the MADMAX interest group



CEA IRFU, Saclay

DESY Hamburg



MPI für Radioastronomie, Bonn



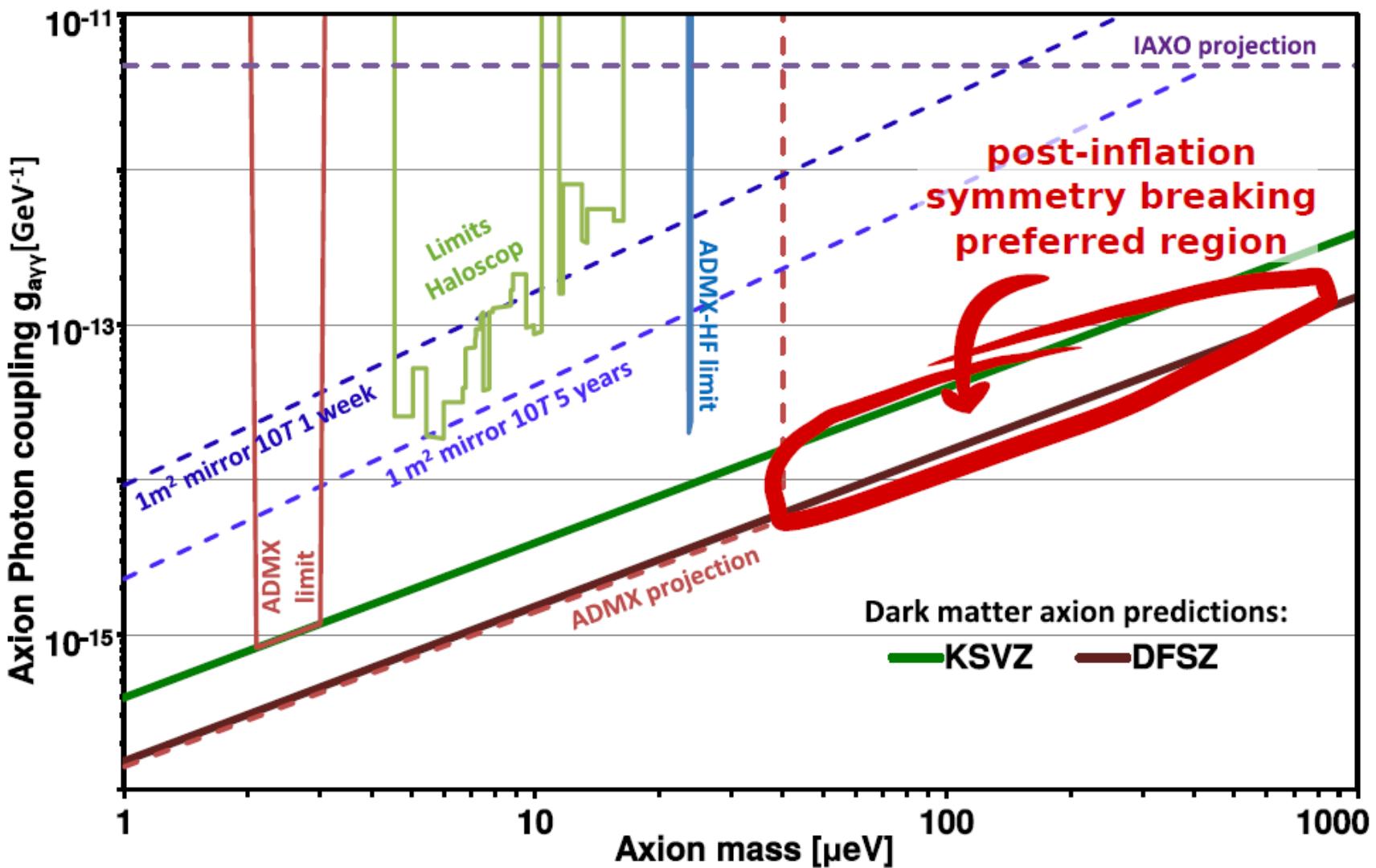
MPI für Physik, München

University of Hamburg

University of Zaragoza

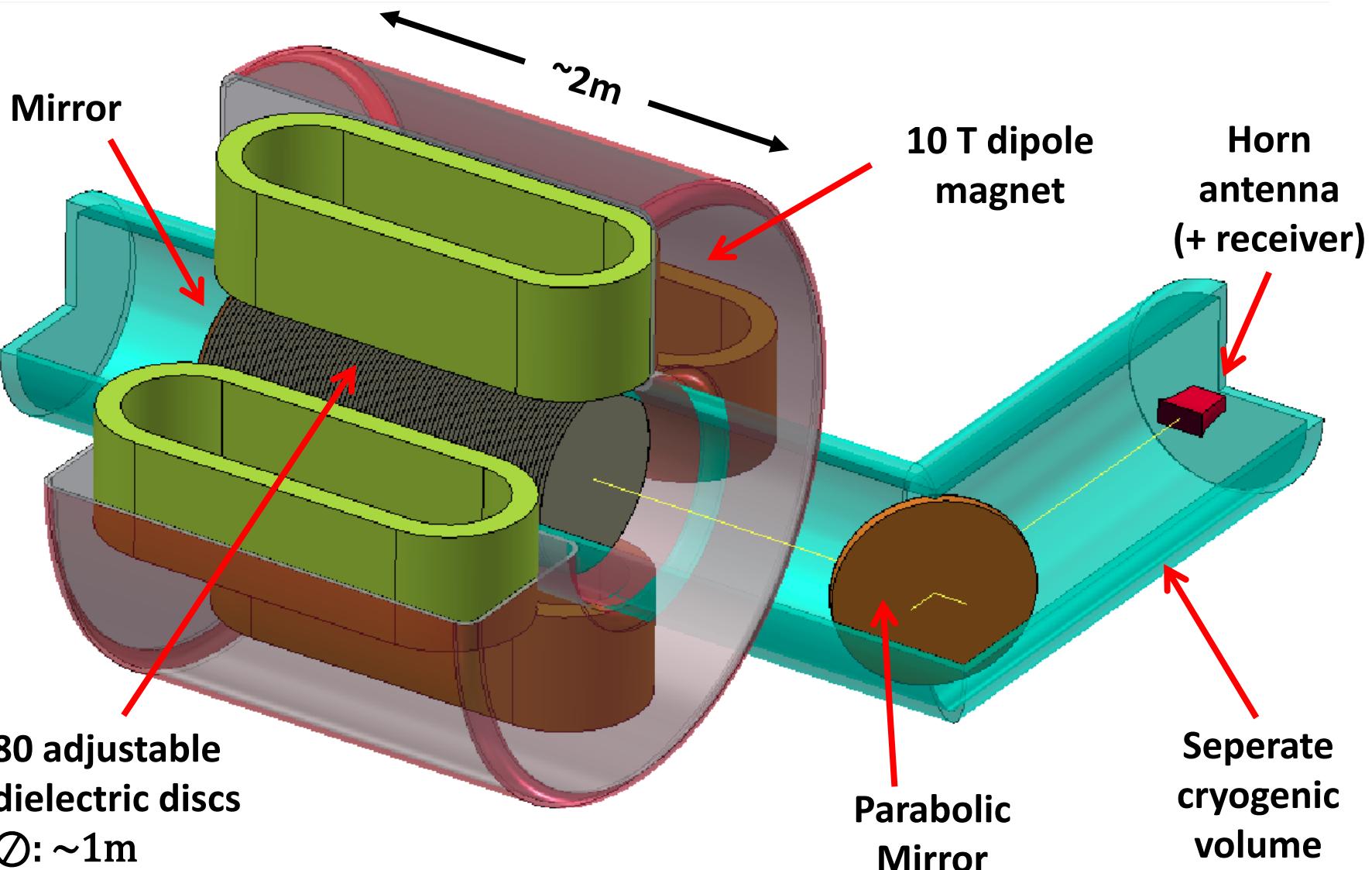


Axion DM: Scenario PQ – Inflation



MADMAX

MAgnitized Disc and Mirror Axion eXperiment



Experimental idea

Chose dielectric material:

- High dielectric constant ϵ (for large boost & conversion)
 - Low loss \rightarrow low $\tan \delta$ (reduce photon losses)
 - Stable
 - Cheap

→ Sapphire (Al_2O_3) @ 300K, 10 GHz:

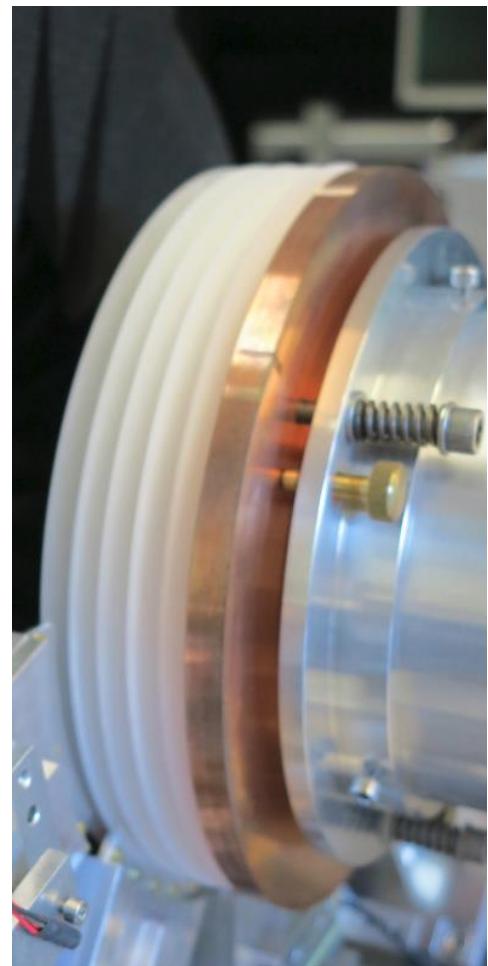
$$\epsilon \sim 10; \quad \tan \delta \sim \text{few} \cdot 10^{-5}$$

→ Lanthanide Aluminate (LaAlO_3) @ 77K

$$\epsilon \sim 24; \quad \tan \delta \sim 3 \cdot 10^{-5}$$

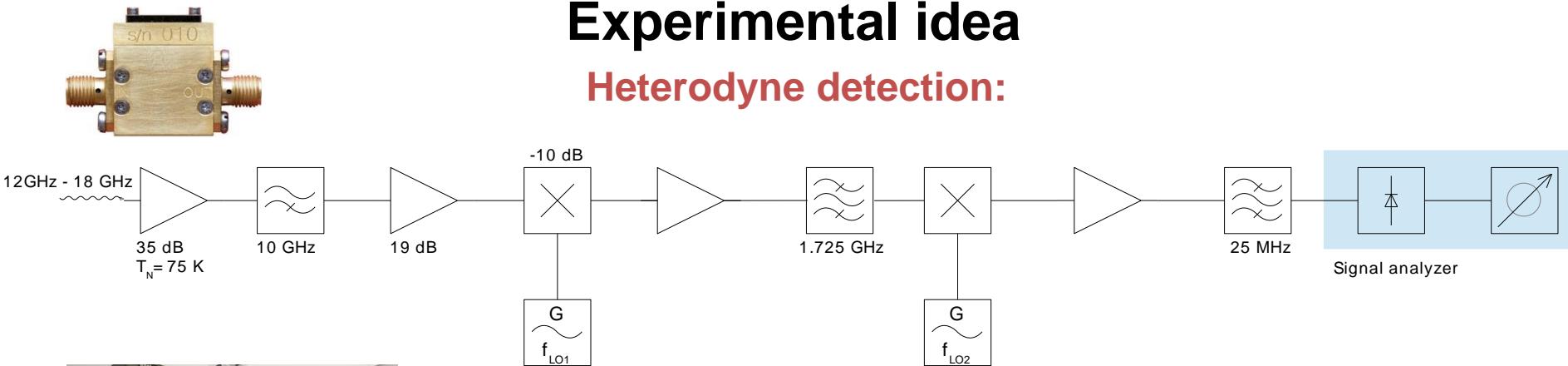
→ Titanium dioxide – Rutil (TiO_2)

$$\epsilon \sim 100; \quad \tan \delta \sim 0.001(?)$$



Experimental idea

Heterodyne detection:



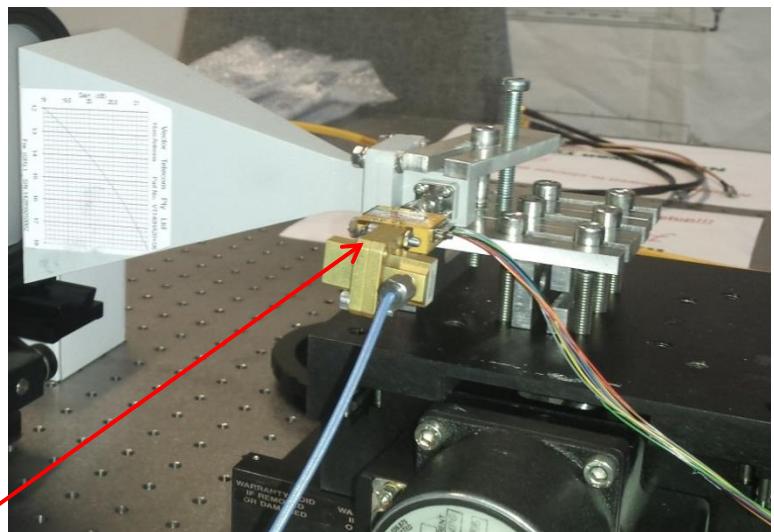
**Signal analyzer
(3 samplers)**

2. local oscillator

1. local oscillator

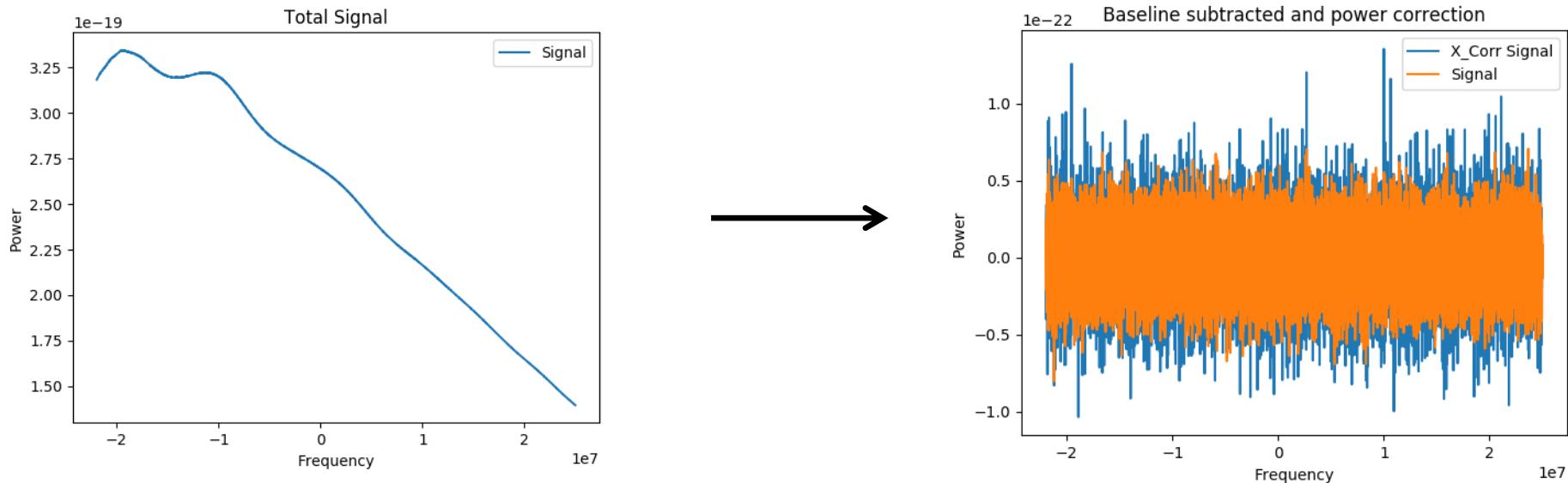
**Rubidium time standard
(oscillator and sampler
synchronization)**

**1. Amplifier +
high pass**



First measurements:

Low noise preamp:

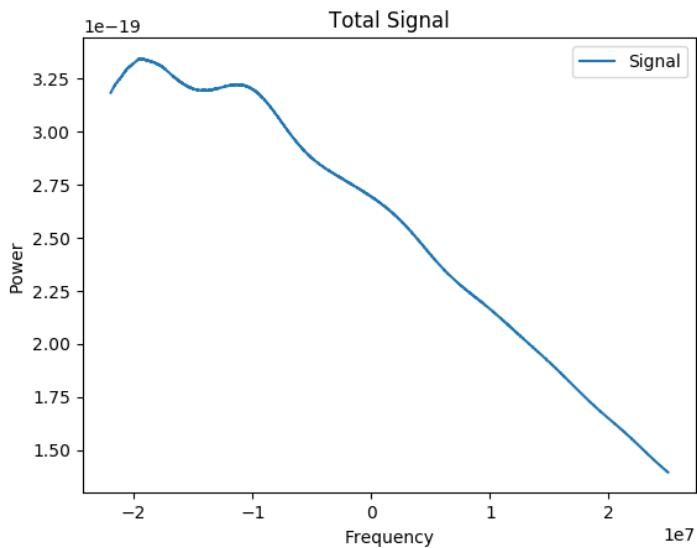


- Inject fake 18GHz axion signal with $1 \cdot 10^{-22}$ W power
 - Measurement for 28 hours (integrate signal):
Receiver at LHe temp.
- Cross correlation analysis (8kHz Lorentz shaped)
 - found $> 6\sigma$ signal successfully
- For 1 week measurement:
expect Sensitivity at the level of \sim few 10^{-23} W (t.b.c.)



First measurements:

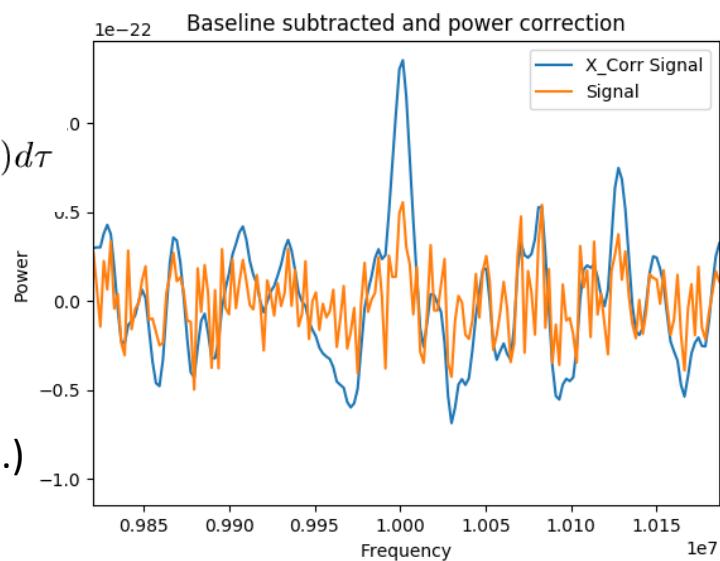
Low noise preamp:



$$X(\tau) = \int s(t) T(t + \tau) d\tau$$

→

s: Signal
T: Testfunction
(Lorentz, Gauss, ...)

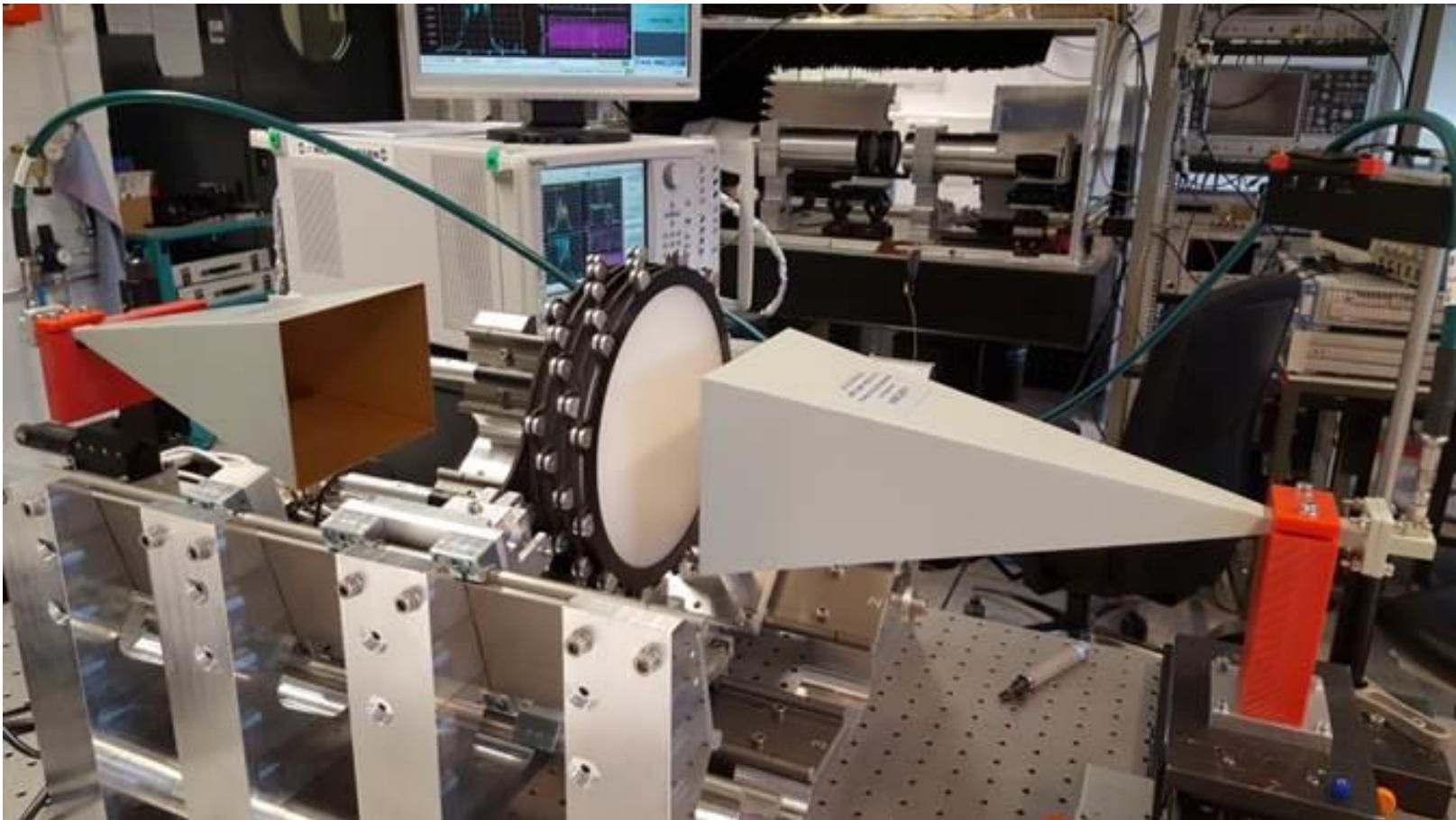


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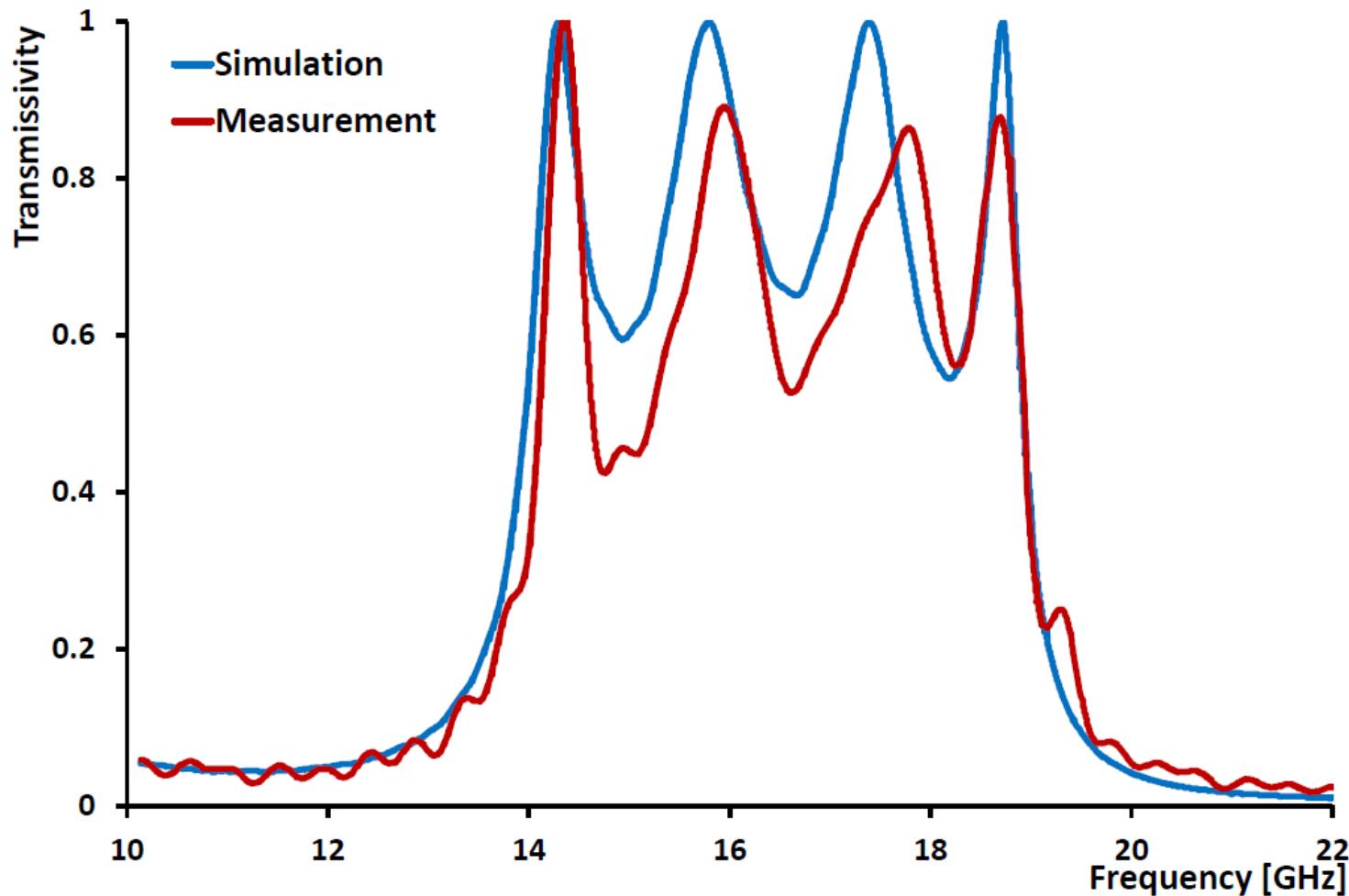
First prototype booster setup:

Transmissivity measurement:



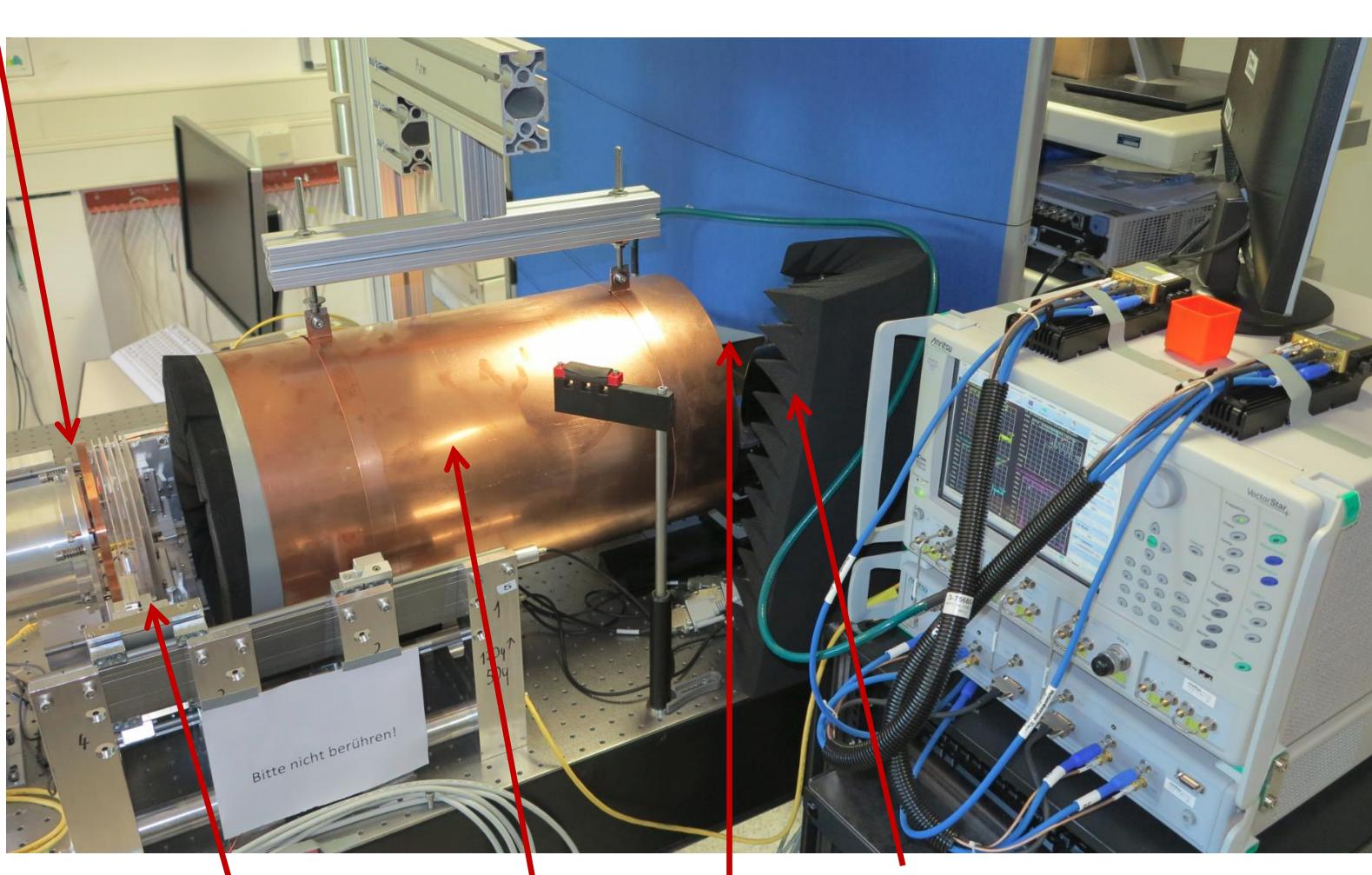
First prototype booster setup:

Transmissivity measurement:



Removable
copper mirror

First prototype booster setup: Reflectivity measurement:



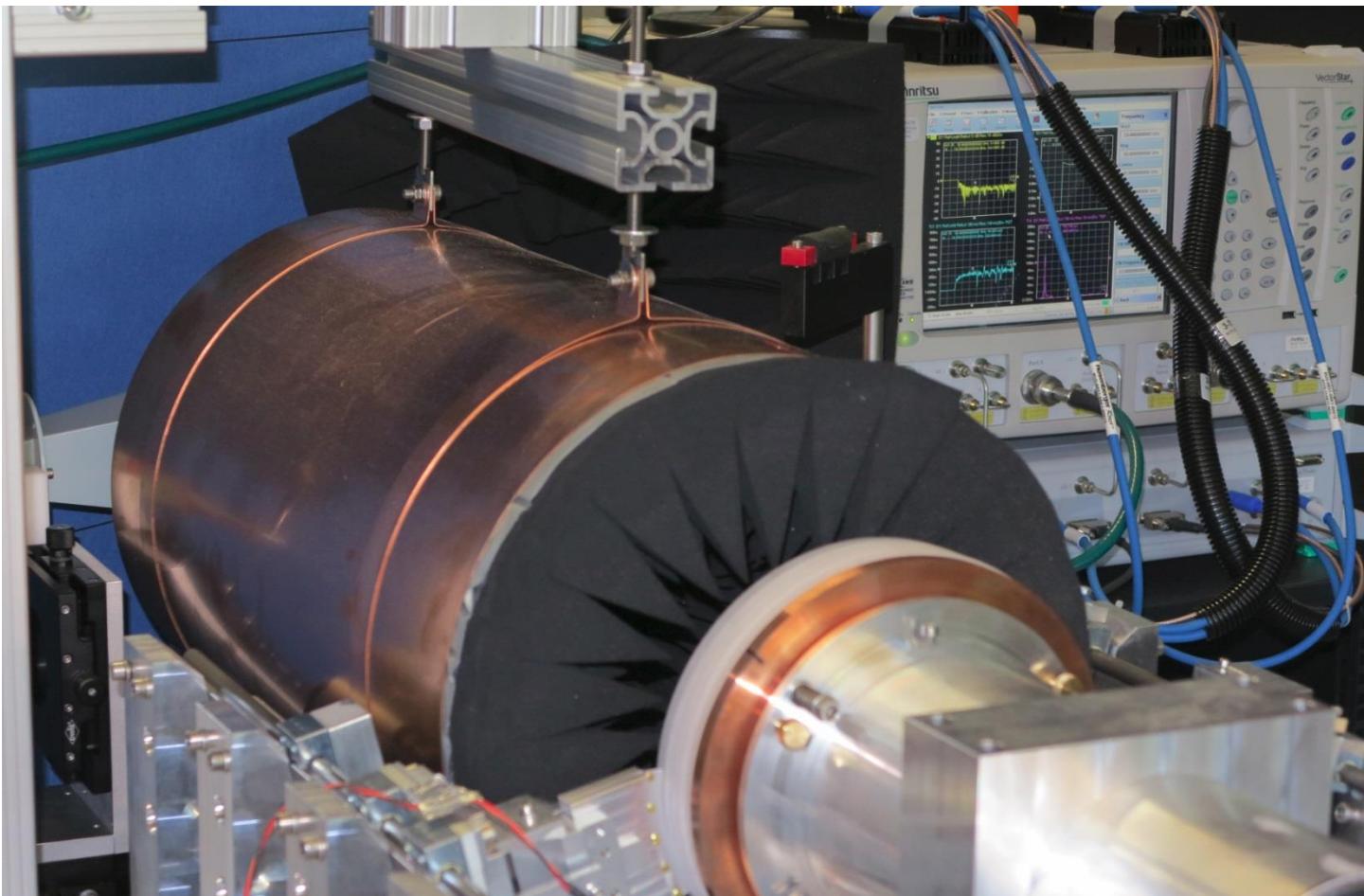
Dielectric discs
(Saphire)

„Wave
guide“
antenna

Mirror

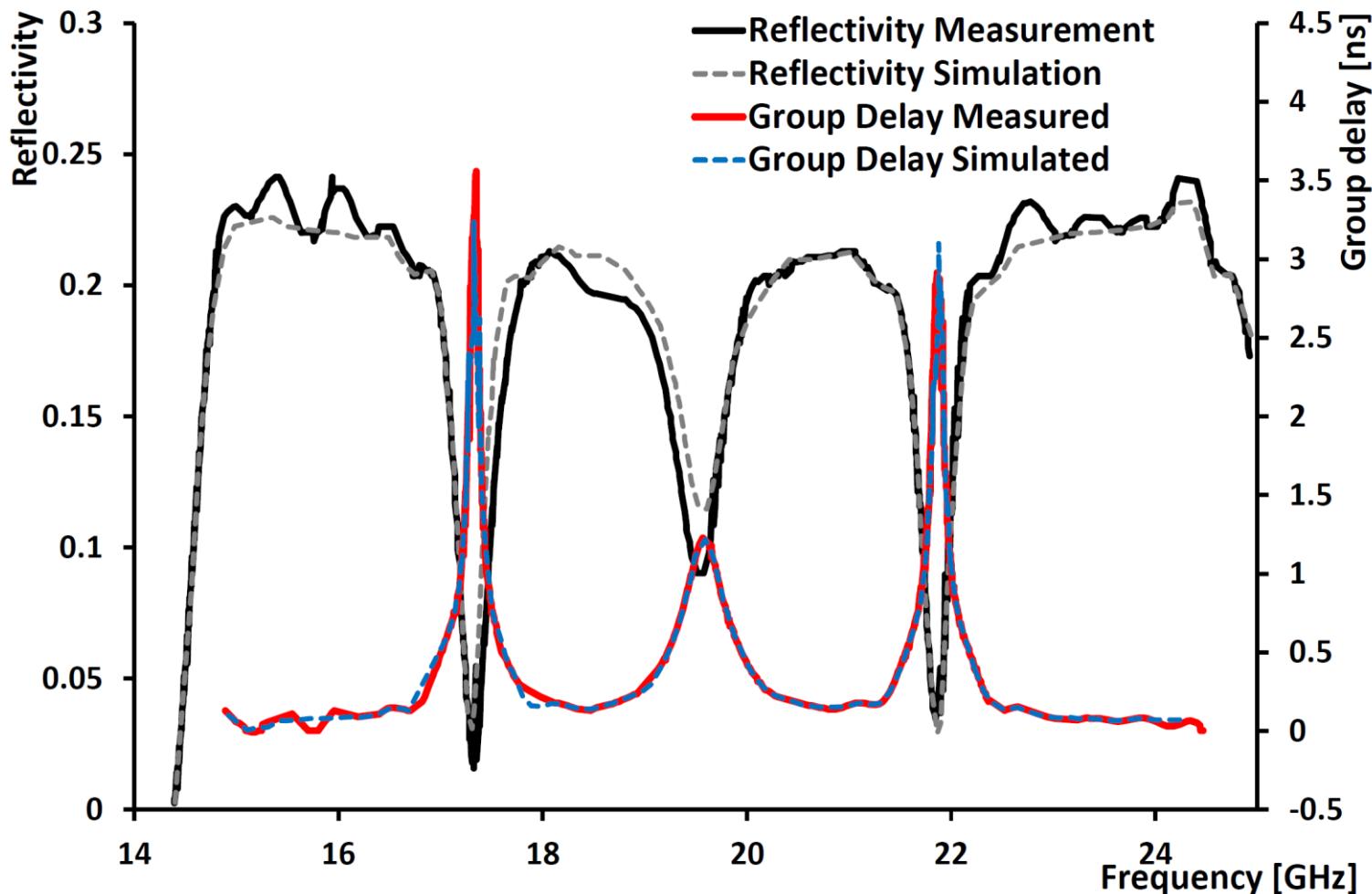
First prototype booster setup:

Reflectivity measurement:



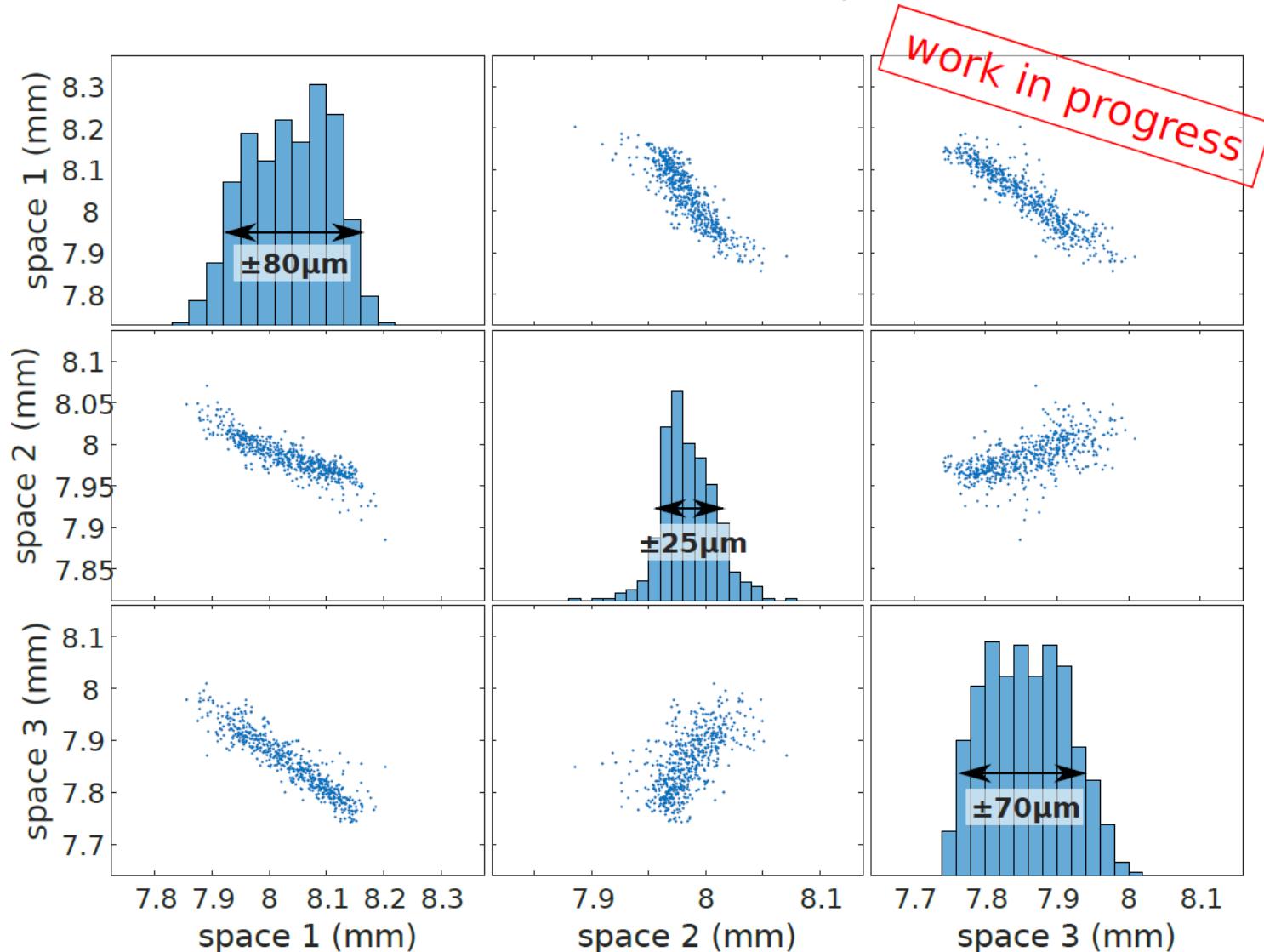
First prototype booster setup:

Reflectivity measurement:



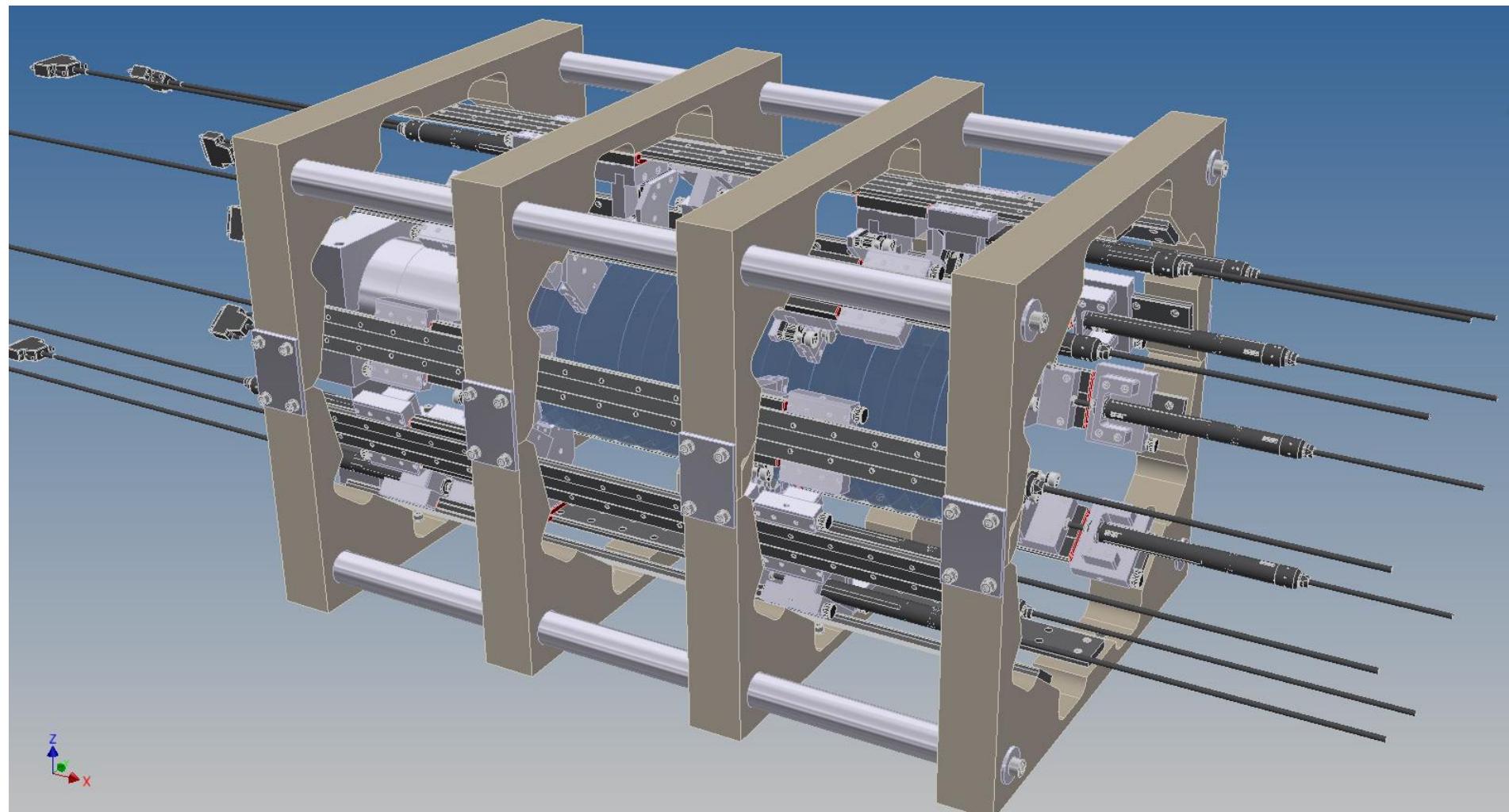
First prototype booster setup:

Position reproducibility:



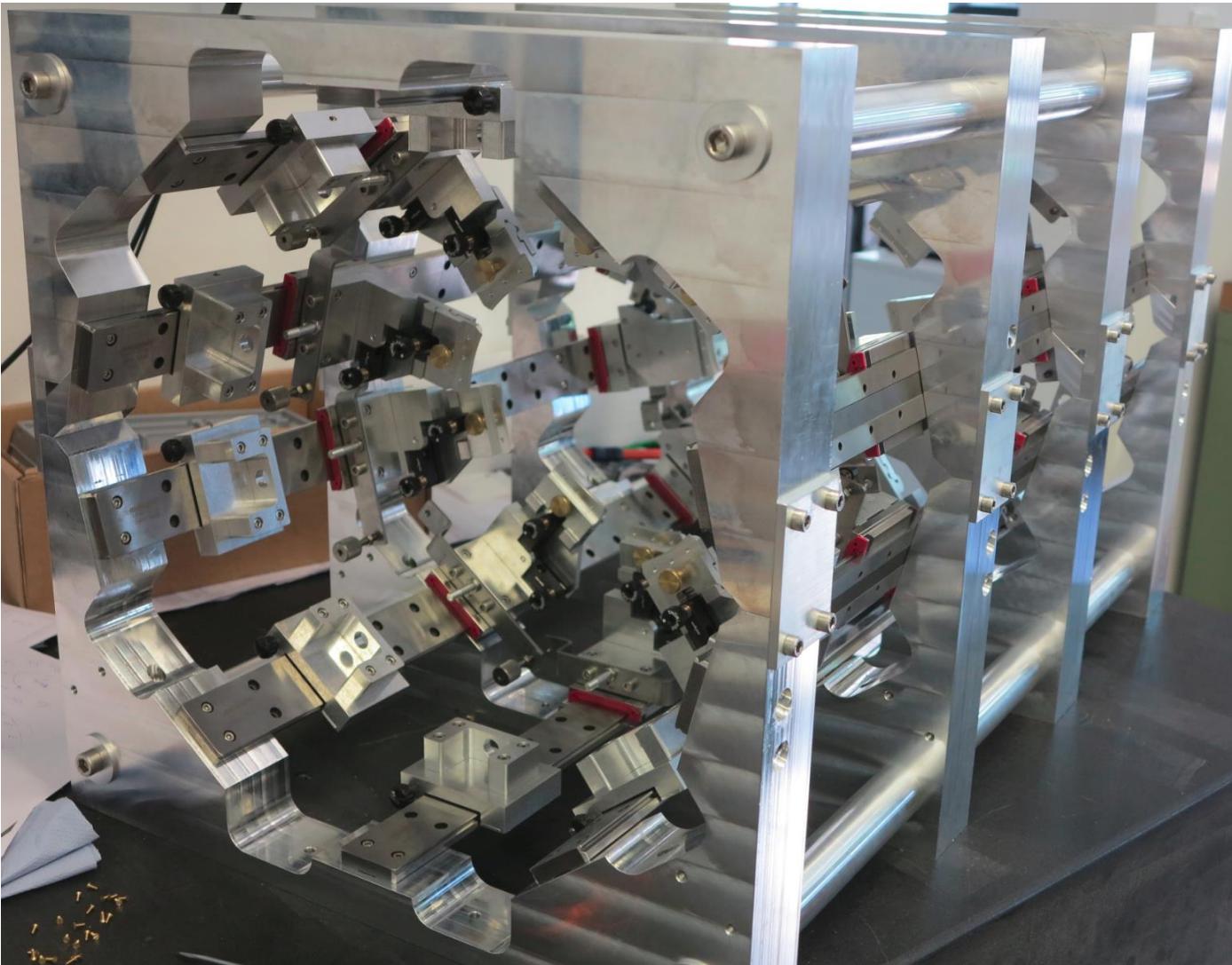
Prototype booster Upgrade:

20 disc setup:



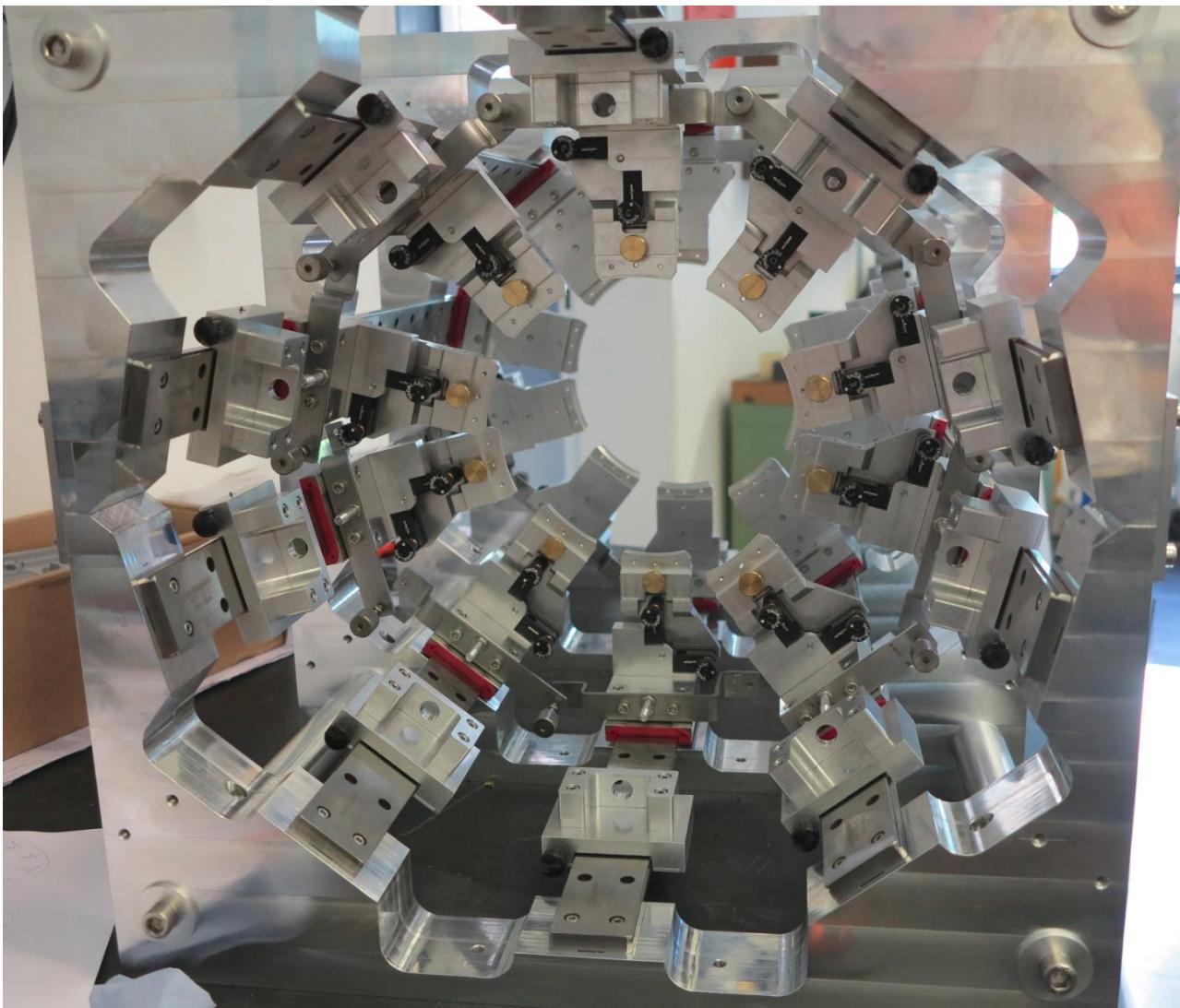
Prototype booster Upgrade:

20 disc setup:



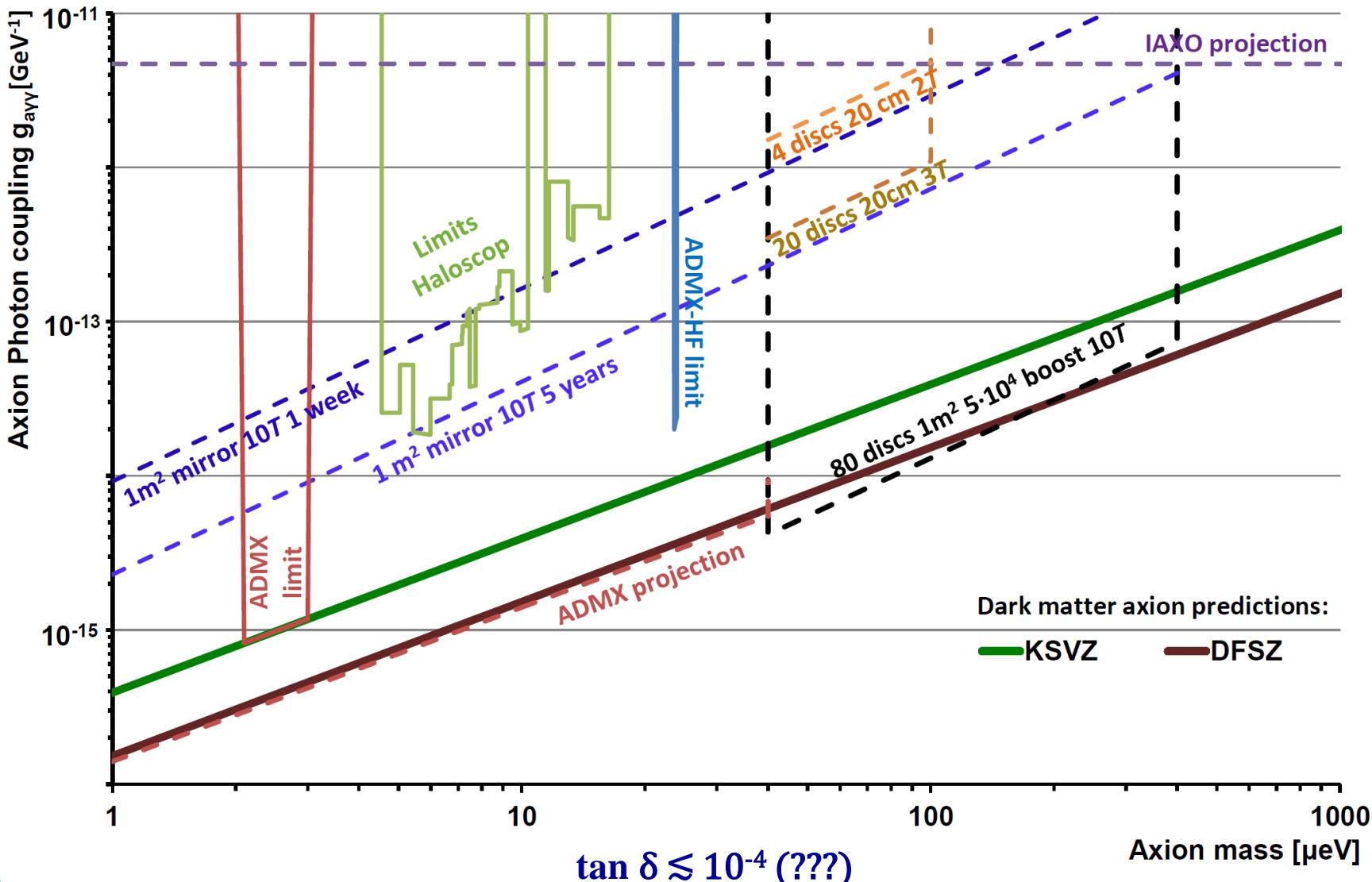
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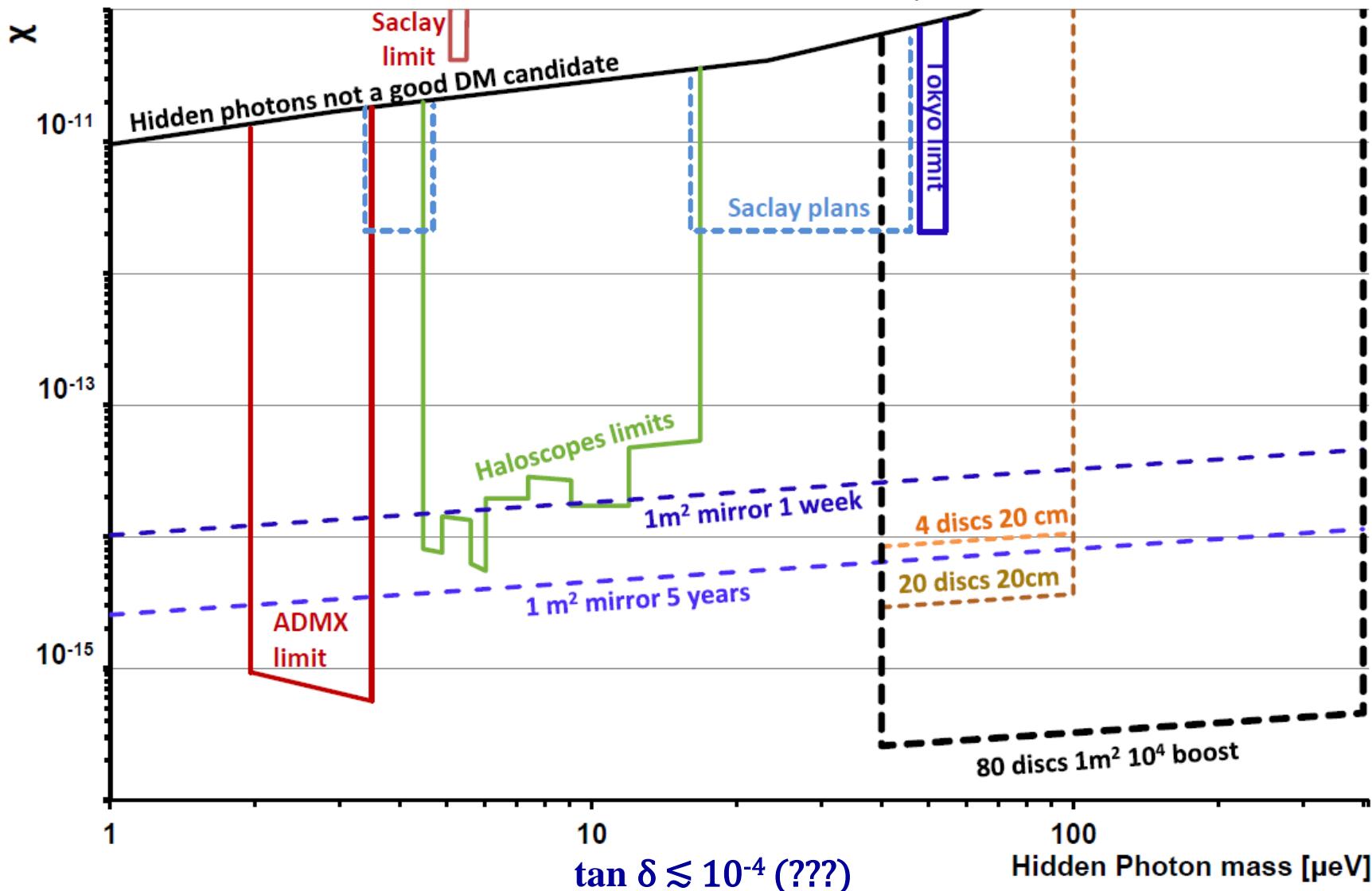
OUTLOOK:

Sensitivity for QCD dark matter axions with $A=1\text{m}^2$, $B_{||}=10\text{T}$, $T_{\text{sys}}=8\text{K}$, $\beta^2=5\cdot10^4$



OUTLOOK:

Sensitivity for hidden photons with $A=1\text{m}^2$, $T_{\text{sys}}=8\text{K}$, $\beta^2=5\cdot10^4$



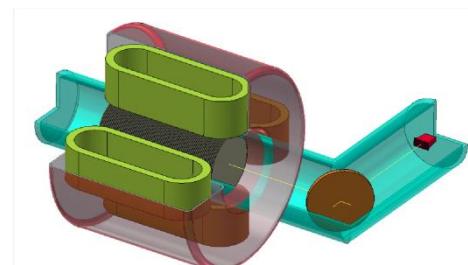


OUTLOOK:

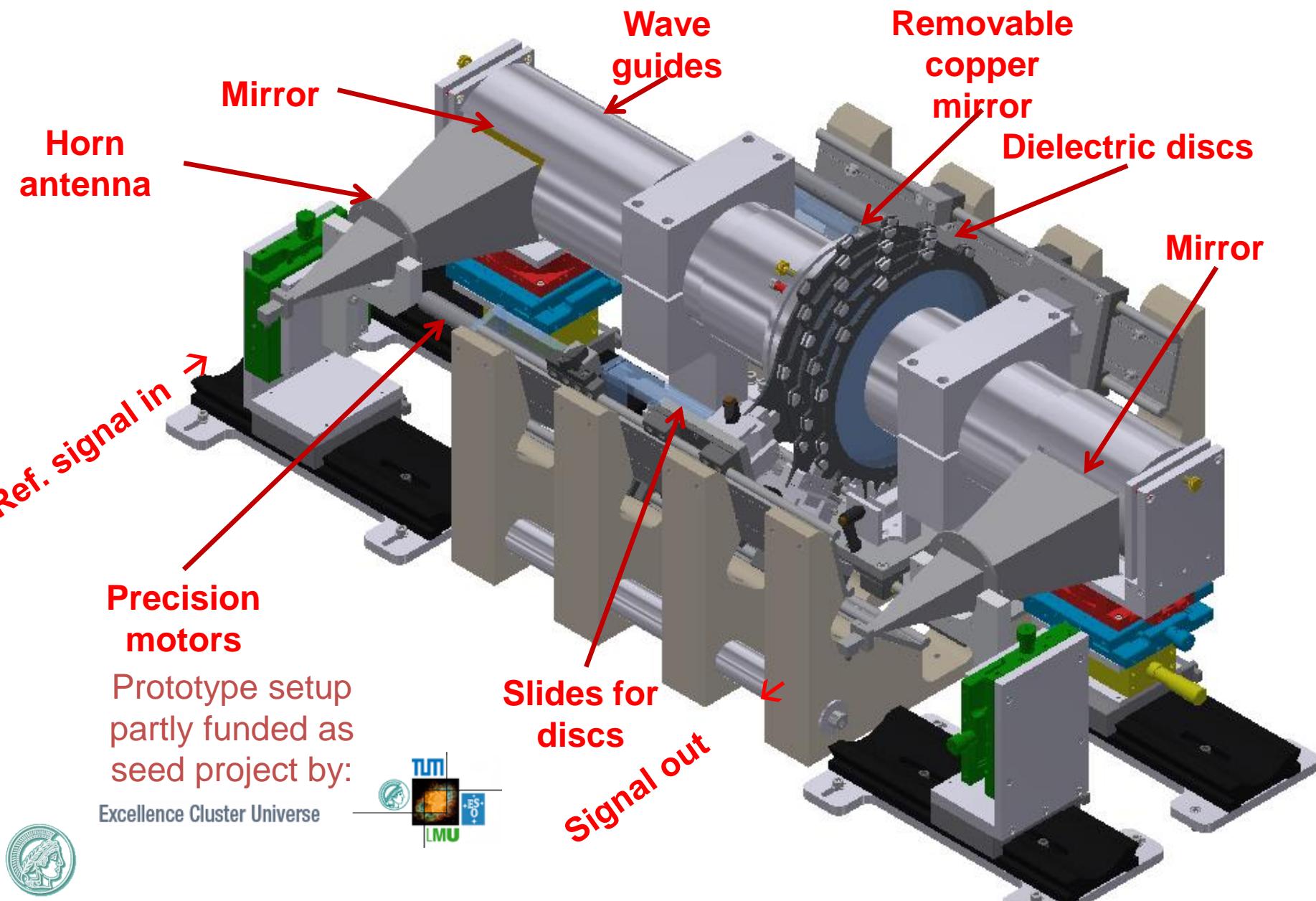
- Sign MoU → officially establish collaboration
- Magnet innovation partnership with (2018)
Bilfinger Bacock Noell
CEA IRFU



- Desin study for booster realization (2018)
- Build prototype 3-4 T magnet & 20 discs 30cm diameter booster (2021?)
→ First QCD axions results 2021
- Build full scale experiment (>2022)
Considering DESY as site



First prototype booster setup:

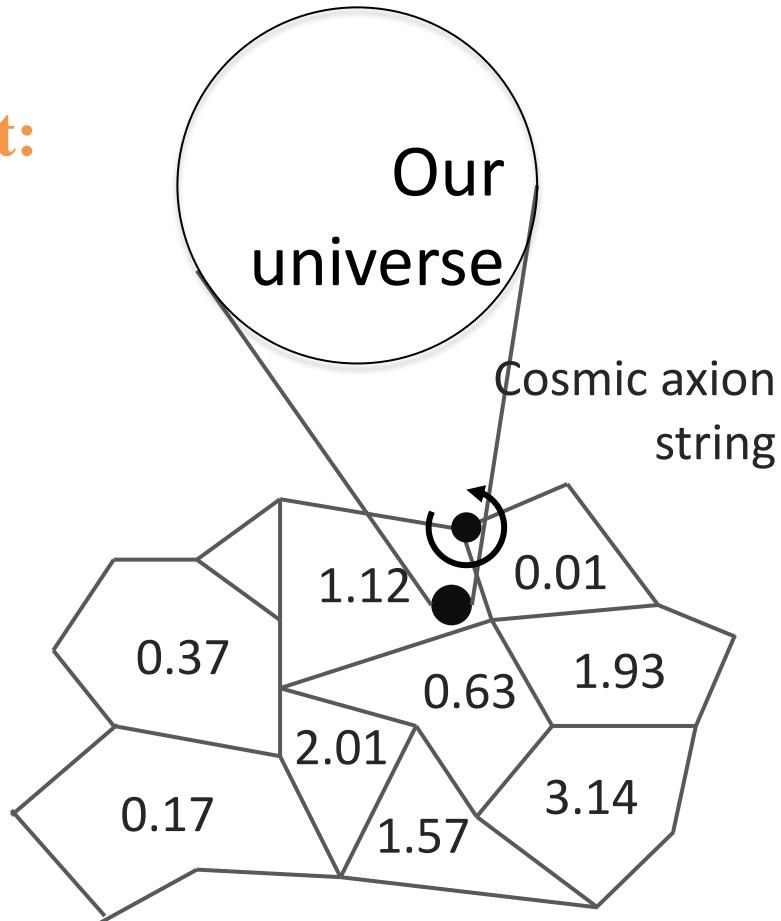


Axion DM: Scenario PQ – Inflation (Pre Inflationary PQ breaking)

Scenario II: PQ symmetry breaking first:

- θ_i has a single random value which determines the dark matter density
- No “topological defects”

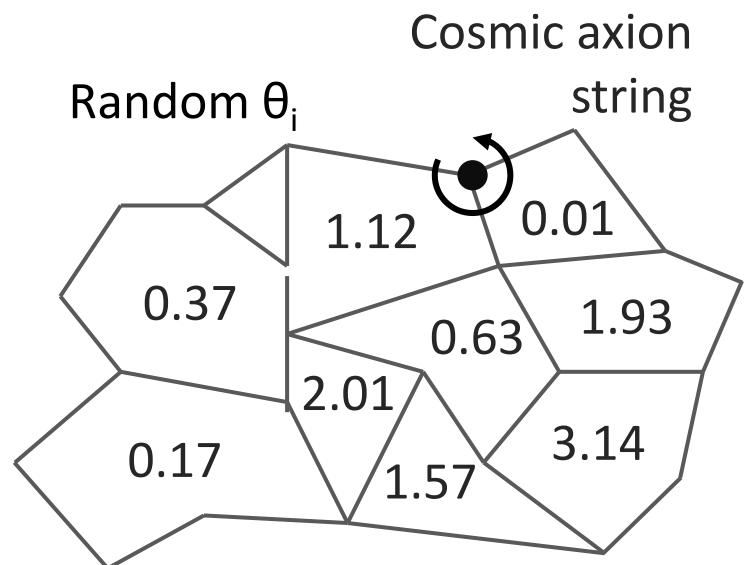
QCD dark matter axions can have any mass $\lesssim 1\text{meV}$!



Axion DM: Scenario Inflation – PQ (Post Inflationary PQ breaking)

Scenario Inflation first:

- PQ broken after inflation
- θ_i has random values in every casual region, with the dark matter density determined by the average
- Topological defects such as strings and domain walls exist in the early universe
 - decay leads to axion production
 - influence axion density



Predicted axion mass $\sim 100 \mu\text{eV}$



First measurements:

Low noise preamp:



InP HEMT preamplifier from LowNoiseFactory

Frequency range: 6-20 GHz

detector noise: $T \sim 7K$ (measured, quick and dirty)

$T_{ds} \sim 6K$ (data sheet)

